

Novel constraints on neutrino physics beyond the standard model from the CONUS experiment

Thomas Rink (on behalf of the CONUS Collaboration)

Max-Planck-Institut für Kernphysik (MPIK), Heidelberg

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Coherent elastic neutrino-nucleus scattering (CEvNS)

CEvNS chronology:

- Daniel Freedman (1974): weak (SM) NC, flavor-blind, threshold-free!
- First detection with π -DAR ν 's: COHERENT CsI (2017) & LAr (2021)
- Reactor experiments (2019 - ...):
CONNIE (Si), CONUS (Ge), NCC-1701 (Ge), ν GEN (Ge)
- Several running and future experiments:
CCM (Ar), Miver (Ge/Si), NEON (NaI[Tl]), ν -cleus (CaWO₄, Al₂O₃),
RED100 (Xe), Ricochet (Ge/Zn), Texono (Ge)

The Channel:

- Coherence = enhancement $\sim N^2$

Upper limit on neutrino energy:

$$E_\nu \leq \frac{1}{2R_A} \approx \frac{197}{2.5 \sqrt[3]{A}} [\text{MeV}]$$

- Observable = nuclear recoil energy T_A

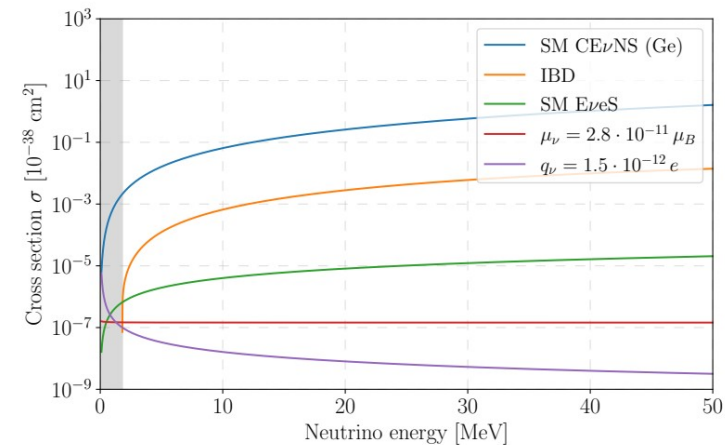
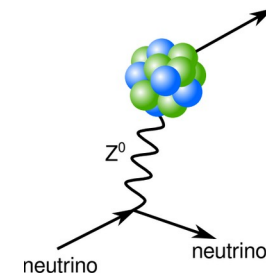
$$\frac{d\sigma}{dT_A} = \frac{G_F^2 m_A}{4\pi} \left[(1 - 4 \sin^2 \theta_W) Z - N \right]^2 \left(1 - \frac{m_A T_A}{2E_\nu^2} \right) F^2(T_A)$$

→ detector material
→ neutrino source
→ energy threshold / neutrino source

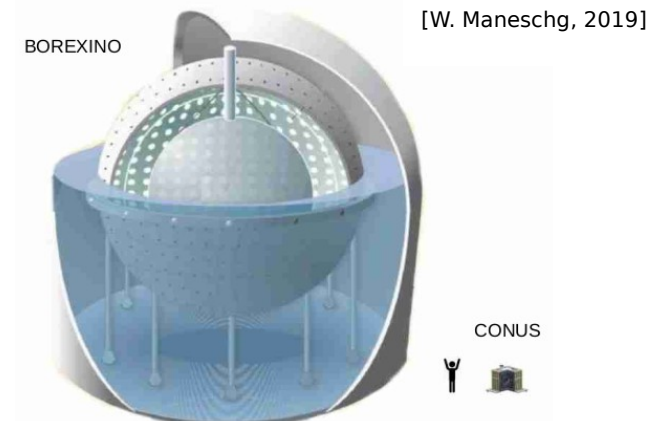
- Very low energy threshold needed:

- $T_A \sim N^{-1}$
- Quenching: $T_A \rightarrow$ "detectable" energy

Cross section σ
vs.
nuclear recoil T_A



[TR, 10.11588/heidok.00031274, 2022]

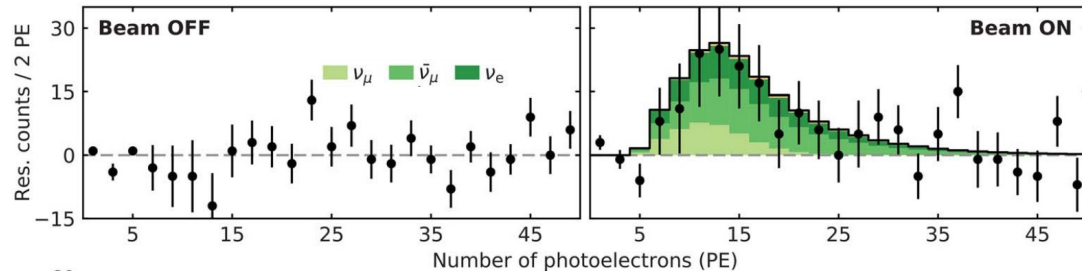


CEvNS observations and constraints

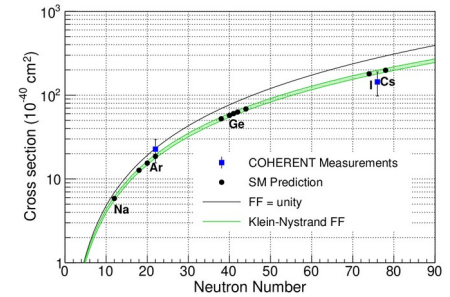
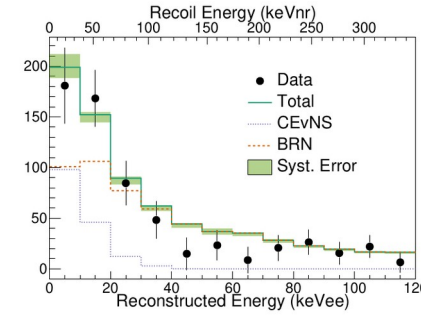
Observations:

COHERENT CsI[Na] (2017): $5 * 10^{20}$ POT
obs. 134+-22, pred. 173+-48

[Akimov et al., [10.1126/science.aao0990](https://doi.org/10.1126/science.aao0990), 2017]



COHERENT LAr (2021): $13.7 * 10^{22}$ POT
obs. 159+-43, pred. 128+-17 [Akimov et al., [10.1103/PhysRevLett.126.012002](https://doi.org/10.1103/PhysRevLett.126.012002), 2021]

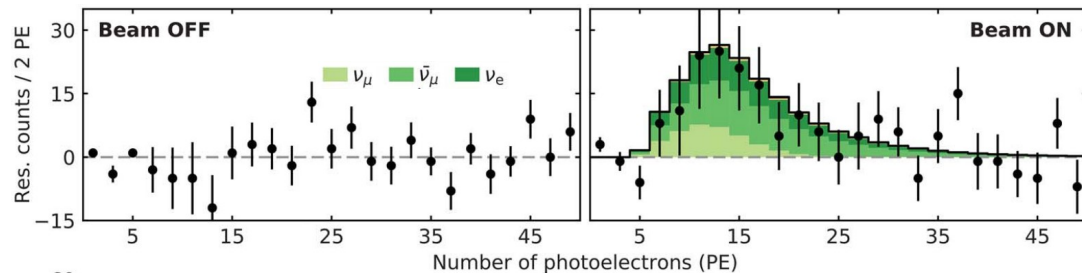


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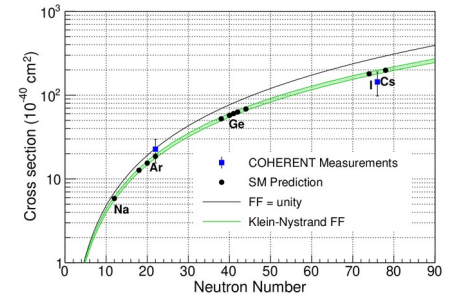
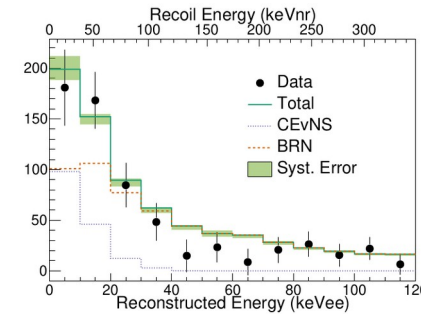
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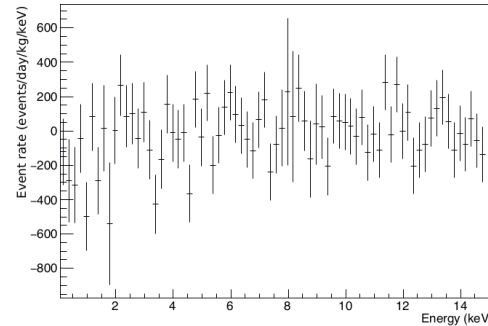


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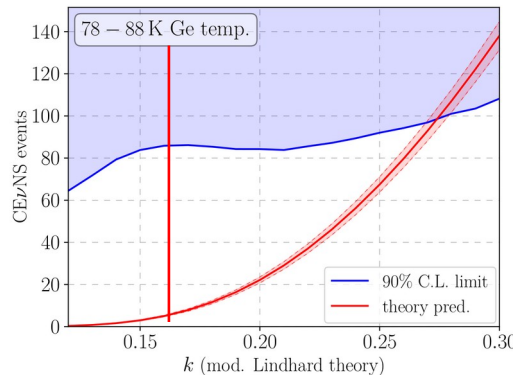


Limits:

CONNIE Si (2019):
 $R_{NP} < 40 * R_{SM}$ @ 95% C.L.
 2.1kg*d ON + 1.6kg*d OFF
 [Aguilar-Arevalo et al., [10.1103/PhysRevD.100.092005](https://doi.org/10.1103/PhysRevD.100.092005), 2019]

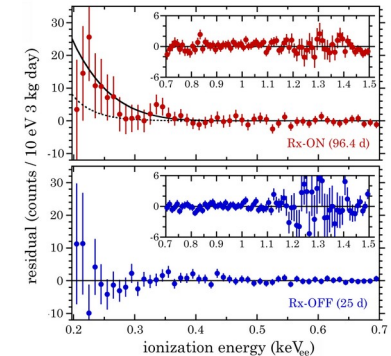


CONUS Ge (2020):
 $R_{SM} < 0.4/\text{kg/d}$ @ 90% C.L.
 (x17 lower than prediction)
 249kg*d ON + 59kg*d OFF
 [Bonet et al., [10.1103/PhysRevLett.126.041804](https://doi.org/10.1103/PhysRevLett.126.041804), 2020]

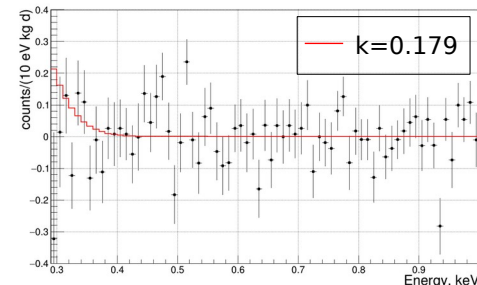


“Suggestive evidence ...”
 [Colaresi et al., [arXiv:2202.09672](https://arxiv.org/abs/2202.09672), 2022]

NCC-1701 Ge:
 289kg*d ON + 75kg*d OFF



ν GEN Ge (2022):
no excess,
 $k < 0.177$ @ 90% C.L.
 133kg*d ON + 66kg*d OFF
 [Alekseev et al., [arXiv:2205.04305](https://arxiv.org/abs/2205.04305), 2022]



CEvNS observations and constraints

Observations:

COHERENT C
obs. 134+2

Coherent Neutrino Nucleus Scattering as a Probe of the Weak Neutral Current

Daniel Z. Freedman (Fermilab and SUNY, Stony Brook)

Oct, 1973

11 pages

Published in: *Phys.Rev.D* 9 (1974) 1389-1392

DOI: [10.1103/PhysRevD.9.1389](https://doi.org/10.1103/PhysRevD.9.1389)

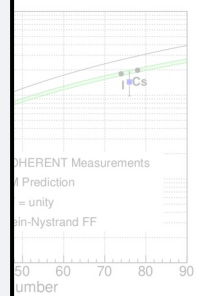
Report number: NAL-PUB-73-76-THY, FERMILAB-PUB-73-076-T

View in: [OSTI Information Bridge Server](#), [ADS Abstract Service](#)

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584 citations

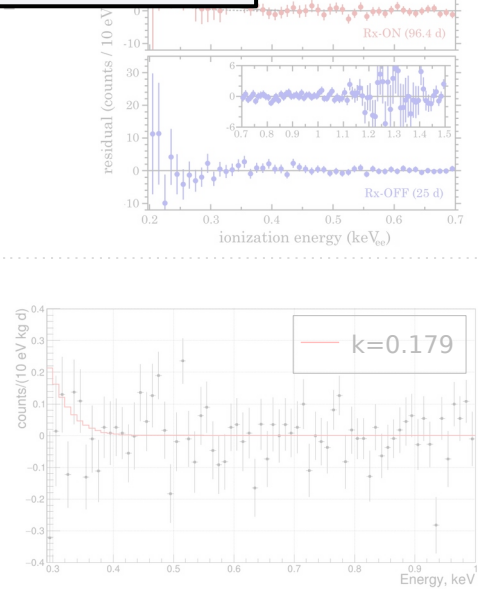
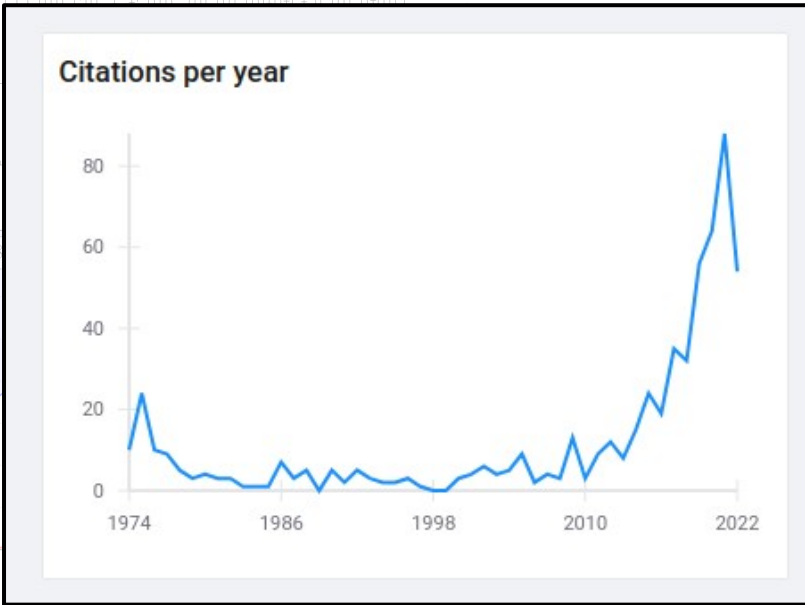
evLett.126.012002,



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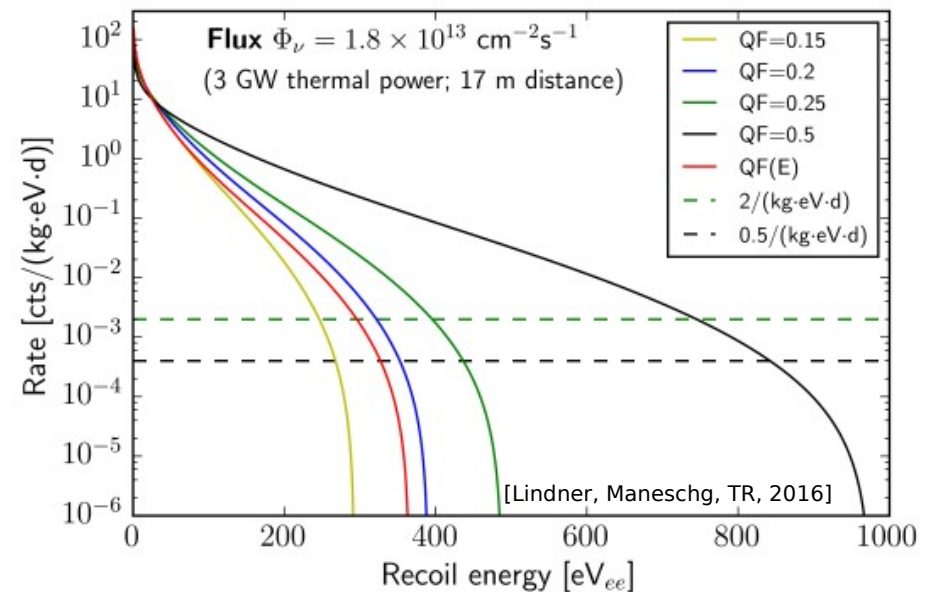
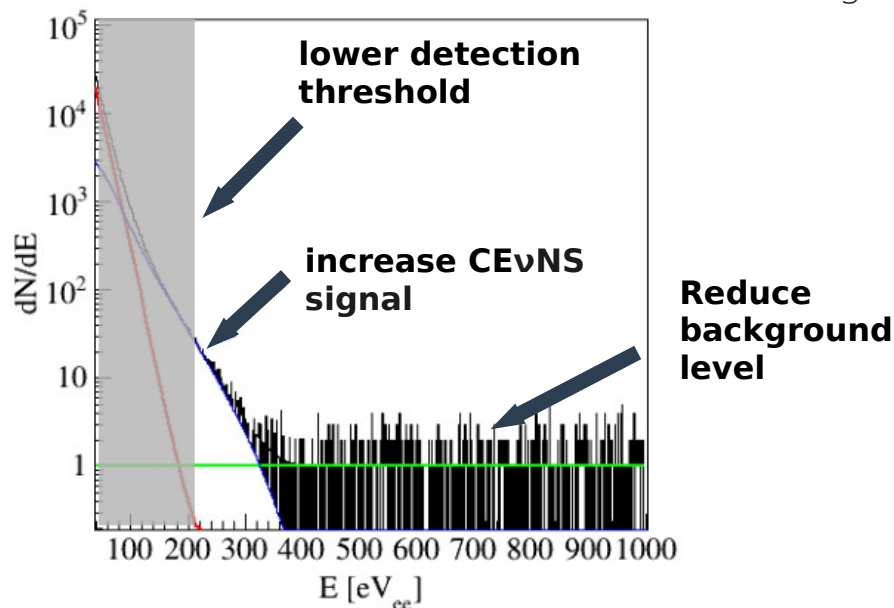
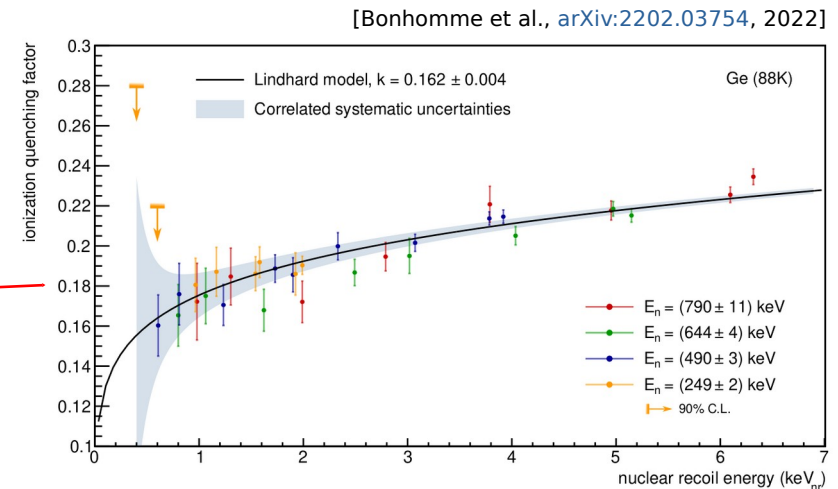
Experimental requirements at reactor site

Goal: Detecting CEvNS with high accuracy!

Several obstacles to overcome:

- 1) Beat $1/R^2$ factor
 - strong (= commercial) power plant, close to reactor core
- 2) Compensate quenching ($E_{\text{recoil}} \rightarrow E_{\text{ion}}$)
 - lowest possible detection threshold
- 3) Low background outside lab condition
 - moderate overburden & limited shielding capacities

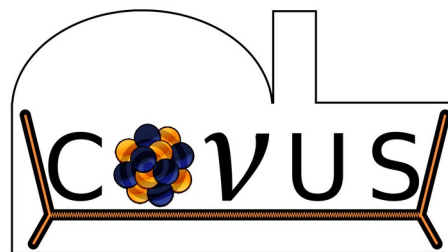
E_ν : 10MeV
 max E_{Recoil} : 3keV
 max E_{ion} : ~600eV



The CONUS Collaboration



Nicola Ackermann, Hannes Bonet, Aurélie Bonhomme, Christian Buck, Kai Fülber, Janina Hakenmüller, Janine Hempfling, Gerd Heusser, Thomas Hugle, Manfred Lindner, Werner Maneschg, Thomas Rink, Edgar Sanchez Garcia, Josef Stauber, Herbert Strecker, Roland Wink



The CONUS experiment

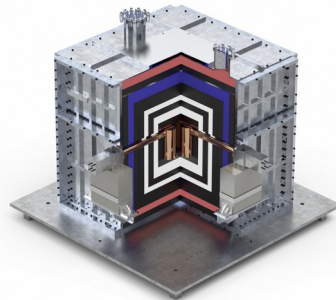
Reactor: Brokdorf nuclear power plant

- @17m to 3.9GW_{th}: inside reactor containment!
- ν flux: $2.3 \cdot 10^{13}/\text{cm}^2/\text{s}$
- Access to reactor information: fuel composition, thermal power

Background & shield

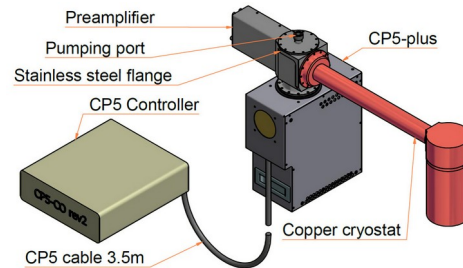
- Overburden: 10-45 m w.e.
- Active and passive components: bkg reduction $\times 10^3$ - 10^4
- Critical reactor-correlated bkg under control!
- Background level $O(10)/\text{keV}/\text{d}/\text{kg}$

[Bonet et al., arXiv:2112.09585, 2021]



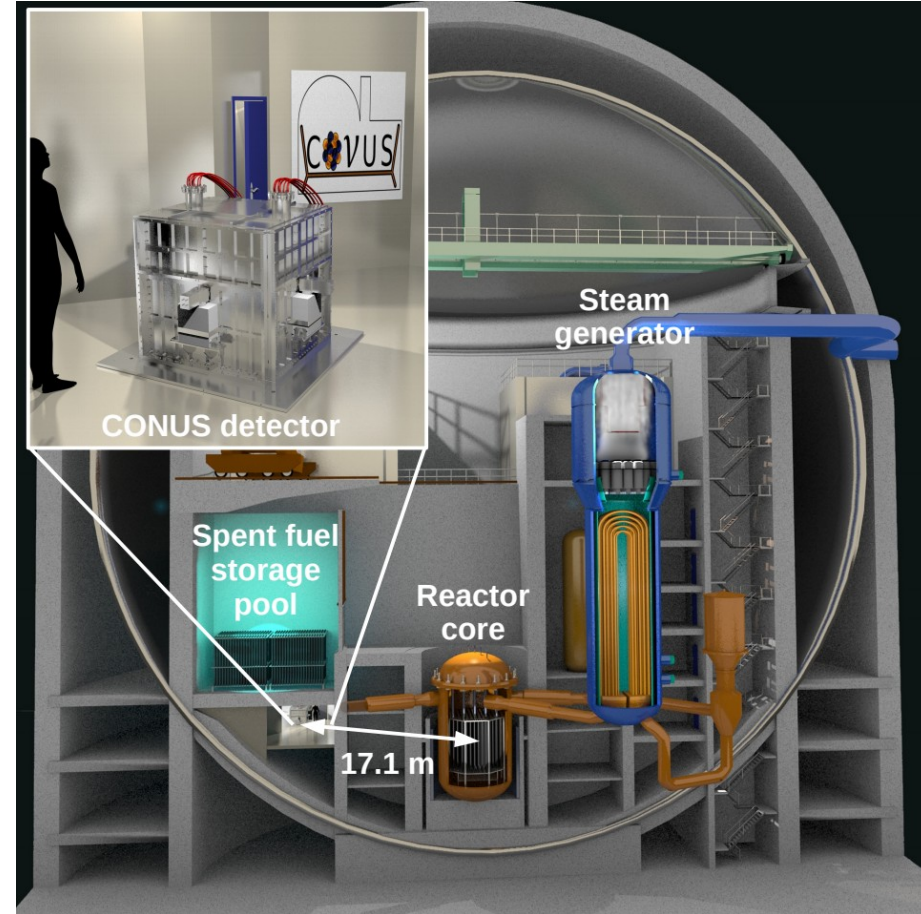
Detectors:

- HPGe PPC: 4*1kg
- $E_{\text{thr}} \sim 300\text{eV} \rightarrow T_A^{\text{thr}} \sim 2\text{keV}_{\text{nr}}$
- Low bkg design
- Electrical cryocooler



[Bonet et al., 10.1140/epjc/s10052-021-09038-3, 2021]

[Bonet et al., 10.1007/JHEP05(2022)085, 2022]



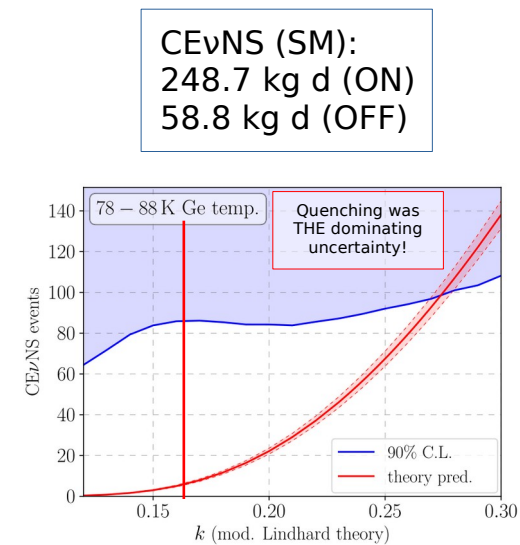
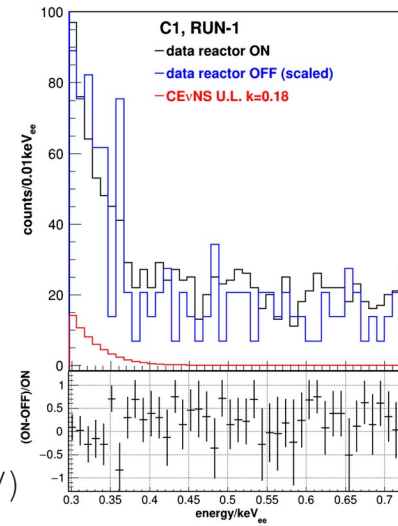
Close to reactor core:
lab conditions not easily realized
→ site monitoring crucial!

Data analyses of experimental data

(Binned log-) Likelihood function:

- ON and OFF data are fitted together:

$$\log \mathcal{L} = \log \mathcal{L}_{\text{ON}} + \log \mathcal{L}_{\text{OFF}} + \text{pull terms}$$
 with $\log \mathcal{L}_{\text{ON}}(s, b, \Theta_{\text{thr}_1}, \Theta_{\text{thr}_2}, \Theta_{\text{rea}}, \Theta_{\text{det}}, \Theta_{\Delta E})$
 $\log \mathcal{L}_{\text{OFF}}(b, \Theta_{\text{thr}_1}, \Theta_{\text{thr}_2}, \Theta_{\text{det}}, \Theta_{\Delta E})$
- Parameter list:
 - s : signal strength (free), b : MC normalisation (free)
 - $\Theta_{\text{thr}_{1/2}}$: electronic noise (free)
 - Θ_{rea} : reactor neutrino flux ($\sim 3\%$)
 - Θ_{det} : detector/DAQ (1-5%), $\Theta_{\Delta E}$: energy scale (10-20 eV)
- Compare signal to background-only hypothesis: **LH ratio**

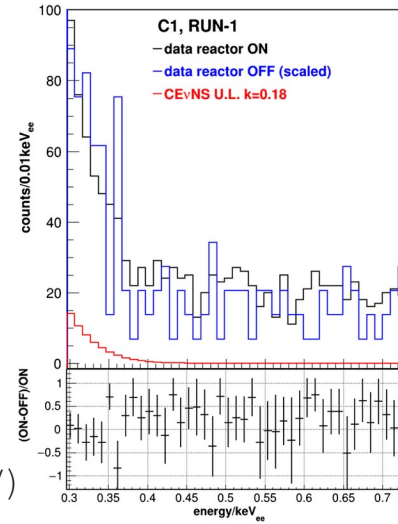


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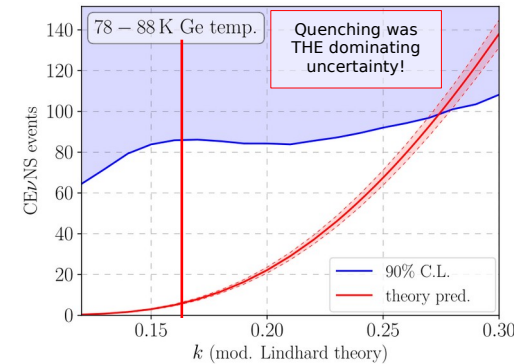
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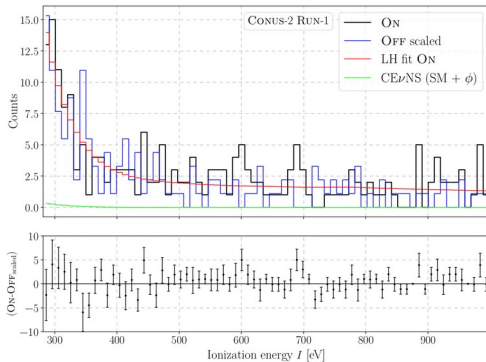
[Bonet et al., 10.1103/PhysRevLett.126.041804, 2020]

CEvNS (SM):
 248.7 kg d (ON)
 58.8 kg d (OFF)



BSM investigations:

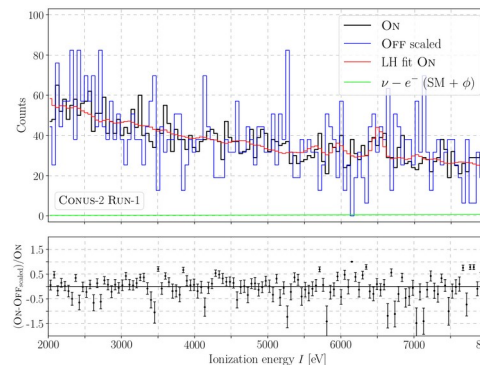
- Extend existing LH with BSM contributions
 \rightarrow SM CEvNS as background



[Bonet et al., 10.1007/JHEP05(2022)085, 2022]

CEvNS (BSM): 208.8 kg d (ON), 37.6 kg d (OFF)

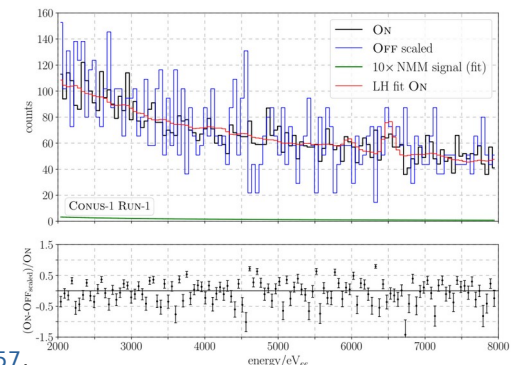
- Add. systematics within [2,8]keV
 \rightarrow 5% shape uncertainty on bkg



[Bonet et al., arXiv:2201.12257, 2022]

EveS (BSM): 648.5 kg d (ON), 93.5 kg d (OFF)

- BSM limits in terms of quenching factor!

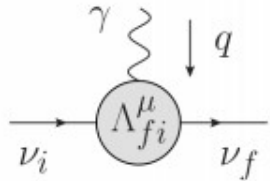


EveS (ν MM): 688.9 kg d (ON), 131.5 kg d (OFF)

Neutrino magnetic moment & millicharge

Loop-induced electromagnetic properties:

[Giunti & Studenikin, 10.1103/RevModPhys.87.531, 2015]



- neutrino charge (radius), electric and magnetic dipole moment, anapole moment
- characteristic dependence on neutrino's fermionic nature
- test for BSM physics!

ν MM from EveS @ 2-8 keV:

$$\left(\frac{d\sigma}{dT}\right)_{\mu\nu}^{e^-} = \frac{\pi\alpha_{em}^2}{m_e^2} \left(\frac{1}{T} - \frac{1}{E_\nu}\right) \left(\frac{\mu_{\nu e}}{\mu_B}\right)^2$$

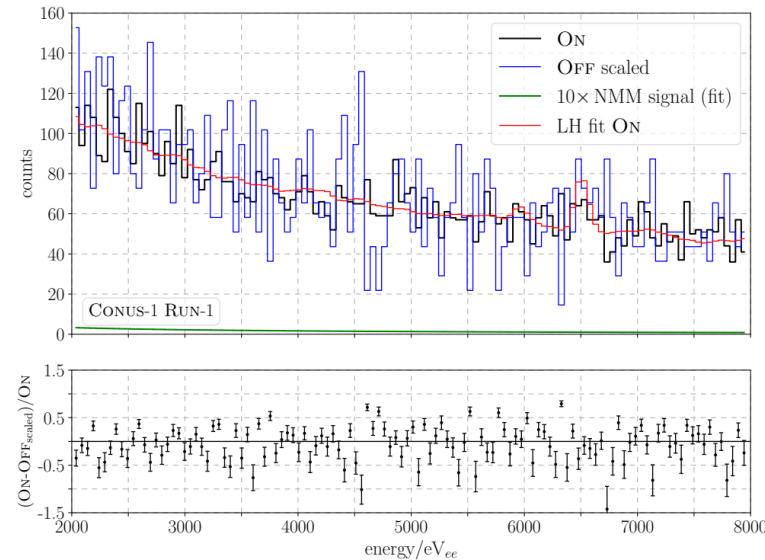
- Model-independent bounds: [Bell et al., 10.1016/j.physletb.2006.09.055, 2006]

$$\mu_{ll'}^D \lesssim 3 \cdot 10^{-15} \mu_B \left(\frac{m_l}{\text{eV}}\right) \left(\frac{\text{TeV}}{\Lambda}\right)^2$$

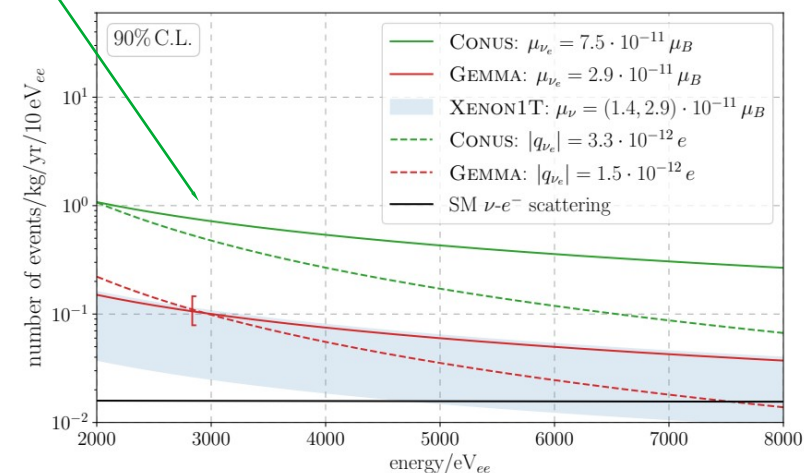
$$\mu_{ll'}^M \lesssim 4 \cdot 10^{-9} \mu_B \left(\frac{M_{ll'}^M}{\text{eV}}\right) \left(\frac{\text{TeV}}{\Lambda}\right)^2 \left|\frac{m_\tau^2}{m_l^2 - m_{l'}^2}\right|$$

- Extended CONUS data sets:

$$\mu_{\nu e} < 7.5 \cdot 10^{-11} \mu_B \text{ (90\% C.L.)}$$



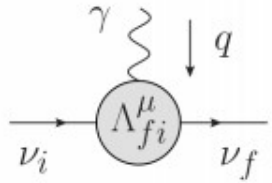
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νMC limit from νMM limit:

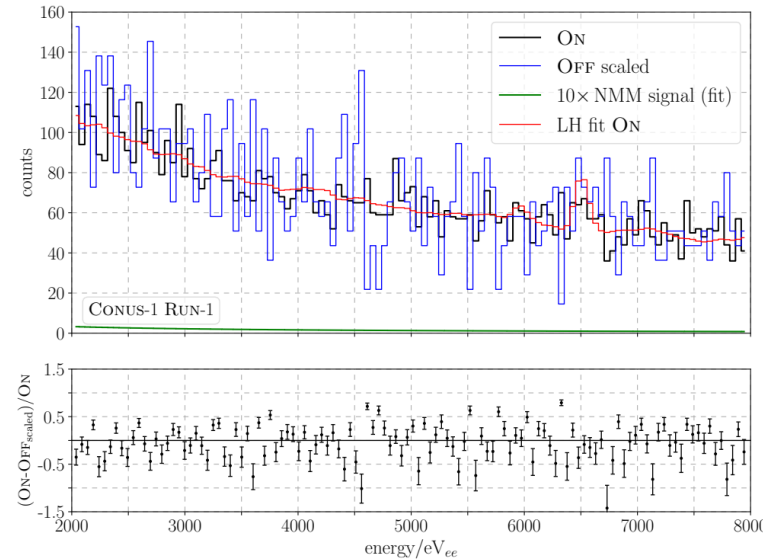
[Giunti & Studenikin, 10.1103/RevModPhys.87.531, 2015]

$$\left(\frac{d\sigma}{dT}\right)_{q\nu}^{e^-} \simeq \frac{2\pi\alpha_{em}^2}{m_e T^2} \left(\frac{q_{\nu e}}{e}\right)^2$$

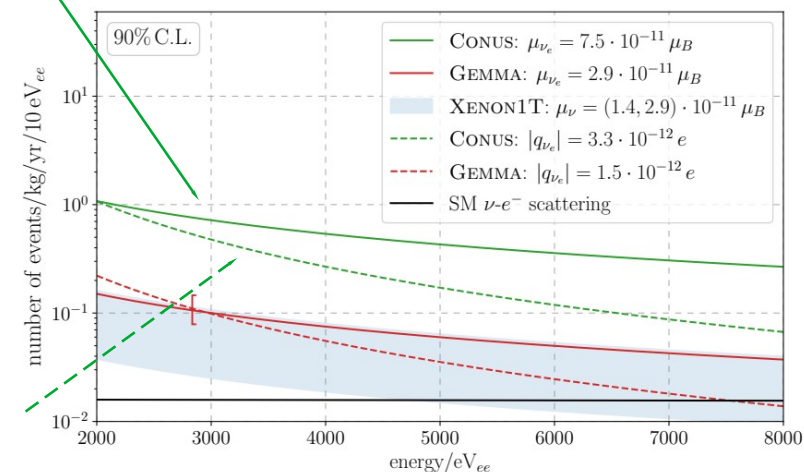
$$\frac{(d\sigma/dT)_q^{e^-}}{(d\sigma/dT)_\mu^{e^-}} \simeq \frac{2m_e}{T} \left(\frac{(q_{\nu e}/e)}{(\mu_{\nu e}/\mu_B)}\right)^2 \leq 1$$

- νMC has different event shape!
- Assume null result within [2,8] keV:
convert νMM into νMC

$$|q_{\nu e}| < 3.3 \cdot 10^{-12} e \text{ (90% C.L.)}$$



[Bonet et al., arXiv:2201.12257, 2022]



Non-standard interactions (NSIs)

Aim: Test for new (high-scale) neutrino interactions

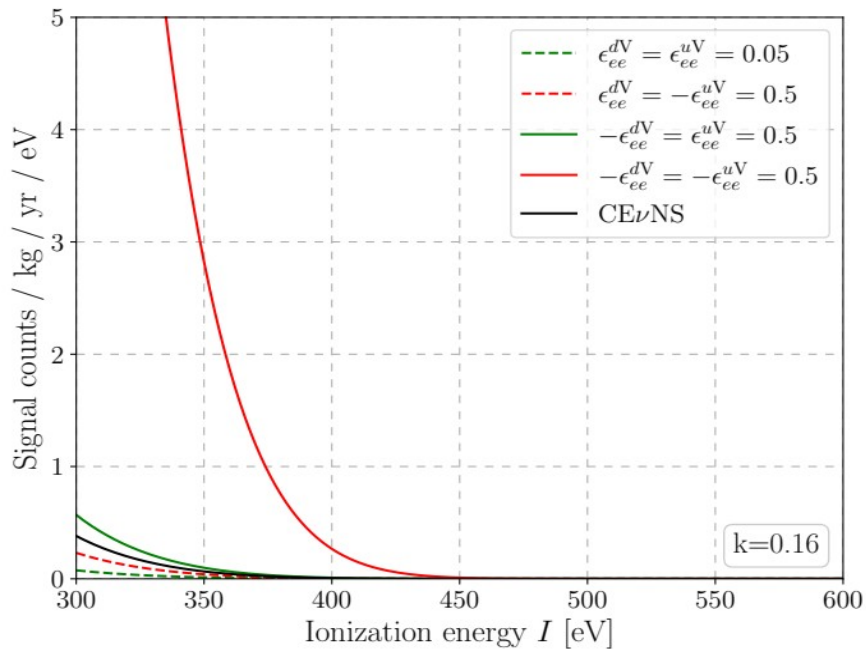
- Vector-type Interaction
→ New feature: destructive interference!
- Modified weak nuclear charge
- New couplings ϵ partially degenerate
→ different isotopes/targets at different sources

[Barranco et al., [10.1088/1126-6708/2005/12/021](https://doi.org/10.1088/1126-6708/2005/12/021), 2005]

$$\mathcal{O}_{\alpha\beta}^{qV} = (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{q} \gamma_\mu P_{L/R} q) + \text{h.c.}$$

$$\mathcal{Q}_{NSI}^V = (2\epsilon_{\alpha\alpha}^{uV} + \epsilon_{\alpha\alpha}^{dV} + g_p^V) Z + (\epsilon_{\alpha\alpha}^{uV} + 2\epsilon_{\alpha\alpha}^{dV} + g_n^V) N + \sum_{\alpha \neq \beta} \left[(2\epsilon_{\alpha\beta}^{uV} + \epsilon_{\alpha\beta}^{dV}) Z + (\epsilon_{\alpha\beta}^{uV} + 2\epsilon_{\alpha\beta}^{dV}) N \right]$$

[Bonet et al., [10.1007/JHEP05\(2022\)085](https://doi.org/10.1007/JHEP05(2022)085), 2022]



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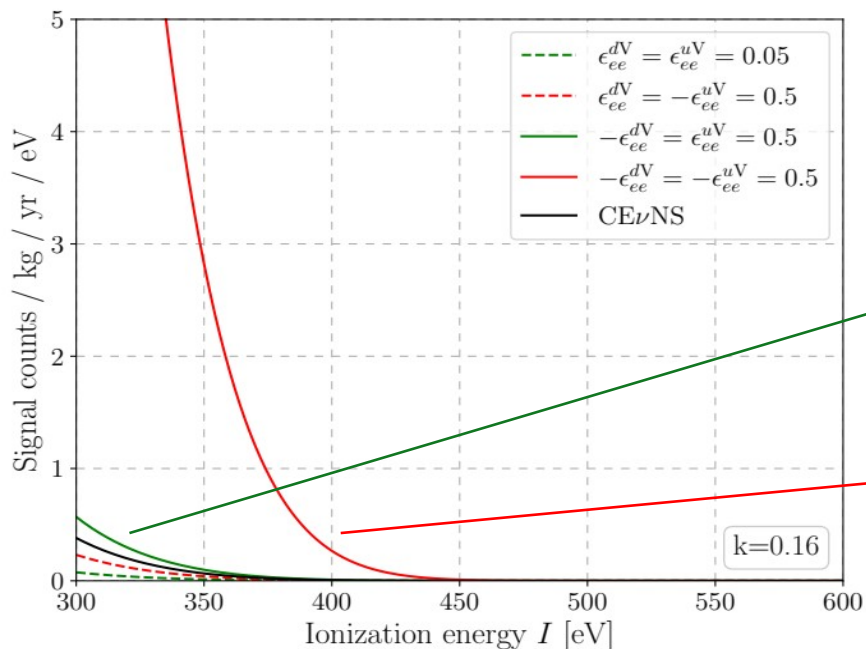
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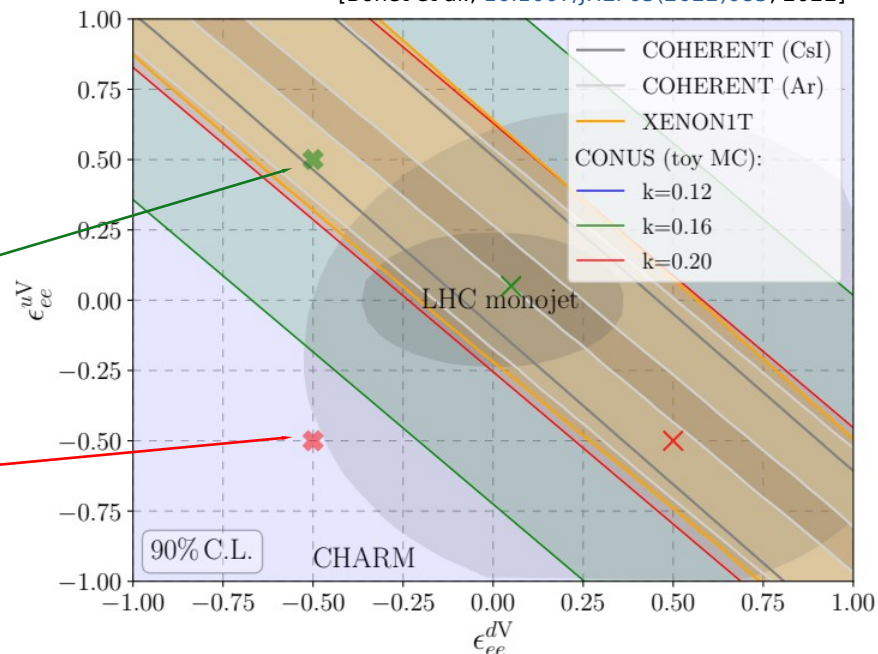
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$$\mathcal{O}_{\alpha\beta}^{qV} = (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{q} \gamma_\mu P_{L/R} q) + \text{h.c.}$$

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[Bonet et al., 10.1007/JHEP05(2022)085, 2022]



If sub-percent sensitivity on ϵ
→ probing **TeV-scale** physics!

[Lindner et al., 10.1007/JHEP03(2017)097, 2017]

$$\epsilon_{\alpha\beta}^{qX} \propto \frac{m_W^2}{\Lambda_{NP}^2}$$

$\Lambda_{NP} > 100 \text{ GeV}$ (CONUS, $k=0.16$)

$\Lambda_{NP} > 240 \text{ GeV}$ (COHERENT, LAr)

Non-standard interactions (NSIs)

Aim: Test for new (high-scale) neutrino interactions

- Tensor-type operator \rightarrow exotic neutral currents!
 - Interaction with higher kinematic cut-off
- \rightarrow quenching less relevant, overall background crucial

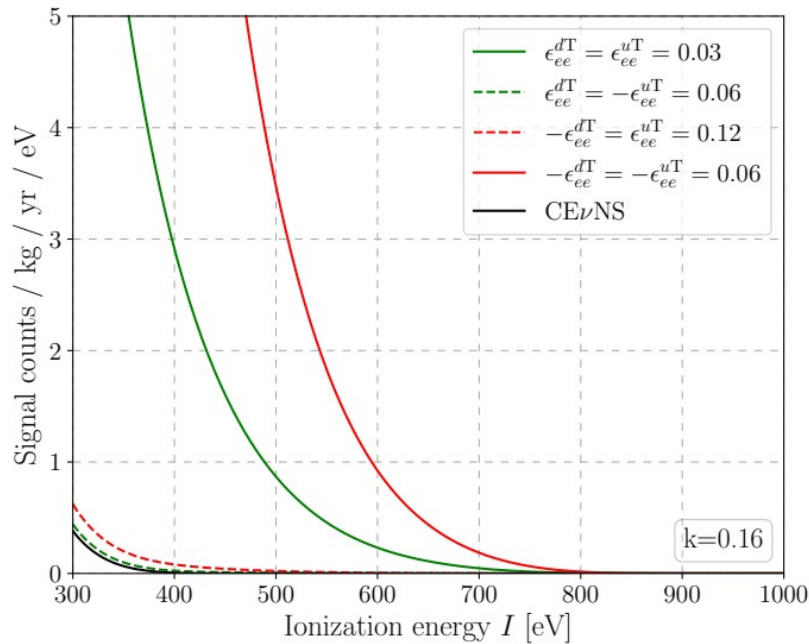
[Barranco et al., [10.1142/S0217751X12501473](https://arxiv.org/abs/10.1142/S0217751X12501473), 2012]

$$O_{\alpha\beta}^{qT} = (\bar{\nu}_\alpha \sigma^{\mu\nu} \nu_\beta) (\bar{q} \sigma_{\mu\nu} q) + \text{h.c.}$$

$$Q_{NSI}^T = (2\epsilon_{\alpha\beta}^{uT} + \epsilon_{\alpha\beta}^{dT}) Z + (\epsilon_{\alpha\beta}^{uT} + 2\epsilon_{\alpha\beta}^{dT}) N$$

$$\left(\frac{d\sigma}{dT_N}\right)_{NSI}^T = \frac{G_F^2 M}{\pi} (Q_{NSI}^T)^2 \left(1 - \frac{MT_N}{4E_\nu^2}\right)$$

[Bonet et al., [10.1007/JHEP05\(2022\)085](https://arxiv.org/abs/10.1007/JHEP05(2022)085), 2022]



Non-standard interactions (NSIs)

Aim: Test for new (high-scale) neutrino interactions

- Tensor-type operator \rightarrow exotic neutral currents!
 - Interaction with higher kinematic cut-off
- \rightarrow quenching less relevant, overall background crucial

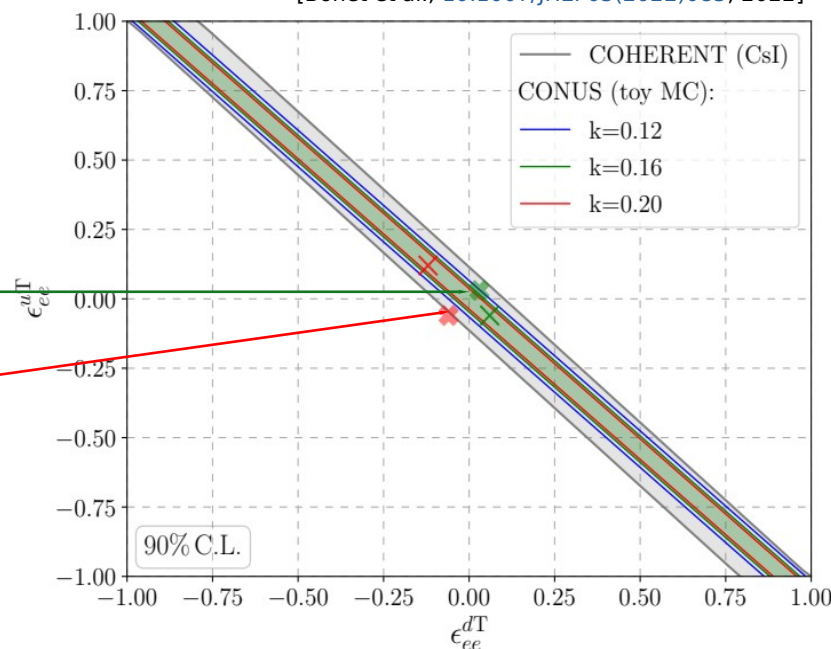
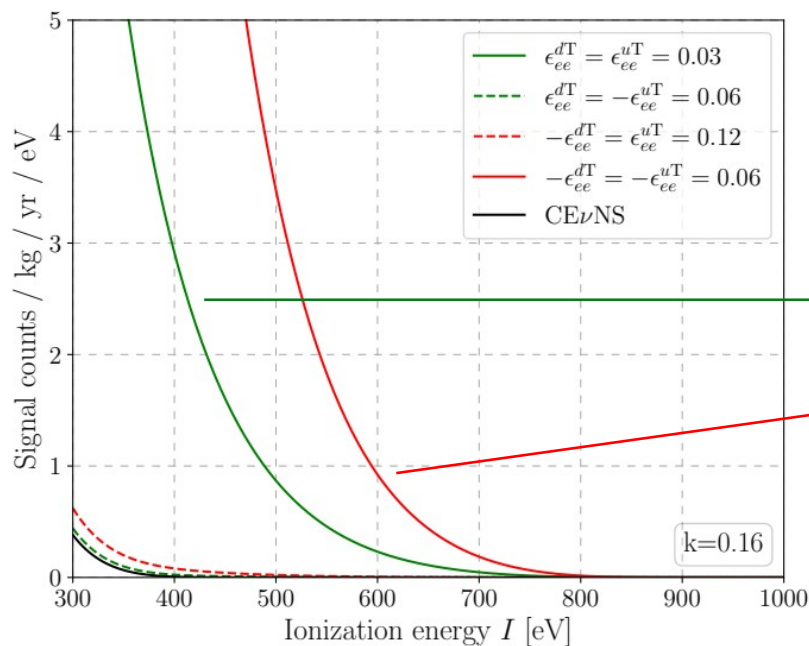
[Barranco et al., 10.1142/S0217751X12501473, 2012]

$$\mathcal{O}_{\alpha\beta}^{qT} = (\bar{\nu}_\alpha \sigma^{\mu\nu} \nu_\beta) (\bar{q} \sigma_{\mu\nu} q) + \text{h.c.}$$

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[Bonet et al., 10.1007/JHEP05(2022)085, 2022]



\rightarrow CONUS data sets allow for competitive bounds with COHERENT!

$$\Lambda_{NP} > 360 \text{ GeV (CONUS, } k=0.16)$$

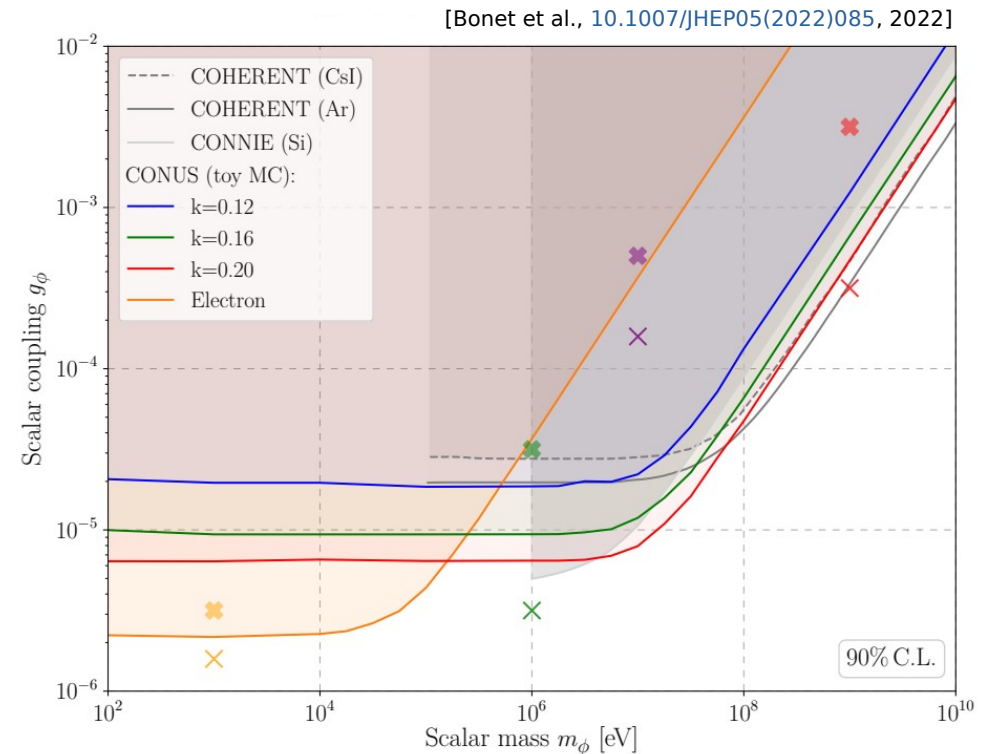
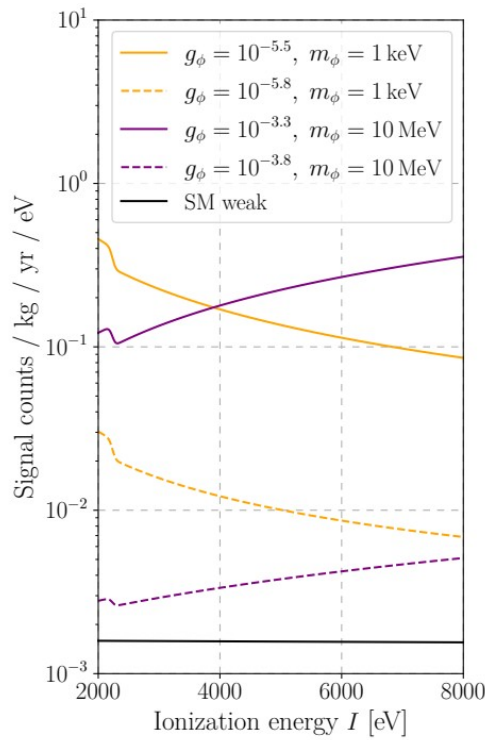
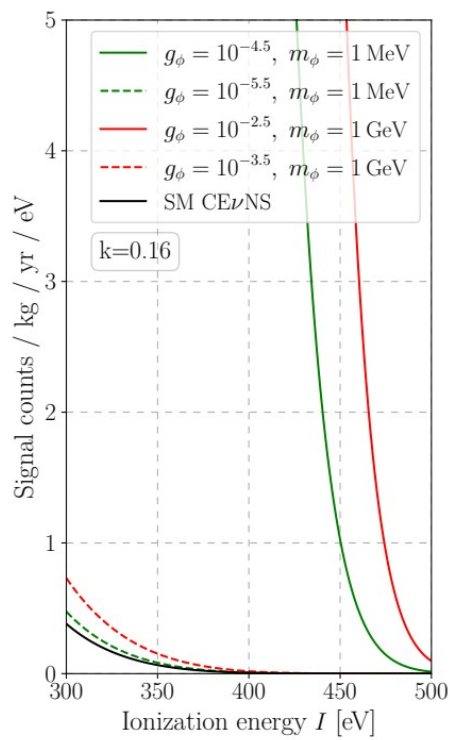
Simplified models: light scalar mediator

Aim: Test specific but simple mediator models that contribute to CEνNS / EveS

[Cerdeño et al., 10.1007/JHEP05(2016)118, 2016]

$$\mathcal{L}_\phi = \phi \left(g_\phi^{qS} \bar{q}q + g_\phi^{eS} \bar{e}e + g_\phi^{\nu S} \bar{\nu}_R \nu_L + \text{h.c.} \right) - \frac{1}{2} m_\phi^2 \phi^2$$

- Assume: universal coupling to quarks / neutrinos: (m_ϕ, g_ϕ)
- Spectral distortions for **small recoil energies**



Reactor neutrinos for low masses \leftrightarrow π -DAR neutrinos for higher masses

Simplified models: light scalar mediator

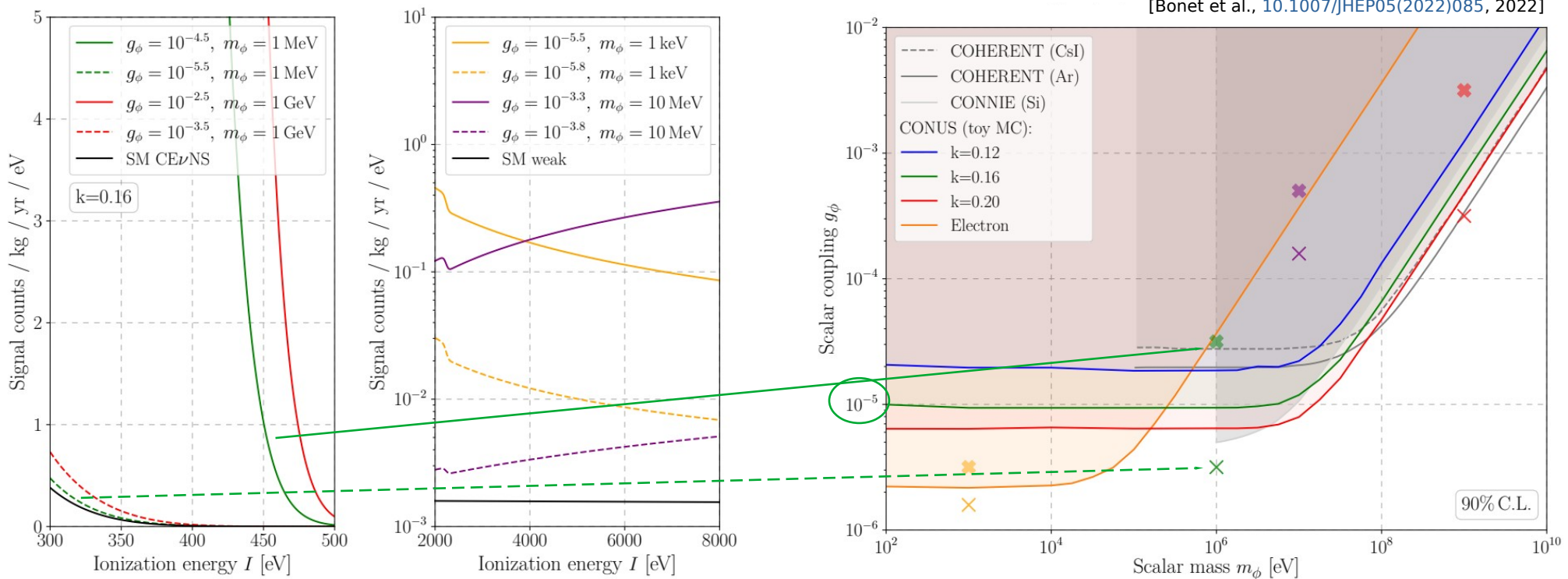
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Limits from CEνNS



[Bonet et al., 10.1007/JHEP05(2022)085, 2022]

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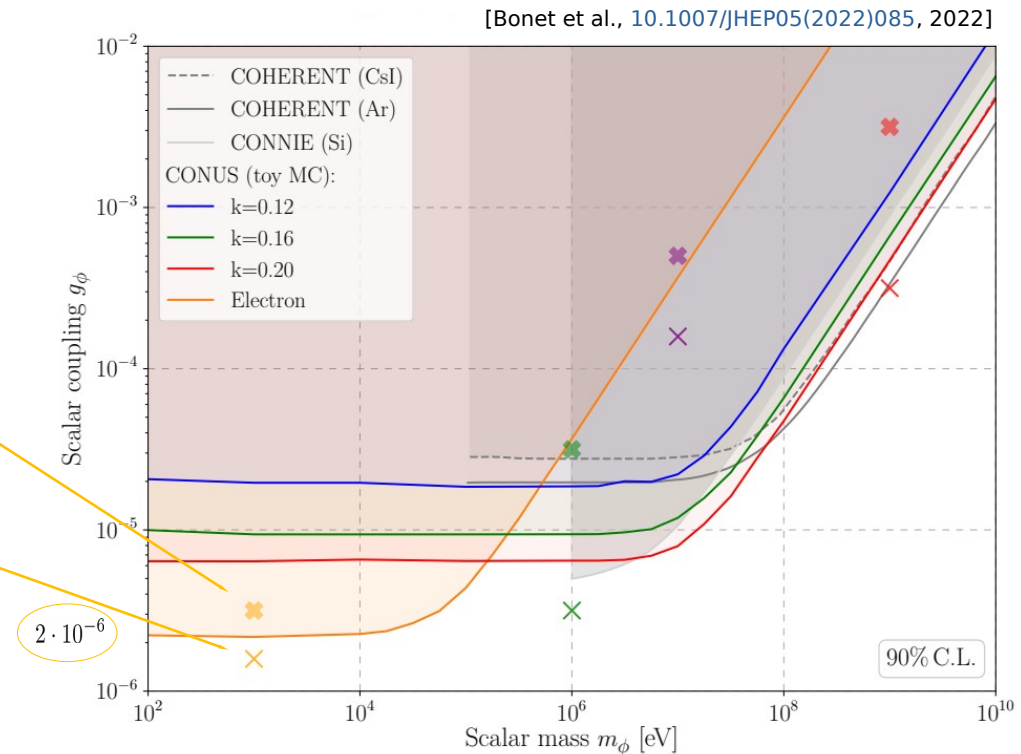
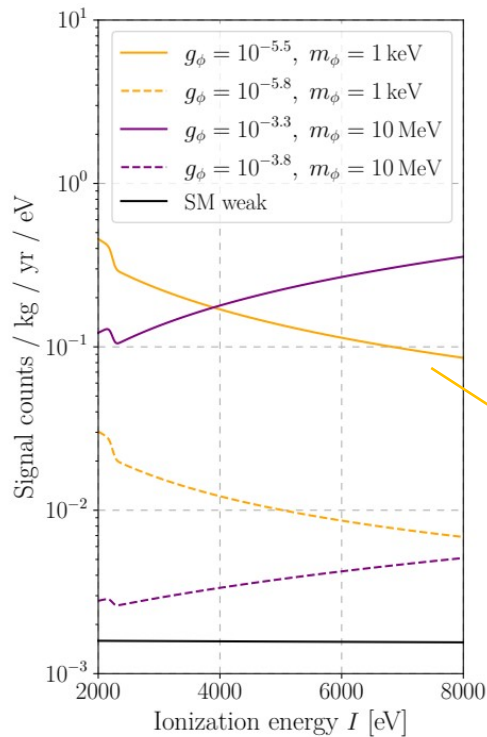
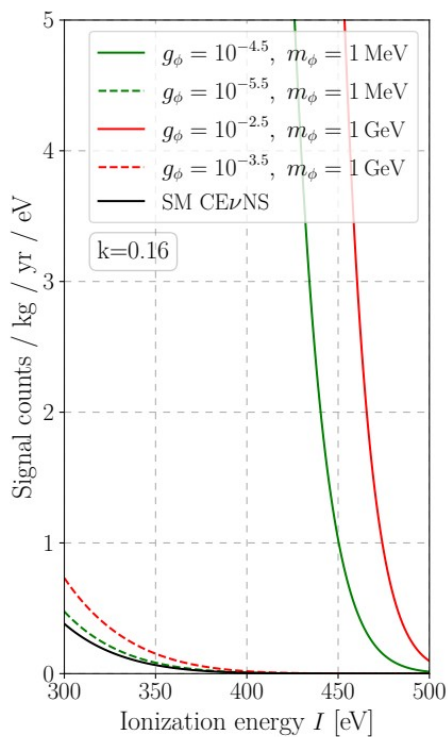
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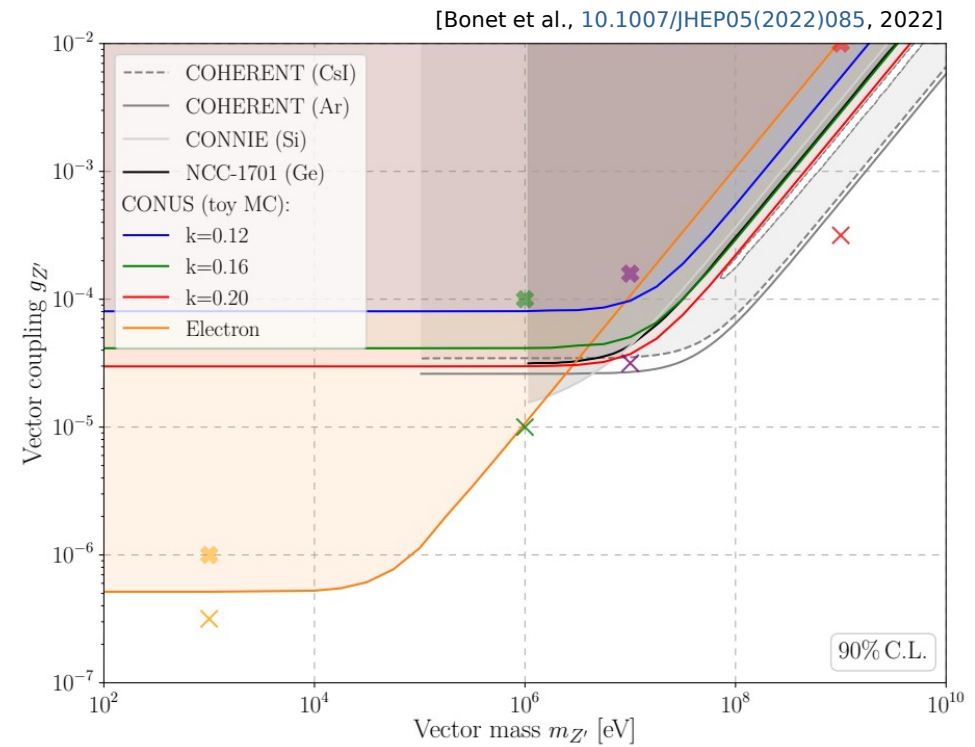
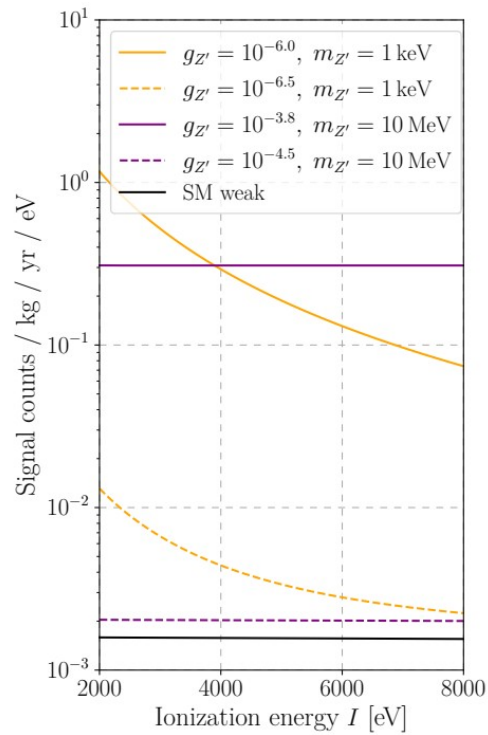
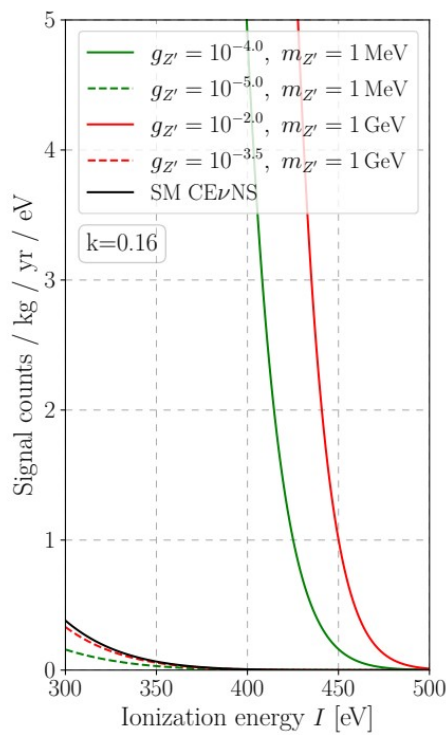
Simplified models: light vector mediator

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$$\mathcal{L}_{Z'} = Z'_\mu \left(g_{Z'}^{\nu V} \bar{\nu}_L \gamma^\mu \nu_L + g_{Z'}^{eV} \bar{e} \gamma^\mu e + g_{Z'}^{qV} \bar{q} \gamma^\mu q \right) + \frac{1}{2} m_{Z'}^2 Z'_\mu Z'^\mu$$

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Reactor neutrinos for low masses \leftrightarrow π -DAR neutrinos for higher masses

Simplified models: light vector mediator

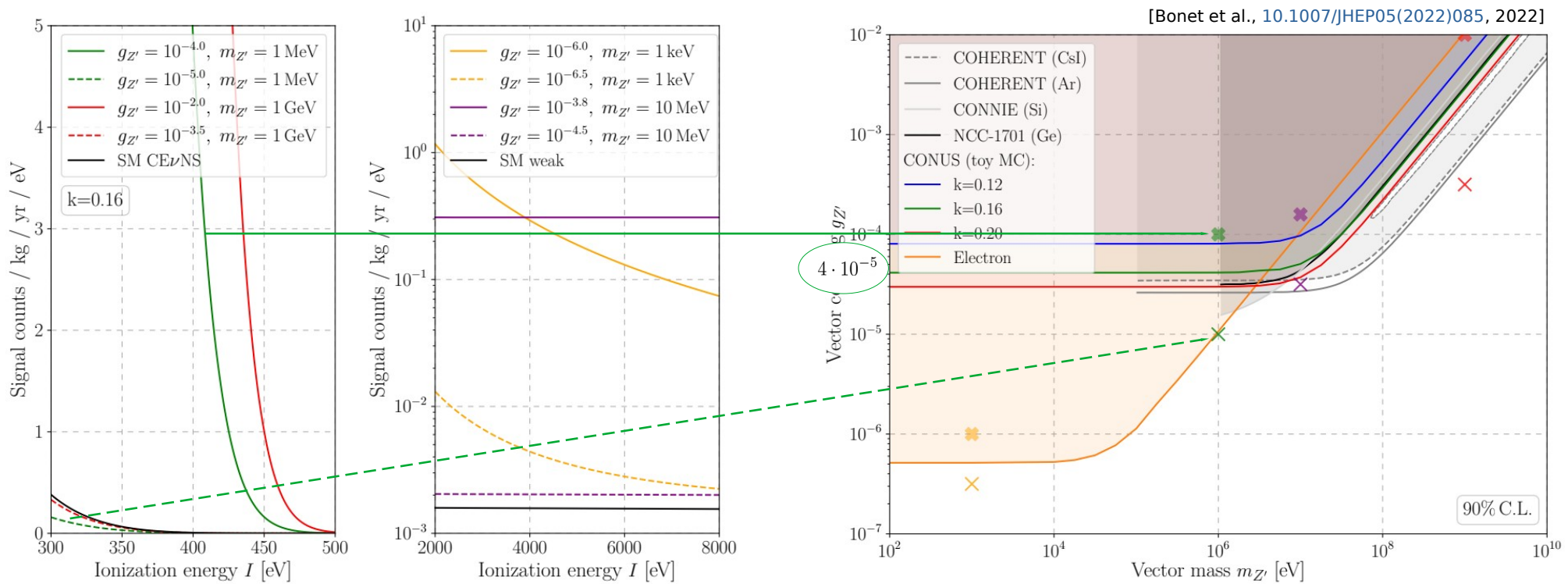
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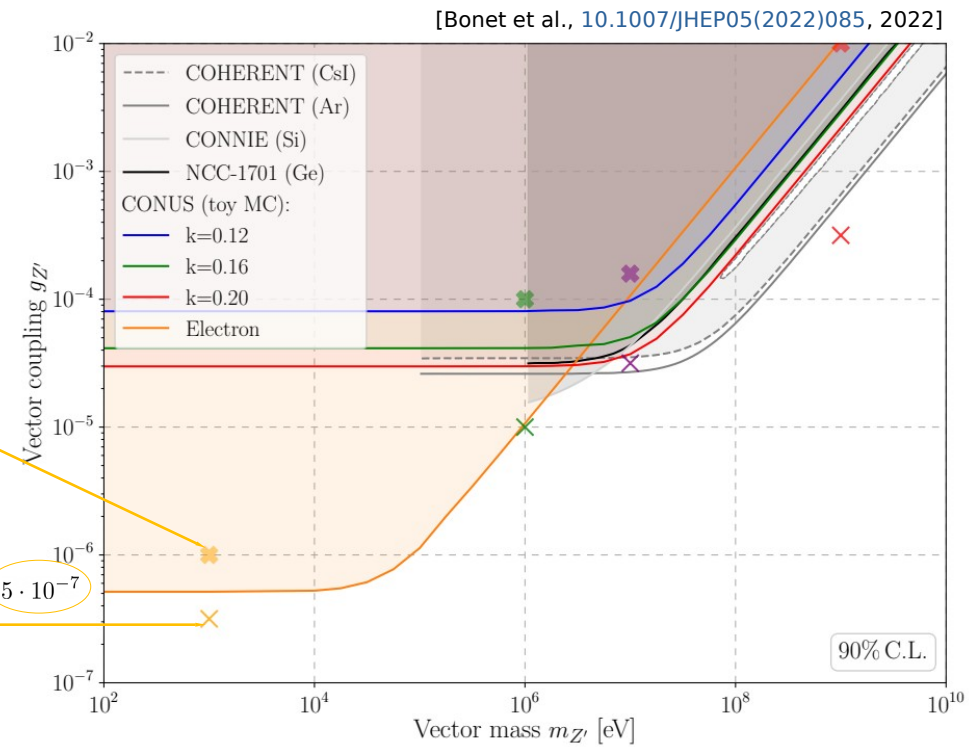
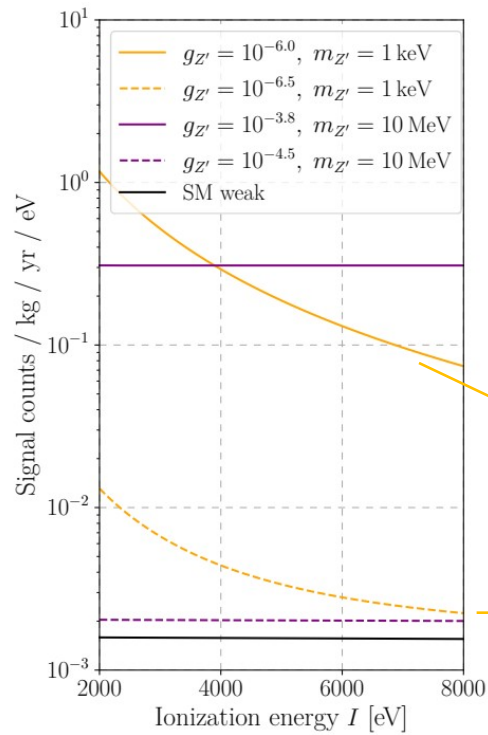
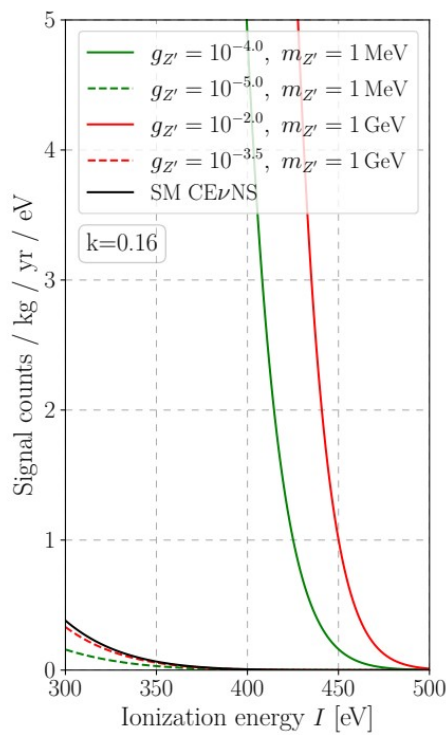
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Reactor neutrinos for low masses \leftrightarrow π -DAR neutrinos for higher masses

Opportunities for CEvNS experiments

SM investigations with enhanced precision:

- **Weinberg angle at low-Q:**

Test for deviations at low energy
 → light Z' , neutrino charge radius

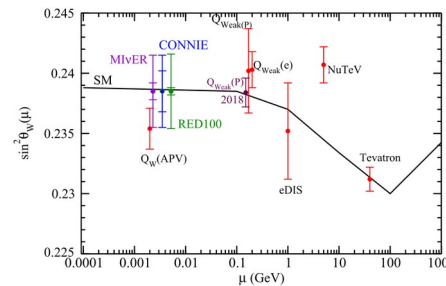
- **Nuclear form factors:**

Model-independent extraction of neutron density distributions
 → Beam-reactor complementarity

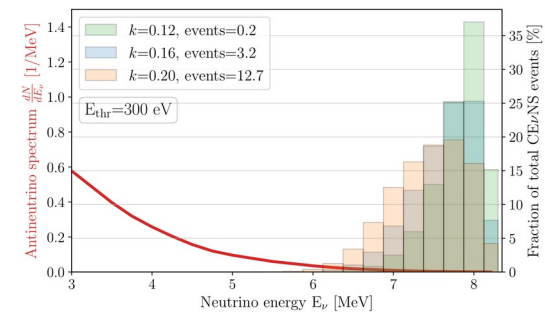
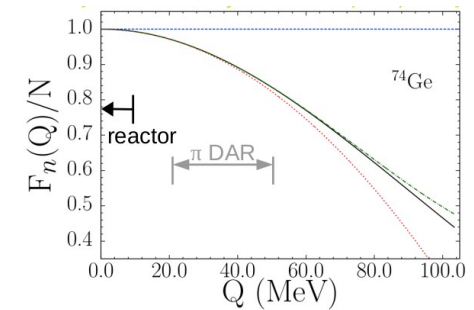
- **Measuring reactor antineutrino spectrum:**

CEvNS sensitive to high-E part where uncertainty is largest

[Cañas et al., 10.1016/j.physletb.2018.07.049, 2018]



[Patton et al., 10.1103/PhysRevC.86.024612, 2012]



[TR, 10.11588/heidok.00031274, 2022]

Opportunities for CEvNS experiments

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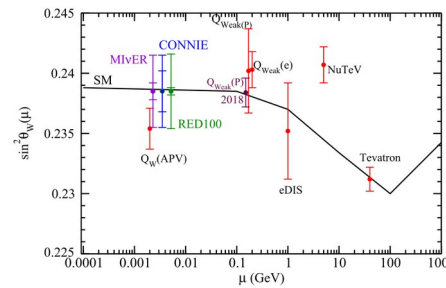
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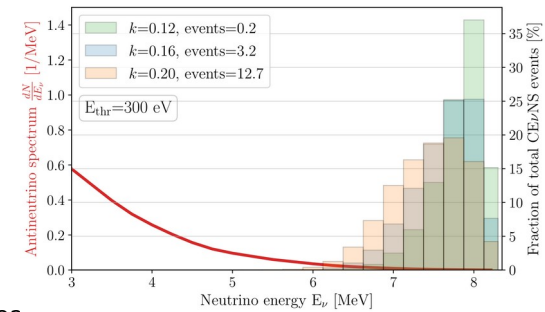
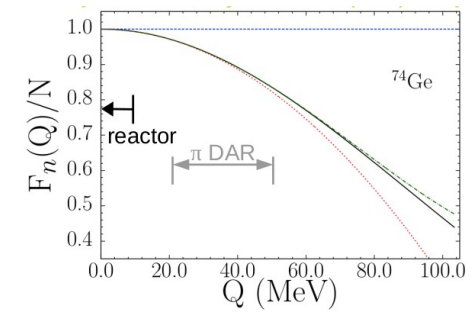
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[TR, 10.11588/heidok.00031274, 2022]

BSM investigations:

- Light sterile neutrinos:**

Use CEvNS for ν flux measurements

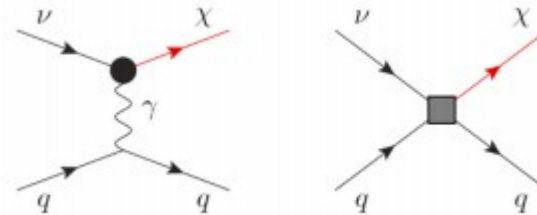
- New fermion searches:**

Test further ν interactions
 → ν mass, DM, ...

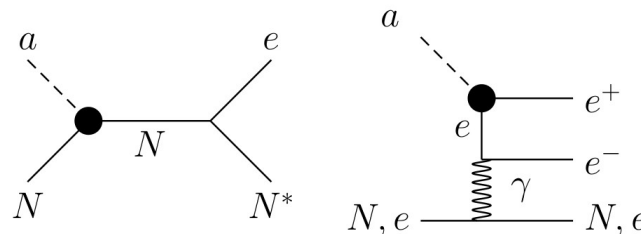
- Probing portals:**

ALPs, dark photons, etc.

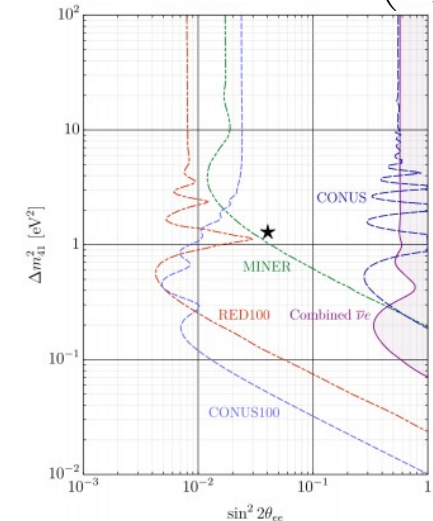
[Chang & Liao, 10.1103/PhysRevD.102.075004, 2022]



[Aristizabal Sierra et al., 10.1007/JHEP03(2021)294, 2021]



$$P_{\nu_\alpha \rightarrow \nu_\alpha} = 1 - \sin^2 2\theta_{\alpha\alpha} \sin^2 \left(\frac{\Delta m_{41}^2 L}{4E_\nu} \right)$$



[Berryman, 10.1103/PhysRevD.100.023540, 2019]

Conclusion

- **CEvNS opens new path to high-statistics neutrino physics**

- beams and reactors go hand in hand!
- full spectrum of modern detection technologies
- “car-size” neutrino detectors!

- **CEvNS data allow various SM & BSM investigations**

- Weinberg angle, nuclear form factors, ...
- *NSIs, light mediators, neutrino properties, DM, ...*

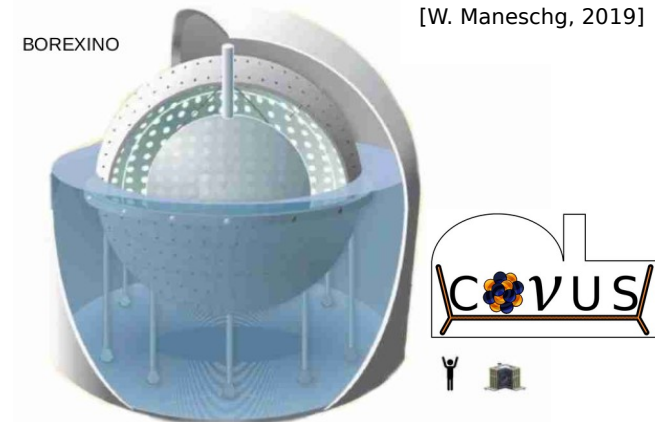
- **Several improvements of CONUS set-up:**

- environmental stability, new DAQ (lower E_{thr} , event shapes)
- pin down CEvNS → deduce more & stronger BSM limits!
- long run data collection ongoing

→ new experimental site under discussion!

- **Future: flavor-blind neutrino astronomy**

- supernovae, neutrino floor, ...



[TR, 10.11588/heidok.00031274, 2022]

| | | |
|----------------|-------|---|
| Vector NSI | CEvNS | $\Lambda_{\text{NP}} \gtrsim 100 \text{ GeV}$ |
| Tensor NSI | CEvNS | $\Lambda_{\text{NP}} \gtrsim 360 \text{ GeV}$ |
| Light scalar | CEvNS | $g_\phi \sim 1 \cdot 10^{-5}$ |
| | EveS | $g_\phi \sim 2 \cdot 10^{-6}$ |
| Light vector | CEvNS | $g_{Z'} \sim 4 \cdot 10^{-5}$ |
| | EveS | $g_{Z'} \sim 5 \cdot 10^{-7}$ |
| νMM | EveS | $\mu_{\nu_e} < 7.5 \cdot 10^{-11} \mu_B$ |
| νMC | EveS | $ q_{\nu_e} < 3.3 \cdot 10^{-12} e$ |

[Hakenmüller et al. 10.1140/epjc/s10052-019-7160-2, 2019]

[Bonet et al., 10.1103/PhysRevLett.126.041804, 2020]

[Bonet et al., 10.1140/epjc/s10052-021-09038-3, 2021]

[Bonet et al., 10.1007/JHEP05(2022)085, 2022]

[Bonet et al., arXiv:2112.09585, 2021]

[Bonet et al., arXiv:2201.12257, 2022]

[Bonhomme et al., arXiv:2202.03754, 2022]

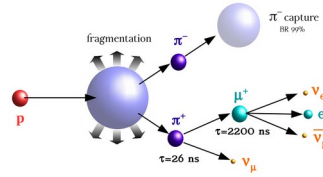
Next-generation experiments (π -DAR, reactors & DM DD!) = huge playground for phenomenology

(Neutrino magnetic moment, neutrino charge radius, vector NSI, tensor NSI, light scalars, light vectors, sterile neutrinos, dark matter, weak mixing angle, nuclear form factors, neutrino-electron-scattering, supernova neutrinos, antineutrino reactor spectra, neutrino floor, etc.)

Backup

Two complementary approaches

π -decay-at-rest neutrinos:



- Pulsed GeV-proton beam hitting heavy target
→ multiple ν flavors
- Time correlation of events
→ background suppression $\times(10^3-10^4)$
- Higher ν energies
→ larger cross section, but reduced coherence

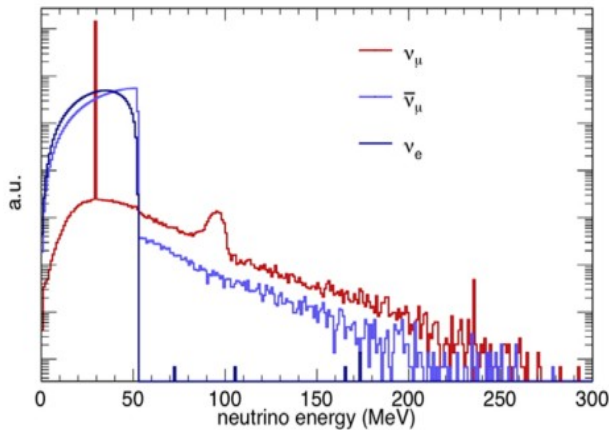
→ **COHERENT, CCM, "CE ν NS@ESS"...**

Reactor antineutrinos:

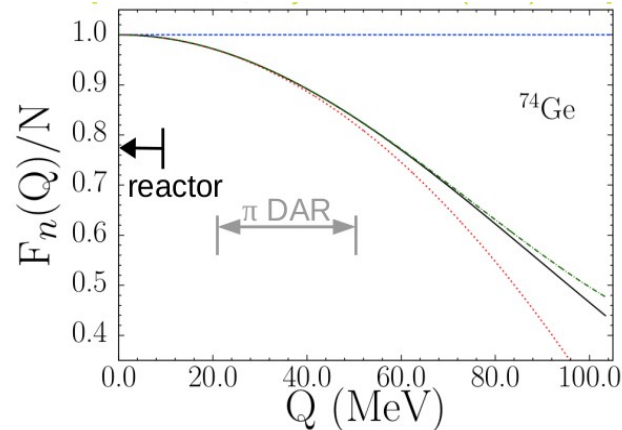
- β decays in nuclear reaction chains → only $\bar{\nu}_e$
- Strongest artificial ν source on earth:
 $\sim 10^{20}$ $\bar{\nu}_e$'s/GW/s
- ν energies up to 10 MeV → coherent regime!
- Close to reactor core: no lab conditions!
→ no cryogenic liquids, no remote control

→ **CONNIE, CONUS, NCC-1701, ν GEN ...**

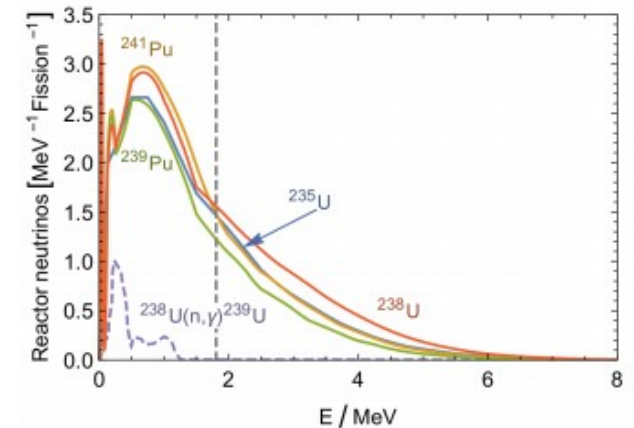
[Akimov et al., arXiv:1509.08702, 2015]



[Patton et al., 10.1103/PhysRevC.86.024612, 2012]



[Vitagliano et al., 10.1103/RevModPhys.92.045006, 2020]



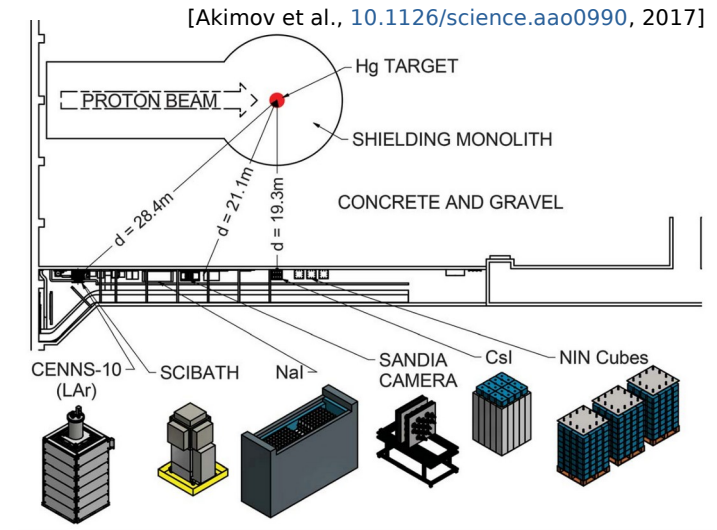
Beam-reactor complementarity:
CE ν NS at reactor site as high statistic baseline for multi-target and multi-flavored beam investigations!

Experimental efforts with first results

COHERENT “neutrino alley” at Oak Ridge National Lab (USA)

- Overburden: 8m w.e.
- ν flux: $4.3 * 10^7/cm^2/s$ @ 20m
- Multiple targets @ different baselines

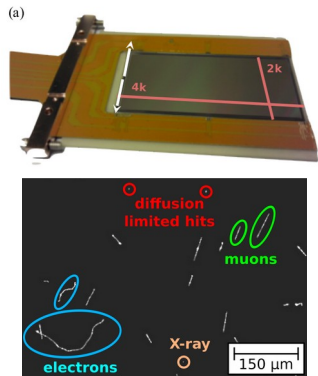
| Target | N | Mass [kg] | T^{thr} [keV _{nr}] |
|---------|----------|-----------|--------------------------------|
| CsI[Na] | 78/74 | 14.6 | ~6.5 |
| LAr | 22 | 22 | ~20 |
| Nal[Tl] | 12/74 | 185 | ~13 |
| Ge | 38/40/42 | 10 | ~5 |



CONNIE (Brazil)

- CCDs (Si, $N=14$) : 47.6g, $T_A^{thr} \sim 1 keV_{nr}$
- @30m to 3.8GW_{th} PWR:
 ν flux: $7.8 * 10^{12}/cm^2/s$

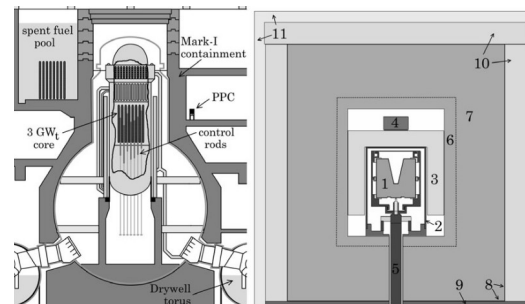
[Aguilar-Arevalo et al., 10.1088/1742-6596/761/1/012057, 2016]



[Aguilar-Arevalo et al., 10.1088/1748-0221/11/07/P07024, 2016]

NCC-1701 (USA)

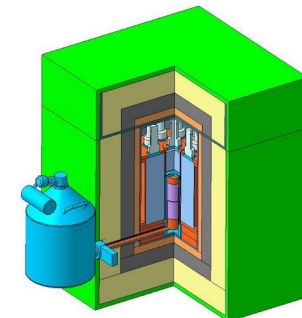
- Ge PPC ($N=\{38,40,42\}$): 3kg, $T_A^{thr} \sim 1-2 keV_{nr}$
- @10.3m to 2.96GW_{th} BWR:
 ν flux: $5 * 10^{13}/cm^2/s$



[Colaresi et al., 10.1103/PhysRevD.104.072003, 2021]

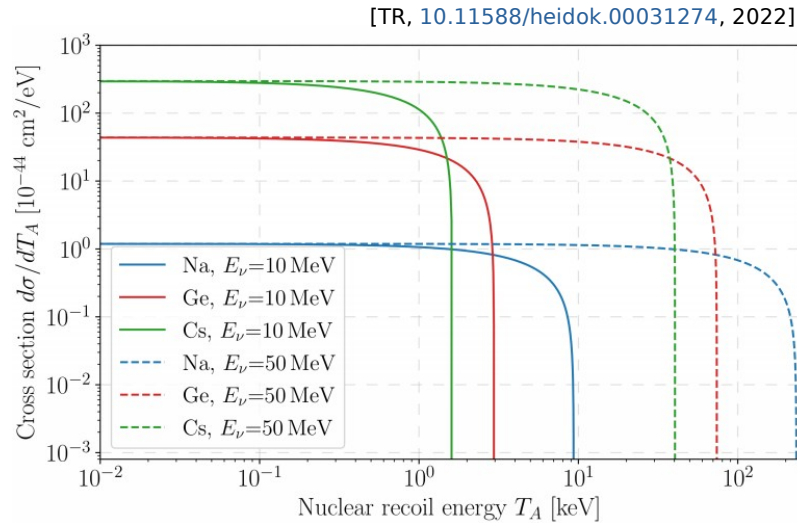
ν GEN (Russia)

- Ge PPC ($N=\{38,40,42\}$): 1.41kg, $T_A^{thr} \sim 1-2 keV_{nr}$
- @11.8m to 3.1GW_{th} PWR:
 ν flux: $3.9 * 10^{13}/cm^2/s$



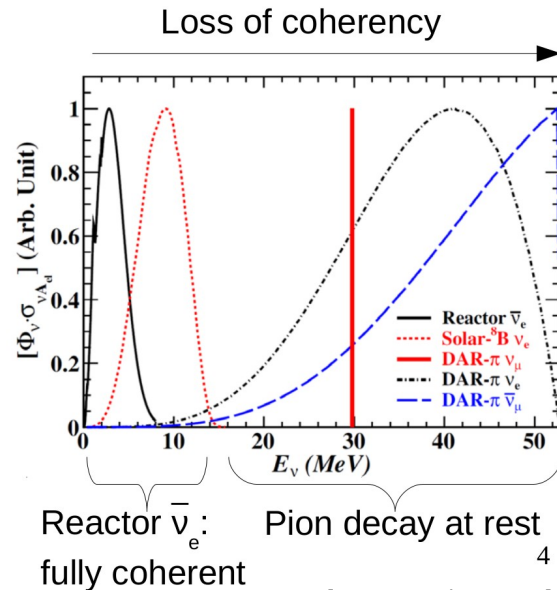
[Belov et al., 10.1088/1748-0221/10/12/P12011, 2015]

CEvNS with different sources and targets



| Element | N | r_A [fm] | E_ν^{\max} [MeV] | T_A^{\max} [keV] |
|---------|----------|------------|----------------------|--------------------|
| Na | 12 | 3.6 | 27.7 | 71.5 |
| Si | 14 | 3.8 | 25.9 | 51.3 |
| Ar | 22 | 4.4 | 23.1 | 28.5 |
| Ge | 38/40/42 | 5.2 | 18.9 | 10.5 |
| I | 74 | 6.3 | 15.7 | 4.16 |
| Xe | 75/77/78 | 6.4 | 15.5 | 3.93 |
| Cs | 78 | 6.4 | 15.4 | 3.85 |

[TR, 10.11588/heidok.00031274, 2022]



[W. Maneschg, 2017]

| Neutrino source | Target | T_A^{\max} [keV] | E (QF \in {0.1, 0.15, 0.2}) [keV] |
|-------------------------------|--------|--------------------|---------------------------------------|
| Nuclear reactor (10 MeV) | Na | 9.33 | 0.93 / 1.40 / 1.87 |
| | Si | 7.64 | 0.76 / 1.15 / 1.53 |
| | Ar | 5.37 | 0.54 / 0.81 / 1.07 |
| | Ge | 2.96 | 0.30 / 0.44 / 0.59 |
| | I | 1.69 | 0.17 / 0.25 / 0.34 |
| | Xe | 1.64 | 0.16 / 0.25 / 0.33 |
| | Cs | 1.62 | 0.16 / 0.24 / 0.32 |
| π -DAR source (50 MeV) | Na | 232.4 | 23.2 / 34.9 / 46.5 |
| | Si | 190.4 | 19.0 / 28.6 / 38.1 |
| | Ar | 134.0 | 13.4 / 20.1 / 26.8 |
| | Ge | 73.8 | 7.38 / 11.1 / 14.8 |
| | I | 42.3 | 4.23 / 6.34 / 8.45 |
| | Xe | 40.9 | 4.09 / 6.13 / 8.17 |
| | Cs | 40.4 | 4.04 / 6.03 / 8.07 |

Antineutrinos from nuclear reaction products

Antineutrino emission in β decays of fuel reaction products

- Mainly from ^{235}U , ^{238}U , ^{239}Pu , ^{241}Pu $\rightarrow >99\%$
- $\sim 6-7$ ν 's/fission up to 10MeV
- Spectral distribution

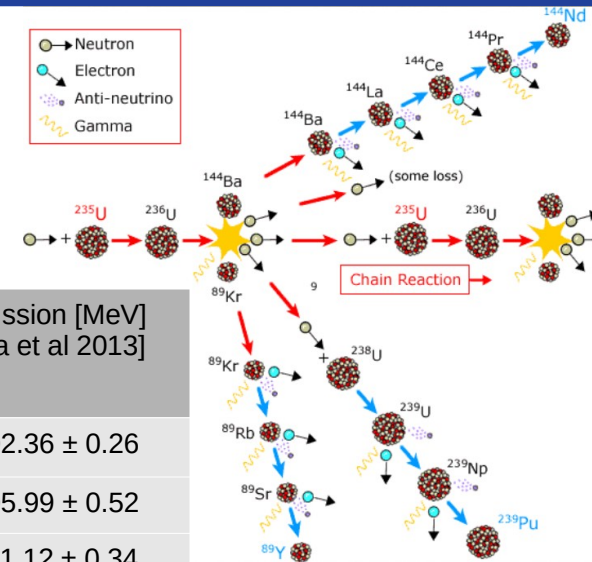
$$S(E_\nu) = \frac{1}{4\pi R^2} \frac{W_{th}}{\sum_i \alpha_i E_i} \sum_i \alpha_i \left(\frac{dN_i}{dE_\nu} \right)$$

Knowledge about a reactors emission spectra

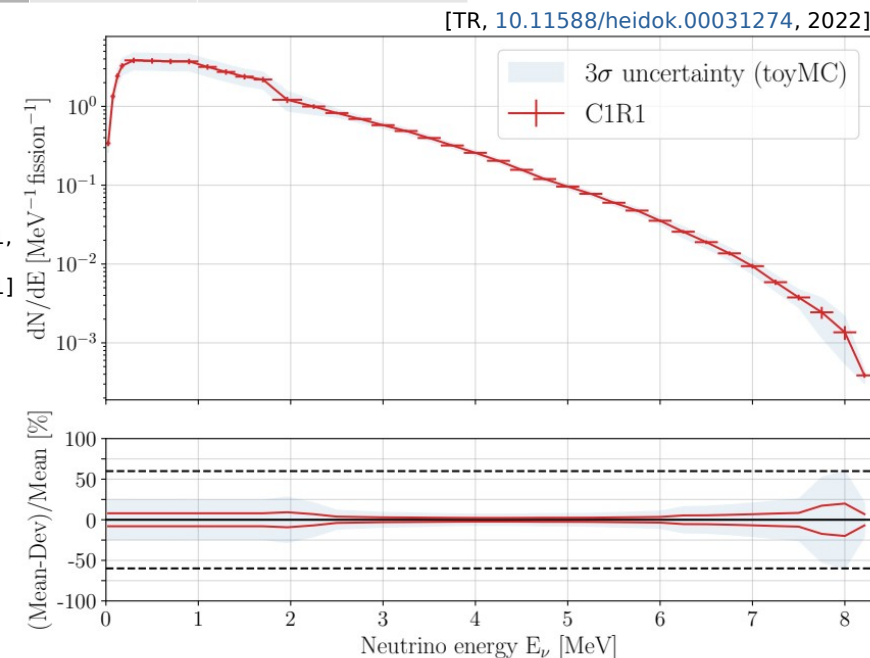
- Summation methods [Kopeikin et al. [10.1134/1.1825513](https://doi.org/10.1134/1.1825513), 2004]
 \rightarrow summing β branches of all fission fragments
- Conversion methods [Haag et al., [10.1103/PhysRevLett.112.122501](https://doi.org/10.1103/PhysRevLett.112.122501), 2014
 Huber, [10.1103/PhysRevC.85.029901](https://doi.org/10.1103/PhysRevC.85.029901), 2011,
 Mueller et al., [10.1103/PhysRevC.83.054615](https://doi.org/10.1103/PhysRevC.83.054615), 2011]
 \rightarrow measure β decay electron spectrum and convert into ν spectrum
- Direct measurements (IBD) [An et al., [10.1088/1674-1137/41/1/013002](https://doi.org/10.1088/1674-1137/41/1/013002), 2017]

Reality much more complicated...

- Varying reactor power $\rightarrow \mathbf{P(t)}$
- Changing fuel composition $\rightarrow \mathbf{\alpha(t)}$



| Isotope | Fission fraction α (PWR) | E/fission [MeV] [Ma et al 2013] |
|-------------------|---------------------------------|---------------------------------|
| ^{235}U | 57% | 202.36 ± 0.26 |
| ^{238}U | 8% | 205.99 ± 0.52 |
| ^{239}Pu | 30% | 211.12 ± 0.34 |
| ^{241}Pu | 5% | 214.26 ± 0.33 |

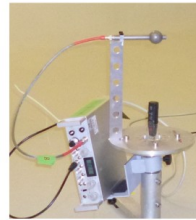
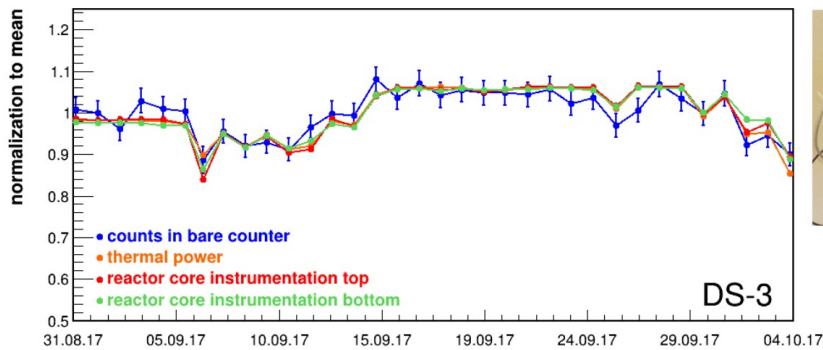


Power-correlated reactor radiation

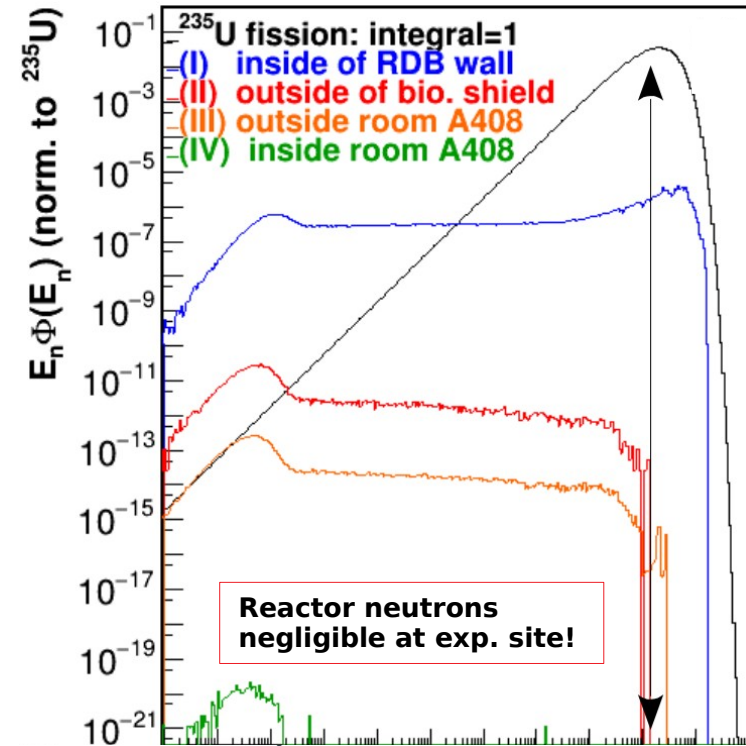
→ Dedicated investigation of reactor-correlated background contributions:

[Hakenmüller et al. 10.1140/epjc/s10052-019-7160-2, 2019]

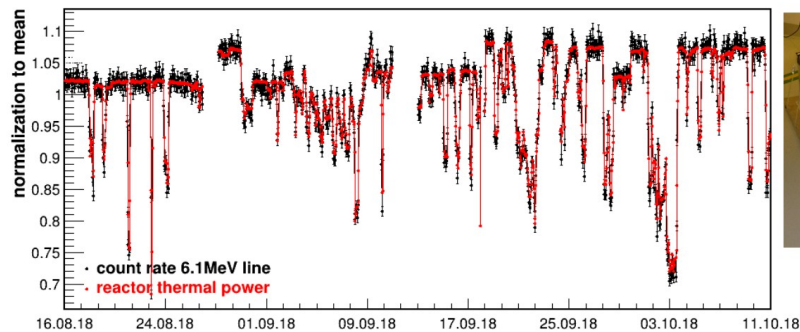
- Simulation and validation of neutrons emitted from reactor core (at CONUS site) → thermal neutron counter



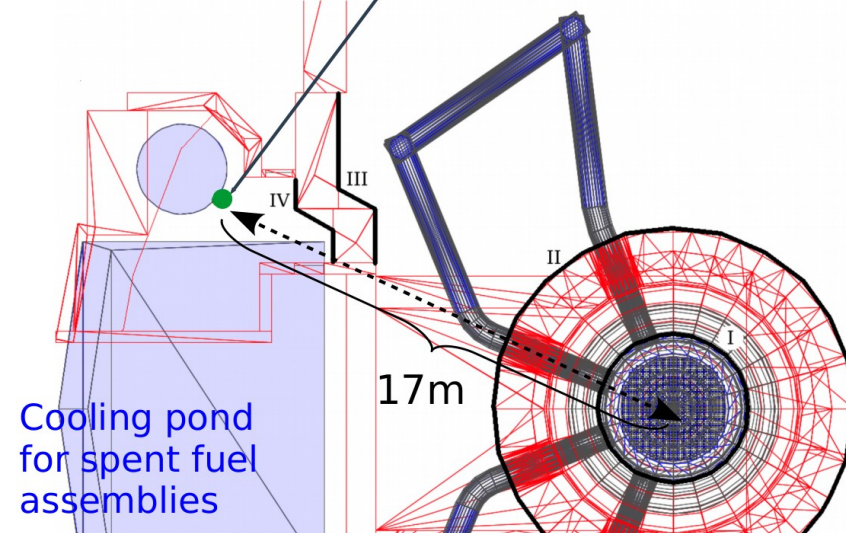
Correlation to incore and excore instrumentation



- γ radiation (6.1 MeV from ^{16}N)

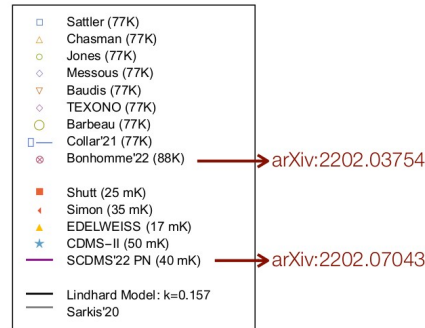
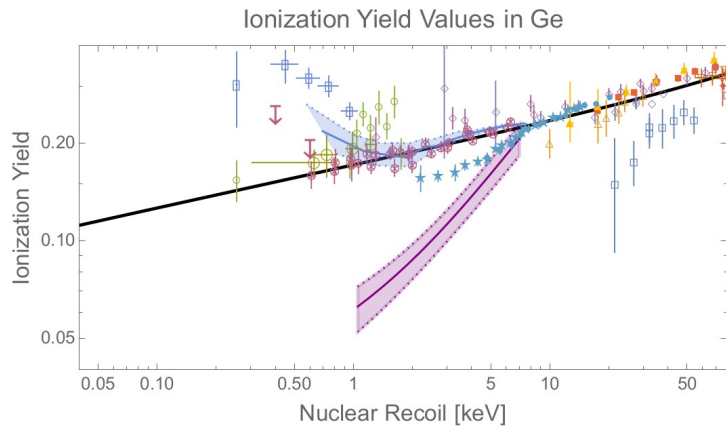


Unshielded HPGGe detector



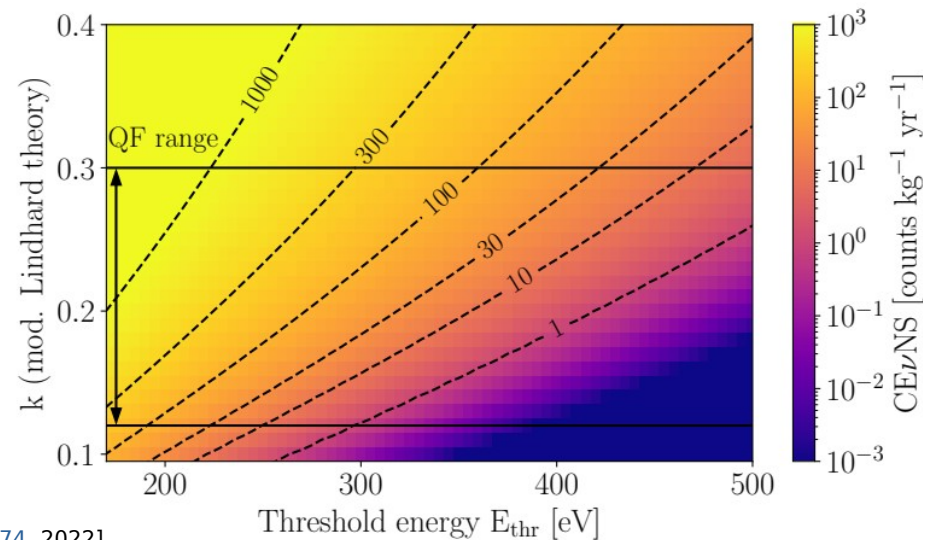
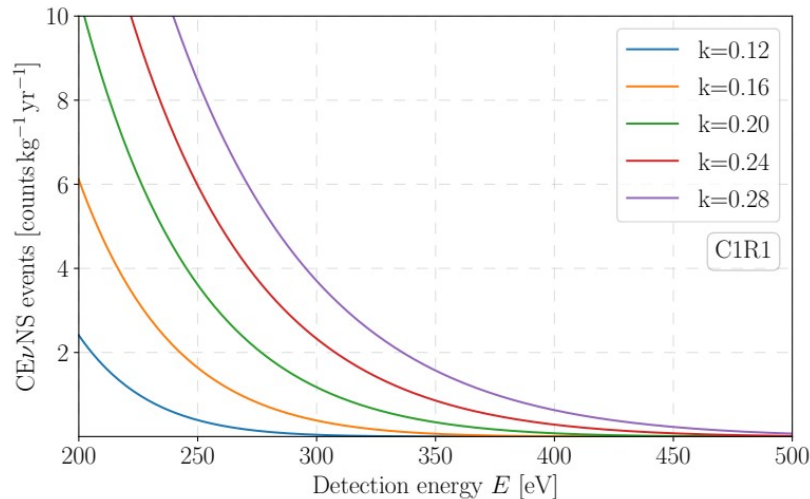
Impact of quenching at low energy

Overview of quenching measurements: T. Saab \ EXCESS 2022 \ February 16, 2022



Affects sensitivity of CEvNS and DM searches:
 (Collar '21) used in [arXiv:2202.09672](https://arxiv.org/abs/2202.09672)
 "Suggestive evidence for Coherent Elastic Neutrino-Nucleus Scattering from reactor antineutrinos"

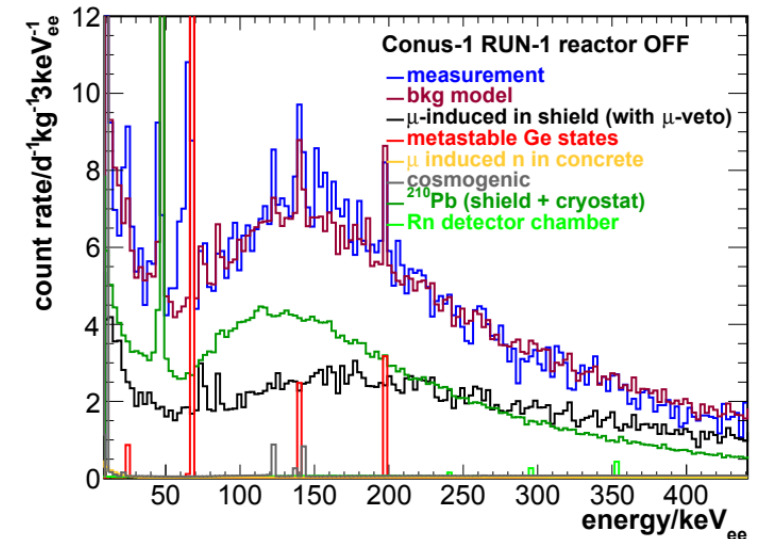
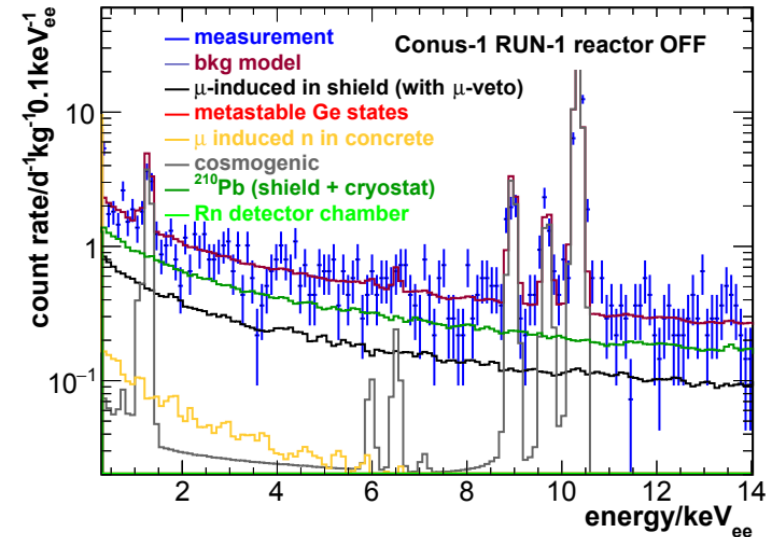
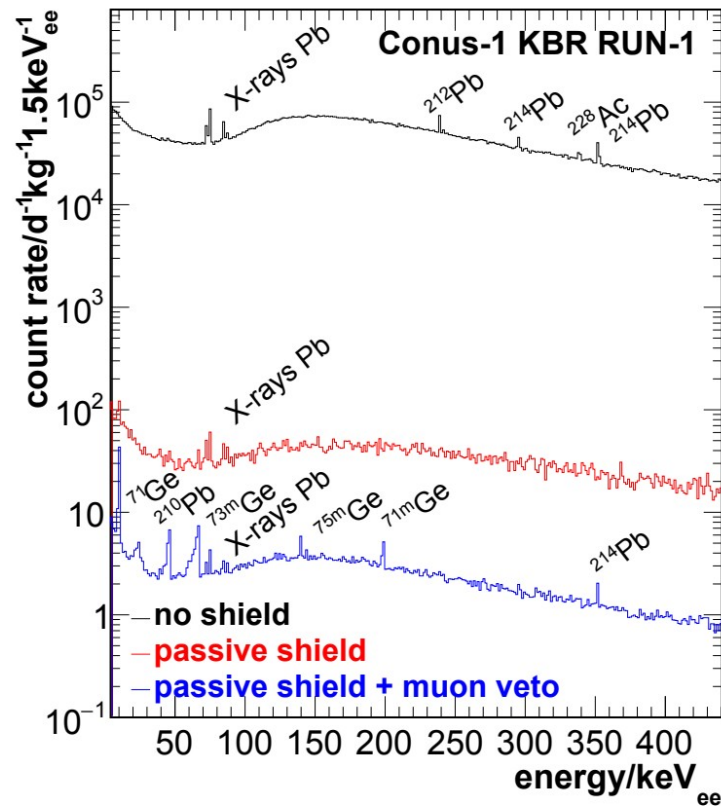
Quenching according to mod. Lindhard model:



[TR, 10.11588/heidok.00031274, 2022]

Background (model) at reactor site

[Bonet et al., arXiv:2112.09585, 2021]



| Fast neutron classes | Corr. with therm. power |
|-------------------------------------|-------------------------|
| μ -ind. in Pb inside shield | No |
| μ -ind. above ceiling | No |
| (α, n) -reactions from walls | No |
| fission n from spent fuel rods | No |
| fission n from reactor core | Yes |

Data analyses of experimental data

(Binned log)-Likelihood in the SM case:

- ON and OFF data are fitted together:

$$\log \mathcal{L} = \log \mathcal{L}_{\text{ON}} + \log \mathcal{L}_{\text{OFF}} + \text{pull terms}$$

$$\text{with } \log \mathcal{L}_{\text{ON}}(s, b, \Theta_{\text{thr}_1}, \Theta_{\text{thr}_2}, \Theta_{\text{rea}}, \Theta_{\text{det}}, \Theta_{\Delta E})$$

$$\log \mathcal{L}_{\text{OFF}}(b, \Theta_{\text{thr}_1}, \Theta_{\text{thr}_2}, \Theta_{\text{det}}, \Theta_{\Delta E})$$

- Parameter list:

- s : signal strength (free), b : MC normalisation (free)
- $\Theta_{\text{thr}_{1/2}}$: electronic noise (free)
- Θ_{rea} : reactor neutrino flux ($\sim 3\%$)
- Θ_{det} : detector/DAQ (1-5%), $\Theta_{\Delta E}$: energy scale (10-20 eV)

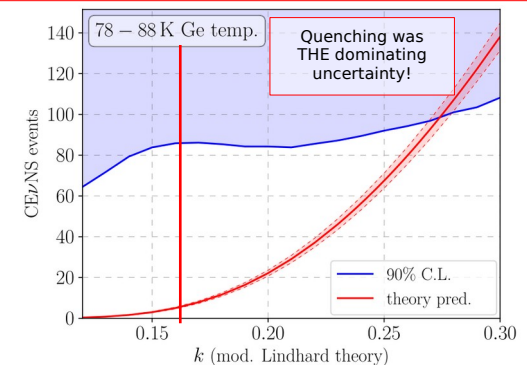
- Compare signal to background-only hypothesis: **LH ratio**

BSM investigations:

- Extend existing LH with BSM contributions
→ CE ν NS now additional background
- Additional systematic due to elastic ν -e- scatterings (**EveS**) @[2,8]keV
→ 5% shape uncertainty on background model
- BSM limits in terms of quenching factor!

Data sets: 5 detectors (in 2 runs)

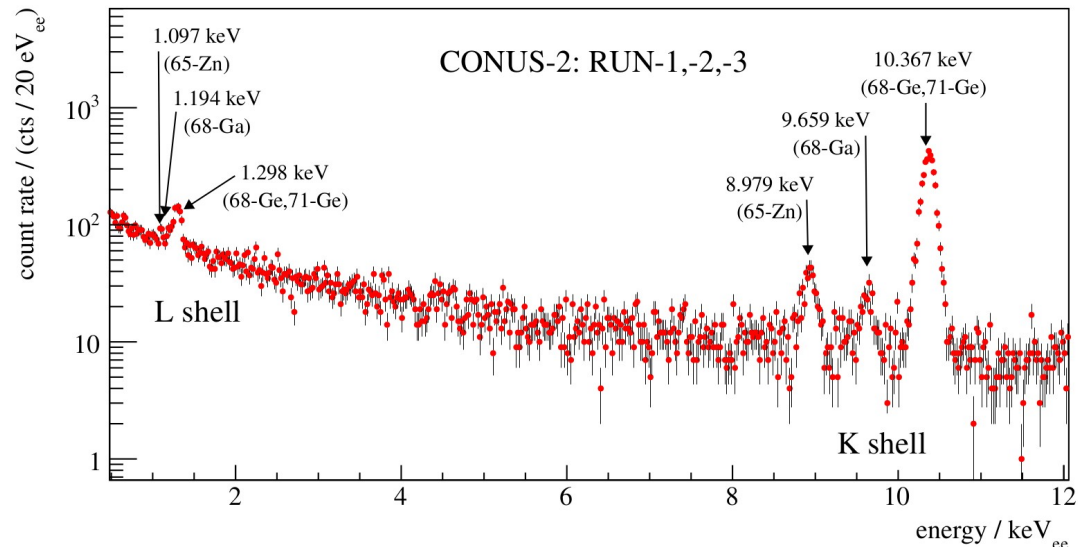
| Data set | ON [kg d] | OFF [kg d] | ROI [keV] |
|----------|-----------|------------|--------------|
| C1R1 | 96.7 | 13.8 | 0.296 - 0.75 |
| C2R1 | 14.6 | 13.4 | 0.311 - 1.00 |
| C3R1 | 97.5 | 10.4 | 0.333 - 1.00 |
| C1R2 | 19.6 | 12.1 | 0.348 - 0.75 |
| C3R2 | 20.2 | 9.1 | 0.343 - 1.00 |
| All | 248.7 | 58.8 | |



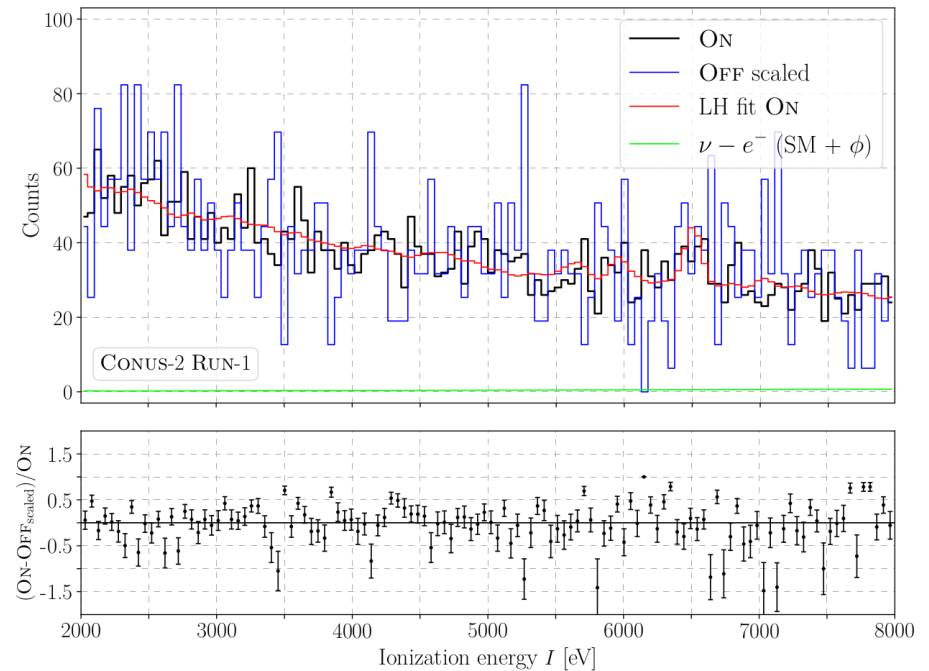
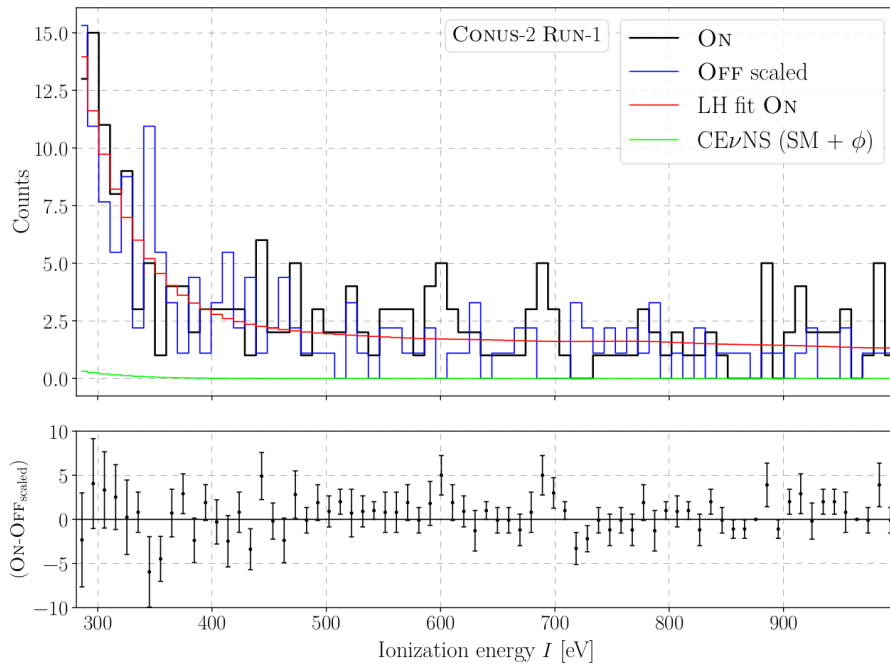
| Analysis | Data set | ON [kg d] | OFF [kg d] | ROI [keV] |
|--|----------|-----------|------------|---------------|
| BSM ($\bar{\nu}_e + A(Z, N)$) | C1R1 | 96.7 | 13.8 | 0.276 - 0.741 |
| | C2R1 | 14.6 | 13.4 | 0.281 - 0.999 |
| | C3R1 | 97.5 | 10.4 | 0.333 - 0.991 |
| | → all | 208.8 | 37.6 | |
| BSM + νMM ($\bar{\nu}_e + e$) | C1R1 | 215.4 | 29.6 | 2.013 - 7.968 |
| | C2R1 | 184.6 | 32.2 | 2.006 - 7.990 |
| | C3R1 | 248.5 | 31.7 | 2.035 - 7.989 |
| | → all | 648.5 | 93.5 | |
| νMM ($\bar{\nu}_e + e$) | C1R2 | 19.6 | 18.5 | 2.010 - 7.955 |
| | C3R2 | 20.8 | 19.0 | 2.007 - 7.991 |
| | → all | 688.9 | 131.5 | |

Exemplary spectra: CONUS-2 RUN-1

[Bonet et al., 10.1140/epjc/s10052-021-09038-3, 2021]



[Bonet et al., 10.1007/JHEP05(2022)085, 2022]



Sensitivity of next and future data

[TR, 10.11588/heidok.00031274, 2022]

Estimate for recent data collection period

- Bkg stability → exposure 2.7 kg*yr
- Quenching factor known! → CONUS measurement
- New DAQ → lower threshold
- PSD → background reduction of ~20%

| E_{thr} [eV] | CONUS [counts] | Z | V |
|-----------------------|----------------|------|-----|
| 300 | 30 | 0.15 | 42 |
| 275 | 49 | 0.25 | 16 |
| 250 | 99 | 0.50 | 4.1 |
| 225 | 178 | 0.88 | 1.3 |
| 200 | 286 | 1.40 | 0.5 |

→ naive counting estimate: no shape, no systematics

$$s: \text{signal events} \quad Z = s/\sqrt{b}$$
$$b: \text{background events} \quad V = (s + b)/s^2$$

Theoretical CONUS upgrade

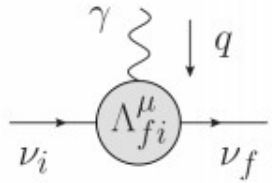
- New experimental site similar to KBR
- Same bkg and detector specifications
- Exposure: 500 kg*yr

→ precision neutrino physics with CEvNS!

| E_{thr} [eV] | CONUS-100 [counts] | Z | V |
|-----------------------|--------------------|------|---------------------|
| 300 | $5.5 \cdot 10^3$ | 2.1 | $227 \cdot 10^{-3}$ |
| 275 | $9.1 \cdot 10^3$ | 3.4 | $85 \cdot 10^{-3}$ |
| 250 | $18 \cdot 10^3$ | 6.8 | $22 \cdot 10^{-3}$ |
| 225 | $33 \cdot 10^3$ | 12.0 | $7 \cdot 10^{-3}$ |
| 200 | $53 \cdot 10^3$ | 19.0 | $3 \cdot 10^{-3}$ |

Electromagnetic neutrino properties

Loop-induced electromagnetic properties: [Giunti & Studenikin, 10.1103/RevModPhys.87.531, 2015]



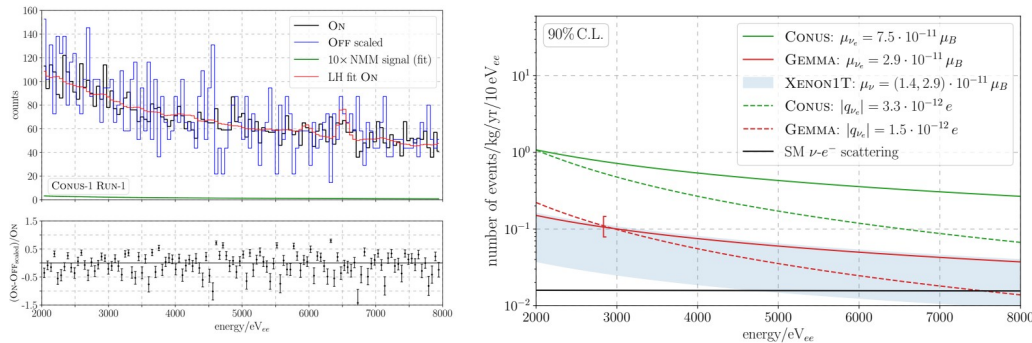
- neutrino charge (radius), electric and magnetic dipole moment, anapole moment
- characteristic dependence on neutrino's fermionic nature
- test for BSM physics!

νMM from electron scattering at reactor site:

$$\left(\frac{d\sigma}{dT_e}\right)_{\nu\text{MM}} = \frac{\pi\alpha_{\text{em}}^2}{m_e^2} \left(\frac{1}{T_e} - \frac{1}{E_\nu}\right) \left(\frac{\mu_{\nu e}}{\mu_B}\right)^2$$

- huge reactor ν flux
- low threshold energies!

[Bonet et al., arXiv:2201.12257, 2022]



$$\mu_{\nu e} < 7.5 \cdot 10^{-11} \mu_B \text{ (90\% C. L.)}$$

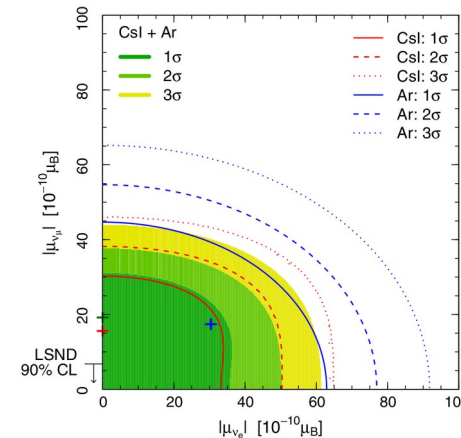
Assuming null result: convert to νMC bound

$$|q_{\nu e}| < 3.3 \cdot 10^{-12} e \text{ (90\% C. L.)}$$

νMM from CEνNS at a π-DAR site:

$$\left(\frac{d\sigma}{dT_A}\right)_{\nu\text{MM}} = \frac{\pi\alpha_{\text{em}}^2}{m_e^2} Z^2 \left(\frac{1}{T_A} - \frac{1}{E_\nu}\right) \left(\frac{\mu_{\nu e}}{\mu_B}\right)^2 F^2(T_A)$$

- higher ν energy
- multiple flavors



[Cadeddu et al., 10.1103/PhysRevD.102.015030, 2020]

νMC bounds from COHERENT:

$$|q_\nu| \sim \mathcal{O}(10^{-8}) e$$

Future studies of transition moments, neutrino charge radii..., etc.