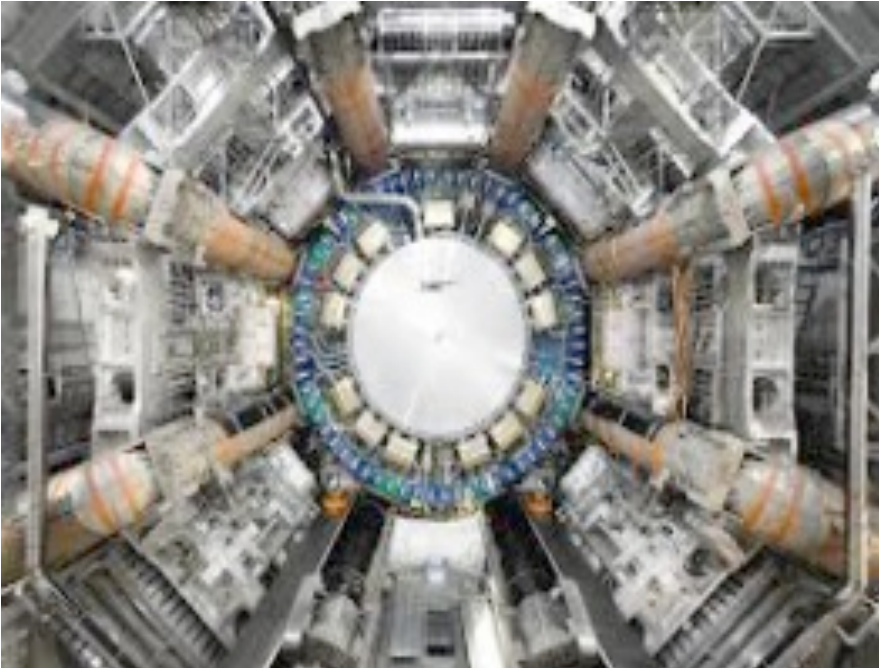


# Origins of the SA-CERN PROGRAMME



Zebon Z. Vilakazi  
(University of the Witwatersrand)

# Layout of the presentation

- Early origins of SA HEP activity
- How did SA-CERN start?
- What has been achieved?

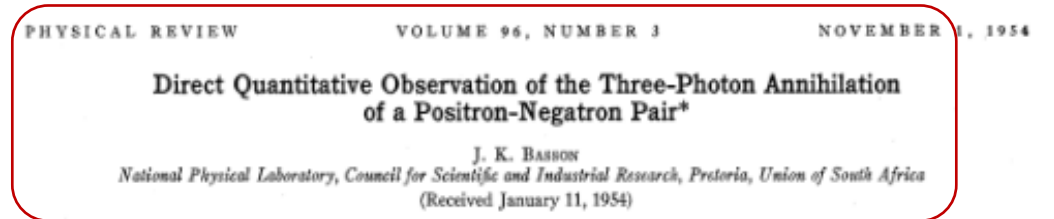


# Context and South Africa's history of HEP

## Some prominent physicists/cosmologists who were educated in RSA

- **Stanley Mandelstam** (Dirac medal); B. Sc. Hons (Witwatersrand)
- **Jonathan Dorfan** [ex SLAC director]; B. Sc. (Cape Town)
- **Werner Israel** [Cosmologist]; B. Sc (Cape Town)
- **Saul Teuklosky** [Dirac Medal] B. Sc Hons (Witwatersrand)
- **Peter Sarnack** (Maths Wolf Prize); B. Sc. Hons (Witwatersrand)
- **George Ellis** (FRS) – still at UCT

All the above – barring Ellis – were based abroad.



Three-photon annihilation of the positron with a negatron has been determined quantitatively as well as qualitatively by the simultaneous observation of the emitted photons with scintillation counters. The ratio of the reaction cross sections for two- and three-photon annihilation has been determined as  $\sigma_{2\gamma}/\sigma_{3\gamma} = 402 \pm 50$ . This is in agreement with the theory of Ore and Powell but definitely differs from the theoretical values obtained by Lifshitz and by Ivanenko and Sokolov.

### INTRODUCTION

THE possibility that an appreciable part of positron-negatron reactions might result in annihilation with the radiation of three photons, was first theoretically postulated by Lifshitz<sup>1</sup> and by Ivanenko and Sokolov<sup>2</sup> and a short while later by Ore and Powell.<sup>3</sup> They all used the time-independent perturbation theory to compute the cross section for three-photon annihilation. The influence of Coulomb binding was neglected and plane wave functions were assumed for the initial and final states of the positron-negatron system. Similar results were obtained but with different numerical values.

When the positron and negatron meet in free space they can be considered to form a bound system similar to that of the hydrogen atom, as suggested by Wheeler.<sup>4</sup> The triplet or singlet state is formed depending on whether the spins of the positron and negatron are parallel or antiparallel. These states are called respectively *ortho*- and *para*-positronium. Transitions between the two are strictly forbidden.<sup>5</sup>

The singlet state is annihilated with the emission

$10^{12}$  per second. This may result in the de-excitation of the triplet state to the singlet state, with resulting two- instead of three-photon annihilation, in a gas (such as NO) where electron exchange takes place easily. The number of delayed ( $\sim 10^{-7}$  sec) coincidences between the emission of the gamma quantum from the decay of the  $\text{Na}^{22}$  nucleus and the appearance of an annihilation quantum when the positron is brought to rest in the gas, has been measured by Deutsch in different gas mixtures. In the case of nitrogen, for example, the number of delayed coincidences—due to the formation of *ortho*-positronium—is markedly decreased by the addition of a few percent of NO. The electrons from the positronium atom are easily exchanged during a collision with an unpaired electron (from the NO) with opposite spin. Furthermore, by observing the number of delayed coincidences from positron capture in freon (where this exchange is almost nonexistent) as a function of the pressure and extrapolating to zero pressure, Deutsch found for the lifetime of the *ortho*-positronium a value in good agreement with the theoretical value of Ore and Powell.

H

SA

# EVIDENCE FOR HIGH-ENERGY COSMIC-RAY NEUTRINO INTERACTIONS\*

F. Reines, M. F. Crouch, T. L. Jenkins, W. R. Kropp, H. S. Gurr, and G. R. Smith

Case Institute of Technology, Cleveland, Ohio

and

J. P. F. Sellschop and B. Meyer

University of the Witwatersrand, Johannesburg, Republic of South Africa

(Received 26 July 1965)

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are favored over the  $3\pi$  and  $\eta+2\pi$  modes, although it probably is still insufficient to account for the vast differences in decay rates between these two types of processes without introducing symmetry-breaking effects. The  $\rho+2\pi$  and  $\omega+2\pi$  modes are found to be comparable. For a detailed list of branching ratios, see reference 2.

<sup>1</sup>H. Harari, H. J. Lipkin, and S. Meshkov, Phys.

other way with equal amounts. Therefore, the statistical average of the  $\rho+3\pi$  processes should not be greatly perturbed.

<sup>2</sup>See reference 1 for a summary of the experimental data.

<sup>3</sup>R. Armenteros et al., Phys. Letters 17, 170 (1965); N. Barash et al., "Antiproton Annihilation in Hydrogen at Rest I. Reaction  $\bar{p}+p \rightarrow K+\bar{K}+\pi$ " (to be published).

## EVIDENCE FOR HIGH-ENERGY COSMIC-RAY NEUTRINO INTERACTIONS\*

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The flux of high-energy neutrinos from the decay of  $K$ ,  $\pi$ , and  $\mu$  mesons produced in the earth's atmosphere by the interaction of primary cosmic rays has been calculated by many authors.<sup>1</sup> In addition, there has been some conjecture<sup>2</sup> as to the much rarer primary flux of high-energy neutrinos originating outside the earth's atmosphere. We present here evidence<sup>3</sup> for the interactions of "natural" high-energy neutrinos obtained with a large area liquid scintillation detector (110 m<sup>2</sup>) located at a depth of 3200 m (8800 meters of water equivalent, average  $Z^2/A \approx 5.0$ ) in a South African gold mine.

The essential idea of the present experiment<sup>3</sup> is to detect the energetic muons produced in neutrino interactions in a mass of rock by means of a large area detector array imbedded in it. Backgrounds are reduced by the large overburden and by utilizing the fact that the angular distribution of the residual muons from the earth's atmosphere is strongly peaked in the vertical direction at this depth. The angular distribution of the muons produced by neutrino interactions should show a slight peaking in the horizontal direction.<sup>1</sup>

The detector array, shown schematically in Fig. 1, consists of two parallel vertical walls made up of 36 detector elements. The array is grouped into 6 "bays" of 6 elements

each. Each detector element, Fig. 2, is a rectangular box of Lucite of wall area 3.07 m<sup>2</sup> containing 380 liters of a mineral-oil based liquid scintillator,<sup>4</sup> and is viewed at each end by two 5-in. photomultiplier tubes. The array constitutes a hodoscope which gives a rough measurement of the zenith angle of a charged particle passing through it. In addition, the event is located along the detector axis by the ratio of the photomultiplier responses at the two ends. The sum of the responses then pro-

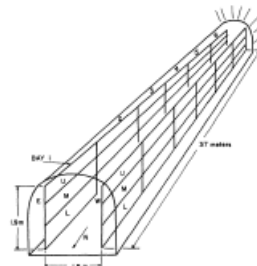
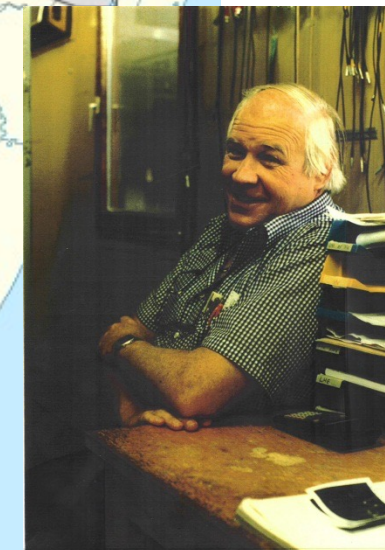


FIG. 1. Schematic of detector array.



Frederick Reines



Friedel Shellschop

# Current issues and the African context

## About 0.5% of CERN users are African Nationals

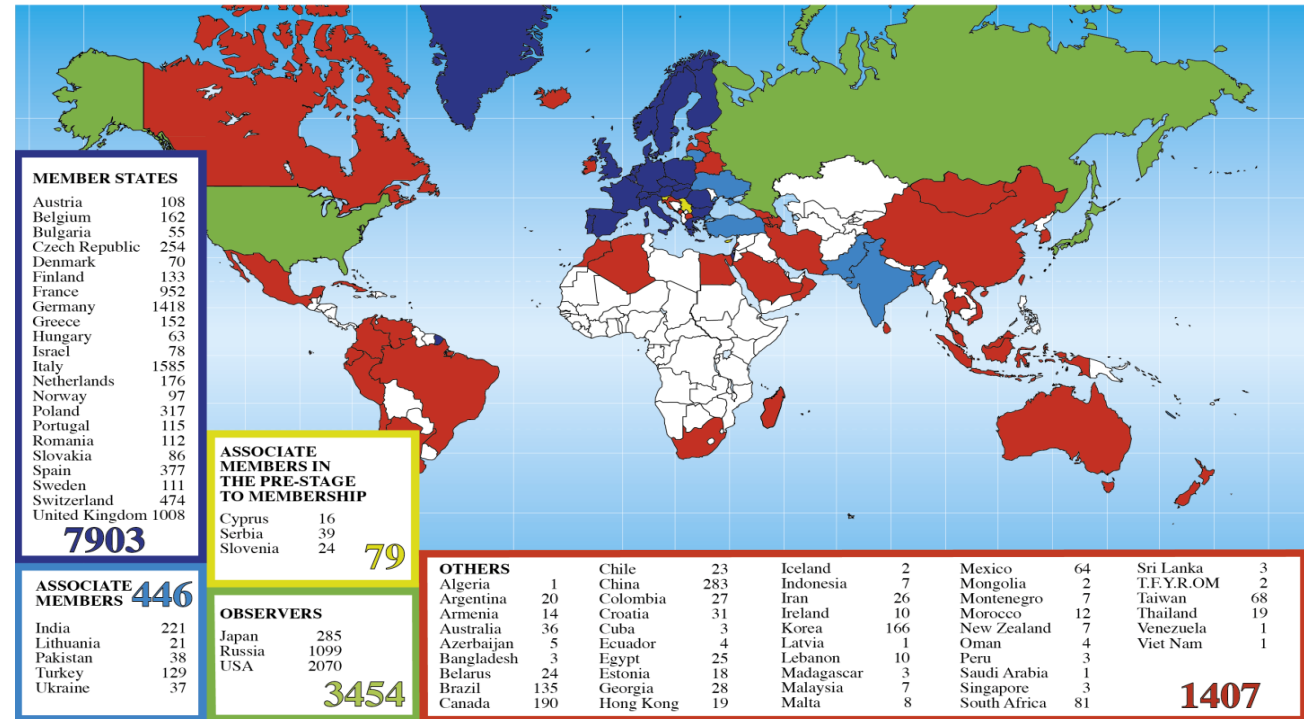
Low participation of African scholars in major research labs around the world. Some examples:

- ▣ CERN users
- ▣ Users of LHC experiments

Not limited to CERN. Broader issue

Opportunity: African Youth Bulge and fast growing economies (albeit from low-base) like Asian economies of past decades.

Distribution of All CERN Users by Location of Institute on 24 January 2018



# Personal Journey



- 1994 —1997 : Ph. D.; CERN- NA43 experiment (Wits/Aarhus)
- 1998 —1999:Post-Doc;CERN- NA59 experiment [M. Velasco (Northwestern)]
- 1999 — 2006 : Faculty at University of Cape Town -> SA-ALICE : initiated HEP
- 2007 —2013 : Director: iThemba LABS —> **SA-CERN programme**
- Deputy V-C Research: university of the Witwatersrand (2014 – ) [ALICE + ATLAS]



# SA OFFICIALS AT CERN

1992: Signing of first Co-operation Agreement by FRD (NRF) President R Arndt and CERN DG C Rubbia.



2005: Visit by Minister of Science of Technology Mr Mangena



2011: Visit by Minister of Science of Technology Mrs N Pandor



# Background: History from UCT-ALICE to UCT-CERN

- UCT joins ALICE (November 2001)
- UCT-CERN Research Centre (August 2003)
- SA-CERN (15 December 2008)
- ALICE: UCT and iThemba LABS
- MoU – Muons & Grid Computing (21 March 2011)
- Wits joins ALICE (March 2014)
- CHPC signs MoU: computing for ALICE&ATLAS (April 2015)
- SA-ALICE joins LS-2 Upgrade (March 2018)

## City boffins join Alice for a Big Bang

The University of Cape Town is among key sites in a world computer linkup on the Net that aims to recreate the beginning of the universe. Education Reporter **Jeanne van der Merwe** reports

When millions upon millions of atoms collide, they need thousands of scientists and enormous computer capacity.

Cleymans, one of the world's leading physicists in the field of nuclear physics, explained: "With this experiment we are trying to get as close as we can to the Big Bang. We can never get it because the Big Bang itself, because that only happened once and can never happen again."

But if we collide different ions together at high enough speed, we might make them "melt" into one, making the different atom nuclei indistinguishable and create that condition of quark-gluon plasma "particle soup" that probably existed a microsecond or so after the Big Bang, long before atoms started defining themselves in electrons, neutrons and protons.

In the experiment, scientists will use a CERN accelerator 26km in circumference to collide lead ions with an electric charge at a fraction slower than light speed.

To ascertain whether they succeeded in making the quark-gluon plasma, they have to track the paths of billions upon billions of atomic particles generated by the atoms' collisions in the accelerator. This is why they need a completely new kind of computer, "a net" with enough collective brainpower to receive and process all information. This revolutionary "data grid" is similar to the World Wide Web, but it can run a program simultaneously on instead of merely exchanging data files on the current Web.

The project will cost more than R600 million, funded mostly by western European countries, with substantial contributions from Japan, the United States and India.

UCT member Zebon Vilakazi, a UCT lecturer who has done post-doctoral studies at Cern, said Cleymans' expertise in nuclear physics was one of the main reasons why the university had been asked to join the programme.

"Over Cleymans has been actively involved in this field for more than 20 years and he is regarded as one of the leading authorities in the world."

The UCT team's job will be in the experiment is to design a way of collecting the data most relevant to the experiment.

"The experiment simply creates so much data, and there are certain events within the experiment that are more important than others. We are finding a way of streamlining the data into a computer programme that can make thousands of decisions a second. This is crucial for the success of the project, because if you lose even a part of an important ion, you could lose a valuable part of the experiment, and finding it again would be like finding a needle in a haystack."



**Pibball wizards:** UCT physics professor Jean Cleymans and colleague Zebon Vilakazi with an electronic instrument while they are colliding in an accelerator.

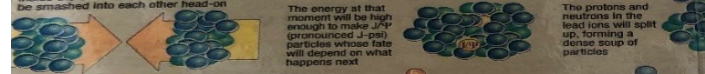
## UNRAVELLING THE MYSTERIES OF THE ORIGIN OF THE UNIVERSE

Scientists believe that there was a Big Bang from which everything in the universe emerged. Fifteen billion years later, the universe is so huge that it would take light millions of years to cross.

To understand the first moments of the Universe's life, the European Organisation for Nuclear Research (CERN) is using particle accelerators to smash ions into each other. The purpose is to find and study matter called Quark-Gluon Plasma (QGP).

### SECTION OF A PARTICLE ACCELERATOR

These machines are housed in a massive complex of circular underground tunnels, the largest being 27km round. Accelerators use powerful electric fields to push energy into a beam of particles which go round and round collecting more energy with each lap.



As the QGP cools down, ordinary matter condenses out, like water condensing from steam. By measuring how many  $J/\psi$ 's, how many particles containing strange quarks and how many electrons emerge from the collision, researchers will be able to tell whether a QGP formed, and learn about how it evolved.

**GLOSSARY OF TERMS**  
 Protons, positively charged sub-atomic particles forming the nucleus of an atom.  
 Neutrons, sub-atomic particles making up the nucleus of an atom.  
 Electrons, sub-atomic particles carrying the negative charge.  
 Quarks, sub-atomic particles carrying the nucleus of an atom.  
 Gluons, sub-atomic particles carrying the force between quarks.  
 Positrons, positively charged sub-atomic particles forming the nucleus of an atom.  
 Neutrons, sub-atomic particles making up the nucleus of an atom.  
 Electrons, sub-atomic particles carrying the negative charge.  
 Quarks, sub-atomic particles carrying the nucleus of an atom.  
 Gluons, sub-atomic particles carrying the force between quarks.

## Explosion of data set to revolutionise the Web

RECREATING the "soup" of sub-atomic particles that existed in the few microseconds after the Big Bang, may be of no consequence at all to ordinary people.

But the computer system that will be used in the experiment could revolutionise the World Wide Web.

The international ALICE collaboration, an experiment involving a thousand scientists all over the world and sponsored by the European Organisation for Nuclear Research (CERN) near Geneva in Switzerland, is the second undertaking of its kind in history.

The first one in 1979 sought to find out why positive-charged protons and neutrons stick together in the nucleus of the atom instead of flying apart like similar poles of two magnets.

In that experiment, led by IBM Nobel physics laureate Carlo Rubbia, scientists developed a system of inter-linked computers able to exchange data files - the start of the World Wide Web, which straddled the world forever.

In the ALICE collaboration, CERN and its participating institutes will form a "data grid" of thousands of inter-linked computers working in concert all over the world. Instead of merely exchanging data files like the Web, the data grid will be capable of running programmes simultaneously.

The system will consist of several tiers of computers linked in parallel to create enormous data processing capacity.

UCT physicist Zebon Vilakazi

explained the need for such powerful computers.

"The experiments in the accelerators generate 40 000 individual tracks of the movements of individual sub-atomic particles as they collide with one another in the accelerator just once a second. To store a day's worth of data, you'd need 50 000 CDs. That is simply too much to store with. We must create a computing system that will be able to make millions of decisions a minute about which tracks to keep and which to discard, and to store them."

The data of the experiment, which takes place in a giant accelerator, will be formed out from CERN's computer "farm" ("The Zoo") to several "Tier-1" centres, of which UCT is one. These centres, all equipped with "farms" of parallel-linked computers, will analyse data or send it through to other linked computers.

UCT physicist professor Jean



**THE CAPE ARGUS ( 1<sup>ST</sup> MAY 2002 )**



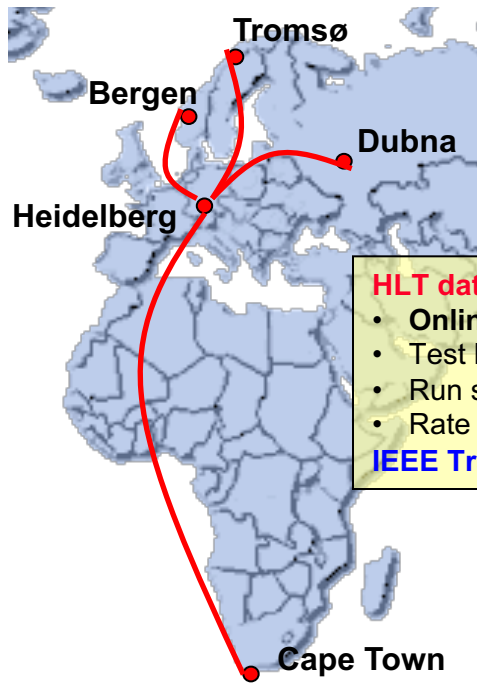
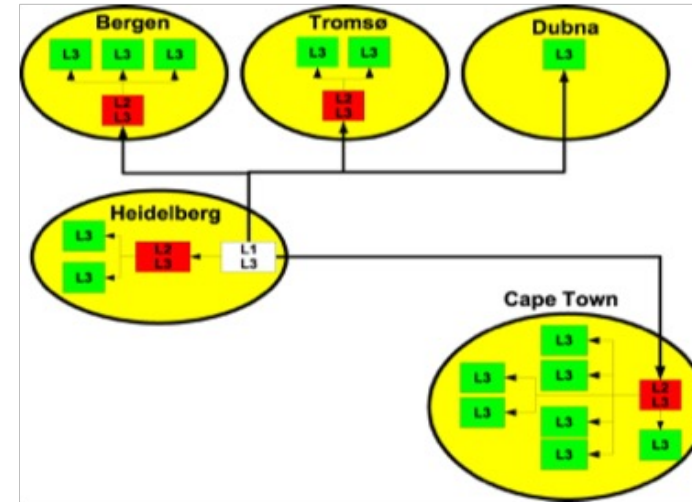
# Participation of SA in ALICE

**Dimuon Arm:** algorithms for online High Level Trigger (dHLT, commissioned in 2008)

- Real-time processing of heavy-ion collisions:
  - online event selection
  - compression
- calibration and quality monitoring
- 25 GB/s, up to 1000 ev/s
- Extreme computing, networking and process control demands in 2008

**Grid Computing:** Computing Cluster integrated into ALICE GRID

**Physics:** Dimuon studies (acceptance, efficiency); W production in pp; ..



**HLT data challenge: Nov. 2004**

- Online test on 'Grid' !
- Test latency tolerance
- Run stable for > 15 hours
- Rate limited by bandwidth

[IEEE Trans.Nucl.Sci.55:703,2008.](#)

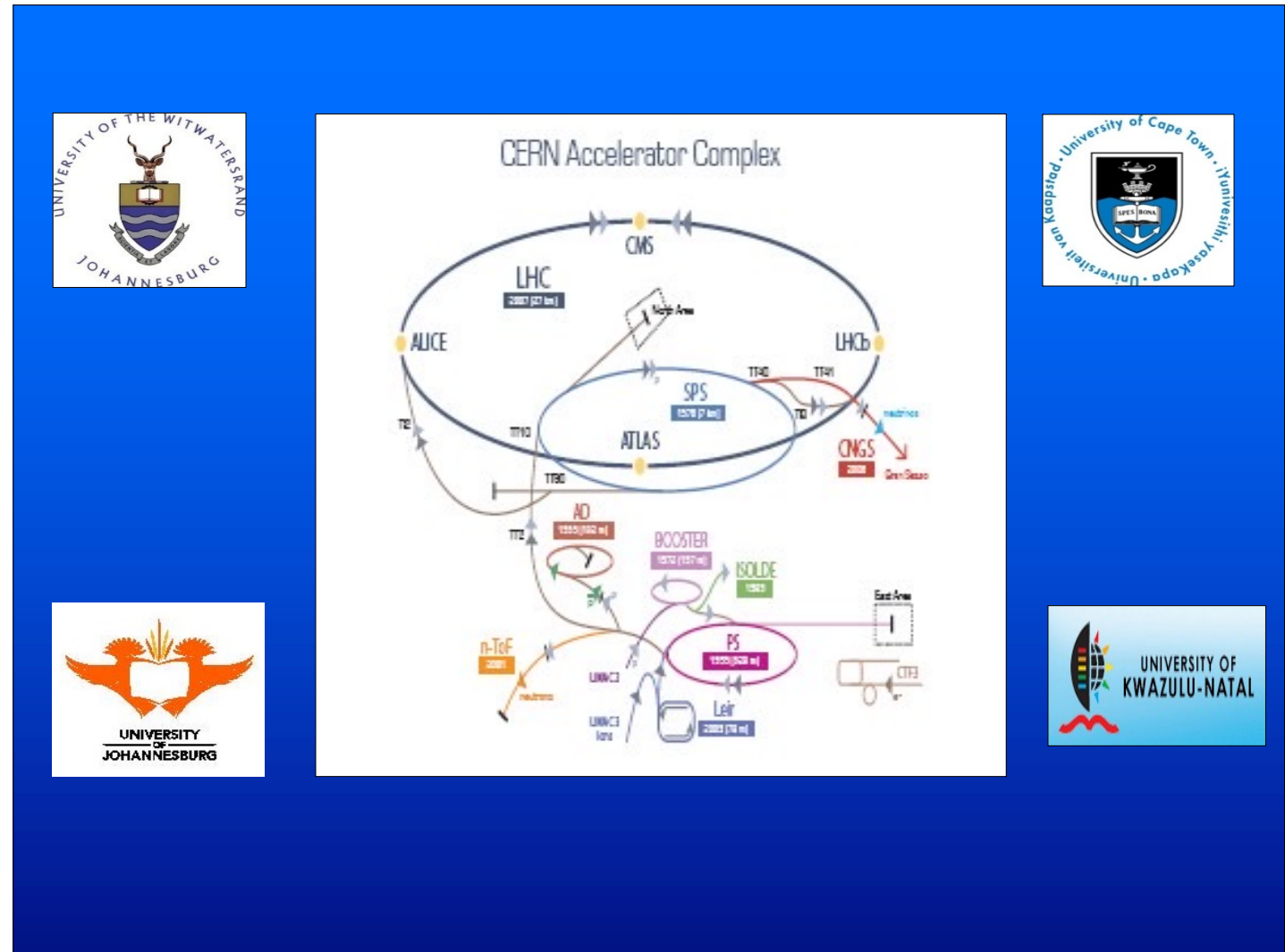




# The SA-CERN Programme



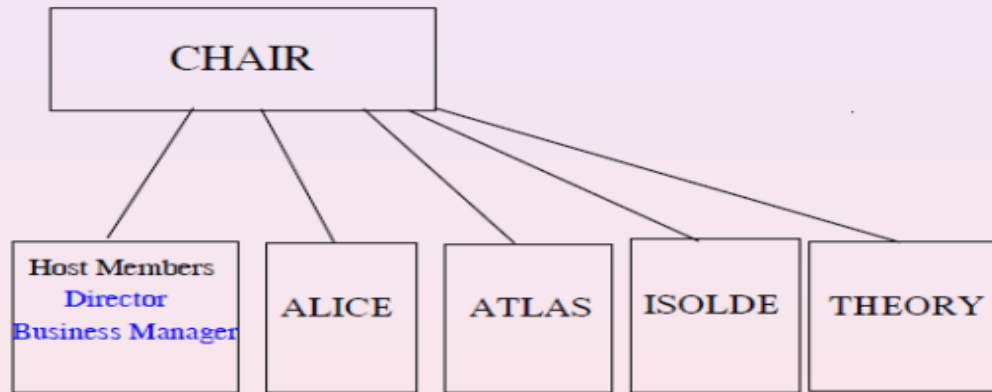
1. Started as a consortium of researchers who had long standing research program with CERN
2. Modelled along the Australian, Indian and Brazilian programs.
  - Allows for central point of coordination and resource allocation.
3. Agreement was that iThemba LABS would act as a neutral institutional hosts for the SA-CERN program:



Has been recognised by the ministry as an exemplar for other multi-lateral collaborations: **SA-JINR**

## SA-CERN Programme

Host Institution: iThemba LABS



Launch of the SA-CERN consortium December 2008.

## Host of the National SA-CERN

## Programme: iThemba LABS

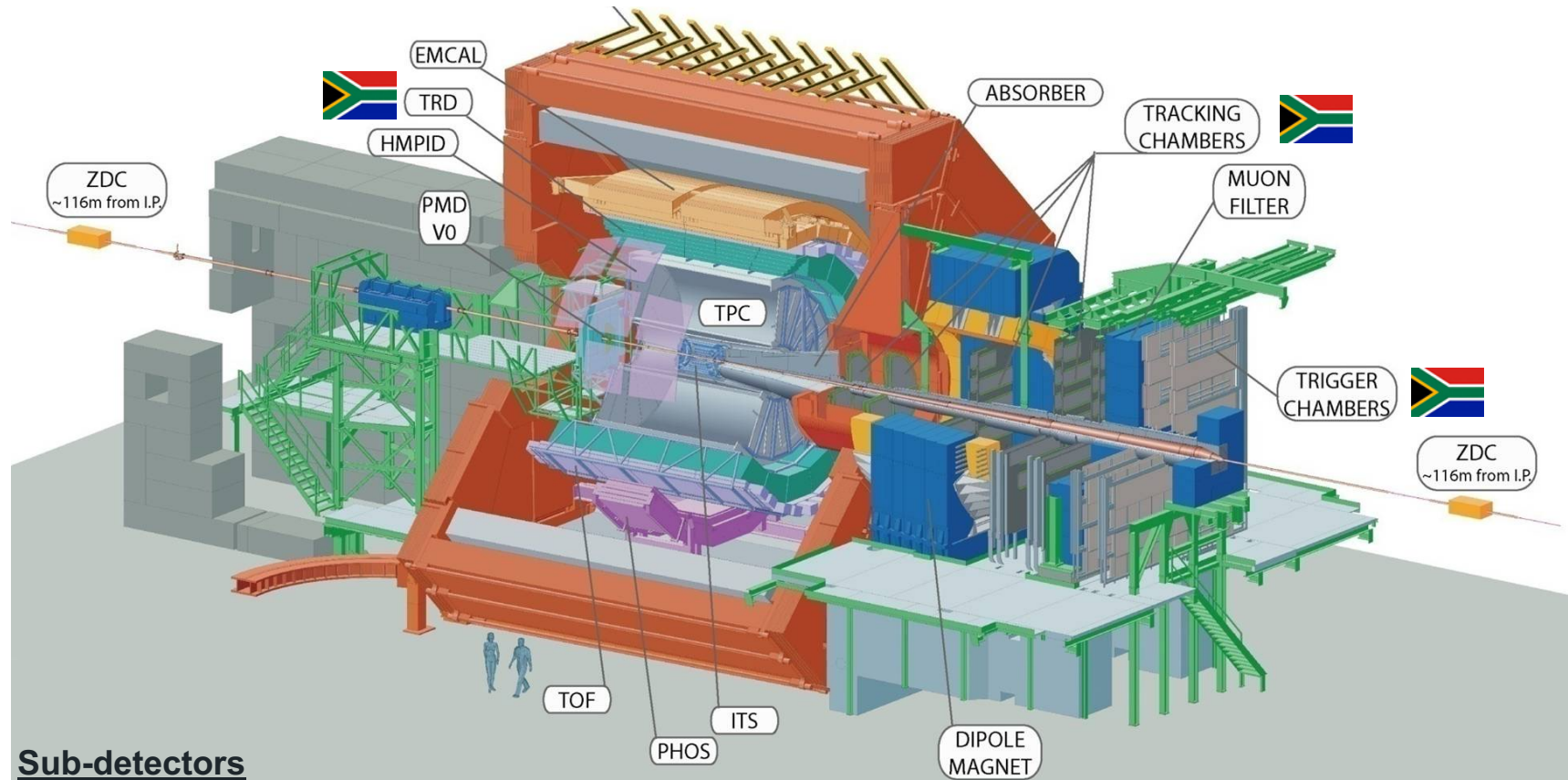
National coordinators are elected.

Chairman is elected by the national coordinators.



# SOUTH AFRICA IN ALICE

## MAINTENANCE AND OPERATIONS



### Sub-detectors

**Muons Tracking and Trigger**

**High Level Trigger (HLT)**

**→ Design and Commissioning of dimuon spectrometer HLT (dHLT)**

# W-Z Boson — Standard Model Test

- **pp**: PDFs at large Q<sup>2</sup>
- **p-Pb**: Cold Nuclear Matter (CNM) effects, nuclear PDFs
- **Pb-Pb**: binary collision/N<sub>coll</sub> scaling of hard processes



PUBLISHED FOR SISSA BY SPRINGER

RECEIVED: November 11, 2016  
ACCEPTED: February 4, 2017  
PUBLISHED: February 15, 2017

W and Z boson production in p-Pb collisions at  $\sqrt{s_{NN}} = 5.02$  TeV

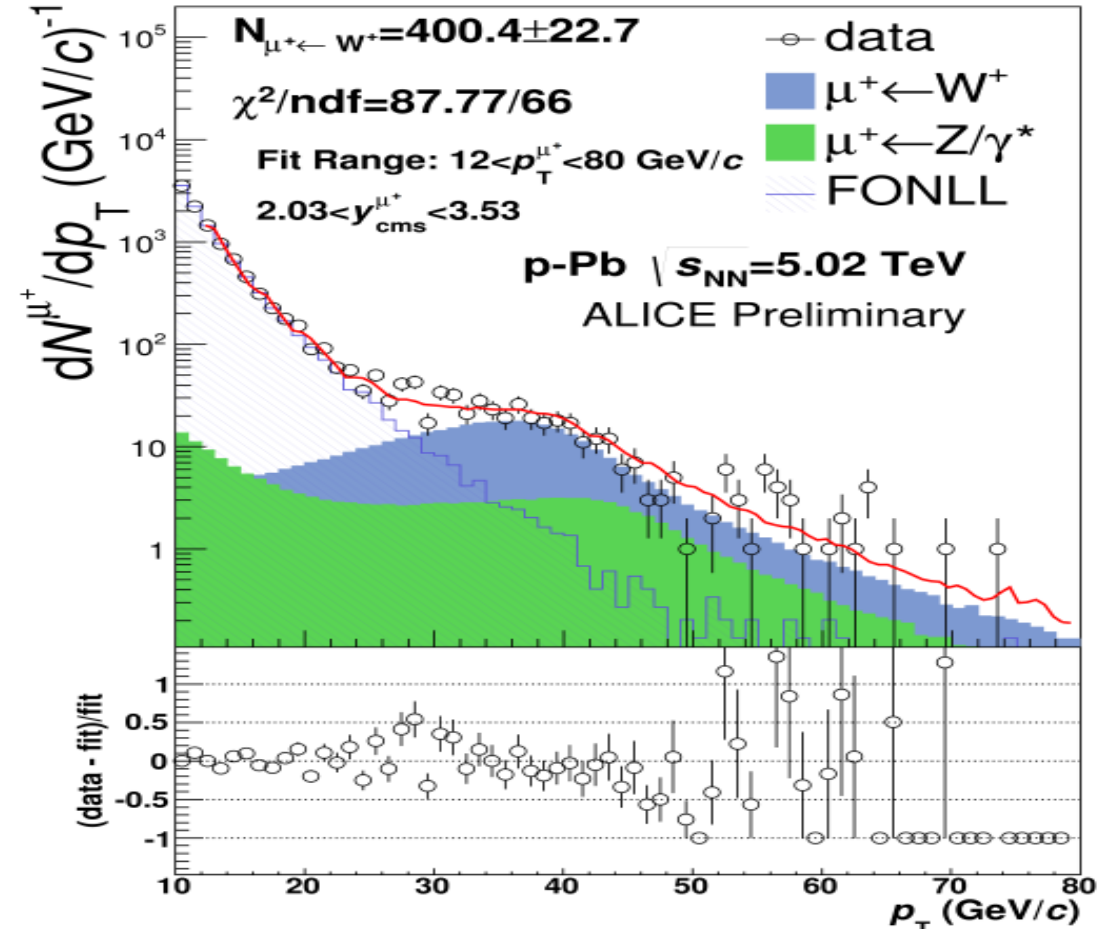


ALICE  
The ALICE collaboration

E-mail: [ALICE-publications@cern.ch](mailto:ALICE-publications@cern.ch)

ABSTRACT: The W and Z boson production was measured via the muonic decay channel in proton-lead collisions at  $\sqrt{s_{NN}} = 5.02$  TeV at the Large Hadron Collider with the ALICE detector. The measurement covers backward ( $-4.46 < y_{cms} < -2.96$ ) and forward ( $2.03 < y_{cms} < 3.53$ ) rapidity regions, corresponding to Pb-going and p-going directions, respectively. The Z-boson production cross section, with dimuon invariant mass of  $60 < m_{\mu\mu} < 120$  GeV/c<sup>2</sup> and muon transverse momentum ( $p_T^\mu$ ) larger than 20 GeV/c is

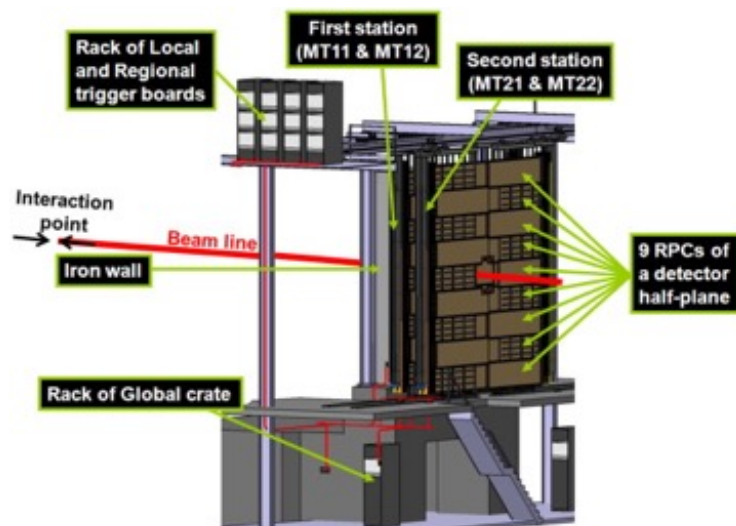
JHEP02(2017)077



ALI-PREL-82168

Work done in collaboration with Subatech (France) and Wuhan (China)

# Muon Tracking Detector



Testbeam for LS-2 Upgrade

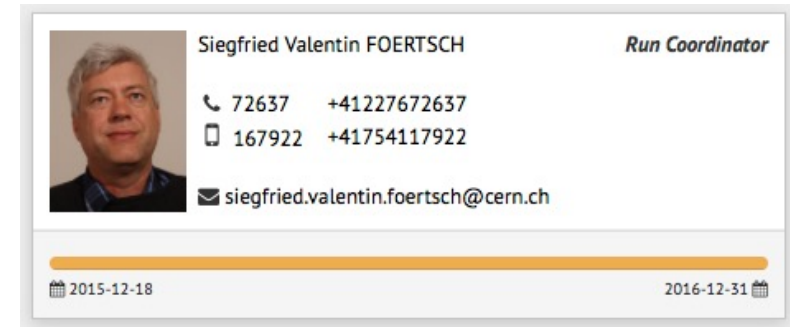


Detector Maintenance during LS-1

F Bossu, S Arkar, D Azmi, J Senosi, R Kuriakose,  
S Mhlanga, C Monteverdi, S Delsanto, D Steyn,  
Z Buthelezi, S Förtsch

# Run Coordination

- Siegfried Förtsch
  - Run Coordinator 2016
- Zinhle Buthelezi
  - Run Manager
  - April 2018: start of data taking
  - November 2018: start of Pb-Pb run

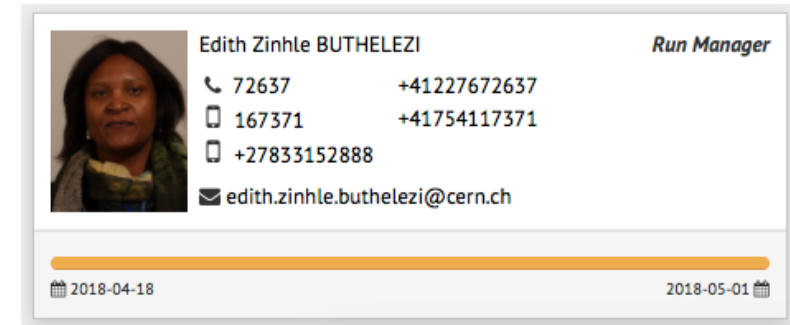


Siegfried Valentin FOERTSCH *Run Coordinator*

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siegfried.valentin.foertsch@cern.ch

2015-12-18 2016-12-31

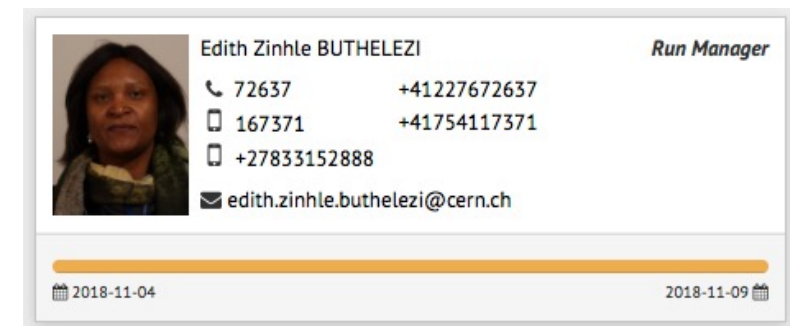


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+27833152888

edith.zinhle.buthelezi@cern.ch

2018-04-18 2018-05-01



Edith Zinhle BUTHELEZI *Run Manager*

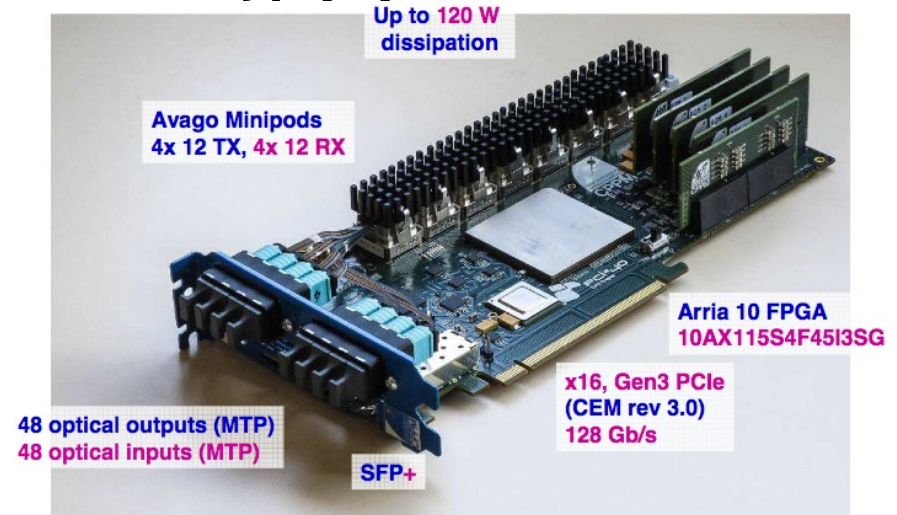
72637 +41227672637  
167371 +41754117371  
+27833152888

edith.zinhle.buthelezi@cern.ch

2018-11-04 2018-11-09

# Muon Arm Upgrade

- Muon Identifier: Common Readout Unit  
On-The-Fly Data Conditioning @ 51.2Gb/s

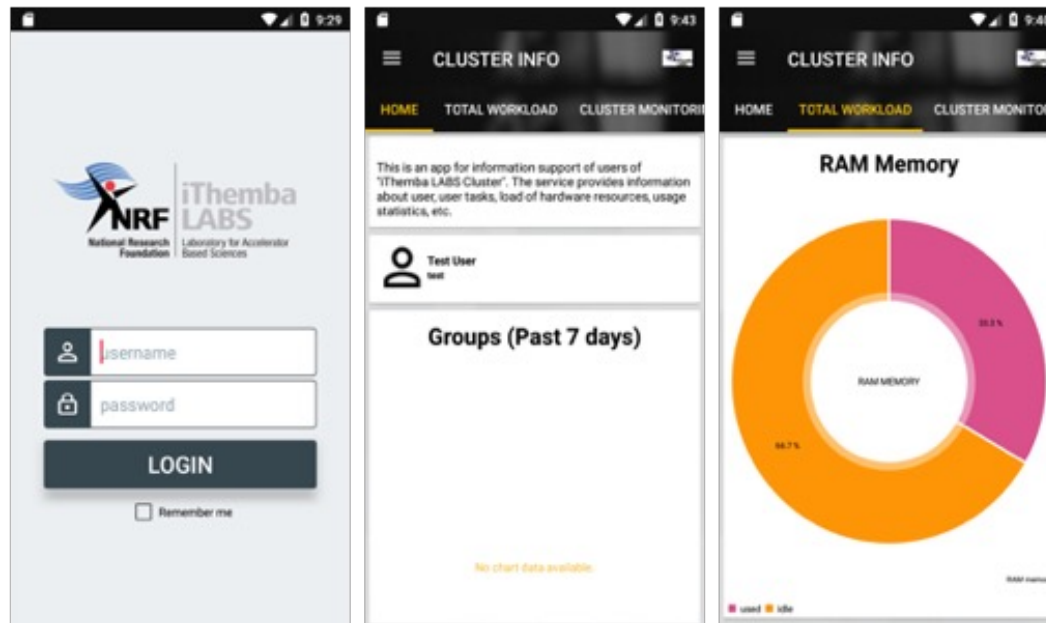
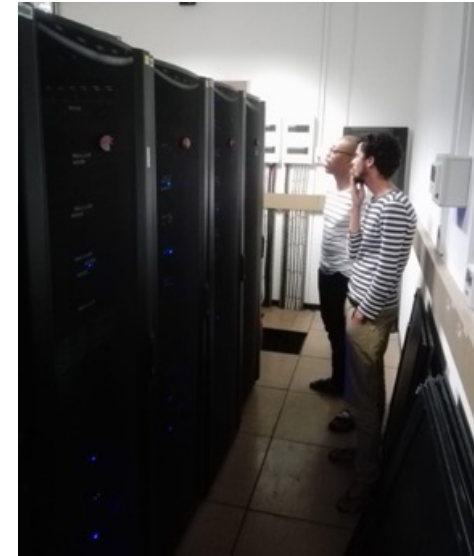


- Muon Tracking:  
Low-Voltage System



# Tier-3 Cluster at iThemba LABS

- 264 CPUs / 528 GB RAM / 32 TB HDD
- used by ALICE + other iThemba groups
- re-using donated Run-1 HLT
- Android monitoring app



# THE ALICE GRID IN SOUTH AFRICA

ALICE grid requires sustained and reliable bandwidth to sites.

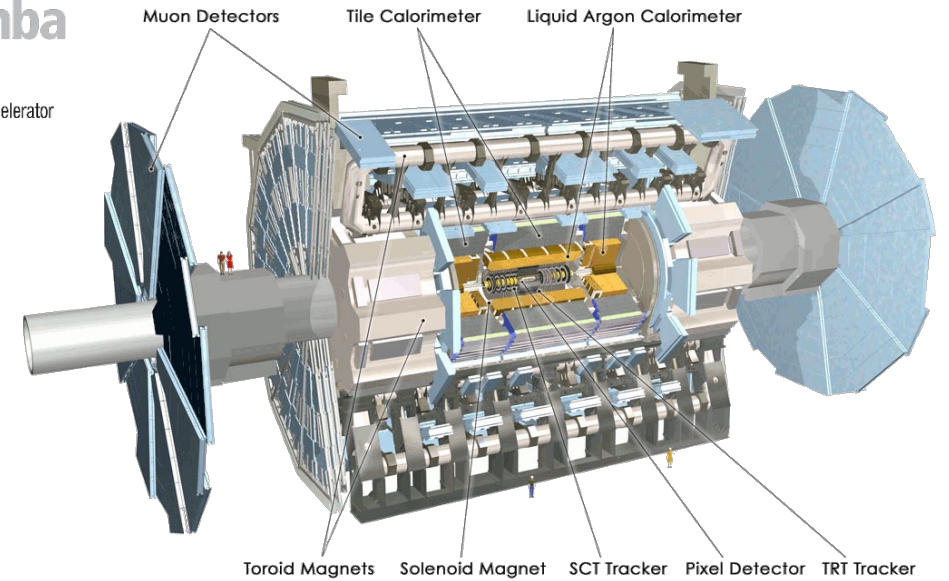
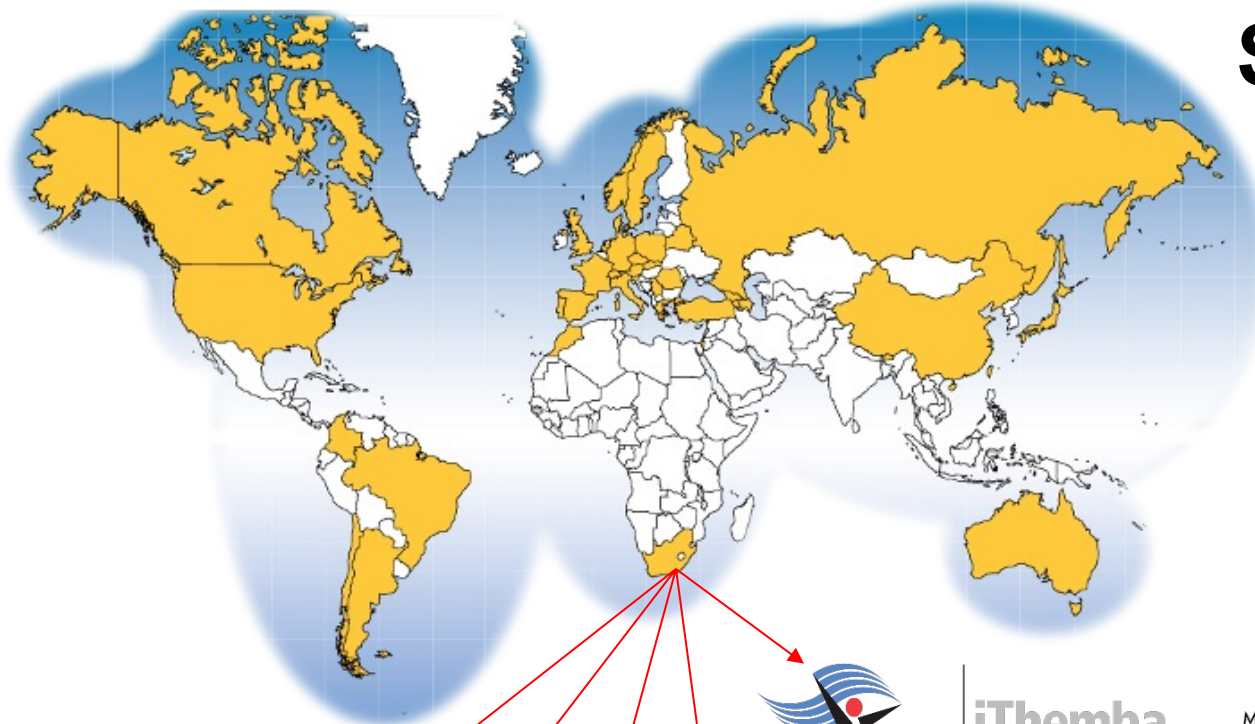
Despite lack, SA Tier2 was processing MC data 2004-2009 with contributed resources.

ALICE monitoring system has been used to “map” the effect of SANReN:

- National network performs very well,
- International bandwidth no longer limiting factor.



# SA@ATLAS



# ATLAS (as of a few years ago...)

- Principal Scientists \*:
  - Simon Connell (UJ),
  - **Bruce Mellado** (Wits), Deepak Kar (Wits), Xifeng Ruan (Wits),
  - Elias Sideras-Haddas (Wits), Sahal Yacoob (UCT)
- Post-doctoral fellows:
  - Debarati Roy (Wits), Yesenia Hernandez (Wits), Loan Truong (UJ),
  - Robert Reed (Wits)
- Students: 26 MSc and PhD Students

\* List not exhaustive

# ISOLDE

## Principal Scientists:

- **Krish Bharuth-Ram** (KwaZulu-Natal and Durban U. of Technology)
- Hillary Masenda & Deena Naidoo (Witwatersrand),
- Nico Orce & S. Trambak (Western Cape)
- Mathis Wiedeking (iThemba LABS), Christine Steenkamp (Stellenbosch)

## • Post-doctoral fellows:

- S. Sithole (Stellenbosch and iThemba LABS) Kumar Raju (Western Cape)

## • Students: 13 MSc and PhD students



# Outreach

## International masterclass April 2014

- **“Looking for Strange Particles in ALICE”**
- ❖ Organized by SA-ALICE and iThemba LABS
- ❖ 24 students from 6 schools
- ❖ International video discussion with Cairo, Geneva & Warsaw



## 2015: CERN Beamline for Schools Competition St. John's College and Barnato Park High School team



Simon H. Connell (University of Johannesburg)

# Conferences in South Africa

June 2004	ALICE Dimuon Workshop
September 2004	Strange Quark Matter
December 2008	Inauguration of SA-CERN
December 2010	First Kruger Workshop
2012	ALICE Dimuon Workshop
December 2012	Kruger Workshop
November 2013	Hard Probes
December 2014	Kruger Workshop
January 2015	Chris Engelbrecht Summer School
December 2016	Kruger Workshop
December 2018	Kruger Workshop
April 2019	ALICE Muon Week



# THE BIENNIAL AFRICAN SCHOOL OF FUNDAMENTAL PHYSICS AND APPLICATIONS (ASP)

- **International Organizing Committee (IOC)**
  - The IOC is the main organizer of the school
  - The IOC Members:
    - B. Acharya (ICTP & King's College London)
    - K. A. Assamagan (BNL)
    - A. Dabrowski (CERN)
    - C. Darve (ESS)
    - J. Ellis (CERN & King's College London)
    - S. Muanza (CNRS-IN2P3)
    - R. Voss (CERN)
- **Local Organizing Committee (LOC)**
  - Local committee in the host country
- **International Advisory Committee (IAC)**
  - Advises on various aspects of the organization including fund raising



<http://www.africanschoolofphysics.org/>





**WITS**

# Thank YOU

South Africa and CERN  
for making this possible.



**UCT**



**iThemba LABS**

