

# GEANT4 simulation package for the GRAPES-3 muon telescope

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**On behalf of GRAPES-3 collaboration**

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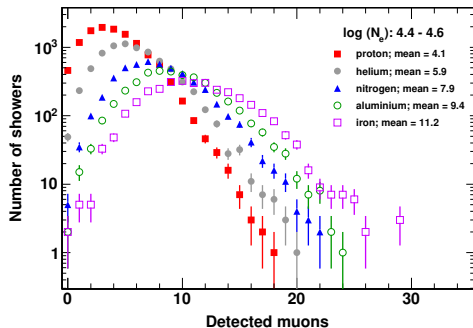
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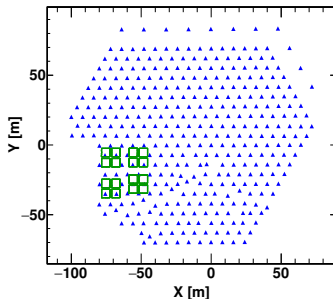
Summary

## Motivation

- Muon component of EAS is sensitive to the nature of PCR.
- Plays a vital role in extracting the nuclear composition of the PCR and separating the  $\gamma$ -rays from the background of the cosmic ray.
- To achieve the desired precision, a thorough understanding of the response of EAS particles in the GRAPES-3 muon telescope (G3MT) is necessary.
- We developed a simulation framework for the G3MT using the GEANT4 toolkit (version 4.10.04) and studied its performance using CORSIKA (version 7.6900) based on the QGSJET-II-04 and FLUKA.



## GRAPES-3 experiment



- 400 plastic scintillation detectors ( $1 \text{ m}^2$  each) cover an area of  $25,000 \text{ m}^2$ .
- Large area ( $560 \text{ m}^2$ ) muon telescope.
- Fraction of detector area covered is 2%.
- Energy range: 1 TeV – 10 PeV.

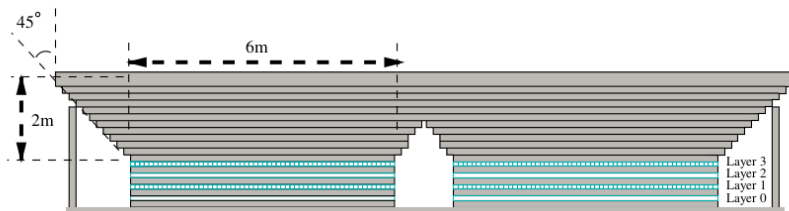
### Location:

- Ooty, south India
- $11.4^\circ \text{ N}$ ,  $76.7^\circ \text{ E}$
- 2200 m a.s.l.

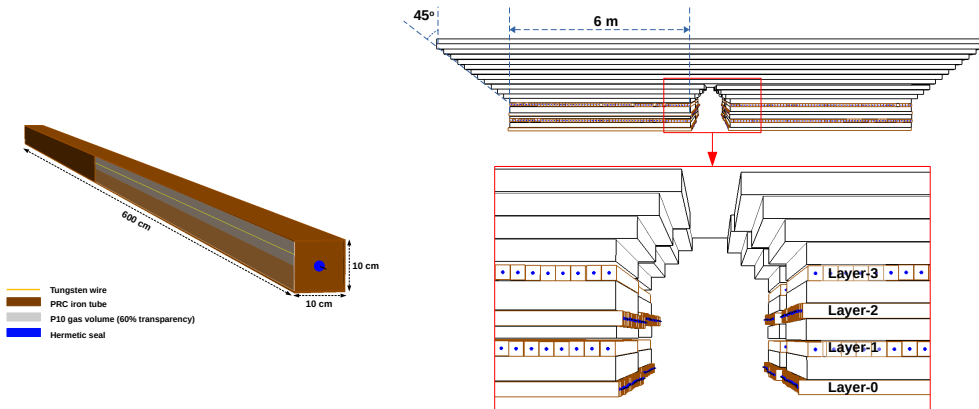


## GRAPES-3 muon telescope

- 16 independent modules (each  $35 \text{ m}^2$ ).
- Total effective area:  $560 \text{ m}^2$ .
- Energy threshold:  $1 \text{ GeV} \times \sec \theta$ .
- Detector unit: proportional counters (PRCs, total 3712) filled with P10 gas.
- Four orthogonal layers of PRCs.
- Concrete absorber ( $550 \text{ g cm}^{-2}$ ): shields electromagnetic & hadronic components.

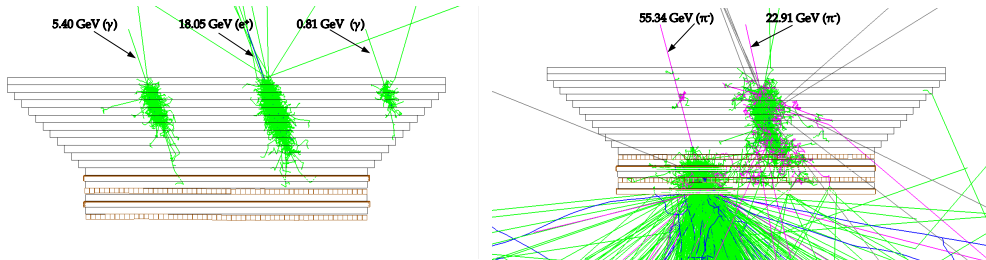


## Geant-4 simulation: Geometric reconstruction of GMT



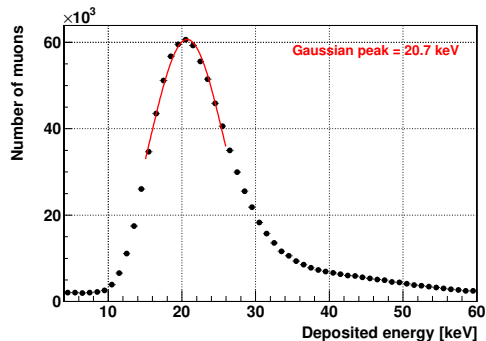
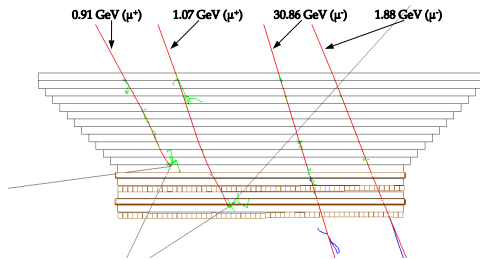
## Geant-4 simulation: EAS particles response

- CORSIKA (version 7.6900) based on the QGSJET-II-04 and FLUKA.



- Electromagnetic (EM) and low energy hadronic components get absorbed.
- High energy hadrons generate an EAS in the absorber and form a cluster of PRCs hit.

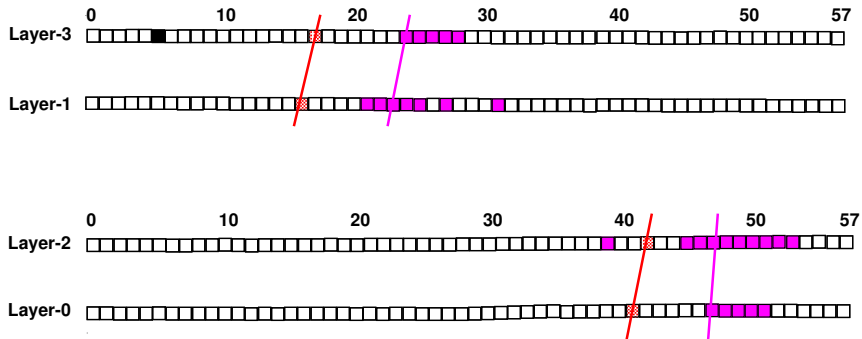
## Geant-4 simulation: EAS particles response



- Muons above  $1 \text{ GeV} \times \sec \theta$  threshold make a clear passage through the module.
- The energy deposition by single muons in PRC peaks at 20.7 keV, which is consistent with the experimental value of  $\sim 20 \text{ keV}$ .

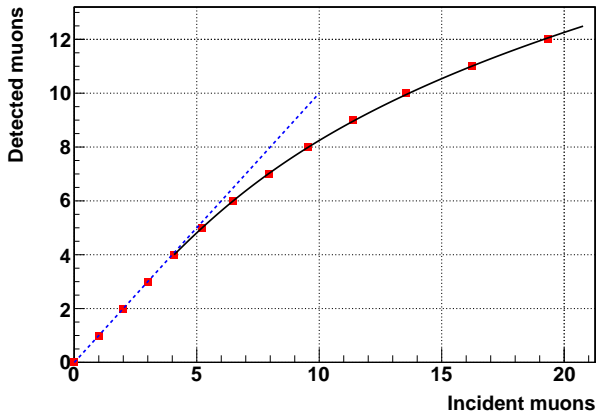


## Reconstruction of muon tracks



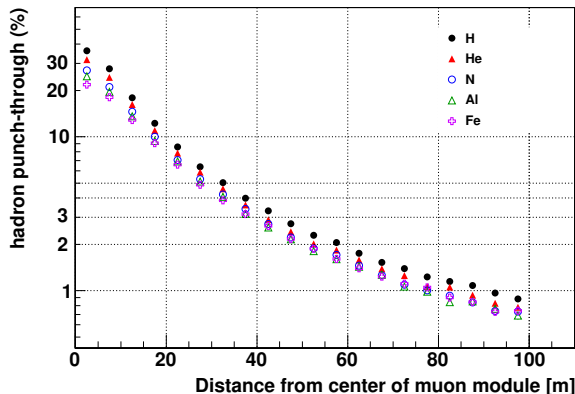
## Muon saturation

- The mean value of incident muons is plotted against the given number of detected muons.
- Curve is modeled with the 3<sup>rd</sup> polynomial and used to correct for the saturation.

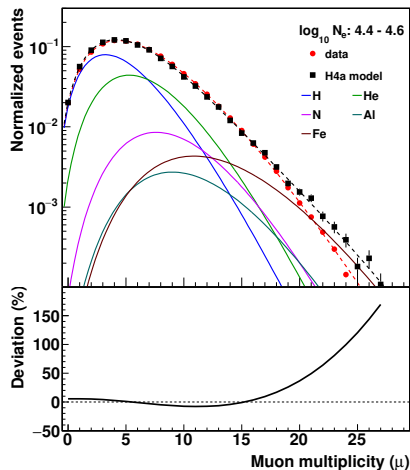
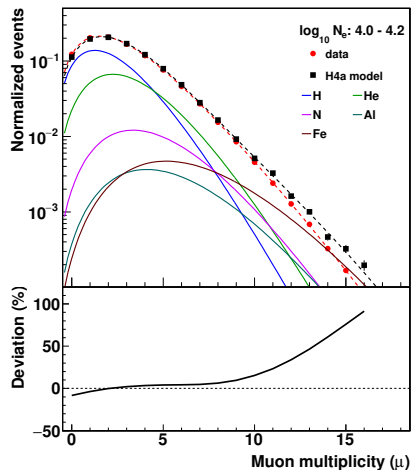


## Hadron punch-through

- Quality cuts:
  - Successful reconstruction of direction and NKG parameters.
  - Reconstructed cores within 50 m from the center of array.
  - Shower age (s):  $[0.2, 1.8]$ .
  - zenith angle :  $[0^\circ, 25^\circ]$ .
  - Shower size ( $N_e$ )  $\geq 10^{4.0}$ .
- event-by-event basis, fraction of number of tracks by hadron is calculated.
- hadron punch-through per EAS is calculated, followed by averaging for all the 16 modules.
- $\sim 10\%$  at 20 m and reduces to  $\sim 2\%$  at 60 m.



## Comparison of observed MMD with the H4a composition model

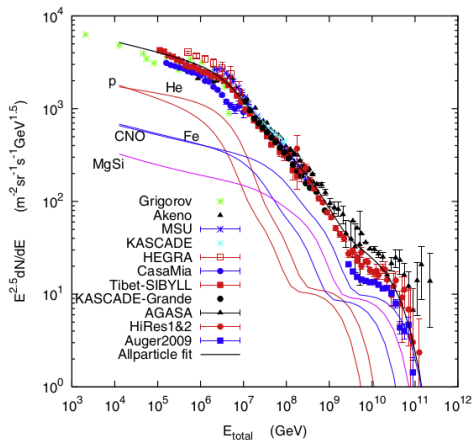


## Summary

- We developed a simulation framework for the G3MT using the GEANT4.
- Performance is studied using CORSIKA based on the QGSJET-II-04 and FLUKA.
- Electromagnetic (EM) and low energy hadronic components get absorbed while high energy hadrons generate an EAS in the absorber and form a cluster of PRCs hit.
- Muons above  $1 \text{ GeV} \times \sec \theta$  threshold make a clear passage through the module.
- The energy deposition by single muons in PRC peaks at 20.7 keV, which is consistent with the experimental value.
- Hadron punch-through is studied and it reduce to less than 2% at 60 m.
- Comparative study shows a reasonably good agreement at lower muon multiplicity and a significant deviation at higher muon multiplicity indicates that the relative composition of heavier nuclei (especially iron) in the H4a model is higher than that expected from the observed data.

# Thank you

## Backup slides: H4A Model



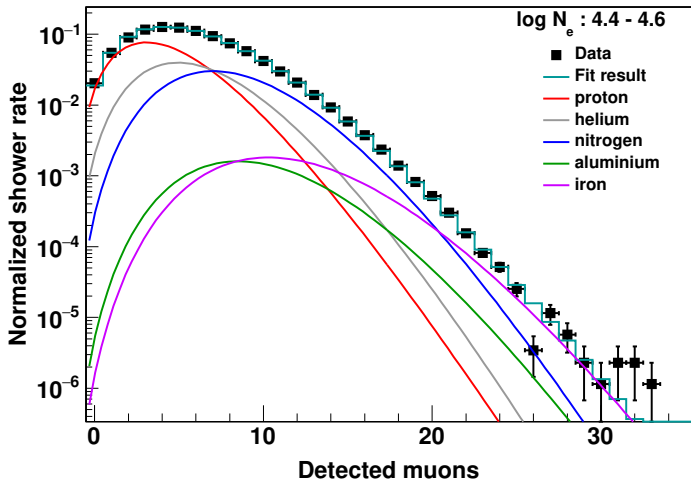
	p	He	CNO	Mg-Si	Fe
Pop. 1:	7860	3550	2200	1430	2120
$R_c = 4$ PV	1.66	1.58	1.63	1.67	1.63
Pop. 2:	20	20	13.4	13.4	13.4
$R_c = 30$ PV	1.4	1.4	1.4	1.4	1.4
Pop. 3:	1.7	1.7	1.14	1.14	1.14
$R_c = 2$ EV	1.4	1.4	1.4	1.4	1.4
Pop. 3(*):	200	0.0	0.0	0.0	0.0
$R_c = 60$ EV	1.6				

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801–806

## Backup slides: Stats and Composition

$\log N_e$	Energy ( $E_H$ )	Number of EAS		Relative composition (%)					Mean of MMD	
		Range	Range [TeV]	Data	H4a	H	He	N	Al	Fe
4.0 - 4.2	35 - 55	4091460	174510	54.7	32.3	7.1	2.4	3.5	2.7	2.8
4.2 - 4.4	55 - 84	2080850	106219	52.3	33.2	7.4	2.7	4.4	4.1	4.1
4.4 - 4.6	84 - 127	1045140	61109	51.1	33.6	7.6	2.7	4.9	6.1	6.0
4.6 - 4.8	127 - 193	527207	35181	50.2	34.4	7.6	2.9	4.9	8.7	8.6
4.8 - 5.0	193 - 294	268389	19729	48.8	35.3	8.1	2.9	4.9	12.3	12.2
5.0 - 5.2	294 - 447	137934	55972	46.4	35.8	8.7	3.3	5.7	17.4	17.7
5.2 - 5.4	447 - 680	71189	31124	43.9	38.4	8.7	3.5	5.4	24.6	25.2
5.4 - 5.6	680 - 1035	37122	17073	43.4	37.7	9.5	3.8	5.6	34.5	35.9

## Backup slides: MMD unfolded



H: 45.4%, He: 33.6%, N: 15.5%, Al: 4.2% and Fe: 1.3%