



Massive quark form factors at three loops

QCD@LHC 2022 | November 28 – December 2, 2022

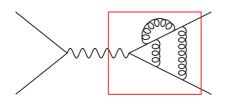
Fabian Lange

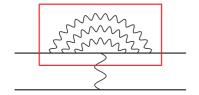
in collaboration with Matteo Fael, Kay Schönwald, Matthias Steinhauser | Nov 30, 2022



Karlsruhe Institute of Technology

Motivation

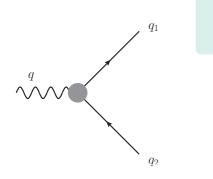




- Form factors are basic building blocks for many physical observables:
 - $t\bar{t}$ production at hadron and e^+e^- colliders
 - µe scattering
 - Higgs production and decay
 - Soft-Higgs approximation for $t\bar{t}H$ production \Rightarrow Chiara Savoini's talk
 - ...
- Form factors exhibit an universal infrared behavior which is interesting to study

The process





$$egin{aligned} X(q) & o Q(q_1) + ar{Q}(q_2) \ q_1^2 = q_2^2 = m^2, \quad q^2 = s = \hat{s} \cdot m^2 \end{aligned}$$

vector : axial-vector : scalar : pseudo-scala

$$j_{\mu}^{\mathsf{v}} = \overline{\psi} \gamma_{\mu} \psi, \qquad \Gamma_{\mu}^{\mathsf{v}} = F_{1}^{\mathsf{v}}(s) \gamma_{\mu} - \frac{\mathsf{i}}{2m} F_{2}^{\mathsf{v}}(s) \sigma_{\mu\nu} q^{\nu}$$

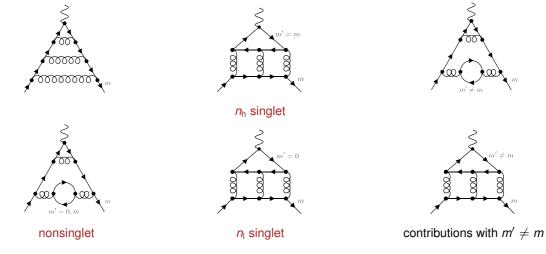
$$: \qquad j_{\mu}^{\mathsf{a}} = \overline{\psi} \gamma_{\mu} \gamma_{5} \psi, \qquad \Gamma_{\mu}^{\mathsf{a}} = F_{1}^{\mathsf{a}}(s) \gamma_{\mu} \gamma_{5} - \frac{1}{2m} F_{2}^{\mathsf{a}}(s) q_{\mu} \gamma_{5}$$

$$j^{\mathsf{s}} = m \overline{\psi} \psi, \qquad \Gamma^{\mathsf{s}} = m F^{\mathsf{s}}(s)$$

$$\mathsf{ar} : \qquad j^{\mathsf{p}} = \mathsf{i} m \overline{\psi} \gamma_{5} \psi, \qquad \Gamma^{\mathsf{p}} = \mathsf{i} m \mathcal{F}^{\mathsf{p}}(s) \gamma_{5}$$



Types of contributions





Status of massive QCD corrections

 $F_i^{(2)}$ (**NNLO**):

nonsinglet:



singlet:



(NNLO):

- fermionic contributions [Hoang, Teubner 1997]
- Complete [Bernreuther, Bonciani, Gehrmann, Heinesch, Leineweber, Mastrolia, Remiddi 2004 2005]
- $F_i^{(3)}$ (NNNLO):
 - nonsinglet large N_c [Henn, Smirnov, Smirnov, Steinhauser 2016; Lee, Smirnov, Smirnov, Steinhauser 2018; Ablinger, Blümlein, Marquard, Rana, Schneider 2 × 2018; Lee, Smirnov, Smirnov, Steinhauser 2018]
 - nonsinglet n [Lee, Smirnov, Smirnov, Steinhauser 2018; Ablinger, Blümlein, Marquard, Rana, Schneider 2 × 2018]
 - nonsinglet n_h (partially) [Blümlein, Marquard, Rana, Schneider 2019]

This talk: full (numerical) results for nonsinglet and singlet contributions at NNNLO



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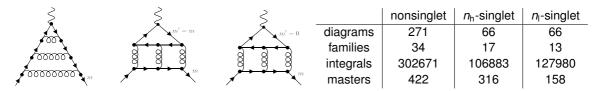
Status for massless form factors:

- F⁽⁴⁾ computed recently [Lee, von Manteuffel, Schabinger, Smirnov, Smirnov, Steinhauser 2022]
- Singlet contributions to $F_a^{(3)}$ with massive quark loop computed in [Chen, Czakon, Niggetiedt 2021]



Setup

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- Generate diagrams with qgraf [Nogueira 1991]
- Map to predefined integral families with q2e/exp [Harlander, Seidensticker, Steinhauser 1998; Seidensticker 1999]
- FORM [Vermaseren 2000; Kuipers, Ueda, Vermaseren, Vollinga 2013; Ruijl, Ueda, Vermaseren 2017] for Lorentz, Dirac, and color algebra [van Ritbergen, Schellekens, Vermaseren 1998]
- Reduction to master integrals with Kira [Maierhöfer, Usovitsch, Uwer 2017; Klappert, FL, Maierhöfer, Usovitsch 2020] and Fermat [Lewis]
 - Construct good basis where denominators factorize in ϵ and \hat{s} with ImproveMasters.m [Smirnov, Smirnov 2020]
- Establish differential equations in ŝ with LiteRed [Lee 2012 + 2013]

Algorithm to solve master integrals I



$$rac{\partial}{\partial \hat{\mathbf{s}}} M_n = A_{nm}(\epsilon, \hat{\mathbf{s}}) M_m$$

• Compute expansion around $\hat{s} = 0$ by:

Inserting an ansatz for the master integrals into the differential equation:

$$M_n(\epsilon, \hat{m{s}}=0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\max}} c_{ij}^{(n)} \, \epsilon^i \, \hat{m{s}}^j$$

• Compare coefficients in ϵ and \hat{s} to establish linear system of equations for $c_{ii}^{(n)}$:

$$c_{12}^{(1)}\epsilon\hat{s}^2 + \cdots = 52c_{33}^{(1)}\epsilon\hat{s}^2 + \cdots + 127c_{14}^{(4)}\epsilon\hat{s}^2 + \ldots$$

Solve system in terms of small number of boundary constants using Kira with FireFly [Klappert, FL 2019; Klappert, Klein, FL 2020]:

$$c_{12}^{(1)} = 52c_{33}^{(1)} + 127c_{14}^{(4)}$$

Compute boundary values to fix remaining constants

Algorithm to solve master integrals II



$$rac{\partial}{\partial \hat{s}} M_n = A_{nm}(\epsilon, \hat{s}) M_m$$

• Repeat for $\hat{s} = \hat{s}_1$:

• Insert an ansatz around $\hat{s} = \hat{s}_1$ into the differential equation:

$$M_n(\epsilon, \hat{\mathbf{s}} = \hat{\mathbf{s}}_1) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\max}} c_{ij}^{(n)} \epsilon^i (\hat{\mathbf{s}} - \hat{\mathbf{s}}_1)^j$$

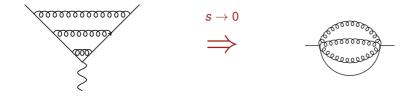
• Compare coefficients in ϵ and \hat{s} and solve system:

$$c_{12}^{(1)}\epsilon(\hat{s}-\hat{s}_1)^2+\cdots = 12c_{44}^{(1)}\epsilon(\hat{s}-\hat{s}_1)^2+\cdots - 23c_{04}^{(4)}\epsilon(\hat{s}-\hat{s}_1)^2+\ldots \qquad \Rightarrow \qquad c_{12}^{(1)} = 12c_{44}^{(1)} - 23c_{04}^{(4)} + 2$$

- Match this new expansion to previous expansion around \$\hat{s} = 0\$ numerically in between, e.g. at \$\hat{s}_1/2\$, to fix the boundary constants
- Repeat

Calculation of boundary conditions

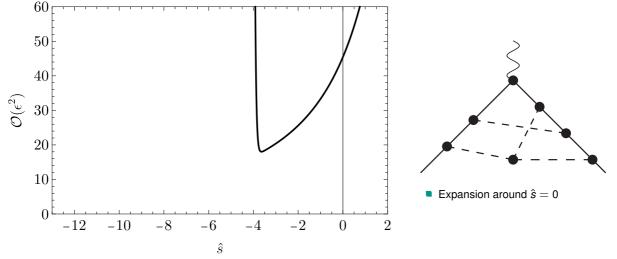




- For s = 0 the nonsinglet master integrals reduce to 3-loop on-shell propagators:
 - Well studied in the literature [Laporta, Remiddi 1996; Melnikov, van Ritbergen 1999; Lee, Smirnov 2010]
 - Some higher-order terms were missing for our application
 - Using the dimensional-recurrence relations from [Lee, Smirnov 2010] we calculated them with SummerTime.m [Lee, Mingulov 2015] and PSLQ [Ferguson, Bailey, Arno 1999]
- n_h singlet boundary conditions similar, but slightly more complicated
- $n_{\rm l}$ singlet boundary conditions: use AMFlow [Liu, Ma 2022] to compute them numerically at $\hat{s} = -1$ with 90 digits



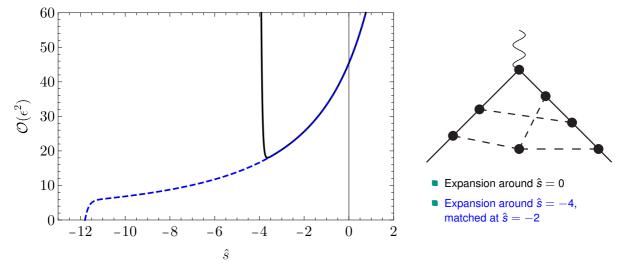
Example



Institute for Theoretical Particle Physics and Institute for Astroparticle Physics

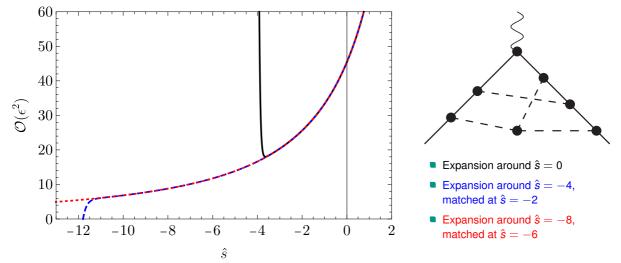


Example





Example



Results – analytic expansion around $\hat{s} = 0$

$$\begin{split} F_{1}^{\text{v,f,(3)}}(\hat{s}=0) &= \left\{ C_{\text{F}}^{3} \Big(-15a_{4} - \frac{17\pi^{2}\zeta_{3}}{24} - \frac{18367\zeta_{3}}{1728} + \frac{25\zeta_{5}}{8} - \frac{5l_{2}^{4}}{8} - \frac{19}{40}\pi^{2}l_{2}^{2} + \frac{4957\pi^{2}l_{2}}{720} + \frac{3037\pi^{4}}{25920} \right. \\ &- \frac{24463\pi^{2}}{7776} + \frac{13135}{20736} \Big) + C_{\text{A}}C_{\text{F}}^{2} \Big(\frac{19a_{4}}{2} - \frac{\pi^{2}\zeta_{3}}{9} + \frac{17725\zeta_{3}}{3456} - \frac{55\zeta_{5}}{32} + \frac{19l_{2}^{4}}{48} - \frac{97}{720}\pi^{2}l_{2}^{2} \\ &+ \frac{29\pi^{2}l_{2}}{240} - \frac{347\pi^{4}}{17280} - \frac{4829\pi^{2}}{10368} + \frac{707}{288} \Big) + C_{\text{A}}^{2}C_{\text{F}} \Big(-a_{4} + \frac{7\pi^{2}\zeta_{3}}{96} + \frac{4045\zeta_{3}}{5184} - \frac{5\zeta_{5}}{64} - \frac{l_{2}^{4}}{24} \\ &+ \frac{67}{360}\pi^{2}l_{2}^{2} - \frac{5131\pi^{2}l_{2}}{2880} + \frac{67\pi^{4}}{8640} + \frac{172285\pi^{2}}{186624} - \frac{7876}{2187} \Big) \Big\} \hat{s} + \text{fermionic corrections} + \mathcal{O}(\hat{s}^{2}) \end{split}$$

•
$$I_2 = \ln(2), a_4 = \text{Li}_4(1/2) \text{ and } C_A = 3, C_F = \frac{4}{3} \text{ for QCD}$$

• Expansions for all currents are available up to $\mathcal{O}(\hat{s}^{67})$



Results – high-energy limit

$$\begin{split} F_{1}^{\text{vl},(3)} \Big|_{s \to -\infty} &= 4.7318C_{\text{F}}^{3} - 20.762C_{\text{F}}^{2}C_{\text{A}} + 8.3501C_{\text{F}}C_{\text{A}}^{2} + \left[3.4586C_{\text{F}}^{3} - 4.0082C_{\text{F}}^{2}C_{\text{A}} - 6.3561C_{\text{F}}C_{\text{A}}^{2}\right]l_{s} \\ &+ \left[1.4025C_{\text{F}}^{3} + 0.51078C_{\text{F}}^{2}C_{\text{A}} - 2.2488C_{\text{F}}C_{\text{A}}^{2}\right]l_{s}^{2} + \left[0.062184C_{\text{F}}^{3} + 0.90267C_{\text{F}}^{2}C_{\text{A}} - 0.42778C_{\text{F}}C_{\text{A}}^{2}\right]l_{s}^{3} \\ &+ \left[-0.075860C_{\text{F}}^{3} + 0.20814C_{\text{F}}^{2}C_{\text{A}} - 0.035011C_{\text{F}}C_{\text{A}}^{2}\right]l_{s}^{4} + \left[-0.023438C_{\text{F}}^{3} + 0.019097C_{\text{F}}^{2}C_{\text{A}}\right]l_{s}^{5} \\ &+ \left[-0.0026042C_{\text{F}}^{3}\right]l_{s}^{6} - \left\{-92.918C_{\text{F}}^{3} + 123.65C_{\text{F}}^{2}C_{\text{A}} - 47.821C_{\text{F}}C_{\text{A}}^{2} + \left[-10.381C_{\text{F}}^{3} + 2.3223C_{\text{F}}^{2}C_{\text{A}} \right]l_{s}^{4} \\ &+ 17.305C_{\text{F}}C_{\text{A}}^{2}\right]l_{s} + \left[4.9856C_{\text{F}}^{3} - 19.097C_{\text{F}}^{2}C_{\text{A}} + 8.0183C_{\text{F}}C_{\text{A}}^{2}\right]l_{s}^{2} + \left[3.0499C_{\text{F}}^{3} - 6.8519C_{\text{F}}^{2}C_{\text{A}} + 1.9149C_{\text{F}}C_{\text{A}}^{2}\right]l_{s}^{5} \\ &+ \left[0.67172C_{\text{F}}^{3} - 0.91213C_{\text{F}}^{2}C_{\text{A}} + 0.24069C_{\text{F}}C_{\text{A}}^{2}\right]l_{s}^{4} + \left[0.13229C_{\text{F}}^{3} - 0.051389C_{\text{F}}^{2}C_{\text{A}} + 0.0043403C_{\text{F}}C_{\text{A}}^{2}\right]l_{s}^{5} \\ &+ \left[0.0041667C_{\text{F}}^{3} - 0.0010417C_{\text{F}}^{2}C_{\text{A}} - 0.00052083C_{\text{F}}C_{\text{A}}^{2}\right]l_{s}^{6} + \mathcal{O}\left(\frac{m^{4}}{s^{2}}\right) + \text{fermionic contributions} \end{split}$$



Results – high-energy limit

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Dedicated calculation of leading logarithms [Liu, Penin, Zerf 2017] :

$$F_{1}^{v,f,(3)} = -\frac{C_{\rm F}^{3}}{384} I_{\rm s}^{6} - \frac{m^{2}}{s} \left(\frac{C_{\rm F}^{3}}{240} - \frac{C_{\rm F}^{2}C_{\rm A}}{960} - \frac{C_{\rm F}C_{\rm A}^{2}}{1920} \right) I_{\rm s}^{6} + \dots, \quad \text{with } I_{\rm s} = \ln\left(\frac{m^{2}}{-s}\right) I_{\rm s}^{6} + \dots,$$

• We reproduce these terms with high precision

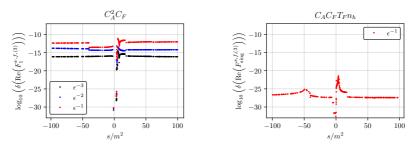


Results – pole cancellation

- We use the pole cancellation to estimate the precision
- To estimate the number of significant digits we use

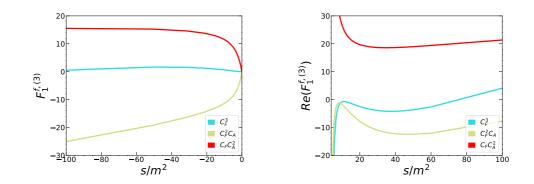
$$\log_{10}\left(\left|\frac{\text{expansion} - \text{analytic CT}}{\text{analytic CT}}\right|\right)$$

- \Rightarrow We estimate at least 8 correct digits for the finite terms
- Most regions for most color factors and especially singlet contributions much more precise



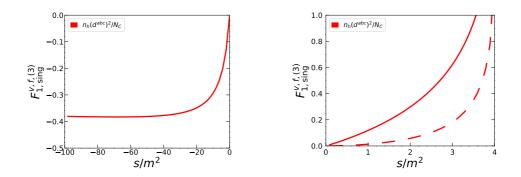


Results – some plots: nonsinglet





Results – some plots: *n*_h **singlet**





Conclusions and outlook

Conclusions

- Calculated massive quark form factors at NNNLO in QCD
 - Nonsinglet and non-anomalous n_h-singlet contributions published in [Fael, FL, Schönwald, Steinhauser 2 × 2022]
- Applied a semianalytic method by constructing series expansions and matching numerically
- Reproduce known results from the literature, e.g.
 - large- N_c limit, n_l and partial n_h contributions
 - static, high-energy, and threshold expansions
- Estimate precision to at least 8 significant digits over the whole real axis
- Extracted matching coefficients between QCD and NRQCD [Egner, Fael, FL, Schönwald, Steinhauser 2022]



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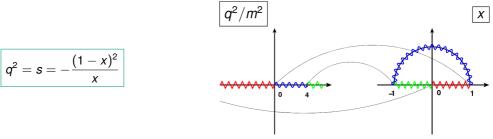
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Outlook

- Anomalous $n_{\rm h}$ -singlet as well as $n_{\rm l}$ -singlet contributions to be published soon
- $\gamma^{\star} \rightarrow \ell \bar{\ell}$ in QED for proposed MUonE experiment most realistic first phenomenological application



Why numerical?



Large-N_c and n_l contributions at NNNLO can be written as iterated integrals over letters

$$\frac{1}{x}, \ \frac{1}{1+x}, \ \frac{1}{1-x}, \ \frac{1}{1-x+x^2}, \ \frac{x}{1-x+x^2}$$

- n_h terms already contain structures beyond iterated integrals (elliptic integrals)
- \Rightarrow No ready-to-use tools available for analytic solution
- \Rightarrow Instead: Full solution through analytic series expansions and numerical matching



Different ansätze for different points:

regular point:

$$M_n(\epsilon, \hat{s} = \hat{s}_0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{max}} c_{ij}^{(n)} \, \epsilon^i \, (\hat{s} - \hat{s}_0)^j$$



 $s = \pm \infty$ (high-energy limit):

Different ansätze for different points:

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$$egin{aligned} &M_n(\epsilon, \hat{m{s}} = \hat{m{s}}_0) = \sum\limits_{i=-3}^\infty \sum\limits_{j=0}^{j_{ ext{max}}} c_{ij}^{(n)} \, \epsilon^i \, (\hat{m{s}} - \hat{m{s}}_0)^j \ &M_n(\epsilon, \hat{m{s}} o \pm\infty) = \sum\limits_{i=-3}^\infty \sum\limits_{j=-s_{ ext{min}}}^{j_{ ext{max}}} \sum\limits_{k=0}^{i+6} c_{ijk}^{(n)} \, \epsilon^i \, \hat{m{s}}^{-j} \, \ln^k(\hat{m{s}}) \end{aligned}$$



Different ansätze for different points:

regular point:

 $s=\pm\infty$ (high-energy limit):

 $s = 4m^2$ (2-particle threshold):

$$\begin{split} & \textit{M}_{\textit{n}}(\epsilon, \hat{s} = \hat{s}_{0}) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{max}} c_{ij}^{(n)} \, \epsilon^{i} \, (\hat{s} - \hat{s}_{0})^{j} \\ & \textit{M}_{\textit{n}}(\epsilon, \hat{s} \to \pm \infty) = \sum_{i=-3}^{\infty} \sum_{j=-s_{min}}^{j_{max}} \sum_{k=0}^{i+6} c_{ijk}^{(n)} \, \epsilon^{i} \, \hat{s}^{-j} \, \ln^{k} \left(\hat{s} \right) \\ & \textit{M}_{\textit{n}}(\epsilon, \hat{s} = 4) = \sum_{i=-3}^{\infty} \sum_{j=-s_{min}}^{j_{max}} \sum_{k=0}^{i+3} c_{ijk}^{(n)} \, \epsilon^{i} \, \left[\sqrt{4 - \hat{s}} \right]^{j} \, \ln^{k} \left(\sqrt{4 - \hat{s}} \right) \end{split}$$



Different ansätze for different points:

regular point:

 $s = \pm \infty$ (high-energy limit):

 $s = 4m^2$ (2-particle threshold):

i = -3 $i = -s_{min}$ $M_n(\epsilon, \hat{s} = 4) = \sum_{i=1}^{\infty} \sum_{j=1}^{j_{\max}} \sum_{i=1}^{i+3} c_{ijk}^{(n)} \epsilon^i \left[\sqrt{4-\hat{s}}
ight]^j \ln^k \left(\sqrt{4-\hat{s}}
ight)$ $s = 16m^2 \text{ (4-particle threshold):} \quad M_n(\epsilon, \hat{s} = 16) = \sum_{i=-3}^{\infty} \sum_{j=-5}^{j_{\text{max}}} \sum_{k=-6}^{i+3} c_{ijk}^{(n)} \epsilon^i \left[\sqrt{16-\hat{s}}\right]^j \ln^k \left(\sqrt{16-\hat{s}}\right)^{i-1}$

 $M_n(\epsilon, \hat{\mathbf{s}} = \hat{\mathbf{s}}_0) = \sum_{i=-3}^{\infty} \sum_{i=0}^{j_{\text{max}}} c_{ij}^{(n)} \epsilon^i (\hat{\mathbf{s}} - \hat{\mathbf{s}}_0)^j$

 $M_n(\epsilon, \hat{s}
ightarrow \pm \infty) = \sum_{i=1}^{\infty} \sum_{j=1}^{j_{max}} \sum_{i=1}^{i+6} c_{ijk}^{(n)} \epsilon^i \, \hat{s}^{-j} \, \ln^k(\hat{s})$



Different ansätze for different points:

regular point:

 $s = \pm \infty$ (h

 $s = 4m^2$ (2)

$$s = \pm \infty \text{ (high-energy limit):} \qquad M_n(\epsilon, \hat{s} \to \pm \infty) = \sum_{i=-3}^{\infty} \sum_{j=-s_{\min}}^{j_{\max}} \sum_{k=0}^{i+b} c_{ijk}^{(n)} \epsilon^i \hat{s}^{-j} \ln^k (\hat{s})$$

$$s = 4m^2 \text{ (2-particle threshold):} \qquad M_n(\epsilon, \hat{s} = 4) = \sum_{i=-3}^{\infty} \sum_{j=-s_{\min}}^{j_{\max}} \sum_{k=0}^{i+3} c_{ijk}^{(n)} \epsilon^i \left[\sqrt{4-\hat{s}}\right]^j \ln^k \left(\sqrt{4-\hat{s}}\right)$$

$$s = 16m^2 \text{ (4-particle threshold):} \qquad M_n(\epsilon, \hat{s} = 16) = \sum_{i=-3}^{\infty} \sum_{j=-s_{\min}}^{j_{\max}} \sum_{k=0}^{i+3} c_{ijk}^{(n)} \epsilon^i \left[\sqrt{16-\hat{s}}\right]^j \ln^k \left(\sqrt{16-\hat{s}}\right)$$

• We construct expansions up to $j_{max} = 50$ around

$$\hat{s} = \{-\infty, -32, -28, -24, -16, -12, -8, -4, 0, 1, 2, \frac{5}{2}, 3, \frac{7}{2}, 4, \frac{9}{2}, 5, 6, 7, 8, 10, 12, 14, 15, 16, 17, 19, 22, 28, 40\}$$

 $M_n(\epsilon, \hat{s} = \hat{s}_0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{max}} c_{ij}^{(n)} \epsilon^i (\hat{s} - \hat{s}_0)^j$

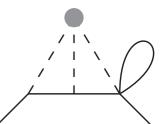
and similar for the $n_{\rm h}$ -singlet contributions

 $(\sqrt{4} - \hat{s})$

Calculation of boundary conditions: *n*_h singlet

- Due to massless cuts we need an asymptotic expansion
- Hard region solved with the same methods as in nonsinglet case
- Other regions:
 - Scalings identified with asy.m [Jantzen, Smirnov, Smirnov 2012]
 - α parameters integrated directly or with HyperInt [Panzer 2014]

$$J_{3} = y^{-4\epsilon} \frac{\Gamma(\epsilon - 1)\Gamma(2\epsilon)}{2} \int_{0}^{\infty} d\alpha_{3} \int_{0}^{\infty} d\alpha_{4} \int_{0}^{\infty} d\alpha_{6} \frac{\alpha_{4}^{-2\epsilon} \alpha_{6}^{-2\epsilon} (\alpha_{6} + \alpha_{4}(1 + \alpha_{6}))^{3\epsilon - 1}}{\alpha_{4} + \alpha_{6} + 2\alpha_{3}\alpha_{6} + \alpha_{3}^{2}(1 + \alpha_{6})}$$
$$= y^{-4\epsilon} \pi^{2} e^{-3\gamma_{E}\epsilon} \left\{ -\frac{1}{6\epsilon^{2}} - \frac{7}{6\epsilon} + \frac{13\pi^{2}}{72} - \frac{43}{6} + \epsilon \left(\frac{59\zeta_{3}}{6} - \frac{259}{6} + \frac{91\pi^{2}}{72} \right) + \mathcal{O}(\epsilon^{2}) \right\}$$





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Moebius Transformations

- The radius of convergence is at most the distance to the closest singularity.
- We can extend the radius of convergence by changing to a new expansion variable.
- If we want to expand around the point x_k with the closest singularities at x_{k-1} and x_{k+1} , we can use:

$$y_k = rac{(x-x_k)(x_{k+1}-x_{k-1})}{(x-x_{k+1})(x_{k-1}-x_k)+(x-x_{k-1})(x_{k+1}-x_k)}$$

• The variable change maps $\{x_{k-1}, x_k, x_{k+1}\} \rightarrow \{-1, 0, 1\}$.

Similar algorithms in the literature



Other approaches based on differential equations and series expansions:

- SolveCoupledSystems.m [Blümlein, Schneider 2017]
- DESS.m [Lee, Smirnov, Smirnov 2017]
- DiffExp.m [Hidding 2020] ⇒ Martijn Hidding's talk
- AMFlow [Liu, Ma 2022]
- SeaSyde.m [Armadillo, Bonciani, Devoto, Rana, Vicini 2022]

• ...

Our approach ...

- ... is tailored to problems with one real-valued kinematic variable
- ... does not require a special form for differential equations (except to be almost pole free on the diagonal)
- ... provides approximations over the whole kinematic range
- ... was successfully applied to physical quantities with 339, 422, and 316 master integrals [Fael, FL, Schönwald, Steinhauser 2021 + 2022]

Renormalization and infrared structure



UV renormalization

- \blacksquare $\overline{\rm MS}$ renormalization of $\alpha_{\rm s}$
- On-shell renormalization of mass Z^{OS}_m, wave function Z^{OS}₂, and (if needed) currents [Chetyrkin, Steinhauser 1999; Melnikov, van Ritbergen 2000]
- Much more involved renormalization for the axial and pseudoscalar singlet contributions

IR subtraction

- Structure of infrared poles given by cusp anomalous dimension Γ_{cusp} [Grozin, Henn, Korchemski, Marquard 2014]
- Define finite form factors $F = Z_{IR}F^{finite}$ with UV-renormalized form factor F and

$$Z_{\rm IR} = 1 - \frac{\alpha_s}{\pi} \frac{1}{2\epsilon} \Gamma_{\rm cusp}^{(1)} - \left(\frac{\alpha_s}{\pi}\right)^2 \left(\frac{\cdots}{\epsilon^2} + \frac{1}{4\epsilon} \Gamma_{\rm cusp}^{(2)}\right) - \left(\frac{\alpha_s}{\pi}\right)^3 \left(\frac{\cdots}{\epsilon^3} + \frac{\cdots}{\epsilon^2} + \frac{1}{6\epsilon} \Gamma_{\rm cusp}^{(3)}\right)$$

- $\Gamma_{\text{cusp}} = \Gamma_{\text{cusp}}(x)$ depends on kinematics
- Γ_{cusp} universal for all currents

Results – threshold expansion around $s = 4m^2$



Close to threshold we can construct cross-sections and decay rates like

$$\sigma(e^+e^- \to Q\bar{Q}) = \sigma_0\beta \underbrace{\left(\left| F_1^{\nu} + F_2^{\nu} \right|^2 + \frac{\left| (1 - \beta^2) F_1^{\nu} + F_2^{\nu} \right|^2}{2(1 - \beta^2)} \right)}_{=3/2\,\Delta^{\nu}}$$

with the quark velocity $\beta = \sqrt{1 - 4m^2/s}$

- Real radiation suppressed by β^3
- ⇒ Direct phenomenological relevance
- We find (with $I_{2\beta} = \ln(2\beta)$)

$$\begin{split} \Delta^{\nu,(3)} &= C_{\mathsf{F}}^3 \Big[-\frac{32.470}{\beta^2} + \frac{1}{\beta} \big(14.998 - 32.470 \mathit{l}_{2\beta} \big) \Big] + C_{\mathsf{A}}^2 C_{\mathsf{F}} \frac{1}{\beta} \Big[16.586 \mathit{l}_{2\beta}^2 - 22.572 \mathit{l}_{2\beta} + 42.936 \Big] \\ &+ C_{\mathsf{A}} C_{\mathsf{F}}^2 \Big[\frac{1}{\beta^2} \big(-29.764 \mathit{l}_{2\beta} - 7.7703 \big) + \frac{1}{\beta} \big(-12.516 \mathit{l}_{2\beta} - 11.435 \big) \Big] \\ &+ \mathcal{O}(\beta^0) + \text{fermionic contributions} \end{split}$$

Agrees with dedicated calculation [Kiyo, Maier, Maierhöfer, Marguard 2009]