

Motivations for neutrino accelerator experiments

ν_e



ν_μ



ν_τ



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History of Neutrino Discoveries

- Atmospheric ν_μ are converted to ν_τ (SK) (98)
- Solar ν_e are converted to either ν_μ or ν_τ (SNO) (02)
- Only the LMA solution left for solar neutrinos (Homestake+Gallium+SK+SNO) (02)
- Reactor anti- ν_e disappear/reappear (KamLAND) (04)
- Accelerator ν_μ disappear (K2K 04 , MINOS 06)
- OPERA sees first tau appearance event (10)
- Message: many discoveries in neutrino physics

■ Impact of Neutrino Discoveries

- Lepton Flavor is not conserved
- Neutrinos have tiny masses, not very hierarchical
- Neutrinos mix a lot
- At least 7 new parameters for SM
- Quite unlike quark mass and mixing
- Of all fermions, neutrinos are least understood
- First new physics beyond the SM
- Message: most important discovery of last 20 years

■ Why Beyond Standard Model?

1. There are no right-handed neutrinos ν_R
 2. There are only Higgs doublets of $SU(2)_L$
 3. There are only renormalizable terms
- } Standard Model

In the **Standard Model** these conditions all apply so neutrinos are **massless**, with ν_e , ν_μ , ν_τ distinguished by separate lepton numbers L_e , L_μ , L_τ

Neutrinos and anti-neutrinos are distinguished by the total conserved lepton number $L=L_e+L_\mu+L_\tau$

To generate neutrino mass we must relax 1 and/or 2 and/or 3

Message: Neutrino Mass and Mixing is first physics BSM

■ Implications for PP and Cosmology

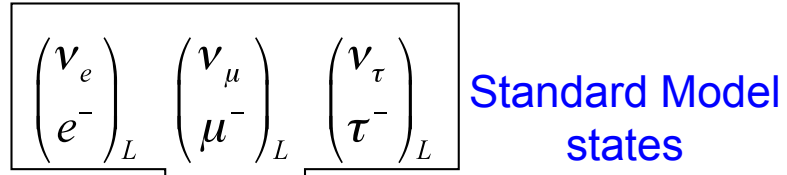
- Origin of tiny neutrino mass
Extra dimensions, See-saw mechanism, SUSY
- Unification of matter, forces and flavour
GUTs, Family Symmetry,...
- Did neutrinos play a role in our existence?
Leptogenesis
- Did neutrinos play a role in forming galaxies?
Hot/Warm Dark matter component
- Did neutrinos play a role in birth of the universe?
Sneutrino inflation
- Can neutrinos shed light on dark energy? $\Lambda \sim m_\nu^4$

Particle
Physics

Cosmology

Message: many interesting implications for PP and Cosmology

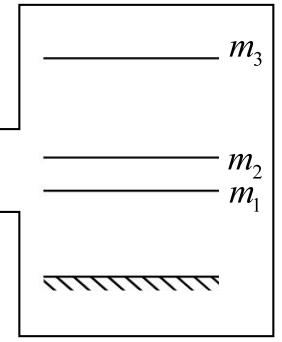
Three Neutrino Mass and Mixing



Pontecorvo
Maki
Nakagawa
Sakata

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Neutrino mass states



$$U_{PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$s_{ij} = \sin \theta_{ij}$$

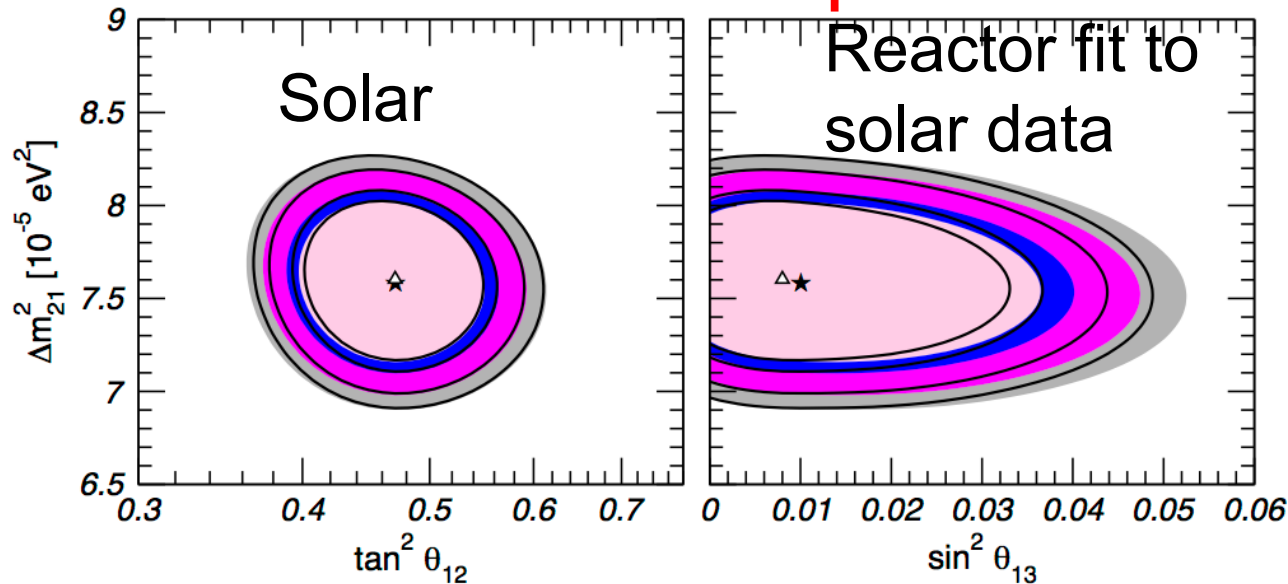
$$c_{ij} = \cos \theta_{ij}$$



Oscillation phase δ
Majorana phases α_1, α_2

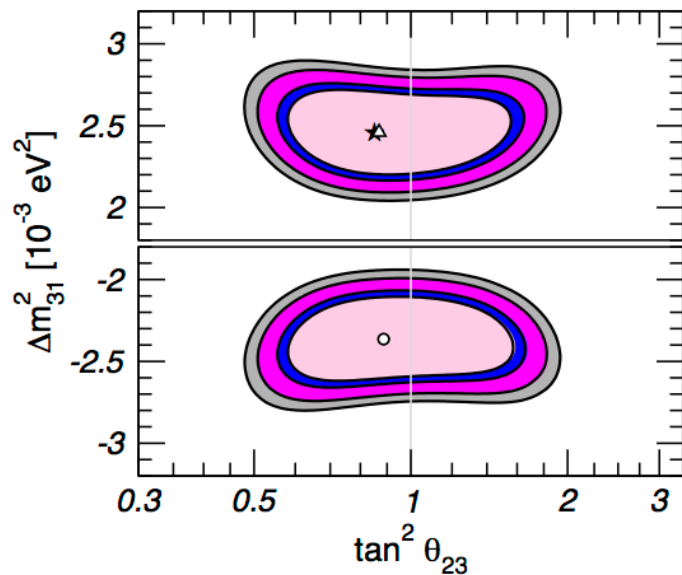
3 masses + 3 angles + 1(3) phase(s)
= 7(9) new parameters for SM

Global Fit to Atmospheric and Solar Data



Gonzalez-Garcia,
Maltoni, Salvado

arXiv:1001.4524



$$\Delta m_{21}^2 = 7.59 \pm 0.20 \left(\begin{smallmatrix} +0.61 \\ -0.69 \end{smallmatrix} \right) \times 10^{-5} \text{ eV}^2$$

$$\Delta m_{31}^2 = \begin{cases} -2.36 \pm 0.11 (\pm 0.37) \times 10^{-3} \text{ eV}^2 \\ +2.46 \pm 0.12 (\pm 0.37) \times 10^{-3} \text{ eV}^2 \end{cases}$$

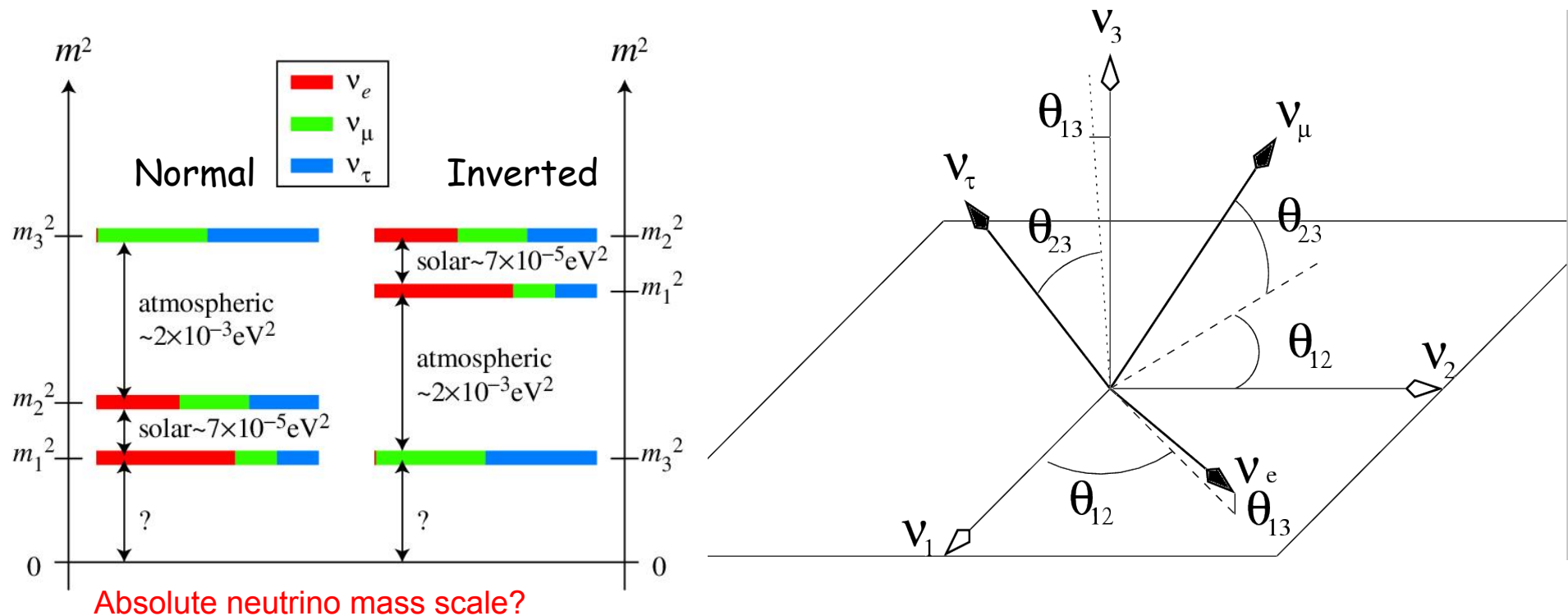
$$\theta_{12} = 34.4 \pm 1.0 \left(\begin{smallmatrix} +3.2 \\ -2.9 \end{smallmatrix} \right)^\circ$$

$$\theta_{23} = 42.8 \begin{smallmatrix} +4.7 \\ -2.9 \end{smallmatrix} \left(\begin{smallmatrix} +10.7 \\ -7.3 \end{smallmatrix} \right)^\circ$$

$$\theta_{13} = 5.6 \begin{smallmatrix} +3.0 \\ -2.7 \end{smallmatrix} (\leq 12.5)^\circ$$

Atmospheric

Three Neutrino Mass and Mixing at a glance



Message: the neutrino mass ordering and mass scale are not yet measured, nor is the reactor mixing angle or CP phase

■ Tri-bimaximal Mixing

$$U_{TB} = \begin{pmatrix} \frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} \quad \text{Harrison, Perkins, Scott}$$

Including
deviations from
TB mixing

$$\theta_{13} = \Delta_{13}^{TB}, \quad \theta_{12} = 35^\circ + \Delta_{12}^{TB}, \quad \theta_{23} = 45^\circ + \Delta_{23}^{TB}.$$

c.f. data

$$\theta_{13} = 5.6_{-2.7}^{+3.0} \quad \theta_{12} = 34.4 \pm 1.0 \quad \theta_{23} = 42.8_{-2.9}^{+4.7}$$

Message: neutrino mixing is consistent with the remarkably simple TB mixing pattern, hinting at an underlying symmetry

■ Tri-bimaximal Parametrisation

SFK arXiv:0710.0530

$$s_{13} = \frac{r}{\sqrt{2}}, \quad s_{12} = \frac{1}{\sqrt{3}}(1 + s), \quad s_{23} = \frac{1}{\sqrt{2}}(1 + a)$$

$$0.07 < r < 0.21, \quad -0.05 < s < 0.003, \quad -0.09 < a < 0.04$$

r = reactor

s = solar

a = atmospheric

$$U \approx \begin{pmatrix} \sqrt{\frac{2}{3}}(1 - \frac{1}{2}s) & \frac{1}{\sqrt{3}}(1 + s) & \frac{1}{\sqrt{2}}re^{-i\delta} \\ -\frac{1}{\sqrt{6}}(1 + s - a + re^{i\delta}) & \frac{1}{\sqrt{3}}(1 - \frac{1}{2}s - a - \frac{1}{2}re^{i\delta}) & \frac{1}{\sqrt{2}}(1 + a) \\ \frac{1}{\sqrt{6}}(1 + s + a - re^{i\delta}) & -\frac{1}{\sqrt{3}}(1 - \frac{1}{2}s + a + \frac{1}{2}re^{i\delta}) & \frac{1}{\sqrt{2}}(1 - a) \end{pmatrix}$$

Present data is consistent with $r,s,a=0 \rightarrow$ tri-bimaximal
so need to measure r,s,a,δ

Oscillation formulae in terms of r, s, a

$$P_{\alpha\beta} = P(\nu_\alpha \rightarrow \nu_\beta)$$

$$\Delta_{ij} = 1.27 \Delta m_{ij}^2 L / E$$

reactor $\left\{ P_{ee} = 1 - 2r^2 \sin^2 \Delta_{31} - \frac{8}{9} \Delta_{21}^2 \right.$ Only sensitive to the reactor parameter r

L
B
L $\left\{ P_{\mu e} = r^2 \sin^2 \Delta_{31} + \frac{4}{9} \Delta_{21}^2 + \frac{4}{3} r \Delta_{21} \sin \Delta_{31} \cos(\Delta_{31} + \delta) \right.$

$\left. P_{\mu\mu} = 1 - (1 - 4a^2) \sin^2 \Delta_{31} - \frac{2}{9} (1 + 3 \cos 2\Delta_{31}) \Delta_{21}^2 \right.$

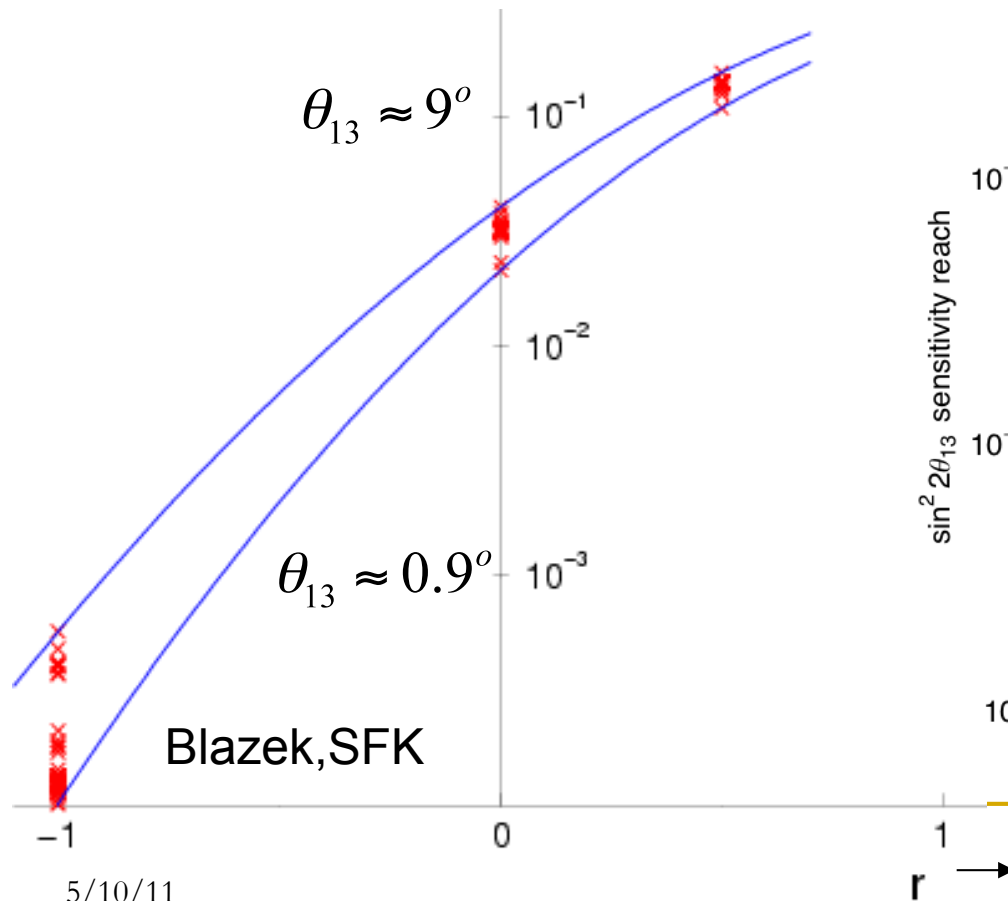
$\left. + \frac{2}{3} (1 - s - r \cos \delta) \Delta_{21} \sin 2\Delta_{31}. \right.$ Sensitive to r, s, a

For a list of formulae including matter effects see SFK [arXiv:0710.0530](https://arxiv.org/abs/0710.0530)

Reactor Angle: Theory vs Experiment

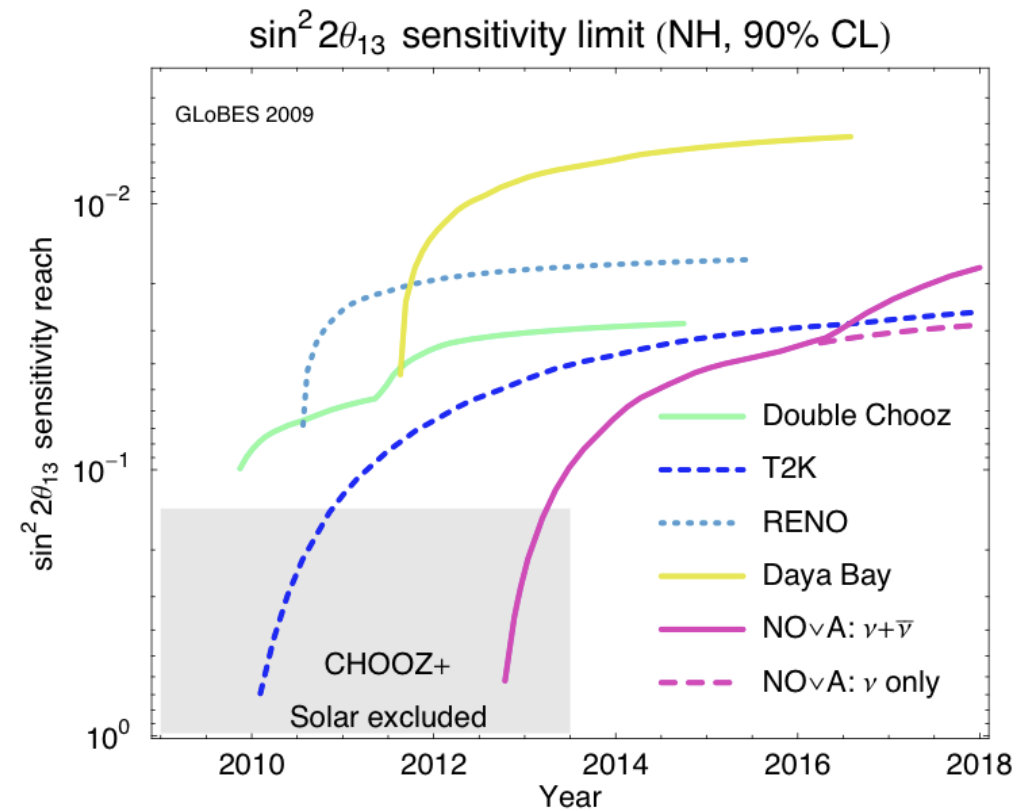
$\sin^2 2\theta_{13}$

Reactor angle in Abelian Family Symmetry Models



5/10/11

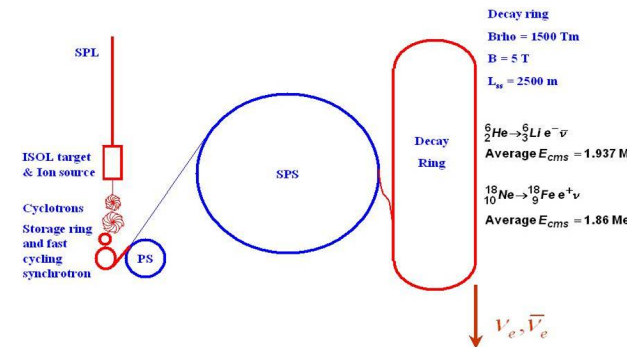
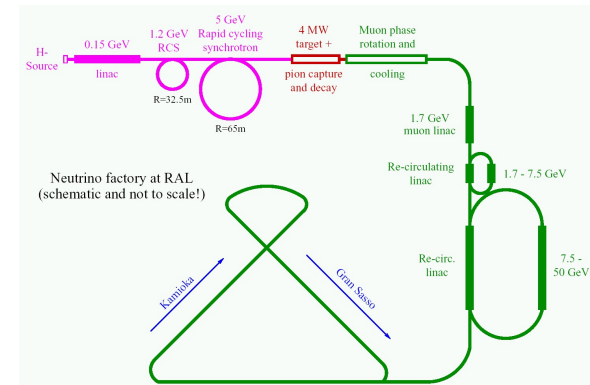
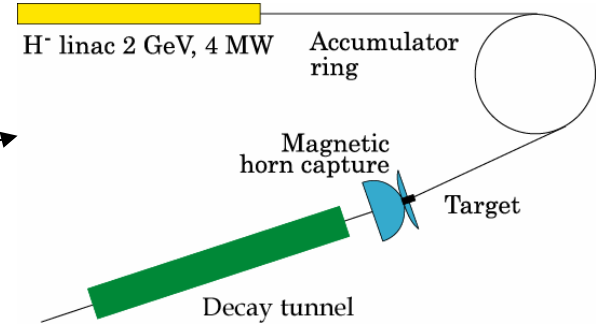
vs Experimental Prospects



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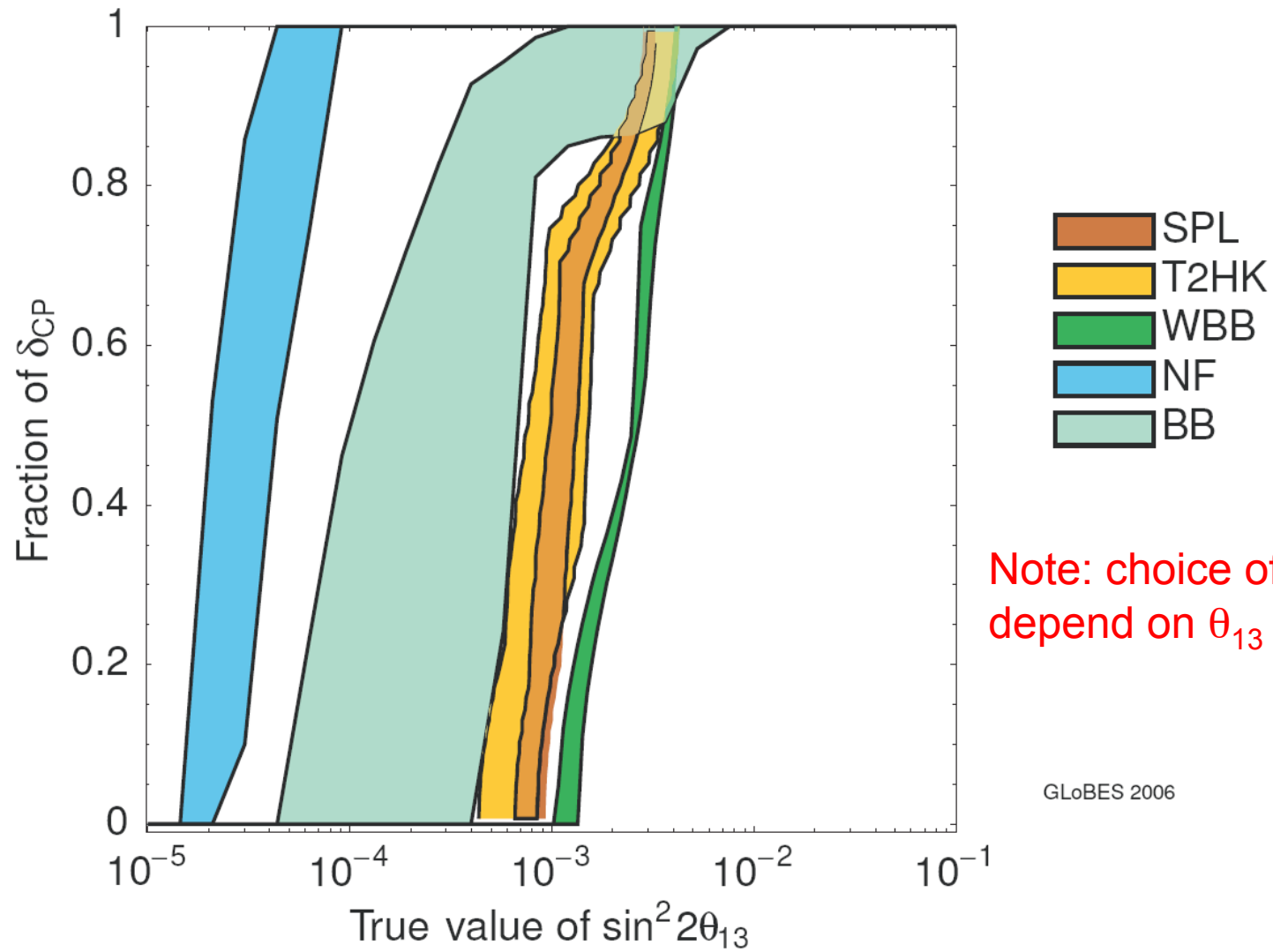
Future LBL Options:

- Second generation super-beam: CERN, LBNE (WBB), T2HK
- Neutrino Factory
- Beta-beam



Prospects to measure θ_{13}

Physics at a future Neutrino Factory and super-beam facility. By ISS Physics Working Group
Rept.Prog.Phys.72:106201,2009
arXiv:0710.4947 [hep-ph]



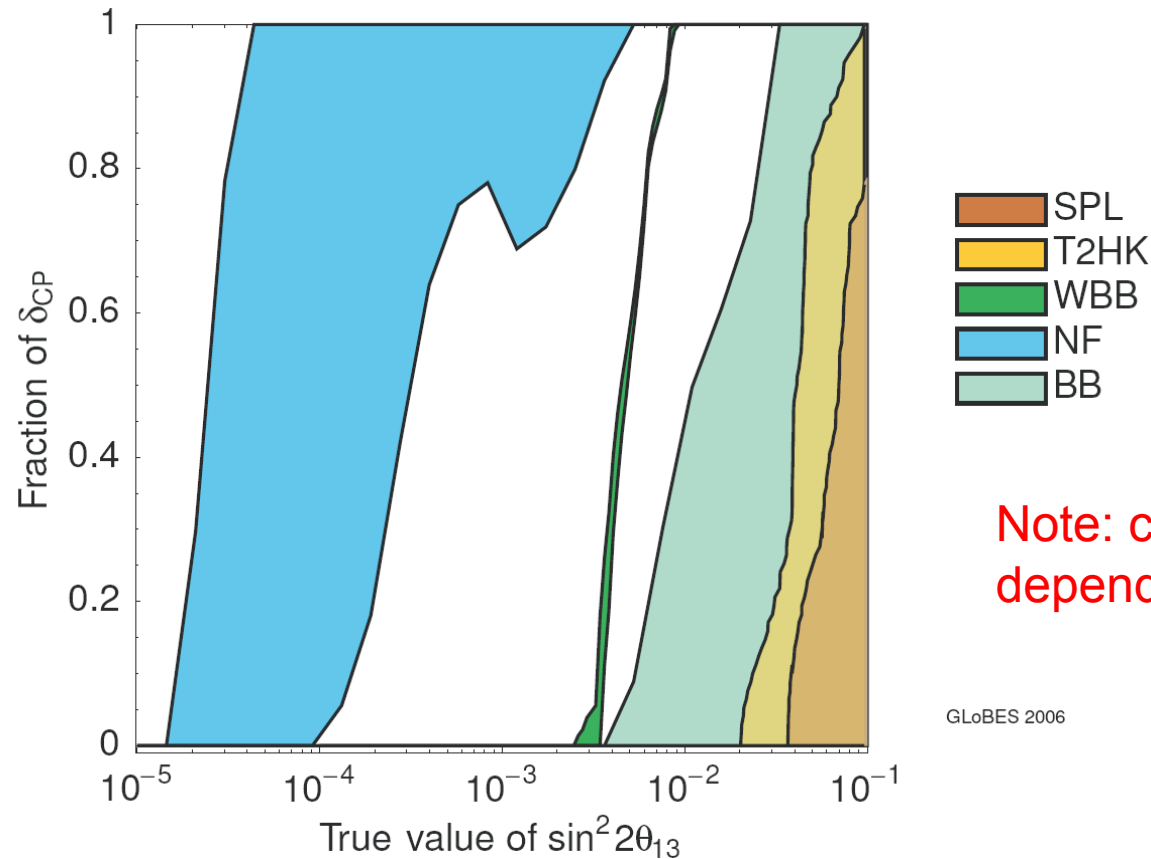
Note: choice of facility will depend on θ_{13}

GLOBES 2006

Prospects to measure the pattern of ν masses

Sensitivity to the Sign of Δm_{23}^2

Physics at a future Neutrino Factory and super-beam facility. By ISS Physics Working Group
Rept.Prog.Phys.72:106201,2009
arXiv:0710.4947 [hep-ph]

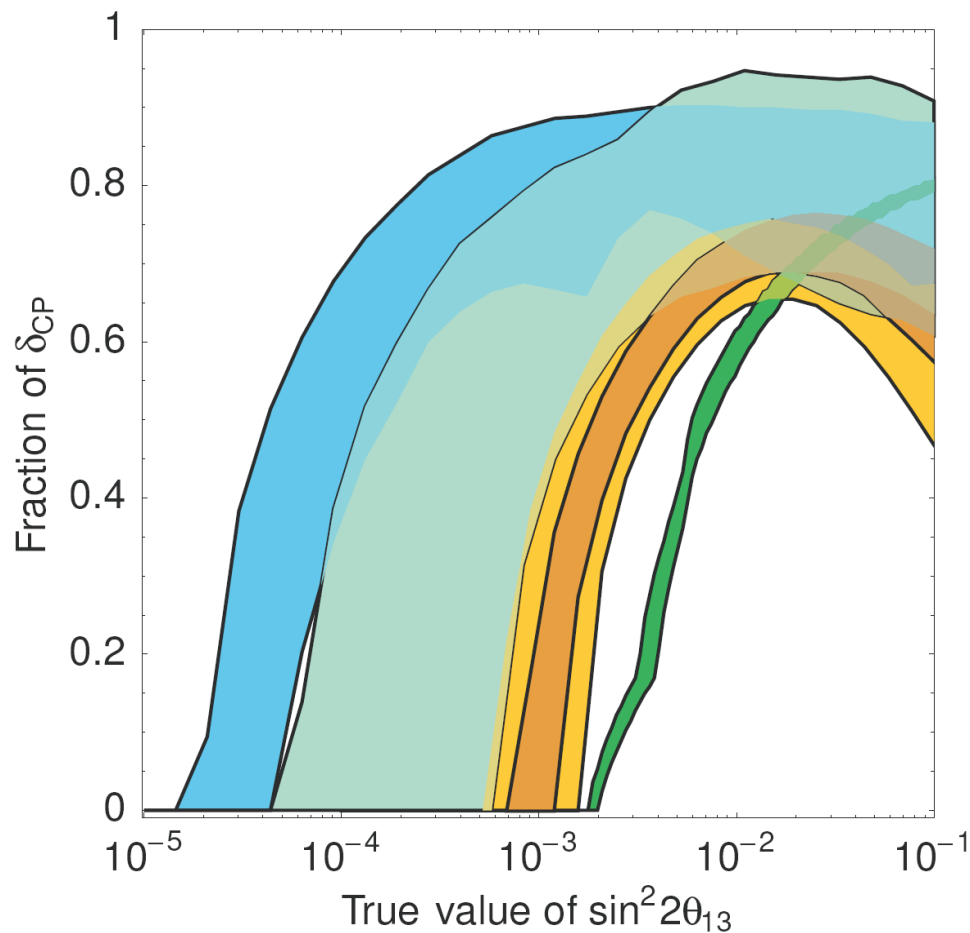


Note: choice of facility will depend on θ_{13}

GLoBES 2006

Prospects to measure CP Violation

Physics at a future Neutrino Factory and super-beam facility. By ISS Physics Working Group
Rept.Prog.Phys.72:106201,2009
arXiv:0710.4947 [hep-ph]

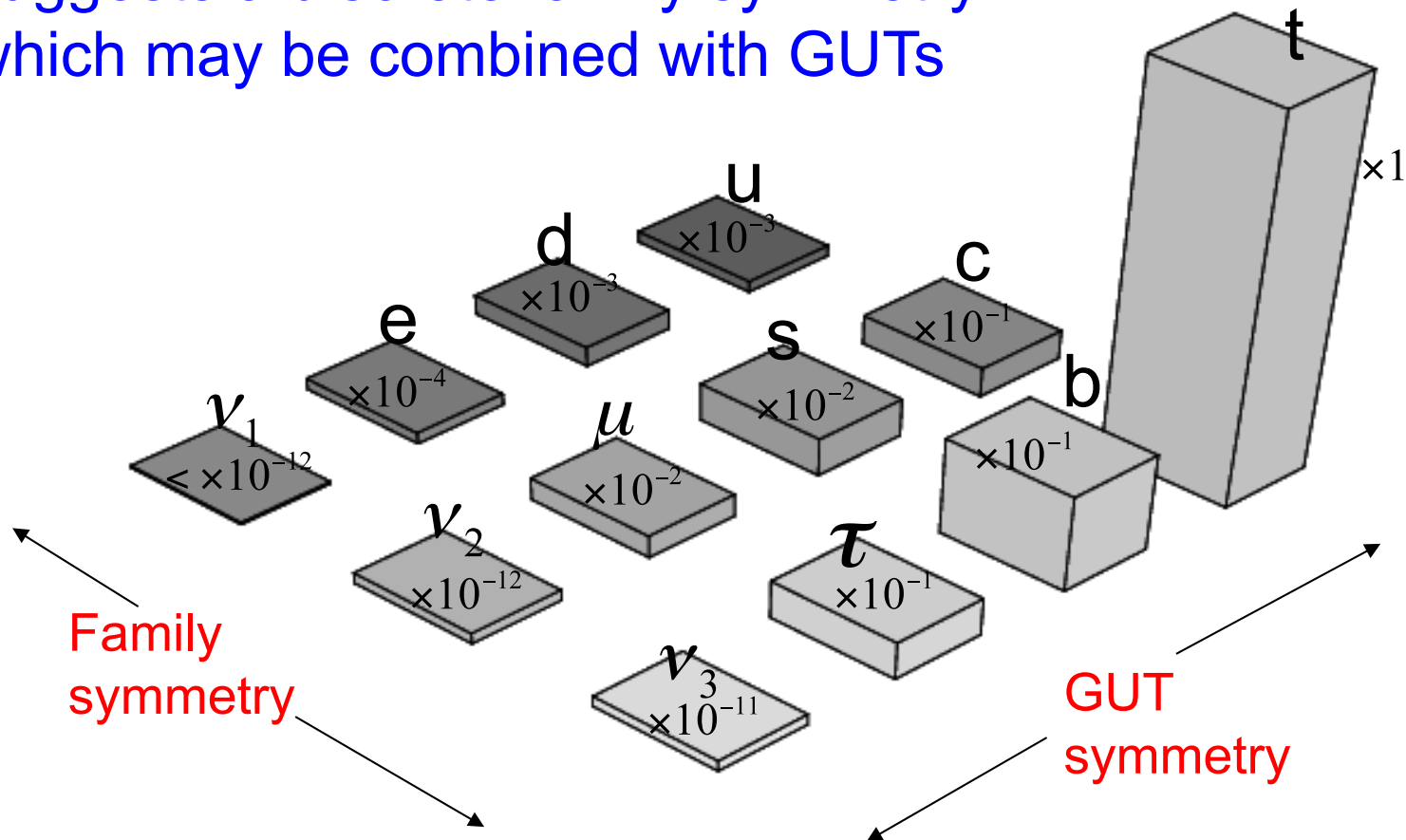


Note: choice of facility will depend on θ_{13}

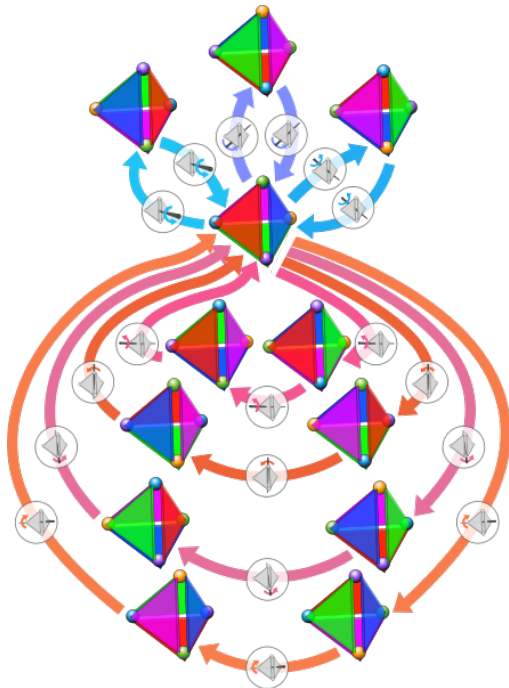
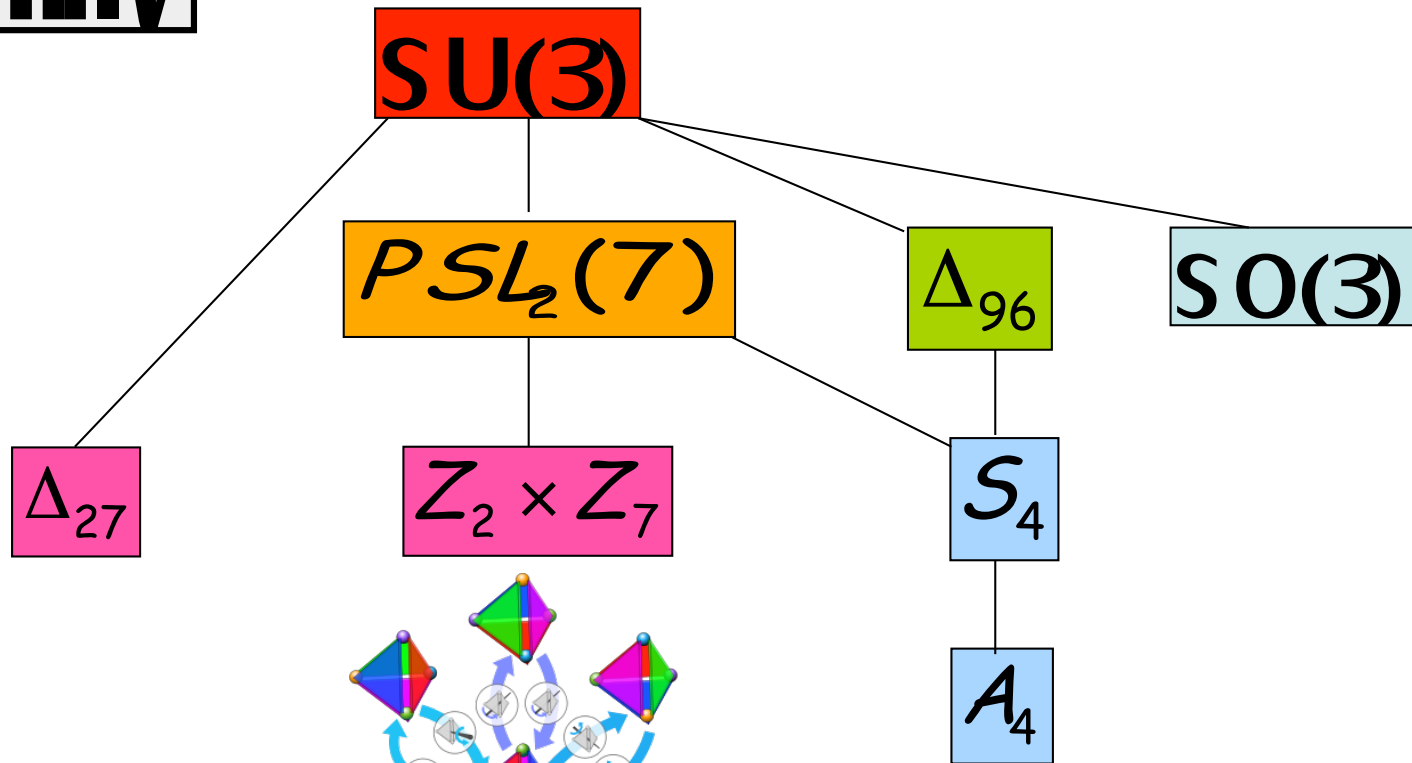
GLoBES 2006

Tri-bimaximal Mixing × GUTs

Message: Tri-bimaximal mixing suggests a discrete family symmetry which may be combined with GUTs

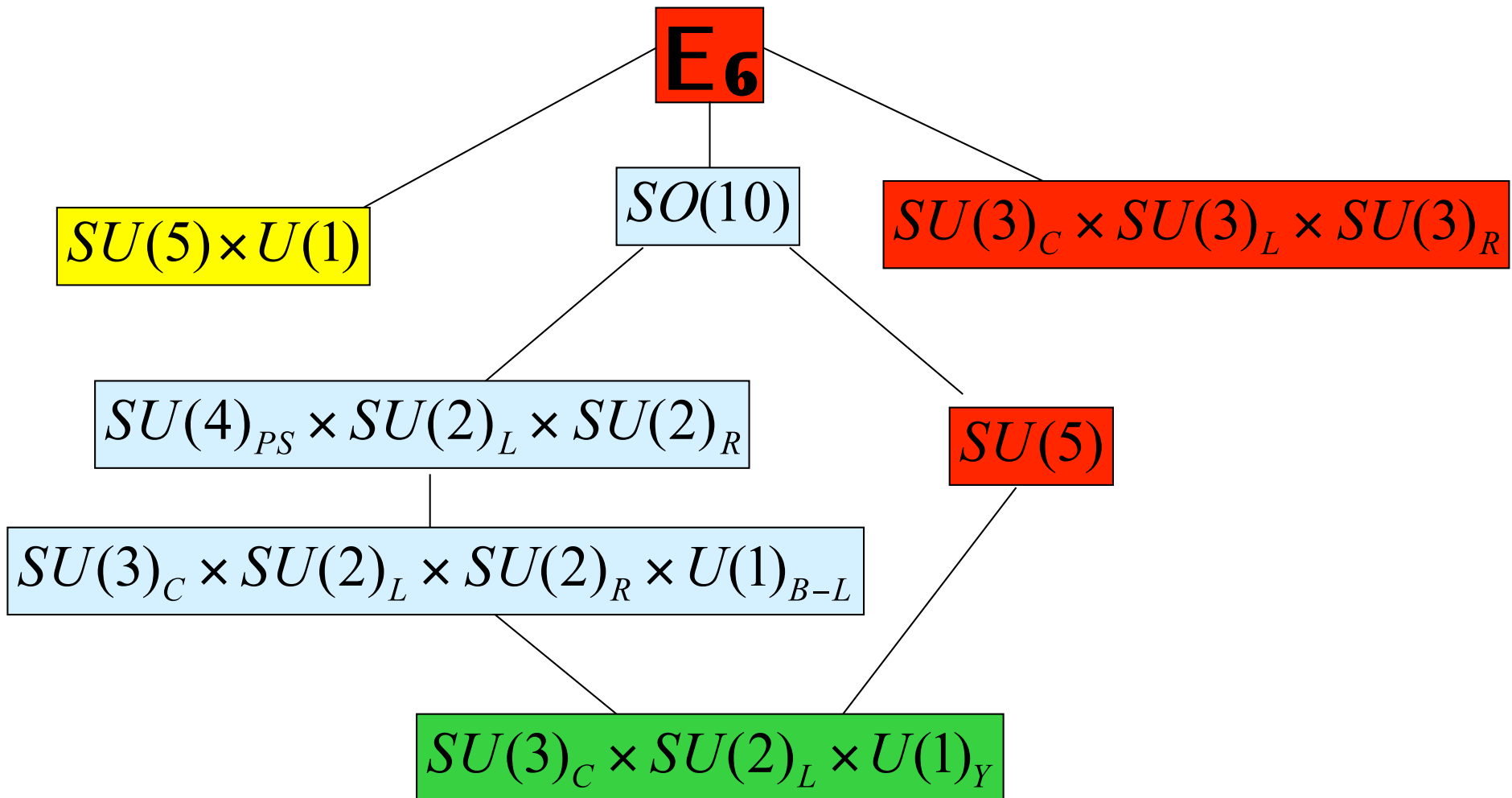


GFamily

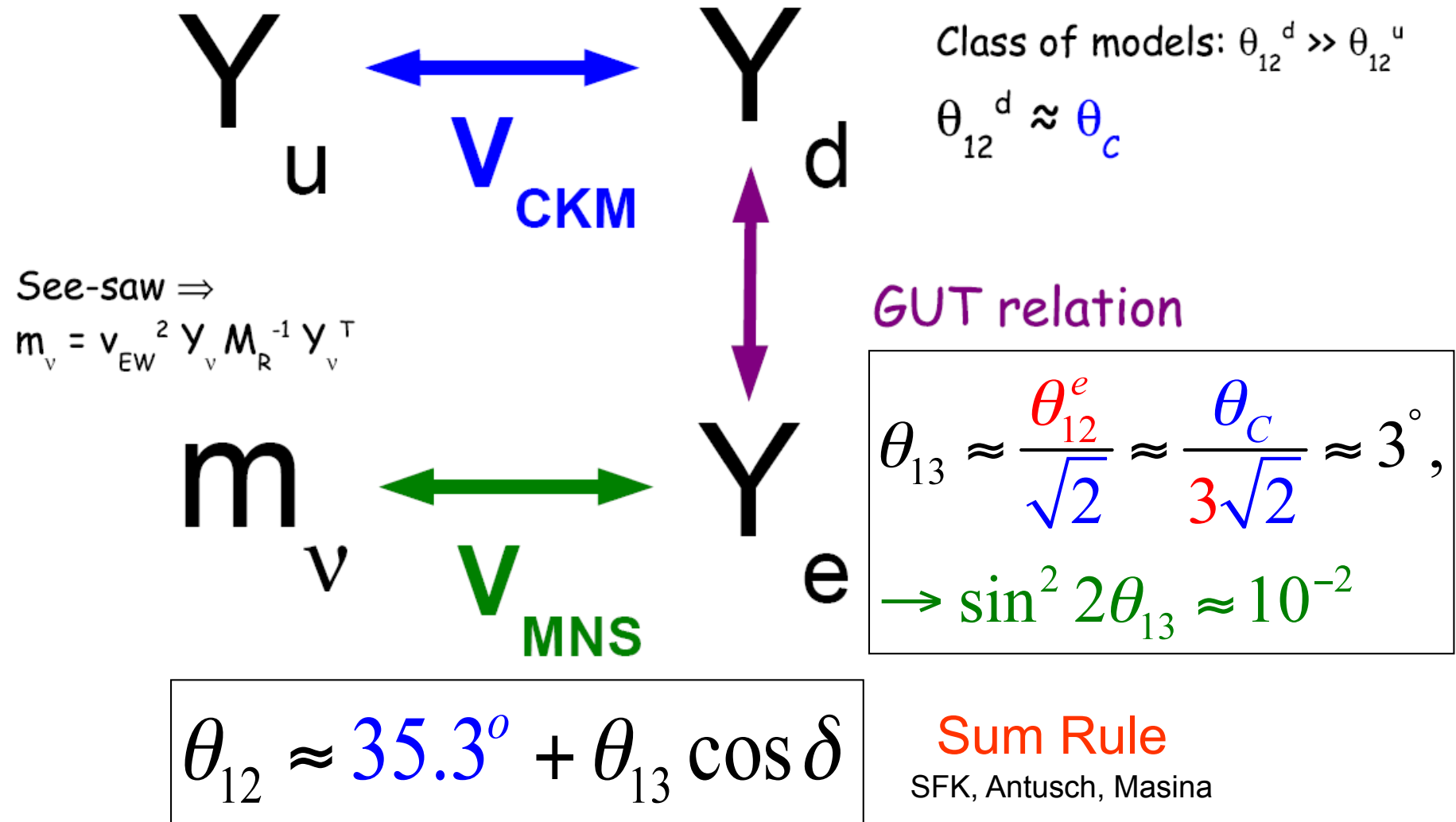


e.g. A_4 is the symmetry of the tetrahedron

GUT

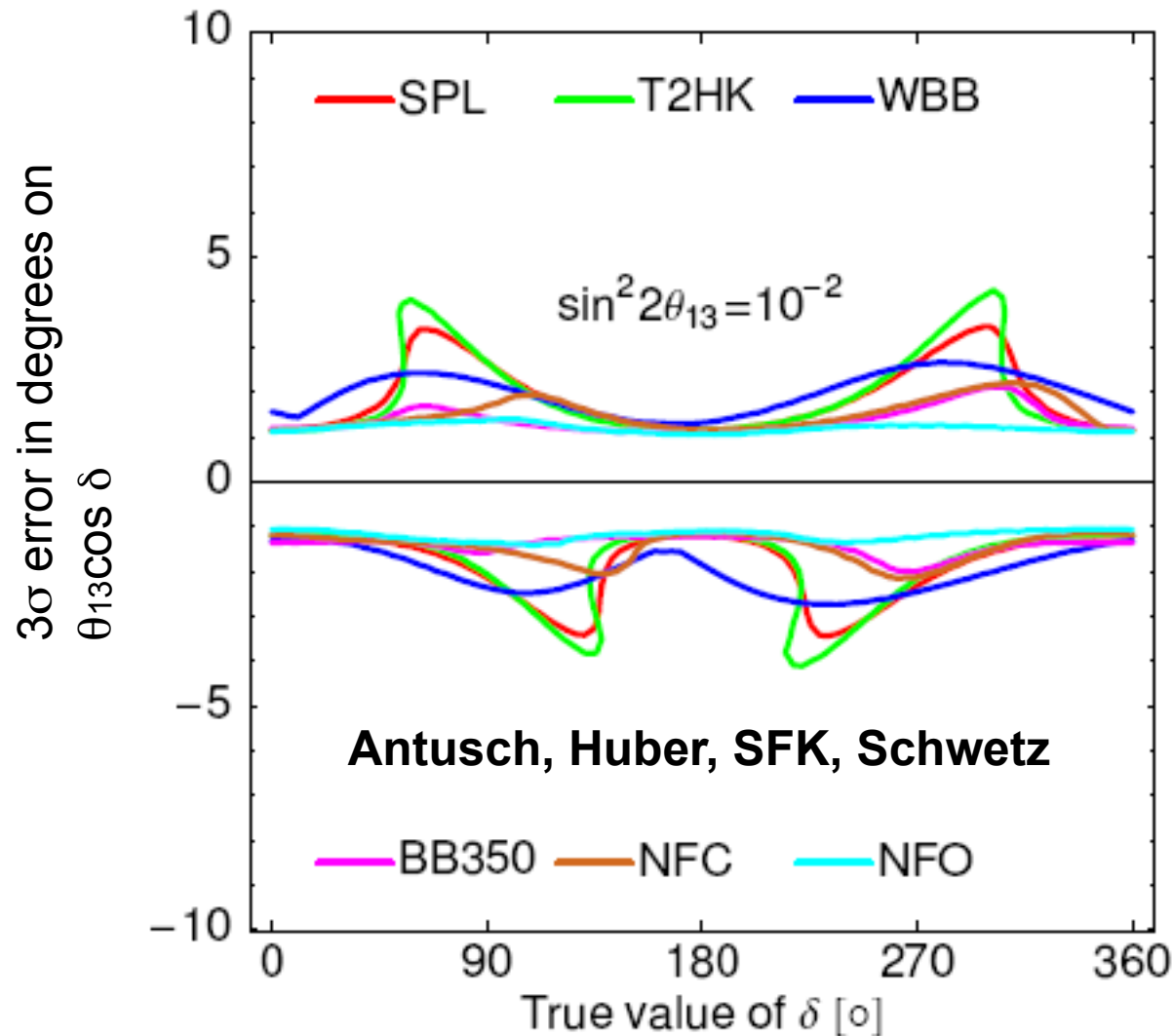


TB Mixing × GUT Predictions



Message: TB mixing can never be exact in GUT models

■ Prospects to measure sum rule combination $\theta_{13} \cos \delta$



■ Summary

- We have witnessed a revolution in neutrino physics
- Yet still do not understand origin of neutrino mass and mixing
- Current data consistent with tri-bimaximal mixing
- Realistic models predict deviations from tri-bimaximal mixing

Benchmark Model	θ_{13}	$ \theta_{23} - 45^\circ $	$ \theta_{12} - 35^\circ $	δ
TBM \otimes GUT [9]	$\frac{\theta_C}{3\sqrt{2}} = 3^\circ$	$\leq 1^\circ$	$\leq 1^\circ$	$90^\circ, 270^\circ$
QLC [10]	$\theta_C = 13^\circ$	$\leq 1^\circ$	large	180°
Abelian [4]	Fig.7	large	large	any

- In order to discriminate between these models to test GUTs of Flavour against Abelian Models or QLC...
- Need to measure the deviations of the reactor, atmospheric and solar angles from their TB values, and also measure δ