Composite Dark Matter and Neutrino Masses from a Light Hidden Sector

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WITH A. AHMED, Z. CHACKO, N. DESAI, S. DOSHI, C. KILIC, in preparation

INTRODUCTION

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- However, there are several puzzles/observations which are unanswered within the SM.
- Two of the most outstanding puzzles are
 - ► Dark matter (DM)
 - ► SM neutrino masses
- Neutrino oscillations have shown that SM neutrinos have tiny but non-zero masses, $m_{\nu} \sim 0.1 \, \mathrm{eV}$.
- However, mechanism to generate such tiny neutrino masses is one of the main research topics of BSM physics.
- At present, the nature of the particles of which DM is composed remain unknown

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- Hidden sector is approximately conformal in the UV, and compositeness scale lies at or below the weak scale.

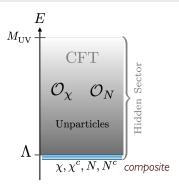
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- In our framework DM candidate arises as a composite state of a strongly coupled hidden sector.
- Hidden sector is approximately conformal in the UV, and compositeness scale lies at or below the weak scale.
- We construct this framework based on 5D AdS geometric setup and explore implication for experiments.

A CONFORMAL HIDDEN SECTOR

 We consider a hidden sector composed of a strongly coupled conformal field theory (CFT) with a relevant deformation O_{def},

$$\mathcal{L}_{\mathrm{UV}} \supset \mathcal{L}_{\mathrm{CFT}} + \lambda_{\mathrm{def}} \mathcal{O}_{\mathrm{def}}$$

• When the deformation grows large in the infrared, it causes the breaking of the conformal dynamics at a scale $\Lambda \lesssim v_{\rm SM}$.

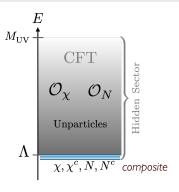


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- Spectrum of hidden sector states includes three composite singlet neutrinos N_i and a composite DM χ along with their Dirac partners N_i^c and χ^c .
- lacksquare Low energy effective Lagrangian contains $(m_N,m_\chi\sim\Lambda)$

$$\mathcal{L}_{\rm IR} \supset i\bar{N}\bar{\sigma}^{\mu}\partial_{\mu}N + i\bar{N}^{c}\bar{\sigma}^{\mu}\partial_{\mu}N^{c} - (m_{N}N^{c}N + \text{h.c.}) + i\bar{\chi}\bar{\sigma}^{\mu}\partial_{\mu}\chi + i\bar{\chi}^{c}\bar{\sigma}^{\mu}\partial_{\mu}\chi^{c} - (m_{\chi}\chi^{c}\chi + \text{h.c.})$$

SM is assumed to be elementary!

Composite Neutrino Portal

Hidden sector interacts with the SM only through the neutrino portal

$$\mathcal{L}_{\mathrm{UV}} \supset -rac{\hat{\lambda}}{M_{\mathrm{UV}}^{\Delta_N-3/2}} LH\mathcal{O}_N + \mathrm{h.c.}$$

- $lacktriangleright \mathcal{O}_N$ is a primary fermionic operator with scaling dimension Δ_N
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- At or below the conformal breaking scale Λ , the portal interaction:

$$\mathcal{L}_{\mathrm{IR}} \supset -\lambda \, LHN + \mathrm{h.c.} \hspace{1cm} ext{with} \hspace{1cm} \lambda \sim \hat{\lambda} igg(rac{\Lambda}{M_{\mathrm{UV}}} igg)^{\Delta_N - 3/2}$$

- lacksquare For $\Delta_N \geq 3/2$, the coupling λ is hierarchically small for $\Lambda \ll M_{\mathrm{UV}}$.
- Naturally small portal coupling λ provides a simple explanation for the both the smallness of the neutrino masses and the observed abundance of DM.

- We assume that the hidden sector possesses a global symmetry such that \mathcal{O}_N , and therefore N, carries charge -1.
- Due to neutrino portal interaction this symmetry can be subsumed into an overall lepton number symmetry, under which N, N^c carry charges -1, +1.

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- To employ the *inverse seesaw* mechanism we add a lepton number violating deformation,

$$\mathcal{L}_{\mathrm{UV}} \supset -\frac{\hat{\mu}^c}{M_{\mathrm{UV}}^{\Delta_{2N^c}-4}} \mathcal{O}_{2N^c} + \mathrm{h.c.}$$

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- We assume \mathcal{O}_{2N^c} carries a charge of +2 under the global symmetry of the hidden sector, so that this deformation violates lepton number by two units.
- In the low-energy effective theory at scale Λ , this deformation gives,

$$\mathcal{L}_{\mathrm{IR}} \supset -rac{\mu^c}{2}ig(N^cig)^2 + \mathrm{h.c.} \qquad ext{with} \qquad \mu^c \sim \hat{\mu}^c \Lambda \left(rac{\Lambda}{M_{\mathrm{UV}}}
ight)^{\Delta_{2N^c-4}}$$

■ The low-energy effective theory now contains all the ingredients required to realize *inverse seesaw* mechanism,

$$\mathcal{L}_{\rm IR} \supset i\bar{N}\bar{\sigma}^{\mu}\partial_{\mu}N + i\bar{N}^{c}\bar{\sigma}^{\mu}\partial_{\mu}N^{c} - \left[m_{N}N^{c}N + \frac{\mu^{c}}{2}(N^{c})^{2} + \lambda LHN + \text{h.c.}\right]$$

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■ By integrating out the composite singlet neutrinos N and N^c we obtain the SM neutrinos masses and their mixing with the composite states N,

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- Smallness of the SM neutrino masses can naturally be explained by either small neutrino mixing λ or small lepton number violating coupling μ^c .
- We require Dirac mass $\lambda v_{\rm EW} \lesssim \Lambda$ and the Majorana mass $\mu^c \lesssim \Lambda$:

$$\frac{m_N}{v_{\rm EW}} \gtrsim \lambda \gtrsim \frac{\sqrt{m_\nu m_N}}{v_{\rm EW}}, \qquad 1 \gtrsim U_{N\ell} \gtrsim \sqrt{\frac{m_\nu}{m_N}}$$

COMPOSITE DM THROUGH THE NEUTRINO PORTAL

- In our framework, the neutrino portal interaction keeps the hidden sector in equilibrium with the SM in the early universe for $|U_{N\ell}|^2 \gtrsim \sqrt{\Lambda/4\pi M_{\rm Pl}}$.
- Composite nature of the fermions χ , χ^c and N, N^c allows non-renormalizable interactions in the low energy theory at the scale Λ ,

$$\mathcal{L}_{\rm IR} \supset -\frac{y_{\rm eff}^2}{\Lambda^2} (\bar{\chi}^c N)^2 + \cdots$$

where $y_{\rm eff} \sim 4\pi$.

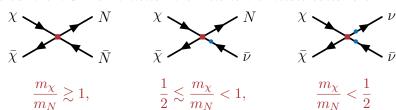
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- DM abundance is set by the standard thermal freeze-out mechanism.
- The dominant DM annihilation channels to the visible sector are



DM Relic Abundance



The thermally averaged DM annihilation cross sections at DM freeze-out, i.e. for $T=T_{\rm fo}\sim m_{\chi}/10$, are

$$\langle \sigma_{\chi\bar\chi\to N\bar N} v \rangle_{\rm fo} \sim \frac{y_{\rm eff}^4}{40\pi\,\Lambda^2}, \quad \langle \sigma_{\chi\bar\chi\to N\bar\nu} v \rangle_{\rm fo} \sim \frac{y_{\rm eff}^4\,U_{N\ell}^2}{40\pi\,\Lambda^2}, \quad \langle \sigma_{\chi\bar\chi\to\nu\bar\nu} v \rangle_{\rm fo} \sim \frac{y_{\rm eff}^4\,U_{N\ell}^4}{40\pi\,\Lambda^2}$$

■ The observed DM relic abundance is produced when $\langle \sigma v \rangle_{\rm fo} \sim 10^{-8} \; {\rm GeV}^{-2}$.

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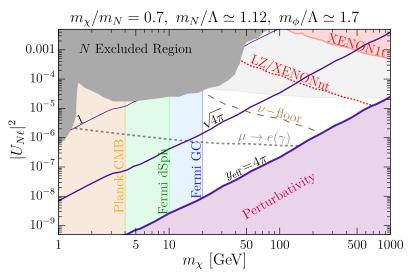
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- The observed DM relic abundance is produced when $\langle \sigma v \rangle_{\rm fo} \sim 10^{-8} \; {\rm GeV}^{-2}$.
- Note $\chi \bar{\chi} \to N \bar{N}$ process leads to DM under-abundance for strong coupling $y_{\rm eff} \sim 4\pi$ and $\Lambda \lesssim \mathcal{O}(100)$ GeV.
- Hence only viable DM production channels are $\chi \bar{\chi} \to N \bar{\nu}$ and $\chi \bar{\chi} \to \nu \bar{\nu}$.

COMPOSITE DM PHENOMENOLOGY

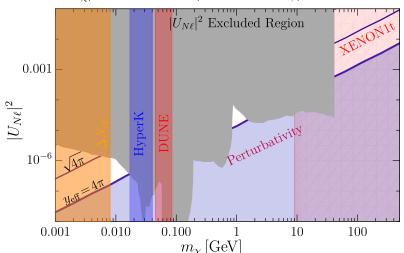
■ Summary for a benchmark in the DM mass range $1/2 \lesssim m_\chi/m_N \lesssim 1$ where the dominant DM annihilation channel is $\chi \bar{\chi} \to N \bar{\nu}$.



Composite DM Phenomenology

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$$m_{\chi}/m_N = 0.4, \ m_N/\Lambda \simeq 1.12, \ m_{\phi}/\Lambda \simeq 1.7$$



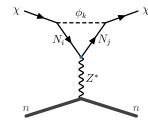
SUMMARY

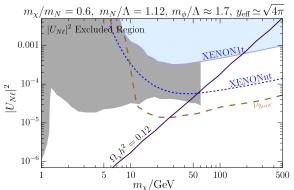
- We presented a class of models in which DM is a composite state of a strongly coupled hidden sector which interacts with the SM through the neutrino portal.
- DM relic abundance is set by annihilation into neutrinos.
- The neutrino portal also leads to the generation of SM neutrino masses through the *inverse seesaw* mechanism.
- We focused on the scenario in which the hidden sector is conformal in the ultraviolet, and the compositeness scale lies at or below the weak scale.
- A holographic realization of this framework is studied based on 5D AdS geometry.
- This scenario can lead to signals in DM detection experiments as well as in colliders in the near future.

DM DIRECT DETECTION

- Dominant contribution to DM-Nucleon arises from Z-boson exchange.
- Spin-independent DM-Nucleon cross-

section is:
$$\sigma_{\chi n} \sim \frac{g^4\,y_{\rm eff}^4\,U_{N\ell}^4}{\pi(4\pi)^4}\frac{\mu_{\chi n}^2}{m_Z^4}$$



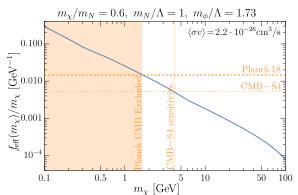


DM INDIRECT DETECTION: CMB CONSTRAINT

- For $\chi \bar{\chi} \to N \bar{\nu}$ channel, the final state N decays to visible end products such as electrons, photons etc., which could alter the CMB measurements.
- From the CMB data Planck collaboration constraints at 95% C.L. on

$$f_{\rm eff}(m_\chi) {\langle \sigma v \rangle \over m_\chi} < 3.2 \times 10^{-28} \, {\rm cm}^3 \, {\rm s}^{-1} \, \, {\rm GeV}^{-1}$$

 $f_{
m eff}(m_\chi)$ is the effective fraction of energy transferred to the IGM.

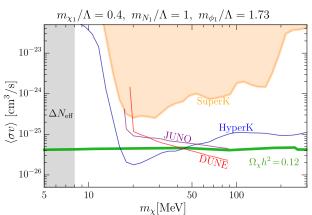


DM INDIRECT DETECTION: NEUTRINO-LINE SIGNALS

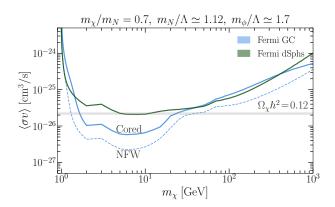
- The dominant annihilation channel $\chi \bar{\chi} \to N \bar{\nu}$ or $\chi \bar{\chi} \to \nu \bar{\nu}$, gives rise to monochromatic neutrinos in the final state.
- In dense DM matter environments e.g. the centre of our Milky Way galaxy such DM annihilations could lead to the possibility of observing neutrinoline signals in neutrino detection experiments.

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GAMMA RAY CONSTRAINTS

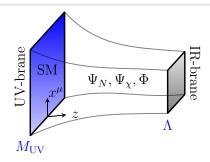


HOLOGRAPHIC REALIZATION

 The holographic model is realized in a 5D anti-de Sitter (AdS) space

$$ds^2 = \left(\frac{R}{z}\right)^2 \eta_{MN} \, dx^M dx^N$$

where
$$x^M = (x^\mu, z)$$
 and $R \le z \le R'$.



- The two branes correspond to the UV and IR scales, $M_{\rm UV} \equiv 1/R$ and $\Lambda \equiv 1/R'$.
- The SM is localized at the UV brane which corresponds to the elementary states in the 4D dual picture.
- New composite states corresponding the strongly coupled hidden sector are in the bulk and at the IR-brane.

HOLOGRAPHIC REALIZATION

Interaction between 5D neutrinos with the SM is

$$S_{\rm UV} \supset \int d^4x \int dz \left(\frac{R}{z}\right)^4 \delta(z-R) \sqrt{R} \,\hat{\lambda} \, LH \, \Psi_N(x,z)$$

- After choosing appropriate boundary conditions and KK-decomposing the bulk fields, 4D effective theory contains KK towers of singlet neutrinos N_n , N_n^c , fermion DM χ_n , χ_n^c , as well as the singlet scalar ϕ_n modes.
- Neutrino portal interaction is

$$S_{\rm UV} \supset \int d^4x \sum_n \lambda_n \, LH \, N_n(x)$$

where λ_n contains the bulk neutrino $\Psi_N(x,R)$ wave-function.

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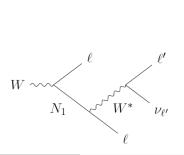
DM and singlet neutrino interact through Yukawa term.

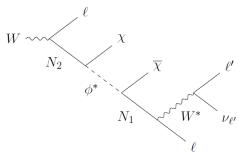
$$S_{\text{bulk}} \supset \int d^4x \int dz \sqrt{g} \,\hat{y} \sqrt{24\pi^3 R} \,\bar{\Psi}_{\chi}^c \Psi_N \,\Phi = \int d^4x \sum_{n,n,q} y_{npq} \,\bar{\chi}_n^c N_p \phi_q$$

Holographic model reproduces our 4D CFT results.

COLLIDER PHENOMENOLOGY

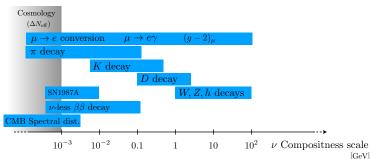
- At colliders and beam-dump experiments, DM can be pair produced in association with one or more composite singlet neutrinos.
- To discover the DM, it is therefore necessary to first discover the composite singlet neutrinos.
- lacksquare Searches for N are broadly divided based on whether N decays promptly in colliders, displaced, or is long lived.
- Collider signal processes of interest for this work are



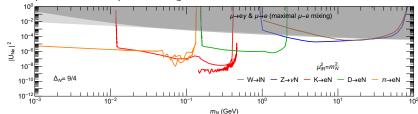


COMPOSITE NEUTRINO SIGNALS

■ There are various probes of neutrino compositeness scales Λ . [Chacko,Fox,Harnik,Liu:2012.01443]



lacksquare Direct probes of composite single neutrino N.



COLLIDER SIGNALS OF COMPOSITE DM AND NEUTRINO

