



Dark Matter Searches at LNF with the PADME detector

F. Ferrarotto for the PADME Coll.

INFN Sez. Roma 1

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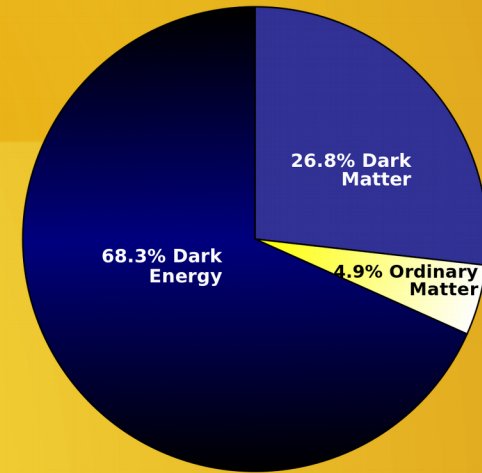


Outline

- A short introduction to the Dark Photon model and its research @PADME
- **P**ositron **A**nnihilation into **D**ark **M**atter **E**xperiment
 - Outlook of Runs I and II and first physics results
- The “ ^8Be anomaly” @ ATOMKI
- The X17 particle resonant search @PADME
 - The Run III
- Conclusions

The Dark Matter issue

From Cosmological and Astrophysical observations of gravitational effects, **something else than ordinary (baryonic) matter should exist.** The abundances of this new entities : **dark matter** and **dark energy** are much larger than SM matter.

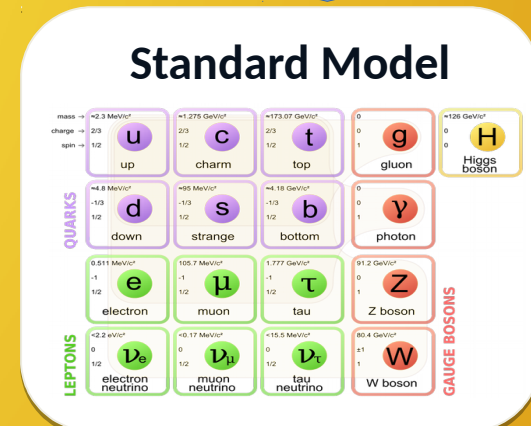
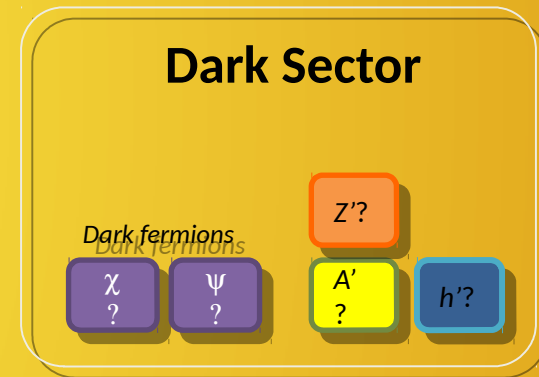


Dark Matter should manifest also in experiment at accelerators ... but up to now **NO clear experimental observation** both at LHC (WIMPs) and at dedicated experiments.

One class of simple models just adds an **additional U(1)** “hidden” symmetry to the SM, with its corresponding **massive neutral boson** : the so-called **“Dark Photon” (A’)**

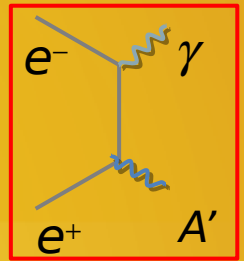
$$U(1)_Y + SU(2)_{\text{Weak}} + SU(3)_{\text{Strong}} [+U(1)_{A'}]$$

A’ could itself be the mediator between the visible and the dark sector. The effective interaction between the fermions and the dark photon is parametrized in term of a factor ϵ representing the mixing strength.

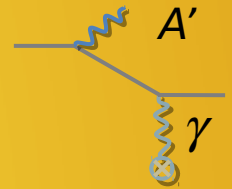


A' production and decay

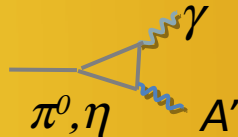
A' production



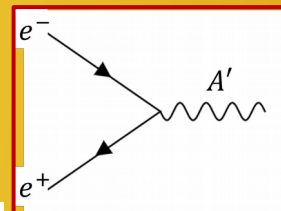
Associated



Bremsstrahlung



Meson decay



Resonant

The effective interaction that can be studied is

$$L \sim g' q_f' \Psi (\gamma_\mu + \alpha'_a \gamma_\mu \gamma^5) \Psi A'^\mu, \quad \text{usually } \alpha'_a = 0 \quad \text{QED-like}$$

Also possible a General U'(1) and kinetic mixing with B(A', Z')

A' couples to SM hypercharge through kinetic mixing operator, acquiring a (small) SM charge

- Universal coupling proportional to the q_{em}
- Just one single additional parameter – ϵ

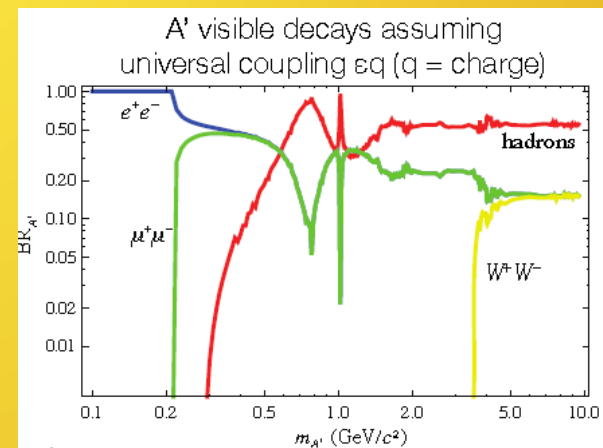


The **associated and resonant** productions can be studied **ONLY** at a positron-beam facility

The **A' decay** can be **visible** or **invisible** if any DM particles exists with $m_{DM} < m_{A'}/2$

YES : $A' \rightarrow$ DM **invisible decays** with (likely) $BR \approx 1$
SM decays suppressed by a factor ϵ^2

NO : $A' \rightarrow$ SM **visible decays** - BR depends on $m_{A'}$



Status of dark photon searches

Positron-based experiments production :

A'-strahlung

Associated production $e^+e^- \rightarrow A'(\gamma)$

Resonant production $e^+e^- \rightarrow e^+e^-$

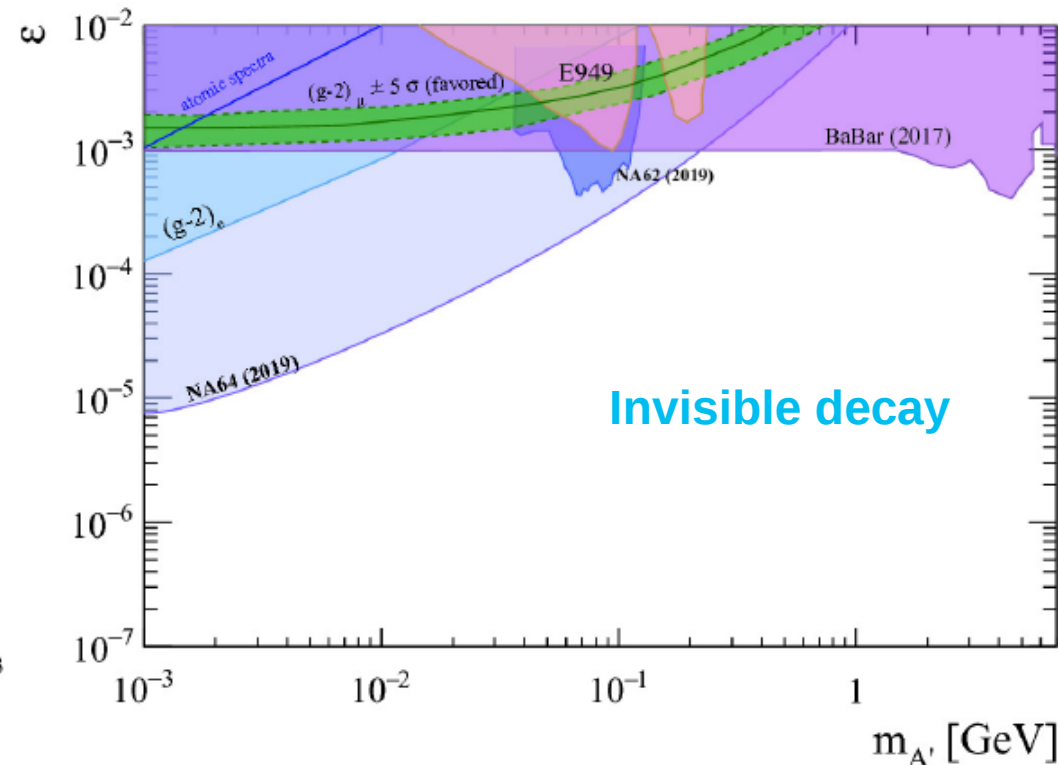
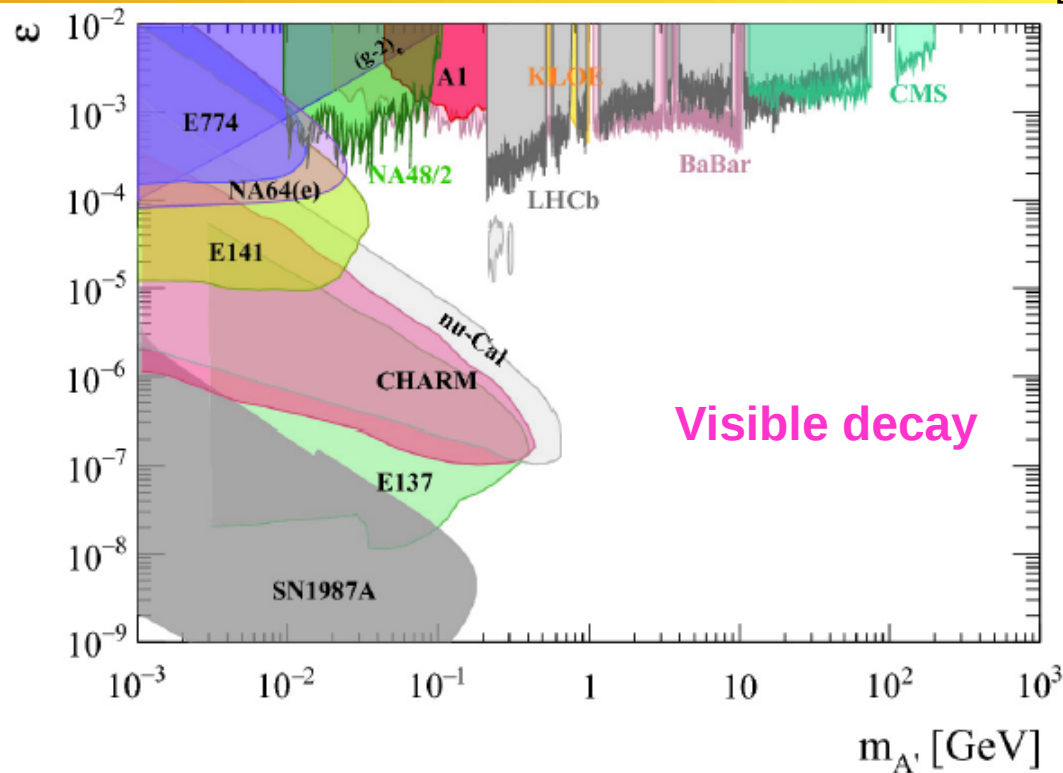
Visible decays : $A' \rightarrow e^+e^-, \mu^+\mu^-$

Thin target : searching bumps in $\ell\ell$ invariant mass

Invisible decays : $A' \rightarrow \chi\chi$

Missing mass : $e^+e^- \rightarrow A'(\gamma)$ search for invisible particle using kinematics

arXiv:2104.10280v1 [hep-ph] 20 Apr 2021



The PADME Experiment

The search for this new mediator A' (invisible decay) is the main goal of the PADME experiment

2016 : INFN approved an experiment at the Linac Beam Test Facility (BTF) at INFN Frascati National Laboratories (LNF) near Rome :

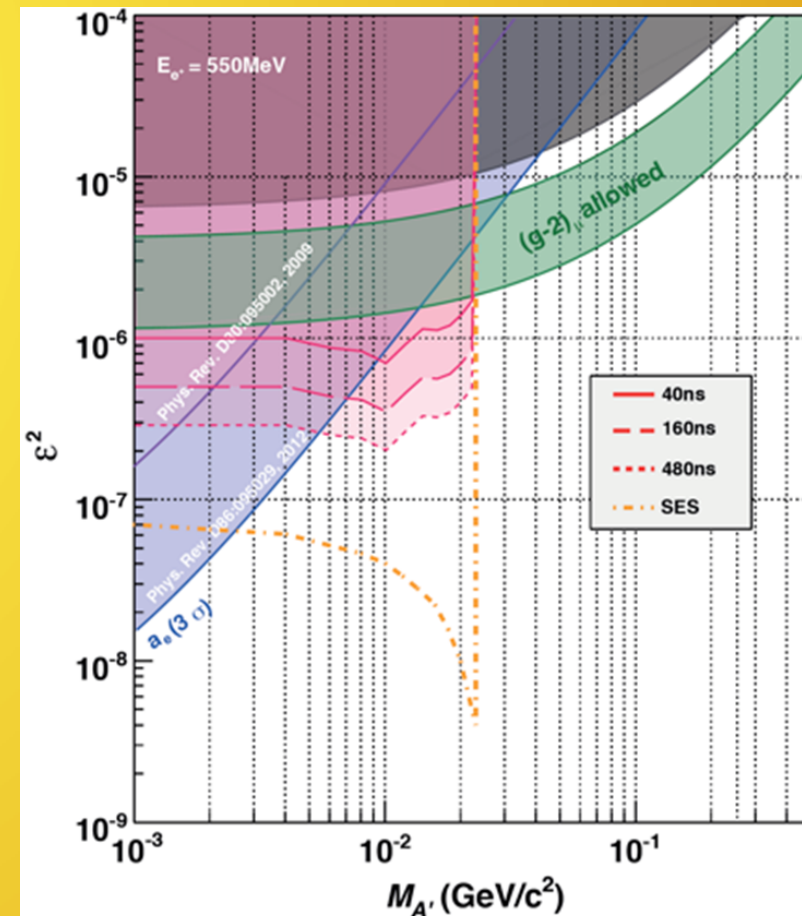
PADME (Positron Annihilation into Dark Matter Experiment)

Invariant Missing Mass peak search over a continuous background

PADME can explore in a model independent way the region down to $\epsilon \approx 10^{-3}$

$m_{A'} < 23.7 \text{ MeV}$ ($E_{\text{beam}} = 550 \text{ MeV}$ - LNF Linac)

INFN Roma, INFN Frascati, INFN Lecce, La Sapienza University, Politecnico di Torino e INFN Sezione di Torino, MTA Atomki Debrecen, University of Sofia, Cornell University, US William and Mary College, Princeton Univ.



The PADME detector

Positron beam of ~ 500 MeV/c@50 Hz

Macro-bunches max length $\Delta t < 300$ ns

Number of annihilations proportional to

$$N_{beam}^{e^+} \times N_{target}^{e^-}$$

Limited intensity (pile-up) $< 3 \times 10^4$ PoT/pulse

Active polycrystalline diamond target

2x2 cm – 100 μ thick

x,y graphitized strips r/out

Beam size, position, time, N_{e^+}

1 m **dipole magnet (0.5-0.6 T)** to :

Sweep away non-interacting positrons

Tag positrons losing energy by Bremsstr

Scintillating bar veto detectors placed

inside vacuum vessel – r/out SiPM

Positron and electron detection inside

magnetic gap

Additional **veto** for e^+ irradiating soft γ

(near beam exit)

BGO EM Calorimeter (ECAL)

616 21x21x230 mm³ BGO - r/out PMT

$\approx 20.5 X_0$ depth

Cylindrical shape with central hole (Bremsstr)

E, Θ , time measurement

Small angle EM Calorimeter (SAC)

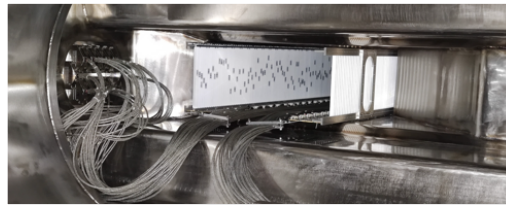
25 30x30x140 mm³ PbF₂ - r/out PMT

E, Θ , time measurement

Silicon pixel Beam Monitor (TimePix3)

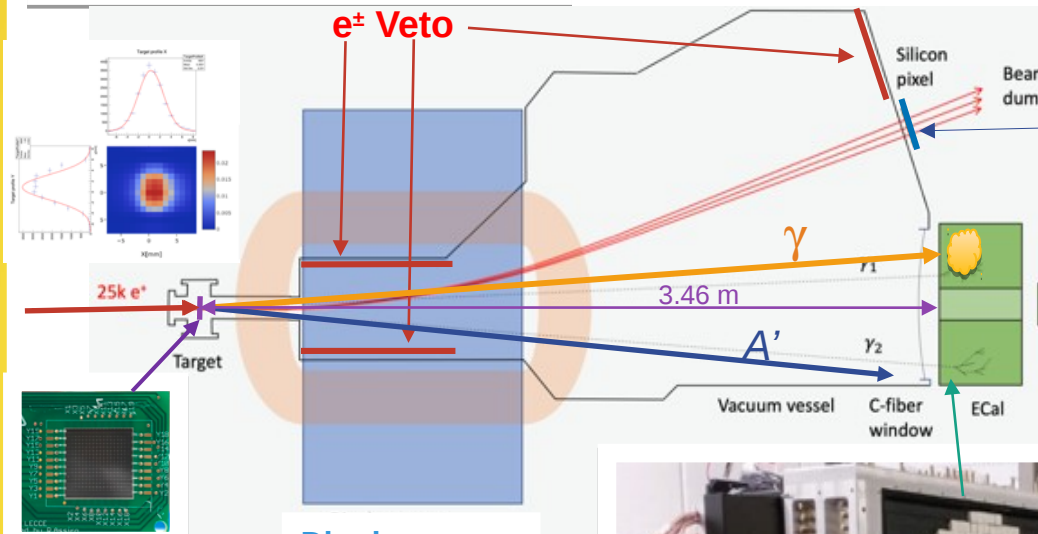
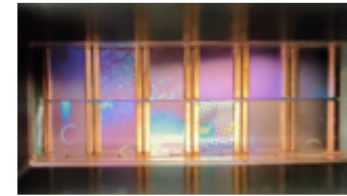
used to tag exiting positrons

(E), x, y, time measurement



“Golden signal” event :
1 single γ in EM calo
 Nothing in all other
 components in ± 2 ns

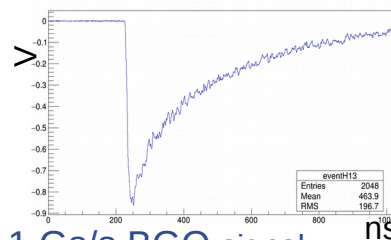
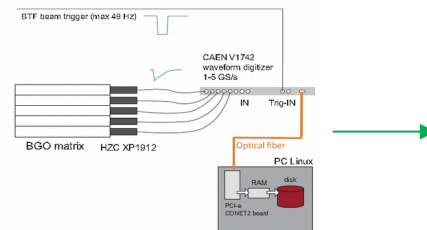
TimePix3



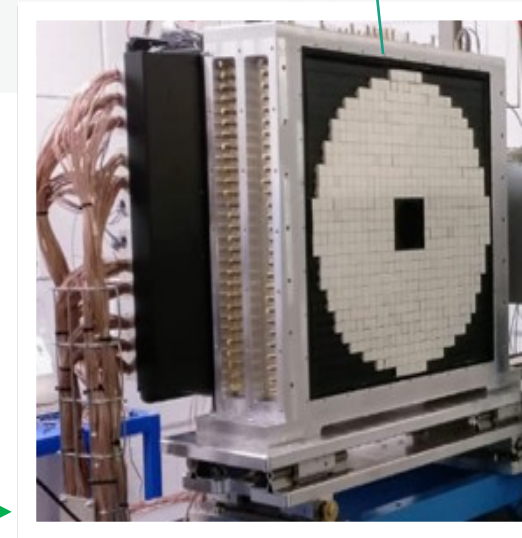
SAC Calo PbF₂

diamond target

Dipole magnet



1 Gs/s BGO signal
 Digitized by CAEN
 V1742



BGO EM Calo
 (ECAL)

$\sigma(E)/E \sim 2\%/\sqrt{E}$
 $\sigma(\theta) < 1$ mrad
 Timing : < 1 ns from signal shape fit
 Linearity OK up to \sim GeV

Run I and Run II data taking

2 runs in 3 configurations between September 2019 and December 2020 (during the pandemics)

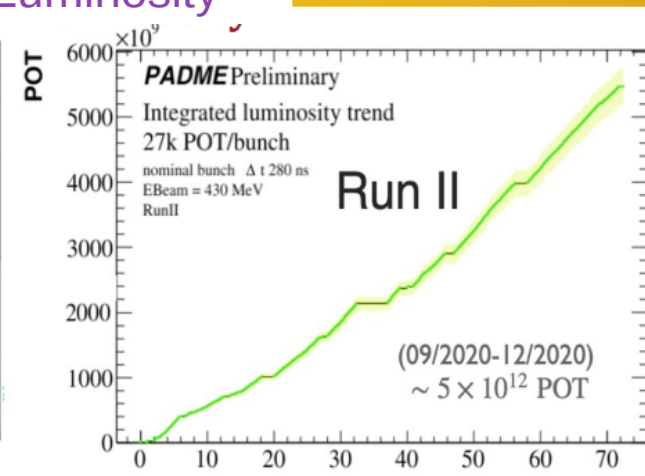
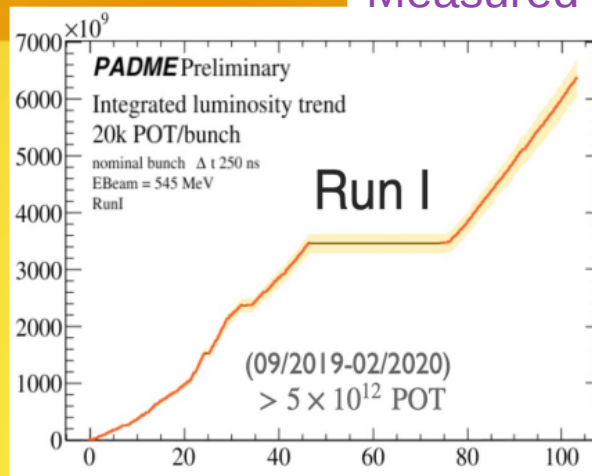
Acquired luminosity measurements :

Run I → 6×10^{12} PoTs

Run II → 5.5×10^{12} PoTs

Luminosity precision : 5%

Measured Luminosity



Changes between the runs :

Run Ia : secondary beam → Run Ib : primary beam

Reduced Background

Beam energy reduced 545 → 490 MeV

Detailed MC simulation of beamline ([JHEP 09 \(2022\), 233](#))

Run Ib → Run II : **changed vacuum separation window**

Reduced background from vacuum window

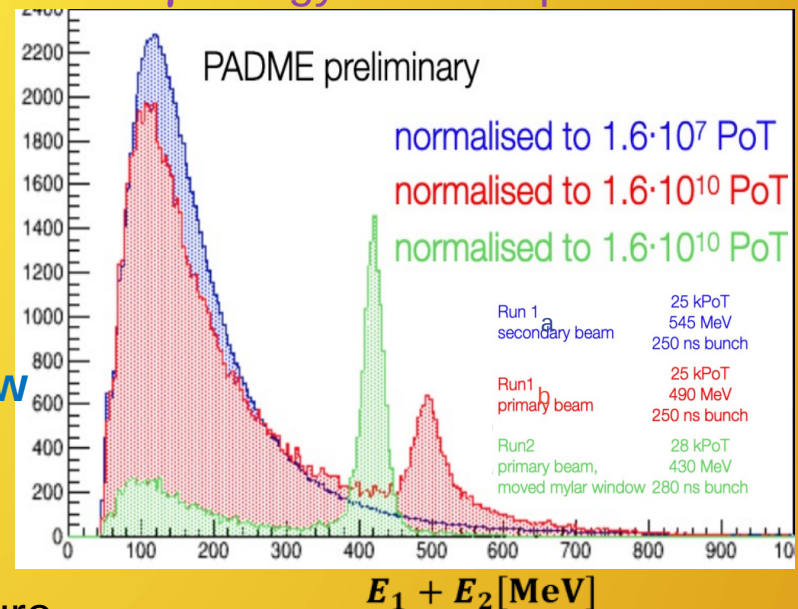
Beam energy reduced 490 → 430 MeV

More PoTs/bunch (20 K → 27 K)

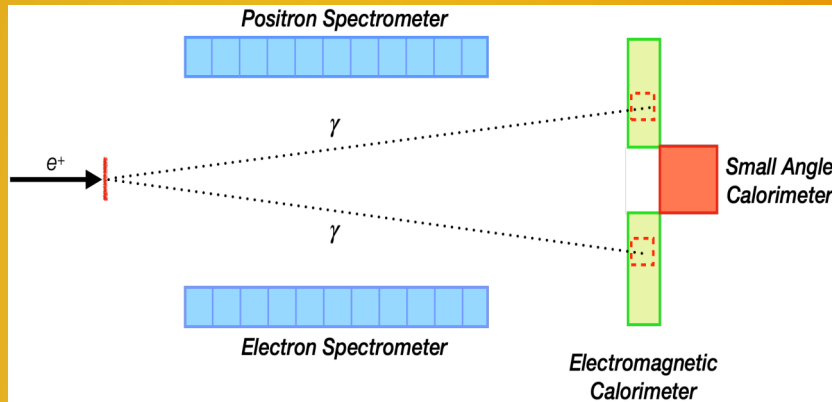
Longer bunches (250 → 280 ns), more stable structure

→ reduced the pileup in the detector

2 γ energy in ECAL per event



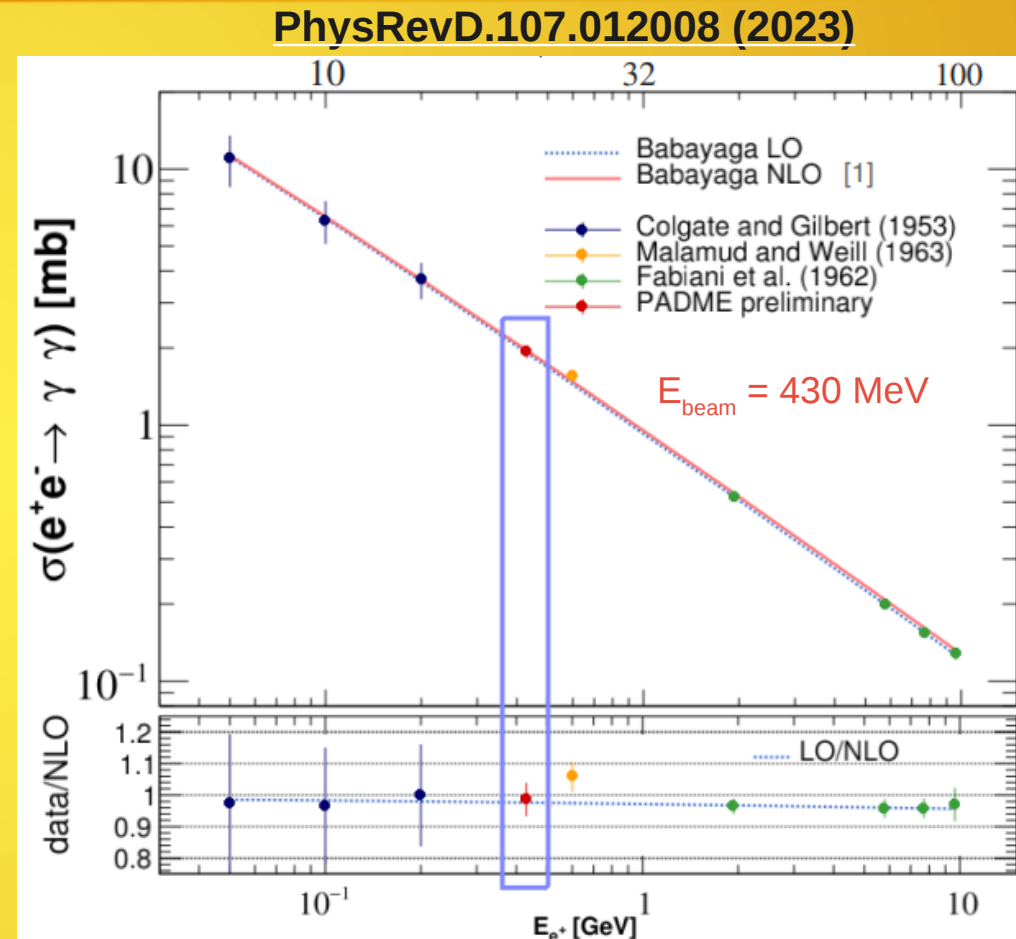
First physics measurement: multi photon annihilation



From PADME Run II (10 % of 2020 data set) :

- Characterisation of ECAL
- Could be sensitive to sub-GeV new physics (e.g. ALPs, ...)

First direct measurement below 500 MeV with ~ 5 % precision (both Gilbert '53 and Malamud '63 measure e^+ disappearance rates)



PADME : $\sigma(e^+e^- \rightarrow \gamma\gamma(\gamma)) = 1.977 \pm 0.018_{\text{stat}} \pm 0.045_{\text{syst}} \pm 0.110_{\text{(n.collisions)}} \text{ mb}$

QED@NLO : $\sigma(e^+e^- \rightarrow \gamma\gamma(\gamma)) = 1.9573 \pm 0.0005_{\text{stat}} \pm 0.0020_{\text{syst}} \text{ mb}$

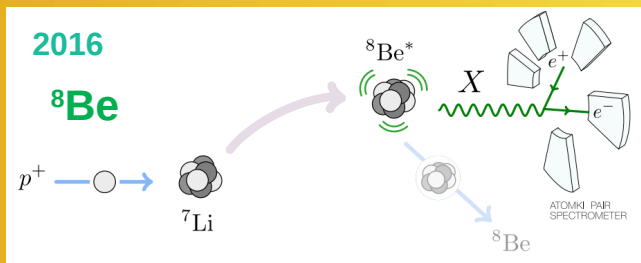
Phys. Lett. B 663 (2008) 209-213

A fundamental step towards the invisible dark photon analysis (ongoing..)

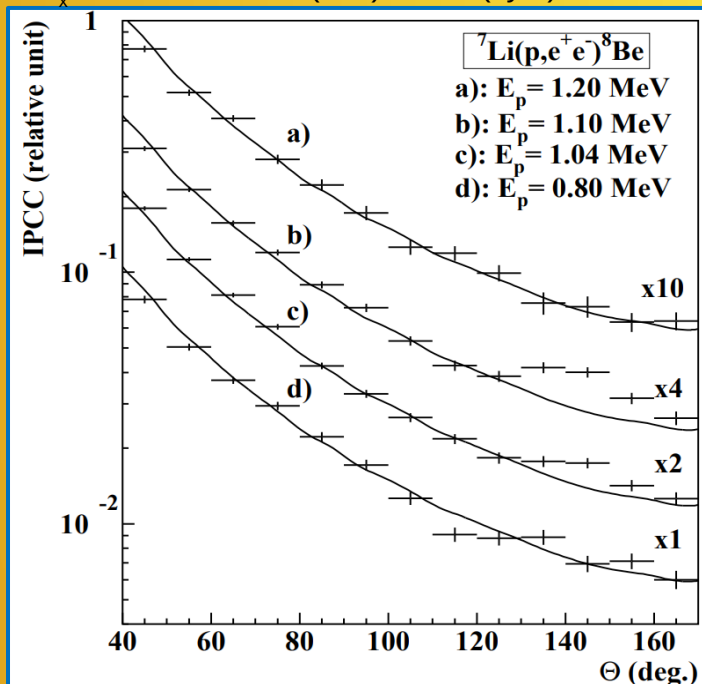
The “ ^8Be anomaly”

Collaboration at ATOMKI institute in Hungary studying IPC decays of excited nuclei in 3 different experiments : ^8Be (2016) / ^4He (2020) / ^{12}C (2022)

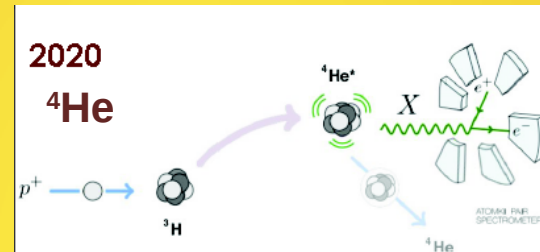
- In **all** 3 experiments found **anomaly** compatible with new particle of ~ 17 MeV mass
- Statistical significance very strong : nearly 7σ for each experiment



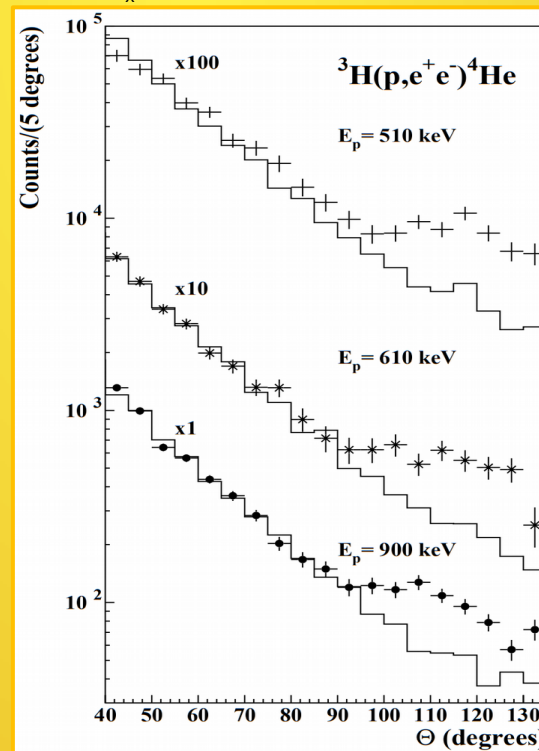
$$M_x c^2 = 17.03 \pm 0.11(\text{stat}) \pm 0.20(\text{syst}) \text{ MeV}$$



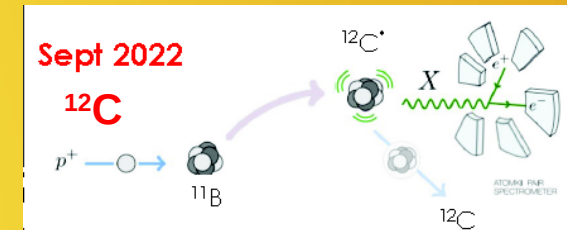
PhysRevLett.116.042501



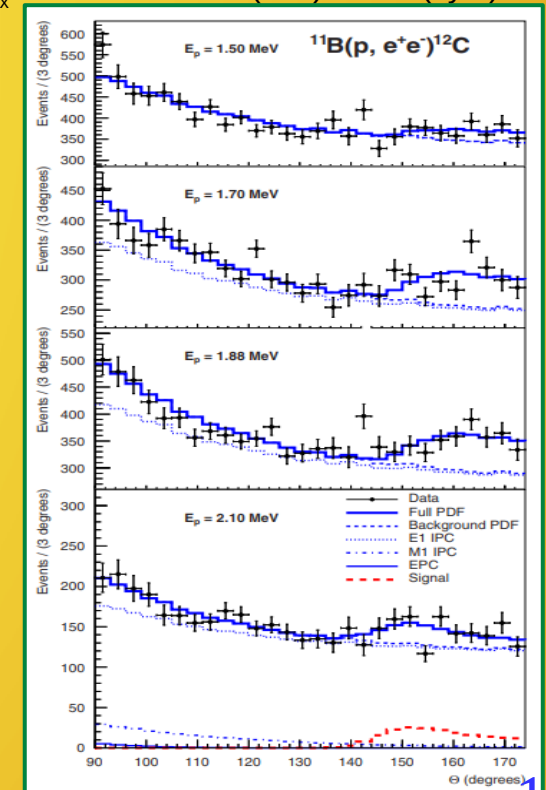
$$M_x c^2 = 17.01 \pm 0.16(\text{tot}) \text{ MeV}$$



Phys.Rev.C, 104(4):044003



$$M_x c^2 = 16.98 \pm 0.16(\text{stat}) \pm 0.20(\text{syst}) \text{ MeV}$$



Phys. Rev. C 106, L061601

The X17 particle

From the ATOMKI observations, the main properties of the **new X₁₇ particle** are :

$$M_{X17} \sim 17 \text{ MeV} \quad \text{proto-phobic}$$

The X17 hypothesis is **kinematically** consistent for all the observed anomalies.

Many proposals for SM explanations, but, in conclusion, no compelling SM explanation so far.

The spin-parity selection rules $J_* = L \oplus J_o \oplus J_x$ and $P_* = (-1)^L P_o P_x$ are required to identify the nature of the new mediator

From the new ¹²C results preferred assignments are a **vector** or an **axial-vector** particle and seem to exclude a scalar or pseudoscalar one.

Phys.Rev.D 102 (2020) 3, 03601

6

N_*	J_*^P	Scalar X17	Pseudoscalar X17	Vector X17	Axial Vector X17
⁸ Be(18.15)	1 ⁺	✗	✓	✓	✓
¹² C(17.23)	1 ⁻	✓	✗	✓	✓
⁴ He(21.01)	0 ⁻	✗	✓	✗	✓
⁴ He(20.21)	0 ⁺	✓	✗	✓	✗

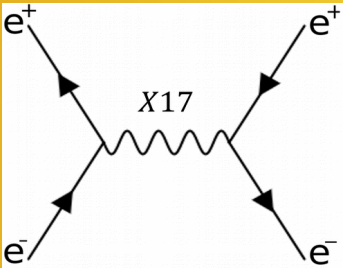
Search for X17 using resonant production on thin target



Planned for 2022 a **dedicated Run** of PADME to study the X17 particle

Idea : use **resonant** production and search for **visible** X_{17} decay into e^+e^-

PADME@LNF is actually the **only** facility in the world capable to do this measurement



$$\sigma_{res} \propto \frac{g_{Ve}^2}{2m_e} \pi Z \delta(E_{res} - E_{beam})$$

The **resonant** production scales only with Z and it's **much larger** than the associated and radiative production

To exploit **resonant** production the center of mass energy should be **as close as possible** to the expected mass : $E_{res} = M_{X17}^2 / m_e \rightarrow$ A **scanning** procedure is needed

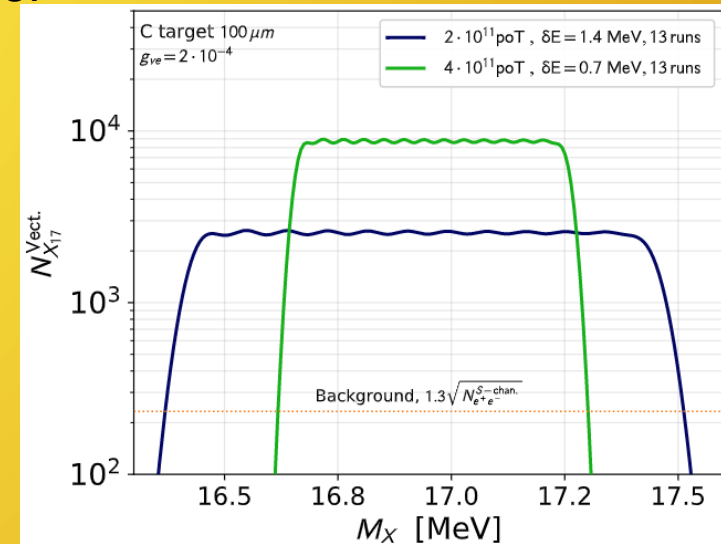
Darmé et al. Phys. Rev. D 106,115036 : **analysis strategy** - vary the beam energy, fit the background, calibrate the luminosity and look for resonance.

$$N_{X17}^{perPoT} \simeq \frac{g_{Ve}^2}{2m_e} \ell_{tar} \frac{N_A \rho Z}{A} f(E_{res}, E_{beam})$$

The resonance shape is exactly the one of the beam energy distribution

$f(E_{res}, E_{beam})$ is the beam spread : **gaussian** distribution with **spread** δ_E

Thousands of events with just **1E11 PoT**



The expected Standard Model background

The main backgrounds are from **Bhabha scattering** and $\gamma\gamma$ production.

They can be fitted directly from data.

$\sigma_{Bhabha} = \sigma_{s\text{-ch}} + \sigma_{t\text{-ch}}$ processes are simulated only at LO

X17 production mechanism is assumed to have the same acceptance of Bhabha s-channel

For : $N^{\text{PoT}} = 1 \times 10^{11}$, $E_{\text{beam}} = 282 \text{ MeV} \rightarrow \sqrt{s} \sim 17 \text{ MeV}$

Cuts on both final state particles :

Azimuthal angle : $25.5 \text{ mrad} < \Theta_{1,2} < 77 \text{ mrad}$

Final energy $E_{1,2} > 100 \text{ MeV}$

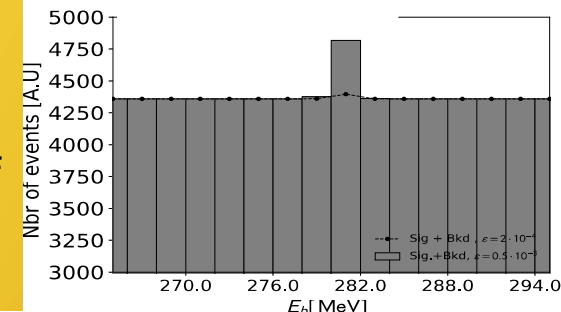
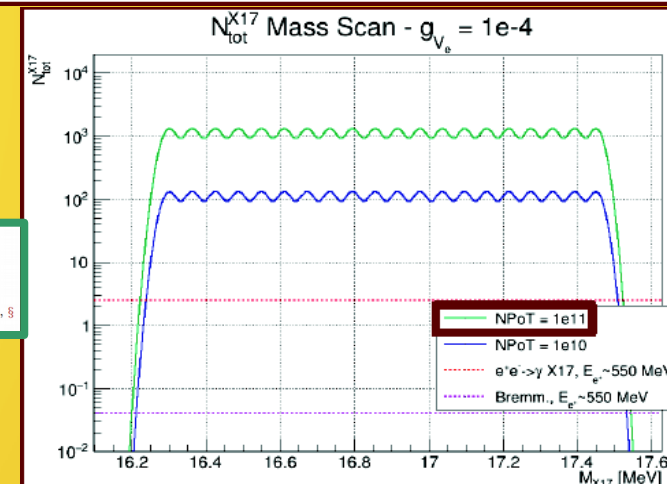
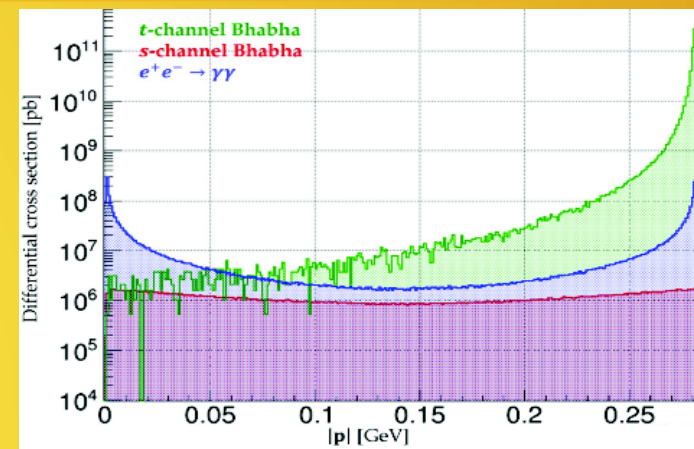
Assume detector efficiency $\sim 100 \%$

$$g_{V_e} = 2 \times 10^4 \text{ and } \delta E = 1.4 \text{ MeV}$$

Resonant search for the X17 boson at PADME

Luc Darmé,^{1,*} Marco Mancini,^{2,†} Enrico Nardi,^{3,‡} and Mauro Raggi^{4,§}

Phys. Rev. D 106, 115036



BG process	No. of Ev.	No. of Ev. in Acc.	Acc.
$e^+e^- \rightarrow e^+e^-$ (t-ch.)	5.4×10^7	6.9×10^4	0.13%
$e^+e^- \rightarrow e^+e^-$ (s-ch.)	3.2×10^4	6.4×10^3	20%
$e^+e^- \rightarrow \gamma\gamma$	2.9×10^5	1.3×10^4	4.5%
$e^+e^- \rightarrow X_{17} \rightarrow e^+e^-$	1250	250	20%

Resonant Signal should emerge on top of **Bhabha BG** in one or more points of the scan.

Expected limits

Challenge : achieve an extremely precise luminosity measurement and systematic errors control (<1%)

Order 10^{10} PoT per each scan point

Under these assumptions, we aim to set limits both on:

Vector model, covering almost the entire free parameter space

Pseudoscalar model, in the case of an ALPs decaying into leptons only.

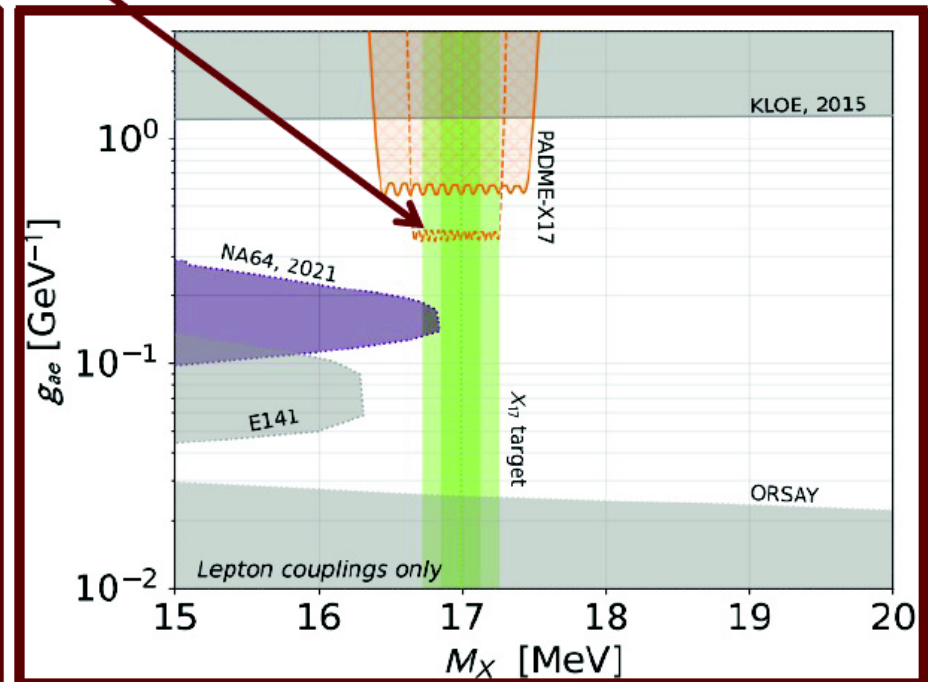
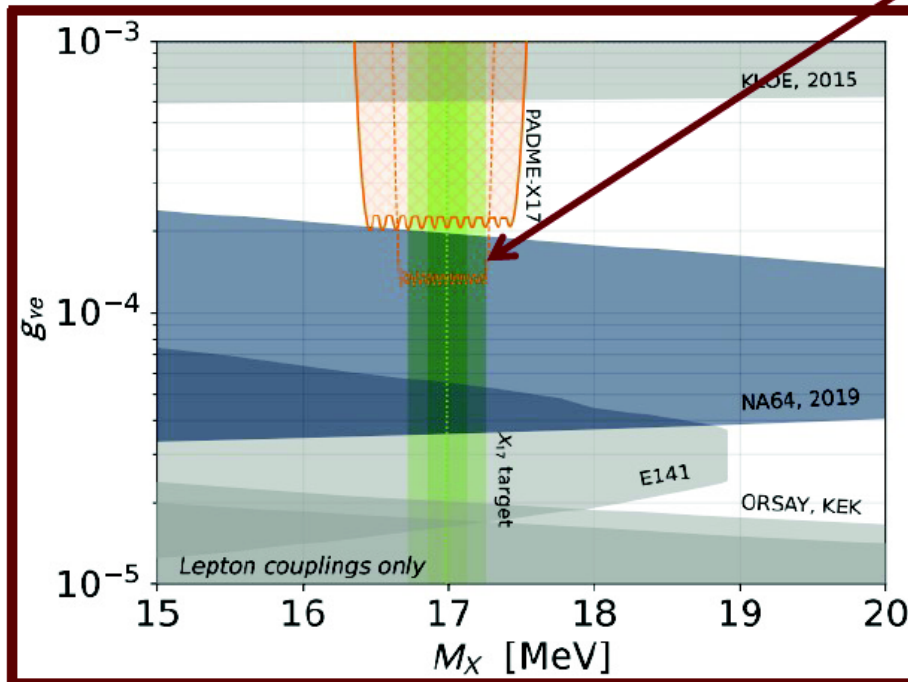
PADME maximum sensitivity is in the **vector** case

PADME limit aim

[Phys.Rev.D 106 \(2022\) 11, 115036](https://arxiv.org/abs/2108.08111)

Vector X17

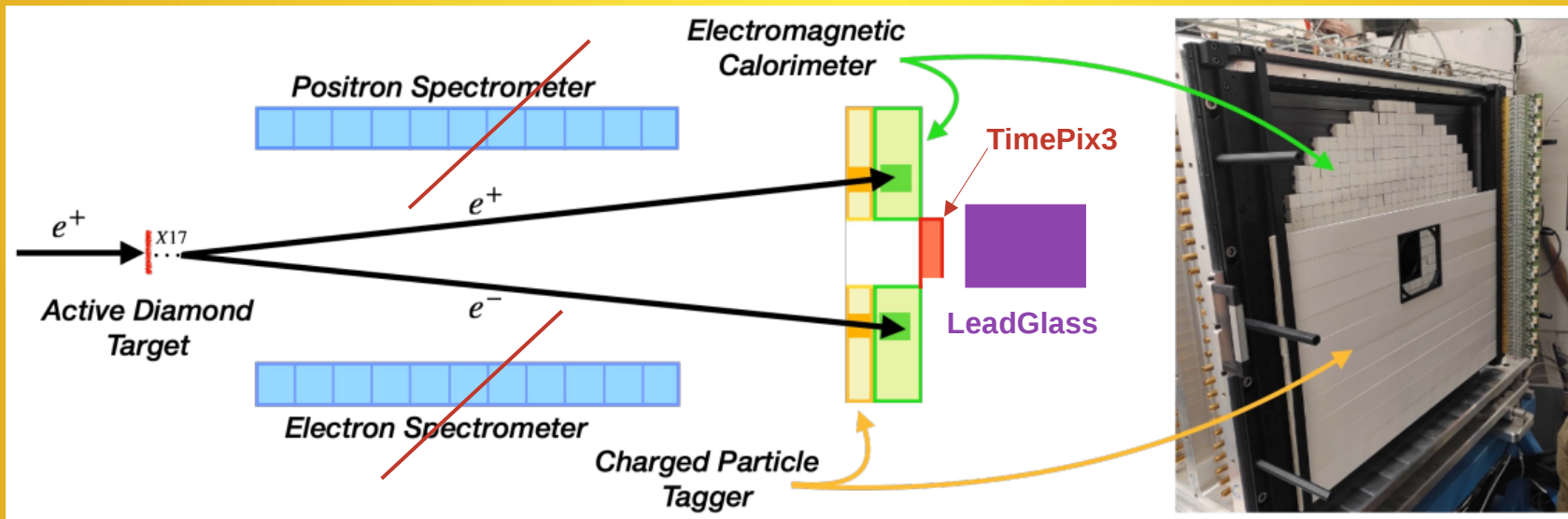
Pseudo scalar X17



The Run III experimental setup

➔ Improvements to the PADME set-up are required for the X17 resonant search !

- Taking confidence from $e^+e^- \rightarrow \gamma\gamma$ measurement we're looking for X17 with the **ECAL**
 → **NO magnetic field** to get both final state particles in **ECAL**
- To distinguish e^+e^- from $\gamma\gamma$: **ETagger** detector - 5 mm thick plastic **scintillators** – r/out SiPM
 Increased **target-ECAL** distance → changes acceptances
- Removed the **SAC** and installed back of hole the **TimePix3 Beam monitor** and a **LeadGlass Detector** with PMT readout (**Luminosity monitors**)



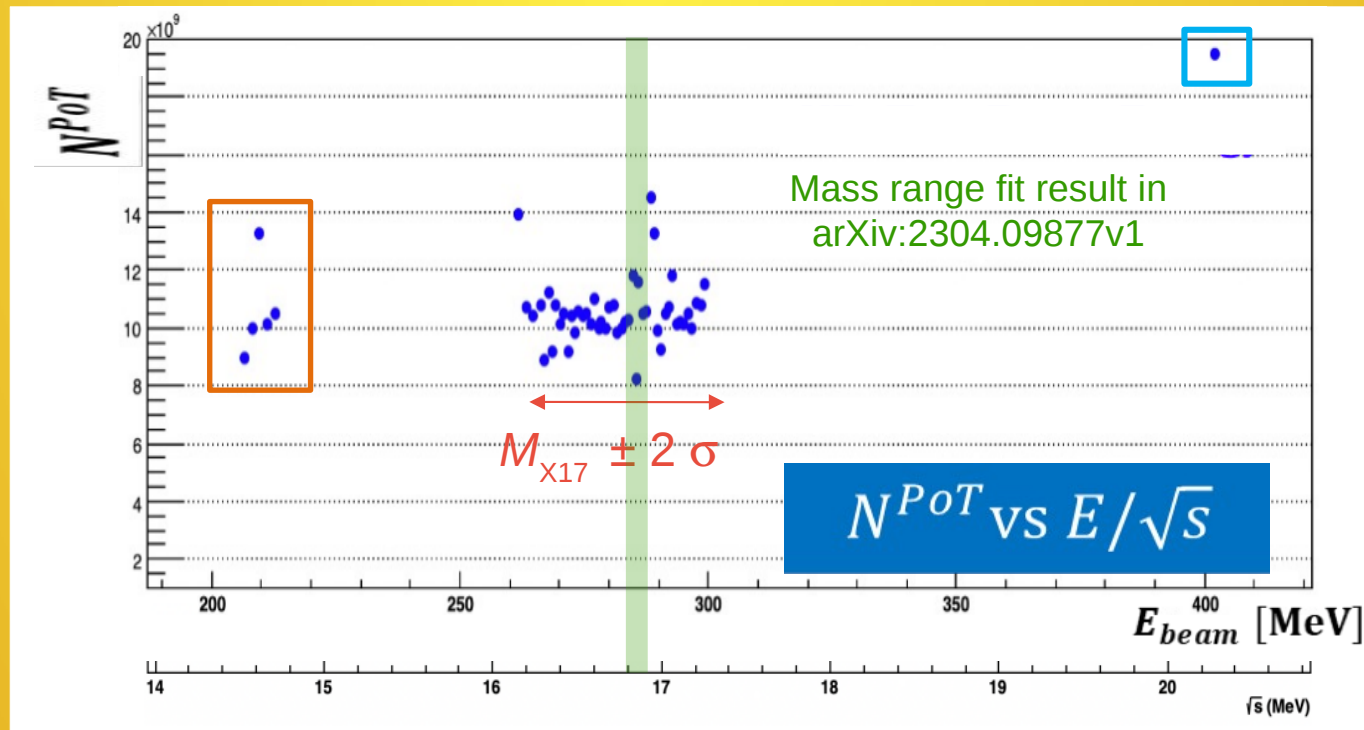
Thanks to the enhanced production cross section can reduce N^{POT}/bunch by factor 10.
 → Much lower pile-up and better energy resolution

The data collected during RunIII

Total amount of data collected $\sim 6 \times 10^{11}$ PoTs (i.e. $\sim 10^{10}$ PoTs per point) :
47 invariant mass points in beam energy range $260 \text{ MeV} < E_{beam} < 300 \text{ MeV}$
 with $\delta E_{beam} \sim 0.75 \text{ MeV}$ ($\pm 2\sigma$ mass around the predicted region by Atomki)
 the precision on the mass measurement will be: $(17.3-16.3)/47 \sim 22 \text{ KeV}$
and 6 points out-of-resonance : **5 points below** + **1 above**
and 3 points without target (beam background studies)

Bunch length $\sim 200 \text{ ns}$, $N_{Bunch}^{PoT} \sim 2500$ at $f \sim 50 \text{ Hz}$

The luminosity and beam energy are measured by combination of **LeadGlass**, **target** and **TimePix3** beam monitors.



The Out-of-Resonance points

Measure the SM cross section **below** and **above** the resonance:

- **5 points** with $N^{PoT} \sim 10^{10}$ events per each point and $205 \text{ MeV} < E_{\text{beam}} < 212 \text{ MeV}$
- **1 point** with $N^{PoT} \sim 2 \times 10^{10}$ events and $E_{\text{beam}} = 402 \text{ MeV}$

Below resonance : X17 production is kinematically **not** allowed

Above : X17 resonance production is suppressed

We will use these datasets to :

- Compare data and MC predictions
- Study the SM backgrounds
- measure Standard Model cross sections
- Tune the search technique
- Establish luminosity measurement precision
- check all systematics

First look at Run III off-resonance data

First selection aimed at $N(e^+e^- + \gamma\gamma)/N_{\text{PoT}}$ studies :

2 clusters in time in ECal ($\Delta t < 5$ ns) + good radial region with reasonable Centre of Gravity

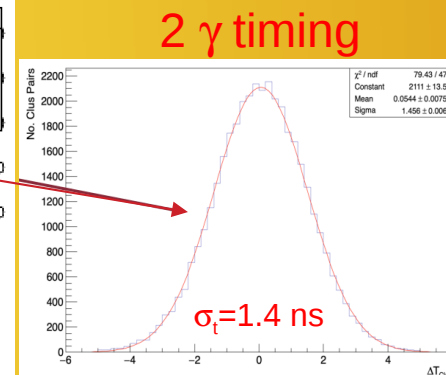
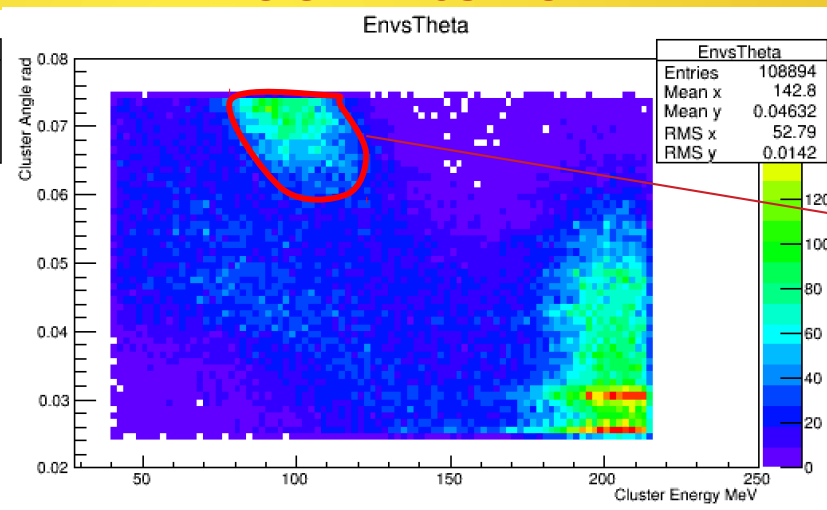
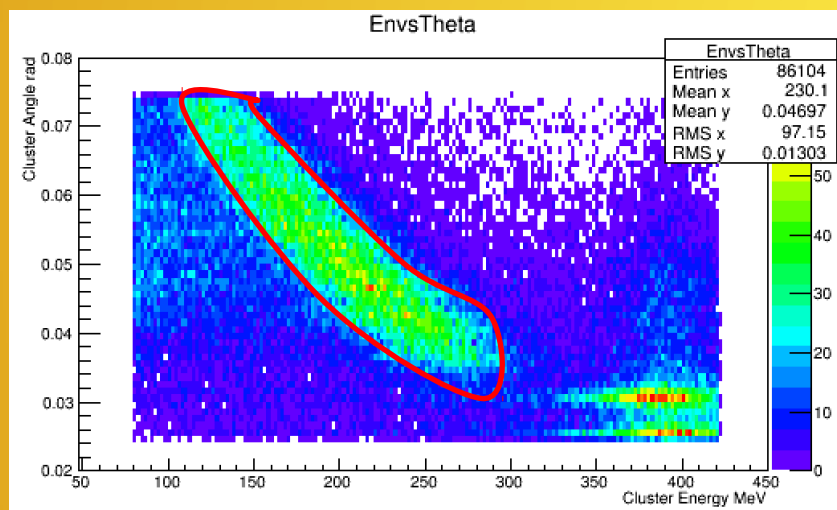
Using kinematic relation between E_γ and $\Theta_\gamma \rightarrow$ very good signal-background separation

compatible with a 2-body final state.

Background on/off resonance data under control

Above : 402 MeV

Below : 205 MeV

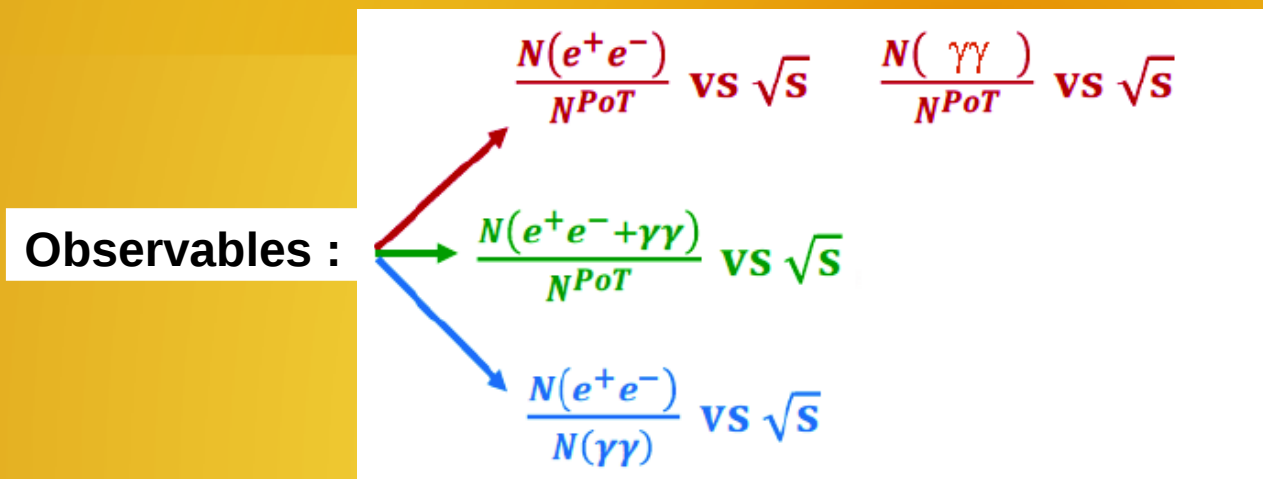


$$g_{V_e} = 2 \times 10^4 \text{ and } \delta E = 0.75 \text{ MeV}$$

Recently, we updated the Toy MC introducing the correct experimental parameters. With respect to preliminary predictions, the BG decreases, while the signal increases

Process	# of Ev.	# of Ev. in Acc.	Acc.
$e^+e^- \rightarrow e^+e^-$ (t-ch.)	$5.4 \cdot 10^7$	$4.3 \cdot 10^4$	0.08%
$e^+e^- \rightarrow e^+e^-$ (s-ch.)	$3.2 \cdot 10^4$	$4.3 \cdot 10^3$	13.6%
$e^+e^- \rightarrow e^+e^-$ (full)	$5.4 \cdot 10^7$	$3.9 \cdot 10^4$	0.07%
$e^+e^- \rightarrow \gamma\gamma$	$2.9 \cdot 10^5$	$8.7 \cdot 10^3$	3%
$e^+e^- \rightarrow X_{17} \rightarrow e^+e^-$	2600	350	13.6%

Observables and possible measurements



Goal: keep at the % level the systematic errors, in particular the luminosity

Several different observables can be used with different outcomes:

- $N(e^+e^- + \gamma\gamma)/N_{PoT}$ = existence of X_{17}
 - High statistical significance
 - No ETag related systematic errors
- $N(e^+e^-)/N(\gamma\gamma)$ = existence of X_{17}
 - ETag efficiency and systematics
 - lower statistical significance due to 2γ cross section
 - Independent from N_{PoT}
- $N(e^+e^-)/N_{PoT}$ = vector nature of X_{17}
 - Systematic errors due to ETag tagging efficiency stability
- $N(\gamma\gamma)/N_{PoT}$ = pseudo-scalar nature of X_{17}
 - Systematic errors due to ETag tagging efficiency stability

Conclusions

In 2019/2020 **PADME** performed 2 physics runs, collecting $> 5 \times 10^{12}$ PoT each

- **Run II** data-set with **primary positron beam** : much better background conditions vs Run I
- Detectors are performing **very well**, a reliable MC simulation, including beamline, is available
- **PADME** delivered its **first physics result**

$\sigma(e+e^- \rightarrow \gamma\gamma) = (1.977 \pm 0.018 \text{ stat} \pm 0.0119 \text{ syst}) \text{ mb}$ - very good agreement with QED NLO

PADME Run III scan for the X_{17} particle successfully made in 2022

- High quality data collected for $16.35 \text{ MeV} < M_{X_{17}} < 17.5 \text{ MeV}$
- Beam Background and Bhabha are under control
- Data quality variables identified allowing to reject beam instabilities
- Strategy to be established to approach the resonance region

Many thanks to the **LNF LINAC team** and all the **accelerator division** for the excellent efficiency and quality of the machine operation during PADME Run III.



STAY TUNED ...

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