

International School on Theory & Analysis in Particle Physics

# Basic Concepts: Relativistic notation, kinematics, and wave equations

Contents: Special Relativity, Lorentz transformations, four-vectors, natural units, collisions, Klein-Gordon equation, Dirac equation

Textbook: Introduction to Elementary Particles, Griffiths

Supplementary textbooks:

Modern Elementary Particle Physics, Kane

Quantum Field Theory, Ryder

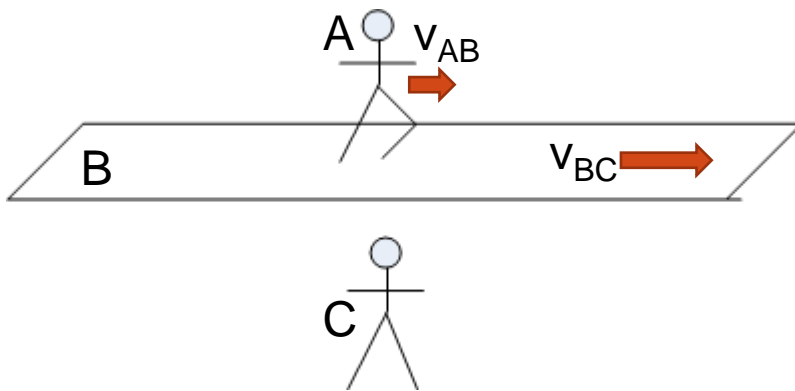
Introducing Einstein's Relativity, d'Inverno

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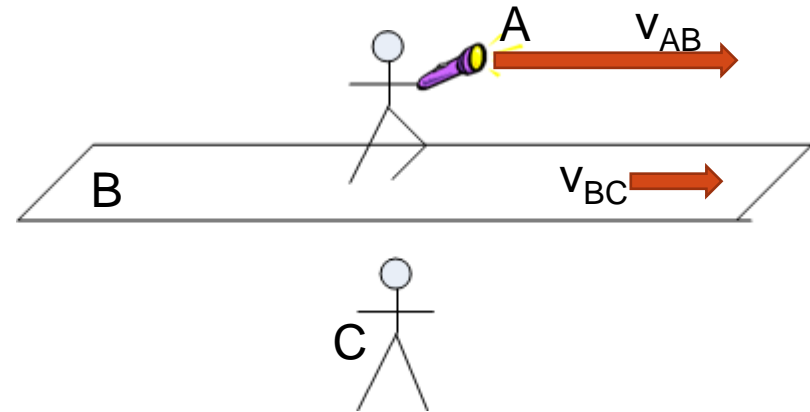
## Special Relativity (Einstein 1905)

1. The principle of relativity  
(The same laws apply in all inertial reference frames)
2. The universal speed of light  
(The speed of light in vacuum ( $c$ ) is the same in all inertial ref. frames)



$$v_{AC} = v_{AB} + v_{BC}$$

$$v_{AC} = \frac{v_{AB} + v_{BC}}{1 + (v_{AB}v_{BC}/c^2)}$$



$$v_{AC} = v_{AB} = c$$

## Consequences of Special Relativity

### i. Relativity of simultaneity

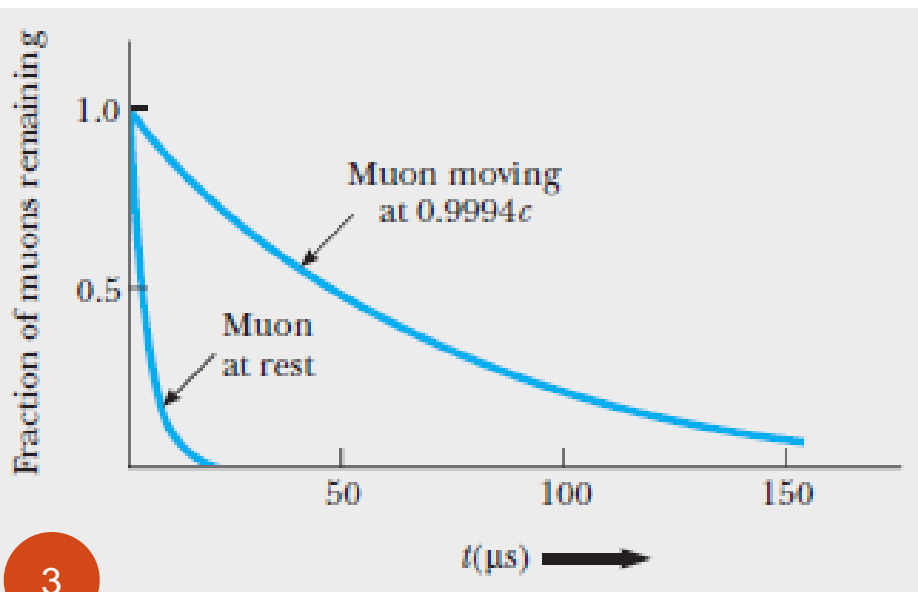
Two events that are simultaneous in S are in general not simultaneous in S'

### ii. Lorentz contraction

Moving objects are shortened by a factor  $\gamma = \frac{1}{\sqrt{1 - v^2/c^2}}$

### iii. Time dilation

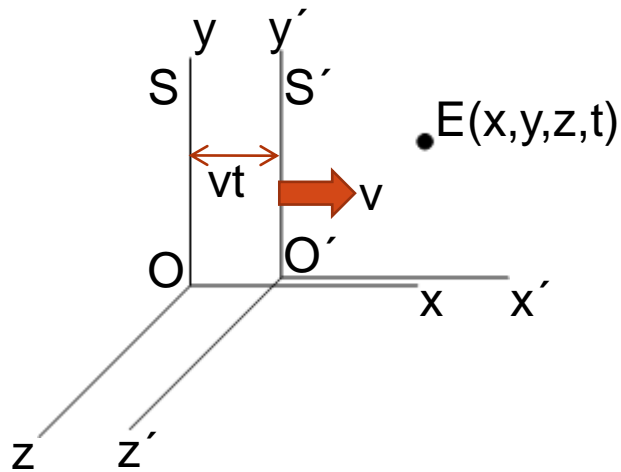
Moving clocks run slow by the same factor  $\gamma$



$$\mu \text{ at rest: } \tau_0 = 2 \times 10^{-6} \mu\text{s}$$

$$\mu \text{ moving at } v = 0.9994c:$$

$$\tau = \gamma \tau_0 = 58 \times 10^{-6} \mu\text{s}$$



## Galilean Transformations

$$x' = x - vt$$

$$y' = y$$

$$z' = z$$

$$t' = t$$

## Lorentz Transformations

$$x' = \gamma(x - vt)$$

$$y' = y$$

$$z' = z$$

$$t' = \gamma \left( t - \frac{v}{c^2} x \right)$$

HMW 1: Verify relativity of simultaneity, time dilation, length contraction, and velocity addition rule, using Lorentz transformations.

Position-time four vector:  $x^\mu$ ,  $\mu=0,1,2,3$

$$x^0=ct, x^1=x, x^2=y, x^3=z$$

Lorentz Transformations

$$t' = \gamma \left( t - \frac{v}{c^2} x \right)$$

$$x' = \gamma(x - vt)$$

$$y' = y$$

$$z' = z$$

$$x'^0 = \gamma(x^0 - \beta x^1)$$

$$x'^1 = \gamma(x^1 - \beta x^0)$$

$$x'^2 = x^2$$

$$x'^3 = x^3$$

, where  $\beta \equiv \frac{v}{c}$

$$x'^\mu = \sum_{\nu=0}^3 \Lambda_\nu^\mu x^\nu \quad (\mu = 0,1,2,3)$$

$$x'^\mu = \Lambda_\nu^\mu x^\nu$$

Lorentz Transformations:  $x'^{\mu} = \Lambda^{\mu}_{\nu} x^{\nu}$

$$\Lambda = \begin{bmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

HMW 2: Rapidity  $\phi = \tanh^{-1} \beta$

a) Show that  $\Lambda = \begin{bmatrix} \cosh \phi & -\sinh \phi & 0 & 0 \\ -\sinh \phi & \cosh \phi & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$

b) Show that  $\phi_{AC} = \phi_{AB} + \phi_{BC}$

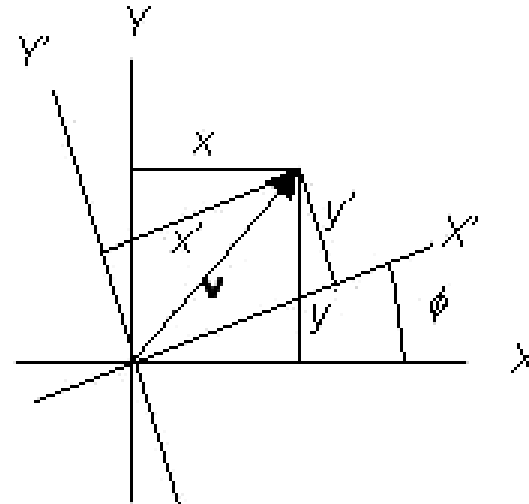
Lorentz Transformation:  $\Lambda = \begin{bmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} = \begin{bmatrix} \cosh \phi & -\sinh \phi & 0 & 0 \\ -\sinh \phi & \cosh \phi & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$

Rotation:

$$\begin{aligned} x' &= x \cdot \cos \phi + y \cdot \sin \phi \\ y' &= y \cdot \cos \phi - x \cdot \sin \phi \\ z' &= z \end{aligned}$$

$$\begin{bmatrix} x' \\ y' \\ z' \end{bmatrix} = \begin{bmatrix} \cos \phi & \sin \phi & 0 \\ -\sin \phi & \cos \phi & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \end{bmatrix}$$

$$\mathbf{R}_z(\phi) = \begin{bmatrix} \cos \phi & \sin \phi & 0 \\ -\sin \phi & \cos \phi & 0 \\ 0 & 0 & 1 \end{bmatrix}$$



$$r^2 = x^2 + y^2 + z^2 = x'^2 + y'^2 + z'^2$$

$r^2$  is invariant under R

$$I \equiv c^2 t^2 - x^2 - y^2 - z^2 = c^2 t'^2 - x'^2 - y'^2 - z'^2$$

$I$  is invariant under  $\Lambda$ : same in any inertial system

HMW 3: Verify this.

$$x^\mu = (ct, x, y, z) = (x^0, x^1, x^2, x^3)$$

$$r^2 = (x^1)^2 + (x^2)^2 + (x^3)^2 = (x'^1)^2 + (x'^2)^2 + (x'^3)^2$$

$$I = (x^0)^2 - (x^1)^2 - (x^2)^2 - (x^3)^2 = (x'^0)^2 - (x'^1)^2 - (x'^2)^2 - (x'^3)^2$$

$$I = g_{\mu\nu} x^\mu x^\nu \quad \text{where } g_{\mu\nu} \text{ are the components of the Metric: } g = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{bmatrix}$$

$$x^\mu = (x^0, x^1, x^2, x^3)$$

$$x_\mu = g_{\mu\nu} x^\nu = (x^0, -x^1, -x^2, -x^3)$$



$$I = x^\mu x_\mu$$

Given *any two* four-vectors, the scalar product

$$a^\mu b_\mu = a^0 b^0 - a^1 b^1 - a^2 b^2 - a^3 b^3 \quad \text{is invariant.}$$



$$a^\mu = (a^0, a^1, a^2, a^3)$$

$$a_\mu = (a^0, -a^1, -a^2, -a^3)$$

$$\partial_\mu \equiv \frac{\partial}{\partial x^\mu} = \left( \frac{\partial}{\partial t}, \frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z} \right) = (\partial_0, \nabla)$$

$$\partial^\mu \equiv \frac{\partial}{\partial x_\mu} = \left( \frac{\partial}{\partial t}, -\frac{\partial}{\partial x}, -\frac{\partial}{\partial y}, -\frac{\partial}{\partial z} \right) = (\partial^0, -\nabla)$$

$$\partial_\mu a^\mu = \partial^\mu a_\mu = \frac{\partial a^0}{\partial t} + \nabla \cdot \vec{a}$$

$$\begin{aligned} \square &= \partial_\mu \partial^\mu = g_{\mu\nu} \partial^\nu \partial^\mu = \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial z^2} \\ &= \frac{\partial^2}{\partial t^2} - \nabla^2 = \frac{\partial^2}{\partial t^2} - \Delta \end{aligned}$$

“Natural” Units for high energy physics:  $\hbar = 1, c = 1$

$$\hbar = 6.6 \times 10^{-25} \text{ GeV}\cdot\text{sec} = 1, \quad c = 3.0 \times 10^8 \text{ m/sec} = 1$$

$$[\text{Energy}] = [\text{Mass}] = [\text{Length}]^{-1} = [\text{Time}]^{-1}$$

HMW 4: a) Show that  $200 \text{ MeV fm} \cong 1$ , where  $1 \text{ fm} = 10^{-15} \text{ m}$ .

b) Show that mass of a proton ( $1.7 \times 10^{-27} \text{ kg}$ ) is  $\cong 1 \text{ GeV}$ .

$$(1 \text{ eV} = 1.6 \times 10^{-19} \text{ J})$$

## Relativistic Energy and Momentum

$$E = \gamma m, \quad \vec{p} = \gamma m \vec{v} \quad \text{where} \quad \gamma = \frac{1}{\sqrt{1 - v^2}}$$

$$\text{for } v \ll 1, \quad \gamma = (1 - v^2)^{-1/2} = 1 + \frac{1}{2}v^2 + \frac{3}{8}v^4 + \dots$$

$$E = \gamma m = m + \frac{1}{2}mv^2 + \dots$$

$$\text{Relativistic Kinetic Energy: } T = E - m = (\gamma - 1)m$$

$$\text{Energy-momentum four vector: } p^\mu = (E, p_x, p_y, p_z)$$

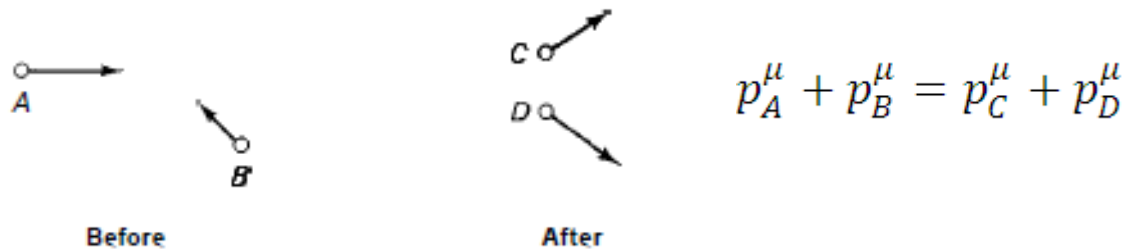
$$\text{invariant: } p^\mu p_\mu = E^2 - \vec{p} \cdot \vec{p} = \gamma^2 m^2 (1 - v^2) = m^2 \quad \longrightarrow \quad E^2 = |\vec{p}|^2 + m^2$$

$$\text{Massless particle: } p^\mu p_\mu = 0, \quad E = |\vec{p}|, \quad v = 1$$

(for a massless particle,  $E = hv$ )

## Relativistic Collisions

Total energy and momentum is conserved.



If collision is elastic,  $T$  and therefore  $m$  is also conserved ( $E = T + m$ )

Conserved: Same before and after the collision.

Invariant: Same in any inertial reference frame.

$p_A^\mu + p_B^\mu$  conserved: Yes, invariant: No

$p_A^\mu p_{A\mu}$  conserved: No, invariant: Yes

$(p_A^\mu + p_B^\mu)(p_{A\mu} + p_{B\mu})$  conserved: Yes, invariant: Yes

Ex. A pion at rest decays into muon and a neutrino ( $\pi^- \rightarrow \mu^- + \bar{\nu}_\mu$ )

On the average, how far would the muon travel in vacuum before decaying?

$$d = v\tau = \gamma v\tau_0$$



Energy-momentum conservation:  $p_\pi = p_\mu + p_\nu$

Solve for  $\mu$  and square both sides:  $p_\mu = p_\pi - p_\nu$

$$m_\mu^2 = m_\pi^2 - 2p_\pi \cdot p_\nu = m_\pi^2 - 2m_\pi E_\nu$$

$$p_\pi = (m_\pi, 0, 0, 0)$$

$E_\nu = |\vec{p}_\nu| = |\vec{p}_\mu|$  since  $\nu$  is massless and total  $\mathbf{p} = 0$

$$|\vec{p}_\mu| = \frac{m_\pi^2 - m_\mu^2}{2m_\pi} = \gamma v m_\mu \quad \longrightarrow \quad d = \gamma v \tau_0 = \frac{m_\pi^2 - m_\mu^2}{2m_\pi m_\mu} \tau_0 = 186 \text{ m}$$

Ex. A proton strikes another proton at rest, creating a proton-antiproton pair.

What is the threshold energy for this reaction? ( $p + p \rightarrow p + p + p + \bar{p}$ )



before, LAB:  $p_{\text{total}}^{\mu} = (E + m, |\vec{p}|, 0, 0)$

after, CM:  $p_{\text{total}}^{\prime\mu} = (4m, 0, 0, 0)$

$p_{\text{total}}^{\mu} p_{\mu\text{total}}$  is both conserved and invariant:  $(E + m)^2 - |\vec{p}|^2 = 16m^2$

$$E = 7m$$

Fixed target: threshold  $T = 6m$

Colliding beams: threshold  $T = m$

HMW 5: In a two-body scattering event,  $A + B \rightarrow C + D$ , it is convenient to introduce the *Mandelstam variables*

$$s \equiv (p_A + p_B)^2/c^2$$

$$t \equiv (p_A - p_C)^2/c^2$$

$$u \equiv (p_A - p_D)^2/c^2$$

The *theoretical* virtue of the Mandelstam variables is that they are Lorentz invariants, with the same value in any inertial system. *Experimentally*, though, the more accessible parameters are energies and scattering angles.

- a) Show that  $s + t + u = m_A^2 + m_B^2 + m_C^2 + m_D^2$
- b) Show that  $E_{\text{TOT}}^{\text{CM}} = \sqrt{s}$

Single particle wave equations:

Schrödinger equation for a free particle

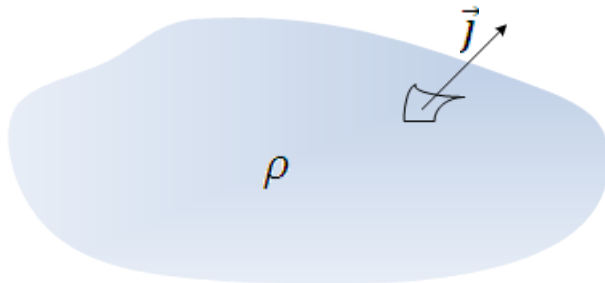
non-relativistic energy-momentum relation:  $E = \frac{|\vec{p}|^2}{2m}$

$$\vec{p} \rightarrow -i\hbar\nabla, \quad E \rightarrow i\hbar\frac{\partial}{\partial t} \quad \longrightarrow \quad i\hbar\frac{\partial\phi}{\partial t} = -\frac{\hbar^2}{2m}\nabla^2\phi$$

probability density:  $\rho = \phi^*\phi$

probability current:  $\vec{j} = -\frac{i\hbar}{2m}(\phi^*\nabla\phi - \phi\nabla\phi^*)$

$$\left. \begin{array}{l} \text{probability density: } \rho = \phi^*\phi \\ \text{probability current: } \vec{j} = -\frac{i\hbar}{2m}(\phi^*\nabla\phi - \phi\nabla\phi^*) \end{array} \right\} \frac{\partial\rho}{\partial t} + \nabla \cdot \vec{j} = 0$$



$$\int_V \frac{\partial\rho}{\partial t} dV + \int_V (\nabla \cdot \vec{j}) dV = 0$$

$$\int_V \frac{\partial\rho}{\partial t} dV + \int_S \vec{j} \cdot d\vec{A} = 0$$



## Klein-Gordon equation

relativistic energy-momentum relation:  $p^\mu p_\mu = E^2 - |\vec{p}|^2 = m^2$

$$\vec{p} \rightarrow -i\hbar\nabla, \quad E \rightarrow i\hbar\frac{\partial}{\partial t} \quad \longrightarrow \quad \left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \nabla^2 \right) \phi + \frac{m^2 c^2}{\hbar^2} \phi = 0$$

$$p^\mu \rightarrow i\partial^\mu \quad \longrightarrow \quad \boxed{(\partial^\mu \partial_\mu + m^2)\phi = 0} \quad \text{or} \quad \boxed{(\square + m^2)\phi = 0}$$

probability density and current:  $j^\mu = (\rho, \vec{j}) = \frac{i\hbar}{2m} (\phi^* \partial^\mu \phi - \phi \partial^\mu \phi^*)$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \vec{j} = \partial_\mu j^\mu = 0$$

$$\rho = \frac{i\hbar}{2m} \left( \phi^* \frac{\partial \phi}{\partial t} - \phi \frac{\partial \phi^*}{\partial t} \right) : \text{negative probability?}$$

Klein-Gordon equation not a single particle wave equation

it is the equation for a *quantum field* whose quanta are scalar (spin 0) particles:

$|\phi|^2 \propto$  number of particles

Dirac Equation: • first order  
• wave eq. for spin 1/2 particles

$$(i\gamma^\mu \partial_\mu - m)\psi = 0$$

$\psi$ : Dirac spinor

$$[-(\gamma^\mu \partial_\mu)(\gamma^\nu \partial_\nu) - i(\gamma^\mu \partial_\mu)m]\psi = 0,$$

$$(\gamma^\mu \gamma^\nu \partial_\mu \partial_\nu + m^2)\psi = 0.$$

Since  $\partial_\mu \partial_\nu = \partial_\nu \partial_\mu$ ,  $\gamma^\mu \gamma^\nu \rightarrow \frac{1}{2}(\gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu) \equiv \frac{1}{2}\{\gamma^\mu, \gamma^\nu\}$

$$\frac{1}{2}\{\gamma^\mu, \gamma^\nu\} \partial_\mu \partial_\nu \psi + m^2 \psi = 0.$$

$$p^\mu p_\mu = E^2 - |\vec{p}|^2 = m^2, \quad p^\mu \rightarrow i\partial^\mu, \quad (\partial^\mu \partial_\mu + m^2)\psi = 0$$

$$(g^{\mu\nu} \partial_\mu \partial_\nu + m^2)\psi = 0$$

$$\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu}$$

$$(\gamma^0)^2 = 1, \quad (\gamma^i)^2 = -1, \quad \gamma^\mu \gamma^\nu = -\gamma^\nu \gamma^\mu \quad (\nu \neq \mu).$$

## Dirac Matrices

$$\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu}$$

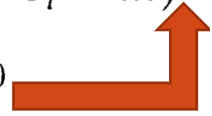
$$(\gamma^0)^2 = 1, \quad (\gamma^i)^2 = -1, \quad \gamma^\mu \gamma^\nu = -\gamma^\nu \gamma^\mu \quad (\nu \neq \mu).$$

$$\gamma^0 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \gamma^i = \begin{pmatrix} 0 & -\sigma^i \\ \sigma^i & 0 \end{pmatrix}$$

$$\gamma^{0\dagger} = \gamma^0, \quad \gamma^{i\dagger} = -\gamma^i \quad \sigma_x = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \sigma_y = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma_z = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

Probability current  $j^\mu$  for Dirac equation:

$$(i\gamma^\mu \partial_\mu - m)\psi = 0 \quad \longrightarrow \quad \psi^\dagger (-i\gamma^0 \overleftarrow{\partial}_0 + i\gamma^i \overleftarrow{\partial}_i - m) = 0.$$

multiply by  $\gamma^0$  

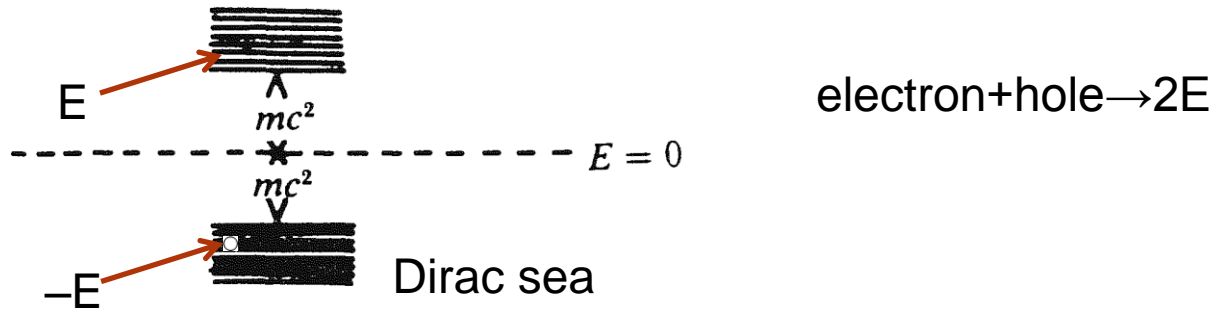
$$\bar{\psi}(i\gamma^\mu \overleftarrow{\partial}_\mu + m) = 0 \quad \text{where} \quad \bar{\psi} = \psi^\dagger \gamma^0 \quad (\text{adjoint spinor})$$

$$j^\mu = \bar{\psi} \gamma^\mu \psi \quad \longrightarrow \quad \partial_\mu j^\mu = 0$$

$$j^0 = \bar{\psi} \gamma^0 \psi = \psi^\dagger \psi = |\psi_1|^2 + |\psi_2|^2 + |\psi_3|^2 + |\psi_4|^2 \quad \longrightarrow \quad \text{no negative probabilities..}$$

$$p^\mu p_\mu = E^2 - |\vec{p}|^2 = m^2$$

$$E = \mp \sqrt{|\vec{p}|^2 + m^2} \quad \rightarrow \text{Negative energies?}$$



LH hole with  $-E$  in Dirac sea corresponds to a RH **antiparticle** (positron) with  $E$

modern (QFT) interpretation:

no Dirac sea required,

negative energy problem solved upon “second” quantization:  $[ , ] \rightarrow \{ , \}$

Dirac’s prediction of antiparticles remains valid

(not as holes but as CPT conjugates)