

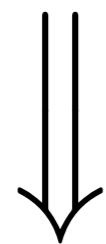
# The Baryon--Dark-Matter Connection

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Oxford University

# Introduction & Motivation

A primary goal of modern physics:

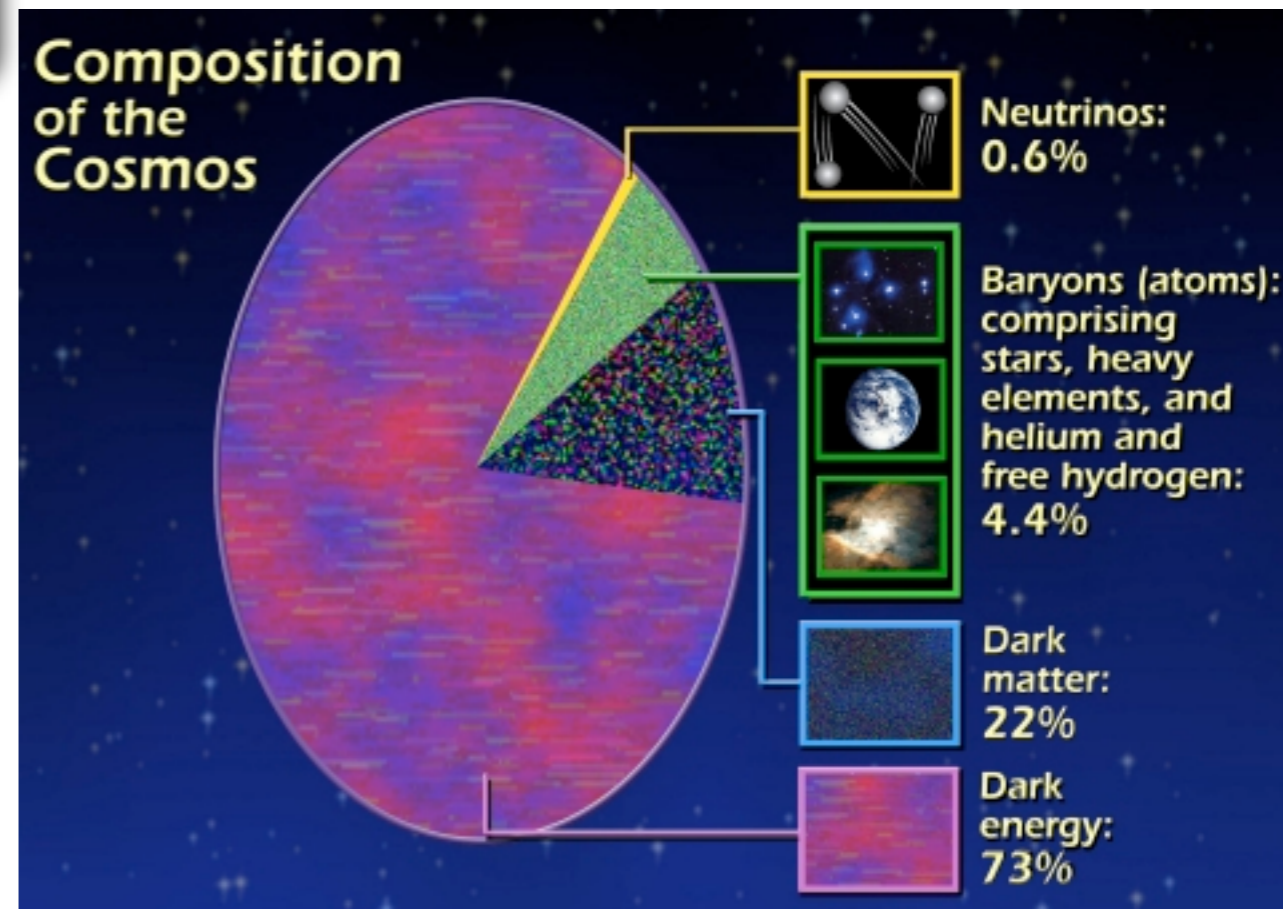
Theory of Particle Interactions  
+  
Big Bang based on GR



compute

Contents of Universe

Composition is complex &  
surprising!



# Introduction & Motivation

Reality is hard to understand

$\Omega_B$   
 $\Omega_{DM}$  } many ideas - can they be **tested?**

$\Omega_{DE}$  remarkable surprise - no good idea (anthropic?)

# Introduction & Motivation

## Matter genesis - leading ideas

- Dark matter

WIMPs: Thermal freeze-out with scale  $v$

Axions: Mis-alignment or thermal production

- Baryons

GUT baryogenesis - problems with cosmo

Affleck-Dine - impossible to test?

Leptogenesis - hard to test, disfavored by cosmo?

Electroweak baryogenesis - hard to implement b/c CP

...but see "Axion-Assisted Electroweak Baryogenesis," Nathaniel Craig & JMR, arXiv:1007.0019

# Introduction & Motivation

- ★ Unrelated origin, involving very different physics, of baryons & DM, make it hard to understand

$$\Omega_{DM}/\Omega_B \simeq 4.86$$

- ★ Freeze-out dominates thinking about DM candidates, detection, and collider phenomenology

Are we being misled?

# Introduction & Motivation

Alternative view:

similar physics underlies both  $\Omega_B$  and  $\Omega_{DM}$

(Nussinov '85; Gelmini, Hall, Lin '87; Barr '91; Kaplan '92; Thomas '95; Hooper, JMR, West '04; explosion in last 2 yrs with work by many people, esp Zurek etal...)

<b>Baryons:</b>	$U(1)_B$	$u, d, s\dots$	$p$ stable	$\Omega_B \propto m_B \eta_B$
<b>DM:</b>	$U(1)_X$	$X_0, X_1, X_2\dots$	$X_0$ stable	$\Omega_X \propto m_X \eta_X$

# Introduction & Motivation

Alternative view:

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**Baryons:**  $U(1)_B$        $u, d, s\dots$        $p$  stable       $\Omega_B \propto m_B \eta_B$

**DM:**       $U(1)_X$        $X_0, X_1, X_2\dots$        $X_0$  stable       $\Omega_X \propto m_X \eta_X$

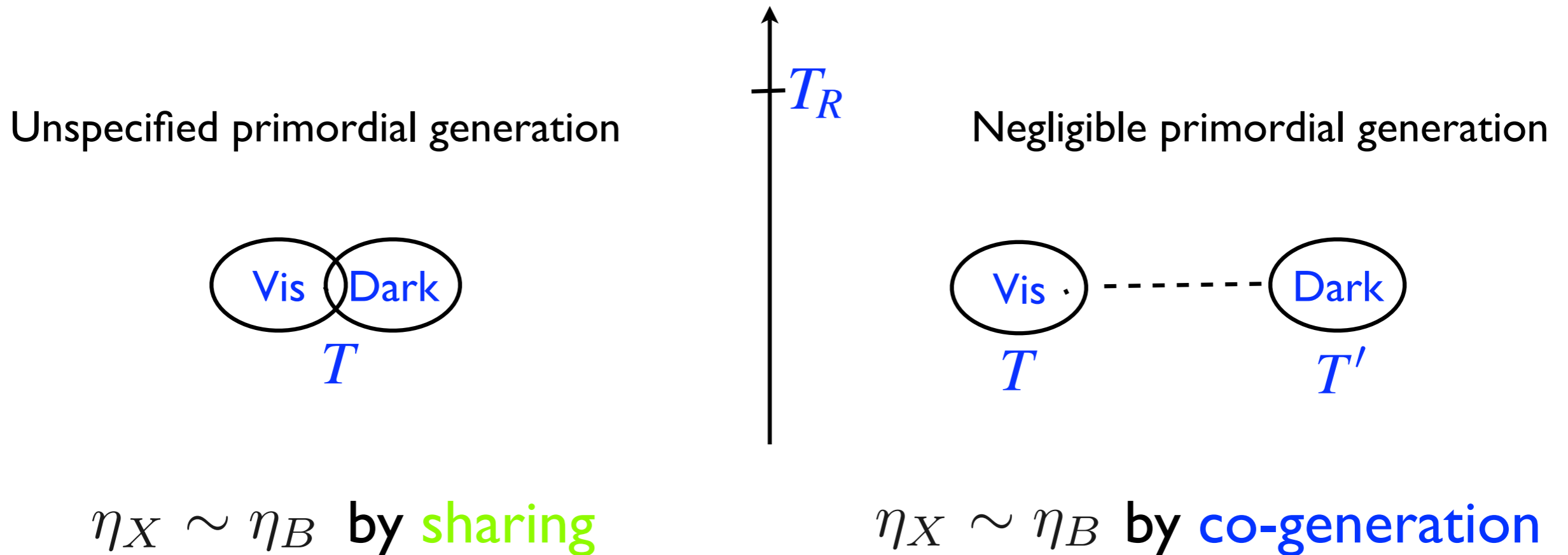
At some era

Interactions violate B and X to yield related values for  $\eta_B$  and  $\eta_X$

$$\frac{\Omega_X}{\Omega_B} = \frac{\eta_X m_X}{\eta_B m_B}$$

# Introduction & Motivation

- Two general categories of theories: “sharing” & “co-generation”



Co-generation is more ambitious: attempts to explain simultaneous origin of B & X asymmetries (at scale  $\sim$  TeV allowing test at LHC...)



# Introduction & Motivation

Alternative view (either sharing or co-generation):

- incompatible with SUSY neutralino DM
- alters expected LHC signals of new physics
- changes one or both direct/indirect DM detection

co-generation appears hard as requires B, X violation & out-of-equilibrium condition at TeV scale

requires a new theory of calculable thermal DM production...

# “Freeze-In” Production

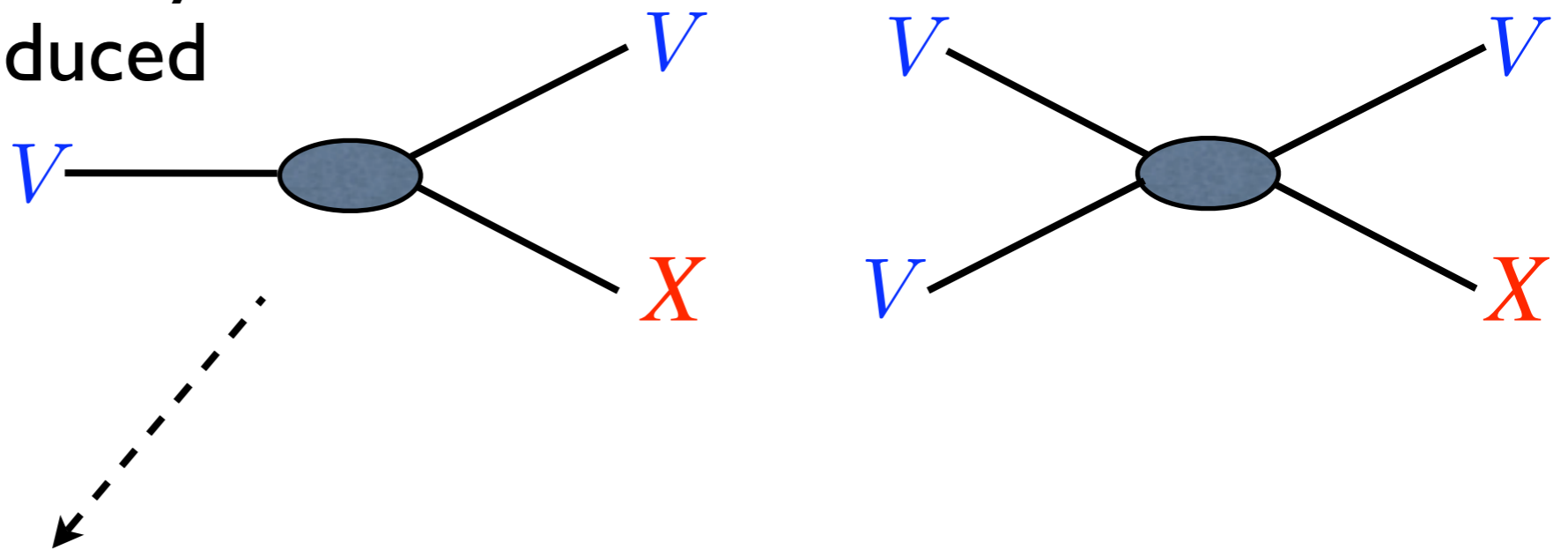
Hall, Jedamzik, JMR, West, arXiv:0911.1120

Hall, JMR, West, arXiv:1010.0245

Suppose

- $X$  only feebly coupled to visible-sector thermal bath particles  $V_i$
- $X$  never in thermal equilibrium with SM

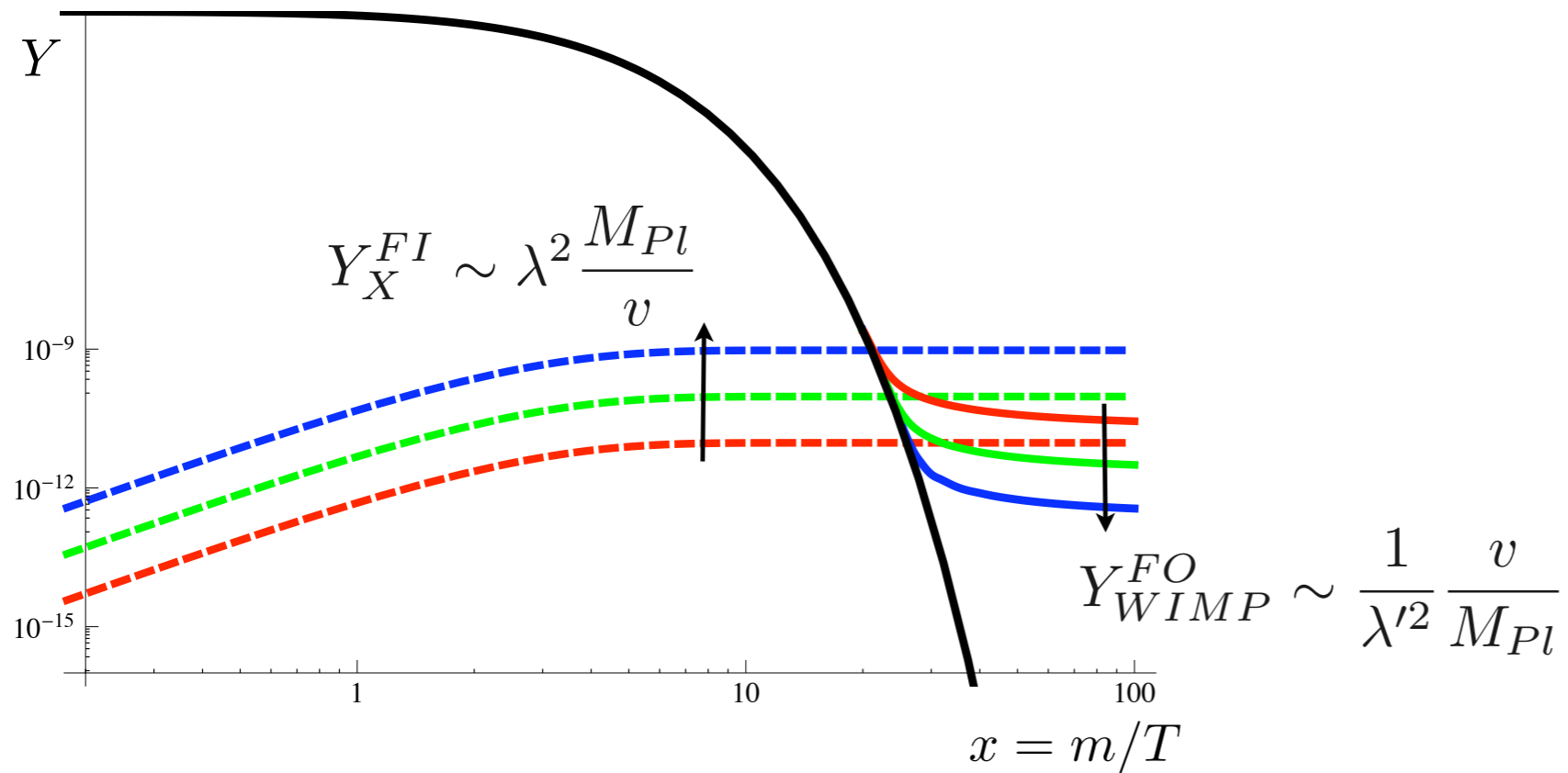
as universe evolves, a tiny  
 $X$  abundance is produced



$$Y_X(t) \sim \Gamma_V t$$

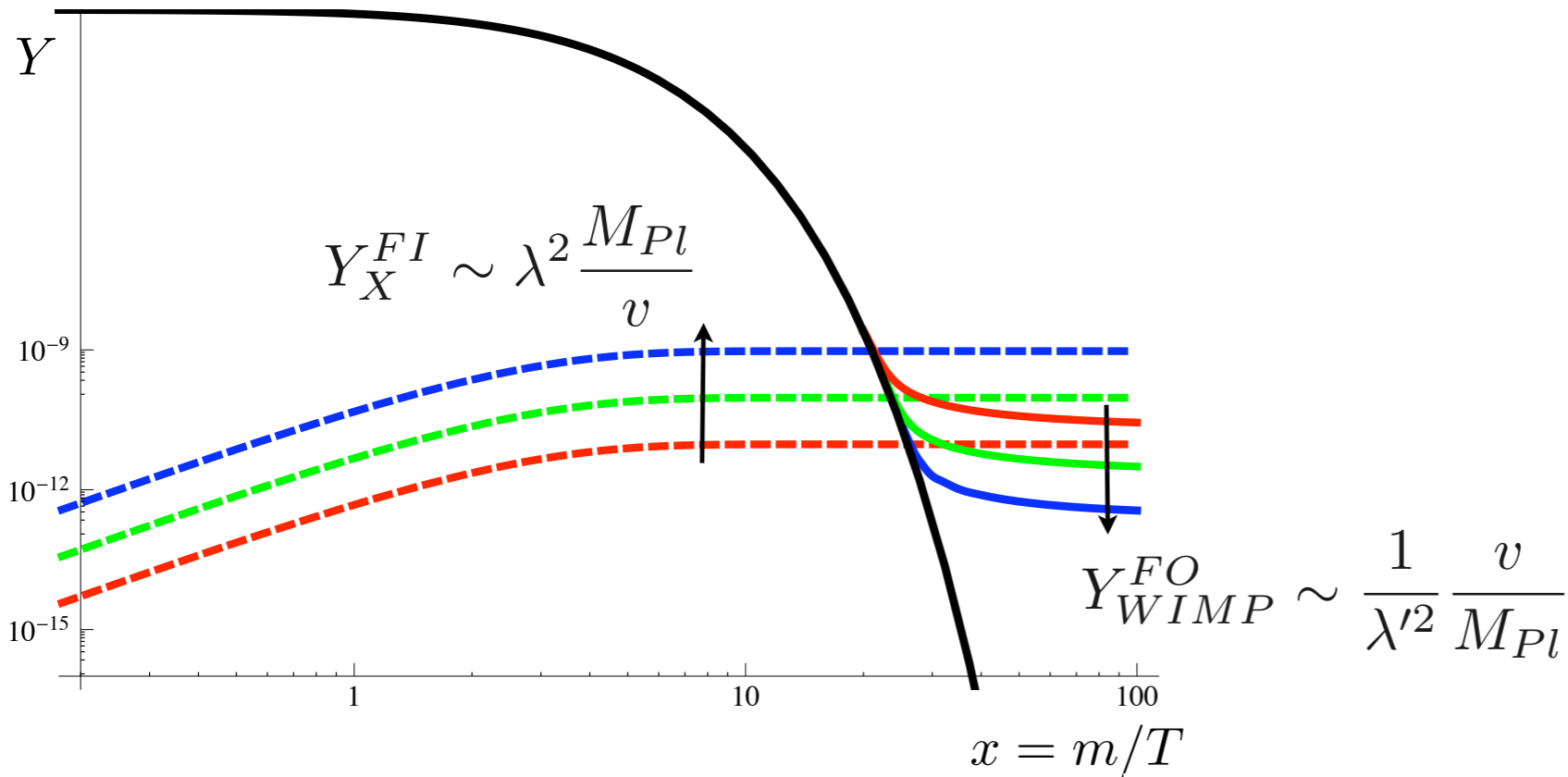
# “Freeze-In” Production

heading either ‘in’ or ‘out’ of equilibrium

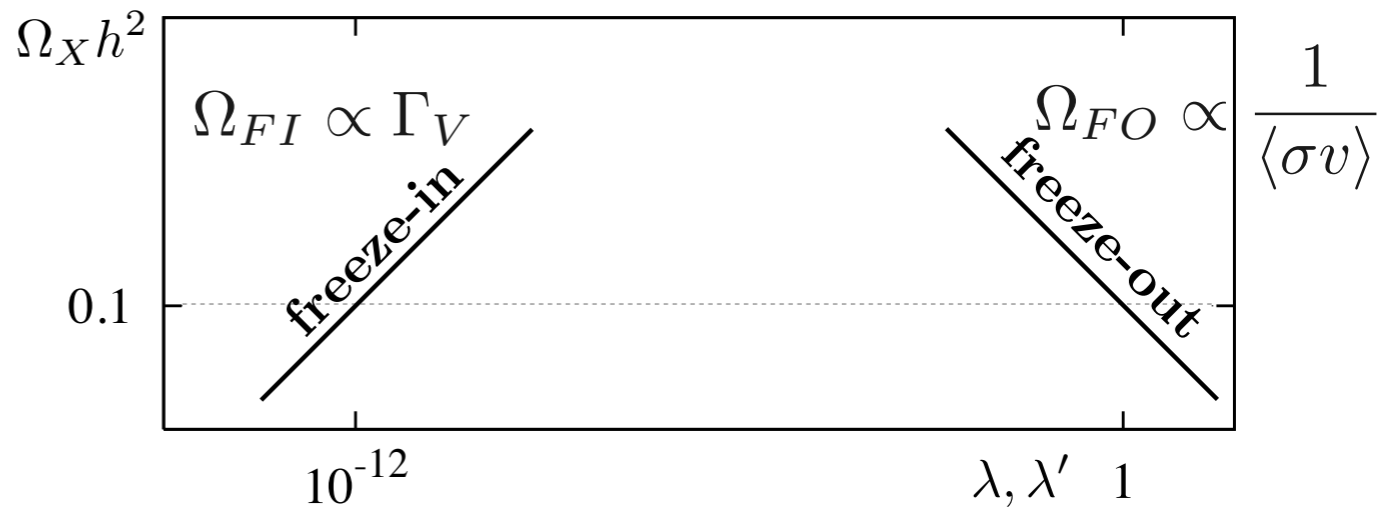


# “Freeze-In” Production

heading either ‘in’ or ‘out’ of equilibrium



**two** thermal mechanisms!



# “Freeze-In” Production

## Comments

- FI yield is **IR-dominated** for renormalizable interactions

$$Y_X^{FI}(T) \sim \lambda^2 \frac{m^2 M_{Pl}}{T^3}$$

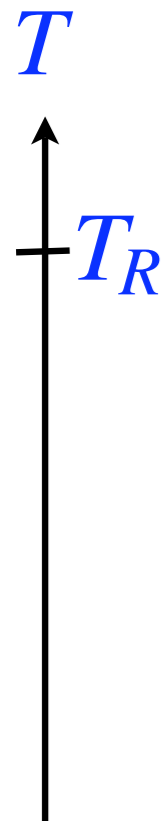
dominant production occurs at  $T \sim m$   
(heaviest particle in vertex)

$$\Delta L = \lambda X V_1 V_2$$

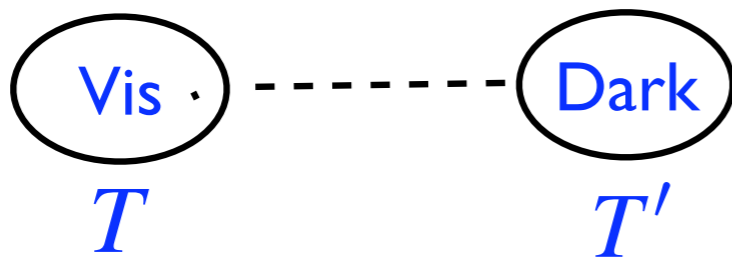
- The lightest ordinary-sector particle (**LOSP**) transforming under the  $X$ -stabilising symmetry is **automatically long-lived**

# Asymmetric “Freeze-In”

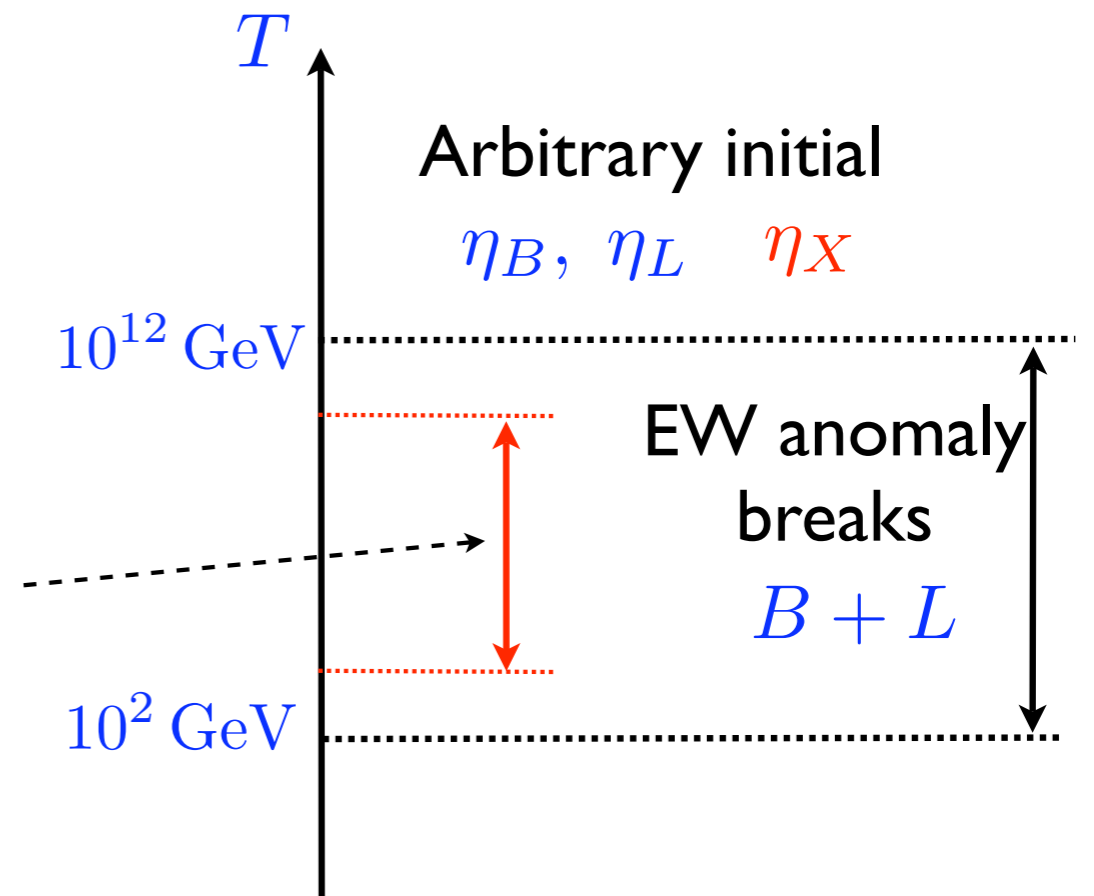
Hall, JMR, West, arXiv:1010.0245



Negligible primordial generation



A feeble “connector interaction” breaks a combination of B/L & X, such that  
There is an era when only conserved U(1) is



$$B - L + X \implies \eta_B : \eta_L : \eta_X = N_1 : N_2 : N_3$$

# Asymmetric “Freeze-In”

Freeze-in production between the sectors via connector interaction leads to related  $\eta_B, \eta_L, \eta_X \neq 0$

if

- temperatures of visible and dark sectors differ (out-of-equilibrium condition)
- CP-violation occurs in decays (and/or scatterings) from visible to dark sectors

# Asymmetric “Freeze-In”

One more condition:

If  $U(1)_{B-L+X}$  only relevant symmetry then depending on relative proton and  $X_0$  masses

either proton or DM not stable

Need additional (discrete) symmetry to stabilise lightest  $X$   
- take to be R-parity in SUSY case

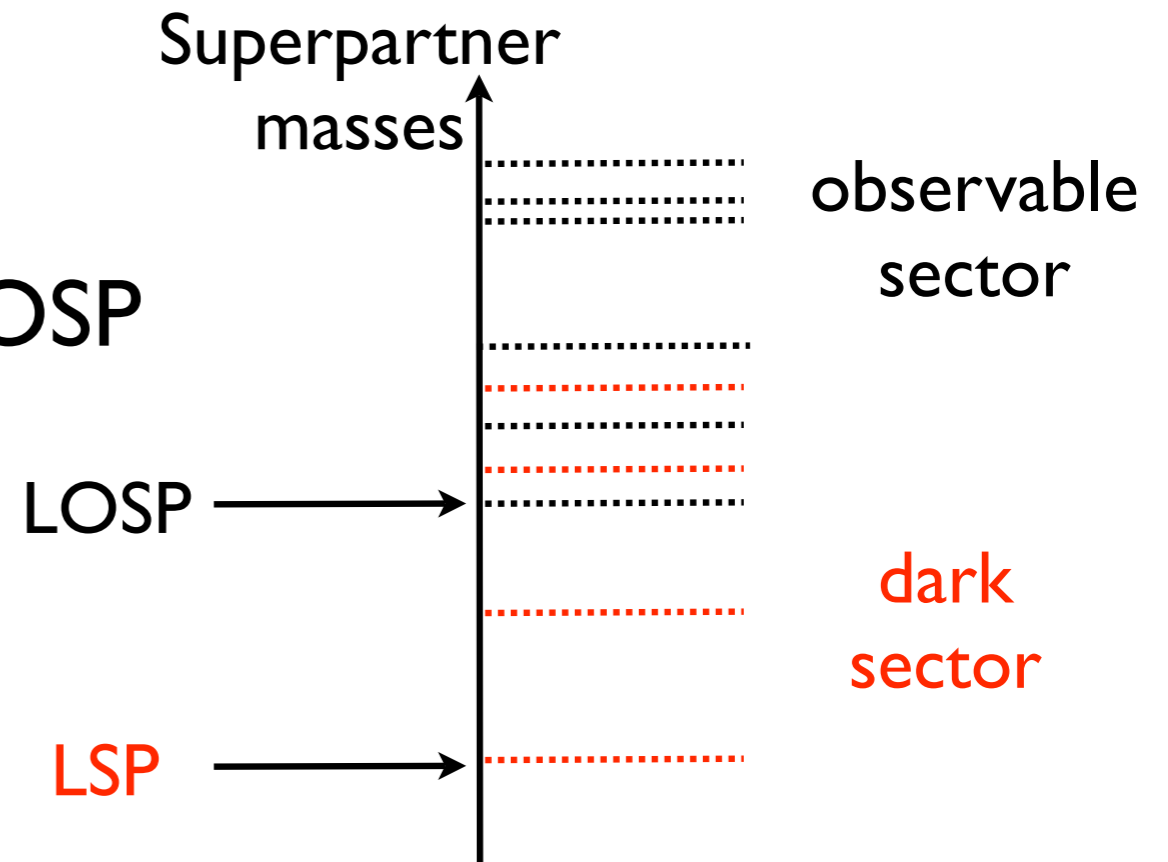


# Asymmetric “Freeze-In”

Supersymmetric example

connector interaction  $\lambda \ll 1$

eg  $W = \lambda L H_u X$  with chargino LOSP

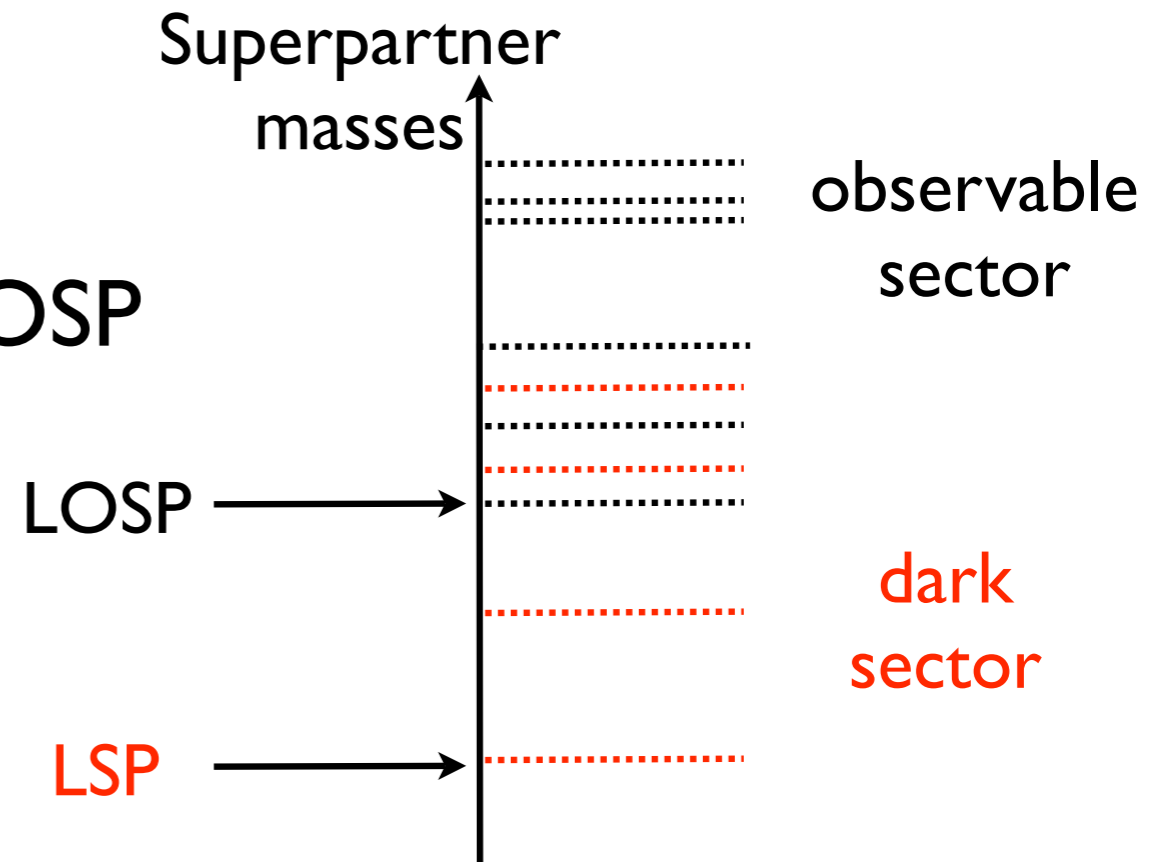


# Asymmetric “Freeze-In”

## Supersymmetric example

connector interaction  $\lambda \ll 1$

eg  $W = \lambda \underbrace{LH_u}_X$  with chargino LOSP



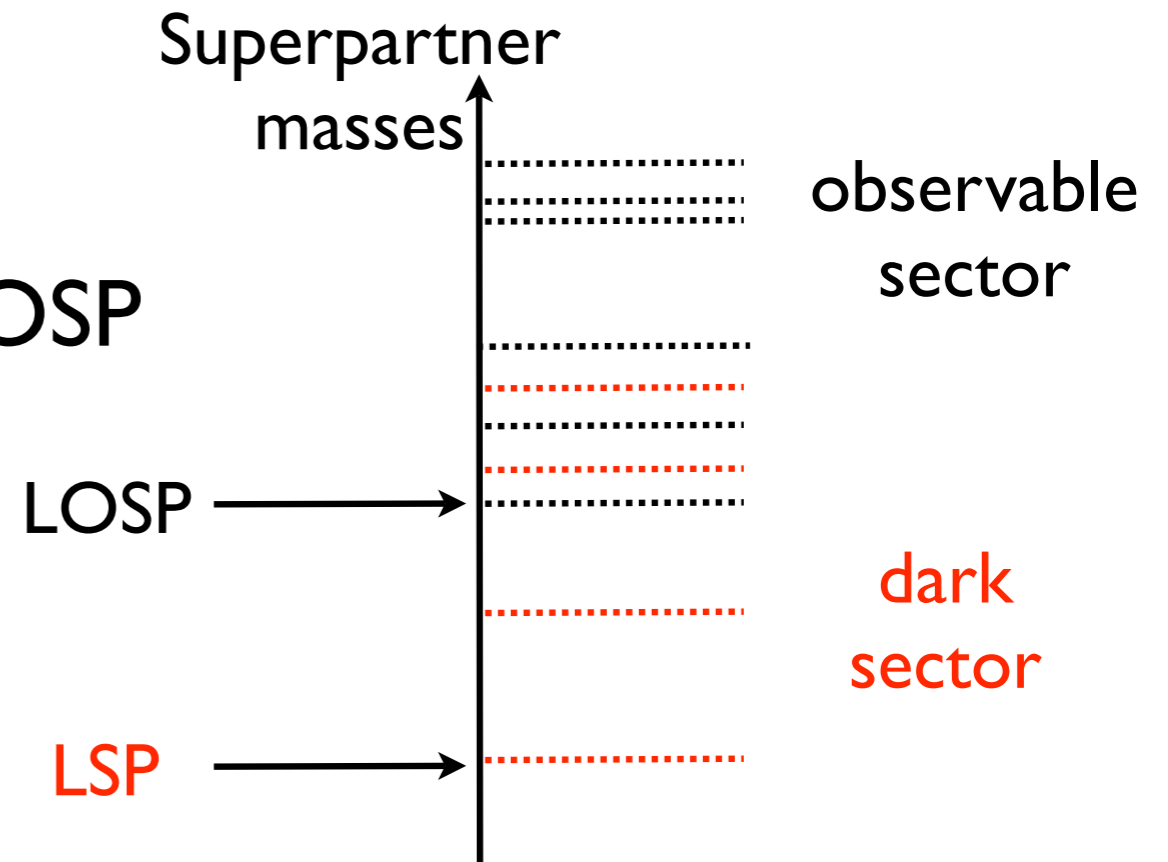
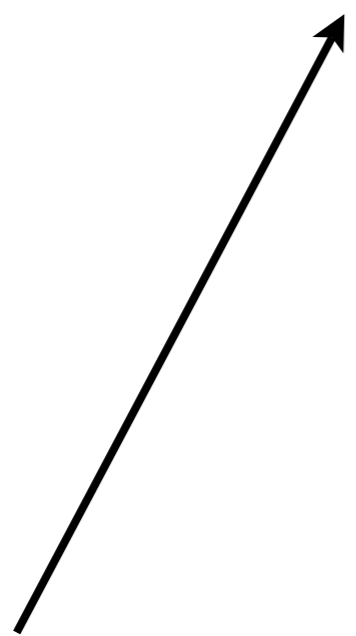
R-parity-odd ops are precisely those that carry B/L - any will do

# Asymmetric “Freeze-In”

## Supersymmetric example

connector interaction  $\lambda \ll 1$

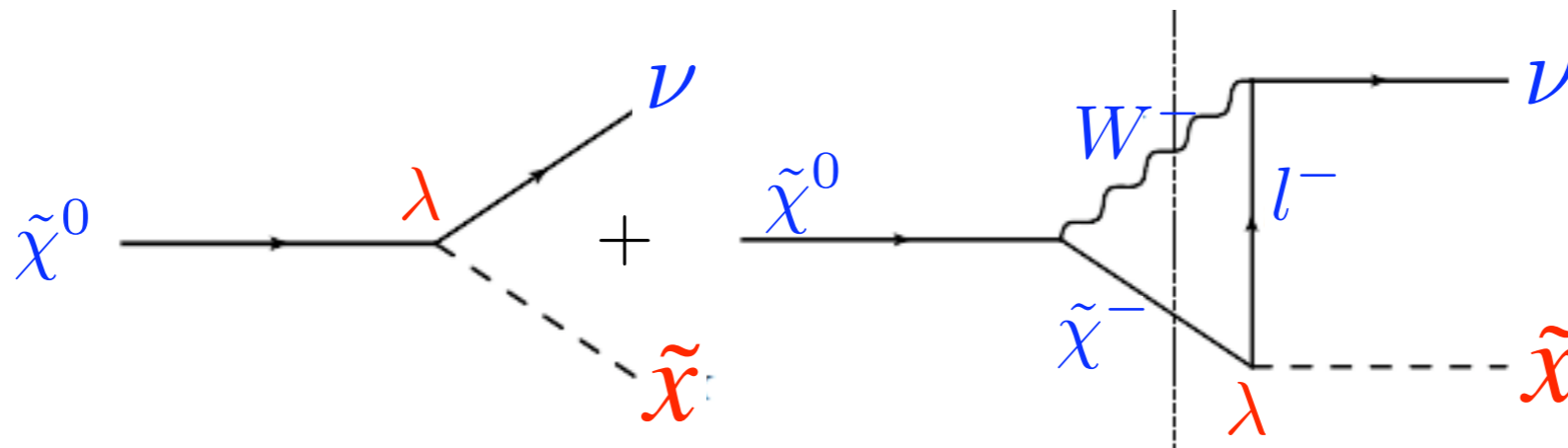
eg  $W = \lambda LH_u X$  with chargino LOSP



linearity in  $X$  is not necessary - other choices change collider & especially DM detection signals

# Asymmetric “Freeze-In”

Decays of NLOSP via on shell LOSP then generate simultaneous L (and thus B) and X asymmetries



$$\eta_L = \eta_X = \epsilon Y_X = \epsilon \Gamma_{\tilde{\chi}^0} \frac{M_{Pl}}{m_{\tilde{\chi}^0}^2}$$

$\epsilon = (\text{loop factor}) \times \sin \phi$  gives fractional CP asymm in decay

# Consequences

DM state can no longer be a real scalar or Majorana fermion

⇒ if DM is the LSP it must be in the dark sector

⇒ must append MSSM with new states and interactions

both very natural from a higher-theory perspective

Observable sector LSP (the LOSP) can be any one of

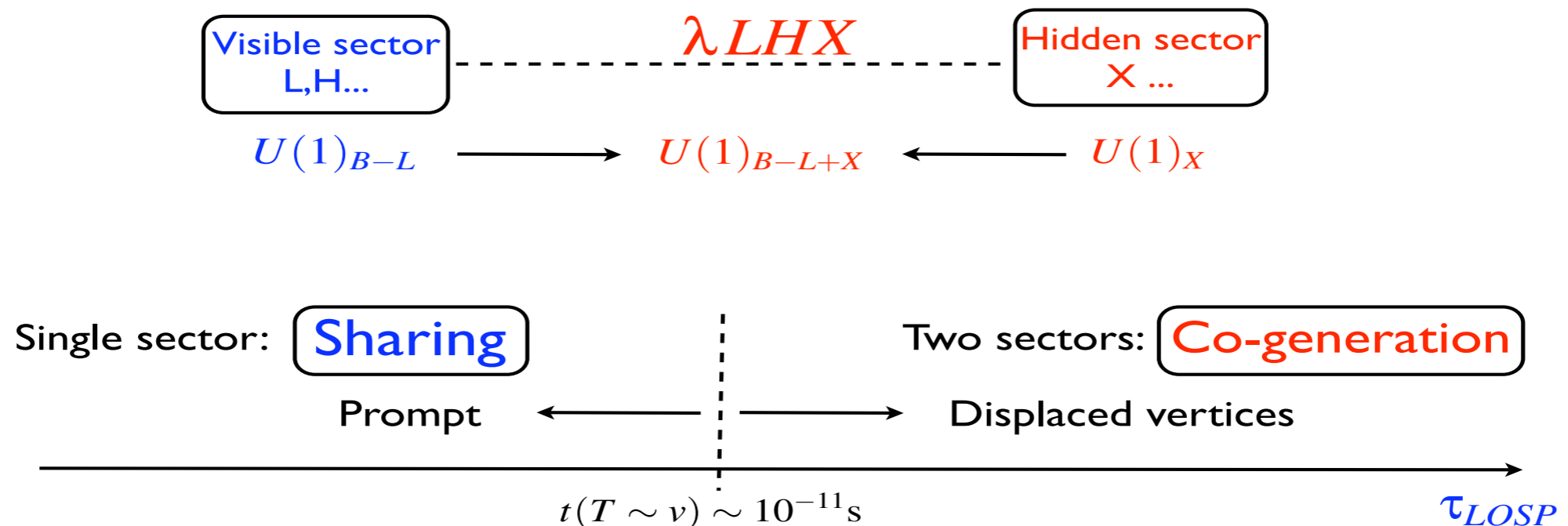
chargino, squark, slepton, neutralino, sneutrino,...

# Consequences

$\epsilon$  is an important parameter

- can be correlated with/deduced from EDM and SUSY-spectrum measurements (work in progress, Hall, JMR, Unwin, West,...)
- determines LOSP lifetime measurable at LHC

$$\tau_{LOSP} \sim 10^{-3} \epsilon r \text{ sec}$$

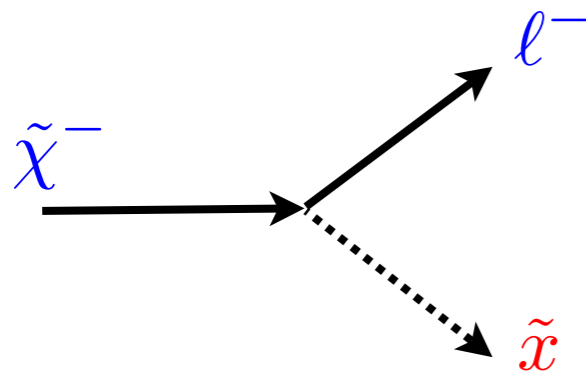


Note that  $10^{-2} \gtrsim \epsilon \gtrsim 10^{-8}$  so sectors don't thermalise

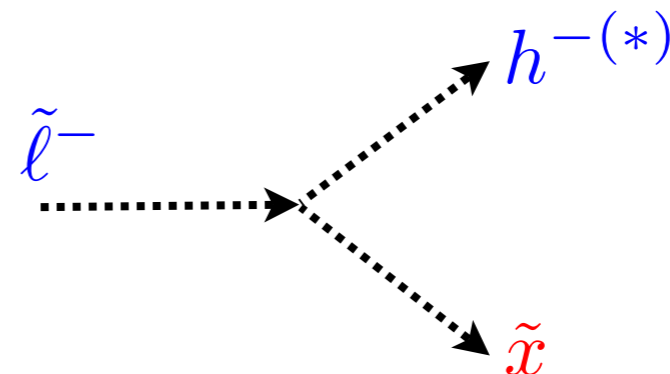
# Consequences

Observable sector LSP (the LOSP) will decay in a **B/L-violating** way

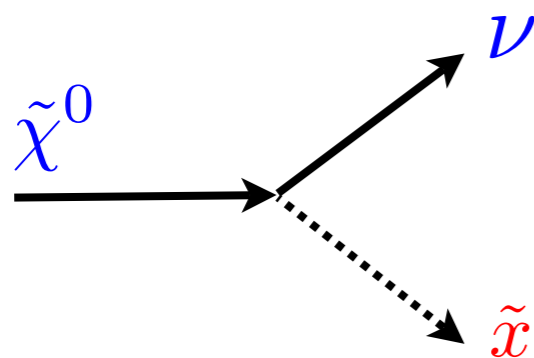
eg, chargino LOSP



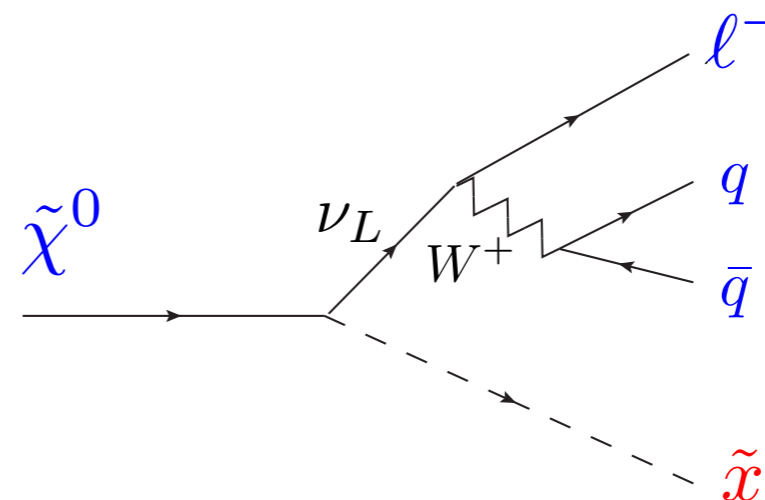
eg, slepton LOSP



eg, neutralino LOSP



but



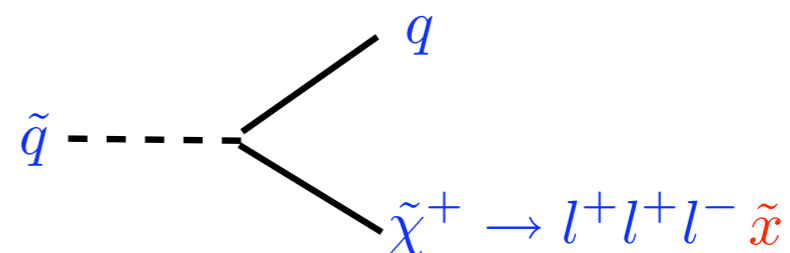
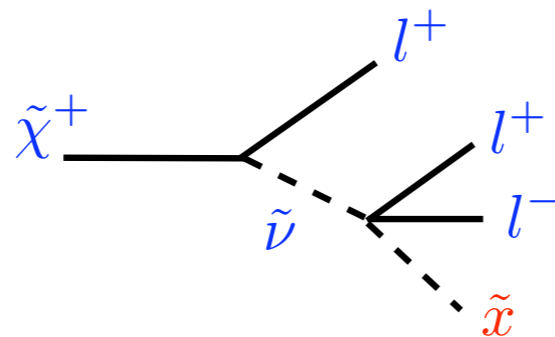
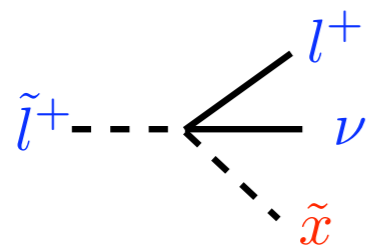
(where we ID the quantum #s of the decaying state by earlier part of cascade chain)

# Consequences

Observable sector LSP (the LOSP) will decay in a **B/L-violating** way

similarly for other connector operators

eg,  $W = \frac{1}{M} L L E^c X$



eg,  $W = \frac{1}{M} U^c D^c D^c X$

$$(\tilde{\chi}^\pm, \tilde{\chi}^0) \rightarrow jjj \quad (\tilde{\ell}^\pm, \tilde{\nu}) \rightarrow \ell^\pm jjj$$

much harder as involves jets in final state and must distinguish  $\dot{J}_q \dot{J}_q \dot{J}_q$  from  $\dot{J}_q \dot{J}_{\bar{q}} \dot{J}_g$

by measuring LOSP mass and lifetime can determine  $\lambda$



# Consequences

Crucial challenge - can we measure  $m_{\tilde{x}}$  ?

Since  $\Omega_{DM}/\Omega_B \simeq 4.86$  and  $\frac{\Omega_X}{\Omega_B} = \frac{\eta_X m_X}{\eta_B m_B}$

$$m_{\tilde{x}} \simeq 4.86 m_p \frac{\eta_X}{\eta_B} \sim 0.5 \div 5 \text{ GeV}$$

depending on the computable asymmetry ratio which is SUSY-spectrum-dependent if connector interaction involves leptons

LHC can bound, but precision of LC likely needed to confirm

# Consequences

## Consequences for DM searches

- light DM  $\sim$  few GeV is strongly favoured
- “sharing” allows normal direct detection, but “co-generation” by asymmetric FI kills direct detection signals (but see JMR+McCullough)
- indirect detection strongly modified - DM can't annihilate to photons but can give rise to anti-B/L final states (see Hall, JMR, Unwin, West,...)

# Consequences

Other generic astro signals

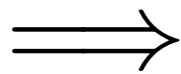
both sharing and co-generation generate an initially dominant symmetric  $(X + \bar{X})$  component

$$\text{eg } Y_X + Y_{\bar{X}} \sim \frac{Y_X - Y_{\bar{X}}}{\epsilon}$$

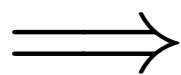
must be efficiently removed - most simply by late annihilation/  
decay to SM states via light X-sector states

Hall, JMR, West, arXiv:1010.0245

& cf Buckley arXiv:1104.1429



late-time energy injection (neutrinos,  
photons, hadronic all possible)



large-scale structure, CMBR, big-bang  
nucleosynthesis signals

# Conclusions

Related baryonic & DM asymmetries, either in sharing or co-generation regime, provides route to understand

$$\Omega_{DM}/\Omega_B \simeq 4.86$$

Asymmetric DM, & especially frozen-in ADM leads to striking collider signals and astro search opportunities

There exists an equally motivated, calculable, and testable mechanism of DM genesis: 'freeze-in'