# Supernova pointing with DUNE

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#### Overview

- DUNE detector and interaction channels
- Interaction channels for supernova (SN) pointing
- SN direction reconstruction (methods developed by AJ Roeth and James Shen)
- SN pointing visualization



## DUNE – The Far Detector

- 40 kt fiducial volume LArTPC
- Events create charge carriers that drift to read out wires, resulting in charge waveforms
- Far detector will be located 1 mile underground to reduce cosmological backgrounds
- LArSoft particle event simulation software
  → simulates neutrino events in the LArTPC
  to generates readout waveforms &
  perform event reconstruction/analysis

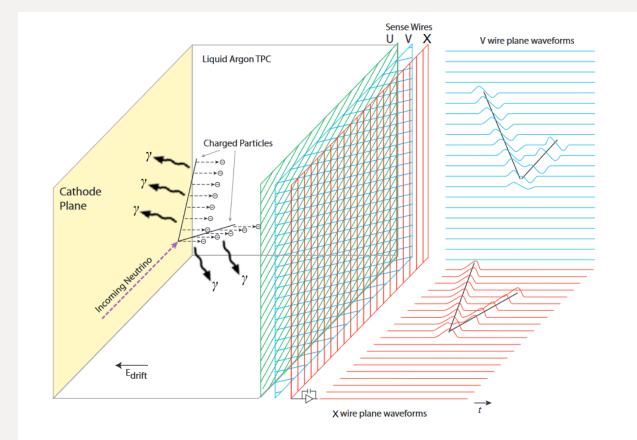
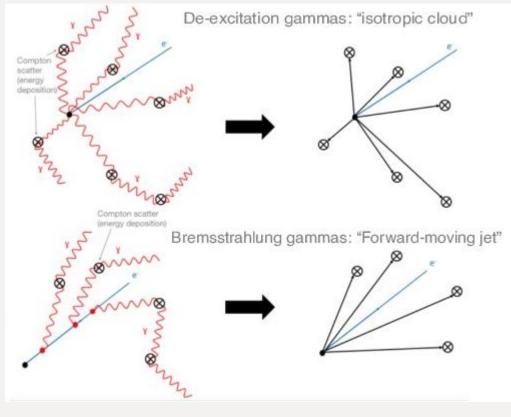


Image from DUNE Technical Design Report. arXiv: 2002.03005



#### DUNE – Interaction channels

- DUNE will look for SN activity through **two channels:** charged current interactions and elastic scattering interactions.
- **Charged current:** electron neutrinos interact with Ar-40 to create an electron and excited K-40, which deexcites through emission of photons
- Elastic scattering: neutrinos of all flavors scatter off electrons, leaving an ionization path in the LAr as the electron loses energy.
- Other Channels: neutral current channel is ignored

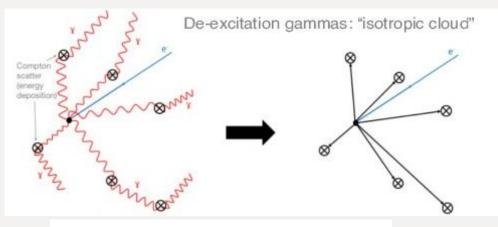


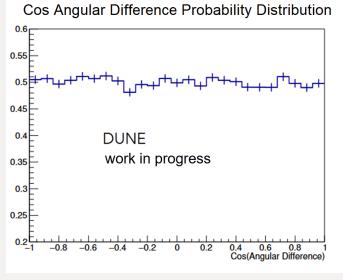


## Charged Current (CC) Interactions

- Large cross section event counts will be useful in triggering real time SN detection
- Gammas and electron are not correlated with the incident neutrino direction, so there is no directional information.

• Define 
$$\cos(\theta) = \hat{d}_{SN} \cdot \hat{d}_{event}$$

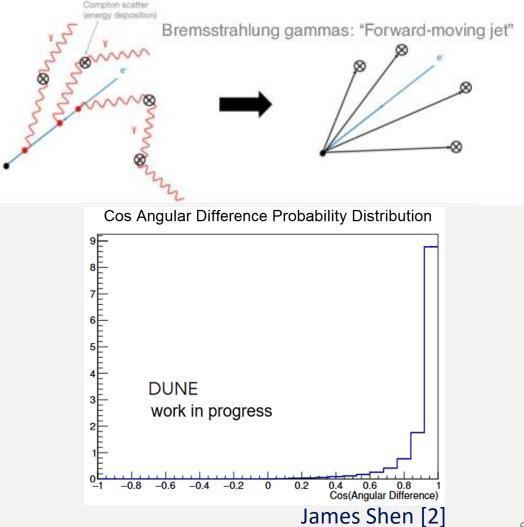




James Shen [2]

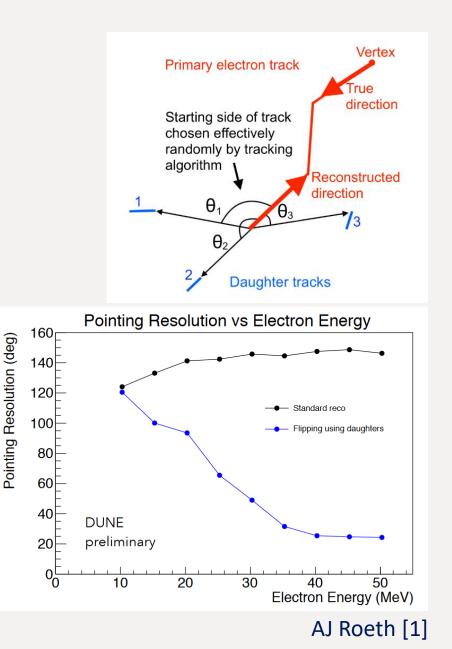
## Elastic Scattering (ES) Interactions

- Smaller cross section than CC
- Emitted electron has strong correlation with the incident neutrino direction – useful for SN pointing
- In reconstruction, starting direction of electron is ambiguous – daughter particles can infer correct direction



# Daughter Flipping

- ES event daughter particles are distributed in the direction of the primary electron
- Higher energy events → more daughter particles and better disambiguation of primary electron direction

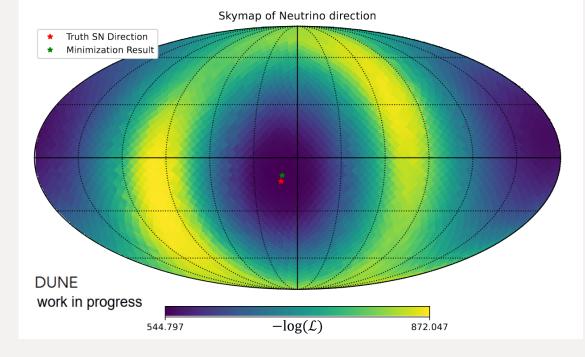


## Determining SN direction

 Use maximum likelihood estimation (MLE) on the collected events

$$\max(\mathcal{L}) = \max \prod_{i} p_i(\cos(\theta)) = \min \sum_{i} -\log(p_i(\cos(\theta)))$$

• Minimizing the negative log likelihood will give the predicted SN direction

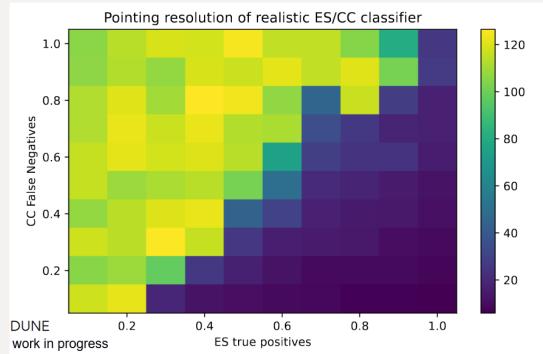


James Shen [2]

• Example skymap: ~300 ES events

## Classification

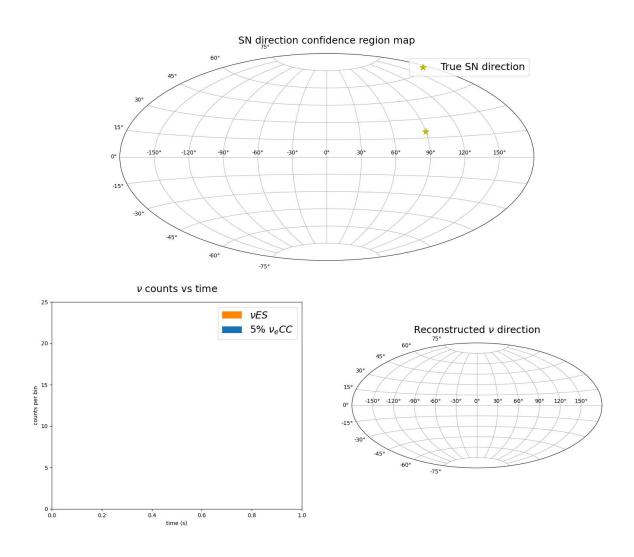
- In order to get the best SN pointing information, we want to only look at the ES events when doing pointing analysis
- Confusion matrix:  $C = \begin{pmatrix} C_{ES} \rightarrow C_{ES} & C_{CC} \rightarrow C_{ES} \\ C_{ES} \rightarrow C_{CC} & C_{CC} \rightarrow C_{CC} \end{pmatrix}$
- Optimistic BDT classification result:  $C_{ES} \rightarrow C_{ES} = 0.86$ ,  $C_{CC} \rightarrow C_{ES} = 0.04$ Pointing resolution: 5.3 degrees [2]



James Shen [2]

## Visualization

- A visualization tool for real time SN neutrino pointing
- Implemented here with a toy model based on previous works ([1] & [2])
- Demonstrates pointing of a 10 kpc SN, with only 5% CC events misclassified as ES



### References

[1] AJ Roeth. Supernova Pointing Resolution of DUNE. *Bulletin of the American Physical society*, 64, 2019.

[2] James Shen. Supernova Pointing Resolution of DUNE. April 2022.

[3] Erin Conley. Using boosted decision trees to identify supernova neutrino interactions in DUNE, June 2020.