

# Fuzzy Dark Matter: Dead or Alive?!?

Boris Zupancic<sup>1,2</sup> — Supervised by Dr. Lawrence Widrow

<sup>1</sup>Queen's University <sup>2</sup>McDonald Institute

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# Cosmology and Lambda Cold Dark Matter

In Cosmology, the standard model is  $\Lambda$ CDM.

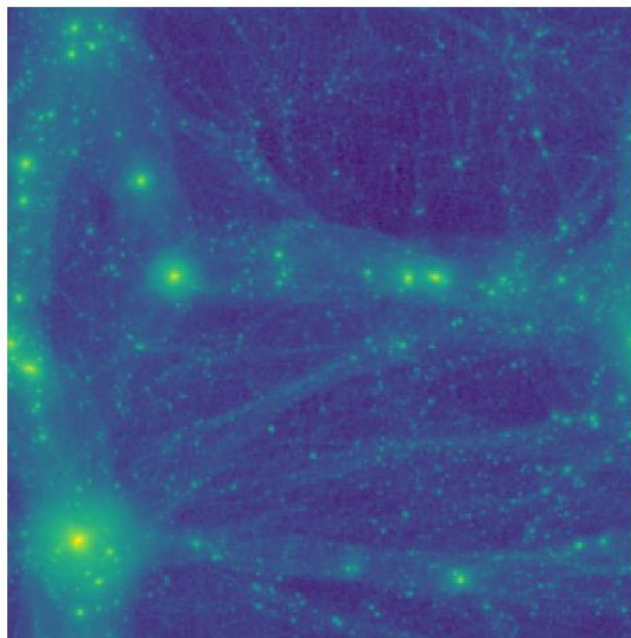


Figure: Simon May and Volker Springel (2021)

Simulations of CDM, both N-Body and fluid, suggest that it's a good theory, on scales (much) bigger than galaxies.

# Problems with CDM

Some small-scale structures (in galaxies) characteristic to CDM are not observed (Hui et al. [2017]).

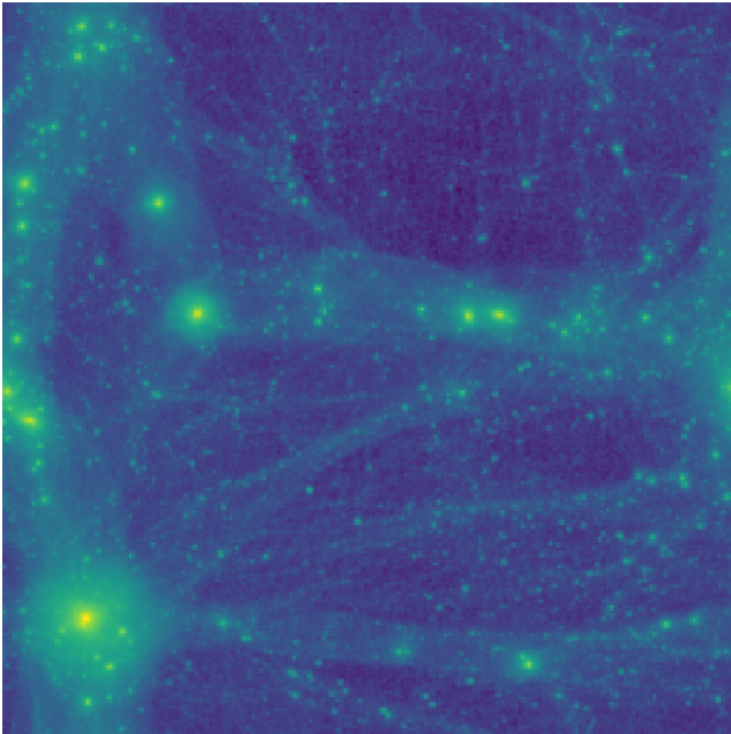


Figure: May and Springel [2021]

**“Cusp-core”** Problem  
(Weinberg et al. [2015]):

Predict cuspy galaxies, but observe more core-like density profiles.

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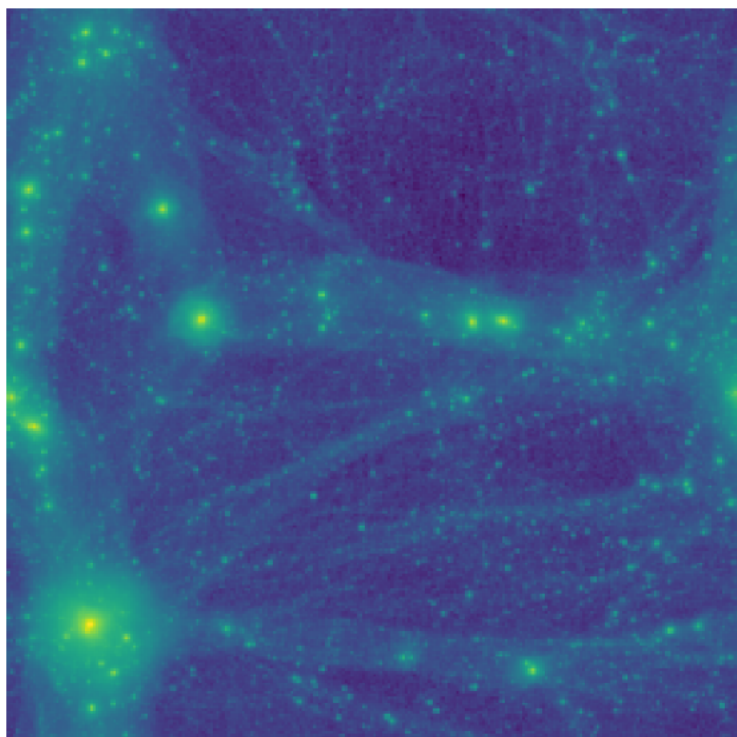


Figure: May and Springel [2021]

**“Missing satellites”**  
(Weinberg et al. [2015]):

Predict thousands of sub-halos, but we observe very few.

# We Want Something Smoother

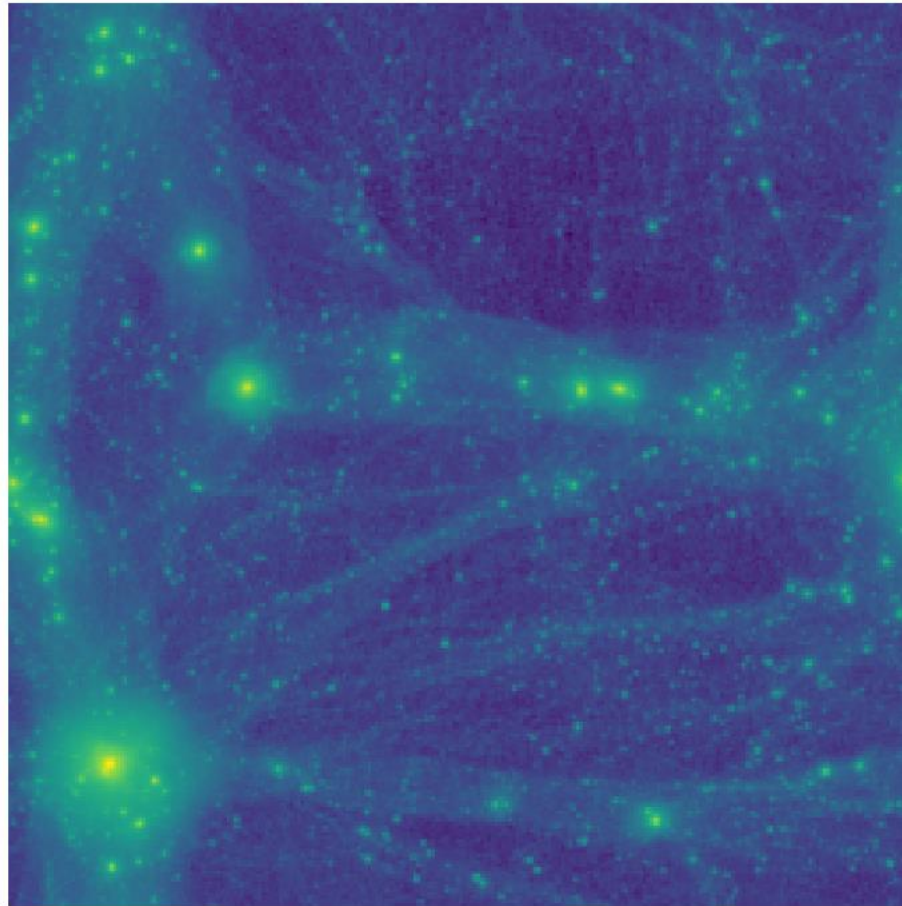


Figure: May and Springel [2021]

# We Want Something Smoother

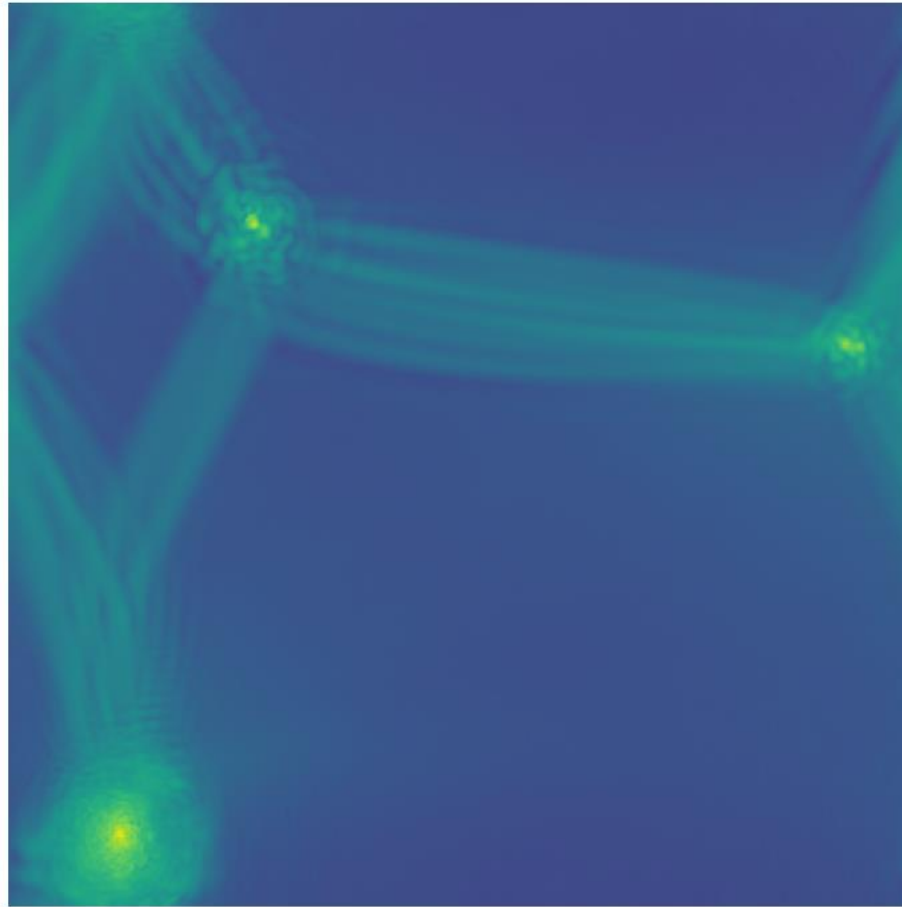


Figure: May and Springel [2021]

Enter **Wave dark matter**.

Consider a particle so light, it's de Broglie wavelength is on the scale of light years.

$$\lambda_{\text{deB}} = \frac{h}{mv} \quad (1)$$

These **ultra-light** particles are referred to as “**axion-like**” and their theoretical motivations arise from string theory and the QCD axion (Hui [2021]).

Typical masses are considered to be  $m \lesssim 30 \text{ eV}$  (Hui [2021]).

# What is FDM?

Fuzzy Dark Matter (FDM) is what we call wave dark matter in the mass range (Hui [2021]):

$$m = 10^{-22} \text{eV} - 10^{-20} \text{eV}$$

At this mass, the de Broglie wavelength becomes astrophysical.



# The Schrodinger-Poisson System

It turns out ultra-light dark matter behaves like a wave:

## Schrodinger-Poisson System (Widrow and Kaiser [1993])

FDM follows two well known equations:

$$\begin{cases} i\hbar \frac{\partial \psi}{\partial t} = \left[ \frac{-\hbar^2}{2m} \nabla^2 + \Phi(\vec{r}) \right] \psi \\ \nabla^2 \Phi = 4\pi G \rho \end{cases} \quad (2)$$

The norm-square of wavefunction is interpreted as the number density, so:

$$\rho = m|\psi|^2$$

# Cosmological Simulations

FDM does a good job of erasing small scale structure in cosmological simulations.

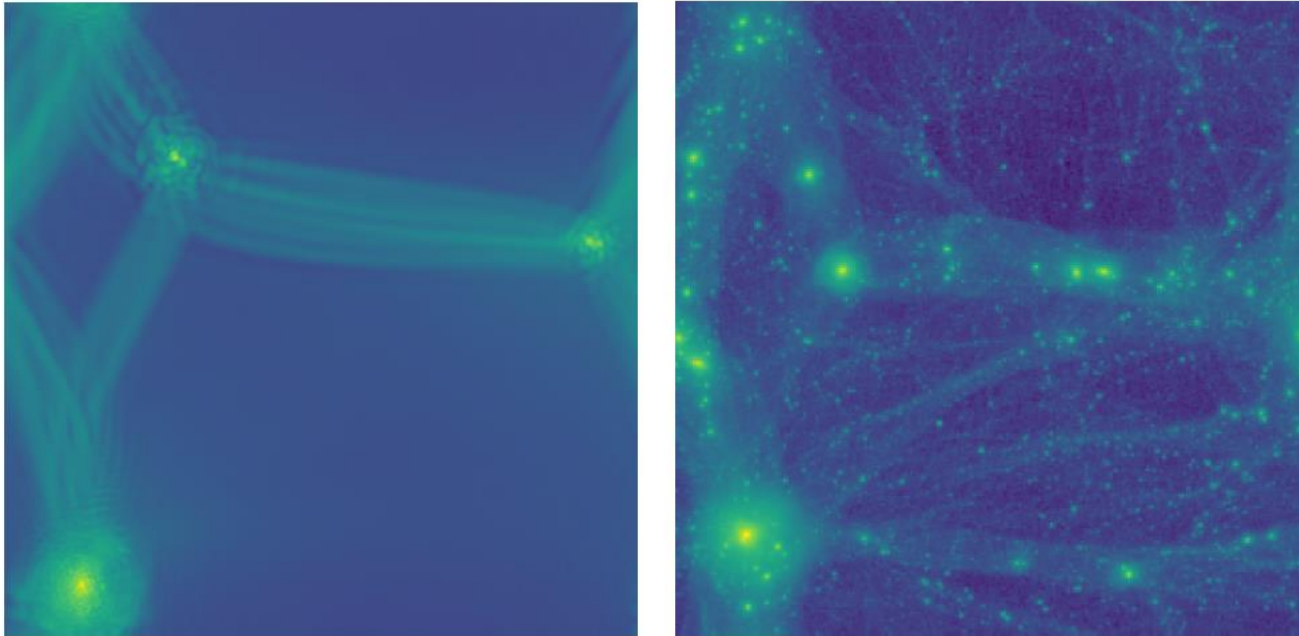


Figure: FDM (left) vs CDM (right). Simon May and Volker Springel (2021)

# The Case Against FDM

Recently, Neal Dalal and Andrey Kravtsov released a paper claiming they had excluded FDM as a possible candidate.

## Their Claim:

Granules (density fluctuations) in FDM dynamically heat test particles enough to cause galaxies to expand over time, and the necessary mass to suppress this phenomenon is  $m \approx 10^{-19} \text{eV}$ . This excludes the FDM mass range  $m = 10^{-22} \text{eV} - 10^{-20} \text{eV}$ . (Dalal and Kravtsov [2022])

# The Case Against FDM: Granules?

**Granules** arise from FDM wave-interference, as seen via **Fourier Expansion**:

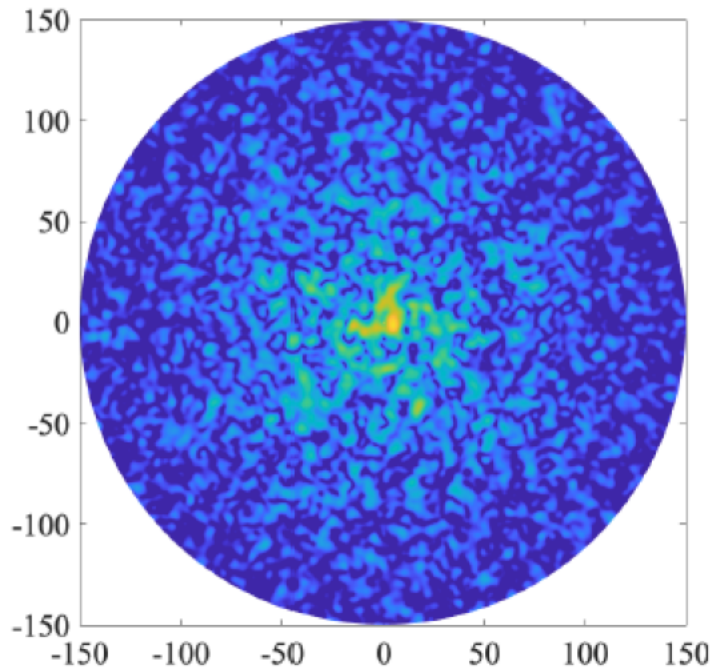


Figure: Dalal and Kravtsov [2022]

(1D and time-independent, for illustration)

$$\begin{aligned}\psi(x) &= \sum_k \hat{\psi}_k e^{ikx} \\ \Rightarrow |\psi|^2 &= \psi^* \psi \\ &= \left( \sum_{k'} \hat{\psi}_{k'} e^{-ik'x} \right) \left( \sum_k \hat{\psi}_k e^{ikx} \right) \\ &= \sum_k \hat{\psi}_k^2 + \sum_{k \neq k'} \hat{\psi}_k \hat{\psi}_{k'} e^{i(k-k')x}\end{aligned}$$

# The Case Against FDM: Dynamical Heating?

Granules “heat” test particles via gravity (Dalal and Kravtsov [2022]):

$$\delta a = \frac{G\delta M}{r^2}$$

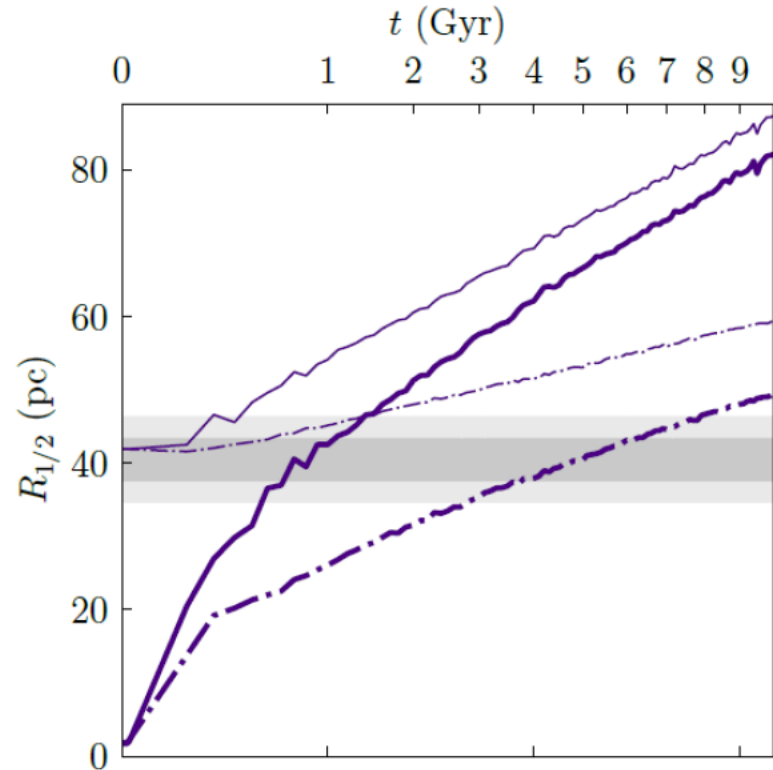
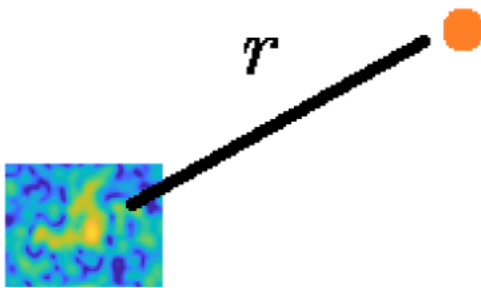


Figure: Dalal and Kravtsov [2022]

# A problem?

In Dalal and Kravtsov [2022], the particles (stars) are being treated as **test particles**.

This may pose a problem, in a system with comparable FDM and particle contributions.

# Our goal

We wish to create **fully self-consistent** simulations of a FDM and particle system. Meaning:

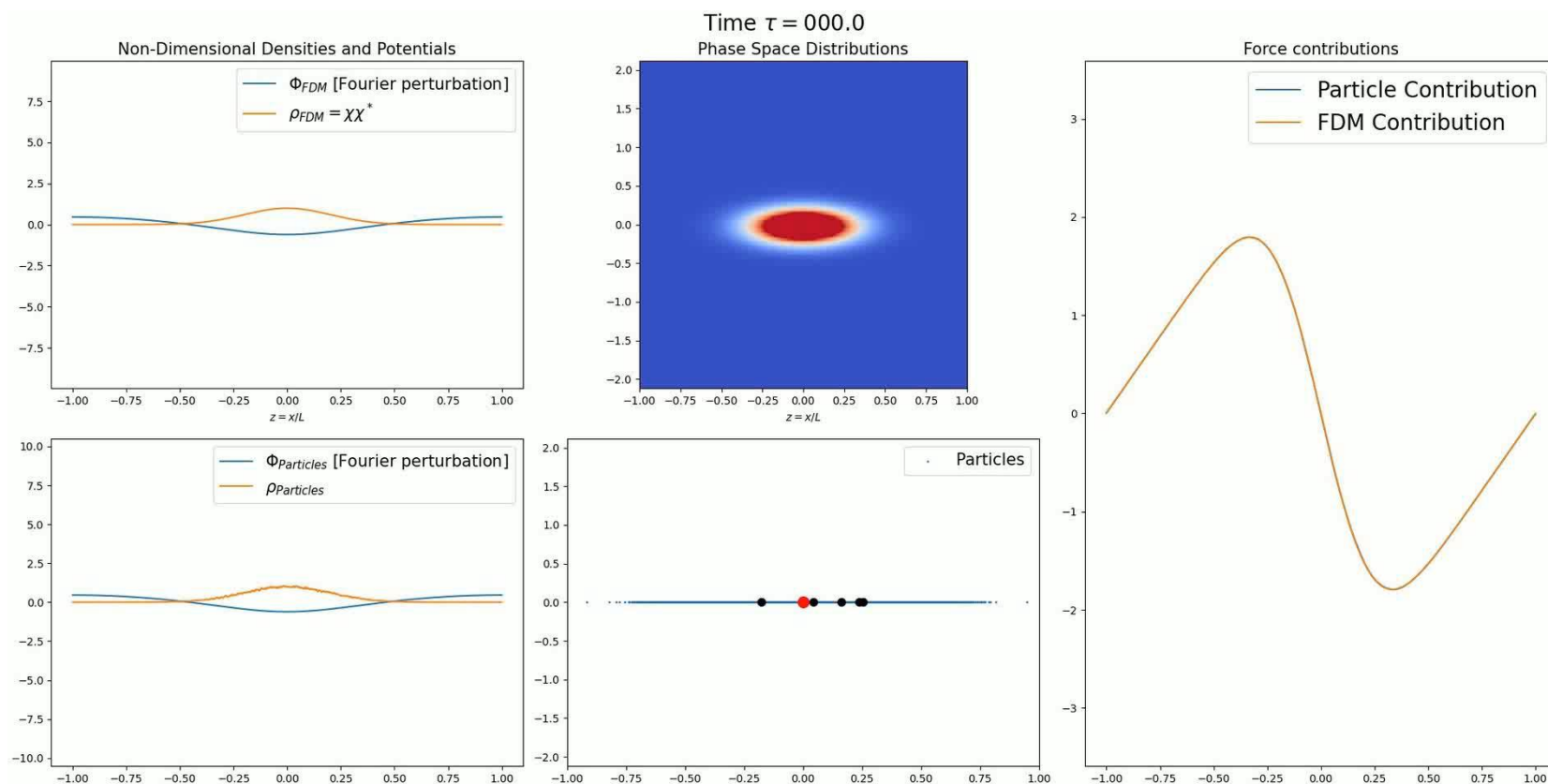
The FDM and particles are both contributing to the potential in the Poisson Equation:

$$\nabla^2\Phi = 4\pi G\rho = 4\pi G(\rho_{\text{FDM}} + \rho_{\text{particles}}) \quad (3)$$

In a system of comparable FDM and particle (total) masses, one should not be negligible.

# Preliminary Results

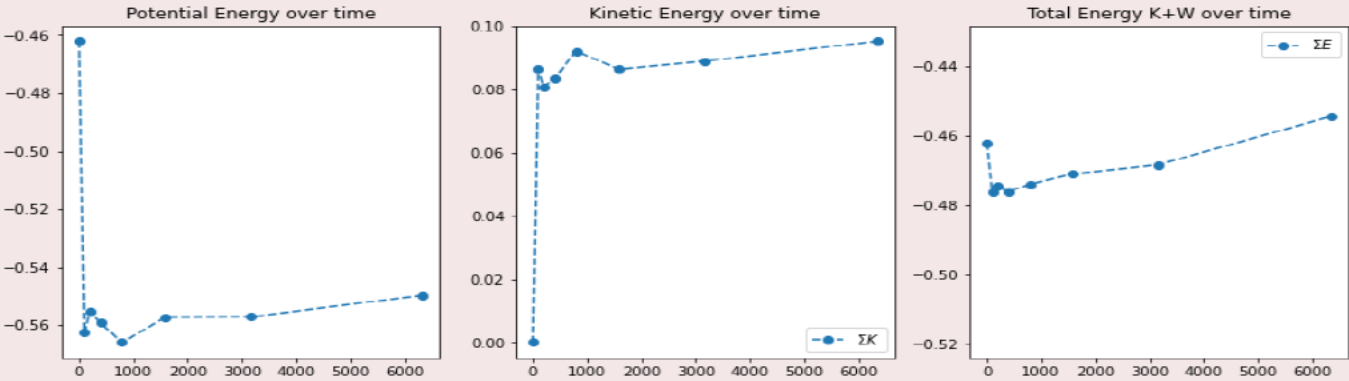
Our time evolution mixes a **“Kick-Drift-Kick”** (for Schrodinger) and a **Fast-Fourier-Transform (FFT)** method (for Poisson).



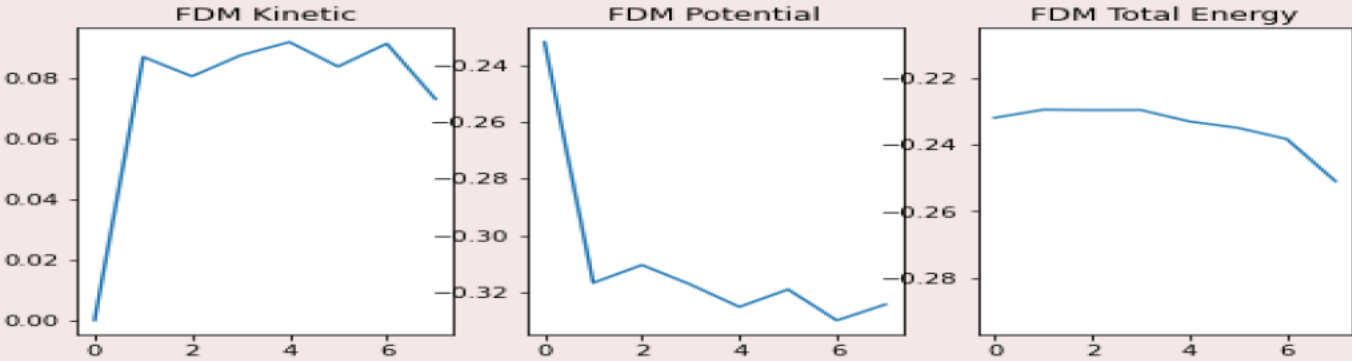


# More Results: Dynamical Heating

## Energy of Particles



## Energy of FDM



# References I

- N. Dalal and A. Kravtsov. Not so fuzzy: excluding fdm with sizes and stellar kinematics of ultra-faint dwarf galaxies, 2022. URL <https://arxiv.org/abs/2203.05750>.
- L. Hui. Wave dark matter. *Annual Review of Astronomy and Astrophysics*, 59(1):247–289, sep 2021. doi: 10.1146/annurev-astro-120920-010024. URL <https://doi.org/10.1146%2Fannurev-astro-120920-010024>.
- L. Hui, J. P. Ostriker, S. Tremaine, and E. Witten. Ultralight scalars as cosmological dark matter. *Physical Review D*, 95(4), feb 2017. doi: 10.1103/physrevd.95.043541. URL <https://doi.org/10.1103%2Fphysrevd.95.043541>.

## References II

- S. May and V. Springel. Structure formation in large-volume cosmological simulations of fuzzy dark matter: impact of the non-linear dynamics. *Monthly Notices of the Royal Astronomical Society*, 506(2):2603–2618, jun 2021. doi: 10.1093/mnras/stab1764. URL <https://doi.org/10.1093%2Fmnras%2Fstab1764>.
- D. H. Weinberg, J. S. Bullock, F. Governato, R. K. de Naray, and A. H. G. Peter. Cold dark matter: Controversies on small scales. *Proceedings of the National Academy of Sciences*, 112(40):12249–12255, feb 2015. doi: 10.1073/pnas.1308716112. URL <https://doi.org/10.1073%2Fpnas.1308716112>.
- L. M. Widrow and N. Kaiser. Using the Schroedinger Equation to Simulate Collisionless Matter. , 416:L71, Oct. 1993. doi: 10.1086/187073.