

# NuFact 2011

## Optimized Neutrino Factory for small and large $\theta_{13}$

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# Oscillation parameters

Parameter	Best fit $\pm 1\sigma$	$3\sigma$ range
$\Delta m_{21}^2$ [ $10^{-5}$ eV <sup>2</sup> ]	$7.59^{+0.20}_{-0.18}$	7.09 - 8.19
$ \Delta m_{31}^2 $ [ $10^{-3}$ eV <sup>2</sup> ]	$2.45 \pm 0.09$	2.18 - 2.73
$\sin^2 \theta_{12}$	$0.312^{+0.017}_{-0.015}$	0.27 - 0.36
$\sin^2 \theta_{23}$	$0.51 \pm 0.06$	0.39 - 0.64
$\sin^2 \theta_{13}$	$0.010^{+0.009}_{-0.006}$	$\leq 0.035$

T. Schwetz, M.A. Tortola, J.W.F. Valle, arXiv:1103.0734v2

Best-fit for normal ordering with new reactor fluxes

Including latest solar (SK II+III), atmospheric (SK I+II+III), KamLAND & MINOS (before June, 2011) data. New reactor fluxes & SBL data

# New T2K and MINOS results

- In June, 2011, new exciting results announced by the T2K and MINOS long-baseline experiments

- T2K experiment observed 6 electron-like events with an estimated background of 1.5 events in the Super-K

*K. Abe et al., [T2K Collaboration], arXiv:1106.2822*

- It indicates towards  $\theta_{13}$  driven  $\nu_{\mu} \rightarrow \nu_e$  appearance signal, rejecting  $\theta_{13} = 0$  at the level of  $2.5 \sigma$

- The MINOS experiment has observed 62 electron-like events with an estimated background of 49 events

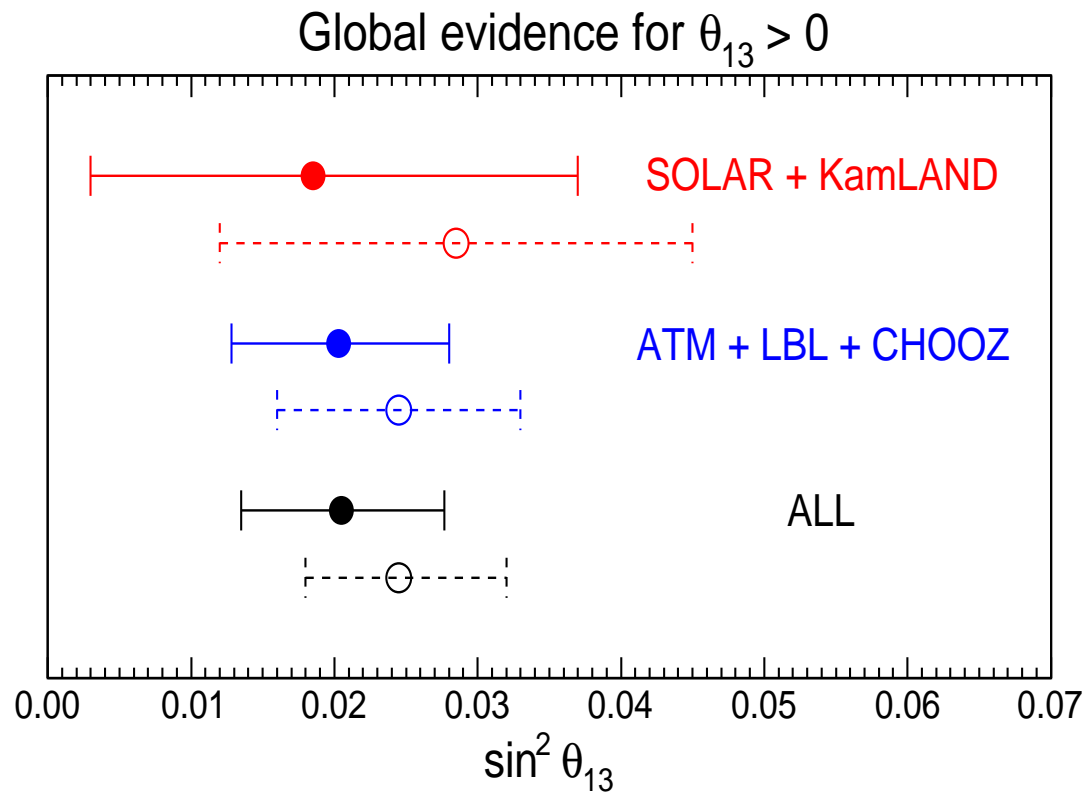
*P. Adamson et al., [MINOS Collaboration], arXiv:1108.0015*

- This favors a non-zero  $\theta_{13}$  at  $1.5 \sigma$  confidence level

- Recent global fit suggests  $> 3 \sigma$  evidence for non-zero  $\theta_{13}$

*G.L. Fogli et al., , arXiv:1106.6028*

# Global Picture of $\theta_{13}$



G.L. Fogli *et al.*, , arXiv:1106.6028

$$\sin^2 2\theta_{13} = \begin{cases} 0.082 \pm 0.028, & \text{old reactor fluxes} \\ 0.098 \pm 0.028, & \text{new reactor fluxes} \end{cases} \quad (1\sigma)$$

# Big Issues in Neutrino Mixing

- The sign of  $\Delta m_{31}^2$  ( $m_3^2 - m_1^2$ ) is not known

It can be normal –  $\Delta m_{31}^2 > 0$

or

inverted hierarchical –  $\Delta m_{31}^2 < 0$

- The value of  $\theta_{13}$

There is evidence of  $\theta_{13} > 0$  from global data

What is the exact value of  $\theta_{13}$ ?

What is the precision in the measurement of  $\theta_{13}$ ?

- CP violation in neutral lepton sector

Do neutrinos exhibit CP violation? Is  $\delta_{CP} \neq 0$ ?

Any value of  $\delta_{CP}$  between 0 to  $2\pi$  is possible

# Upcoming experiments

## Upcoming superbeam and reactor experiments

Setup	$t_\nu$ [yr]	$t_{\bar{\nu}}$ [yr]	$P_{\text{Th}}$ or $P_{\text{Target}}$	$L$ [km]	Detector tech	$m_{\text{Det}}$
Double Chooz	-	3	8.6 GW	1.05	Liquid scint	8.3 t
Daya Bay	-	3	17.4 GW	1.7	Liquid scint	80 t
RENO	-	3	16.4 GW	1.4	Liquid scint	15.4 t
T2K	5	-	0.75 MW	295	Water Cerenkov	22.5 kt
NO $\nu$ A	3	3	0.7 MW	810	TASD	15 kt

## Standard set-ups at their nominal luminosities

*P. Huber et al., JHEP 11 044 (2009)*

Can play a crucial role in planning a possible next generation of LBL neutrino experiments to address CPV and MH

# Ultimate Message

Expectation in 2025 without new facilities at  $3\sigma$  C.L.

- Size of  $\theta_{13} \Rightarrow$  if  $\sin^2 2\theta_{13} > 0.01$
- Mass ordering  $\Rightarrow$  if  $\sin^2 2\theta_{13} > 0.04$  for at most 30% of all CP phases
- CP violation in leptons  $\Rightarrow$  if  $\sin^2 2\theta_{13} > 0.02$  for at most 20% of all CP phases

Even for the largest currently allowed  $\theta_{13}$  more than 70% of parameter space are not accessible

# Ultimate Machine

Neutrino Factory  $\rightarrow$  Ultimate facility

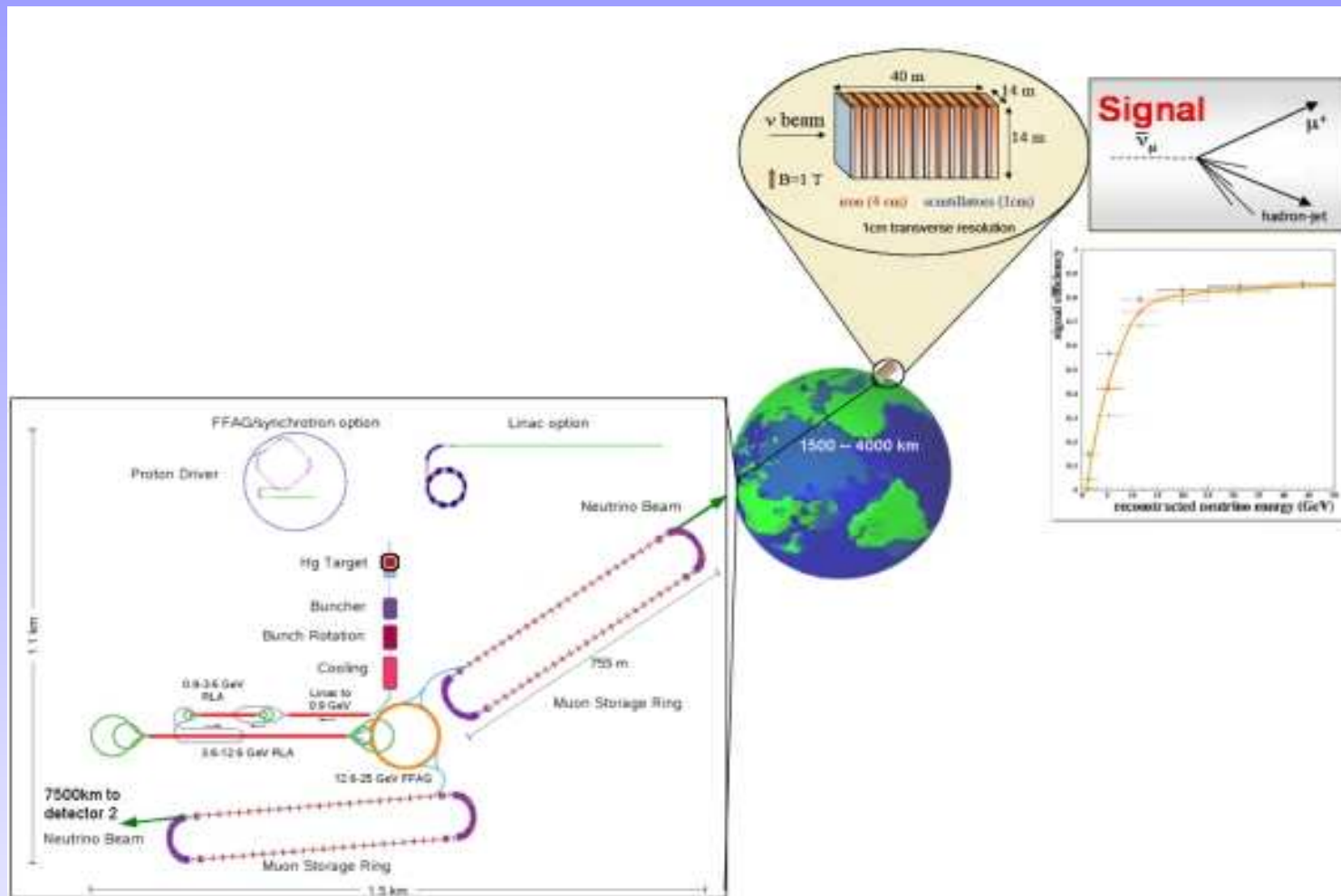
Powerful tool for CP violation discovery for small & large  $\theta_{13}$

Excellent sensitivity to  $Sgn(\Delta m_{31}^2)$  for 100% values of  $\delta_{CP}$

Neutrino Factory may be the first step towards  
the high energy frontier in form of a muon collider



# Baseline setup



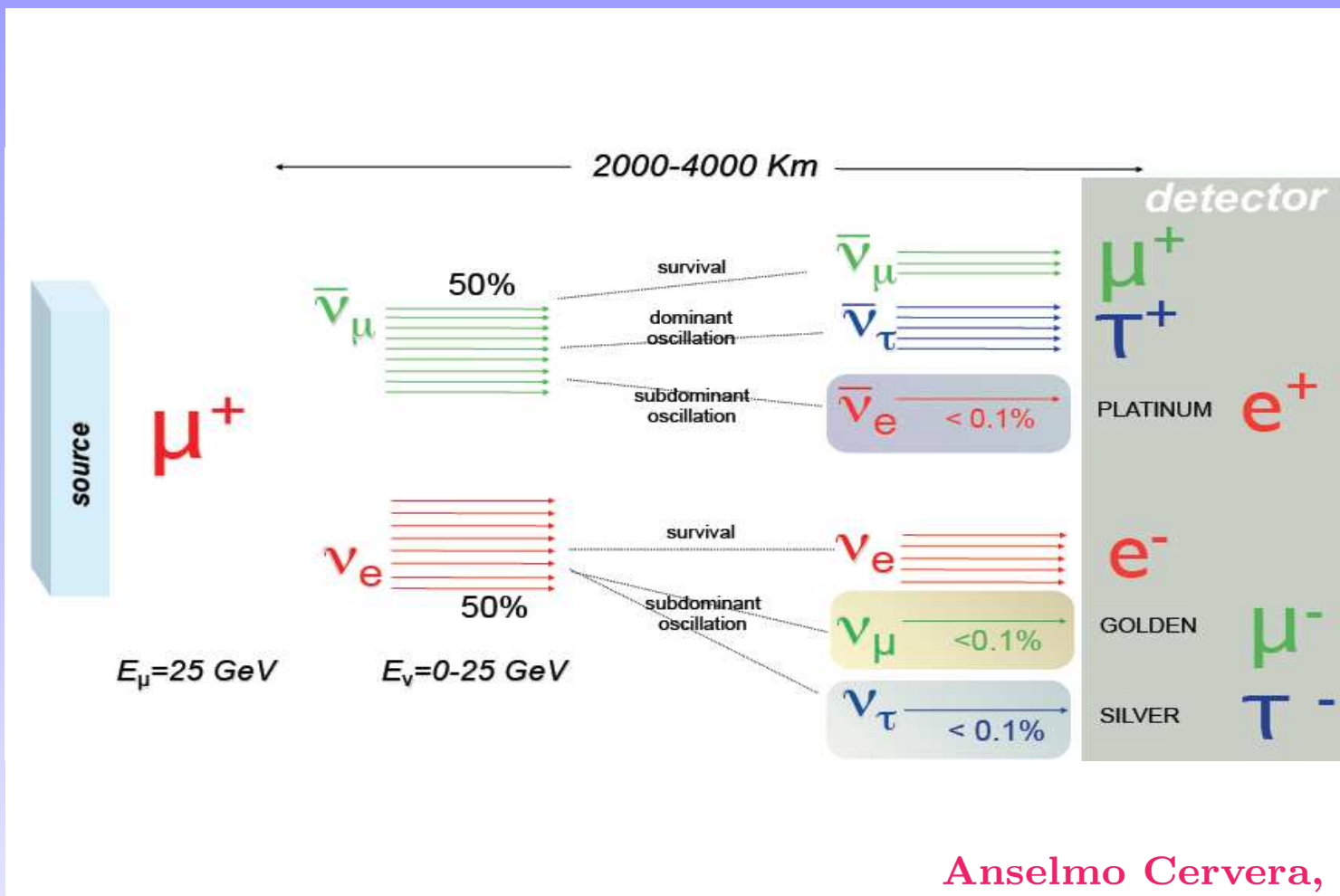
International Design Study for the Neutrino Factory

IDS-NF gives a 1st version baseline setup of HENF

# IDS-NF 1.0

- Two magnetized iron calorimeters (fiducial mass 50 kt) at  $L = 4000$  km and  $L = 7500$  km
- Two racetrack-shaped storage rings pointing towards these detectors
- $2.5 \times 10^{20}$  useful muon decays per polarity, decay straight, and year, *i.e.*,  $10^{21}$  useful muon decay per year
- Total run time of 10 years, *i.e.*,  $10^{22}$  useful muon decay in total
- The parent muon energy is assumed to be  $E_{\mu} = 25$  GeV

# Signal



Anselmo Cervera, WIN'11

Requires a detector which can distinguish  $\mu^-$  from  $\mu^+$

MIND can do that with a magnetic field of around 1 T

# Osc. channels & Backgrounds

- $\nu_\mu$  appearance:  $\nu_e \rightarrow \nu_\mu$  for  $\mu^+$  stored
- $\bar{\nu}_\mu$  appearance:  $\bar{\nu}_e \rightarrow \bar{\nu}_\mu$  for  $\mu^-$  stored
- $\nu_\mu$  disappearance:  $\nu_\mu \rightarrow \nu_\mu$  for  $\mu^-$  stored
- $\bar{\nu}_\mu$  disappearance:  $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$  for  $\mu^+$  stored
- Include backgrounds from  $\Rightarrow$ 
  1. charge mis-identification
  2. (electron) flavor mis-identification
  3. neutral current
- We use the GLoBES software for the simulation

P. Huber *et al.*, hep-ph/0407333 and hep-ph/0701187

# Golden Channel ( $P_{e\mu}$ )

The appearance probability ( $\nu_e \rightarrow \nu_\mu$ ) in matter, upto second order in the small parameters  $\alpha \equiv \Delta m_{21}^2 / \Delta m_{31}^2$  and  $\sin 2\theta_{13}$ ,

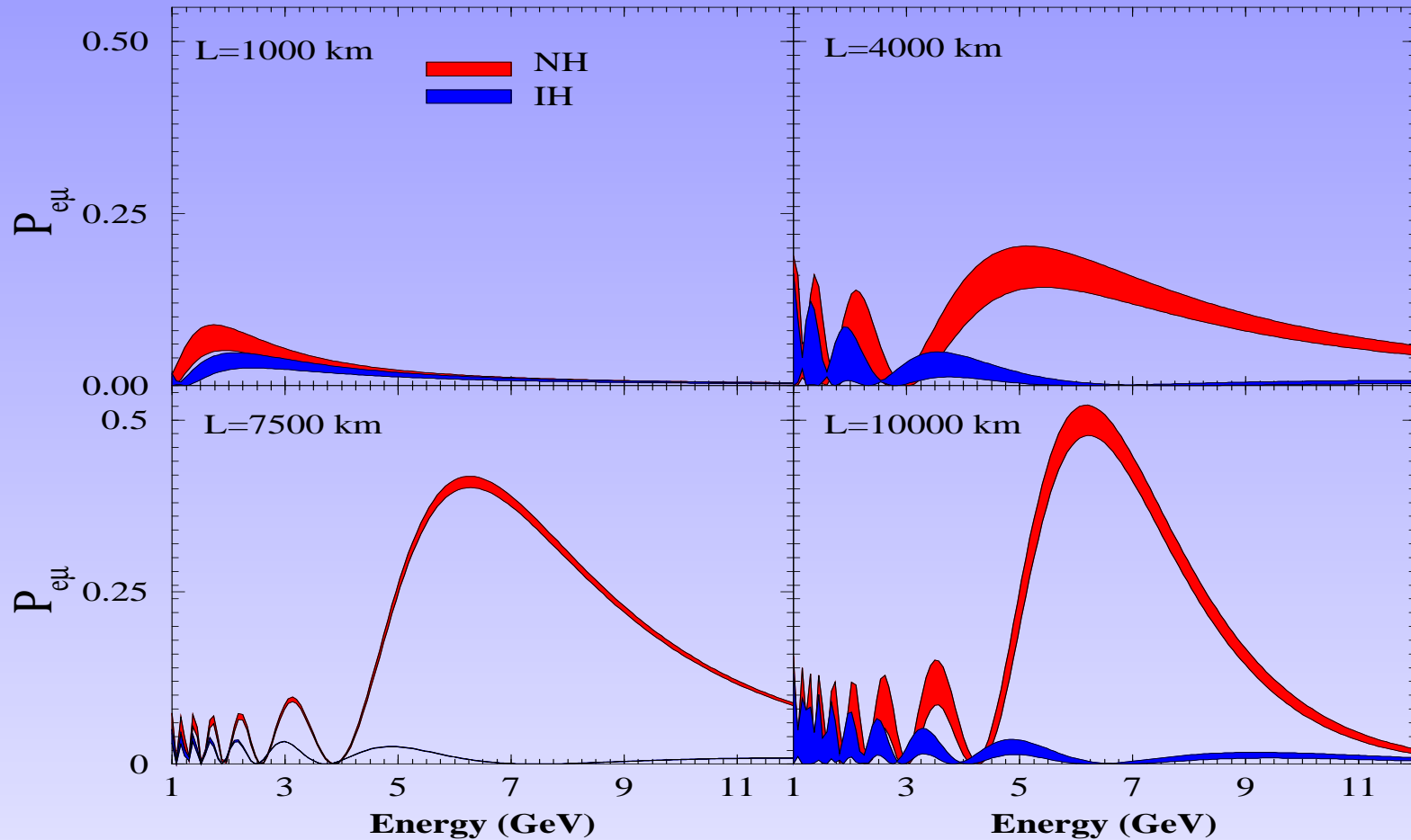
$$\begin{aligned} P_{e\mu} &\simeq \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2[(1 - \hat{A})\Delta]}{(1 - \hat{A})^2} \\ &+ \alpha \sin 2\theta_{13} \xi \sin \delta_{CP} \sin(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1 - \hat{A})\Delta]}{(1 - \hat{A})} \\ &+ \alpha \sin 2\theta_{13} \xi \cos \delta_{CP} \cos(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1 - \hat{A})\Delta]}{(1 - \hat{A})} \\ &+ \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}; \end{aligned}$$

where  $\Delta \equiv \Delta m_{31}^2 L / (4E)$ ,  $\xi \equiv \cos \theta_{13} \sin 2\theta_{21} \sin 2\theta_{23}$ ,  
and  $\hat{A} \equiv \pm(2\sqrt{2}G_F n_e E) / \Delta m_{31}^2$

Cervera *et al.*, hep-ph/0002108

Freund, Huber, Lindner, hep-ph/0105071

# Transition Probability $P_{e\mu}$

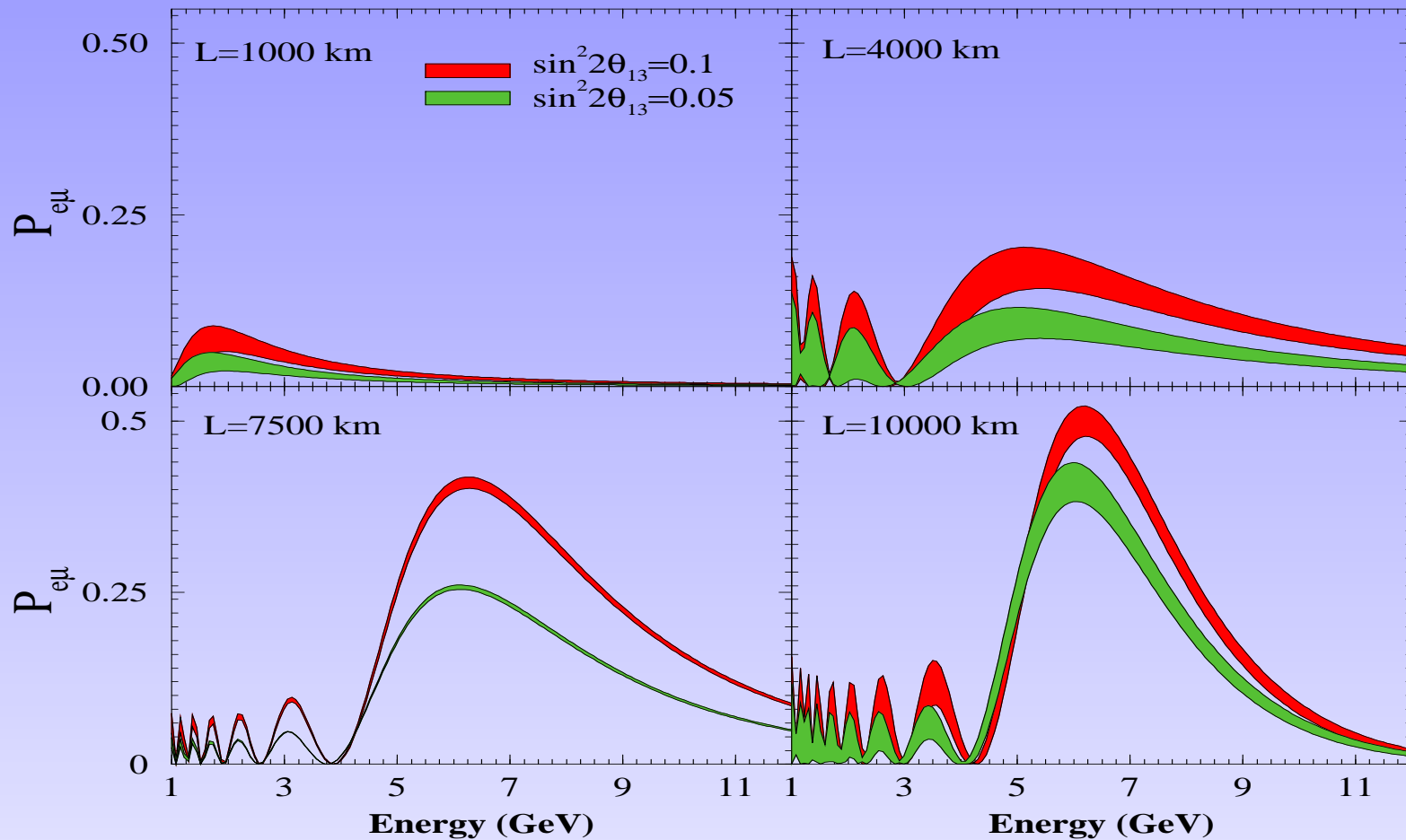


Agarwalla, Choubey, Raychaudhuri, hep-ph/0610333

Normal .vs. Inverted hierarchy

$\sin^2 2\theta_{13} = 0.1$

# Transition Probability $P_{e\mu}$



Agarwalla, Choubey, Raychaudhuri, hep-ph/0610333

Two different values of  $\sin^2 2\theta_{13}$

Normal hierarchy

# Eight-fold Degeneracy

■  $(\theta_{13}, \delta_{CP})$  intrinsic degeneracy

Burguet-Castell, Gavela, Gomez-Cadenas, Hernandez, Mena,  
hep-ph/0103258

■  $(\text{sgn}(\Delta m_{31}^2), \delta_{CP})$  degeneracy

Minakata, Nunokawa, hep-ph/0108085

■  $(\theta_{23}, \pi/2 - \theta_{23})$  degeneracy

Fogli, Lisi, hep-ph/9604415

Severely deteriorates the sensitivity



# Recent MIND Simulations

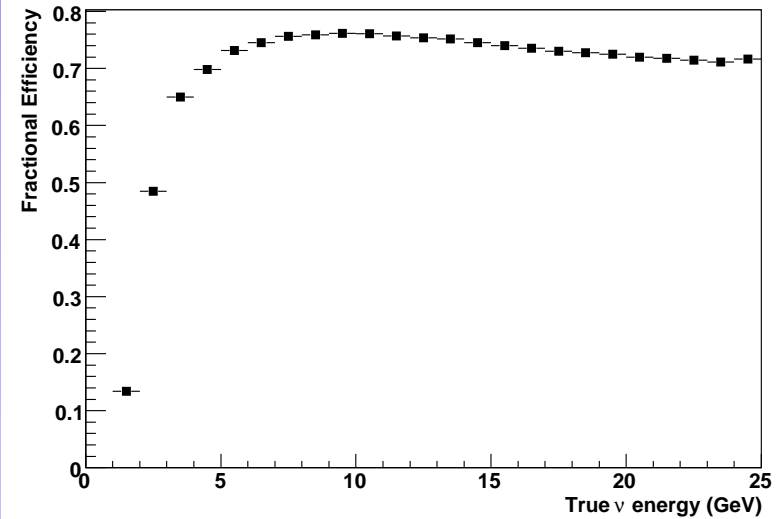
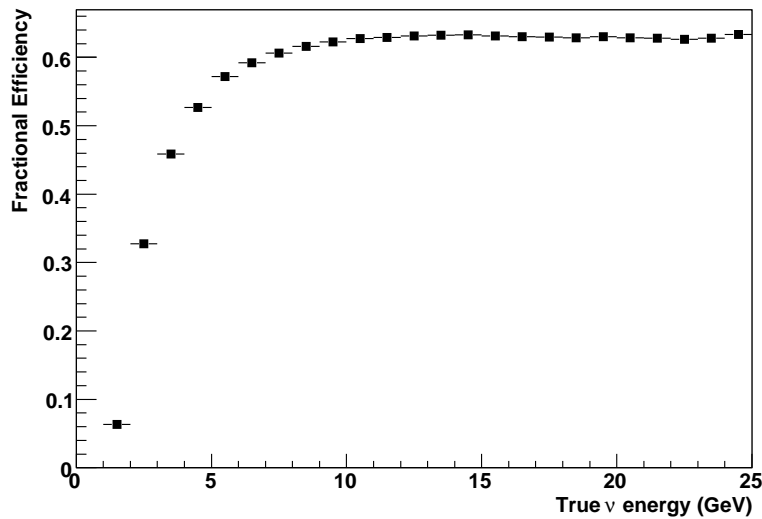
- Migration matrices for MIND are available  $\Rightarrow$  map the incident to the reconstructed neutrino energy for all individual signal and background channels

Cervera, Laing, Martin-Albo, Soler, arXiv:1004.0358 [hep-ex]

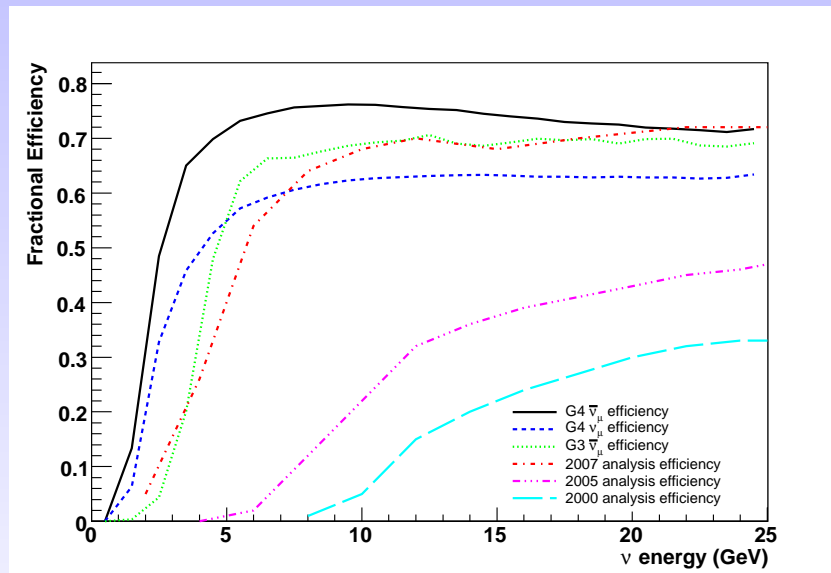
A. Laing's Ph.D. thesis, Glasgow university (2010)

- Optimized cuts have lead to a  $\Rightarrow$  lower threshold and higher signal efficiencies than in previous versions, while the background level has been maintained in the most recent analysis
- Separate response functions for  $\nu$  and  $\bar{\nu}$  are available  $\Rightarrow$  detection efficiency is better for  $\bar{\nu}_\mu$  compared to  $\nu_\mu$

# Improved Signal Efficiency

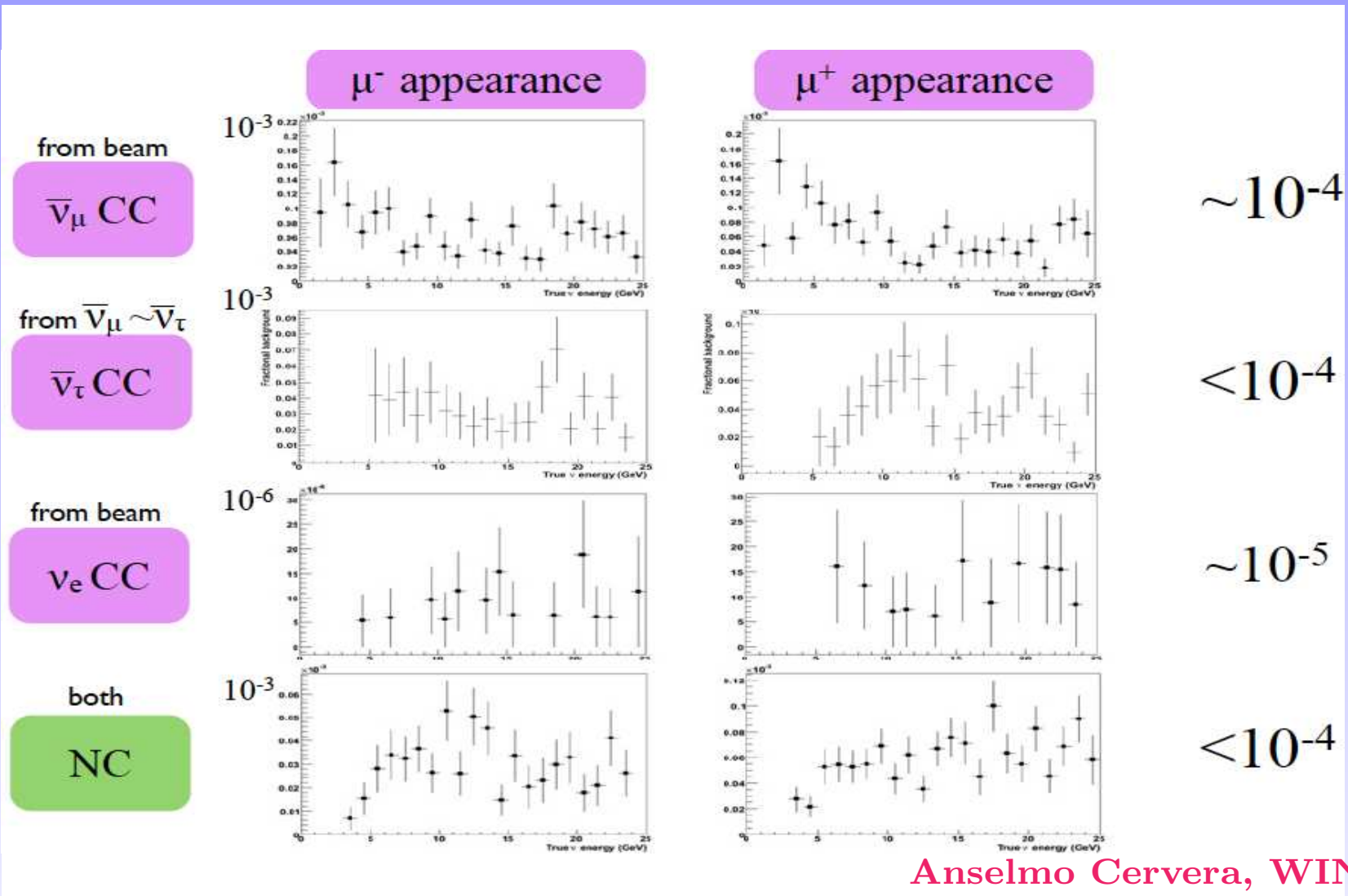


Left :  $\mu^-$  appearance & Right :  $\mu^+$  appearance



QES & RES events added, threshold  $\sim 2$  GeV, plateau  $\sim 5$  GeV

# Fractional Backgrounds



Background levels expected for the appearance channels

# $\nu_\tau$ contamination

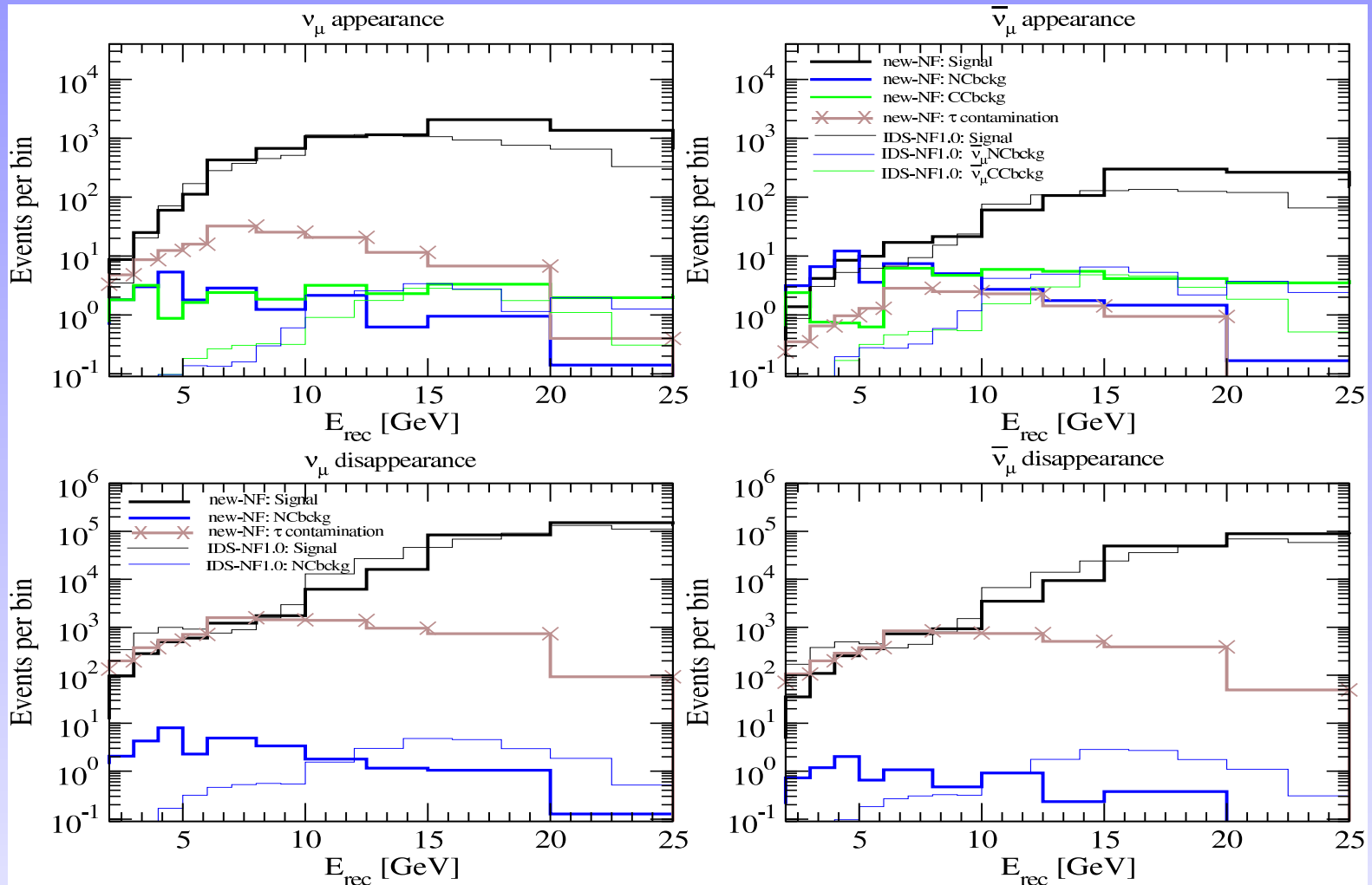
## Issue of $\nu_\tau$ contamination

- App.:  $\nu_e \rightarrow \nu_\tau \rightarrow \tau^- \xrightarrow{17\%} \mu^-$  (background) versus  $\nu_e \rightarrow \nu_\mu \rightarrow \mu^-$  (signal)
- Disapp.:  $\bar{\nu}_\mu \rightarrow \bar{\nu}_\tau \rightarrow \tau^+ \xrightarrow{17\%} \mu^+$  (background) versus  $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu \rightarrow \mu^+$  (signal)
- MIND cannot resolve the second vertex from the  $\tau$  decay, in contrast to OPERA-like emulsion cloud chamber
- For the  $\nu_\tau$  contamination ( $\nu_e \rightarrow \nu_\tau$  and  $\nu_\mu \rightarrow \nu_\tau$  channels), we use the migration matrix from

A. Donini *et al.*, arXiv:1005.2275

See also, D. Indumathi *et al.*, arXiv:0910.2020

# Event rate comparison



Agarwalla, Huber, Tang, Winter, JHEP 01 120 (2011)

Thin curves : IDS-NF 1.0 and thick curves : new-NF including backgrounds from  $\nu_\tau$

Muon energy = 25 GeV, detector mass = 50kt,  $L = 4000$  km,  $\theta_{13} = 5.6^\circ$  &  $\delta_{CP} = 0$

# Event rates

	Signal	NC bckg	CC bckg	$\nu_\tau$ bckg
$\nu_\mu$ (app)	7521	20	25	142
$\bar{\nu}_\mu$ (app)	924	45	39	13
$\nu_\mu$ (disapp)	$4.0 \times 10^5$	31	-	8154
$\bar{\nu}_\mu$ (disapp)	$2.4 \times 10^5$	8	-	4337

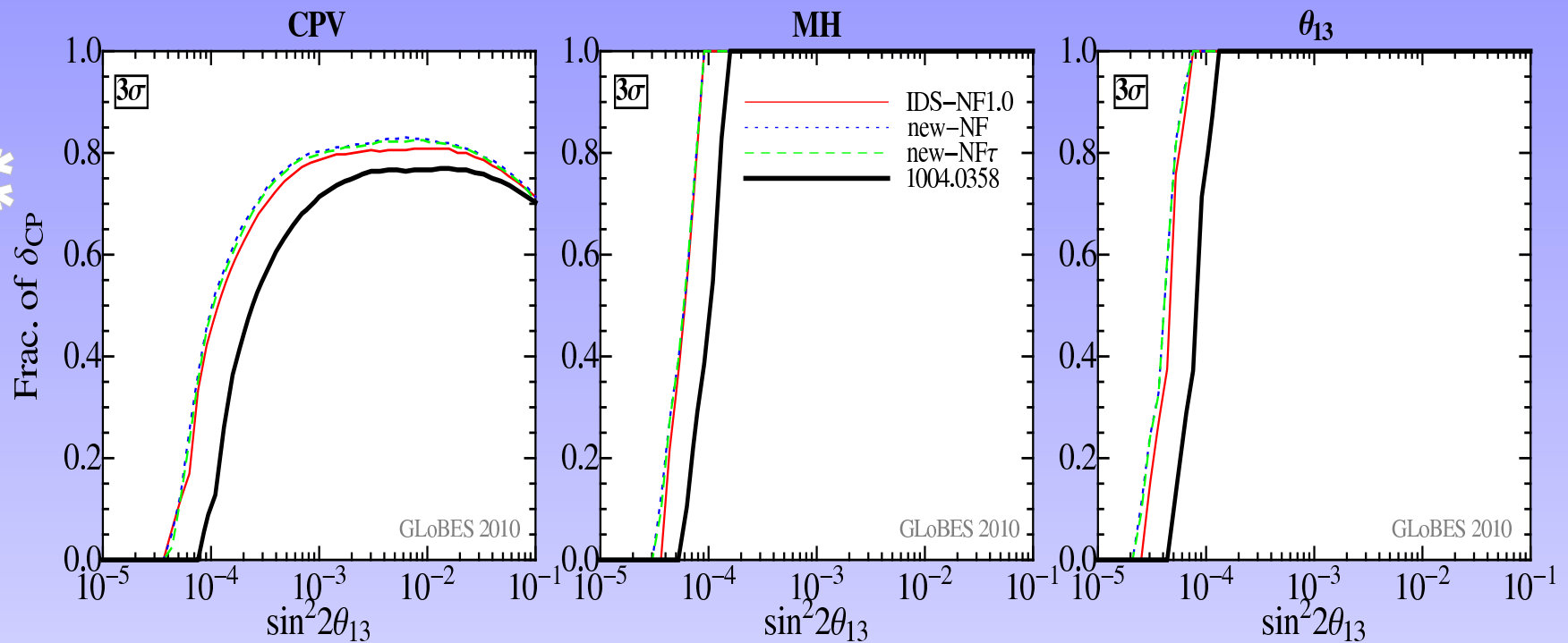
## Event rates for new- $\text{NF}_\tau$

50kt detector,  $L = 4000$  km, muon energy of 25 GeV

NH,  $\theta_{13} = 5.6^\circ$  and  $\delta_{\text{CP}} = 0$

Agarwalla, Huber, Tang, Winter, JHEP 01 120 (2011)

# Comparison of discovery reach



Agarwalla, Huber, Tang, Winter, JHEP 01 120 (2011)

“IDS-NF” refers to the detector in the IDS-NF baseline setup 1.0

Results with the migration matrices from arXiv:1004.0358 is indicated by the label “1004.0358”

The label “new-NF” refers to most up-to-date detector simulation in A. Laing’s thesis

$\nu_\tau$  contaminations in the app. and disapp. channels are, in addition, included in “new-NF $\tau$ ” following arXiv:1005.2275

Two baselines 4 000 km and 7 500 km with two 50 kt MIND detectors and  $E_\mu = 25$  GeV

# More Results

## Optimization of a green-field setup

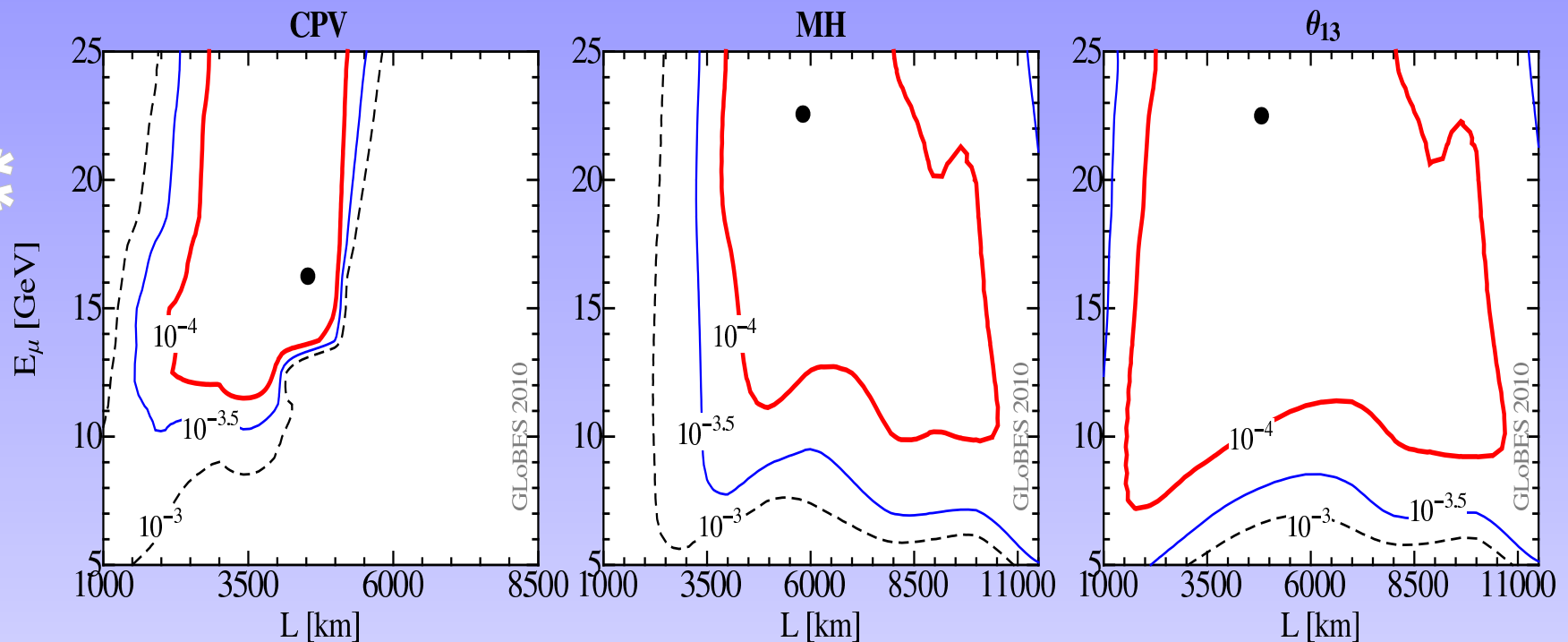
Agarwalla, Huber, Tang, Winter, JHEP 01 120 (2011)

### LENF .vs. HENF

- no particular accelerator and detector sites
- no constraints on the baselines and muon energy
- optimization is performed using the migration matrices from A. Laing's Ph.D. thesis
- Since the detection threshold has improved, can MIND detector interpolate between LENF and HENF?



# Assume $\sin^2 2\theta_{13}$ is not known



Agarwalla, Huber, Tang, Winter, JHEP 01 120 (2011)

Discovery reach in  $\sin^2 2\theta_{13}$  for maximal CPV, MH, and  $\theta_{13}$  as a function of  $L$  and  $E_\mu$

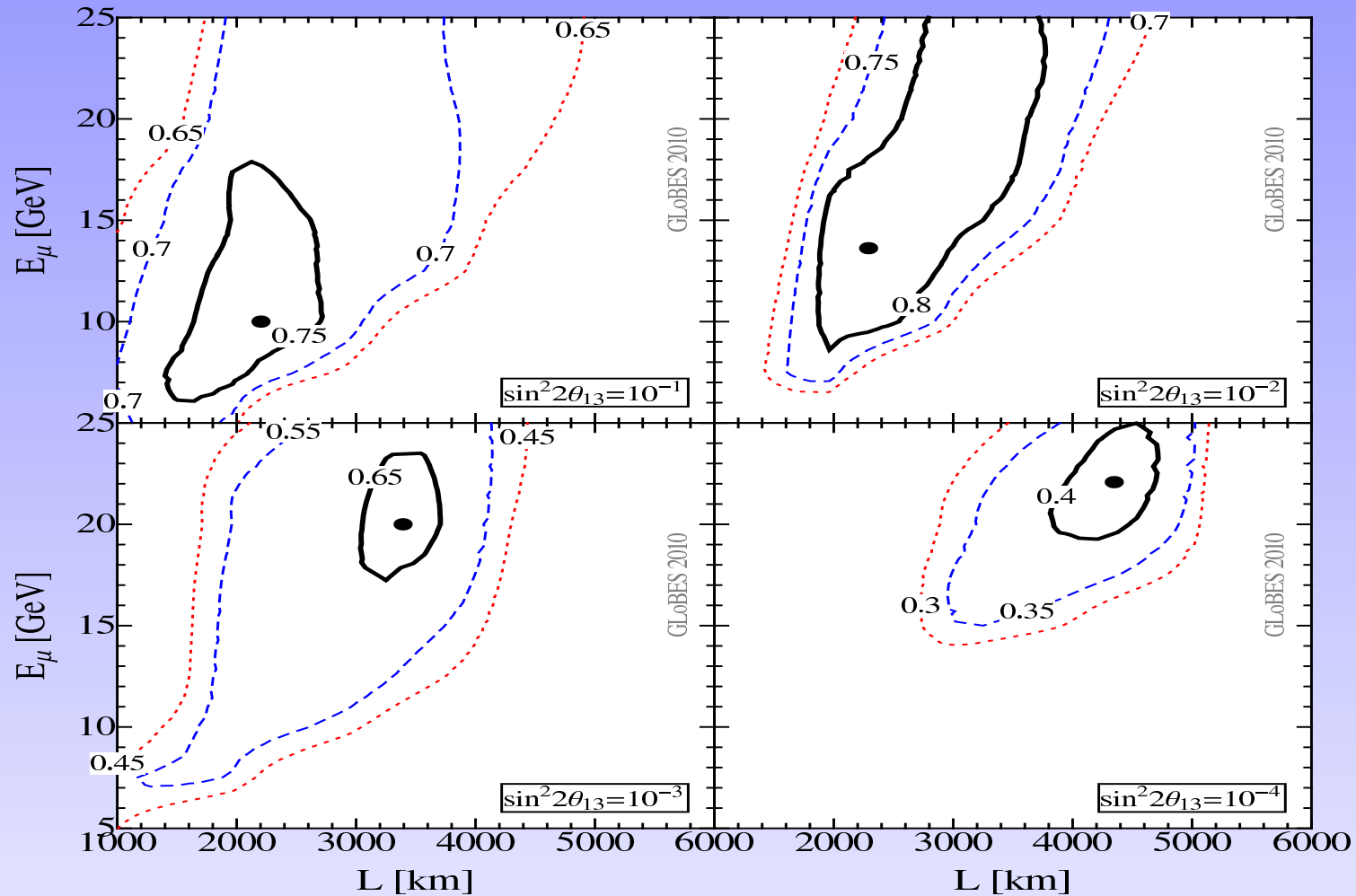
The contours show for how small (true)  $\sin^2 2\theta_{13}$  the different quantities will be discovered at the  $3\sigma$  CL, where  $\delta_{CP} = \pi/2$  is chosen as a true value in all cases

Optimal performance: (4519,16.25), (5805,22.57) and (4800,22.50). Here SF=1 is used with one 50 kt detector

# What do we learn?

- CPV : 2500-5000 km baseline and  $E_\mu$  above 12 GeV
- For  $\delta_{\text{CP}} = 3\pi/2$ , a second baseline may be required
- Relatively low  $E_\mu$  are allowed because of the low detection threshold
- MH : Baselines longer than 4000 km and  $E_\mu$  larger than about 10-12 GeV are needed to cover the MSW resonance energy of about 8 GeV. Magic baseline (7500 km) is very useful for degeneracy resolution
- $\theta_{13}$  : Extremely wide baseline and energy range, posing the least constraints
- Note : All these results depend on the choice of  $\delta_{\text{CP}}$

# Assume $\sin^2 2\theta_{13}$ is known



Agarwalla, Huber, Tang, Winter, JHEP 01 120 (2011)

Frac. of  $\delta_{CP}$  for which CPV will be discovered ( $3\sigma$  CL) as a function of  $L$  and  $E_\mu$  for the single

baseline NF. Here SF=1 is used with a 50 kt detector. Optimal performance: (2200,10.00),

(2288,13.62), (3390,20.00) & (4345,22.08)

# What is the message?

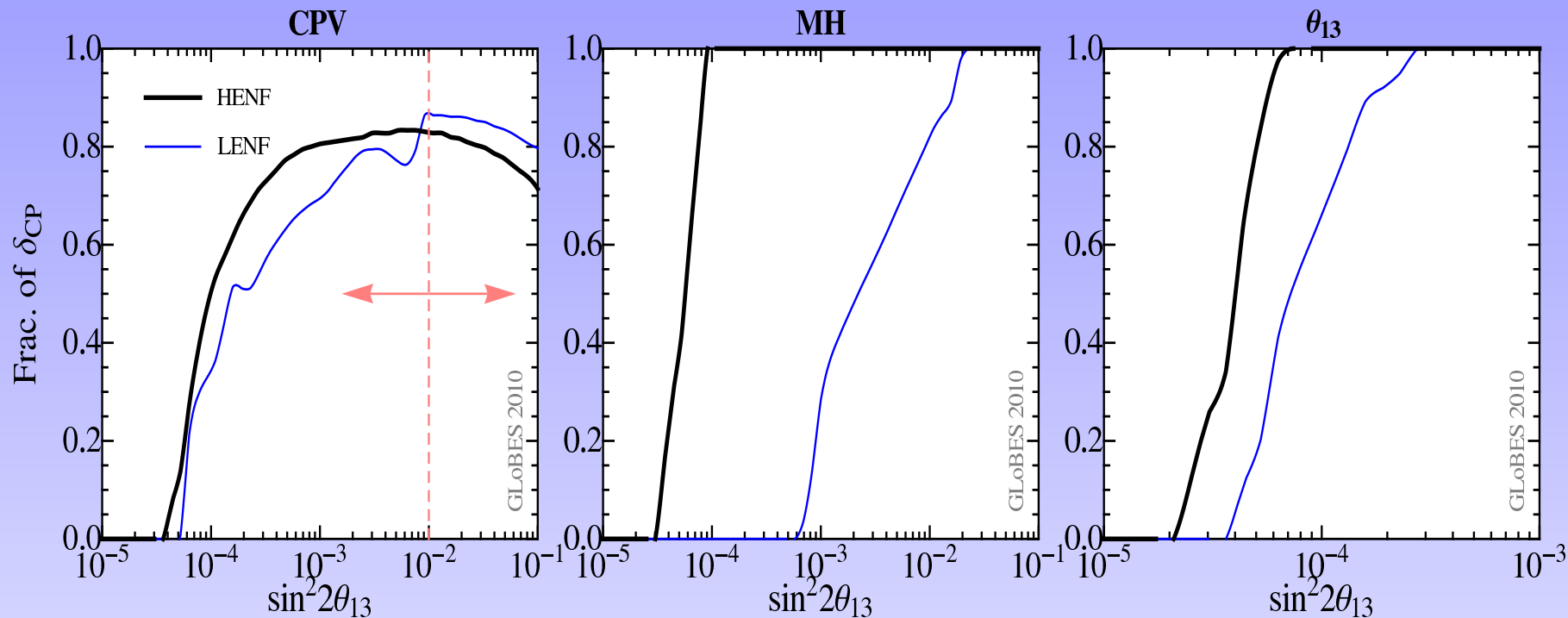
- Most interesting : optimization of the fraction of  $\delta_{\text{CP}}$  for which CPV can be discovered
- This strongly depends on the value of  $\sin^2 2\theta_{13}$  chosen
- Large  $\sin^2 2\theta_{13} \simeq 10^{-1}$  : shorter baselines and lower energies are preferred. Even  $E_\mu$  as low as 5 GeV at the FNAL-Homestake baseline of about 1300 km is not far from optimal
- MIND approaches the T ASD performance of the L ENF. Compared to earlier analyses without background migration, too high  $E_\mu$  are in fact disfavored in the large  $\sin^2 2\theta_{13}$  case



# LENF .vs. HENF

LENF and HENF are just two versions of the same  
experiment in different optimization regions

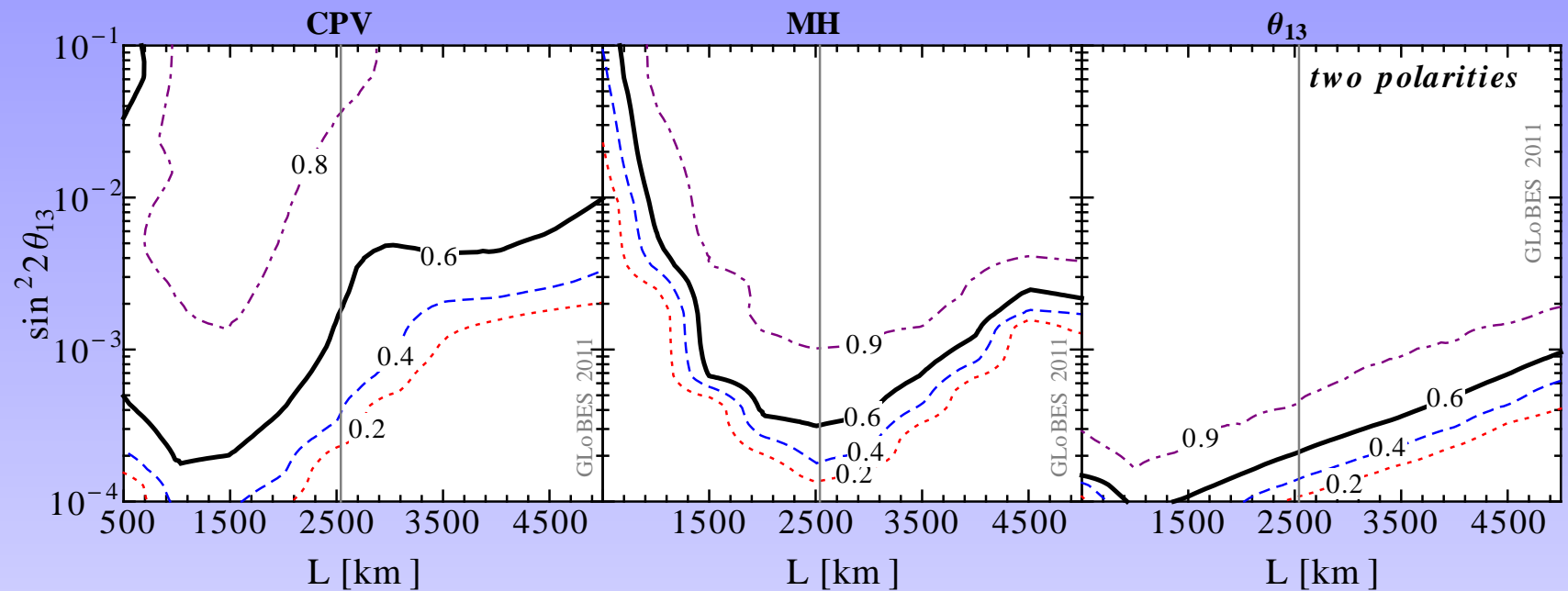
# LENF .vs. HENF with MIND



Agarwalla, Huber, Tang, Winter, JHEP 01 120 (2011)

Comparison of the discovery reaches of CPV, MH, and  $\theta_{13}$  at the  $3\sigma$  CL between a single baseline LENF with  $E_\mu = 10$  GeV and  $L = 2000$  km, and a two baseline HENF with  $E_\mu = 25$  GeV,  $L_1 = 4000$  km, and  $L_2 = 7500$  km

# LENF with Magnetized TASD



Agarwalla, Huber, Tang, Winter, JHEP 01 120 (2011)

Total  $1.25 \times 10^{22}$  muons per polarity &  $E_\mu = 5$  GeV

25 kt magnetized TASD with 94% efficiency above 1 GeV,  
10% energy resolution and  $10^{-3}$  background level

Talk by Peter Ballett

# Concluding remarks

- Lower threshold and higher efficiencies compared to earlier simulations imply that the MIND detector characteristics are getting more similar to the characteristics of the detectors proposed for the LENF (magnetized T ASD)
- We recover the  $L$ - $E_\mu$ -optimization of the LENF for large  $\sin^2 2\theta_{13}$  with MIND : a single baseline NF with  $E_\mu$  as low as 5 GeV and a baseline as short as FNAL-Homestake (about 1300 km) might be sufficient

Thank You