PHASE TRANSITIONS AND GRAVITATIONAL WAVES

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27th International Symposium on Particles, Strings and Cosmology (PASCOS 2022)

MPIK Heidelberg

July 28, 2022

Phase transitions and gravitational waves

Gravitational waves as windows into the early Universe

Gravitational waves from first order phase transitions

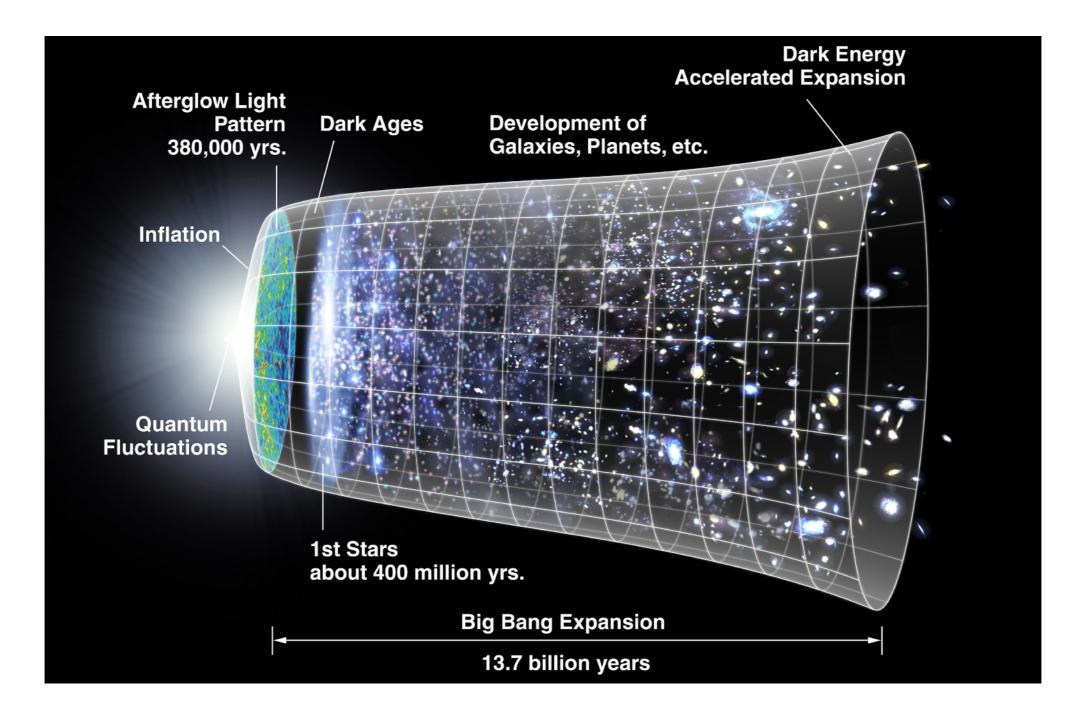
- The basics
- Some recent (and upcoming) results

The NANOGrav signal and GWs from PTs



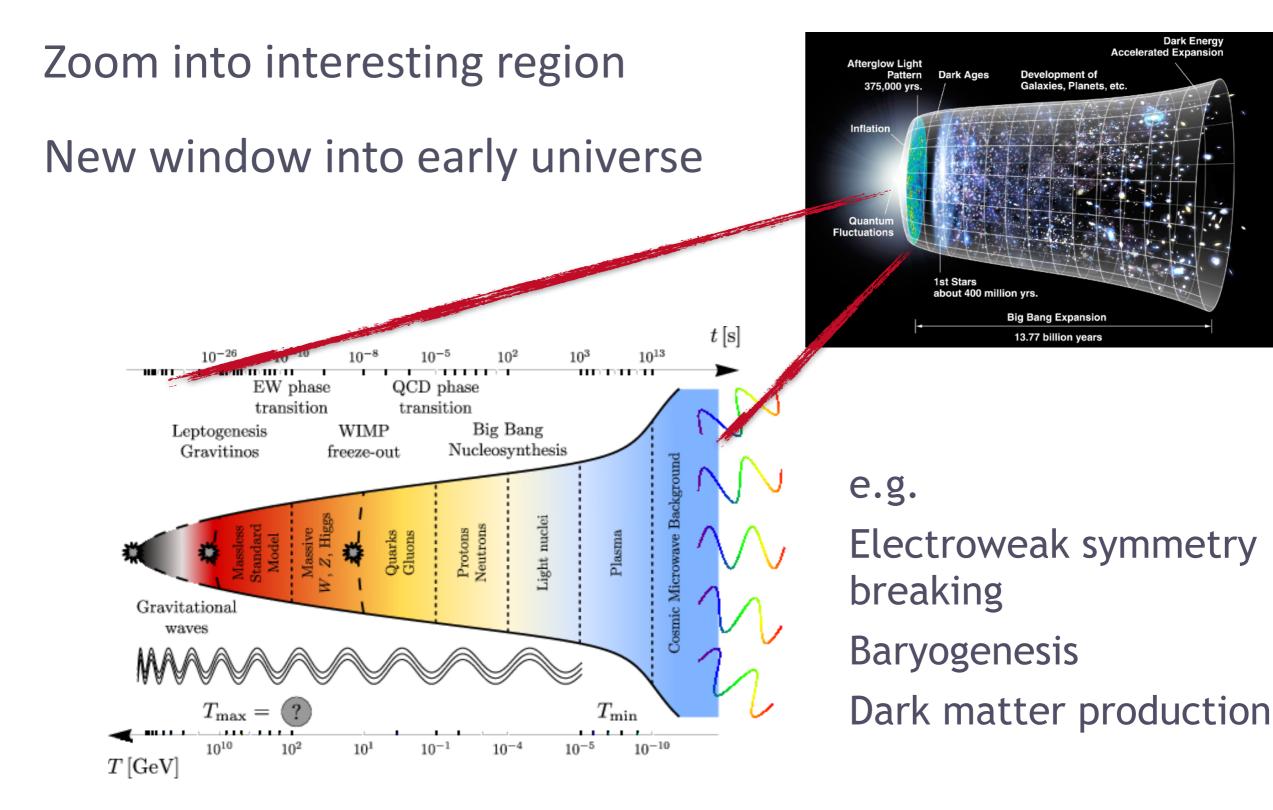
What do we know about the early Universe?

Thermal History





Gravitational Waves?





GWs & early Universe

CMB encodes information about the state of the universe at the time of its emission

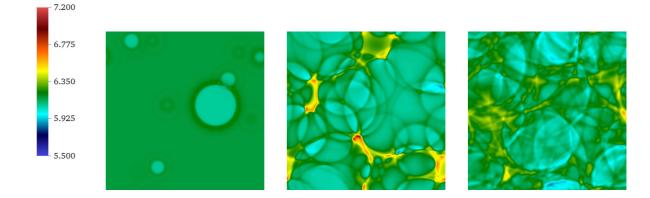
- Densities of matter, radiation, dark matter
- ► Fluctuations, Hubble rate, ...
- GWs could do the same
 - For earlier times
 - ► For different periods!
 - ▶ Need a strong source (CMB photons are just there!)





Primordial sources of GWs

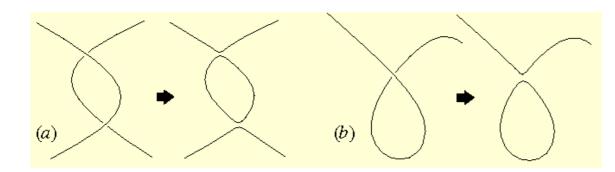
First order phase transitions (symmetry breaking)

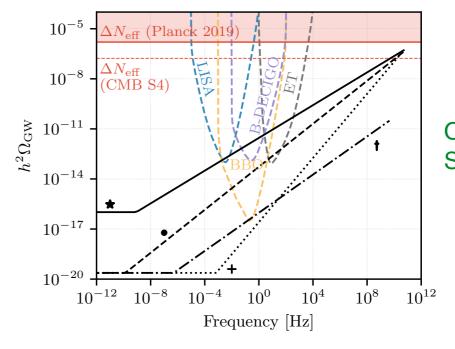


from Hindmarsh et al

Inflation/Reheating

Cosmic strings



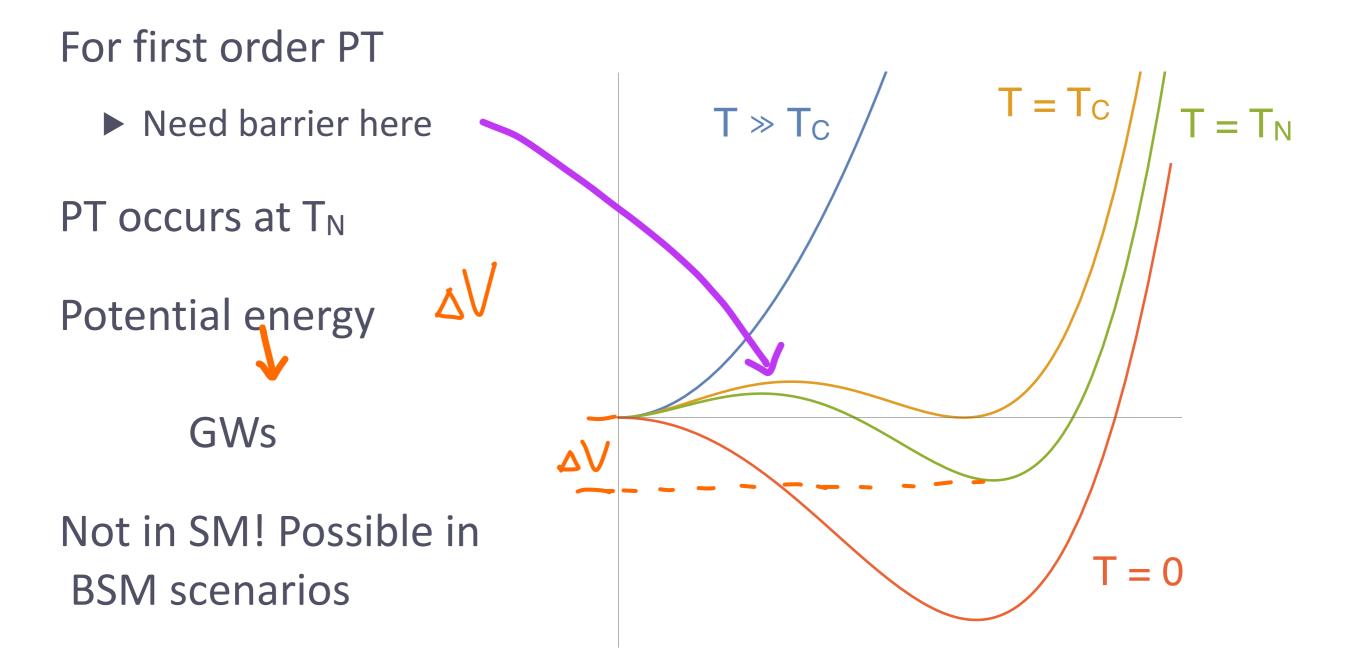


Opferkuch, PS, Stefanek, 2019



GWs from Phase Transitions

QFT at finite temperature → symmetry restoration

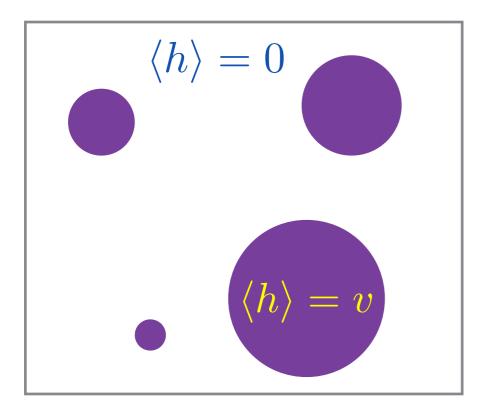


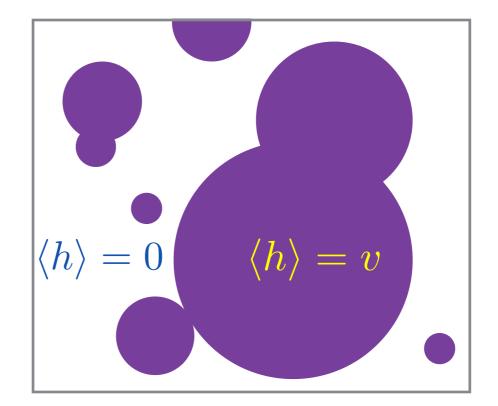


GWs from Phase Transitions

First order PT \rightarrow Bubbles nucleate, expand

Bubble collisions → Gravitational Waves







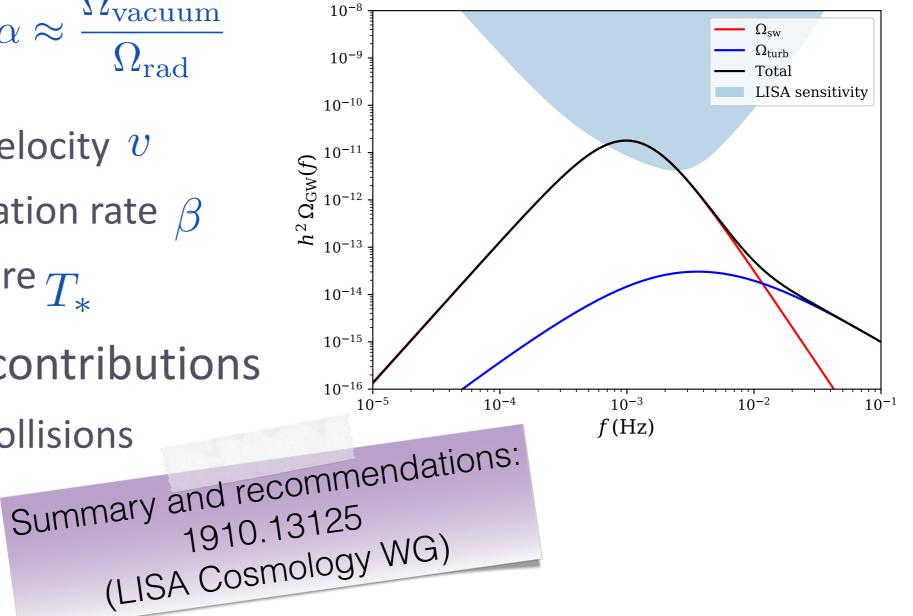
PT signal

PT characterised by few parameters:

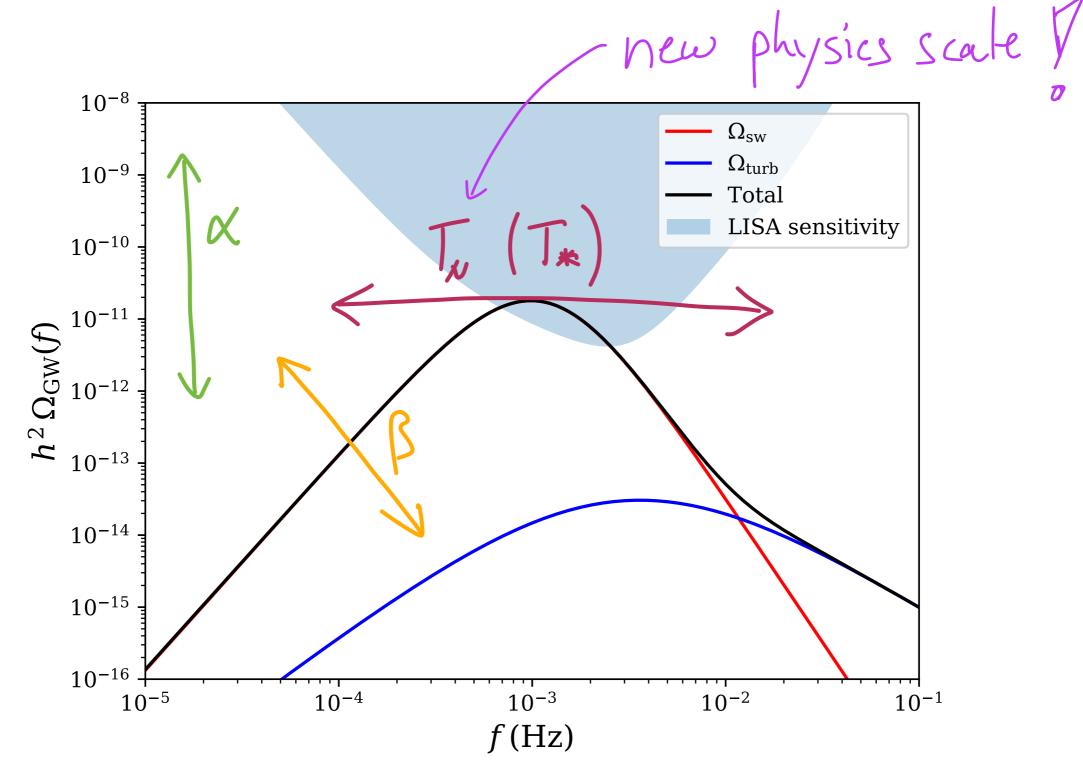
- Latent heat $\alpha \approx \frac{\Omega_{\text{vacuum}}}{\Omega_{\text{vacuum}}}$
- Bubble wall velocity v
- Bubble nucleation rate β
- PT temperature T_*

Three physical contributions

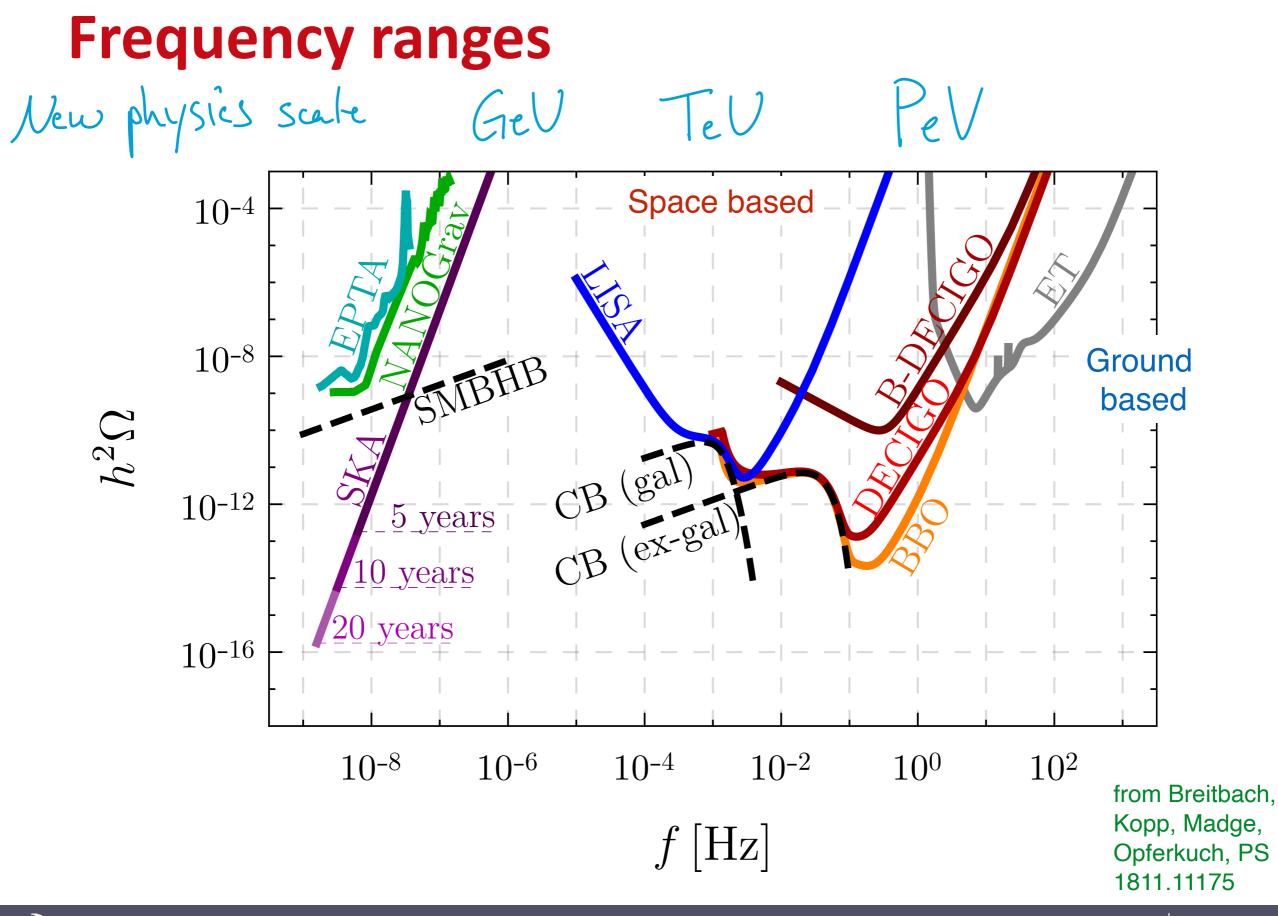
- Bubble wall collisions
- Turbulence
- Sound waves



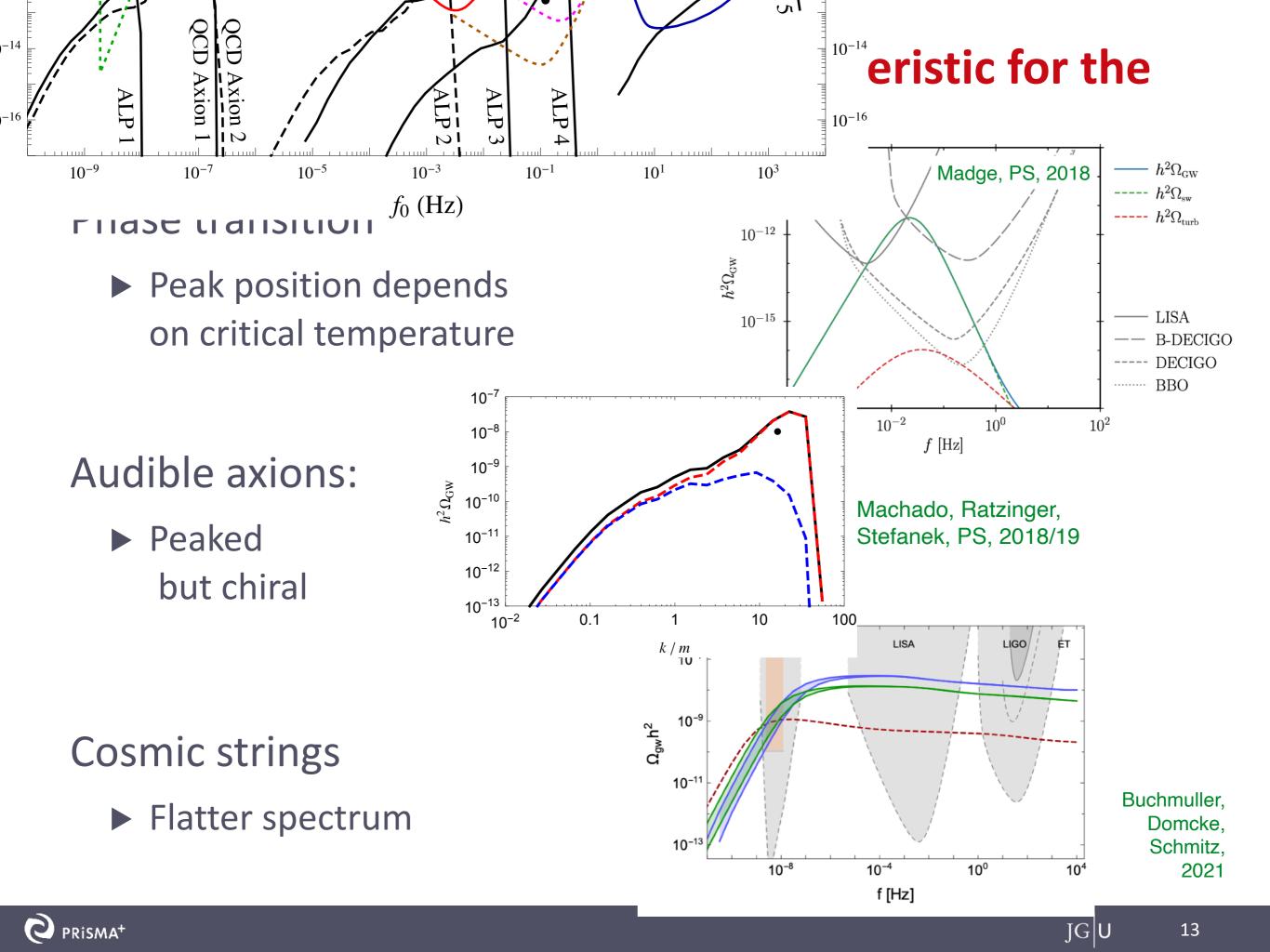
Signal properties







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Some recent progress

(with my personal bias, sorry!)

Computing the GWs from PTs is hard I

Gauge fields are finite temperature are problematic

Slow convergence of perturbation theory, nonperturbative methods partially required

Dimensional reduction + 3D lattice can overcome this

HIP-2022-11/TH NORDITA 2022-030

DRalgo: a package for effective field theory approach for thermal phase transitions

Andreas Ekstedt^{a,b,c,*}, Philipp Schicho^{d,†}, and Tuomas V. I. Tenkanen^{e,f,g,‡}

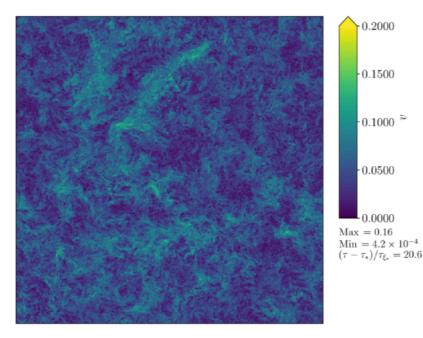




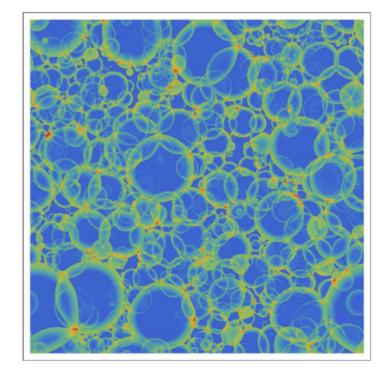
Computing the GWs from PTs is hard II

Expanding bubbles source sound waves in plasma

- Requires hydro simulations
- ► For each parameter point!



Auclair, Caprini, Cutting, Hindmarsh, Rummukainen, Steer, Weir, 2022



Jinno, Konstandin, Rubira, 2021

Many simulations, signal models are improving!



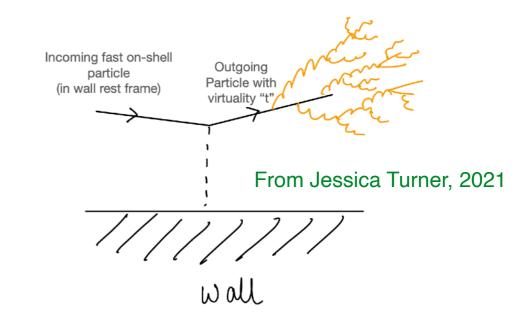
Computing the GWs from PTs is hard III

Bubble wall velocity still open question

- Strictly non-equilibrium
- Runaway?
 - Yes (Boedeker and Moore, 2009 [leading order])
 - No (Boedeker and Moore, 2017 [NLO])

New attempts using:

- Resumation Höche, Kozaczuk, Long, Turner, Wang, 2020
- N-Body simulations Lewicki, Vaskonen, Veermäe, 2022
- ► To be continued... :)







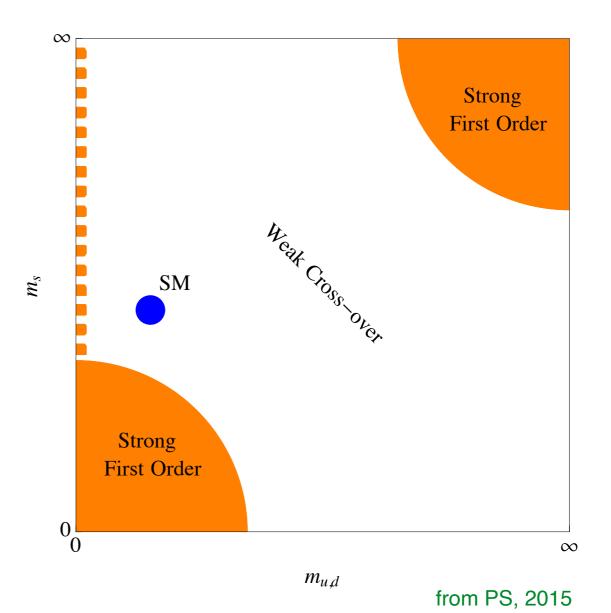
Strongly coupled PTs are also difficult

Nonabelian SU(N) dark sectors appear in many BSM scenarios (composite DM, twin Higgs, string theory)

Confinement PT is first order for many choices of N and N_f

Non-perturbative PT

Difficult to predict GW signal





Strongly coupled PTs are also difficult

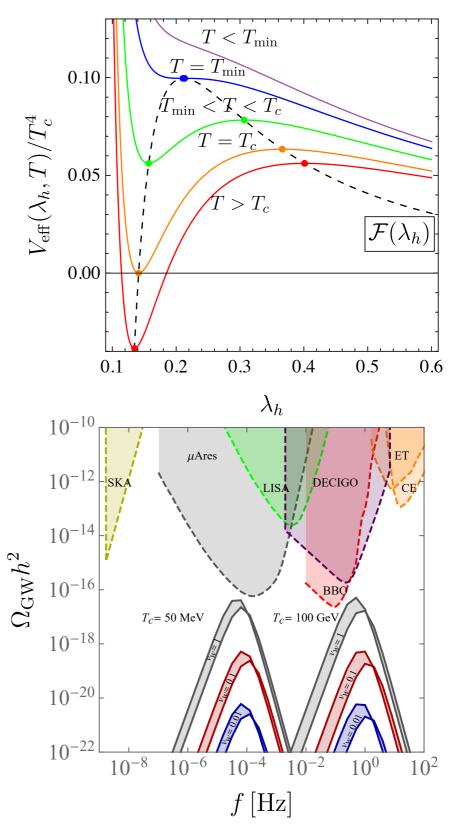
Computed thermal effective potential in improved holographic QCD

 Fit to reproduce finite T lattice data

First prediction for GW spectra of QCD-like dark sectors from holography

Enrico Morgante, Nicklas Ramberg, PS, in preparation

except for the wall velocity...

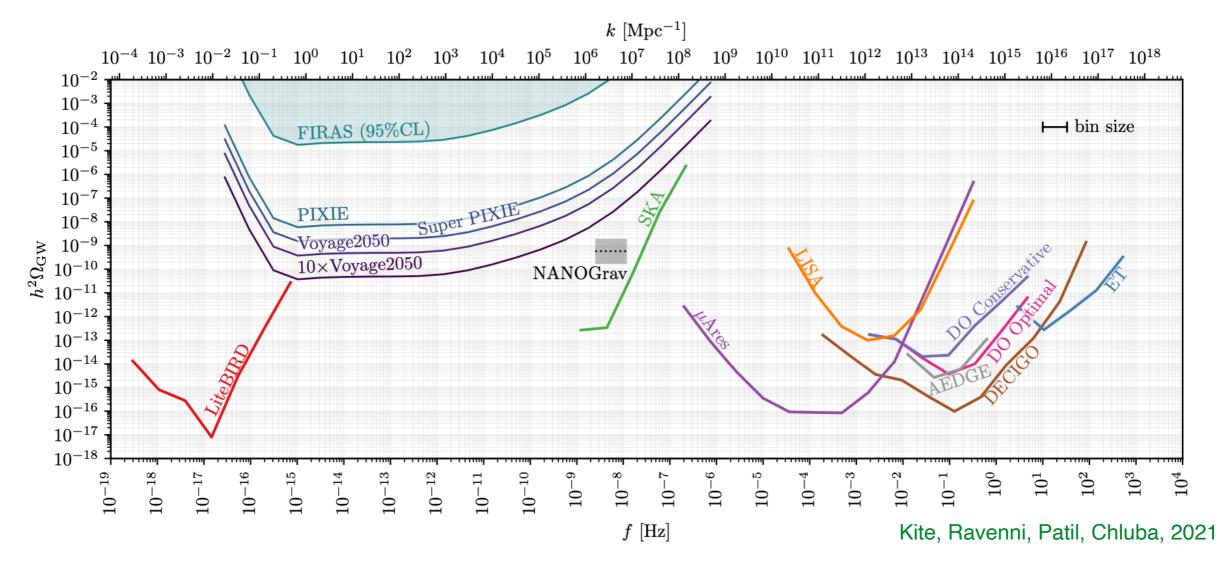


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Probing sub-MeV phase transitions

Very low frequency GWs induce CMB spectral distortions

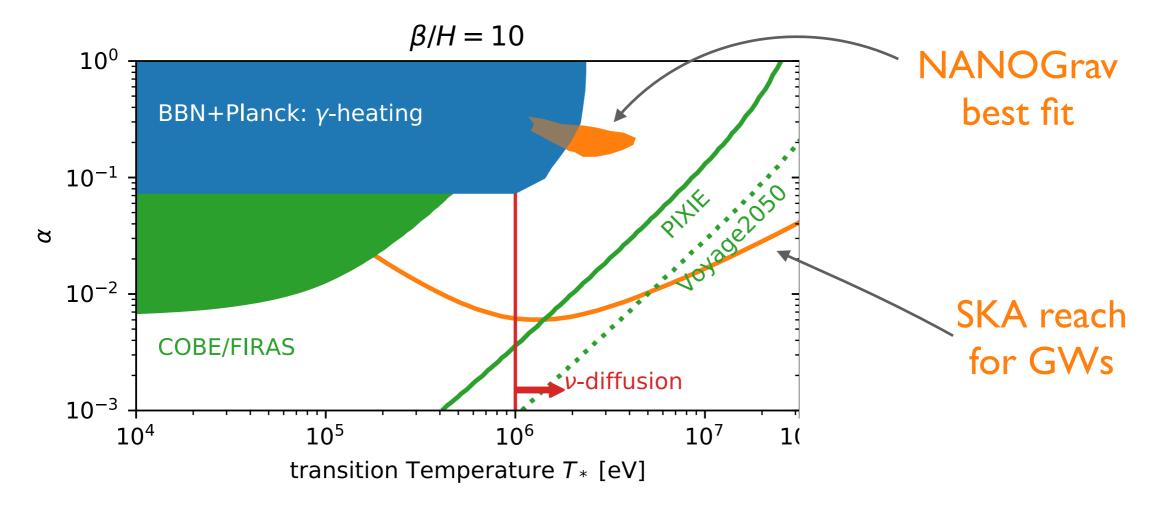


Probe sources that give peaked GW spectra (like PTs)



Probing sub-MeV phase transitions

Can also directly probe the scalar (density) fluctuations induced by PTs in a dark or visible sector



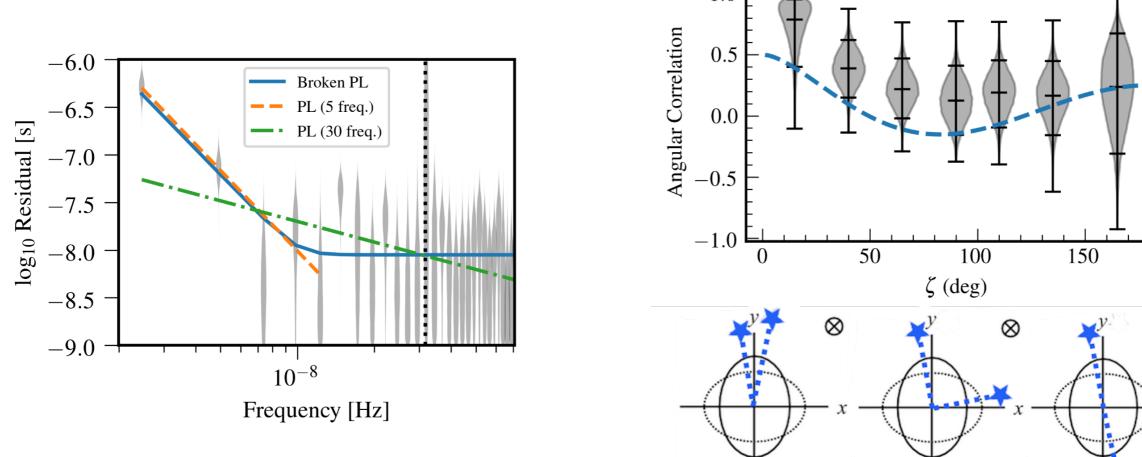
More sensitive! Multi-messenger cosmology!

Nicklas Ramberg, Wolfram Ratzinger, PS, to appear



NANOGrav saw something!

No 4σ evidence for Quadrupole



1.0

From NANOGrav collaboration, 2009.04496 Now also consistent signals in PPTA, EPTA and IPTA - still not fully conclusive though

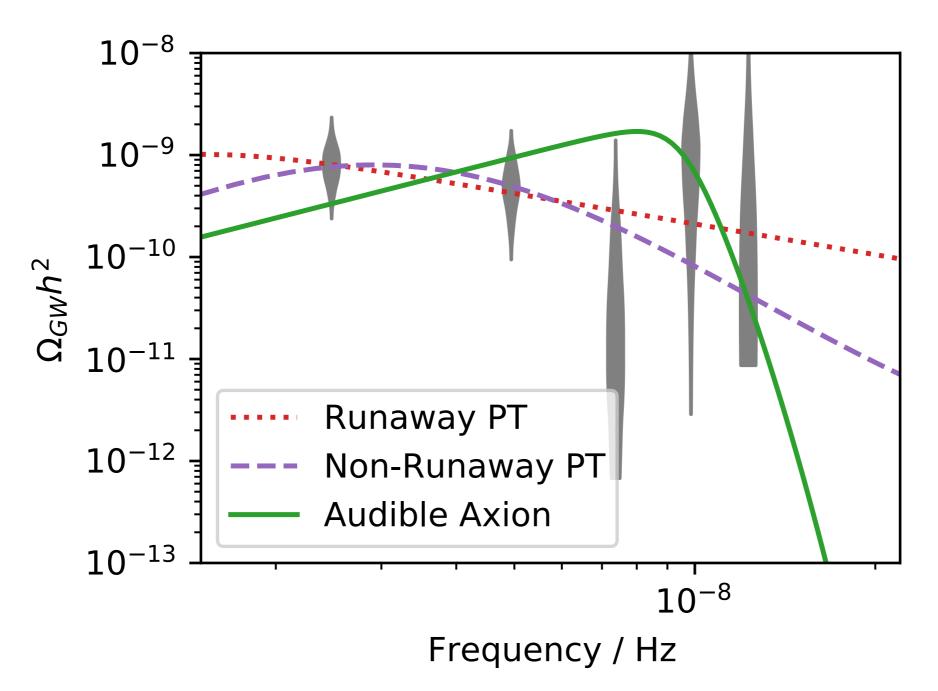




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Significant Strain at low frequencies

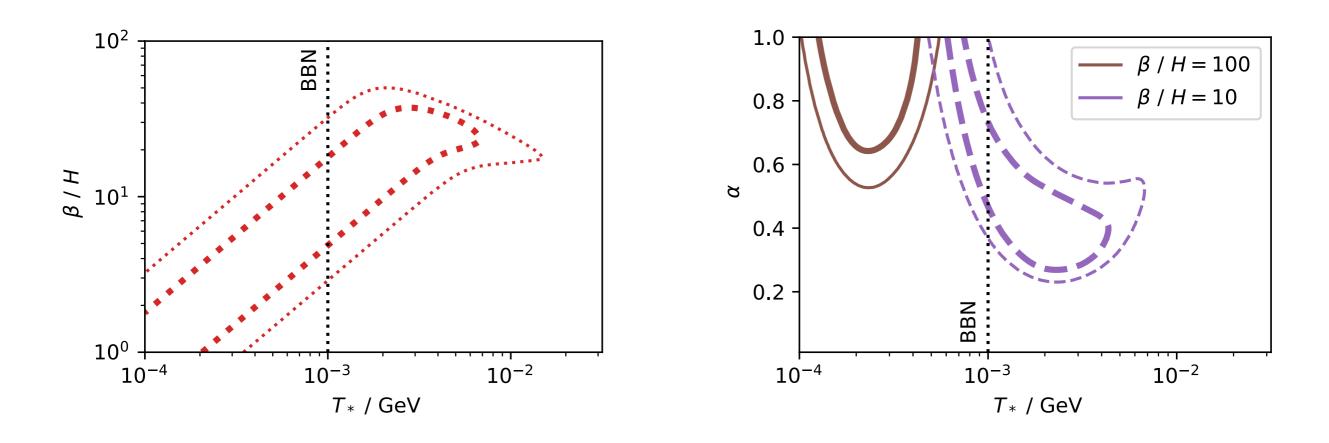
Fit with broken power law signals



Wolfram Ratzinger & PS, 2009.11875



Fit with Phase Transition

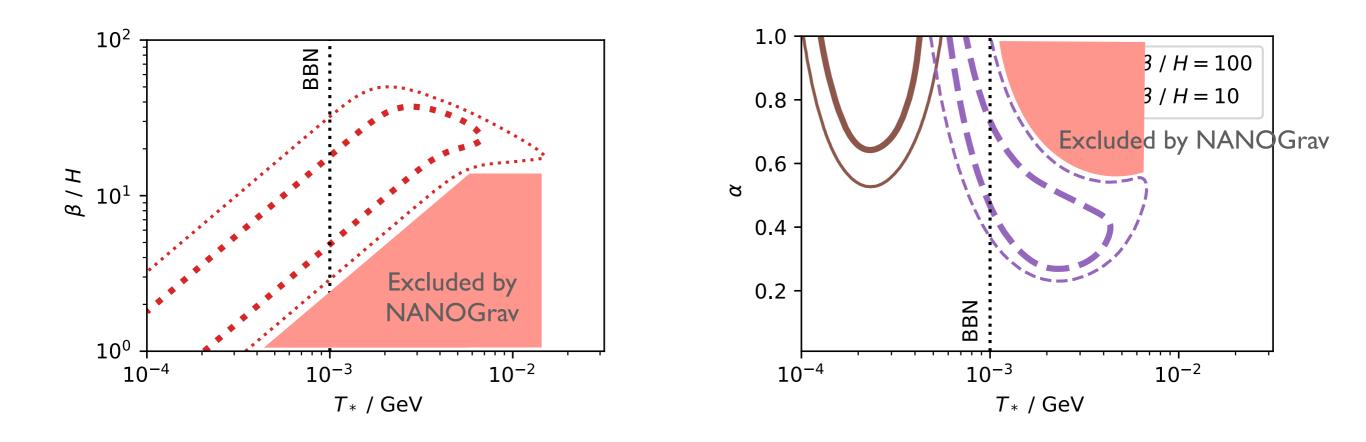


Generic PT parameterisation, best fit with PT at temperatures in few MeV range

Also here, challenging to build model that does not break cosmology (BBN and/or N_{eff})



Fit with Phase Transition



Generic PT parameterisation, best fit with PT at temperatures in few MeV range

Some model parameters excluded by PTA data now!



Summary

GWs offer new window into the early Universe

Can probe dynamics of otherwise invisible dark sectors, purely through gravitational interactions

GWs from phase transitions probe presence of new symmetries (and their breaking)

► Large theory effort to improve predictions on many frontiers

PTA data already constrains new physics models

Possibly first hint for primordial GW background

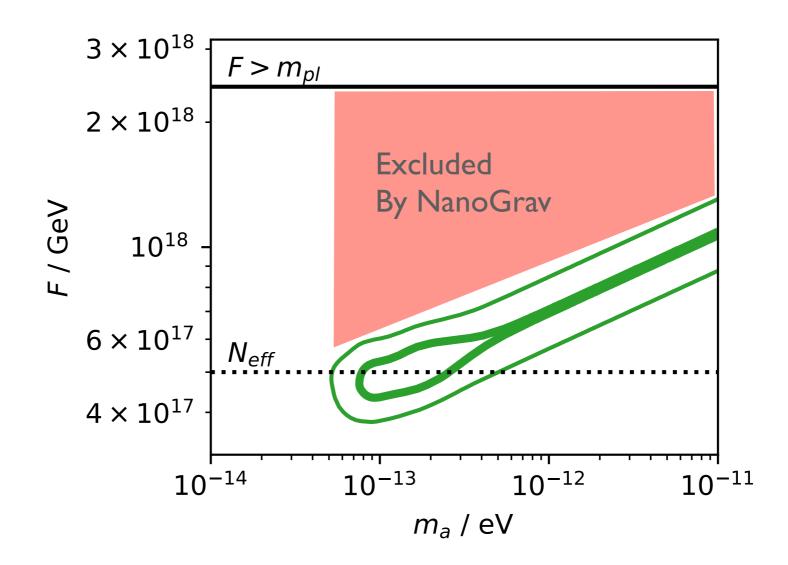
New data expected in near and far future!







Example: Audible Axion

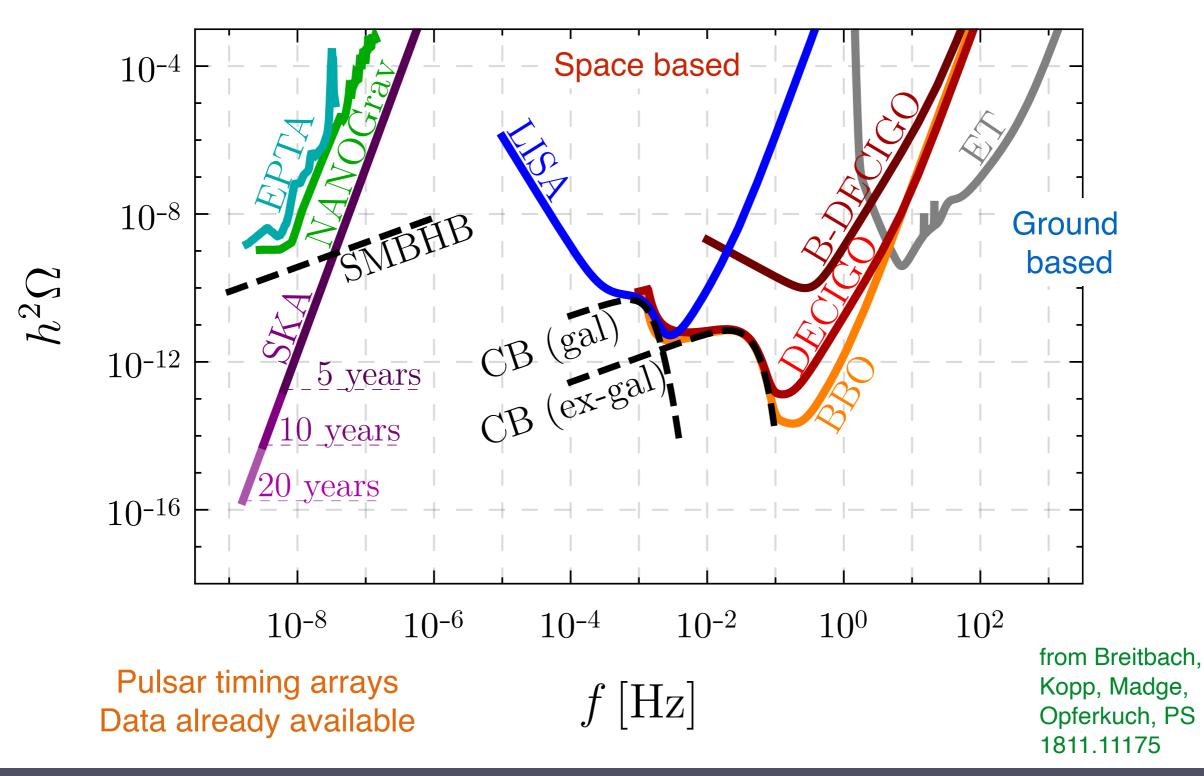


Parameter reconstruction already possible

Non-trivial constraints from cosmology (N_{eff})



Frequency ranges



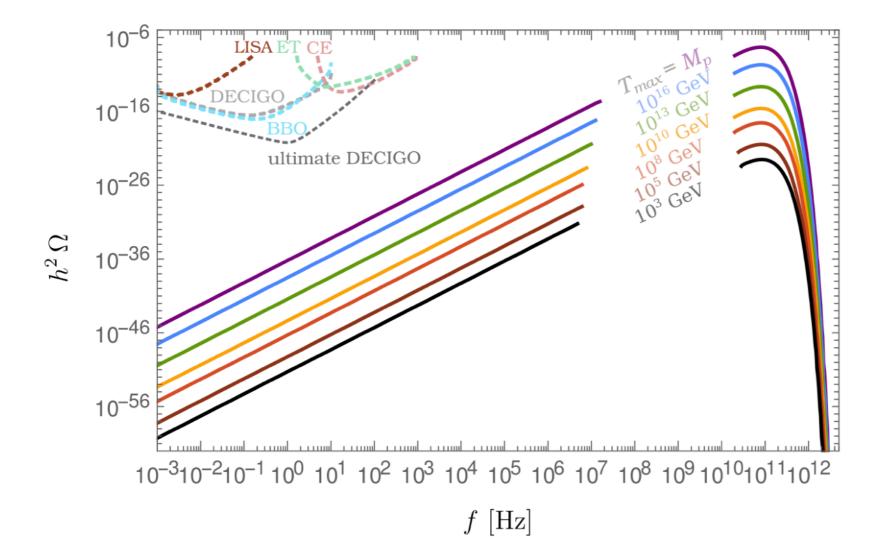
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Standard model

The hot early Universe sources GWs!

- Classical picture: thermal fluctuations source tensor fluctuations
- Quantum picture: gluon + gluon -> graviton



From Ringwald, Schütte-Engel, Tamarit, 2020

Original computations: Ghiglieri, Laine, 2015 Ghiglieri, Jackson, Laine, Zhu, 2020