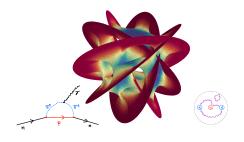
PQ AXIVERSE



Calabi-Yau image by Geoff Fatin

based on [2204.06566] with Mehmet Demirtas, Naomi Gendler, Cody Long, and Liam McAllister

> 07/28/2022 at PASCOS 2022 Jakob Moritz (Cornell)

KEY TAKE AWAYS

We study the Peccei-Quinn mechanism and its quality problem in a vast landscape of type IIB models on Calabi-Yau hypersurfaces in toric varieties.

We find that the quality problem is generically absent for $N \gtrsim 20$ axions, in the geometric regime.

Hierarchical large-N structures of divisor volumes appear to solve the quality problem.

GENERAL REMARKS

general goal: make string theory predictions accessible to low-energy observers (such as us).

Obviously difficult due to $\frac{m_Z}{M_P} \sim 10^{-17} \ll 1$.

Instead of tackling QG feature directly, study the landscape of low-energy theories arising in string theory.

One of the most ubiquitous features of string compactifications:



promising candidate: inflaton, dark matter, ..., solution to strong CP problem

PLAN

- 1. Strong CP problem, PQ-Axion solution, PQ quality problem and all that.
- 2. Evading a quality problem from small QCD instantons.
- 3. Evading a quality problem from stringy instantons.
- 4. Conclusions

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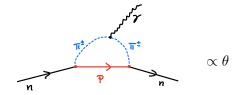
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STRONG CP PROBLEM

Strong CP problem is a tuning problem of the standard model:

$$S_{QCD} \supset \frac{1}{8\pi} \int \frac{\theta}{2\pi} \text{Tr} (G \wedge G)$$
 breaks CP.

Gives electric dipole moment to neutron:



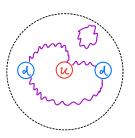
STRONG CP PROBLEM (continued)

Experiment constrains neutron EDM, leading to sharp bound:

$$\theta \lesssim 10^{-10}$$

Strong CP problem:

Why is the neutron so incredibly neutral?!



THE PQ MECHANISM

Axion-solution [Peccei,Quinn'77]: θ is vev of dynamical pseudo-scalar:

$$\mathcal{L} \supset -\frac{f_{\theta}^2}{2} (\partial \theta)^2$$

As θ -dependent non-perturbative potential induced by renormalizable QCD has $\theta = 0$ as its minimum [Vafa, Witten'84],

$$V_{QCD}(\theta) = \frac{1}{2} \frac{m_u m_d}{(m_u + m_d)^2} \Lambda_{QCD}^4 \theta^2 + \mathcal{O}(\theta^4)$$

axion will relax dynamically to

$$\langle \theta \rangle = 0$$

THE PQ QUALITY PROBLEM

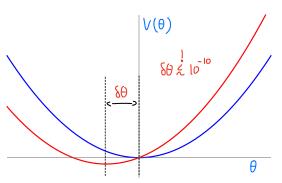
The axion solution comes with its own problem: one has to assume that the "pre-QCD" axion-potential is very small:

$$V(\theta) = V_{QCD}(\theta)$$
 , $V(\theta)$

THE PQ QUALITY PROBLEM

The axion solution comes with its own problem: one has to assume that the "pre-QCD" axion-potential is very small:

$$V(\theta) = V_{QCD}(\theta) + V_{hidden}(\theta)$$
,



THE PQ QUALITY PROBLEM

$$V_{hidden}(\theta) \stackrel{!}{\lesssim} 10^{-10} \Lambda_{QCD}^4 \sim e^{-200} M_P^4$$

or highly CP preserving, (or a combination of both).

At first glance, this looks just as bad as the original strong CP problem...

STRING THEORY TO THE RESCUE?

In string theory, dynamical axions coupling to non-abelian gauge groups are ubiquitous.

Under mild restrictions axion shift symmetries are broken only non-perturbatively, thus abating the PQ quality problem

[Dine'86; Conlon'06; Svrcek, Witten'06;...]

But: Need to check that this actually works!

Goal for the rest of this talk:

Show that PQ quality problem is generically evaded in type IIB Calabi-Yau orientifold models, with QCD hosted on stacks of seven-branes.

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STRING THEORY TO THE RESCUE?

PQ quality problem can come from two (related) sources:

- 1. UV limit of QCD gauge instanton. Rather model-independent, i.e. can study without explicit reference to string compactifications.
- 2. Stringy instantons. Meaningful analysis requires string theory.

 \longrightarrow will study both in turn.

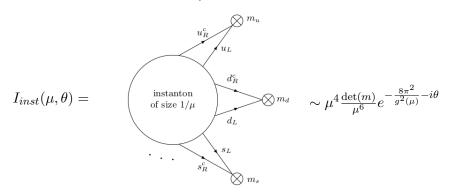
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A QUALITY PROBLEM FROM QCD ITSELF ?

Gauge-instantons in renormalizable QCD induce potential

$$V_{QCD}(\theta) = \int \frac{d\mu}{\mu} I_{inst}(\mu, \theta) + c.c.$$

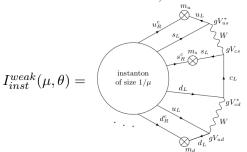


and this is minimized at $\theta = 0$:

A QUALITY PROBLEM FROM QCD ITSELF ?

But QCD as an EFT below weak scale is not renormalizable.

CP breaking in the weak interactions already shifts the θ -minimum,

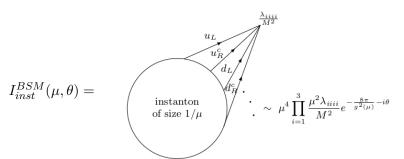


Happily, this is within the allowed region: $\langle \theta \rangle \to 10^{-19} \ll 10^{-10}$

A QUALITY PROBLEM FROM QCD ITSELF ?

Viewing the SM as an EFT with cutoff M we are allowed to consider further sources of CP breaking from dimension six operators, e.g.

$$S_{eff} \supset \int d^4x \frac{\lambda_{ijkl}}{M^2} \mathcal{O}_6^{ijkl} + c.c. \,, \quad \mathcal{O}_6^{ijkl} = \epsilon_{ab} D^i Q^{j,a} U^k Q^{l,b}$$



Whether or not this generates a sufficiently small $V_{hidden}(\theta)$ depends on the UV!

THE UV QCD POTENTIAL

Let us make the following (mild?) assumptions:

- 1. The standard model is supersymmetrically completed at some scale M_{SUSY} . No further light states charged under QCD.
- 2. 4d EFT breaks down at some UV scale M, where CP becomes fully broken.

"SUSY-desert scenario"

Following the (one-loop) RG flow we find

[Demirtas, Gendler, Long, McAllister, JM'21]

$$\delta V_{hidden}(\theta) \sim 10^{-12} \left(\frac{1 \text{ TeV}}{M_{SUSY}}\right)^3 \Lambda_{QCD}^4 \left(1 - \cos(\theta + \phi_0)\right)$$

(up to factor logarithmic in M/m_Z)

THE UV QCD POTENTIAL (continued)

Some upshots:

- ► Generic CP breaking in the deep UV almost spoils the PQ mechanism for low SUSY breaking scale!
- ▶ High SUSY breaking scale suppresses the effects of small instantons because the QCD coupling runs faster without SUSY.
- ► Adding light vector-like pairs endangers the PQ solution more, and can spoil it.

In SUSY-desert scenario small instantons do not lead to a quality problem.

Rest of talk: Assume this scenario, and inquire about stringy instantons in the type IIB landscape.

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A STRINGY SETUP

To fully assess PQ quality, need to restrict to a workable part of the string landscape:

► Consider type IIB O3/O7 orientifolds of Calabi-Yau threefolds X, with QCD realized on stack of seven-branes wrapping divisor D, at large volume:

▶ Chern-Simons action of D-branes implies that

$$\theta = 2\pi \int_D C_4,,$$

so presence of QCD axion is automatic.

- ▶ In general, θ is linear combination of $h^{1,1}(X)$ independent axions $\theta = \sum_{A=1}^{h^{1,1}} q_A \xi^A$.
- $V_{hidden}(\xi)$ is generated by euclidean D3-branes wrapped on divisors (this includes the UV-limit of the QCD instanton).

STRATEGY

- ightharpoonup Consider the Kreuzer-Skarke dataset of Calabi-Yau hypersurfaces X in toric fourfolds V.
- For each CY scan over distinct choices of QCD-divisors D in set of $h^{1,1}(X) + 4$ toric divisors intersecting hypersurface.
- ▶ Assume $M_{SUSY} = 1$ TeV to endanger PQ mechanism most.
- ▶ Dilate overall volume to reproduce correct QCD coupling at low energies.
- ▶ compute all BPS instanton actions and demand that they do not collectively spoil $\langle \theta \rangle \lesssim 10^{-10}$.

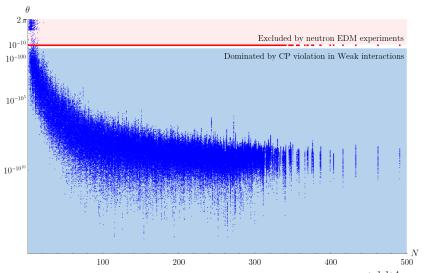
OUR ENSEMBLE OF GEOMETRIES

With above scheme we studied:

- ▶ 32,040 CY's with $2 \le h^{1,1} \le 491$.
- ▶ up to 10 polytopes per $h^{1,1}$, up to 10 triangulations per CY.
- ▶ up to 5 choices of QCD divisor.

Total ensemble size: 136,659 distinct models.

RESULTS



No PQ quality problem for $h^{1,1} \gtrsim 20$: $\delta\theta_{\rm stringy} \sim e^{-1.8 \times (h^{1,1})^4}$

COMMENTS

Why does the PQ mechanism work so well at large $h^{1,1}$?

- ▶ Generating a quality problem in a many-axion theory requires existence of instantons with charges q_1, \ldots, q_n , all more relevant than UV-QCD gauge instanton, and such that $q_{QCD} \in \text{span}(q_1, \ldots, q_n)$.
- ▶ This does happen sometimes, even at large $h^{1,1}$, but once this occurs the smallest divisor volume is very small, thus outside of regime of computational control.

Absence of quality problem explained by hierarchical structures found in Calabi-Yau's at large $h^{1,1}$!

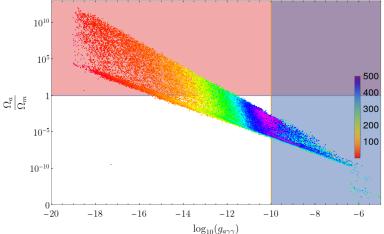
CONCLUSIONS

- ▶ Studied the Peccei-Quinn mechanism and its potential quality problem in type IIB landscape.
- ▶ PQ quality is endangered most for low SUSY breaking scale. For $M_{SUSY} = 1$ TeV, expect $\theta_{QCD} \sim 10^{-12}$ due to small gauge instantons.
- For $N \gtrsim 20$ axions stringy instantons are negligible and thus do not lead to a quality problem.
- Aside: Many models appear to be consistent with both dark matter and astrophysical exclusion bounds.

THANKS!

COSMO AND ASTROBOUNDS

Aside: In our ensemble we have also computed dark matter relic abundances, and checked astrophysical constraints (CAST):



Many models appear not ruled out by either!

VECTORLIKE PAIRS

More generally, consider adding n vector-like pairs in $3 + \overline{3}$. Then, small instantons can lead to a quality problem:

