

Neutrinos and the LHC

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with Maiezza, Nesti, Senjanović, Tello, Vissani and Zhang

NuFact '11, CERN

LHC a neutrino machine?

Neutrino mass is an experimentally established fact for BSM physics.

CERN colloquium by Senjanović

• Might as well be @ TeV (hierarchy)

• A phenomenological hint may already be here

Neutrino mass and BSM

Facts

- $\Delta m_{S,A}^2 \,\&\, \theta_{A,S,13}$
- At least two massive light neutrinos

Questions

- Mass scale & hierarchy
- CP phases
- Dirac or Majorana

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Origin of m_{ν}

• Standard Model ν_L only, $m_{\nu} = 0$

• Effective theory
$$\mathcal{O}_W = y \frac{\ell h \ell h}{\Lambda} \Rightarrow m_\nu = \frac{y^2 v^2}{\Lambda}$$

High scale
Typical in GUTs
Indirect, LHC of little use

• Theory: GUT remnant

TeV

• Phenomenological

Neutrinos @ LHC

UV completions of \mathcal{O}_W

• Renormalizable theory = seesaw scenarios

Fermionic singlet

Bosonic triplet

Fermionic triplet







Type I+III

- Triplet predicted at TeV by a minimal SU(5) with 24_F
- Neutrino mass matrix through decays
- Oscillations-collider connection

Bajc, Senjanović '06 Bajc, MN, Senjanović '07

UV completions of \mathcal{O}_W

Fermionic singlet

S

ν

Bosonic triplet

Fermionic triplet



Type I+II

 ν

- Predicted by a minimal LR symmetric theory
- Singlet is gauged, type I required by anomalies

ν

- Triplet breaks the LR symmetry type II
- May need to be light \hookrightarrow

ν

Left-Right symmetry Talk by Senjanović

$$\begin{pmatrix} \nu_L \\ e_L \end{pmatrix}, W_L \leftrightarrow \begin{pmatrix} \nu_R \\ e_R \end{pmatrix}, W_R$$

• Parity restoration at high scales

Pati, Salam '74 Mohapatra, Pati '75

• Spontaneously broken

Mohapatra, Senjanović '75

• Origin of the seesaw mechanism

Minkowski '77 Senjanović '79 Senjanović, Mohapatra '80

Minimal LR Model

$$SU(2)_L \times SU(2)_R \times U(1)_{B-L}$$

• LR symmetric Higgs sector, parity invariant \mathcal{L} $\Phi(2,2,0), \Delta_L(3,1,2), \Delta_R(1,3,2)$



Mass spectra

 $\mathcal{L}_{Y} = \overline{\ell}_{L} (Y_{\Phi} \Phi + \tilde{Y}_{\Phi} \tilde{\Phi}) \ell_{R} + \ell_{L}^{T} Y_{\Delta_{L}} \Delta_{L} \ell_{L} + \ell_{R}^{T} Y_{\Delta_{R}} \Delta_{R} \ell_{R} + \text{h.c.}$ • Dirac terms • Majorana terms $M_D = v_1 Y_{\Phi} + v_2 e^{-i\alpha} \tilde{Y}_{\Phi}, \quad M_{\nu_R} = v_R Y_{\Delta_R}$ $M_{\ell} = v_2 e^{i\alpha} Y_{\Phi} + v_1 \tilde{Y}_{\Phi} \quad M_{\nu_L} = v_L Y_{\Delta_L} - M_D^T M_{\nu_B}^{-1} M_D$ $\mathcal{C}: \quad Y_{\Phi} = Y_{\Phi}^{T}, \quad Y_{\Delta_{L,R}} = Y_{\Delta_{R,L}}^{*}$ • Mass eigenstate basis

 $M_{\ell} = U_{\ell L} \, m_{\ell} \, U_{\ell R}^{\dagger}, \ M_{\nu_{L}} = U_{\nu L}^{*} \, m_{\nu} \, U_{\nu L}^{\dagger}, \ M_{\nu_{R}} = U_{\nu R}^{*} \, m_{N} \, U_{\nu R}^{\dagger}$ $\mathcal{C} : U_{\ell L} = U_{\ell R}^{*}$

The Interactions

• New Gauge:
$$\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} \left(\bar{\nu}_L V_L^{\dagger} W \ell_L + \bar{\nu}_R V_R^{\dagger} W_R \ell_R \right)$$

• New Scalar: $\mathcal{L}_{\Delta^{++}} = e_R^T Y \Delta_R^{++} e_R$

$$Y = \frac{g}{M_{W_R}} V_R^* m_N V_R^\dagger$$

- Flavor fixed in type II: $V_R^* = V_L$
 - Remember; two angles, one limit, no phases

The Interactions

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Flavor fixed in type II: V teonly!
Remember; two Example only!

• LHC could eventually measure these

Bounds on the LR scale

• Theoretical bounds since '81

Beall et al. '81...Zhang et al. '07

• A recent detailed study, including CP violation

Maiezza, Nesti, MN, Senjanović '10

 $C: M_{W_R} > 2.5 \,\text{TeV}$ $\mathcal{P}: M_{W_R} > 3.2 - 4.2 \,\text{TeV}$

• Direct searches

• Dijets

• Light neutrino

 $M_{W_R} > 1.51 \text{ TeV}$ $M_{W_R} > 2.27 \text{ TeV}$

CMS 1107.4771

CMS PAS EXO-11-024

LHC is here!

Left-Right @ LHC

• Most interesting channel: l l j j

Keung, Senjanović '83

• LNV at colliders



Gauge production: s-channel, W nearly on-shell
Clean, no missing energy
Reconstructs W and N masses
Information on the flavor

Limits from the LHC

MN, Nesti, Senjanović, Zhang '11



High N inaccessible due to jets, low N due to isolation
This year: L = 0.1(1) fb⁻¹ : M_{W_R} > 1.6(2.2) TeV
Electron and muon channel essentially the same

Limits from the LHC

CMS PAS EXO-11-002

MN, Nesti, Senjanović, Zhang '11



Electron and muon channel essentially the same 0

Low/high energy interplay

- Majorana mass term dominance
 - $M_{\nu_R} = v_R Y \qquad \qquad M_{\nu_L} = v_L Y^*$
- Implications for masses and mixings
 - $m_N \propto m_{\nu} \qquad \qquad U_{\nu R} = U_{\nu L}^* \Rightarrow V_R = V_L^*$

•

$$\frac{m_{N_2}^2 - m_{N_1}^2}{m_{N_3}^2 - m_{N_1}^2} = \frac{m_{\nu_2}^2 - m_{\nu_1}^2}{m_{\nu_3}^2 - m_{\nu_1}^2} \simeq \pm 0.03 \qquad \circ \begin{array}{c} \text{Hierarchy probe} \\ @ LHC \end{array}$$

$$m_{\rm cosm} = \sqrt{\Delta m_A^2} \frac{\sum_i m_{N_i}}{\sqrt{|m_{N_3}^2 - m_{N_2}^2|}}$$

• Cosmo-oscillations -LHC link

$0\nu 2\beta$ and cosmology



• Majorana mass implies $0\nu 2\beta$ Talk by Lopez-Pavon

Gerda, Cuore, Majorana,...

Review by Rodejohann '11

• Claim of observation

 $|m_{\nu}^{ee}| \simeq 0.4 \text{ eV}$

Klapdor-Kleingrothaus '06, '09

$0\nu 2\beta$ and cosmology



Cosmological bounds

 $\sum m_{
u} < 0.17 \text{ eV}$ Seljak et al. '06

WMAP alone

$$\sum m_{\nu} < 0.44 \text{ eV}$$

Hannestad et al. '10

• If confirmed, a hint for NP Vissani '02

New Diagrams for $0\nu 2\beta$

Mohapatra, Senjanović '81



• In agreement with cosmology?

$0\nu 2\beta$: the new contribution



$$\mathcal{H}_{\rm NP} = G_F^2 \sqrt{\frac{2}{N_{ej}}} \left[\frac{1}{m_{Nj}} + \frac{2m_{Nj}}{m_{\Delta}^2} \right] \frac{M_W^4}{M_{W_R}^4} J_{R\mu} J_R^\mu \overline{e_R} e_R^c$$

• Type II: $V_R^* = V_L$ • Δ_L suppressed: $m_{\nu}/m_{\Delta} \ll 1$

- LFV: $m_N/m_{\Delta_R} < 1$
- The gauge contribution is dominant

$0\nu 2\beta$: the total contribution

• LHC accessible regime



$$|m_{\nu+N} e^{e}|^{2} \equiv |m_{\nu} e^{e}|^{2} + |m_{N} e^{e}|^{2}$$
• Interference small
• Reversed role of hierarchies
• No tension with cosmology
• A light $m_{N} < M_{W}$
• No cancellations

R

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R

LHC and $0\nu 2\beta$

• Suppose NP is a must for $0\nu 2\beta$; $\theta_{13} \simeq 0 \ \alpha = 2(\varphi_2 - \varphi_1)$

 $|M_{\nu+N}^{ee}|^2 = |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_1} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12} + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12}|^2 + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta_{12}|^2 + m_{\nu_2} e^{i\alpha} \sin^2 \theta_{12}|^2 + |m_{\nu_2} \cos^2 \theta$ $\left| p^2 \frac{M_W^4}{M_{W_{\rm P}}^4} \left(\frac{1}{m_{N_1}} \cos^2 \theta_{12} + \frac{1}{m_{N_2}} e^{i\alpha} \sin^2 \theta_{12} \right) \right|^2$

LHC and $0\nu 2\beta$

• Suppose NP is a must for $0\nu 2\beta$; $\theta_{13} \simeq 0 \ \alpha = 2(\varphi_2 - \varphi_1)$



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Lepton Flavor Violation



Cirigliano et al '04 Raidal, Santamaria '97



Combined LFV constraints

• Muonic and tau channels

Tello, MN, Senjanović, Vissani '10

• Varying PMNS constrains



$$m_N/m_\Delta < 1$$



Role of θ_{13}

• Flavor structure in muon-electron transitions

$$\mathcal{A}(\mu \to 3e) \propto \left(-m_{N_1} + \exp^{i\alpha} m_{N_2}\right) \sin 2\theta_{12} \cos \theta_{23}$$
$$\left(m_{N_1} \cos^2 \theta_{12} + \exp^{i\alpha} m_{N_2} \sin^2 \theta_{12}\right) / m_{\Delta^{++}}^2$$

• $\mu \rightarrow 3e$ insensitive to θ_{13}

$$\mathcal{A}(\mu \to e) \propto \frac{\Delta m_{N_{13}}^2}{m_{\Delta^{++}}^2} \left(\frac{\Delta m_S^2}{2\Delta m_A^2} \sin 2\theta_{12} \cos \theta_{23} + e^{i\delta} \sin \theta_{13} \sin \theta_{23} \right)$$

• $\mu - e$ conversion and $\mu \rightarrow e\gamma$ depend a lot

Role of θ_{13}



Conclusions

- $0\nu 2\beta$ signal may require new physics at TeV
- Left-Right symmetry @ LHC a natural example
 - no tension with cosmology, precision data or LFV
 - links oscillations, cosmology and the LHC
- Exclusions with fairly low luminosity and 7 TeV
- LHC may observe LNV and in turn connect it to low energy rates, both LNV and LFV