

The QCD Equation of State

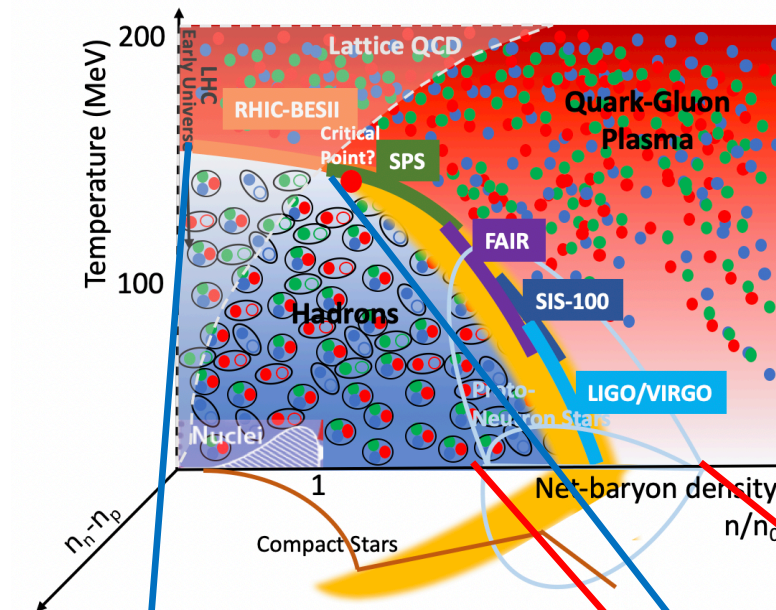
CLAUDIA RATTI

UNIVERSITY OF HOUSTON

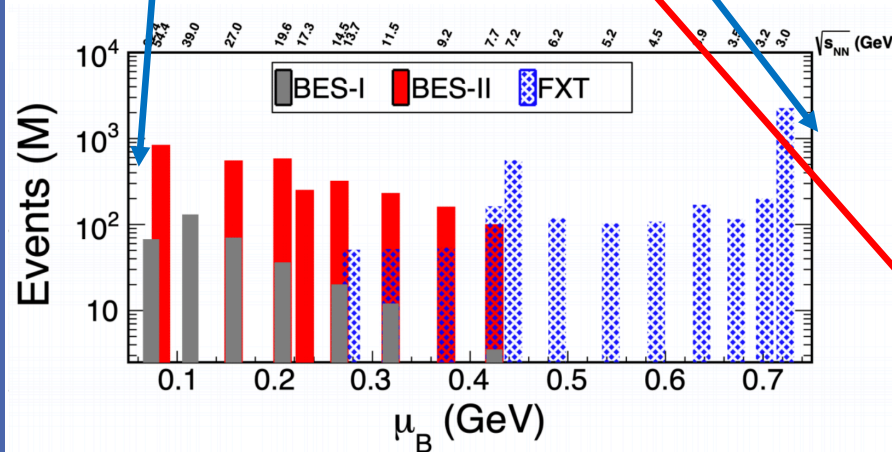


Motivating science goals

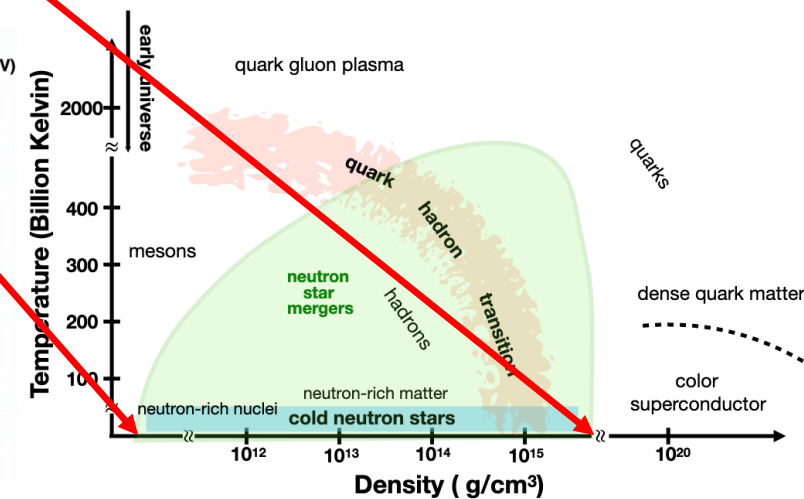
- Is there a critical point in the QCD phase diagram?
- What are the degrees of freedom in the vicinity of the phase transition?
- Where is the transition line at high density?
- What are the phases of QCD at high density?
- Are we creating a thermal medium in experiments?



- Run 2019:
 - Collider: $v_{NN}=14.6, 19.6, 200$ GeV
 - Fixed target: $v_{NN}=3.2$ GeV
- Run 2020:
 - Collider: $v_{NN}=9.2, 11.5$ GeV
 - Fixed target: $v_{NN}=3.5, 3.9, 4.5, 5.2, 6.2, 7.2, 7.7$ GeV
- Run 2021:
 - Collider: $v_{NN}=7.7, 17.3$ GeV
 - Fixed target: $v_{NN}=3.0, 9.2, 11.5, 13.7$ GeV



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Comparison of the facilities

Compilation by D. Cebra

CP=Critical Point

OD= Onset of
Deconfinement

DHM=Dense
Hadronic Matter

Facility	RHIC BESII	SPS	NICA	SIS-100 SIS-300	J-PARC HI
Exp.:	STAR +FXT	NA61	MPD + BM@N	CBM	JHITS
Start:	2019-2021	2009	2023	2022	2025
Energy:	7.7– 19.6	4.9-17.3	2.7 - 11	2.7-8.2	2.0-6.2
$\sqrt{s_{NN}}$ (GeV)	2.5-7.7		2.0-3.5		
Rate:	100 HZ	100 HZ	<10 kHz	<10 MHZ	100 MHZ
At 8 GeV	2000 Hz				
Physics:	CP&OD	CP&OD	OD&DHM	OD&DHM	OD&DHM

Collider
Fixed target

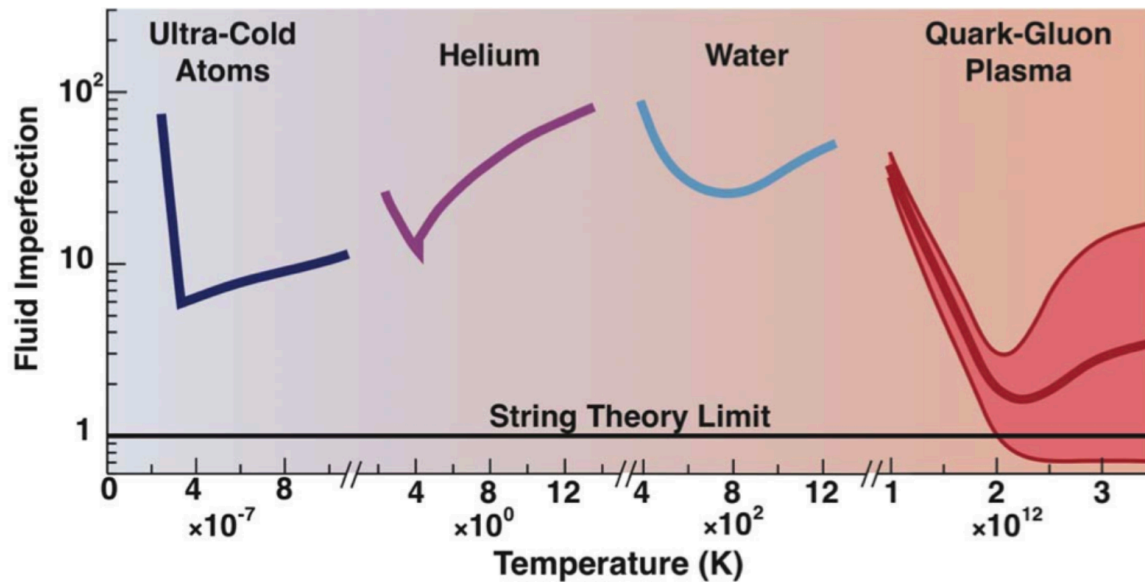
Fixed target
Lighter ion
collisions

Collider
Fixed target

Fixed target

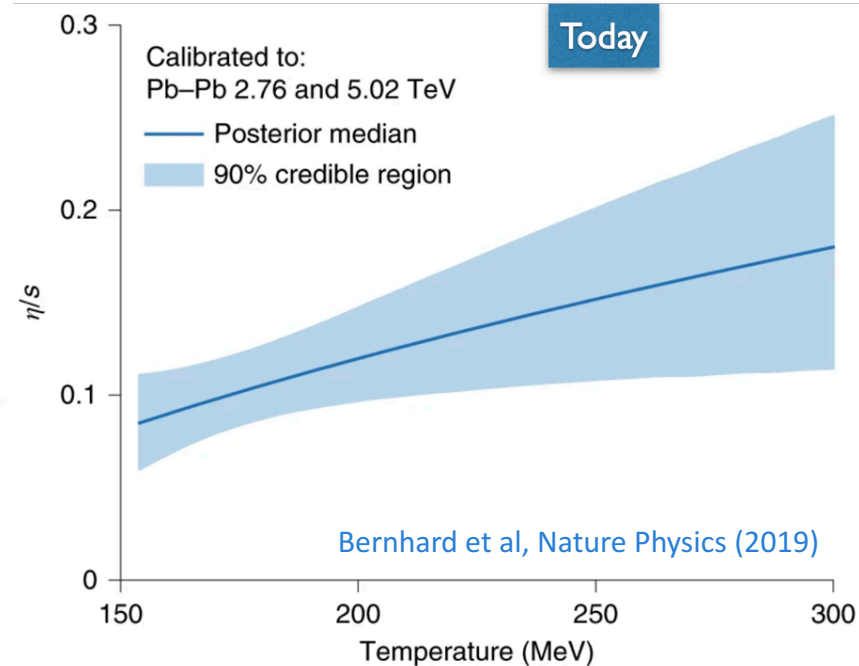
Fixed target

Nearly perfect fluidity



Kovtun, Son, Starinets,
PRL (2005)

$$\eta/s = 1/4\pi$$

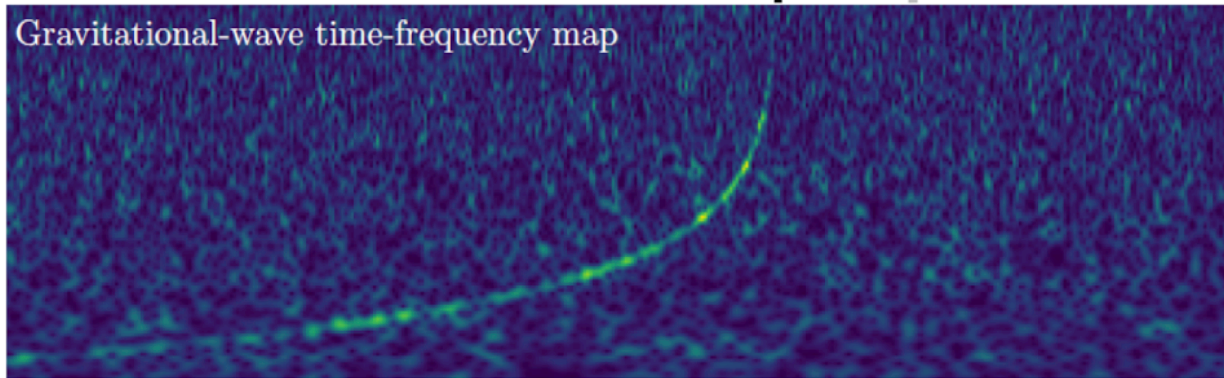
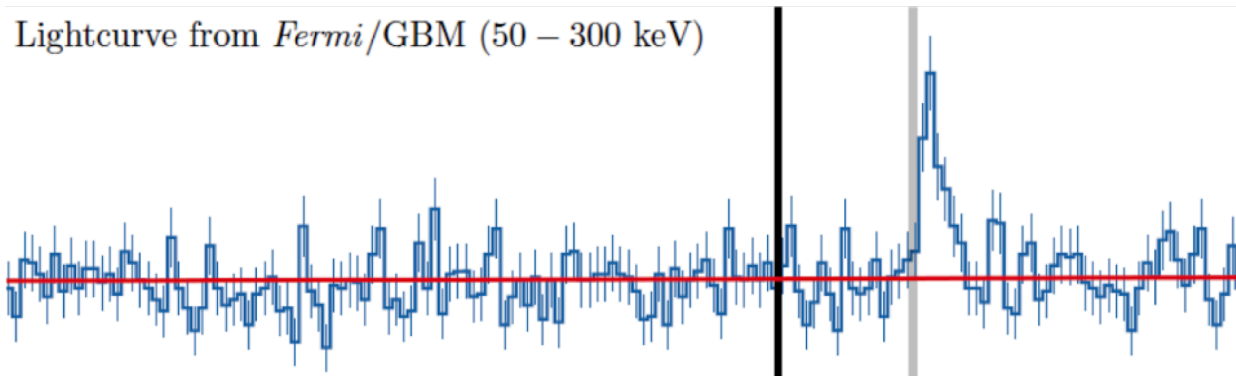


- Hydrodynamic description of the system created in heavy-ion collisions works extremely well
- It needs an equation of state as input

GW170817

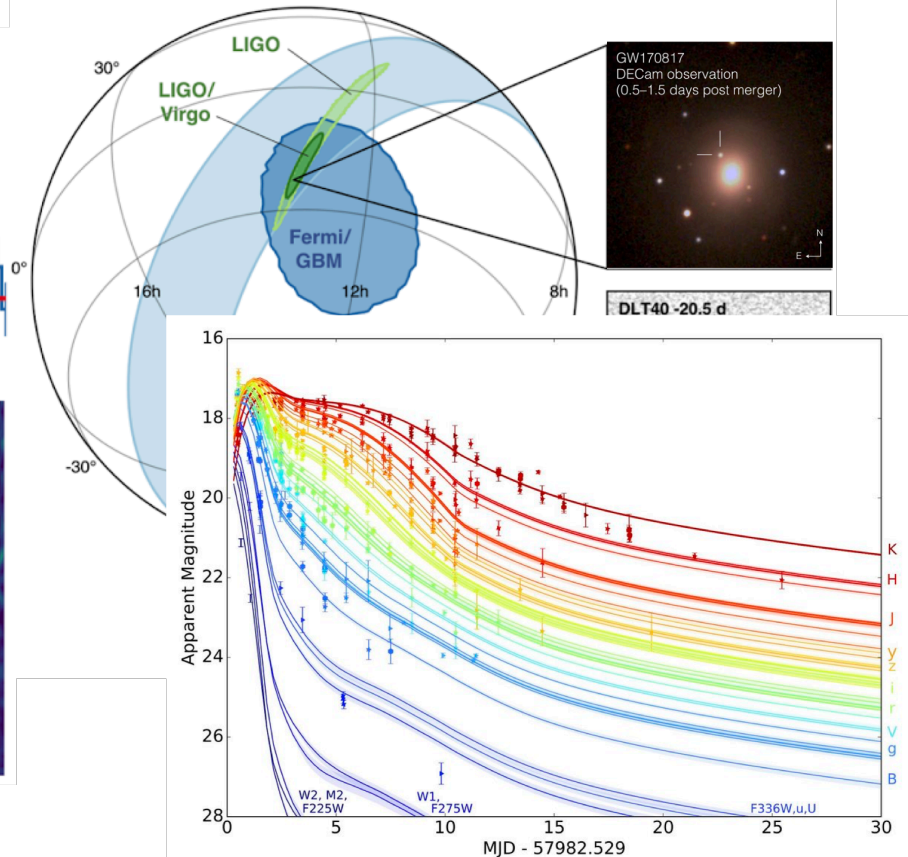
Demonstrated the ability of mergers to advance Nuclear Physics

Lightcurve from *Fermi*/GBM (50 – 300 keV)



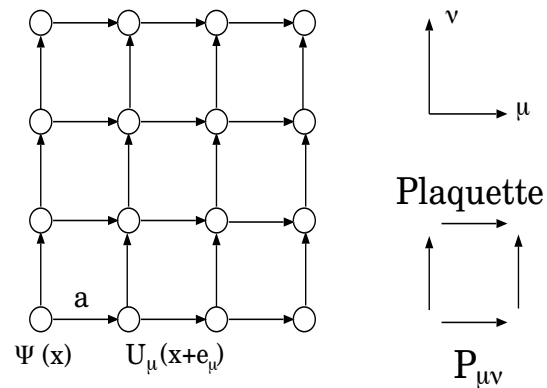
LIGO/Virgo PRL (2017)

P.S. Cowperthwaite et al., *Astrophys. J. Lett.* (2017)



Simulating strongly interacting matter

- ✧ Analytic solutions of **QCD** are not possible in the non-perturbative regime
- ✧ **Numerical** approach to solve QCD: discretization of 4D space-time
- ✧ Uncertainties:
 - Statistical: finite sample
 - Systematic: finite box size, unphysical quark masses
- Results from different groups, adopting different discretizations, must converge to consistent results



Fundamental fields



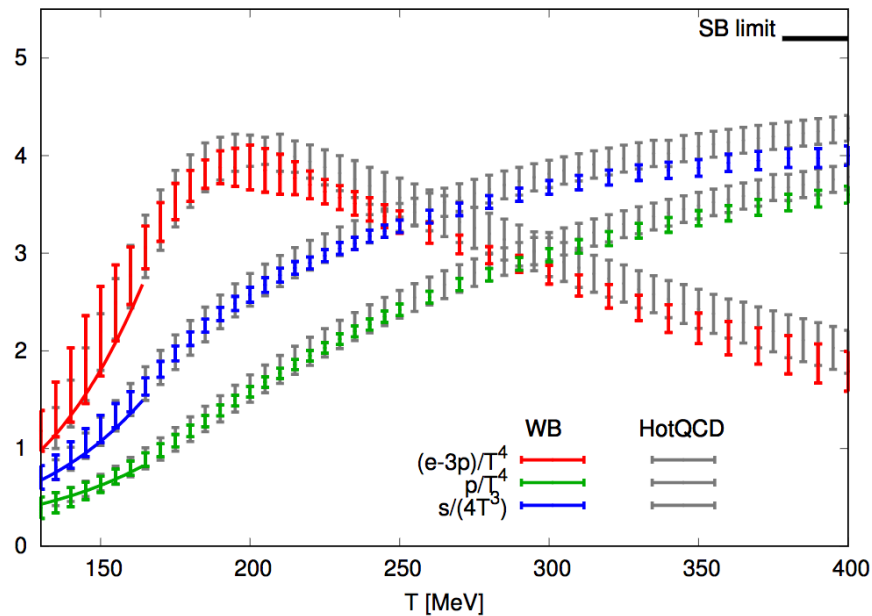
QCD Equation of State at finite density

TAYLOR EXPANSION

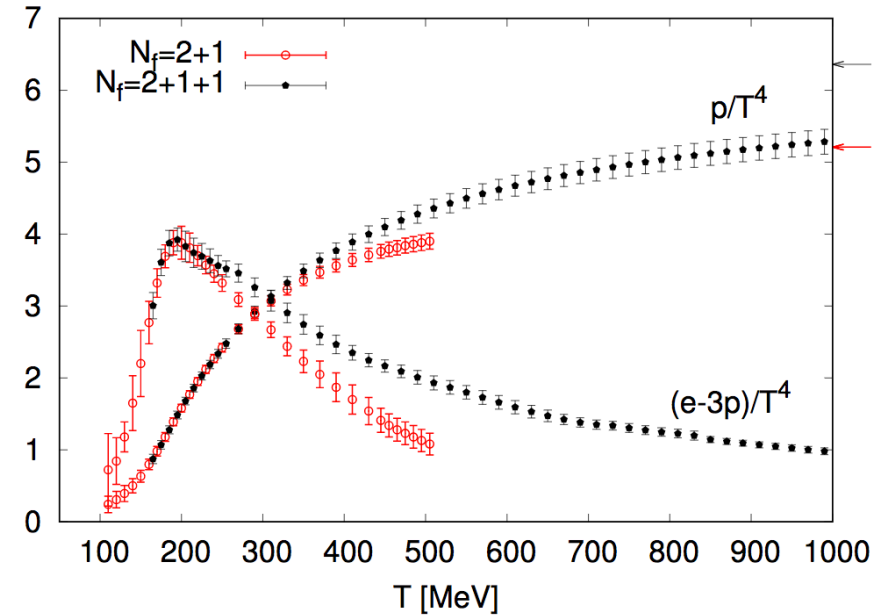
NEW EXPANSION SCHEME

QCD EoS at $\mu_B=0$

WB: PLB (2014); HotQCD: PRD (2014)

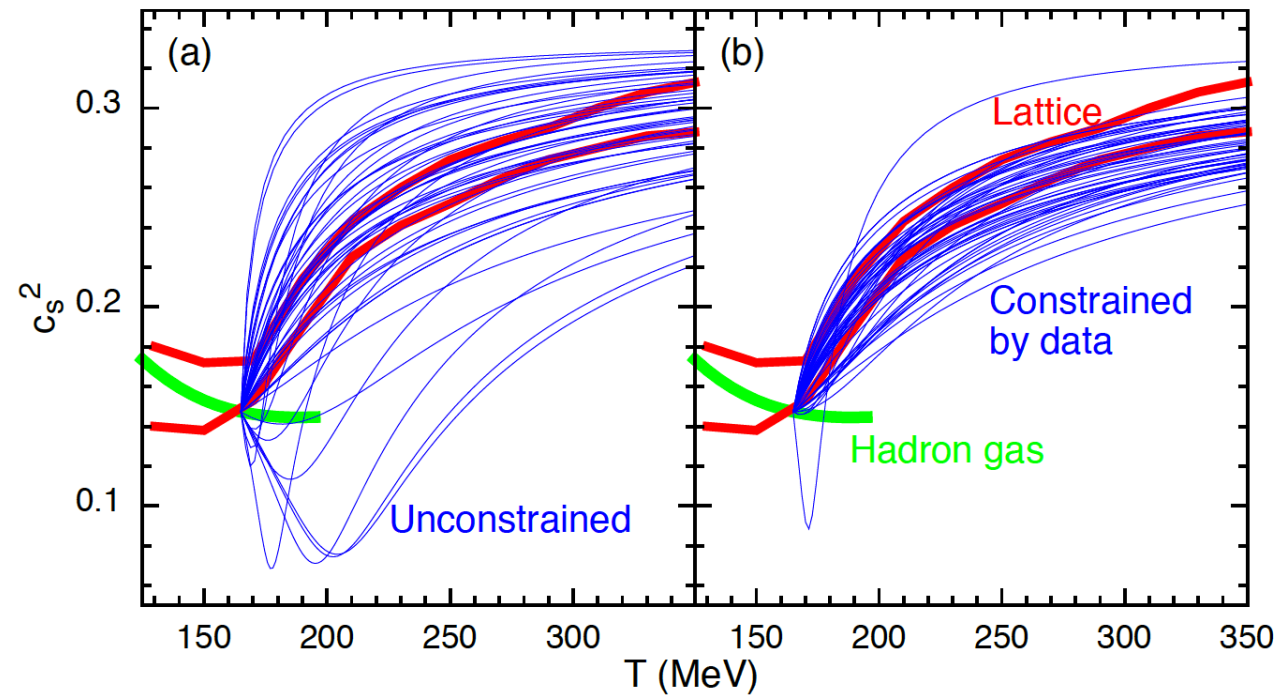


WB: Nature (2016)



- EoS for $N_f=2+1$ known in the continuum limit since 2013
- Good agreement with the HRG model at low temperature
- Charm quark relevant degree of freedom already at $T \sim 250$ MeV

Constraints on the EoS from the experiments



S. Pratt et al., PRL (2015)

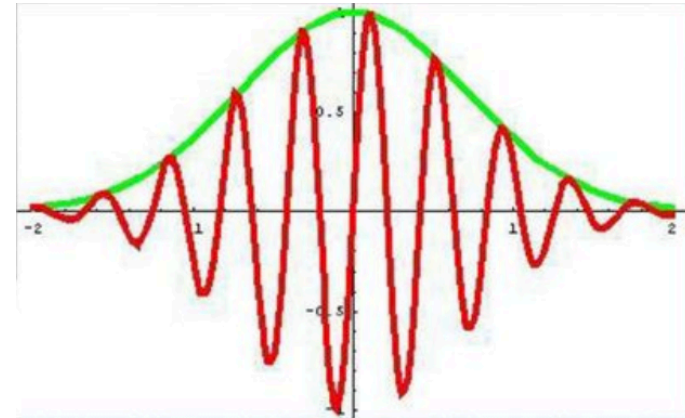
- Comparison of data from RHIC and LHC to theoretical models through Bayesian analysis
- The posterior distribution of EoS is consistent with the lattice QCD one

Fermionic sign problem

- The QCD path integral is computed by Monte Carlo algorithms which sample field configurations with a weight proportional to the exponential of the action

$$Z(\mu_B, T) = \text{Tr} \left(e^{-\frac{H_{\text{QCD}} - \mu_B N_B}{T}} \right) = \int \mathcal{D}U e^{-S_G[U]} \det M[U, \mu_B]$$

- $\det M[\mu_B]$ complex \rightarrow Monte Carlo simulations are not feasible
- We can rely on a few approximate methods, viable for small μ_B/T :
 - Taylor expansion of physical quantities around $\mu_B=0$
Bielefeld-Swansea collaboration 2002; R. Gavai, S. Gupta 2003
 - Simulations at imaginary chemical potentials
Alford, Kapustin, Wilczek, 1999; de Forcrand, Philipsen, 2002; D'Elia, Lombardo 2003



Taylor expansion of EoS

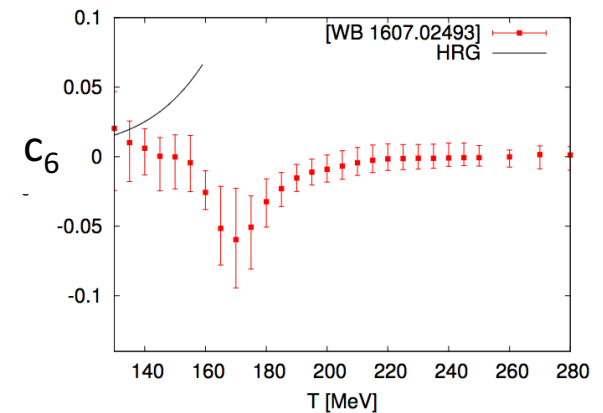
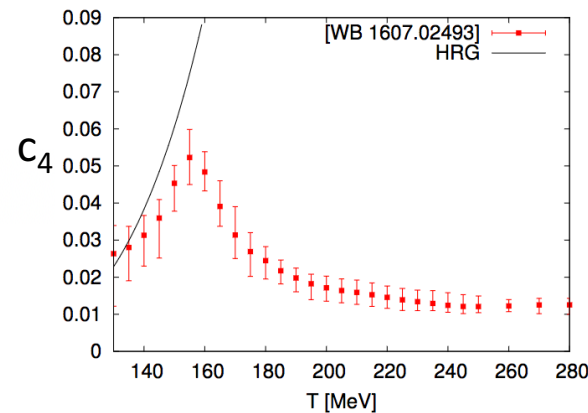
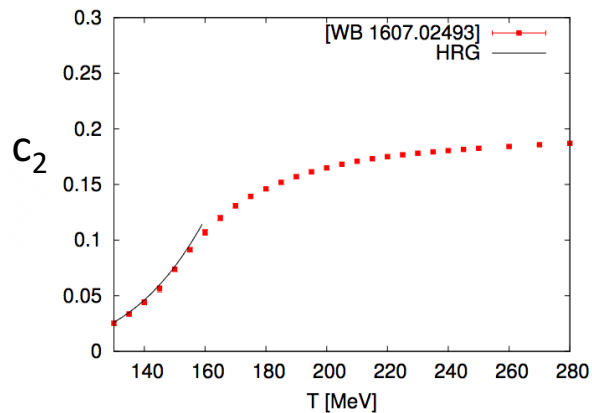
- Taylor expansion of the pressure:

$$\frac{p(T, \mu_B)}{T^4} = \frac{p(T, 0)}{T^4} + \sum_{n=1}^{\infty} \frac{1}{(2n)!} \chi_{2n}^B \left. \frac{d^{2n}(p/T^4)}{d(\frac{\mu_B}{T})^{2n}} \right|_{\mu_B=0} \left(\frac{\mu_B}{T}\right)^{2n} = \sum_{n=0}^{\infty} c_{2n}(T) \left(\frac{\mu_B}{T}\right)^{2n}$$

Simulations at imaginary μ_B :

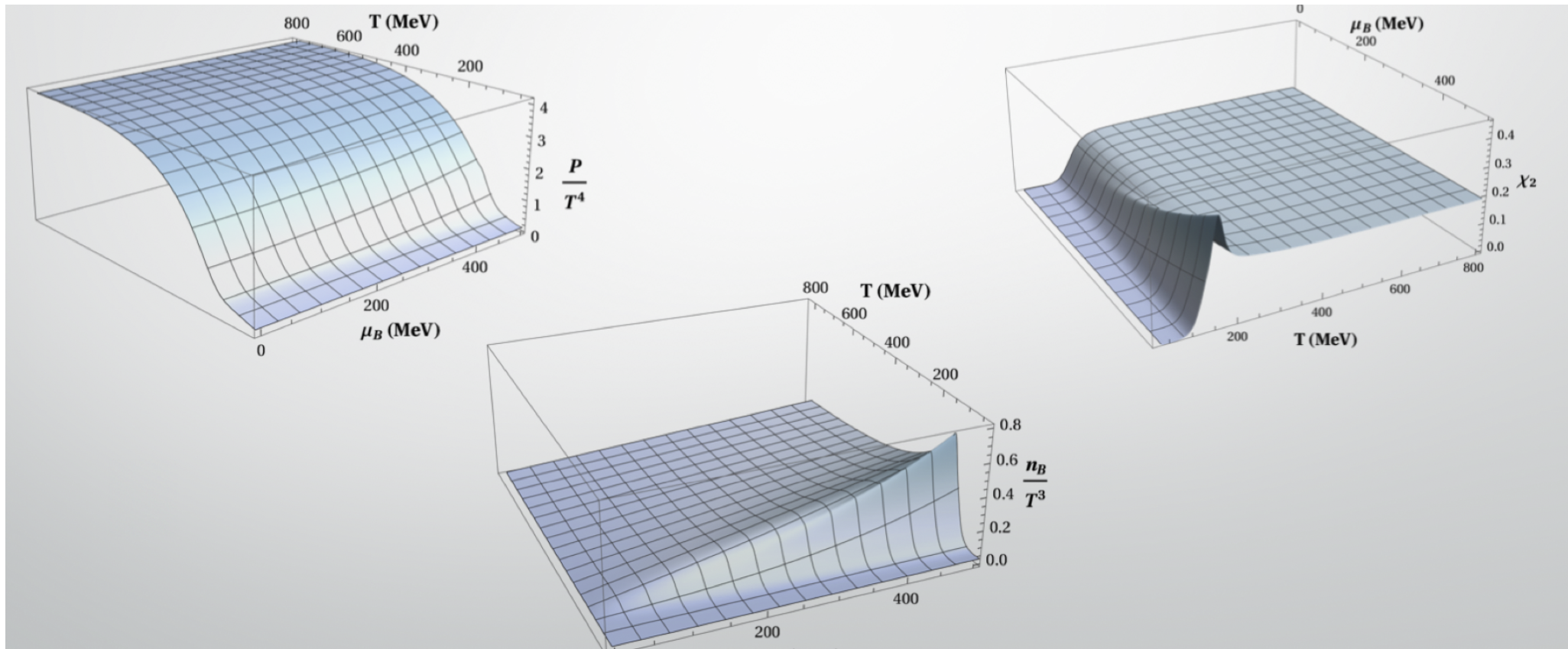
Continuum, $O(10^4)$ configurations, errors include systematics

WB: NPA (2017)



See also: HotQCD, PRD (2017), PRD (2022)

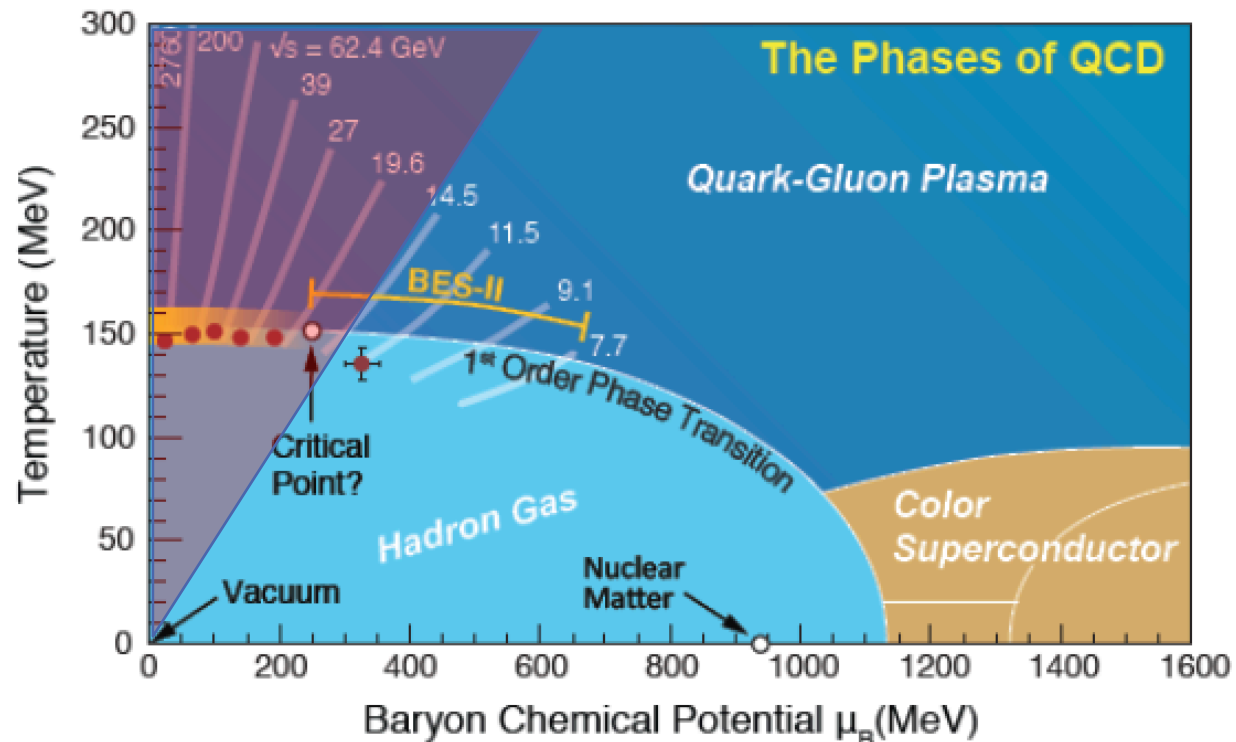
Taylor expansion of EoS



Range of validity of equation of state

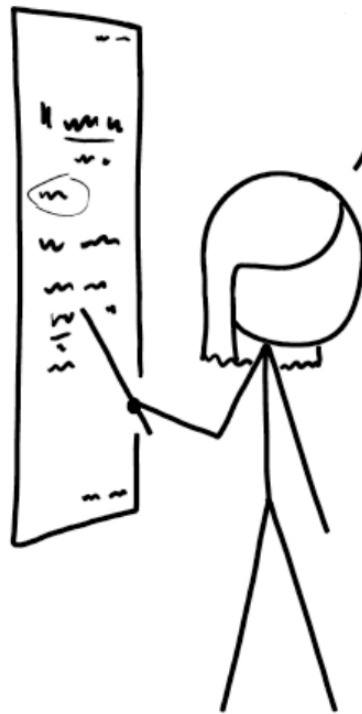
- We now have the equation of state for $\mu_B/T \leq 2$ or in terms of the RHIC energy scan:

$$\sqrt{s} = 200, 62.4, 39, 27, 19.6, 14.5 \text{ GeV}$$



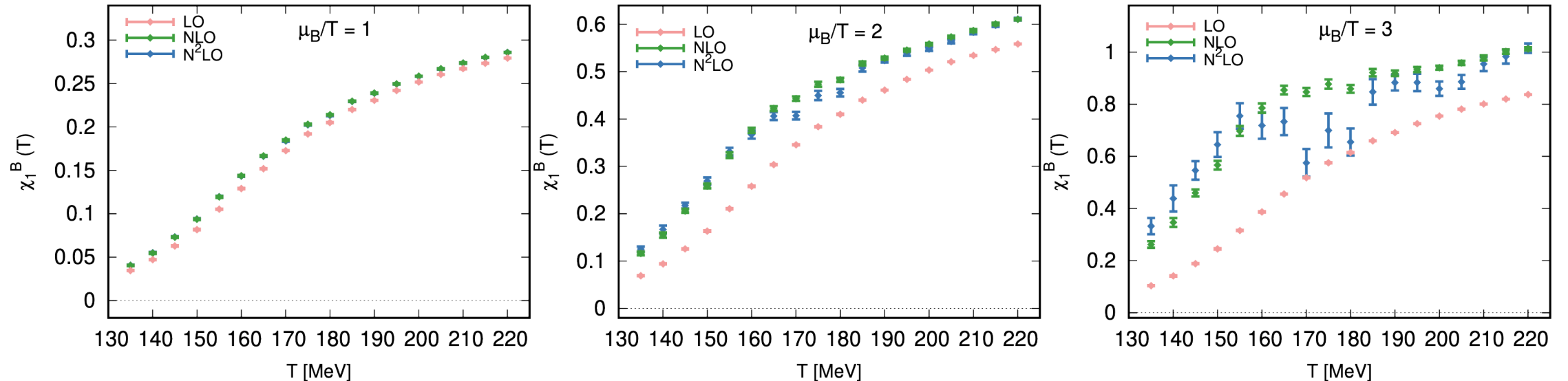
AT THIS POINT, YOU'RE PROBABLY
THINKING, "I LOVE THIS EQUATION
AND WISH IT WOULD NEVER END!"

WELL, GOOD NEWS!



TAYLOR SERIES EXPANSION IS THE WORST.

Problems with Taylor series

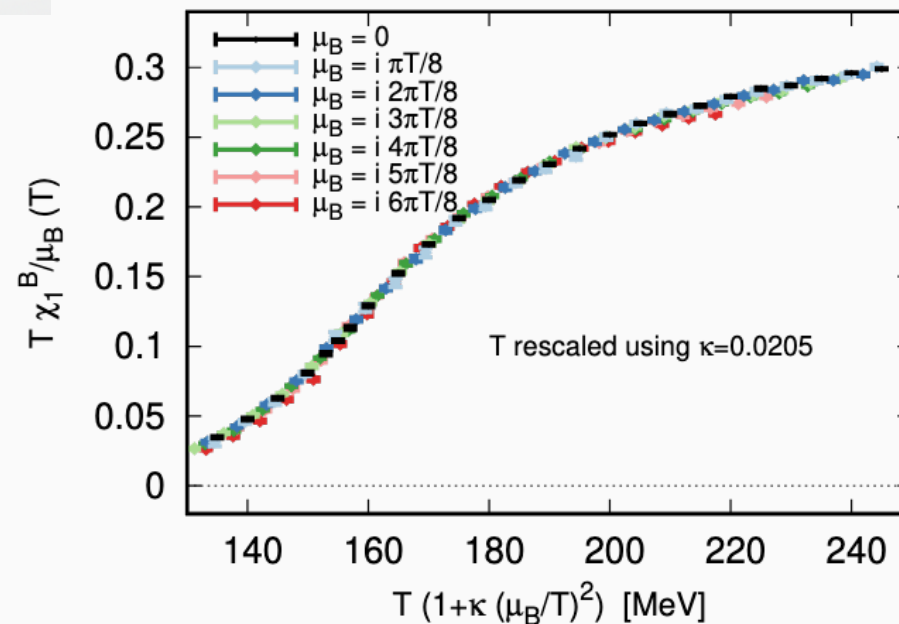
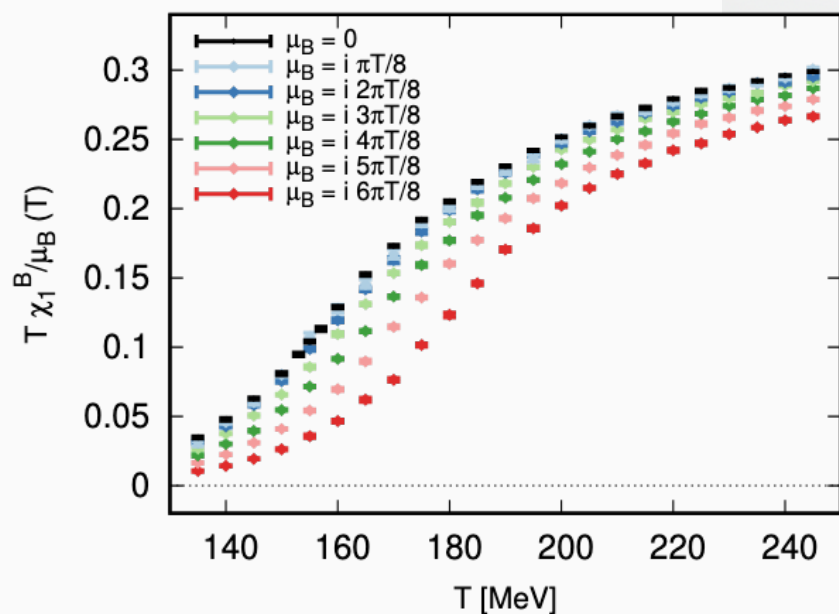


- ❑ Poor convergence of Taylor series: need to sum many terms to reach high μ_B
- ❑ Oscillatory/non-monotonic behavior in some observables at high μ_B
 - Unphysical, due to truncation of Taylor series

An alternative approach

From simulations at imaginary μ_B we observe that $\chi_1^B(T, \hat{\mu}_B)$ at (imaginary) $\hat{\mu}_B$ appears to be differing from $\chi_2^B(T, 0)$ mostly by a rescaling of T :

$$\frac{\chi_1^B(T, \hat{\mu}_B)}{\hat{\mu}_B} = \chi_2^B(T', 0), \quad T' = T (1 + \kappa \hat{\mu}_B^2)$$



Formulation

- We have observed the $\hat{\mu}_B$ -dependence seems to amount to a simple T - rescaling
- A simplistic scenario with a single T - independent parameter κ does not provide a systematic treatment which can serve as an alternative expansion scheme
- We allow for more than $\mathcal{O}(\hat{\mu}^2)$ expansion of T' and let the coefficients be T -dependent:

$$\frac{\chi_1^B(T, \hat{\mu}_B)}{\hat{\mu}_B} = \chi_2^B(T', 0) , \quad T' = T \left(1 + \kappa_2(T) \hat{\mu}_B^2 + \kappa_4(T) \hat{\mu}_B^4 + \mathcal{O}(\hat{\mu}_B^6) \right)$$

- **Important:** we are simply re-organizing the Taylor expansion via an expansion in the shift

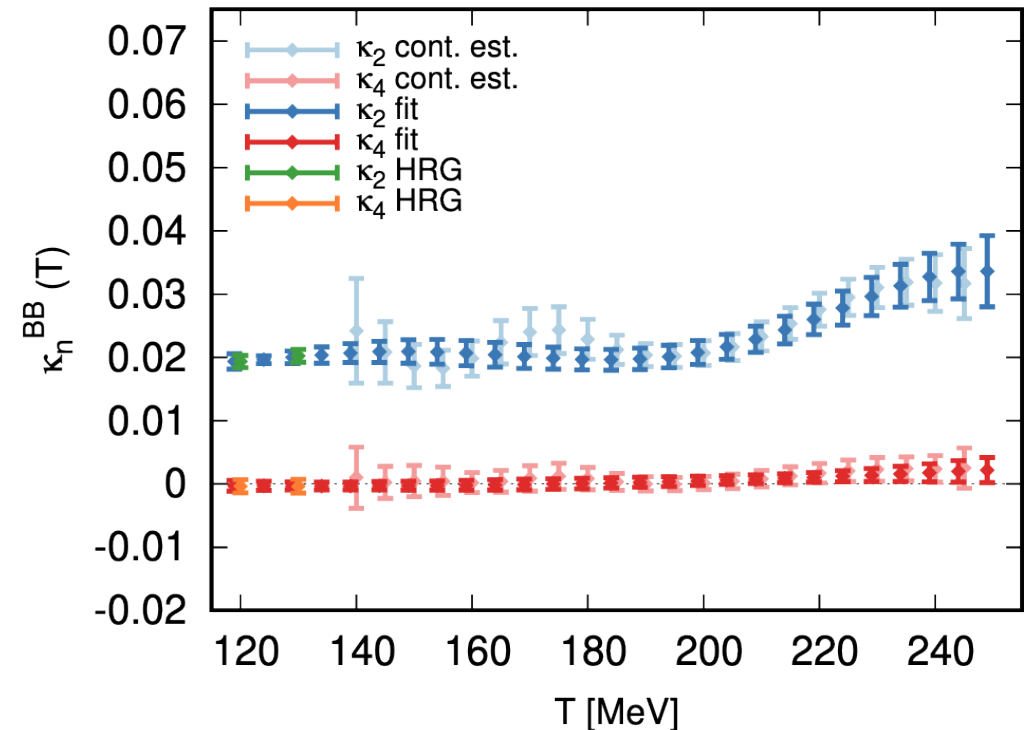
$$\Delta T = T - T' = \left(\kappa_2(T) \hat{\mu}_B^2 + \kappa_4(T) \hat{\mu}_B^4 + \mathcal{O}(\hat{\mu}_B^6) \right)$$

- Comparing the (Taylor) expansion in $\hat{\mu}_B$ and our expansion in ΔT order by order, we can relate $\chi_n^B(T)$ and $\kappa_n(T)$

Results for the coefficients

Our initial guess was not far-off:

- Fairly constant $\kappa_2(T)$ over a large T -range
- Clear separation in magnitude between $\kappa_2(T)$ and $\kappa_4(T)$ hints at better convergence
- Agreement with the HRG model results at low temperatures
- Polynomial fits of $\kappa_2(T)$ and $\kappa_4(T)$ before use in thermodynamics (good fit qualities)



NOTE: polynomial fits take into account both statistical and systematic correlations.

Thermodynamics at finite μ_B

Thermodynamic quantities at finite (real) μ_B can be reconstructed from the same ansatz:

$$\frac{n_B(T, \hat{\mu}_B)}{T^3} = \hat{\mu}_B \chi_2^B(T', 0)$$

with $T' = T(1 + \kappa_2^{BB}(T) \hat{\mu}_B^2 + \kappa_4^{BB}(T) \hat{\mu}_B^4)$.

From the baryon density n_B one finds the pressure:

$$\frac{p(T, \hat{\mu}_B)}{T^4} = \frac{p(T, 0)}{T^4} + \int_0^{\hat{\mu}_B} d\hat{\mu}'_B \frac{n_B(T, \hat{\mu}'_B)}{T^3}$$

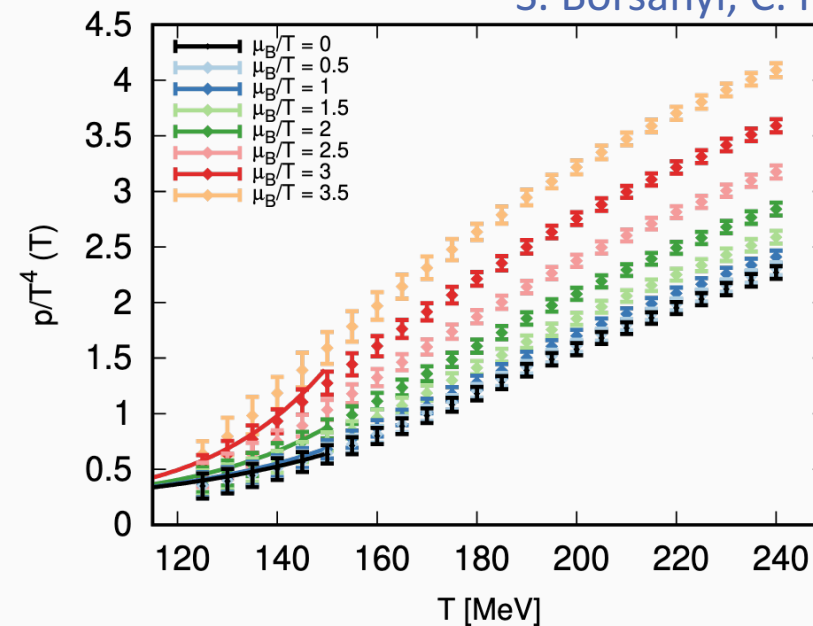
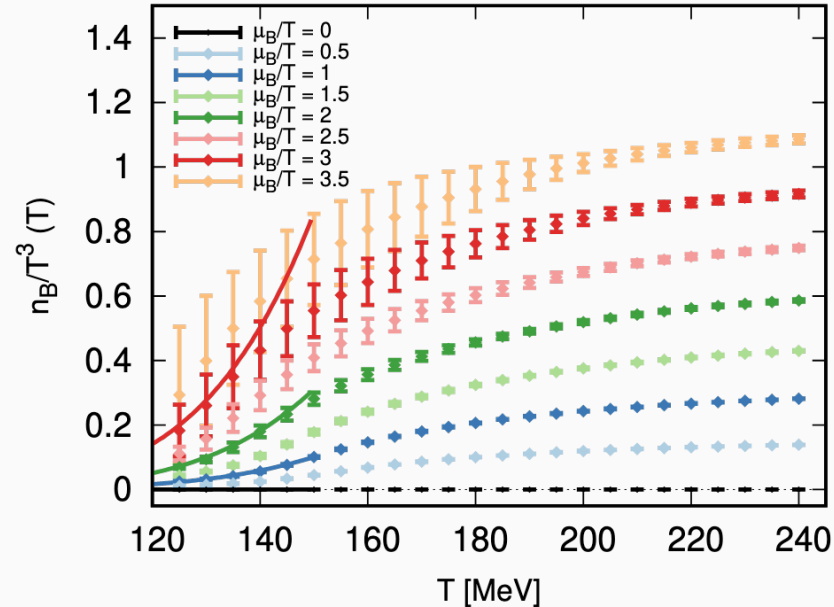
then the entropy, energy density:

$$\frac{s(T, \hat{\mu}_B)}{T^4} = 4 \frac{p(T, \hat{\mu}_B)}{T^4} + T \left. \frac{\partial p(T, \hat{\mu}_B)}{\partial T} \right|_{\hat{\mu}_B} - \hat{\mu}_B \frac{n_B(T, \hat{\mu}_B)}{T^3}$$
$$\frac{\epsilon(T, \hat{\mu}_B)}{T^4} = \frac{s(T, \hat{\mu}_B)}{T^3} - \frac{p(T, \hat{\mu}_B)}{T^4} + \hat{\mu}_B \frac{n_B(T, \hat{\mu}_B)}{T^3}$$

Thermodynamics at finite μ_B : results

- We reconstruct thermodynamic quantities up to $\hat{\mu}_B \simeq 3.5$ with uncertainties well under control
- Agreement with HRG model calculations at small temperatures
- No pathological (non-monotonic) behavior is present

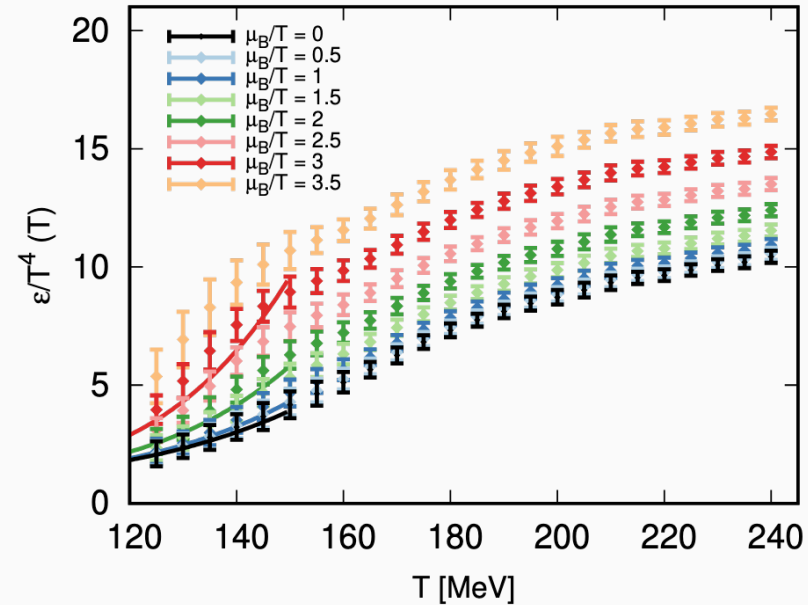
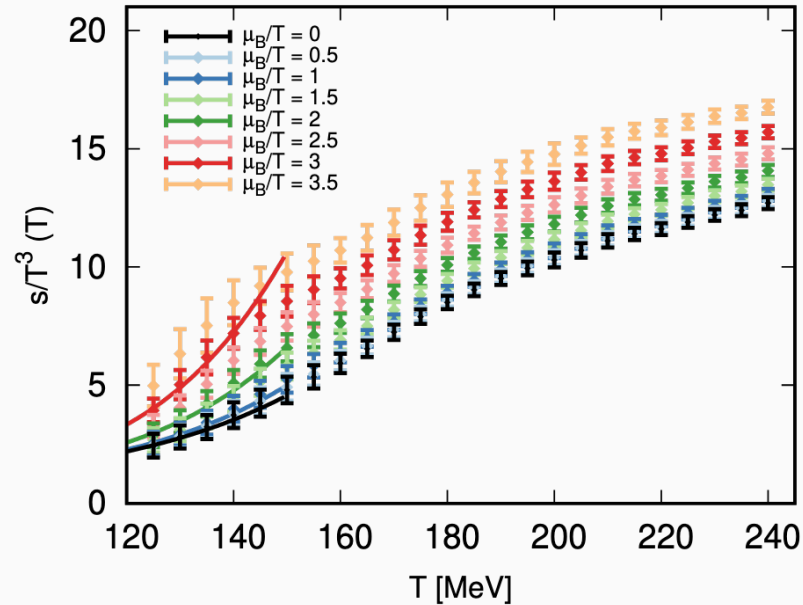
S. Borsanyi, C. R. et al., PRL (2021)



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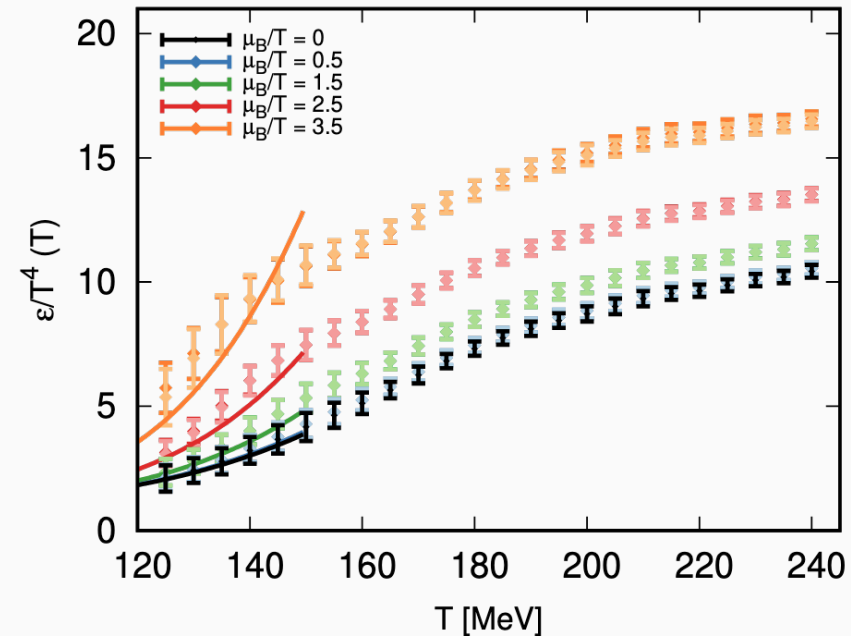
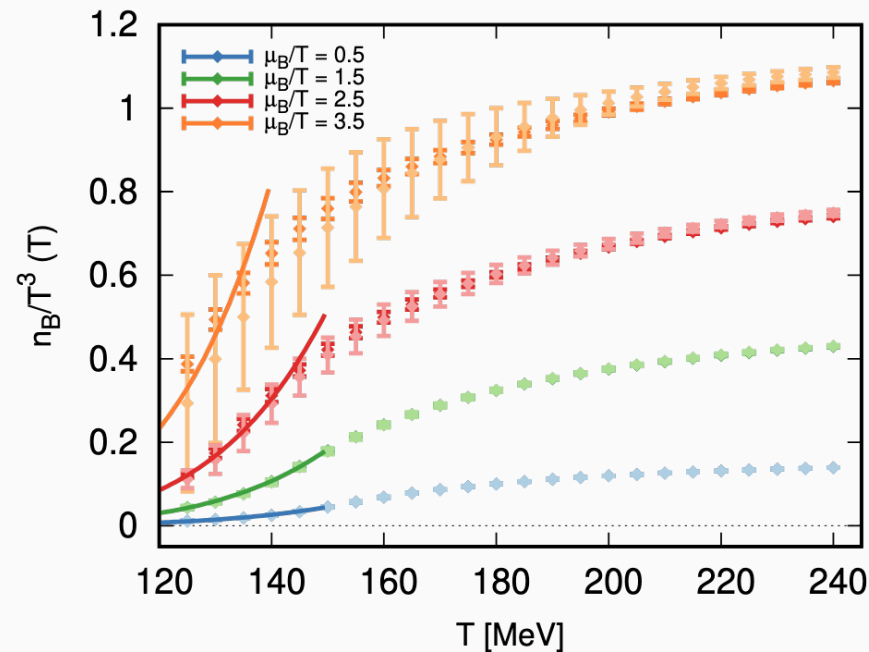
S. Borsanyi, C. R. et al., PRL (2021)



Convergence check

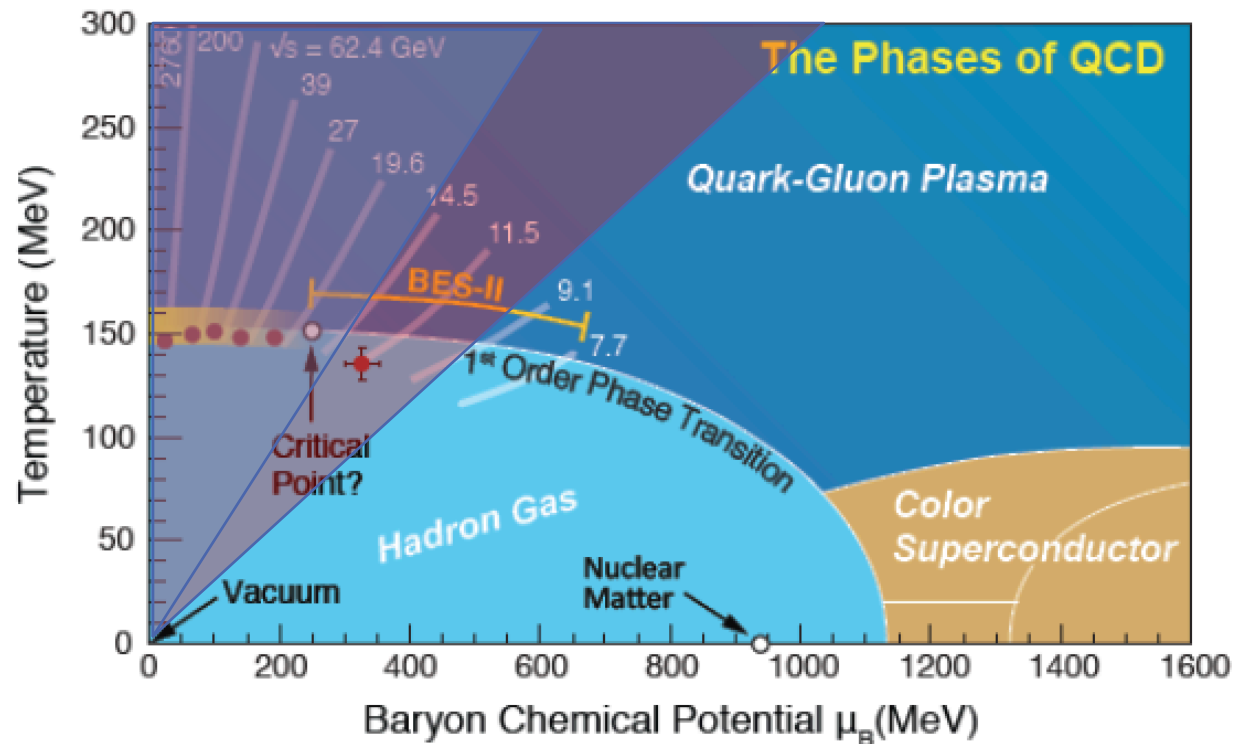
- We also check the results without the inclusion of $\kappa_4(T)$ (darker shades)
- Including $\kappa_4(T)$ only results in added error, but does not “move” the results

→ Good convergence



New range of validity of equation of state

- We now have the equation of state for $\mu_B/T \leq 3.5$



Other interesting results

➤ Recent development in reweighting schemes

[M. Giordano et al., JHEP 05, 088](#); [Borsanyi et al., PRD \(2022\)](#)

➤ Alternative ways to resum the Taylor series

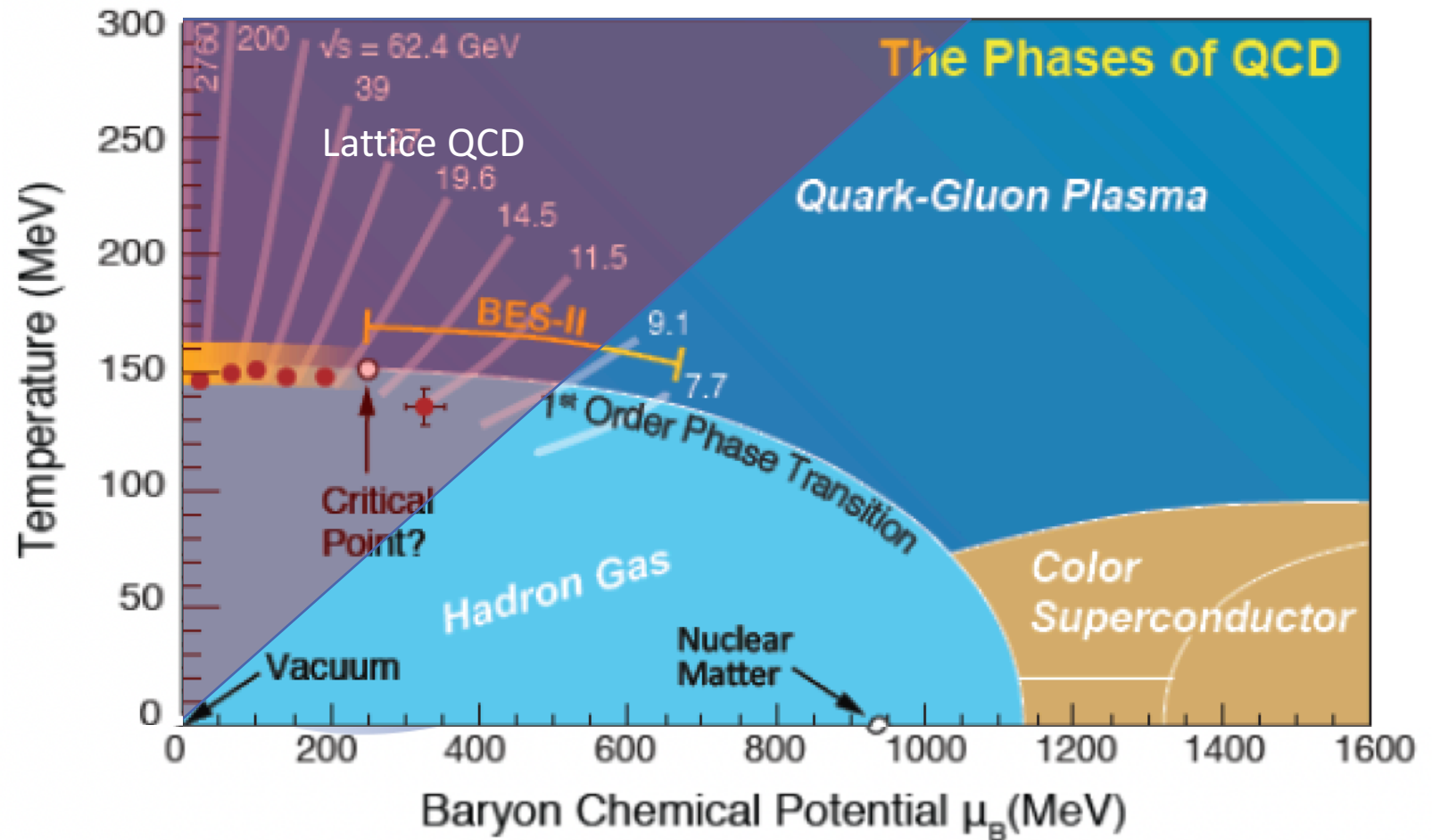
[S. Mondal et al., PRL \(2022\)](#); [S. Mukherjee et al., PRD \(2022\)](#); [S. Mitra et al., 2205.08517](#)

➤ Recent improvement on Taylor expansion at finite μ_B

[D. Bollweg et al., PRD \(2022\)](#)

What happens at large densities?

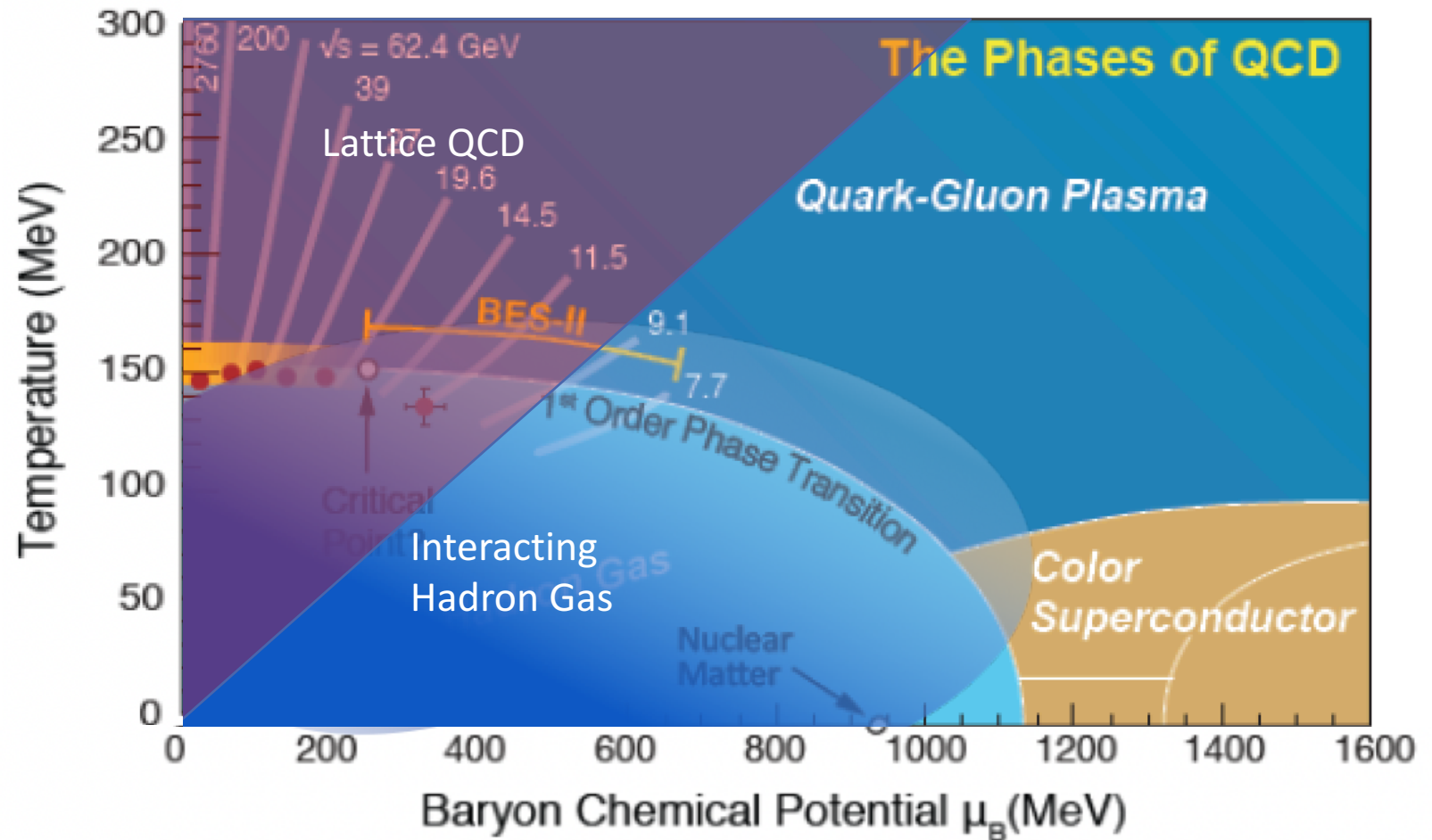
- We need to merge the lattice QCD equation of state with other effective theories
- Careful study of their respective range of validity
- Constrain the parameters to reproduce known limits
- Test different possibilities and validate/exclude them



Lattice QCD: WB: PRL (2021)

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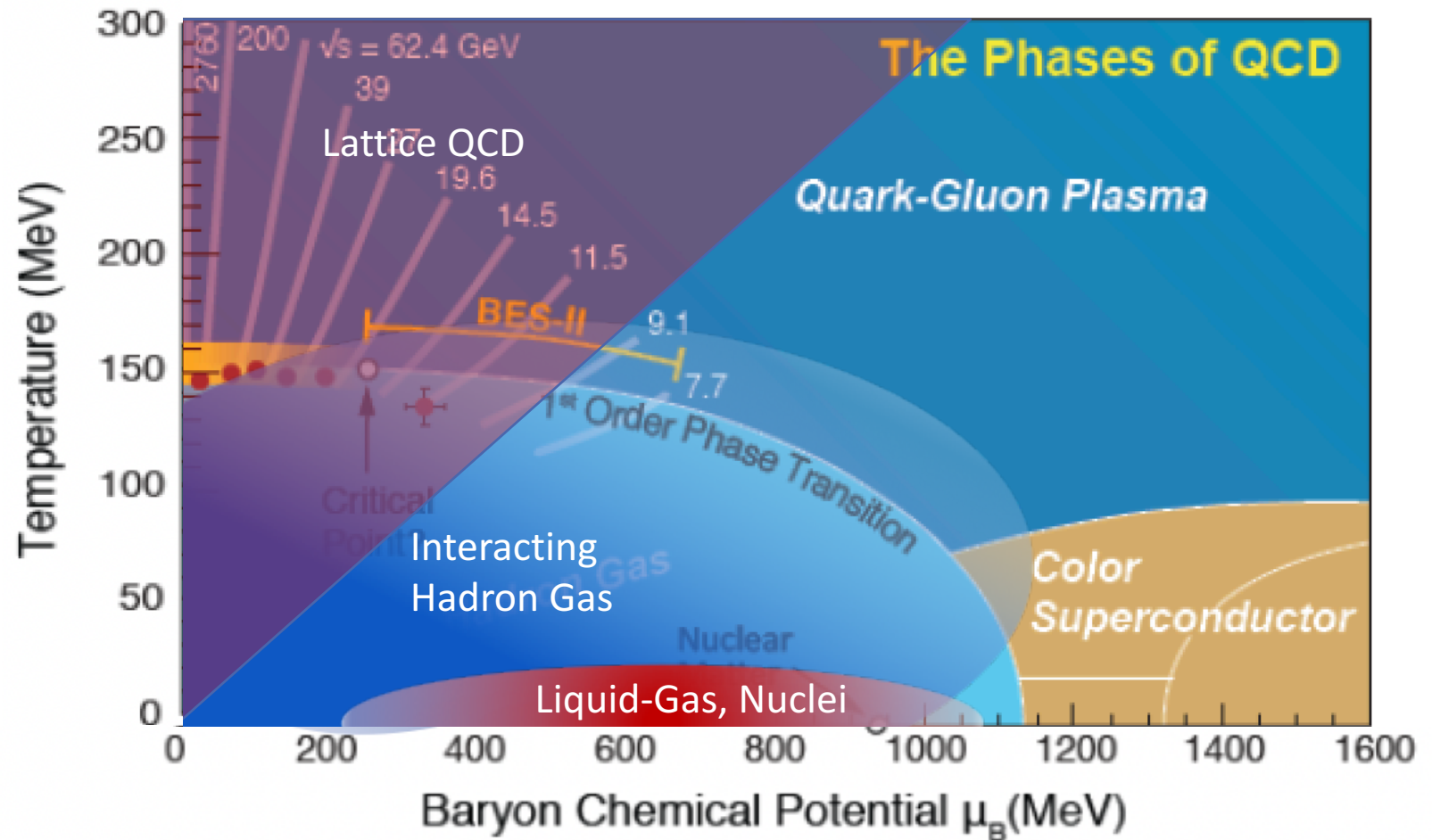


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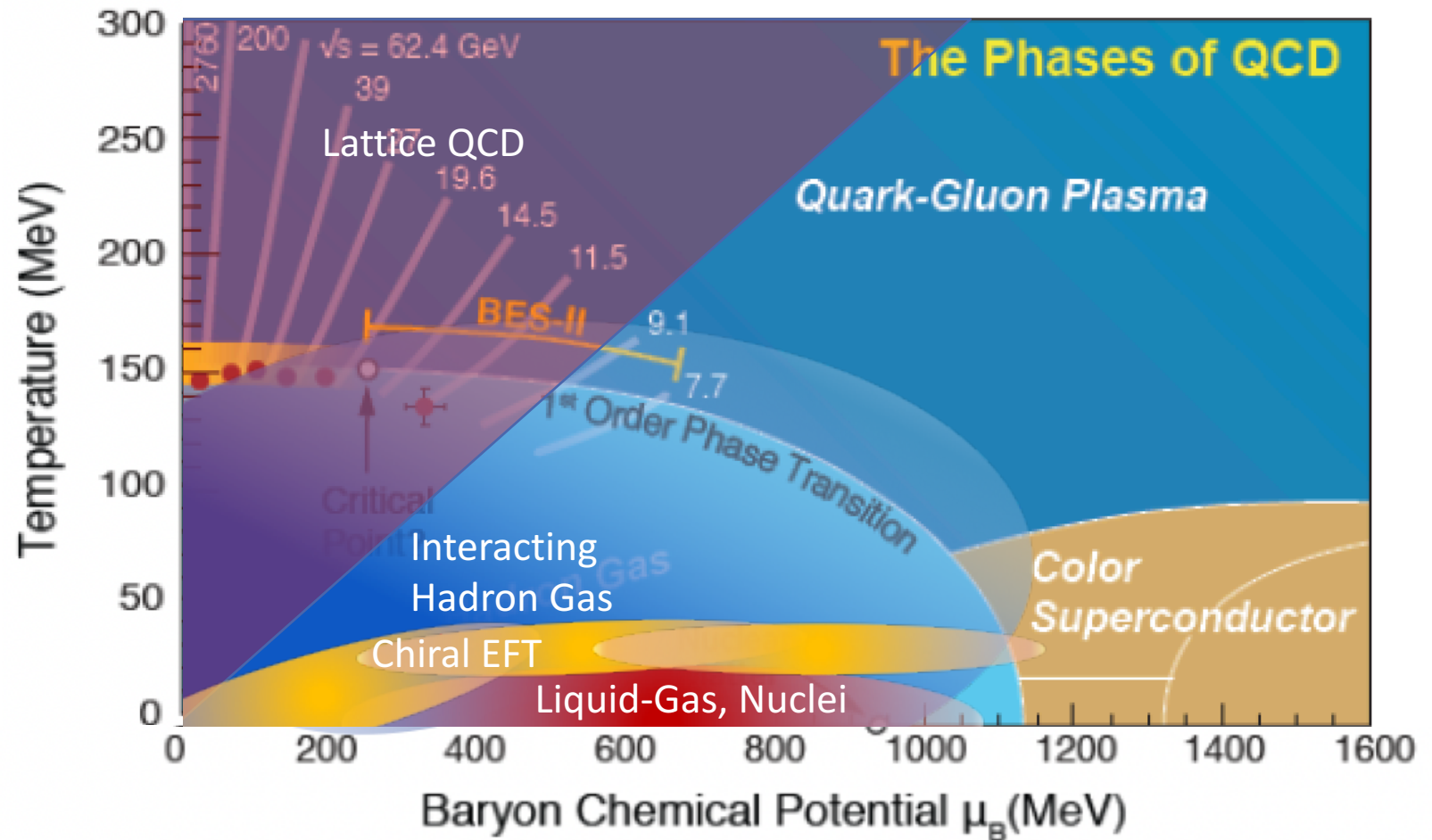
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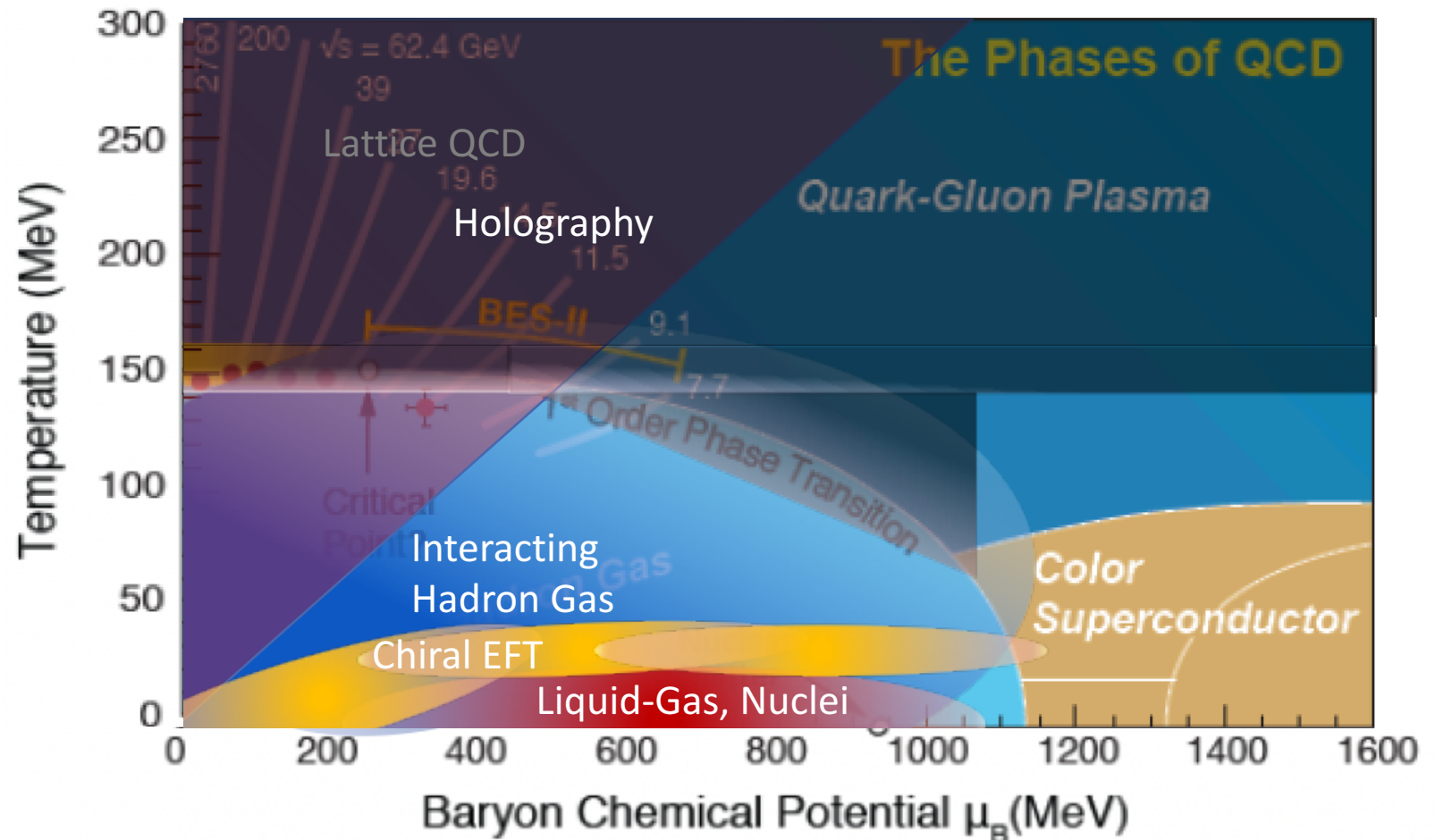
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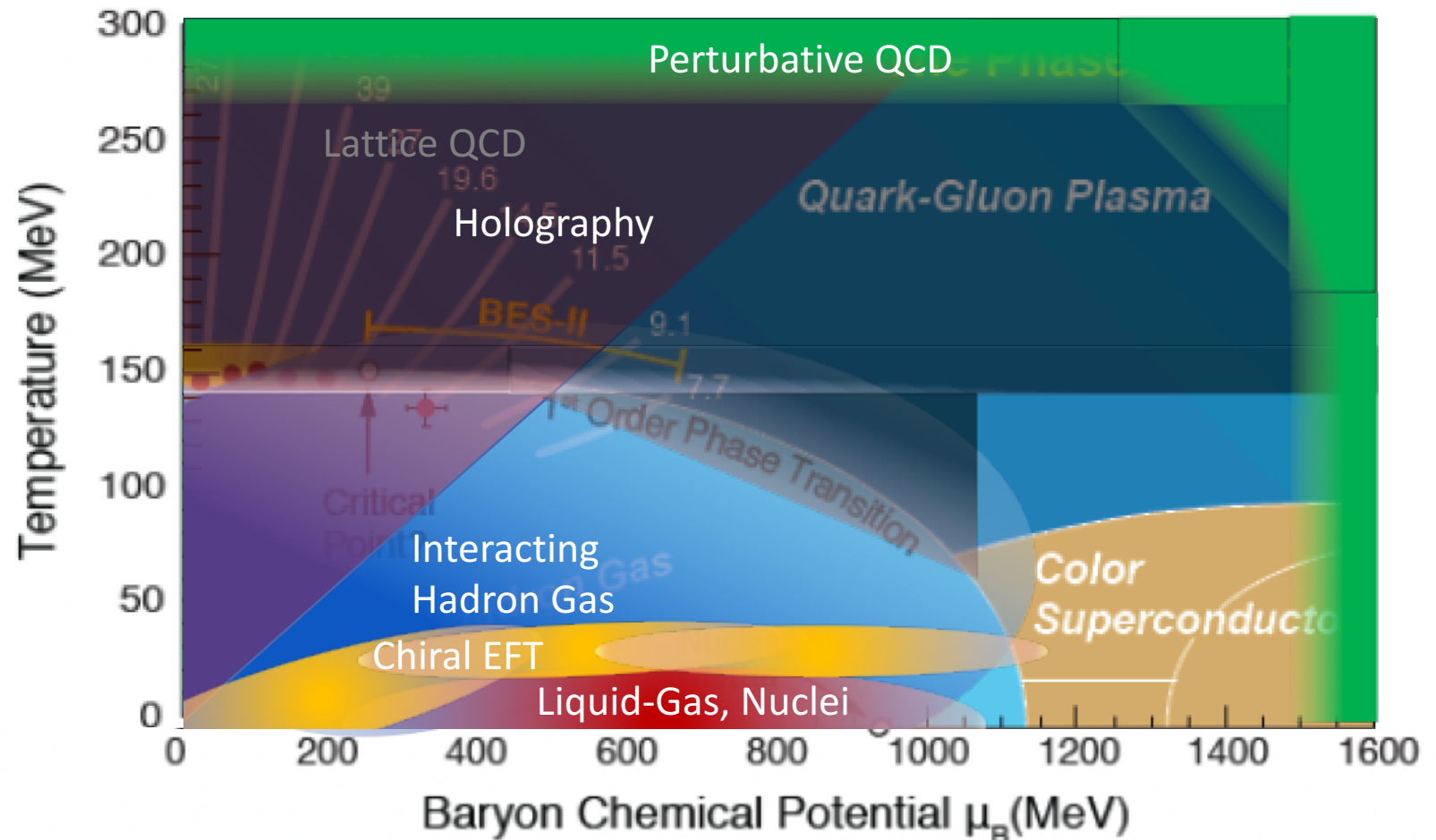
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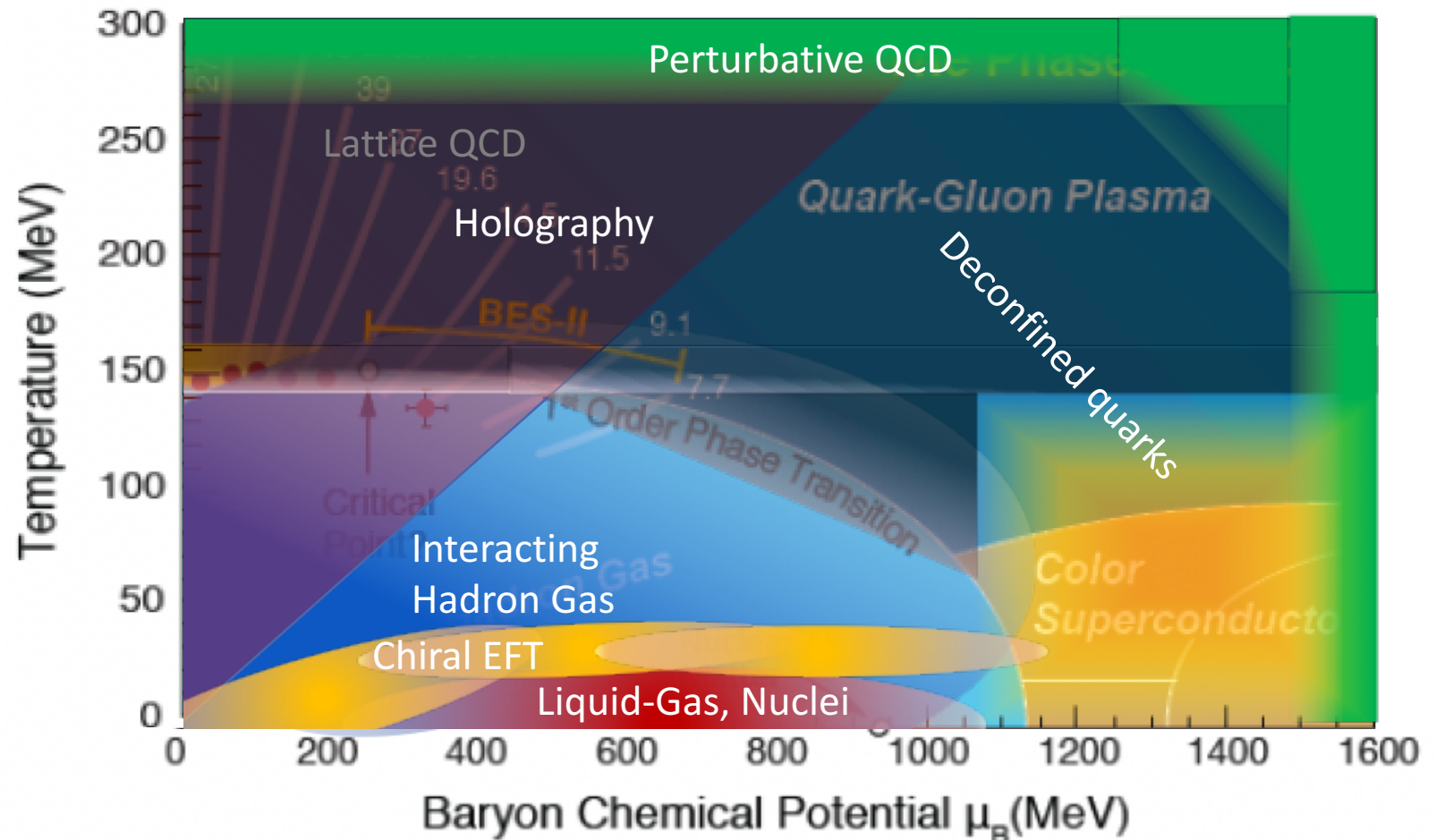
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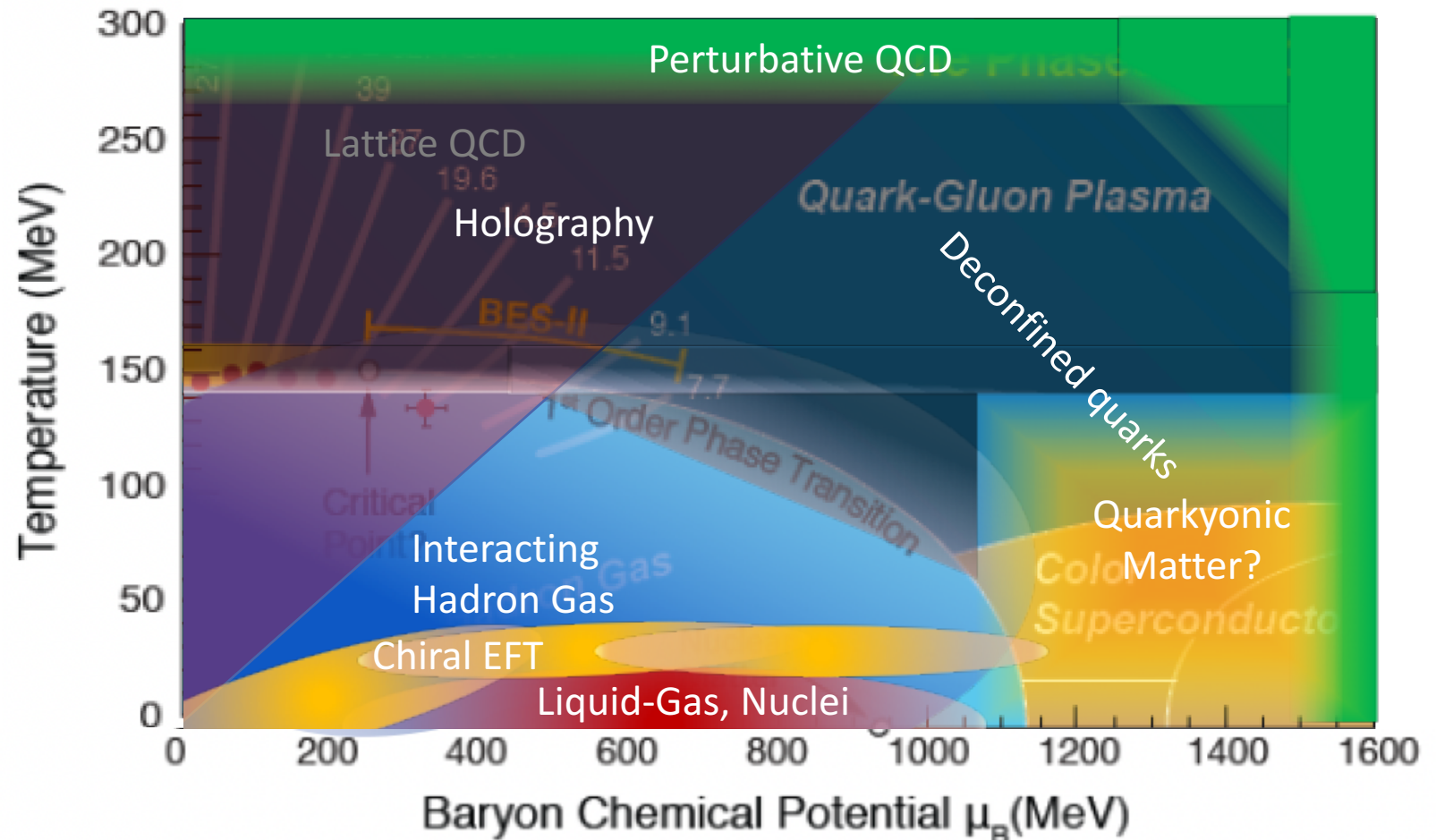
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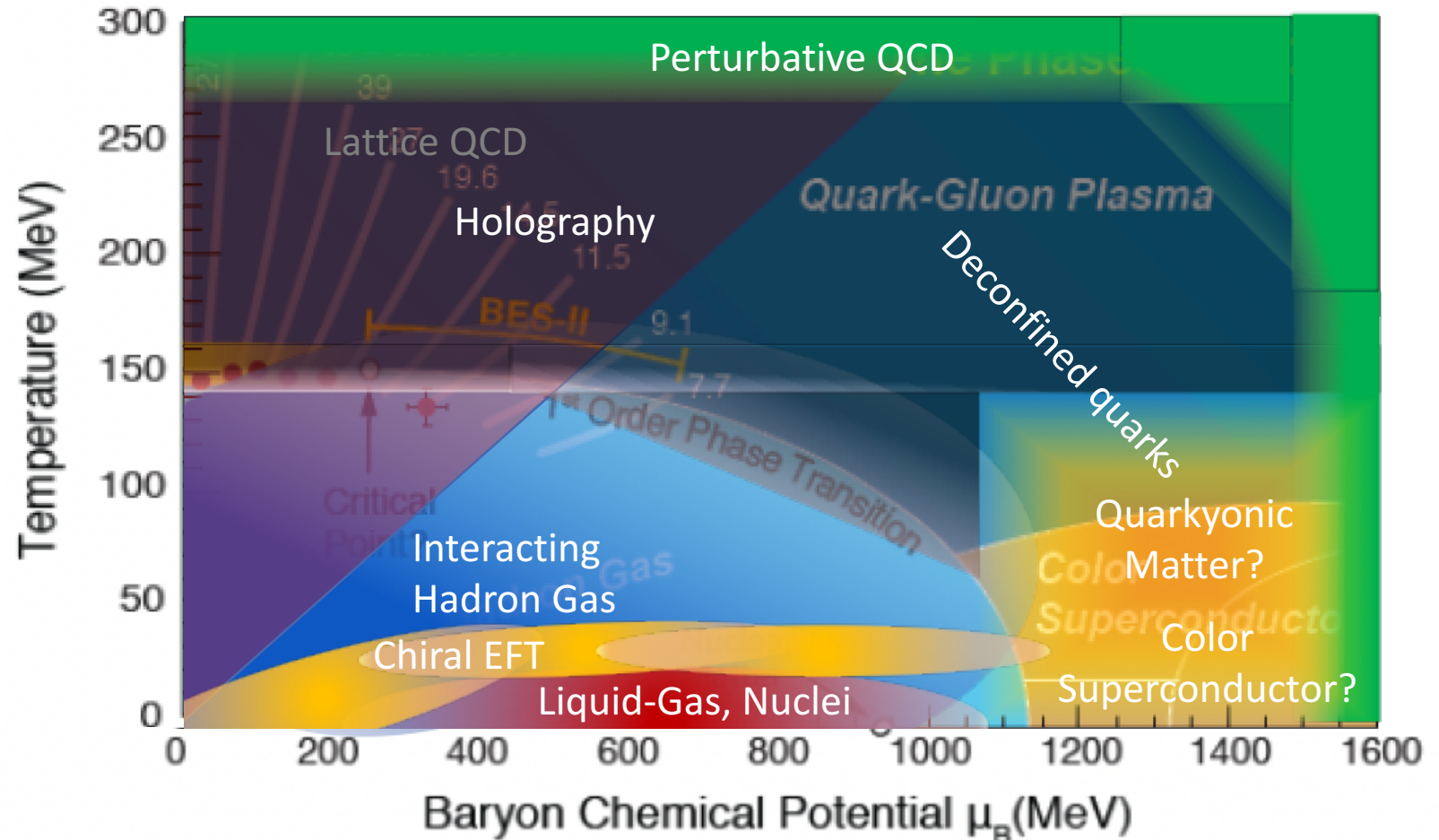
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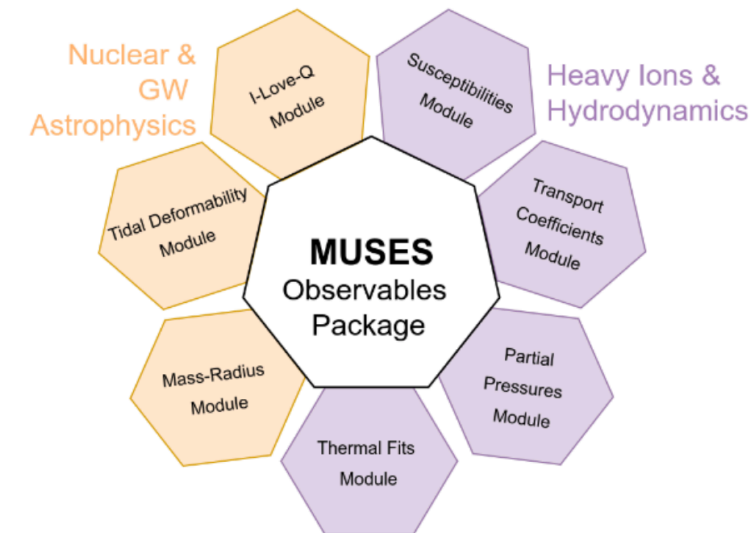
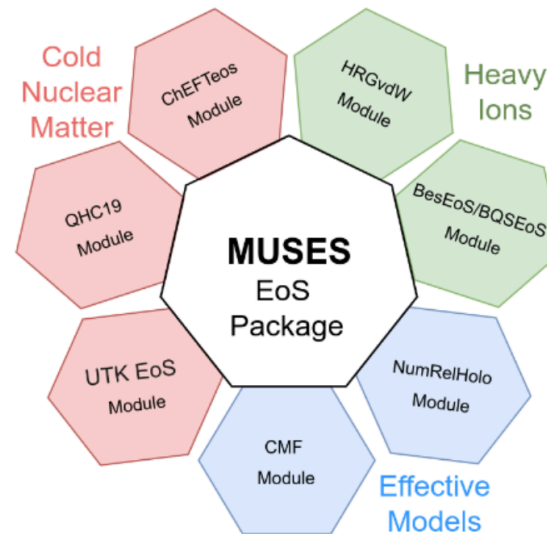
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CSC: Alford et al., PLB (1998); Rapp et al., PRL (1998).

Modular Unified Solver of the Equation of State collaboration

- Funded by NSF through CSSI program
- **Developers** and **Users** are working together to create a sustainable software to generate equations of state in the whole phase space
- **Modular**: Different models (“modules”) to describe the EoS in different regimes of phase space
- **Unified**: Modules smoothly integrated to (i) ensure maximal coverage of phase space, and (ii) respects constraints



A possible approach at large densities

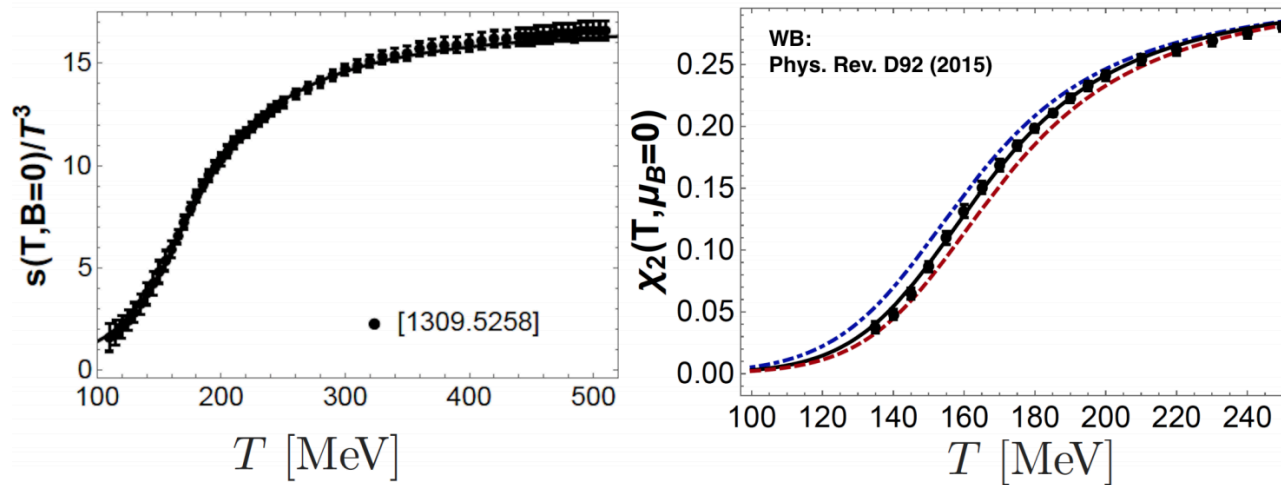
Use AdS/CFT correspondence

Fix the parameters to reproduce everything we know from the lattice

Calculate thermodynamics at finite density

Gravitational action

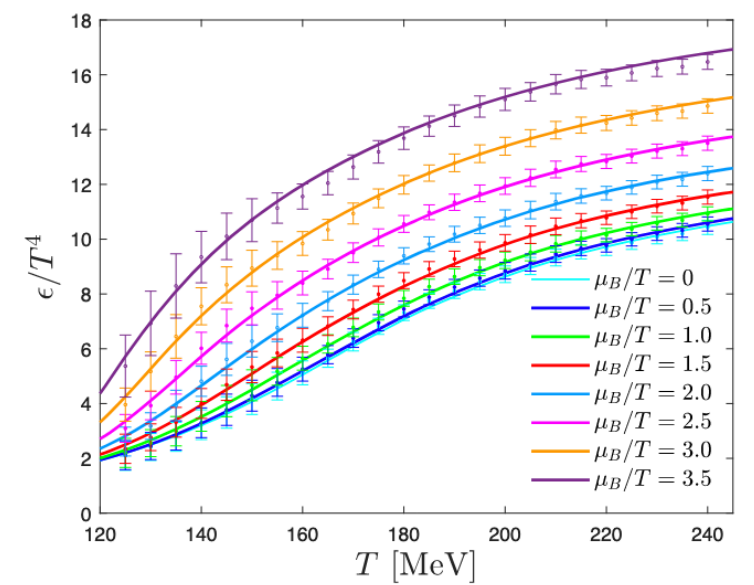
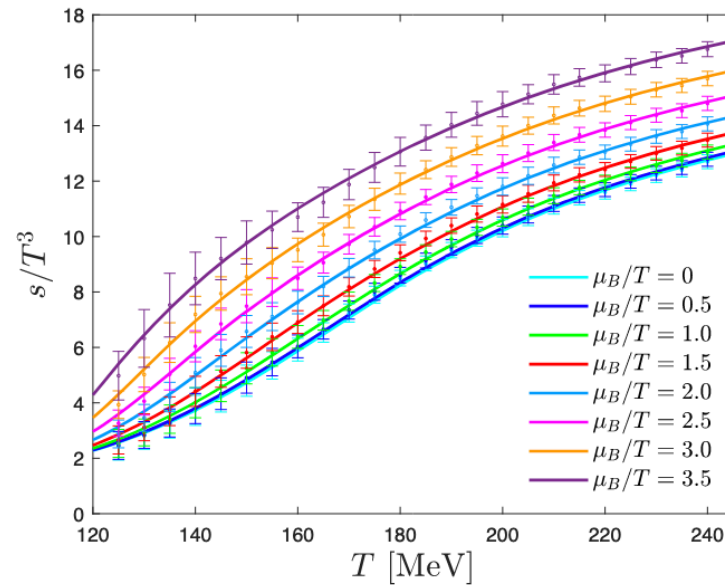
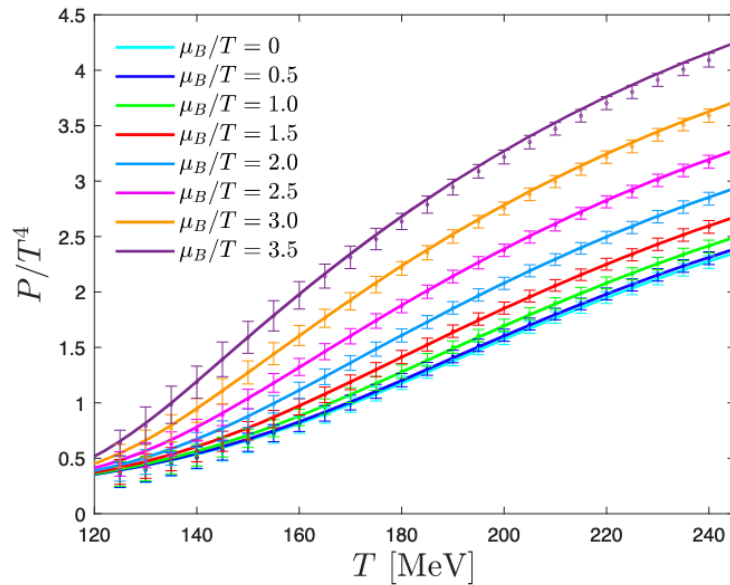
$$S = \frac{1}{2\kappa_5^2} \int_{M_5} d^5x \sqrt{-g} \left[R - \frac{(\partial_\mu \phi)^2}{2} - \underbrace{V(\phi)}_{\text{nonconformal}} - \underbrace{\frac{f(\phi)F_{\mu\nu}^2}{4}}_{\mu_B \neq 0} \right]$$



$$\chi_2(T, \mu_B) = \frac{\partial^2(P/T^4)}{\partial(\mu_B/T)^2} = \frac{\partial(\rho_B/T^3)}{\partial(\mu_B/T)}$$

O. DeWolfe et al., PRD (2011)
 R. Rougemont et al., JHEP (2016)
 R. Critelli et al., PRD (2017)

Comparison with lattice QCD thermodynamics

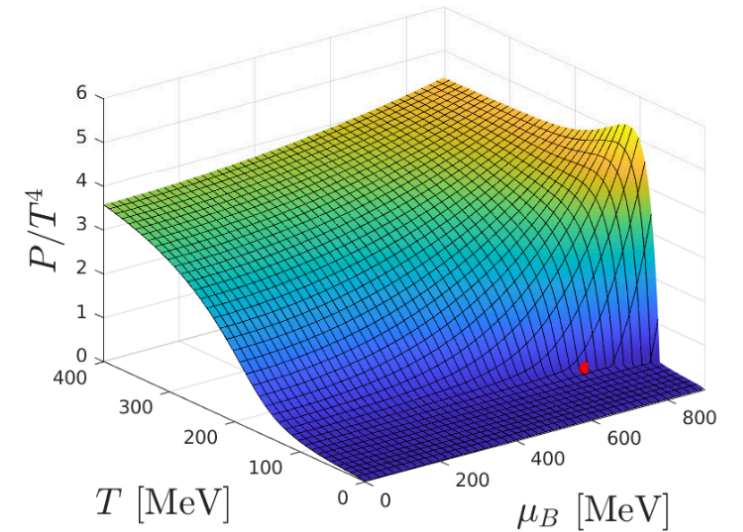
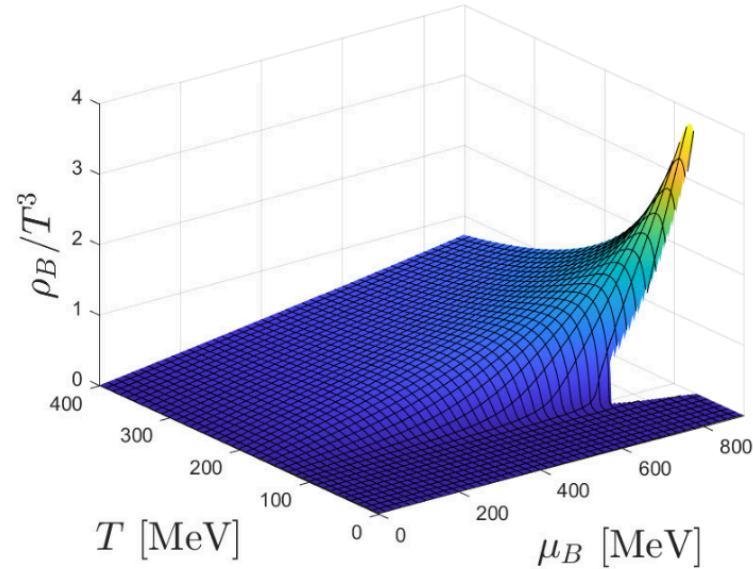
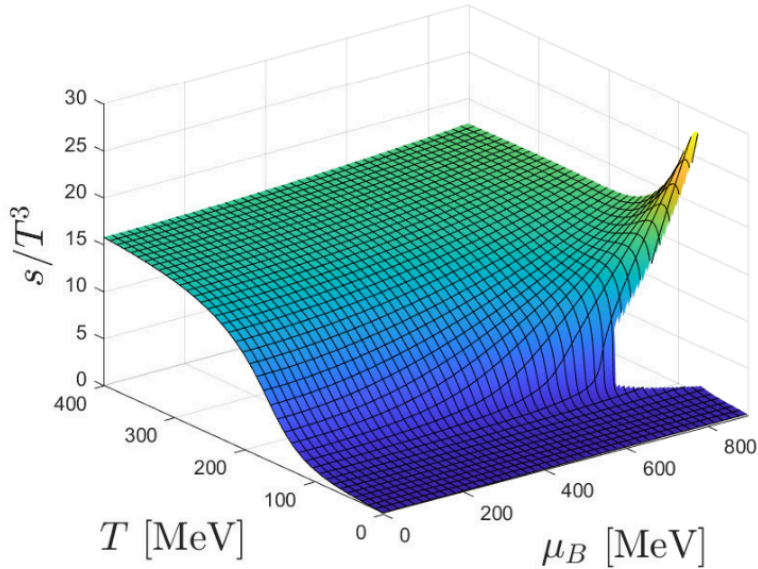


Good agreement for all thermodynamic quantities up to large μ_B

Makes us confident in our extrapolation to larger μ_B

J. Grefa et al., PRD (2021)
Lattice data: S. Borsanyi et al., PRL (2021)

Equation of State

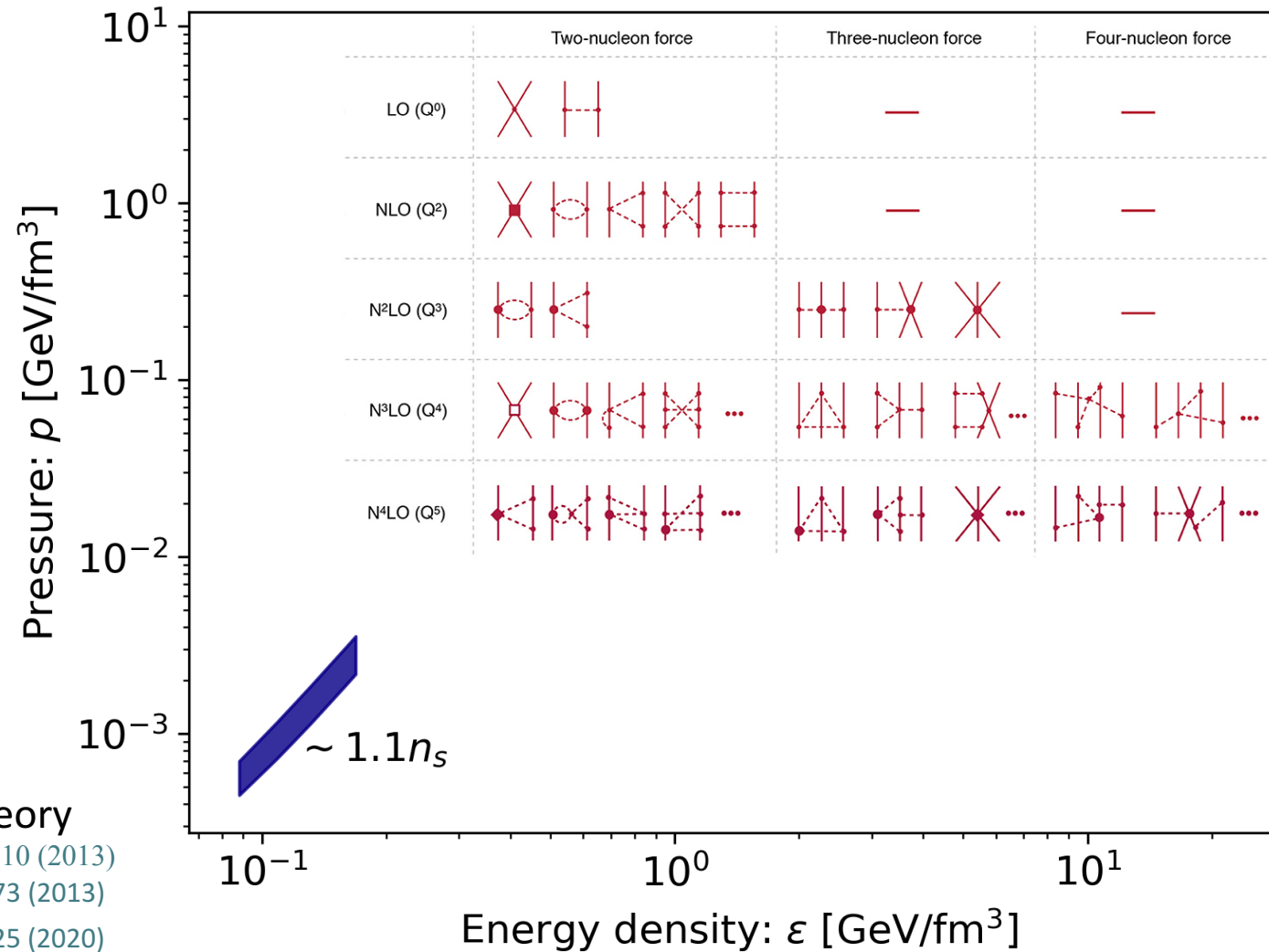


Model has a critical point at $\mu_B \sim 723$ MeV and $T \sim 89$ MeV

Equation of state can cover up to $\mu_B \sim 1$ GeV

Thermodynamic quantities seem too small at low T

Neutron stars and mergers



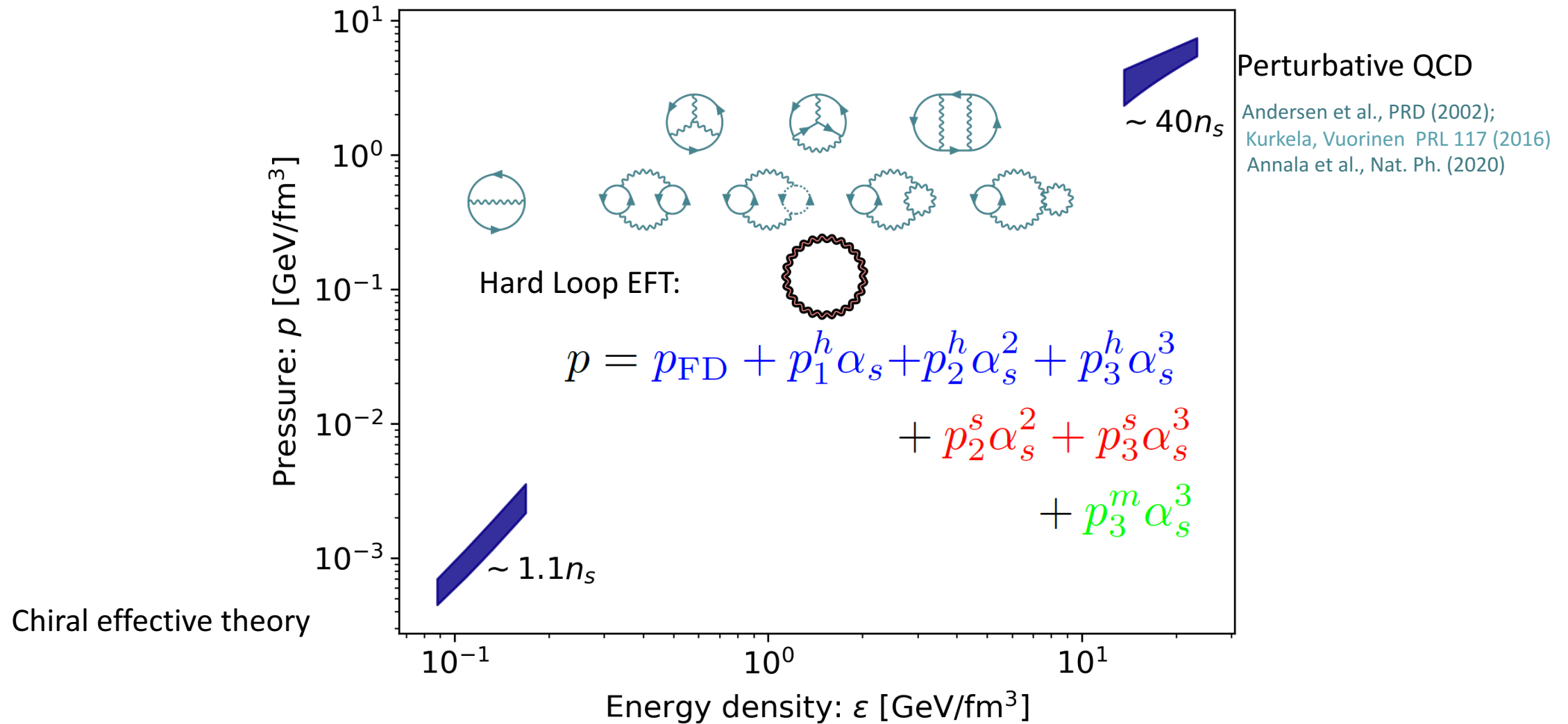
Chiral effective theory

Tews et al. PRL 110 (2013)

Hebeler, Lattimer et.al. APJ 773 (2013)

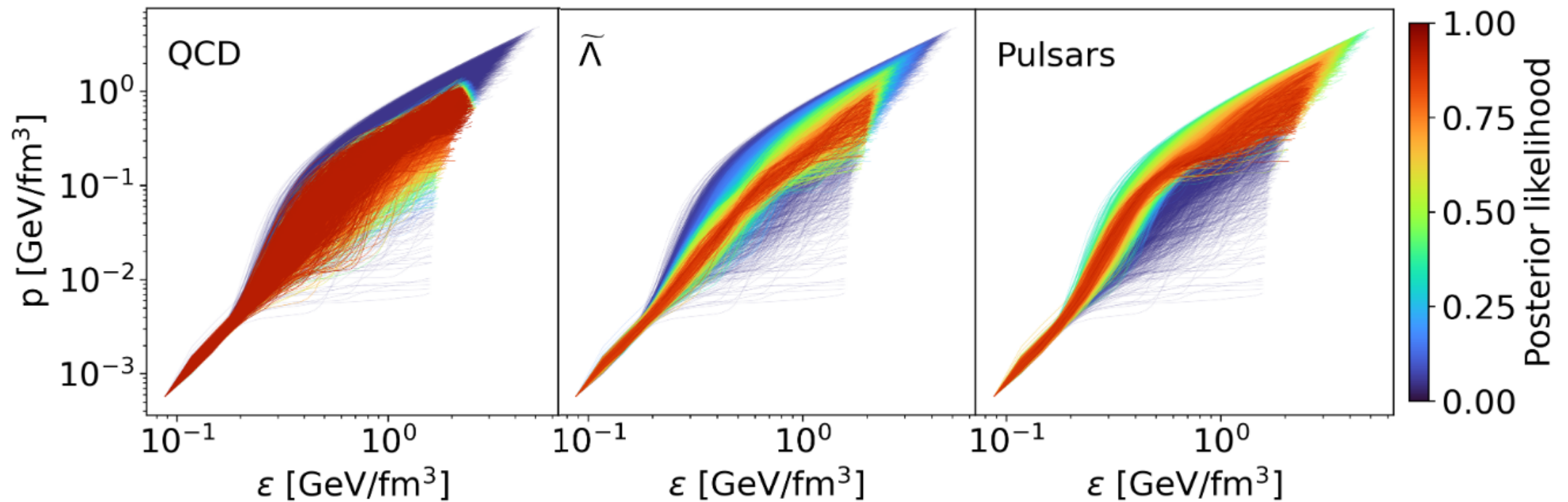
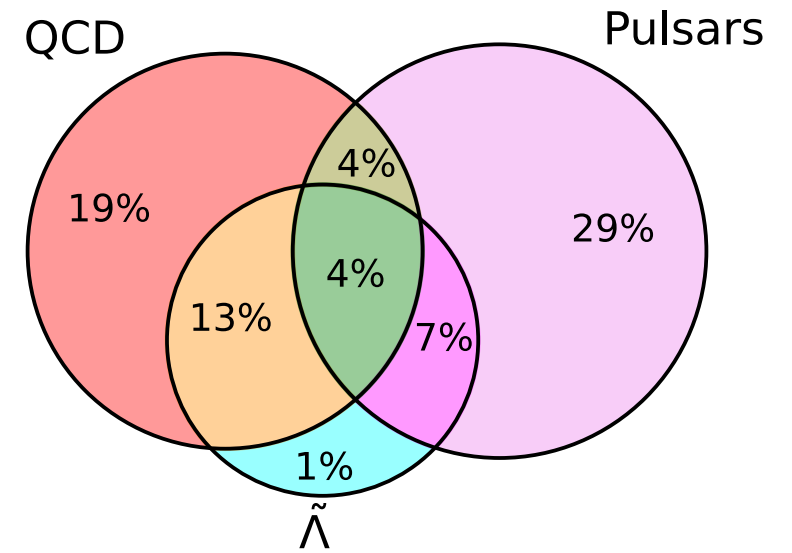
Drischler, Furnstahl et.al. PRL 125 (2020)

Neutron stars and mergers



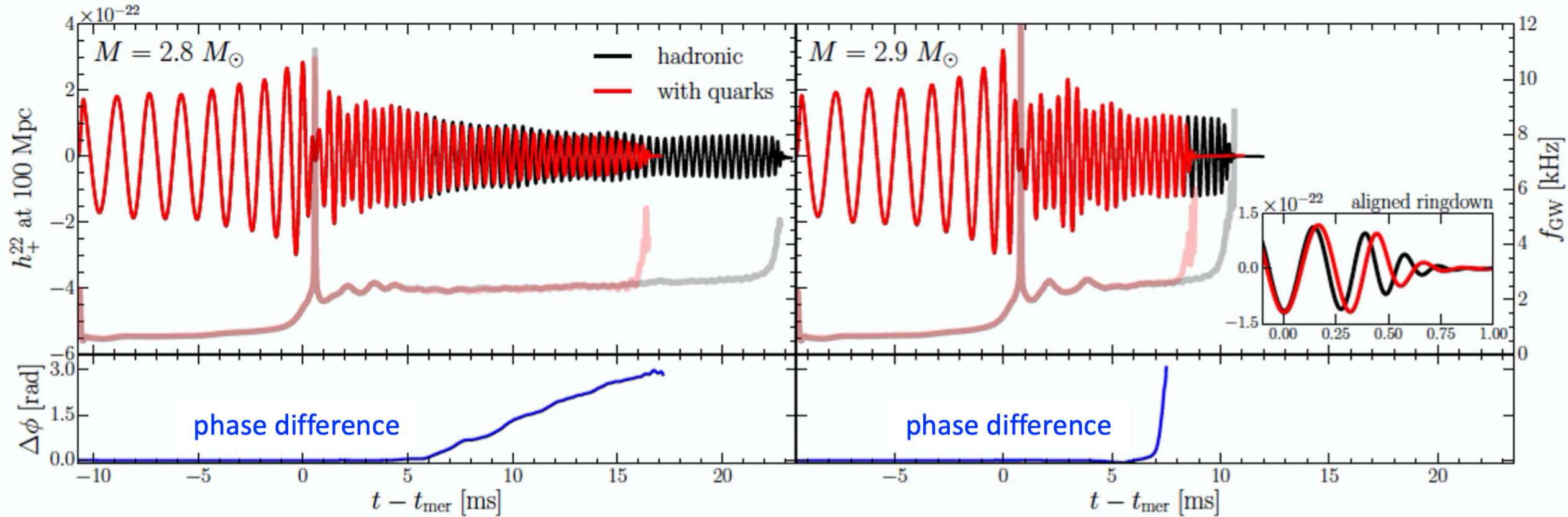
Neutron stars and mergers

QCD input complements NS observations

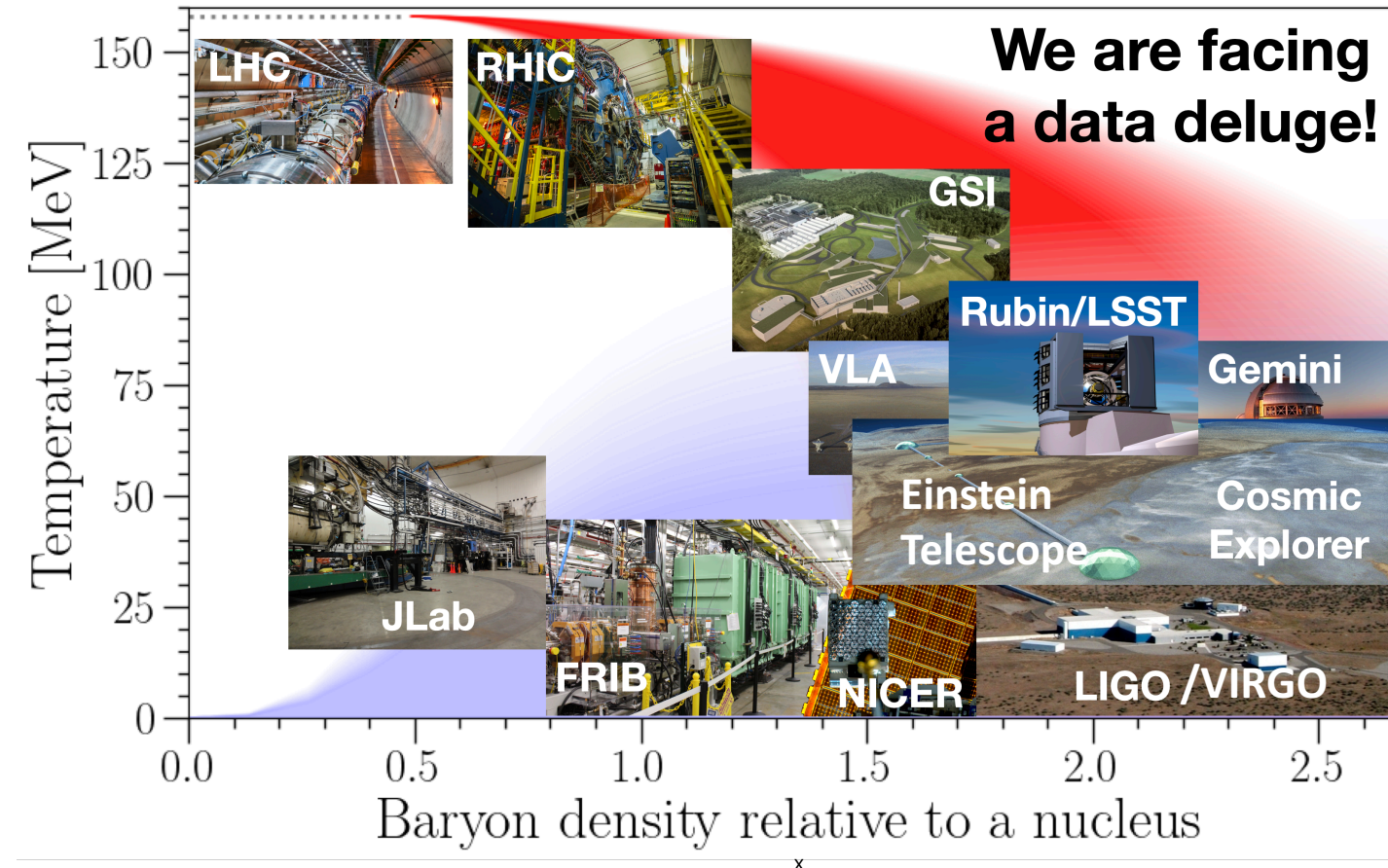


Neutron stars and mergers

- Post-merger signal sensitive to order of the phase transition
- Next generation observatories will be able to detect it!



Future directions



- ALICE will continue operating till at least 2036
- CBM at FAIR (GSI) will hopefully start soon
- New generation of gravitational wave observatories will be able to see the post-merger signal (sensitive to the EoS)
- FRIB & NUSTAR (FAIR) will study nuclear structure
- Theory needs to continue to support these experimental programs

Conclusions

- Need for quantitative results at finite-density to support the experimental programs and reach out to the Neutron Star merger regime
- Current lattice results for thermodynamics available up to $\mu_B/T \leq 3.5$
- Extensions to higher densities by means of effective theories

How can lattice QCD support the experiments?

Equation of state

- Needed for **hydrodynamic** description of the QGP

QCD phase diagram

- Transition line at finite density
- Constraints on the location of the critical point

Fluctuations of conserved charges

- Can be **simulated** on the lattice and **measured** in experiments
- Can give information on the **evolution** of heavy-ion collisions
- Can give information on the **critical point**

Formulation

S. Borsanyi, C. R. et al., PRL (2021)

- We have observed the $\hat{\mu}_B$ -dependence seems to amount to a simple T - rescaling
- A simplistic scenario with a single T - independent parameter κ does not provide a systematic treatment which can serve as an alternative expansion scheme
- We allow for more than $\mathcal{O}(\hat{\mu}^2)$ expansion of T' and let the coefficients be T -dependent:

$$\frac{\chi_1^B(T, \hat{\mu}_B)}{\hat{\mu}_B} = \chi_2^B(T', 0) , \quad T' = T (1 + \kappa_2(T) \hat{\mu}_B^2 + \kappa_4(T) \hat{\mu}_B^4 + \mathcal{O}(\hat{\mu}_B^6))$$

- **Important:** we are simply re-organizing the Taylor expansion via an expansion in the shift

$$\Delta T = T - T' = (\kappa_2(T) \hat{\mu}_B^2 + \kappa_4(T) \hat{\mu}_B^4 + \mathcal{O}(\hat{\mu}_B^6))$$

- Comparing the (Taylor) expansion in $\hat{\mu}_B$ and our expansion in ΔT order by order, we can relate $\chi_n^B(T)$ and $\kappa_n(T)$

QCD phase diagram

TRANSITION TEMPERATURE

TRANSITION LINE

TRANSITION WIDTH

QCD matter under extreme conditions

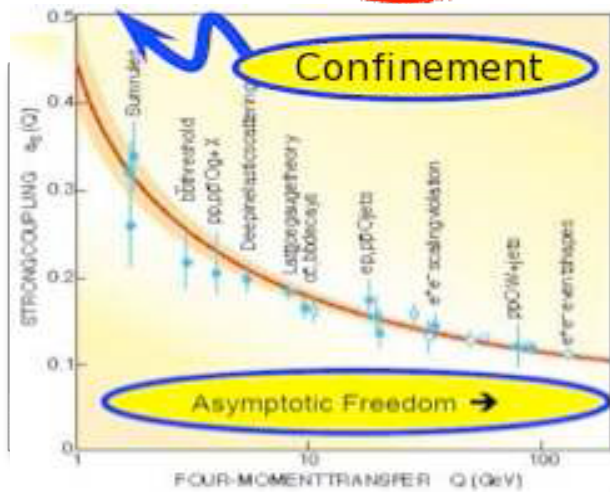
To address these questions, we need fundamental theory and experiment

Theory: Quantum Chromodynamics

- ▶ QCD is the fundamental theory of strong interactions
- ▶ It describes interactions among quarks and gluons

$$L_{QCD} = \sum_{i=1}^{n_f} \bar{\psi}_i \gamma_\mu \left(i \partial^\mu - g A_a^\mu \frac{\lambda_a}{2} \right) \psi_i - m_i \bar{\psi}_i \psi_i - \frac{1}{4} \sum_a F_a^{\mu\nu} F_a^{\mu\nu}$$

$$F_a^{\mu\nu} = \partial^\mu A_a^\nu - \partial^\nu A_a^\mu + i f_{abc} A_b^\mu A_c^\nu$$



Experiment: heavy-ion collisions



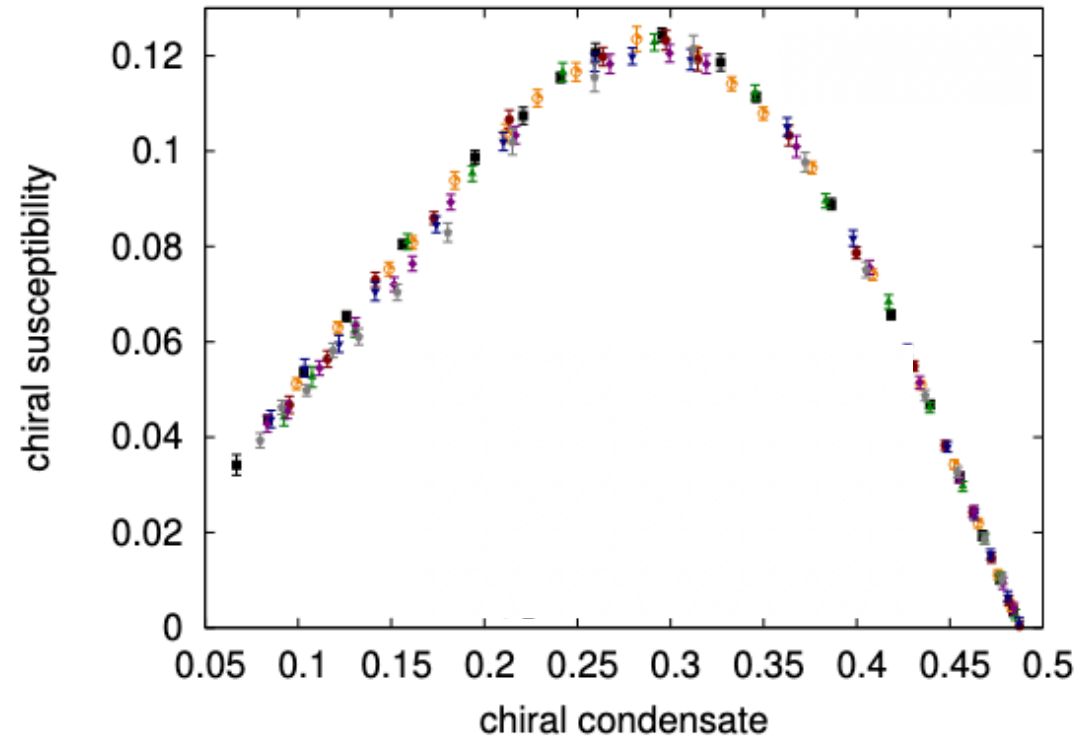
- ▶ Quark-Gluon Plasma (QGP) discovery at RHIC and LHC:
- ▶ SURPRISE!!! QGP is a **PERFECT FLUID**
- ▶ Changes our idea of QGP (no weak coupling)
- ▶ Microscopic origin still unknown



Phase Diagram from Lattice QCD

The transition at $\mu_B=0$ is a smooth crossover

Aoki et al., Nature (2006)
Borsanyi et al., JHEP (2010)
Bazavov et al., PRD (2012)



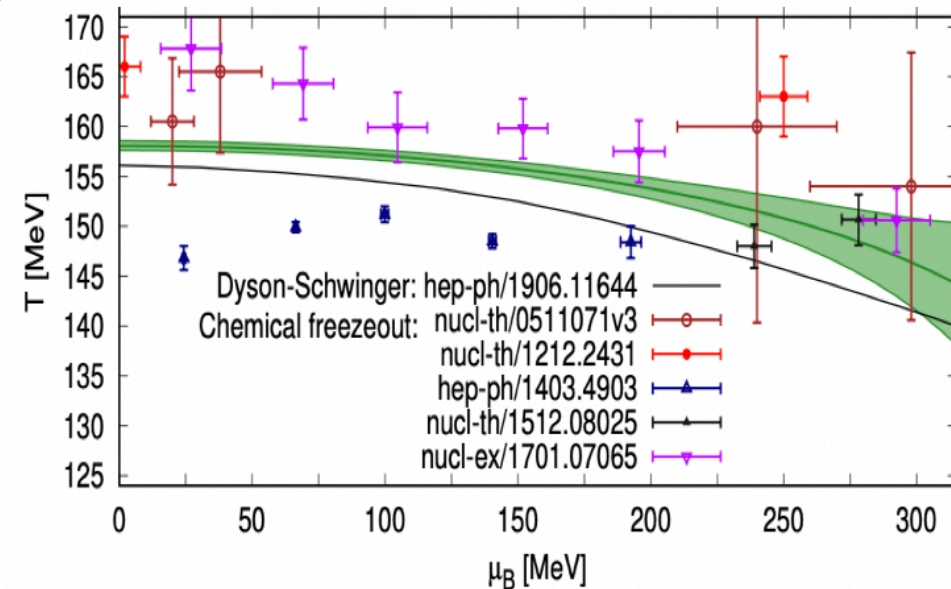
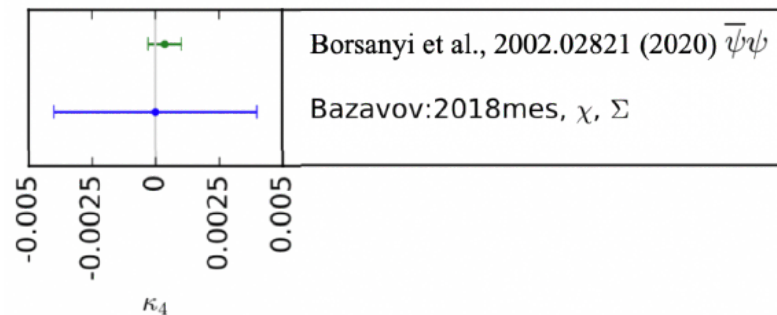
QCD transition temperature and curvature

$$\frac{T_c(\mu_B)}{T_0} = 1 - \kappa_2 \left(\frac{\mu_B}{T_0}\right)^2 - \kappa_4 \left(\frac{\mu_B}{T_0}\right)^4 + O(\mu_B^6)$$

Borsanyi, C. R. et al. PRL (2020)

- Latest results on T_0 from WB collaboration based on subtracted chiral condensate and chiral susceptibility

$T_0 = 158.0 \pm 0.6$ MeV

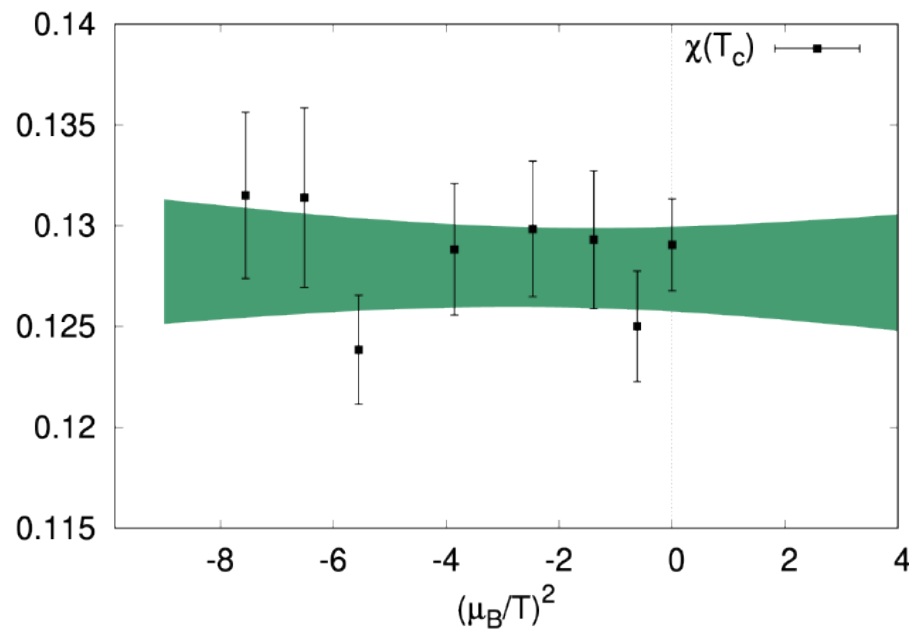


$\kappa_2 = 0.0153 \pm 0.0018$,
 $\kappa_4 = 0.00032 \pm 0.00067$

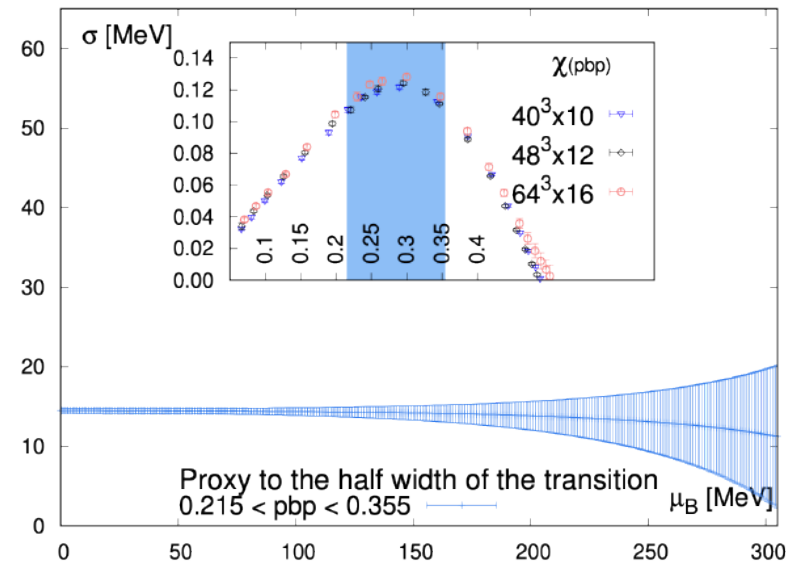
Limit on the location of the critical point

For a genuine phase transition, the height of the peak of the chiral susceptibility diverges and the width shrinks to zero

Height of chiral susceptibility peak



Width of chiral susceptibility peak



No sign of criticality for $\mu_B < 300$ MeV

Borsanyi, C. R. et al. PRL (2020)

Fluctuations of conserved charges

COMPARISON TO EXPERIMENT

CHEMICAL FREEZE-OUT PARAMETERS

Fluctuations of conserved charges

Definition:

$$\chi_{lmn}^{BSQ} = \frac{\partial^{l+m+n} p / T^4}{\partial(\mu_B/T)^l \partial(\mu_S/T)^m \partial(\mu_Q/T)^n}.$$

Relationship between chemical potentials:

$$\mu_u = \frac{1}{3}\mu_B + \frac{2}{3}\mu_Q;$$

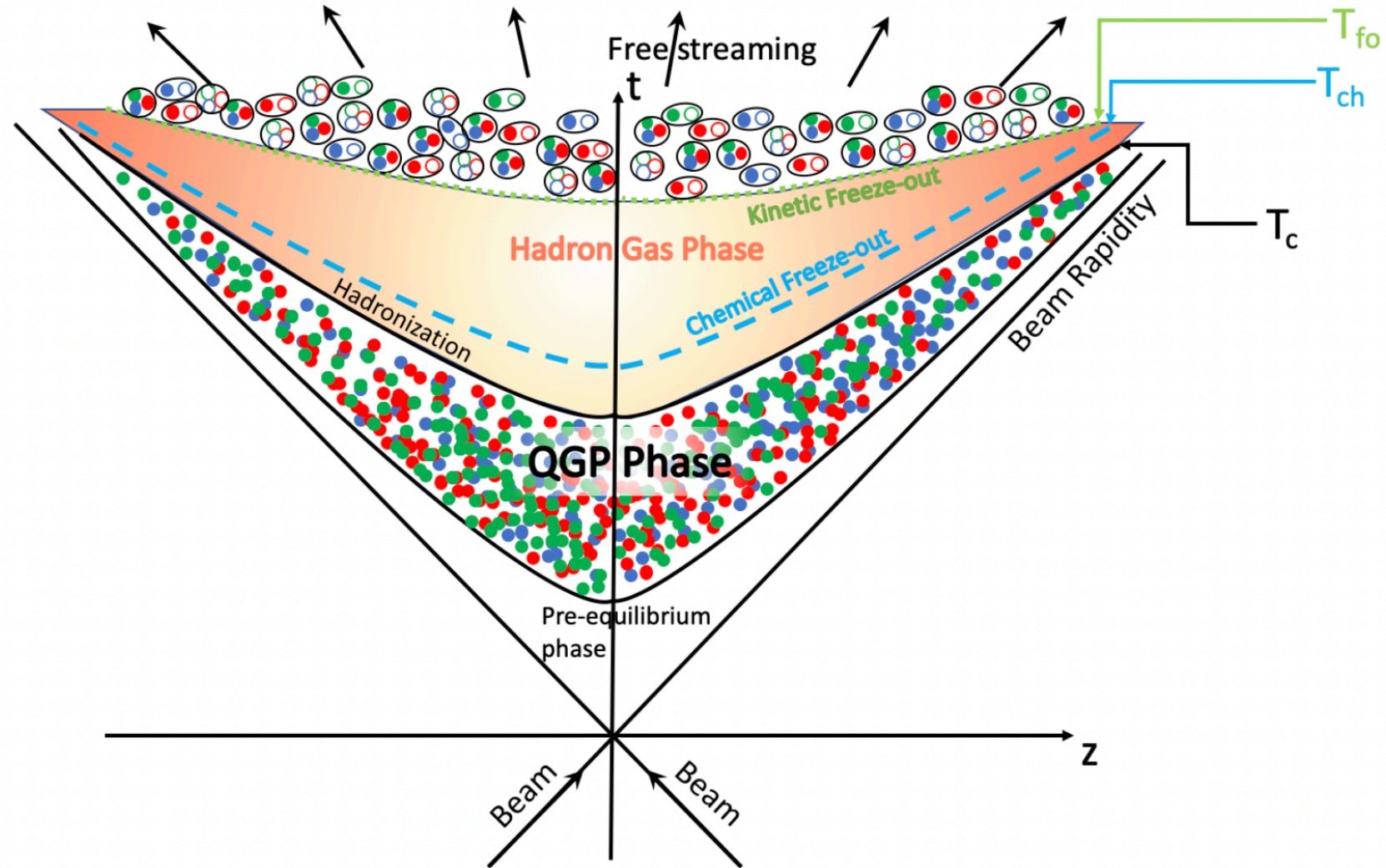
$$\mu_d = \frac{1}{3}\mu_B - \frac{1}{3}\mu_Q;$$

$$\mu_s = \frac{1}{3}\mu_B - \frac{1}{3}\mu_Q - \mu_S.$$

They can be calculated on the lattice and compared to experiment

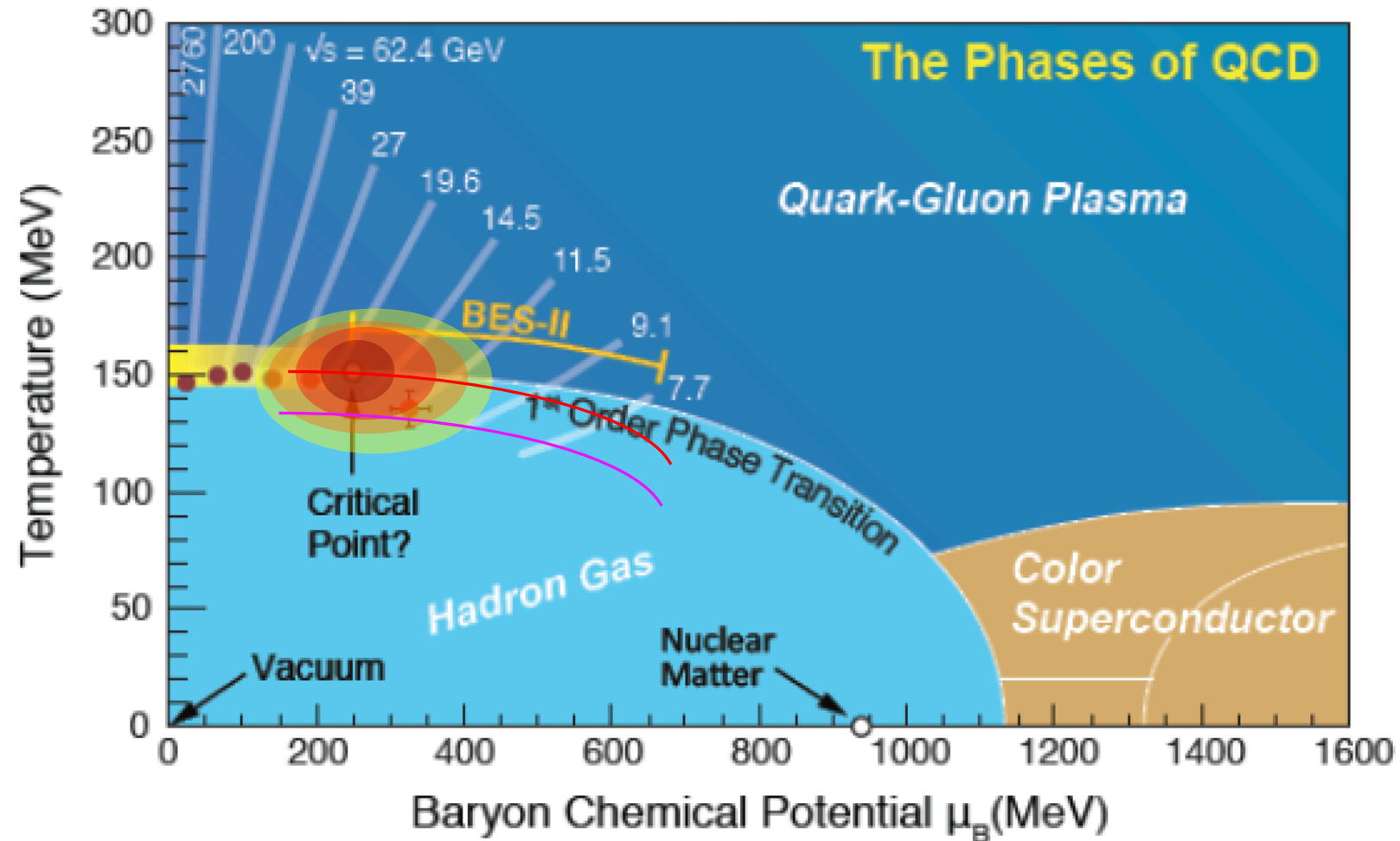
Evolution of a heavy-ion collision

- **Chemical freeze-out:** inelastic reactions cease: the chemical composition of the system is fixed (particle yields and fluctuations)
- **Kinetic freeze-out:** elastic reactions cease: spectra and correlations are frozen (free streaming of hadrons)
- Hadrons reach the detector



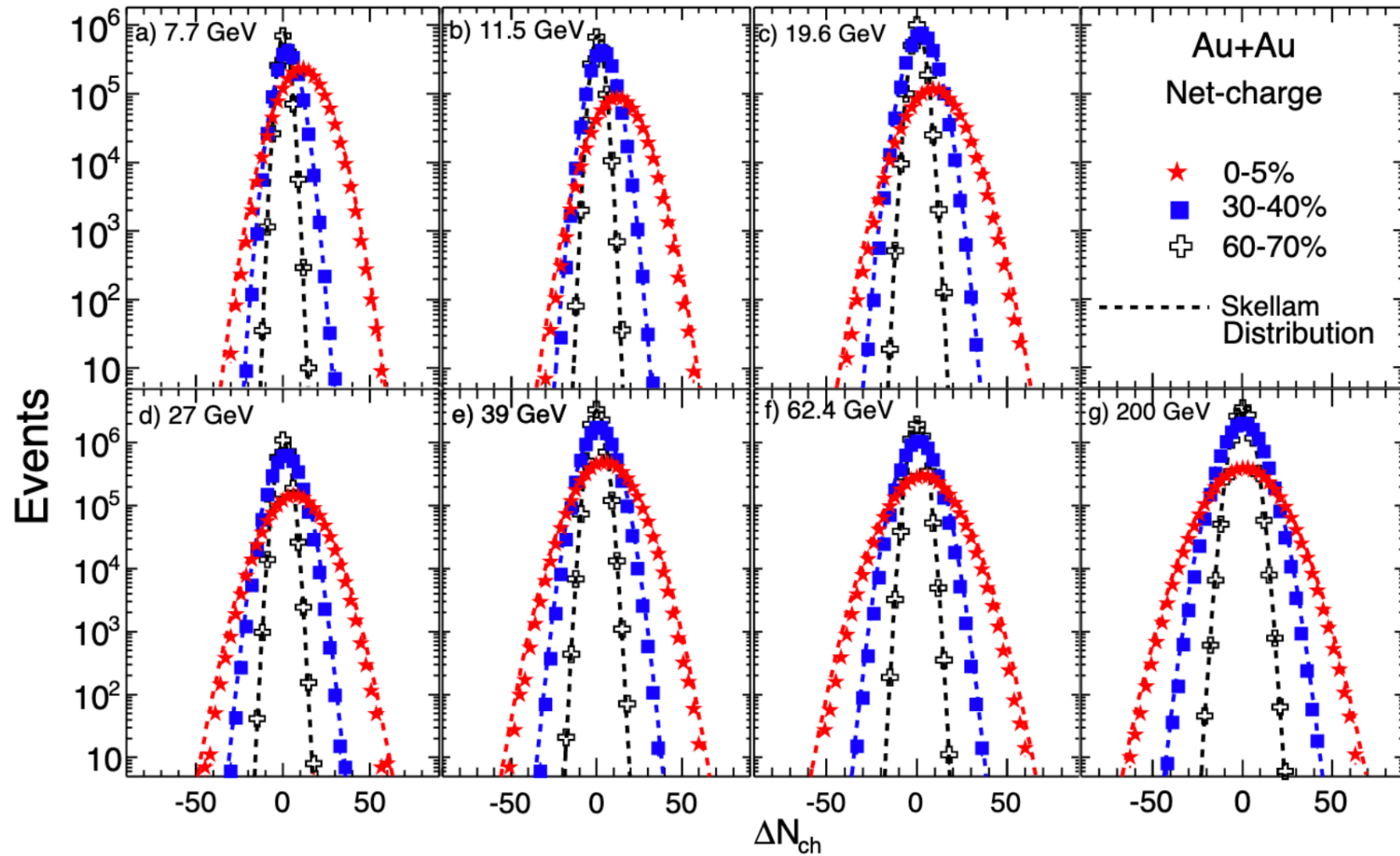
Evolution of a heavy-ion collision

- **Chemical freeze-out:** inelastic reactions cease: the chemical composition of the system is fixed (particle yields and fluctuations)
- **Kinetic freeze-out:** elastic reactions cease: spectra and correlations are frozen (free streaming of hadrons)
- Hadrons reach the detector



Connection to experiment

- Consider the number of electrically charged particles N_Q
- Its average value over the whole ensemble of events is $\langle N_Q \rangle$
- In experiments it is possible to measure its **event-by-event distribution**



Connection to experiment

Fluctuations of conserved charges are the **cumulants** of their event-by-event distribution

$$\text{mean : } M = \chi_1$$

$$\text{variance : } \sigma^2 = \chi_2$$

$$\text{skewness : } S = \chi_3/\chi_2^{3/2}$$

$$\text{kurtosis : } \kappa = \chi_4/\chi_2^2$$

$$S\sigma = \chi_3/\chi_2$$

$$\kappa\sigma^2 = \chi_4/\chi_2$$

$$M/\sigma^2 = \chi_1/\chi_2$$

$$S\sigma^3/M = \chi_3/\chi_1$$

F. Karsch: Centr. Eur. J. Phys. (2012)

The chemical potentials are not independent: fixed to match the experimental conditions:

$$\langle n_S \rangle = 0$$

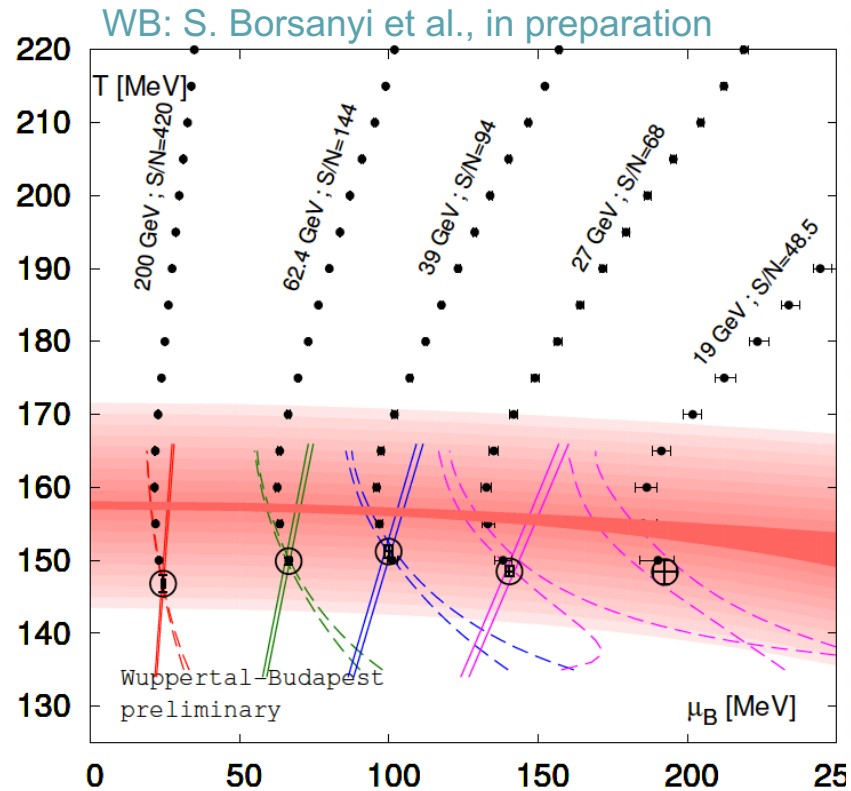
$$\langle n_Q \rangle = 0.4 \langle n_B \rangle$$

Freeze-out line from first principles

Use T - and μ_B -dependence of R_{12}^Q and R_{12}^B for a combined fit:

$$R_{12}^Q(T, \mu_B) = \frac{\chi_1^Q(T, \mu_B)}{\chi_2^Q(T, \mu_B)} = \frac{\chi_{11}^{QB}(T, 0) + \chi_2^Q(T, 0)q_1(T) + \chi_{11}^{QS}(T, 0)s_1(T)}{\chi_2^Q(T, 0)} \frac{\mu_B}{T} + \mathcal{O}(\mu_B^3).$$

$$R_{12}^B(T, \mu_B) = \frac{\chi_1^B(T, \mu_B)}{\chi_2^B(T, \mu_B)} = \frac{\chi_2^B(T, 0) + \chi_{11}^{BQ}(T, 0)q_1(T) + \chi_{11}^{BS}(T, 0)s_1(T)}{\chi_2^B(T, 0)} \frac{\mu_B}{T} + \mathcal{O}(\mu_B^3)$$



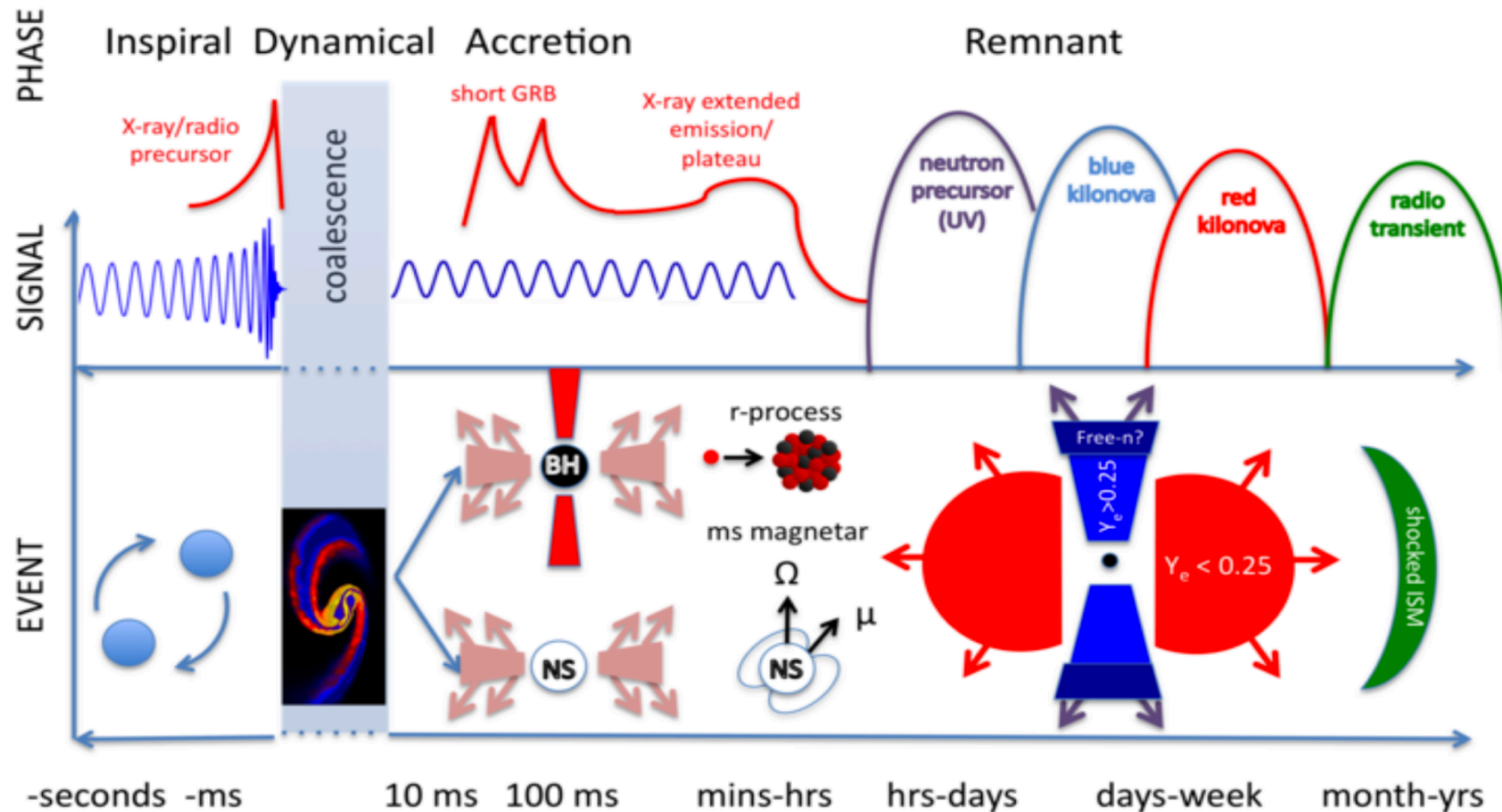
S/N=const from lattice EOS [WB 2015] \bullet
 HRG analysis [Alba et al] \circ
 T_c from lattice [WB 1507.07510] ■



Conclusions

- Need for quantitative results at finite-density to support the experimental programs
 - Equation of state
 - Phase transition line
 - Fluctuations of conserved charges
- Current lattice results for thermodynamics up to $\mu_B/T \leq 3.5$
- Extensions to higher densities by means of lattice-based models
- No indication of Critical Point from lattice QCD in the explored μ_B range

Anatomy of a multi-messenger merger



Nuclear Physics is encoded in all phases!

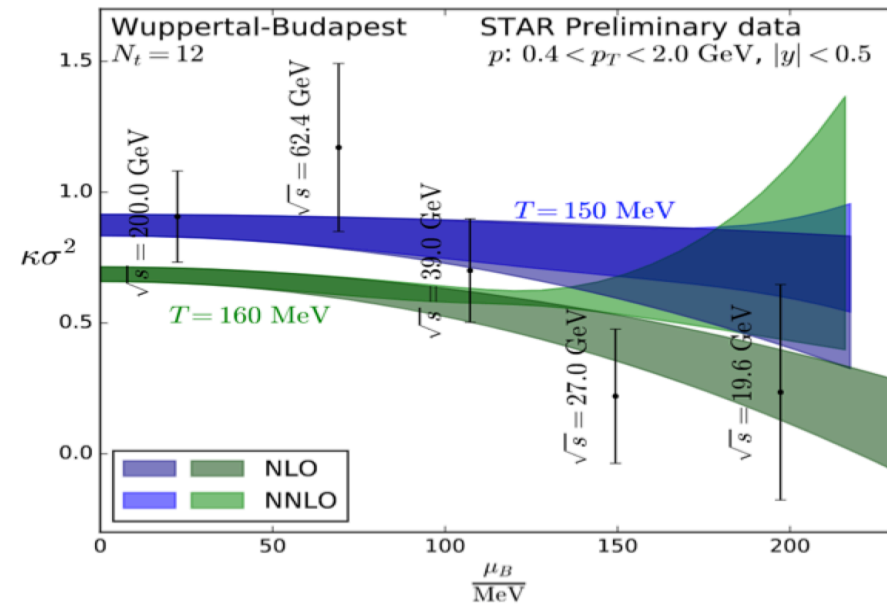
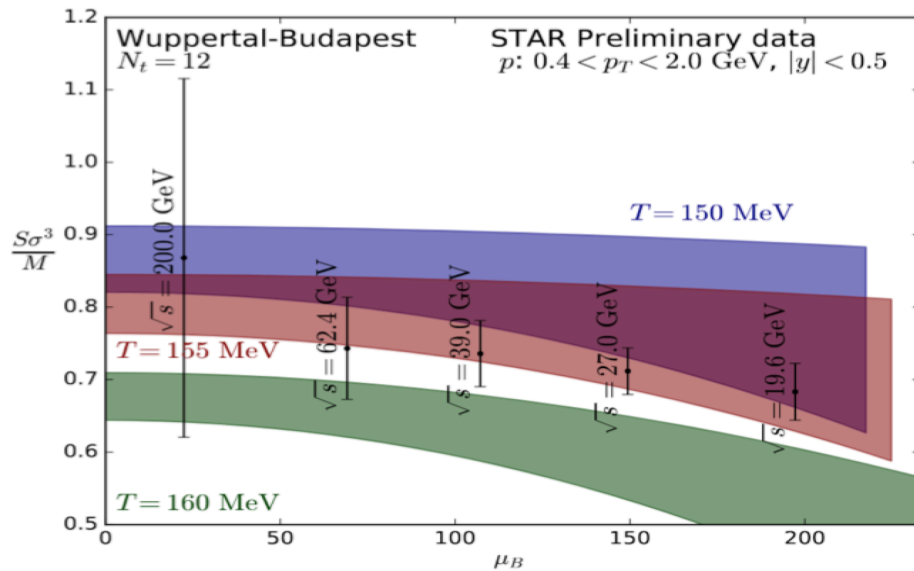


Higher order fluctuations

$$\frac{S_B \sigma_B^3}{M_B} = \frac{\chi_3^B(T, \mu_B)}{\chi_1^B(T, \mu_B)} = \frac{\chi_4^B + s_1 \chi_{31}^{BS} + q_1 \chi_{31}^{BQ}}{\chi_2^B + s_1 \chi_{11}^{BS} + q_1 \chi_{11}^{BQ}} + \mathcal{O}(\mu_B^2) \equiv r_{31}^{B,0} + r_{31}^{B,2} \hat{\mu}_B^2 + \mathcal{O}(\mu_B^4)$$

$$\kappa_B \sigma_B^2 = \frac{\chi_4^B(T, \mu_B)}{\chi_2^B(T, \mu_B)} = \frac{\chi_4^B}{\chi_2^B} + \mathcal{O}(\mu_B^2) \equiv r_{42}^{B,0} + r_{42}^{B,2} \hat{\mu}_B^2 + \mathcal{O}(\mu_B^4),$$

WB, JHEP (2018)



See also HotQCD, PRD (2017)

“Baryometer and Thermometer”

Let us look at the Taylor expansion of $R^{B_{31}}$

$$R_{31}^B(T, \mu_B) = \frac{\chi_3^B(T, \mu_B)}{\chi_1^B(T, \mu_B)} = \frac{\chi_4^B(T, 0) + \chi_{31}^{BQ}(T, 0)q_1(T) + \chi_{31}^{BS}(T, 0)s_1(T)}{\chi_2^B(T, 0) + \chi_{11}^{BQ}(T, 0)q_1(T) + \chi_{11}^{BS}(T, 0)s_1(T)} + \mathcal{O}(\mu_B^2)$$

- To order μ_B^2 it is independent of μ_B : it can be used as a **thermometer**
- Let us look at the Taylor expansion of $R^{B_{12}}$

$$R_{12}^B(T, \mu_B) = \frac{\chi_1^B(T, \mu_B)}{\chi_2^B(T, \mu_B)} = \frac{\chi_2^B(T, 0) + \chi_{11}^{BQ}(T, 0)q_1(T) + \chi_{11}^{BS}(T, 0)s_1(T)}{\chi_2^B(T, 0)} \frac{\mu_B}{T} + \mathcal{O}(\mu_B^3)$$

- Once we extract T from $R^{B_{31}}$, we can use $R^{B_{12}}$ to extract μ_B

The highest man-made temperature



5.5×10^{12} °C: 340.000 times the temperature at the center of the sun!!!

The produced energy density is $5 \text{ GeV}/\text{fm}^3$

In one year in America $\sim 10^{20}$ J of energy are used

$$10^{20} \text{ J} \times 1\text{eV}/(1.6 \times 10^{-19} \text{ J}) = 6.6 \times 10^{38} \text{ eV}$$

At $5 \text{ GeV}/\text{fm}^3$ this would correspond to a volume:

$$6.6 \times 10^{38} \text{ eV} \div \frac{5 \times 10^9 \text{ eV}}{\text{fm}^3} = 1.3 \times 10^{29} \text{ fm}^3$$

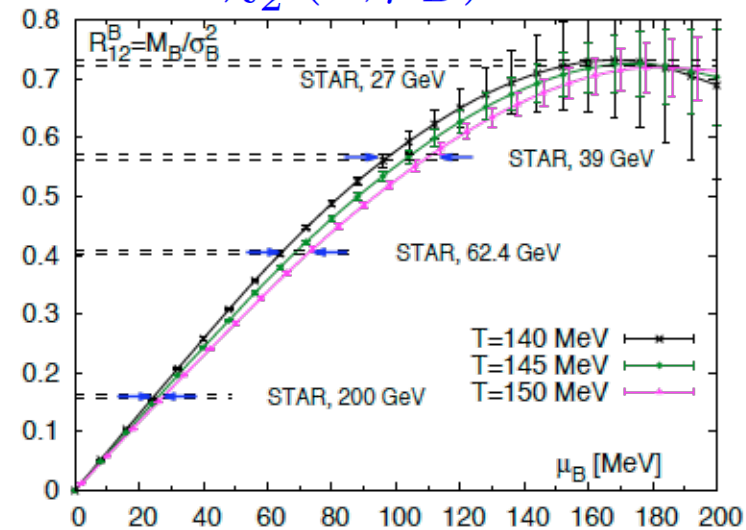
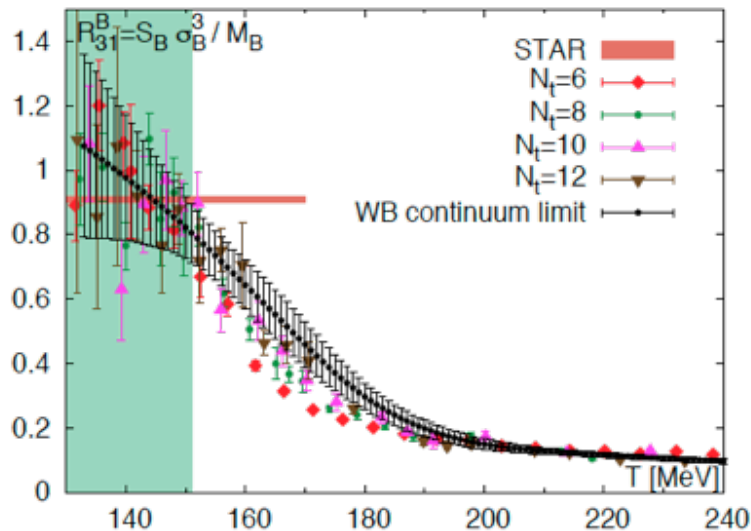
Or, equivalently, to a box of size:

$$\sqrt[3]{1.3 \times 10^{29} \text{ fm}^3} = 5 \times 10^9 \text{ fm} \times \frac{1\text{m}}{10^{15} \text{ fm}} \times \frac{10^6 \mu\text{m}}{1\text{m}} = 5 \mu\text{m}$$

Freeze-out parameters from B fluctuations

➤ Thermometer: $\frac{\chi_3^B(T, \mu_B)}{\chi_1^B(T, \mu_B)} = S_B \sigma_B^3 / M_B$

Baryometer: $\frac{\chi_1^B(T, \mu_B)}{\chi_2^B(T, \mu_B)} = \sigma_B^2 / M_B$



Upper limit: $T_f \leq 151 \pm 4$ MeV

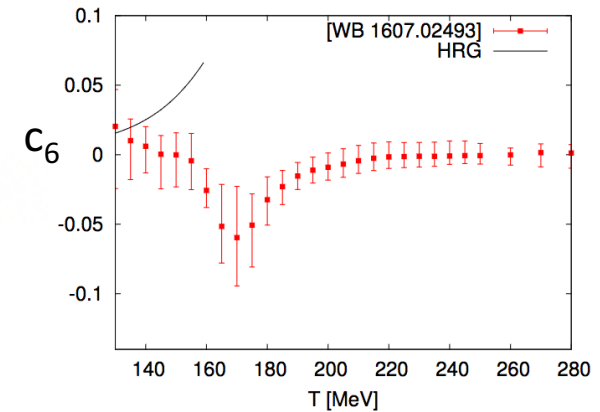
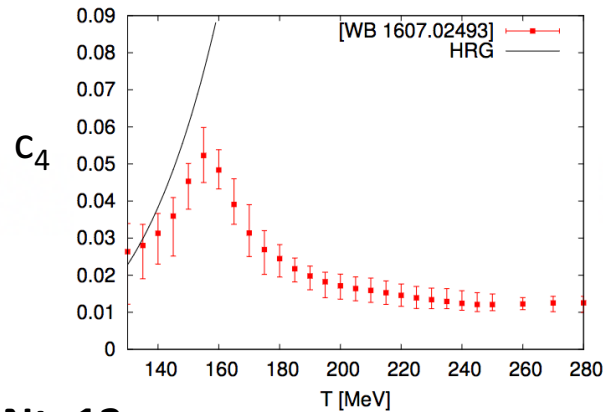
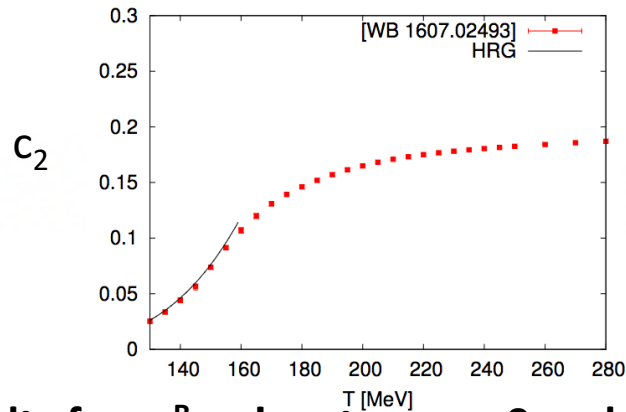
WB: S. Borsanyi et al., PRL (2014)
STAR collaboration, PRL (2014)

Consistency between freeze-out chemical potential from electric charge and baryon number is found.

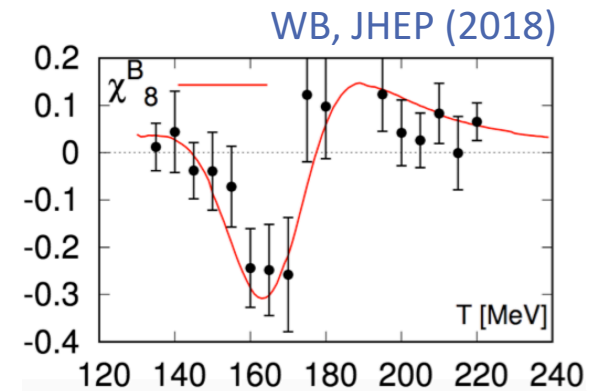
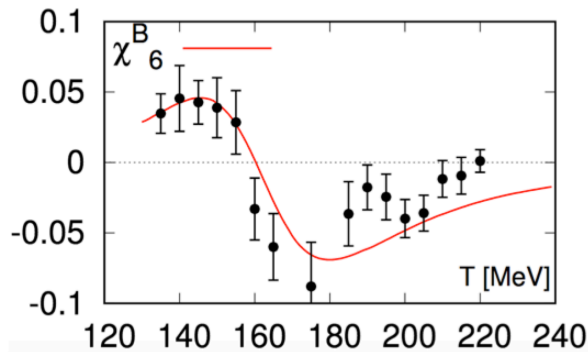
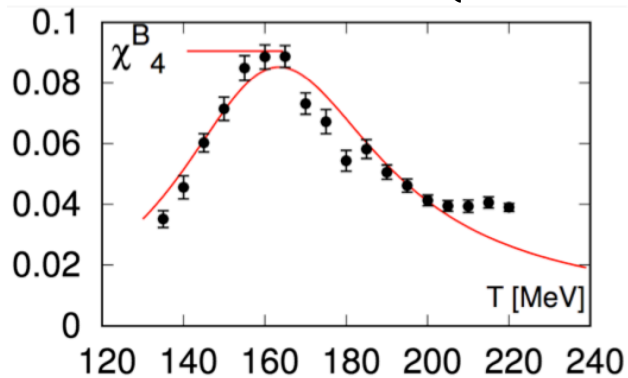
Pressure coefficients

Simulations at imaginary μ_B :

Continuum, $O(10^4)$ configurations, errors include systematics (WB: NPA (2017)) **Strangeness neutrality**



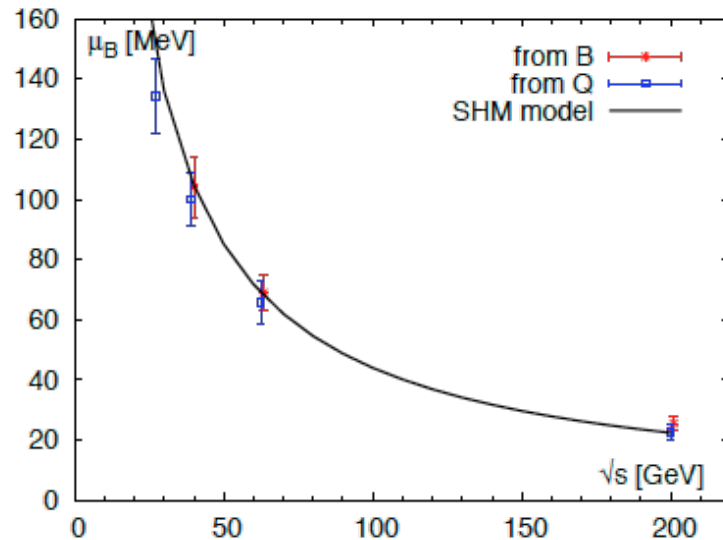
Results for $\chi_n^B = n!c_n$ at $\mu_S = \mu_Q = 0$ and $Nt=12$



Freeze-out parameters from B fluctuations

➤ Thermometer: $\frac{\chi_3^B(T, \mu_B)}{\chi_1^B(T, \mu_B)} = S_B \sigma_B^3 / M_B$

Baryometer: $\frac{\chi_1^B(T, \mu_B)}{\chi_2^B(T, \mu_B)} = \sigma_B^2 / M_B$



\sqrt{s} [GeV]	μ_B^f [MeV] (from B)	μ_B^f [MeV] (from Q)
200	25.8 ± 2.7	22.8 ± 2.6
62.4	69.7 ± 6.4	66.6 ± 7.9
39	105 ± 11	101 ± 10
27	-	136 ± 13.8

WB: S. Borsanyi et al., PRL (2014)
STAR collaboration, PRL (2014)

Upper limit: $T_f \leq 151 \pm 4$ MeV

Consistency between freeze-out chemical potential from electric charge and baryon number is found.

A few Lessons learned

- Heavy ion collisions:
 - Phase transition at small μ_B is a smooth crossover
 - If a critical point exists, it is in the 3D-Ising model universality class
 - Equation of state and phase diagram are known from 1st principles at $\mu_B/T < 3.5$
 - Quark-Gluon Plasma is a strongly coupled fluid with very small viscosity/entropy
- Neutron star mergers:
 - GWs travel essentially at the speed of light
 - binary neutron star mergers are progenitors of short gamma ray bursts
 - they are prolific sites for the formation of heavy elements
 - constrained neutron-star radii to be between 9.5 and 13 km

Anatomy of a heavy-ion collision

