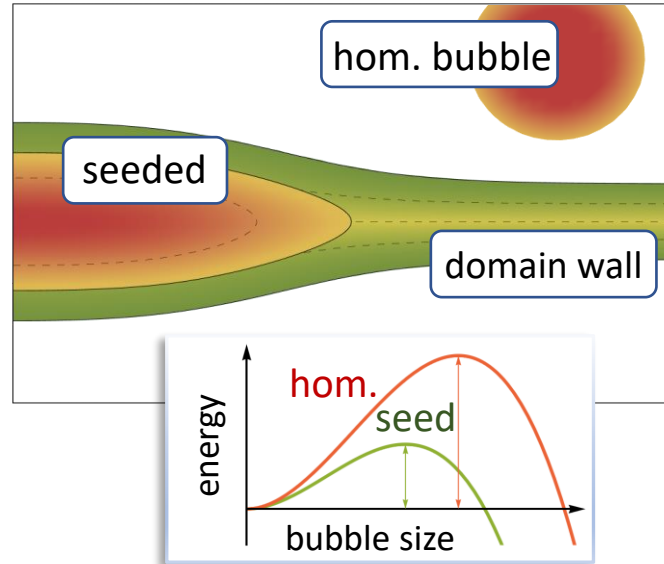


# Domain walls seeding the electroweak phase transition in the xSM



Simone Blasi

work in collaboration with A. Mariotti

Based on 2203.16450 [hep-ph]

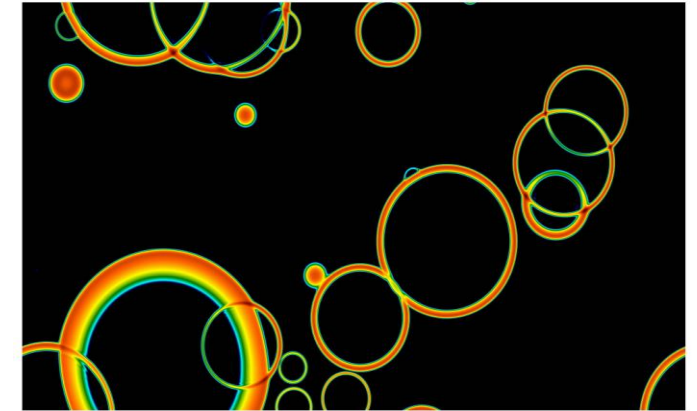


Be.HEP Summer Solstice, 21.06.22

# Introduction

- In the SM, QCD and electroweak phase transitions are not first-order (given measured  $m_h$ )

Fig. from  
Weir's review  
[1705.01783]



Departure from thermal equilibrium, **baryogenesis**

Have there been any first order transitions in the early Universe?

Stochastic background of **gravitational waves**

**New phase transitions,**  
e.g. secluded sector

Breitbach, Kopp, Madge, Opferkuch,  
Schwaller [1811.11175] JCAP

Ertas, Kahlhoefer, Tasillo  
[2109.06208] JCAP

**New physics** can affect the expected SM transitions

- Electroweak **extended scalar sectors**, composite Higgs, SUSY...
- QCD 1st order with  $N_f = 6$  (active)

e.g. Ipek and Tait  
[1811.00559] PRL

# The SM + $Z_2$ -odd scalar singlet

$h$  boson as the **portal** to new physics:

$$V = -\frac{1}{2}\mu^2 h^2 + \frac{1}{4}\lambda h^4 \\ -\frac{1}{2}m^2 S^2 + \frac{1}{4}\eta S^4 \\ + \frac{1}{2}\kappa h^2 S^2$$

When  $Z_2$  symmetry  $S \rightarrow -S$  imposed difficult to test at colliders (nightmare scenario)

Curtin, Meade, Yu [1409.0005] JHEP

- Simplest new physics scenario with strong first order EWPT, tree-level barrier

Espinosa, Konstandin, Riva [1107.5441] NPB

- Minimal mechanism for EW baryogenesis when  $Z_2 = CP$

Espinosa, Gripaios, Konstandin, Riva [1110.2876] JCAP

- New confining group as natural UV completion, next-to-minimal Composite Higgs

Gripaios, Pomarol, Riva, Serra [0902.1483] JHEP

- Standard benchmark for gravitational wave signals

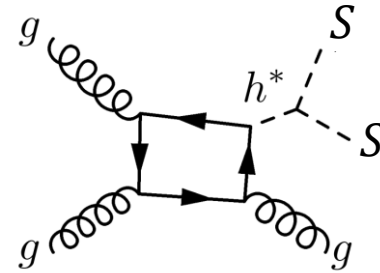
Caprini et al. [1512.06239] JCAP

# The SM + $Z_2$ -odd scalar singlet: into the nightmare

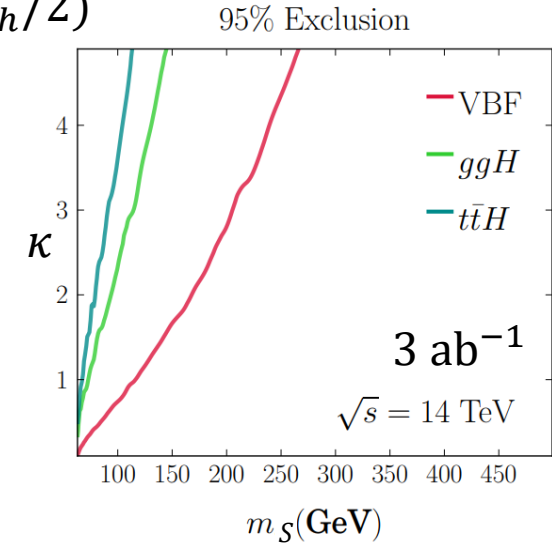
HL-LHC

$$V \supset +\frac{1}{2}\kappa h^2 S^2$$

- Non-resonant pair production ( $m_S > m_h/2$ )



$$pp \rightarrow h^{(*)} + X \rightarrow SS + X$$

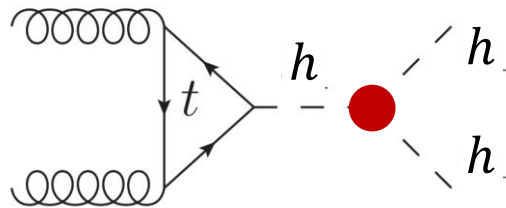


Figs. from  
Craig, Lou,  
McCullough,  
Thalapillild,  
[1412.0258],  
JHEP

- No single production of  $S$ , missing energy
- $m_S < m_h/2$  strongly constrained by invisible Higgs decays

Englert, Jaeckel,  
Khoze, Spannowsky,  
[1111.1719], PRD

- Loop corrections to trilinear  $h$  couplings



$$\frac{\kappa^2}{16\pi^2} \frac{v^2}{m_S^2} > 30\% \lambda_{SM}$$

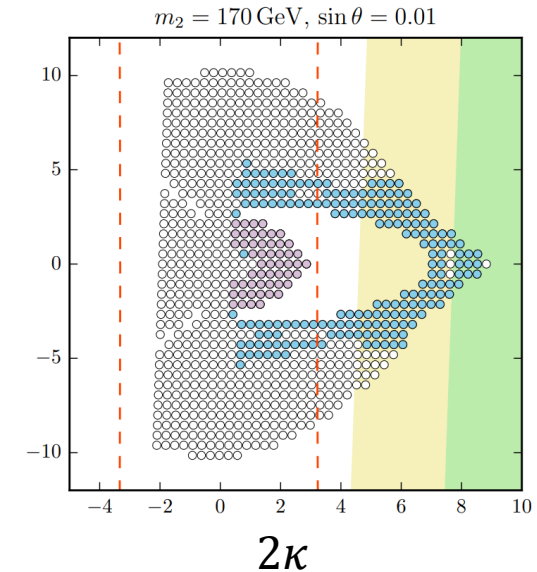
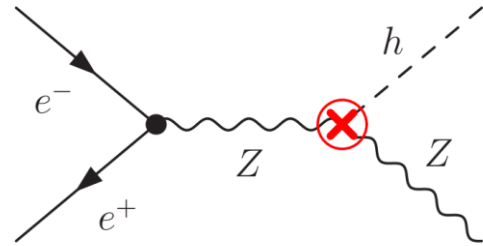


Fig. from  
Chen,  
Kozaczuk,  
Lewis,  
[1704.05844],  
JHEP

# The SM + $Z_2$ -odd scalar singlet: into the nightmare

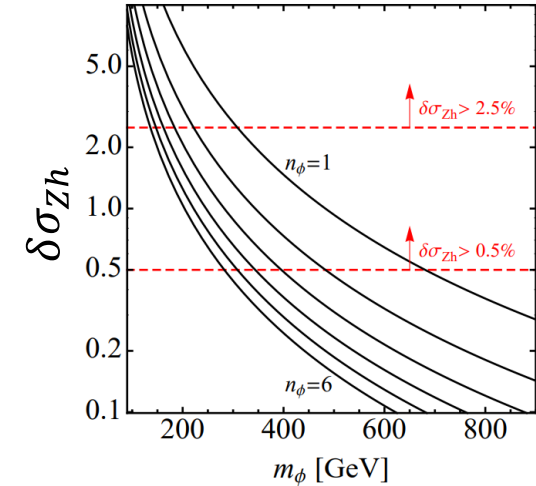
- Associated production at lepton colliders (Higgs-strahlung)

ILC,  
FCC-ee



$$\frac{\delta\sigma_{Zh}}{\sigma_{Zh}} \sim \frac{\kappa^2}{16\pi^2} \frac{v^2}{m_S^2} \sim O(1\%)$$

peaks @  $\sqrt{s} \approx 250$  GeV



Figs. from Craig, Englert, McCullough, [1305.5251], PRL

- Vector boson fusion at high-energy lepton colliders

CLIC,  
muon collider

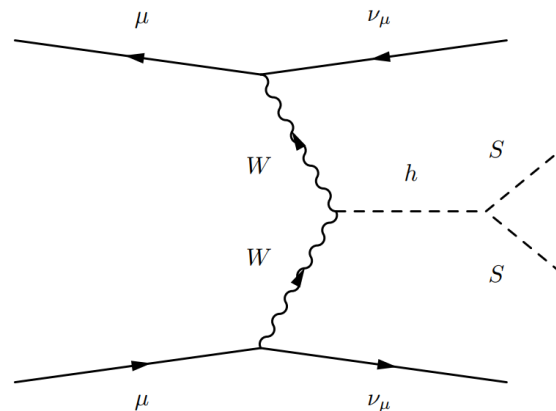


Fig. from Constantini, De Lillo, Maltoni, Mantani, Mattelaer, Ruiz, Zhao [2005.10289], JHEP

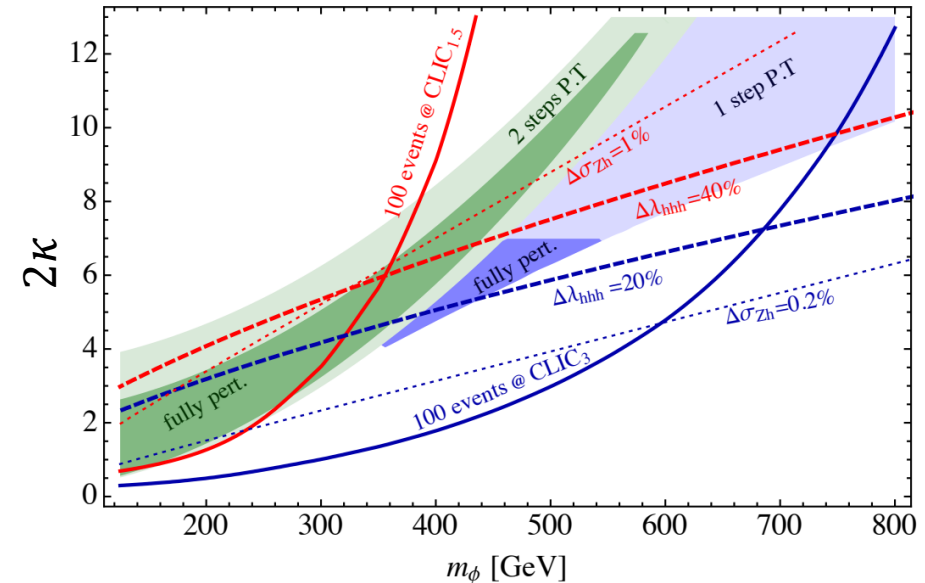
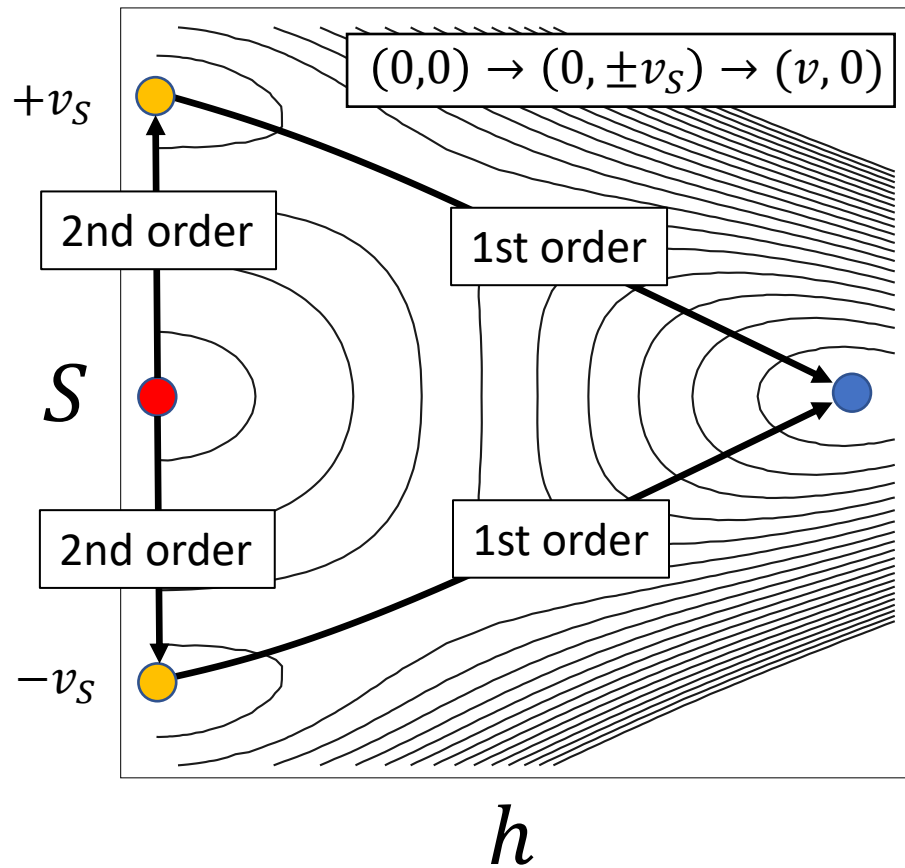


Fig. from Buttazzo, Redigolo, Sala, Tesi, [1807.04743], JHEP

# The electroweak PT in the SM + scalar singlet

- Two-step electroweak phase transition



- The second step is 1st order owing to non-zero  $S$  vev

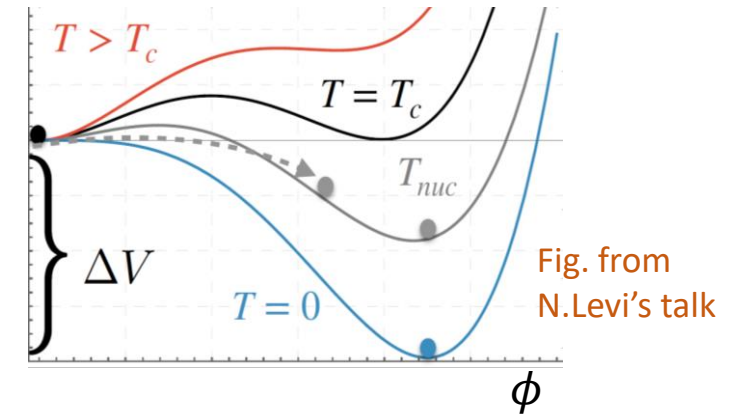
$$\gamma_V \sim T^4 \exp(-S_3/T) \sim H^4$$

spherical bubbles,  
homogeneous spacetime

- Minimal EW baryogenesis

$$L \supset i \frac{1}{\Lambda} S H \bar{q} \gamma_5 t$$

Espinosa, Gripaio, Konstandin,  
Riva [1110.2876] JCAP



- Detectable GW signal at LISA

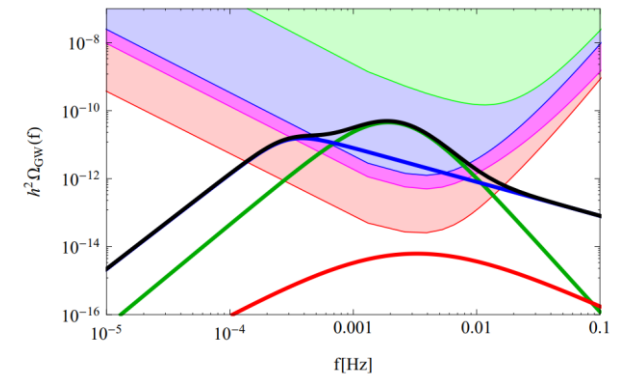


Fig. from Caprini et al.  
[1512.06239] JCAP

# However...

## MONOPOLE AND VORTEX DISSOCIATION AND DECAY OF THE FALSE VACUUM

Paul Joseph STEINHARDT

*Lyman Laboratory of Physics, Harvard University, Cambridge, Massachusetts 02138, USA*

Received 17 February 1981

“If **monopole** (or vortex) **solutions exist** for a metastable or false vacuum, **a finite density of monopoles** (or vortices) **can act as impurity sites that trigger inhomogeneous nucleation** and decay of the false vacuum.”

## Impurities in the early universe

Yutaka Hosotani

*Department of Physics, University of Pennsylvania, Philadelphia, Pennsylvania 19104*

(Received 1 November 1982)

“Now one has to ask the following question: **Is the early universe really sufficiently pure in order for supercooling to take place?** The aim of this paper is to show that in most cases the early universe is very pure. [...] In this paper we consider **ordinary particles as impurities.**”

## Cosmic separation of phases

Edward Witten\*

*Institute for Advanced Study, Princeton, New Jersey 08540*

(Received 9 April 1984)

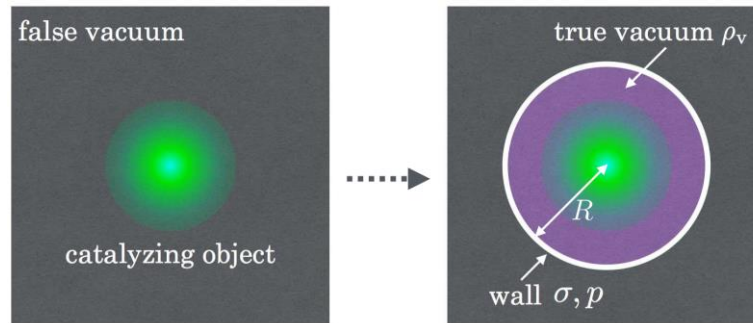
“In particle physics it is often assumed that phase transitions are nucleated by thermal fluctuations. In practice, [...] except in very pure, homogeneous samples, **phase transitions are often nucleated by various forms of impurities and inhomogeneities of nonthermal origin.**”

“What if the transition was nucleated by impurities? In this case **the mean spacing between bubbles has nothing to do with free energies** of nucleation and is simply the spacing between the relevant impurities.”

# The nature of impurities

- Compact objects, (not only) gravitational effects

Fig. from Oshita,  
Yamada, Yamaguchi  
[1808.01382], PLB



- Topological defects (strings and monopoles)

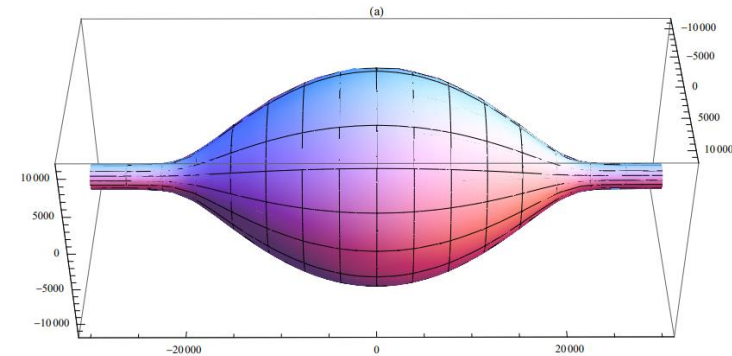
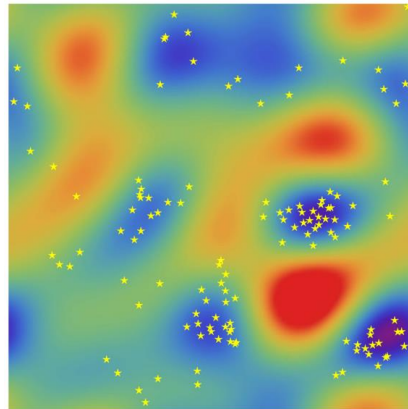


Fig. from  
Lee et al.,  
[1310.3005]  
PRD

- Primordial density fluctuations

Fig. from Jinno,  
Konstandin, Rubira,  
van de Vis,  
[2108.11947], JCAP



- What about domain walls?

This talk:

- ✓ Higgs + Singlet (xSM)
- ✓ Thermal history
- ✓ New method for bounce

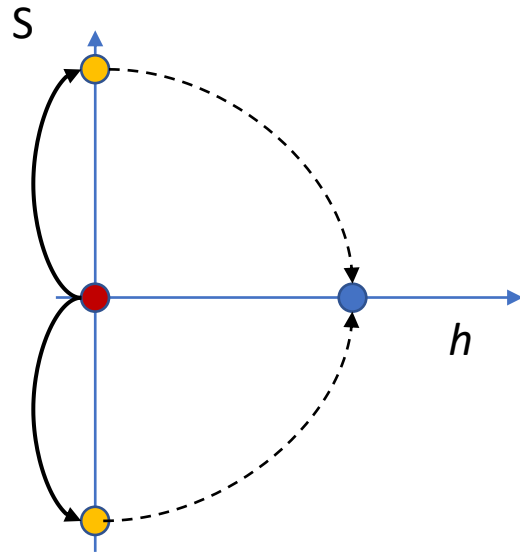
SB & Mariotti, [2203.16450]



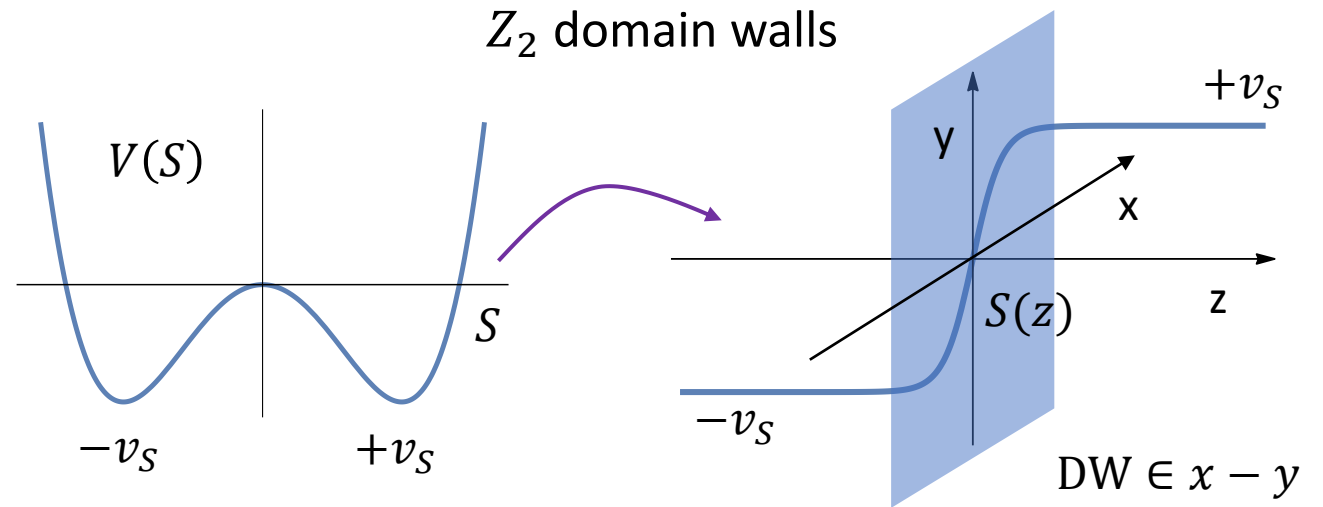
# Topological defects

[Zel'dovich et al. '74, Kibble '76]

- Defects as relics of phase transitions depending on topology of vacuum manifold  $M$ , not on the strength of transition



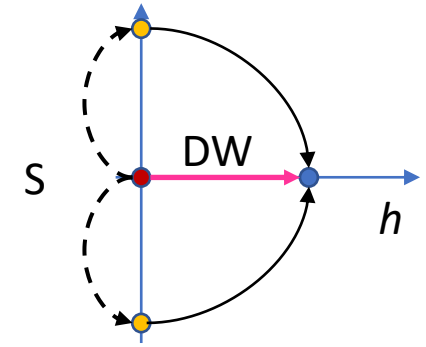
- The first step entails the formation of a **domain wall network**



- Vacuum manifold is disconnected  
 $\pi_0(M) \neq I$

Cosmic strings and other topological defects, Vilenkin and Shellard '94

# Seeded vs homogeneous tunneling



- Nucleation prob. no longer the same everywhere, enhanced at DW location ( $S = 0$ )

- Seeded, or inhomogeneous, tunneling probability per unit surface:

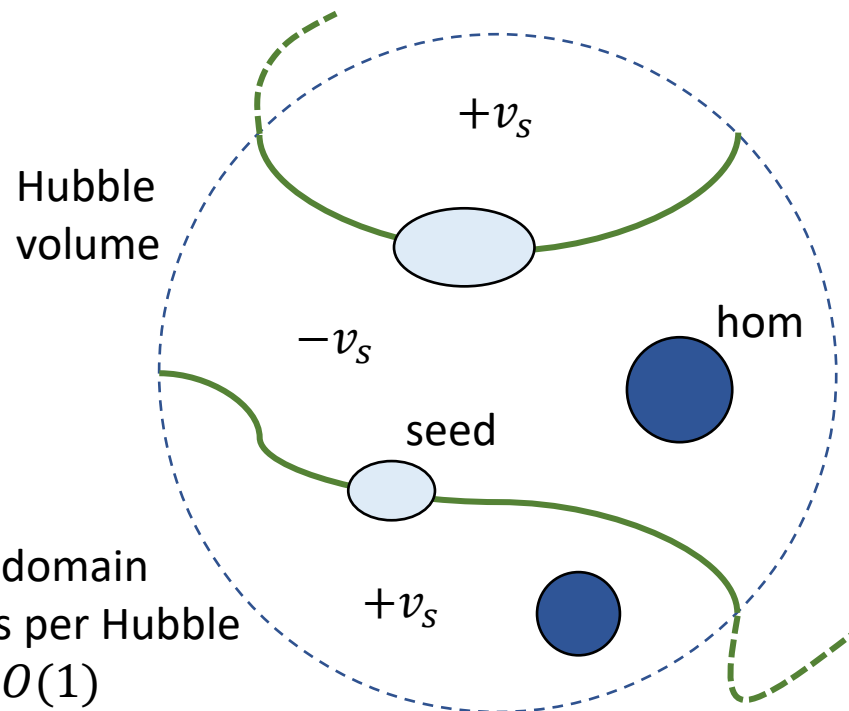
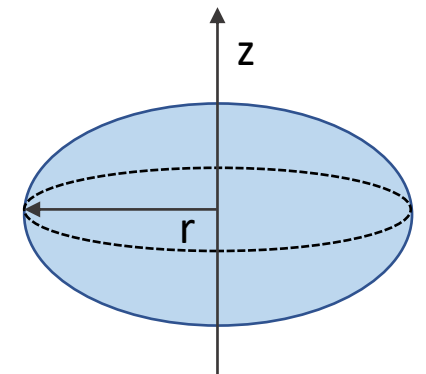
Lazarides, Shafi, Kibble 1982, PRD  
Perkins, Vilenkin 1992, PRD

$$\gamma_S \sim v_S^3 \exp(-S_2/T) \quad [\gamma_V \sim T^4 \exp(-S_3/T)]$$

- Stricter nucleation condition (only on submanifold)

$$S_2/T \sim 100 + \log \xi \quad [S_3/T \sim 140]$$

- Critical bubble will have only  $O(2)$  symmetry on the domain wall plane

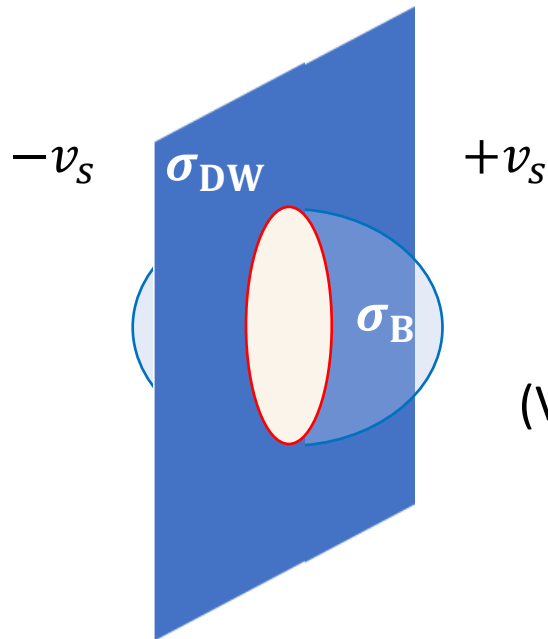


# of domain walls per Hubble  
 $\xi \sim O(1)$

# Seeded tunneling rate: thin wall limit

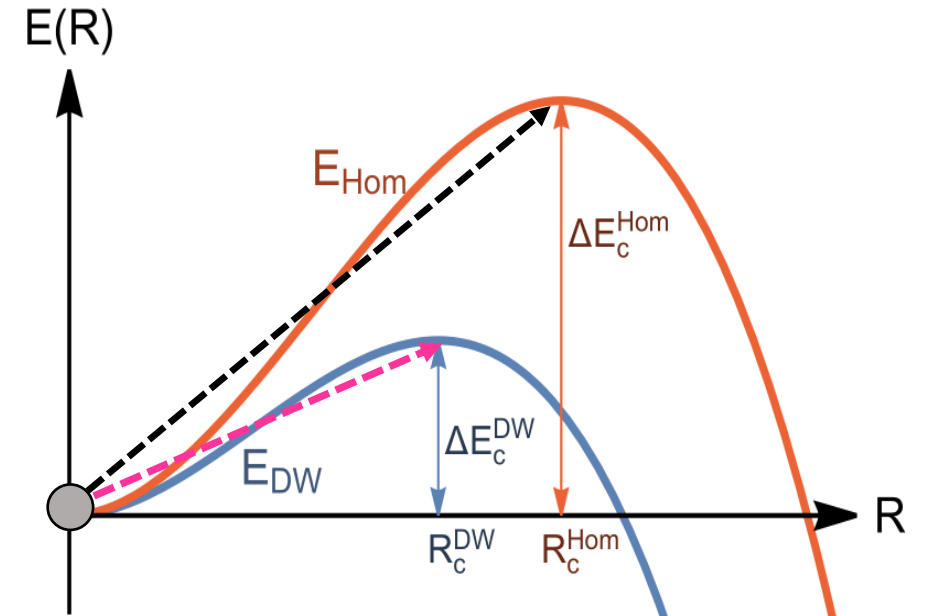
- Geometrical approach to estimate the energy of the critical bubble configuration

$$E(R) \simeq 4\pi R^2 \sigma_B - \frac{4}{3}\pi R^3 \epsilon - \pi R^2 \sigma_{DW}$$



Gain by eating up domain wall surface

(Valid at  $T \sim T_c$ )



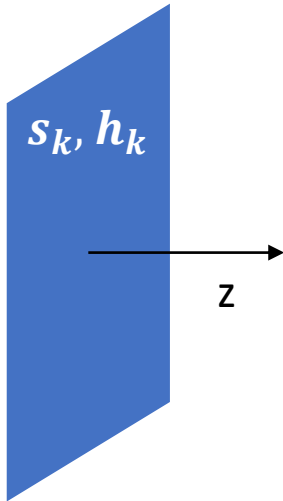
- Domain walls do catalyze the phase transition  $S_2/T \sim \Delta E/T$

# Seeded tunneling rate: Kaluza-Klein decomposition

- Expand the fields around the domain wall background, and integrate over the orthogonal direction

$$S = S_{\text{DW}}(z) + \sum_k s_k(x_\mu) \sigma_k(z)$$

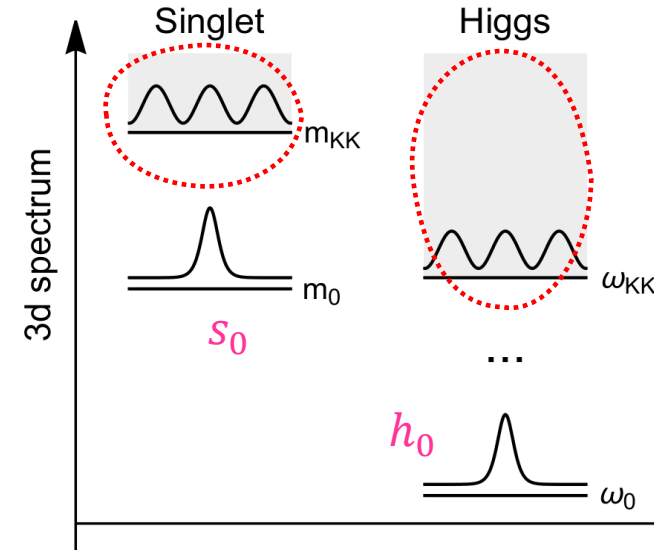
$$h = \sum_k h_k(x_\mu) \phi_k(z)$$

$$x_\mu = t, x, y$$


- Profiles are chosen in order to have canonical 3d fields

$$-\sigma_k''(z) + (3\eta S_{\text{DW}}^2(z) - m^2)\sigma_k(z) = m_k^2 \sigma_k(z)$$

$$-\phi_k''(z) + (\kappa S_{\text{DW}}^2(z) - \mu^2)\phi_k(z) = \omega_k^2 \phi_k(z)$$

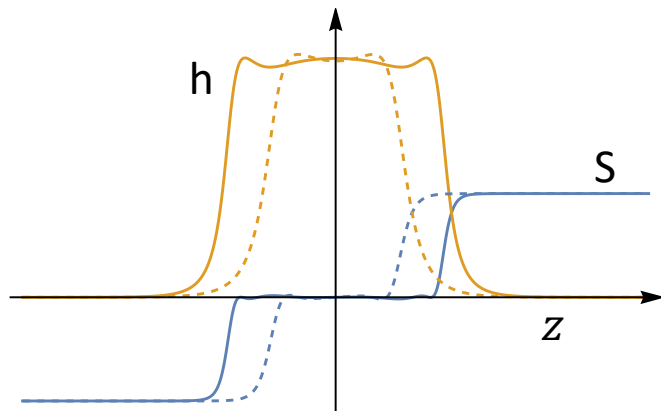


- Study the 3d theory on the DW plane, interaction from overlap integrals with  $V(h, S)$
- Take advantage of the gap to integrate continuum states out in favor of  $(h_0, s_0)$  EFT

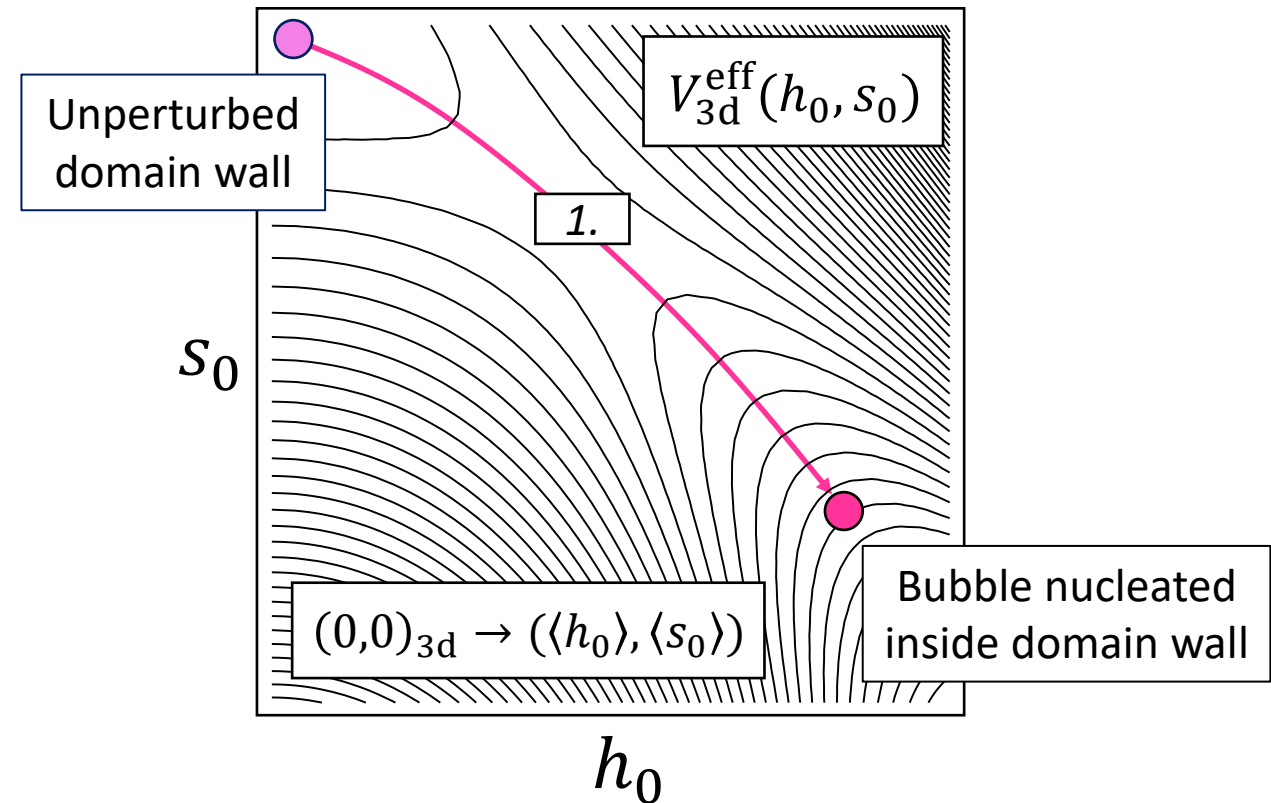
# Seeded tunneling rate: Kaluza-Klein decomposition

- Metastability of DW (= barrier at the origin) controlled by the **3d  $h$  mass,  $\omega_0^2$** :

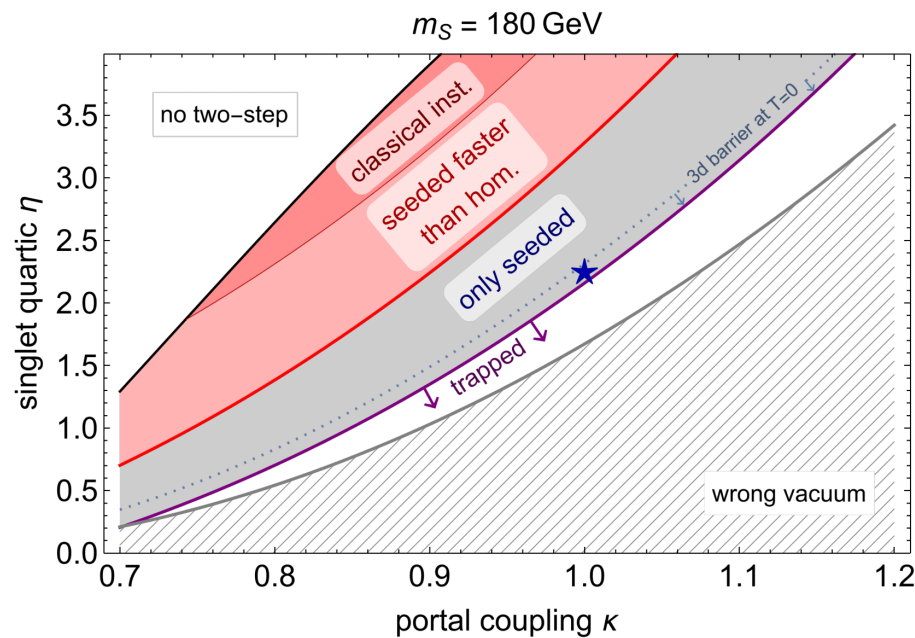
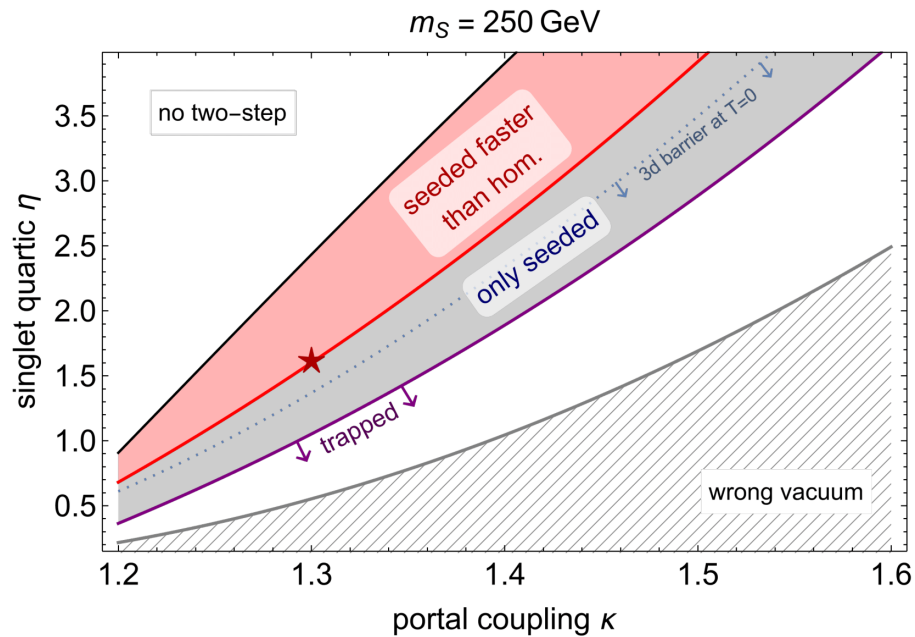
1.  $\omega_0^2 > 0$  for  $T < T_c$ : **seeded tunnelling** (origin is a minimum)
2.  $\omega_0^2 < 0$  for  $T > T_c$ : classical instability (origin is a saddle)



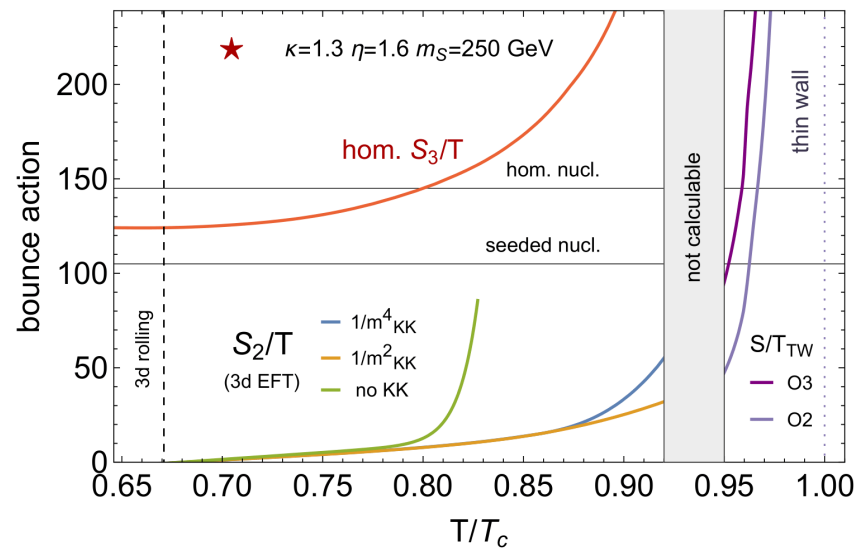
- **Seeded tunneling** as standard homogeneous problem in lower dimension,  $S_2/T$  from CosmoTransition



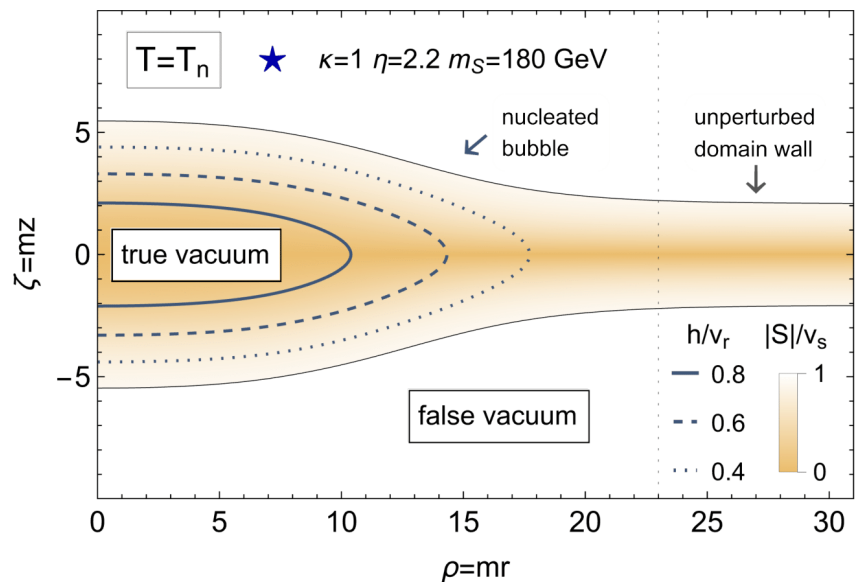
# Results



## Nucleation rate: homogeneous vs seeded

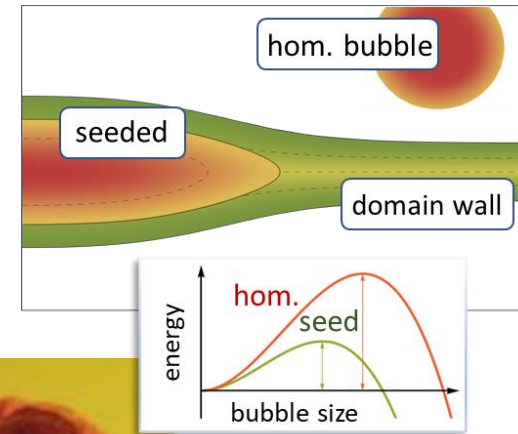


## Seeded critical bubble



# Conclusion

- The xSM is a compelling model for new physics with interesting properties and challenges at colliders
- The electroweak phase transition in the xSM is typically first order and proceeds as a two-step process
- In the  $Z_2$  limit, the first step entails the formation of domain walls: these act as local impurities which supersede the homogeneous nucleation
- The phenomenology of the seeded phase transition is a new target of study, e.g. for gravitational waves
- General ideas for multi-step phase transitions, possibly involving other types of defects



**HOMOGENEOUS  
NUCLEATION**

**NUCLEATION  
AROUND  
IMPURITIES**