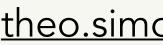
# Constraining cosmological models with the Effective Field Theory of Large-Scale Structures







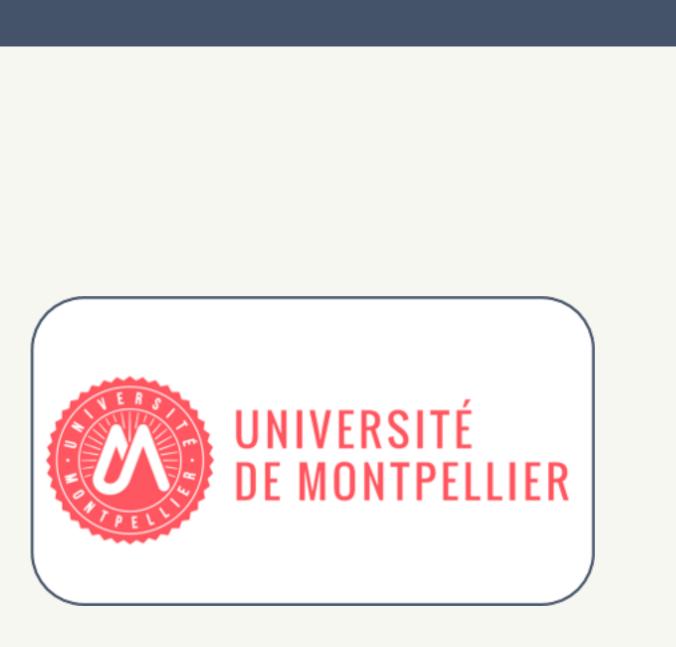


[Cosmological inference from the EFTofLSS: the eBOSS QSO full-shape analysis]

PONT - 04/05/2023

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#### Based on arXiv:2210.14931 **TS**, Pierre Zhang and Vivian Poulin

## The Effective Field Theory of Large-Scale Structures (EFTofLSS) Main motivations

In **linear perturbation theory**, there are two popular ways to use LSS data: 1. Extract information from the full galaxy power spectrum:

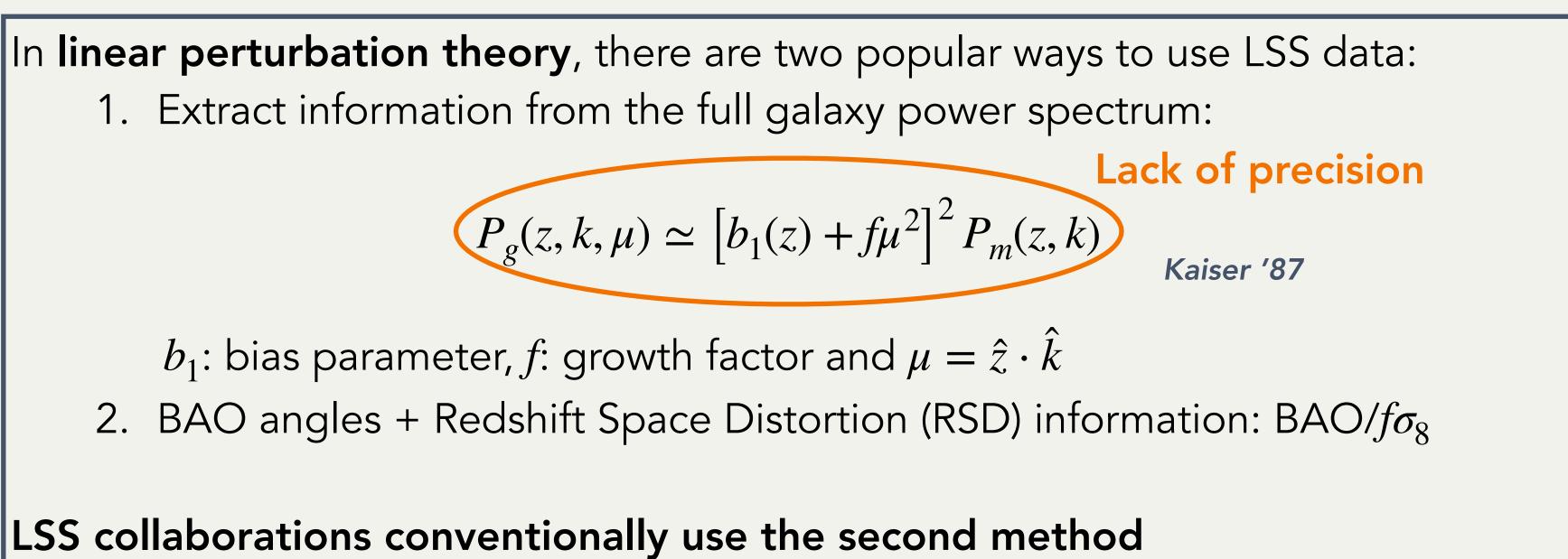
$$P_g(z,k,\mu) \simeq \left[b_1(z) + f\mu^2\right]^2 P_m(z,k)$$
 Kaiser '87

 $b_1$ : bias parameter, f: growth factor and  $\mu = \hat{z} \cdot \hat{k}$ 

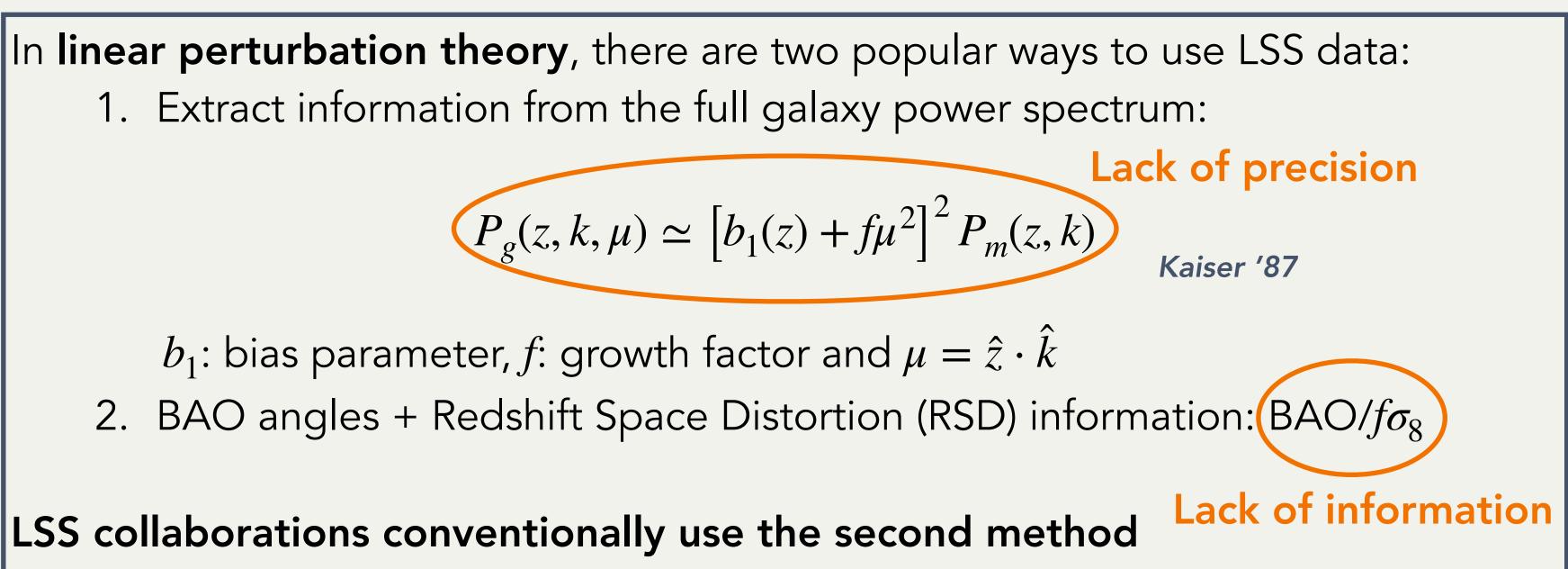
LSS collaborations conventionally use the second method

2. BAO angles + Redshift Space Distortion (RSD) information: BAO/ $f\sigma_8$ 

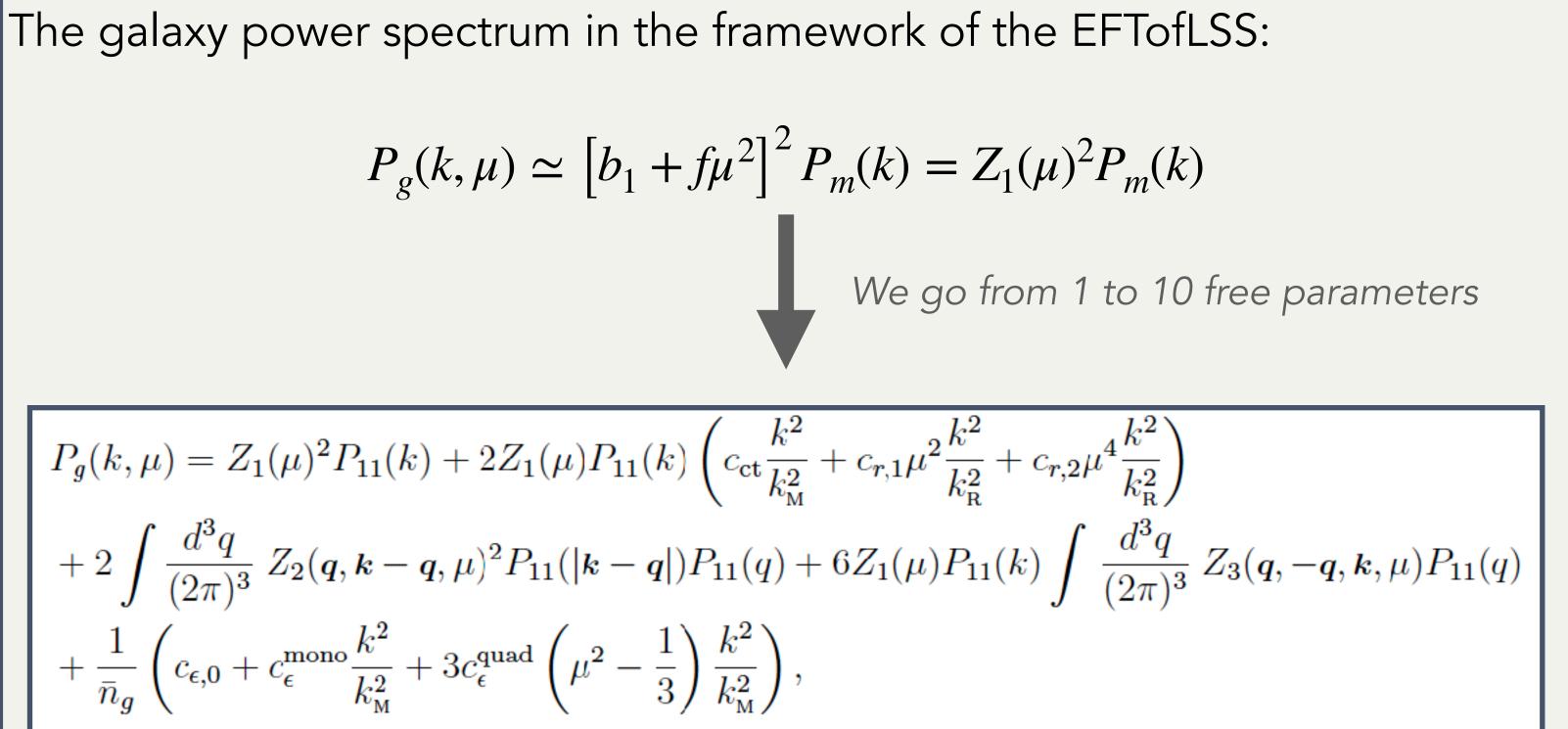
### The Effective Field Theory of Large-Scale Structures (EFTofLSS) Main motivations



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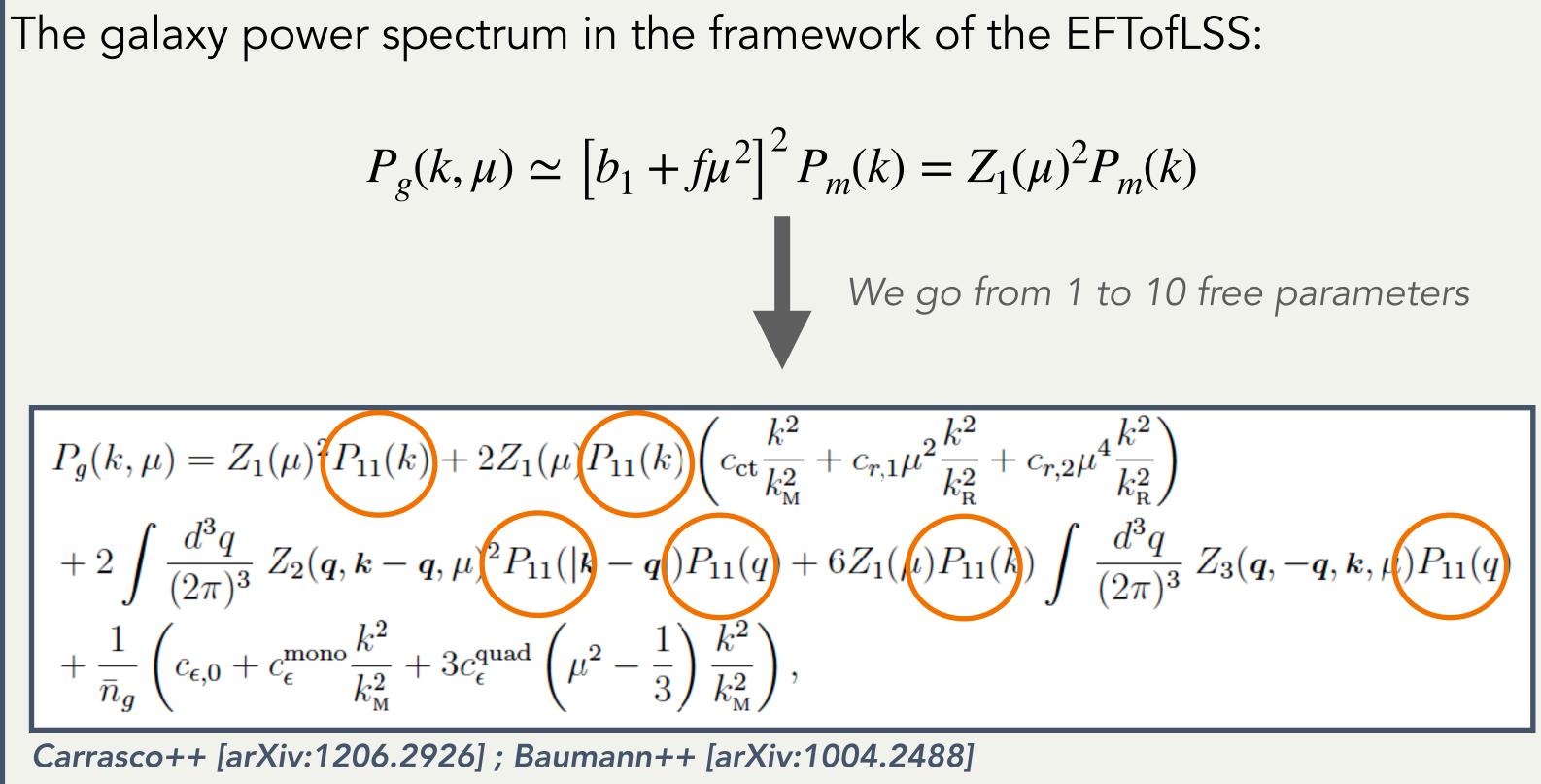
### The effective field theory of large-scale structures (EFTofLSS) Main motivations



Carrasco++ [arXiv:1206.2926] ; Baumann++ [arXiv:1004.2488] Senatore [arXiv:1406.7843] ; Perko++ [arXiv:1610.09321]

#### See Guido D'Amico's talk

## The effective field theory of large-scale structures (EFTofLSS) Main motivations



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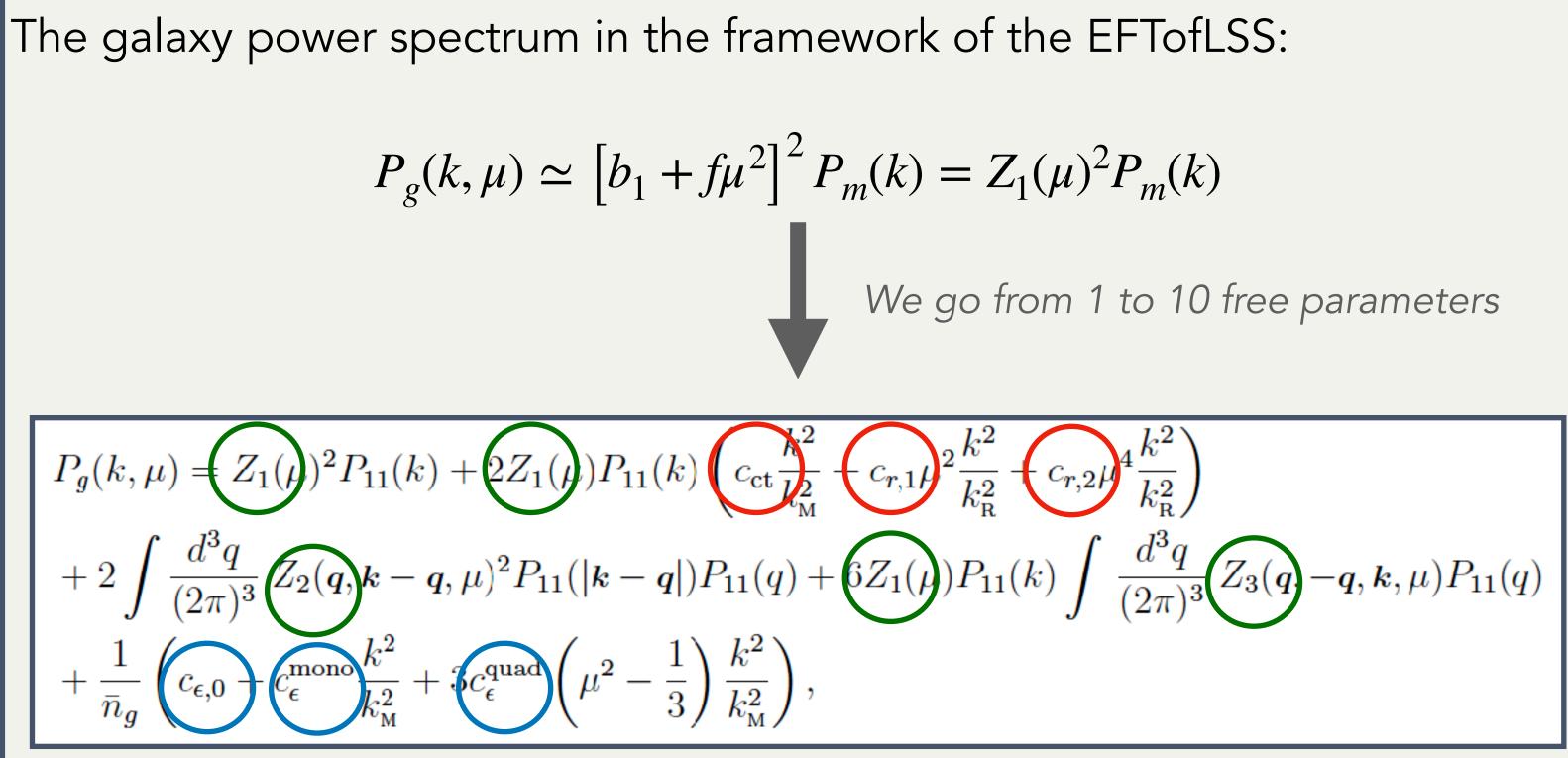
#### See Guido D'Amico's talk

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 $P_g(k,\mu)$  can be determined directly from  $P_{11}(k) = P_m^{lin}(k)$ 



## The effective field theory of large-scale structures (EFTofLSS) Main motivations



Carrasco++ [arXiv:1206.2926] ; Baumann++ [arXiv:1004.2488] Senatore [arXiv:1406.7843] ; Perko++ [arXiv:1610.09321]

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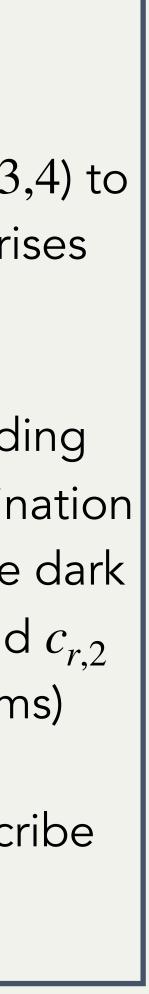
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#### **10 parameters**

**4 parameters**  $b_i$  (i = 1, 2, 3, 4) to describe the galaxy bias which arises from the one-loop contributions

3 parameters corresponding to **counterterms** ( $c_{ct}$  linear combination of a higher derivative bias and the dark matter sound speed, while  $c_{r,1}$  and  $c_{r,2}$ are the redshift-space counterterms)

3 parameters which describe **stochastic** terms

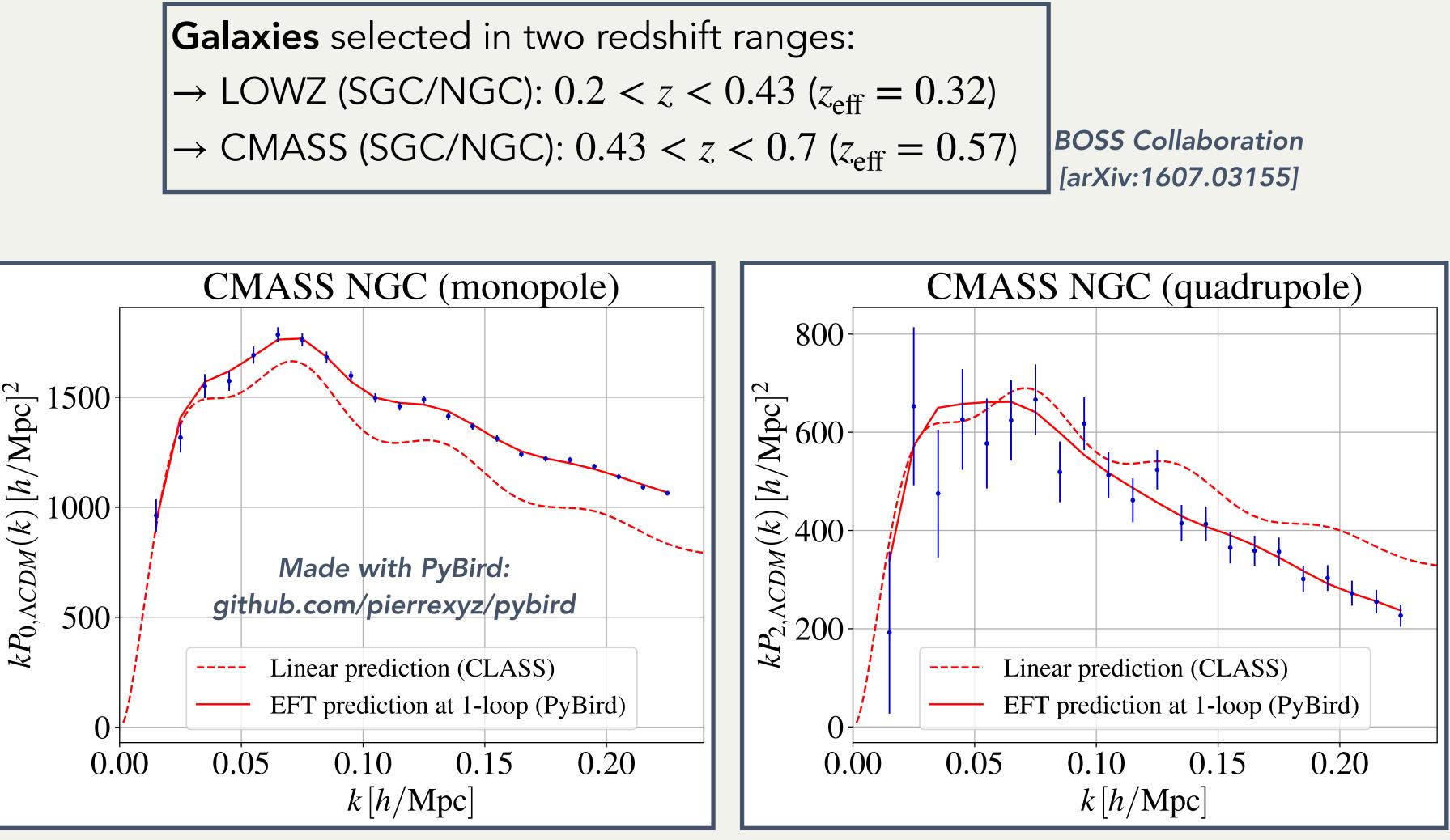


## The effective field theory of large-scale structures (EFTofLSS) Application to BOSS data

**Multipoles** of the galaxy power spectrum, obtained through a **Legendre** polynomials (
$$\mathscr{L}_{\ell}$$
) decomposition:

$$P_g(z,k,\mu) = \sum_{\substack{\ell \text{ even}}} \mathscr{L}_\ell(\mu) P_\ell(z,k)$$

 $\rightarrow$  the two main contributions to  $P_g(z, k, \mu)$  are the **monopole**  $(\ell = 0)$  and the **quadrupole**  $(\ell = 2)$ 



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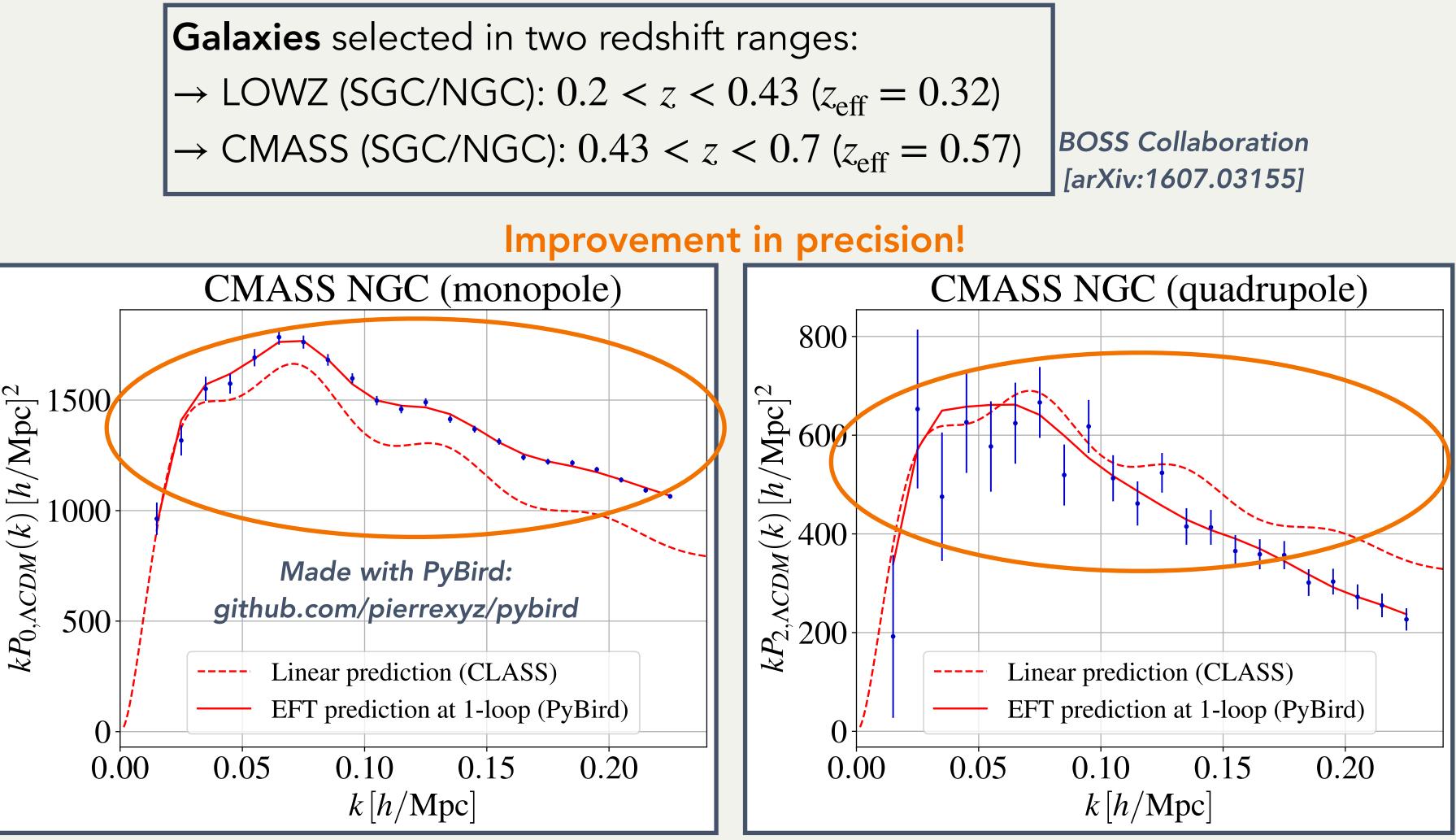
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## The effective field theory of large-scale structures (EFTofLSS) Application to BOSS data

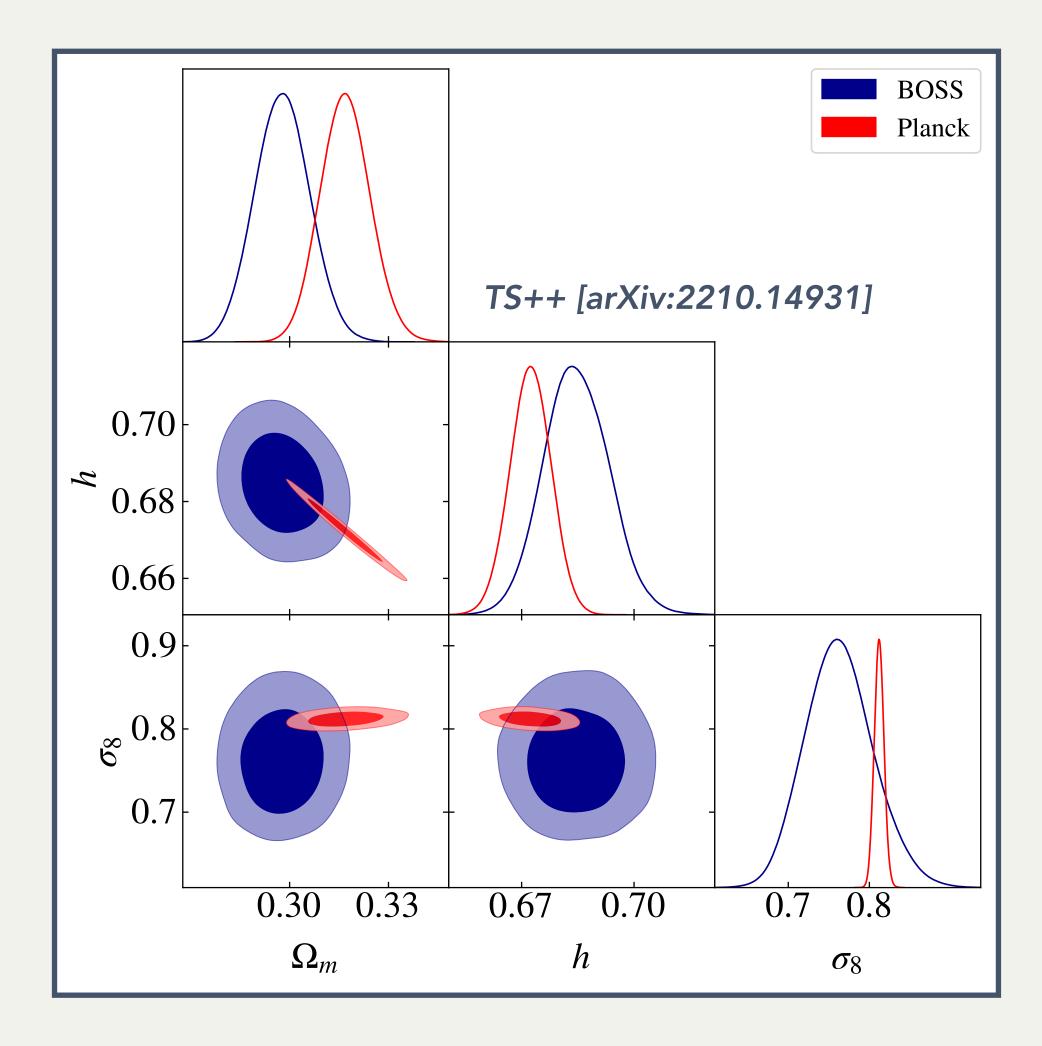
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## The effective field theory of large-scale structures (EFTofLSS) Application to BOSS data



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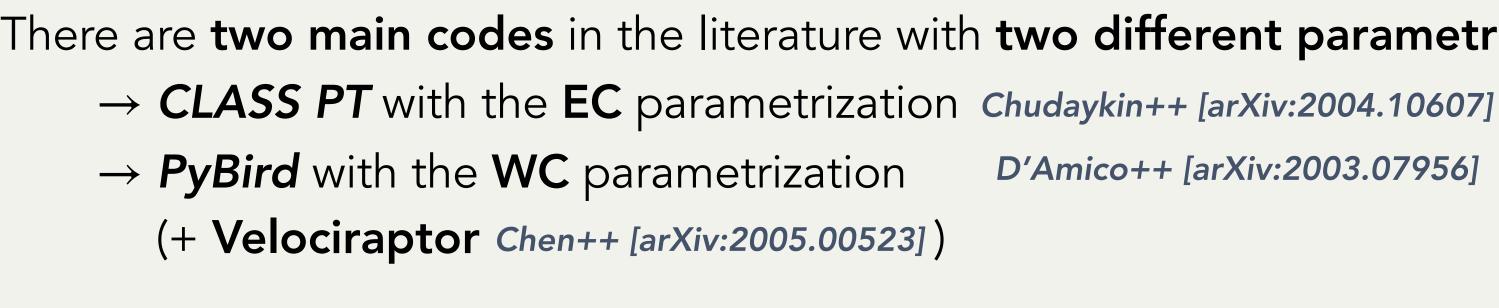
The EFTofLSS analysis of BOSS data allows to determine  $\Omega_m$  and h with a **precision of only** 10% and 60% lower than Planck

See also D'Amico++ [arXiv:1909.05271] ; Philcox++ [arXiv:2002.04035]

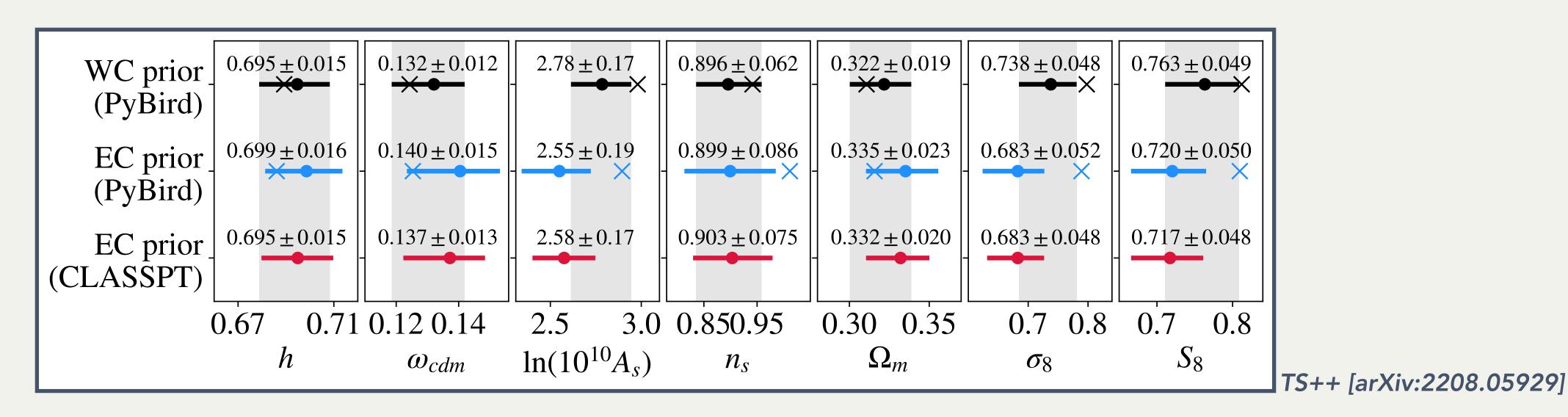




## On the consistency of EFTofLSS The EFT prior issue



#### $\rightarrow$ these two codes use **two different sets of priors** on EFT parameters



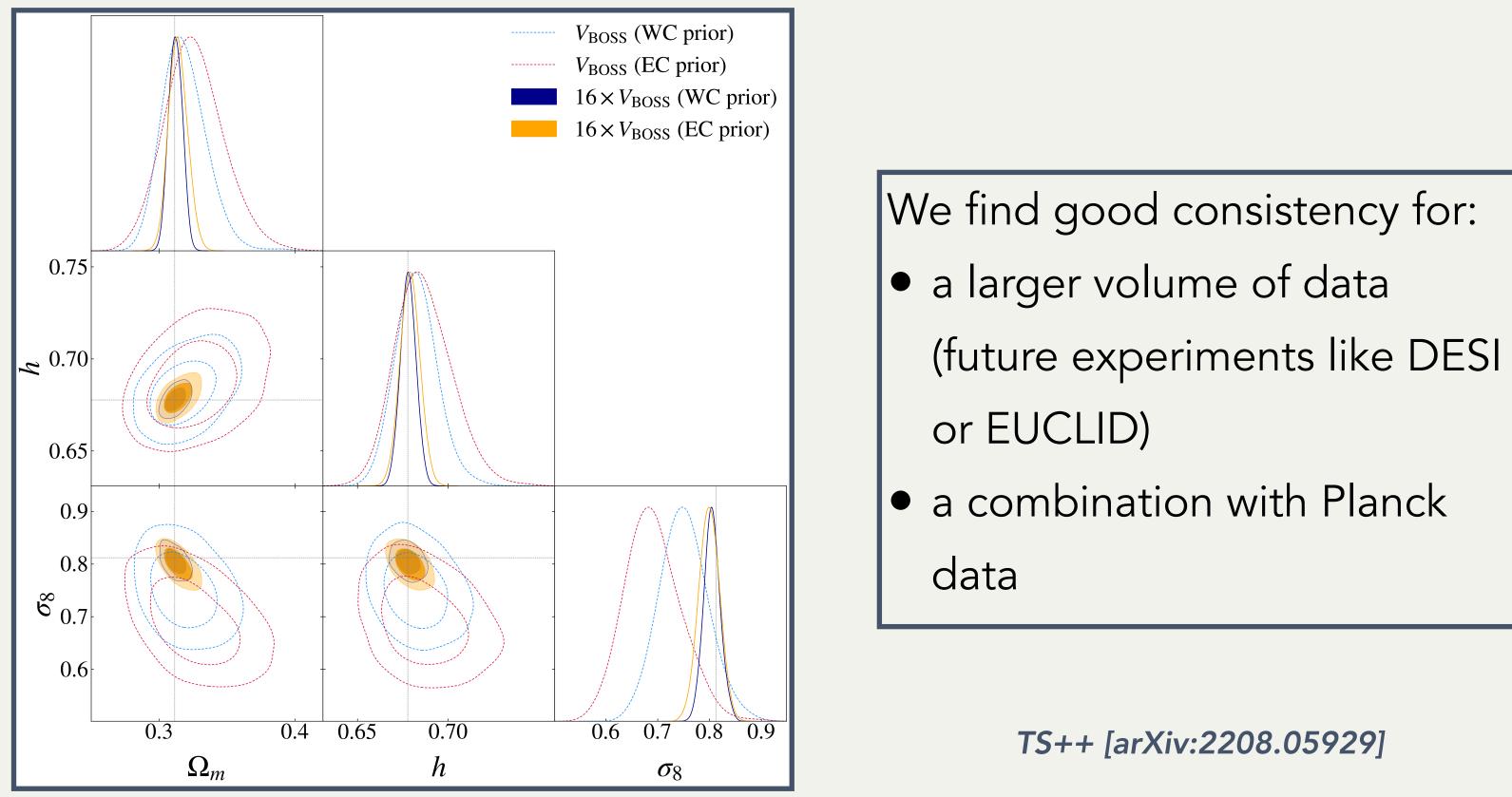
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There are two main codes in the literature with two different parametrizations:

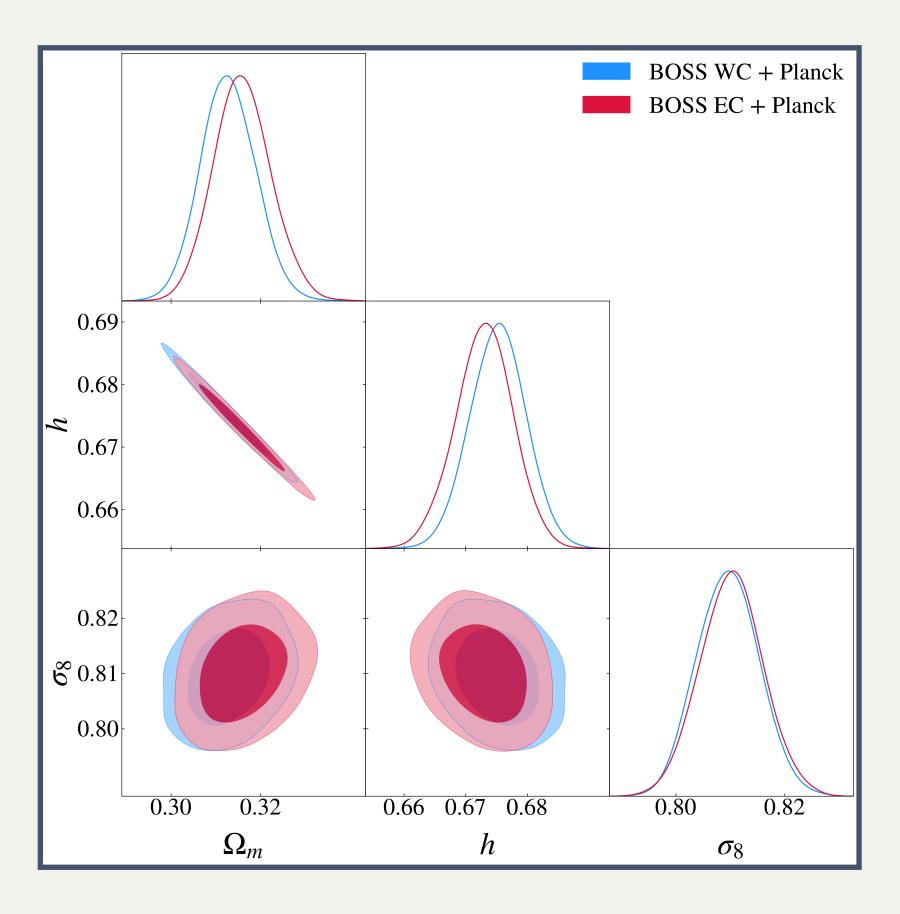
- D'Amico++ [arXiv:2003.07956]



#### On the consistency of EFTofLSS How to overcome this problem?

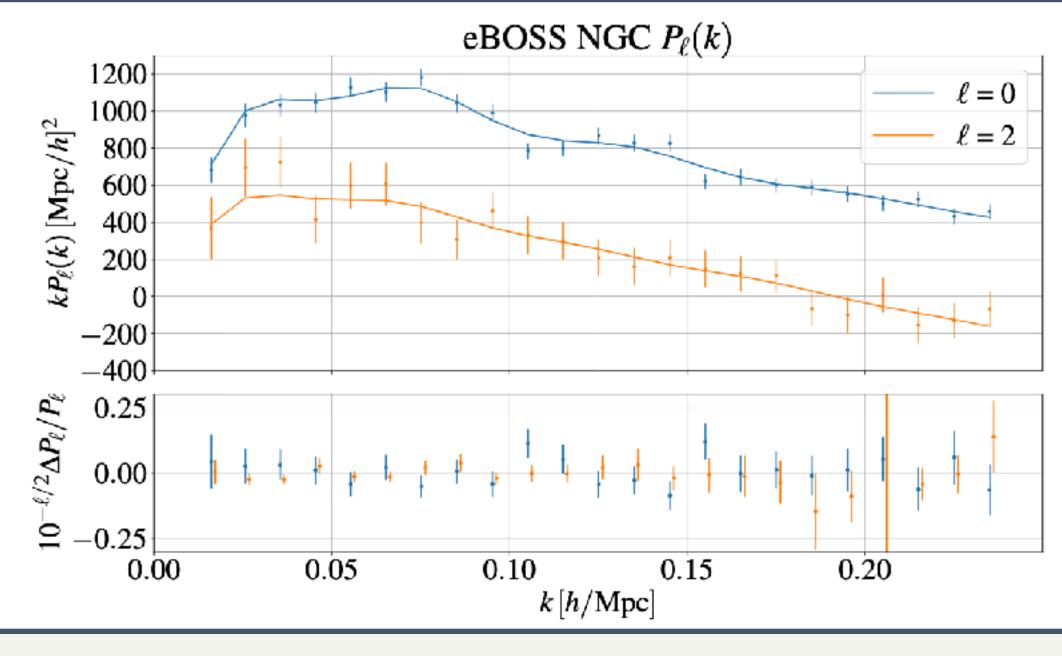


Work in progress: profile likelihood of the EFTofLSS analysis of BOSS data



## EFTofLSS applied to eBOSS QSO data

- $z_{\rm eff} = 1.5$
- 2 skycuts: NGC and SGC

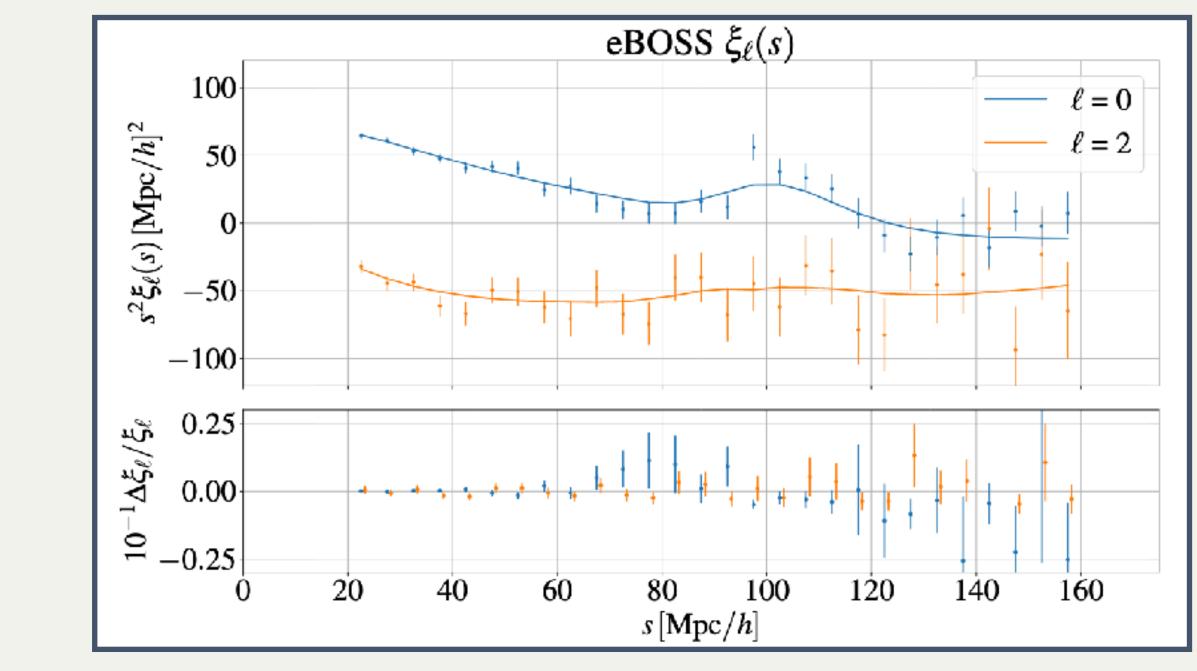


TS++ [arXiv:2210.14931]

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• 343 708 quasars selected in the redshift range 0.8 < z < 2.2

eBOSS Collaboration [arXiv:2007.08991]



#### Determination of the cut-off scale $k_{max}$ of the one-loop prediction The next-to-next-to-leading order (NNLO) terms

At one-loop order, the galaxy power spectrum reads:

$$\begin{split} P_g(k,\mu) &= Z_1(\mu)^2 P_{11}(k) + 2Z_1(\mu) P_{11}(k) \left( c_{\rm ct} \frac{k^2}{k_{\rm M}^2} + c_{r,1} \mu^2 \frac{k^2}{k_{\rm M}^2} + c_{r,2} \mu^4 \frac{k^2}{k_{\rm M}^2} \right) \\ &+ 2 \int \frac{d^3 q}{(2\pi)^3} \, Z_2(q, \mathbf{k} - q, \mu)^2 P_{11}(|\mathbf{k} - q|) P_{11}(q) + 6Z_1(\mu) P_{11}(k) \int \frac{d^3 q}{(2\pi)^3} \, Z_3(q, -q, \mathbf{k}, \mu) P_{11}(q) \\ &+ \frac{1}{\bar{n}_g} \left( c_{\epsilon,0} + c_{\epsilon,1} \frac{k^2}{k_{\rm M}^2} + c_{\epsilon,2} f \mu^2 \frac{k^2}{k_{\rm M}^2} \right), \end{split}$$

One can add the **NNLO terms** (*i.e.*, the dominant two-loop terms):

$$P_{\text{NNLO}}(k,\mu) = \frac{1}{4} b_1 \left( c_{r,4} b_1 + c_{r,6} \mu^2 \right) \mu^4 \frac{k^4}{k_{\text{R}}^4} P_{11}(k)$$

If the contribution of  $P_{\text{NNLO}}(k,\mu)$  becomes **too large,** the one-loop prediction is **not accurate enough**  $\rightarrow$  this determines the **cut-off scale**  $k_{\text{max}}$  of the prediction

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Zhang++ [arXiv:2110.07539]

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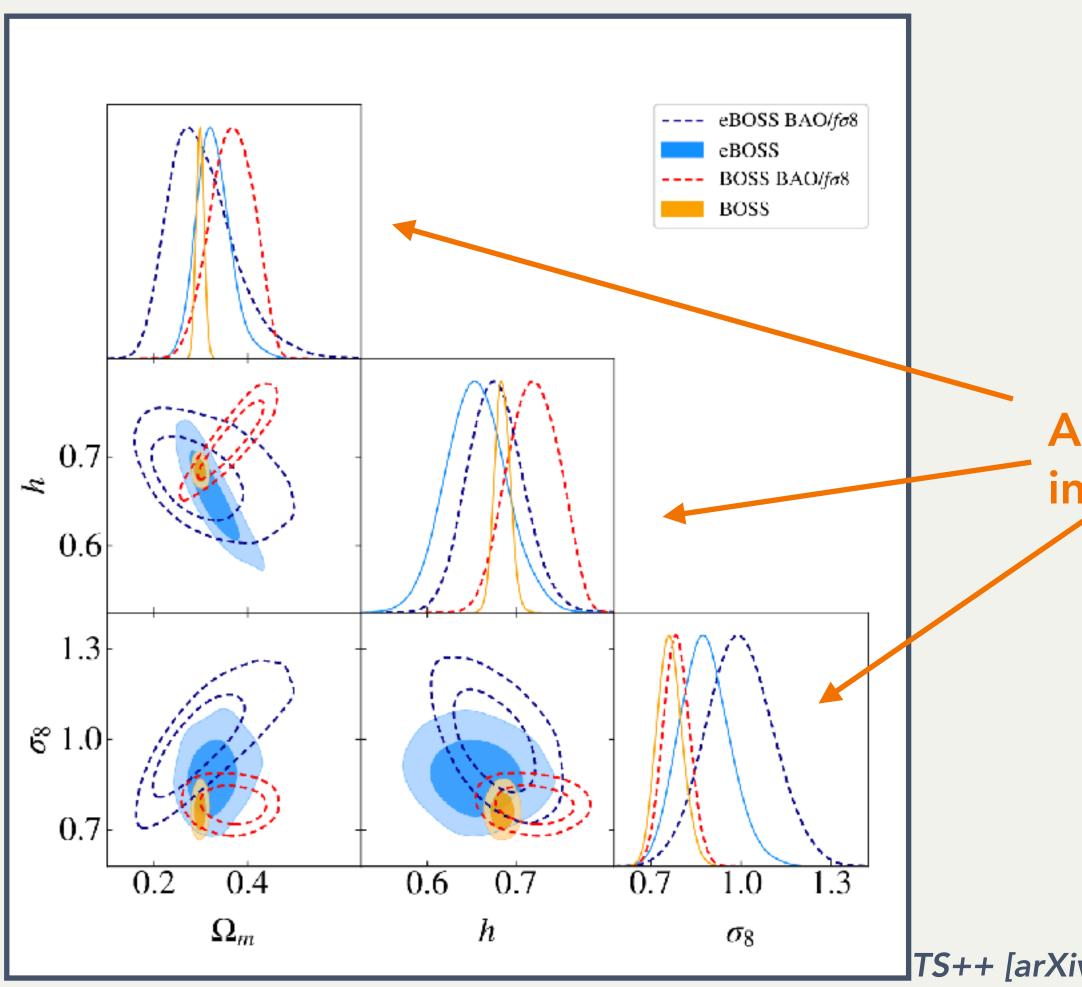
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Zhang++ [arXiv:2110.07539]

2 new EFT terms



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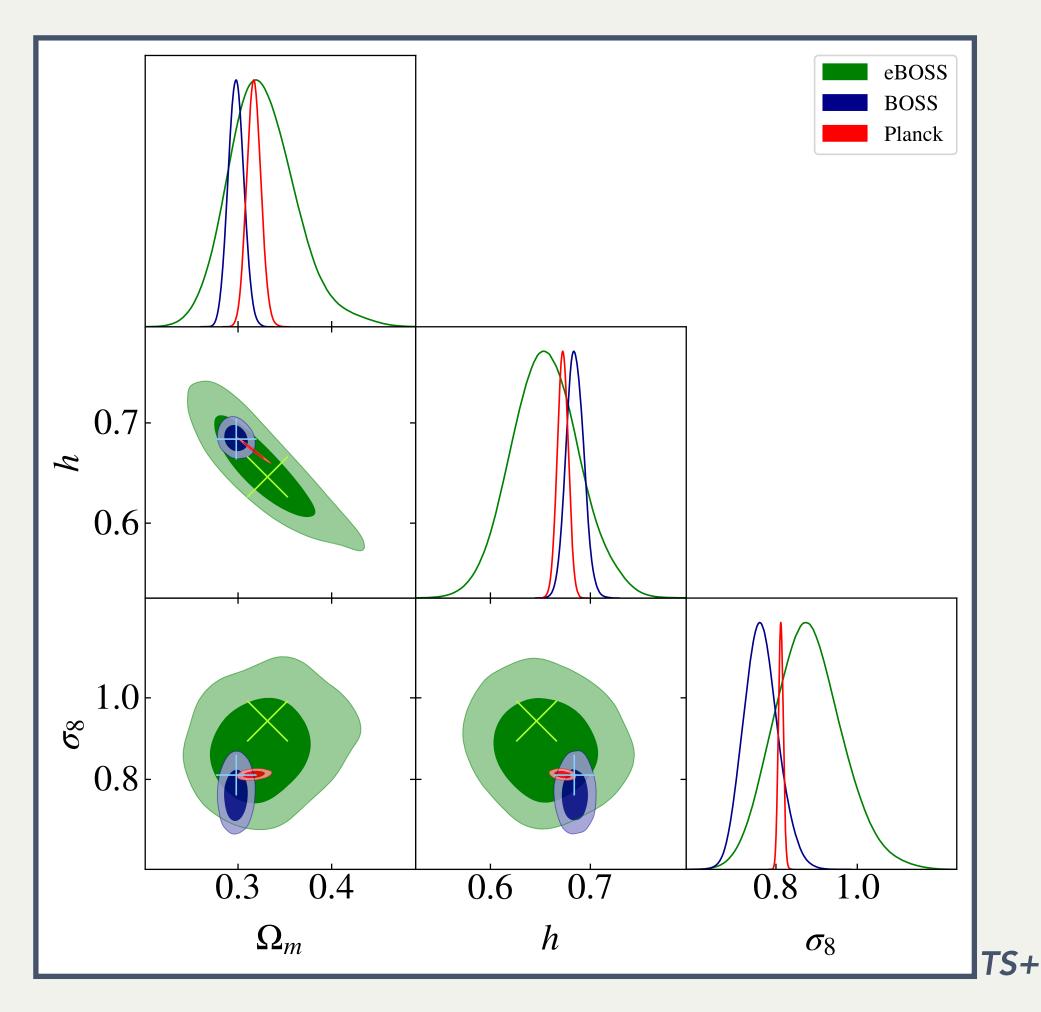
Additional information

- For **eBOSS**, the error bars of  $\Omega_m$  and  $\sigma_8$  are reduced by a factor  $\sim 2.0$  and  $\sim 1.3$
- For **BOSS**, the error bars of  $\Omega_m$  and h are reduced by a factor ~ 5.4 and ~ 3.2

TS++ [arXiv:2210.14931]



## LSS data vs Planck

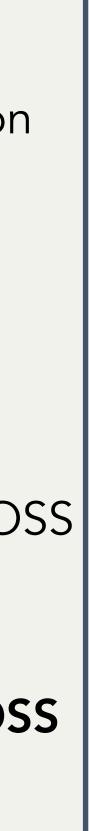


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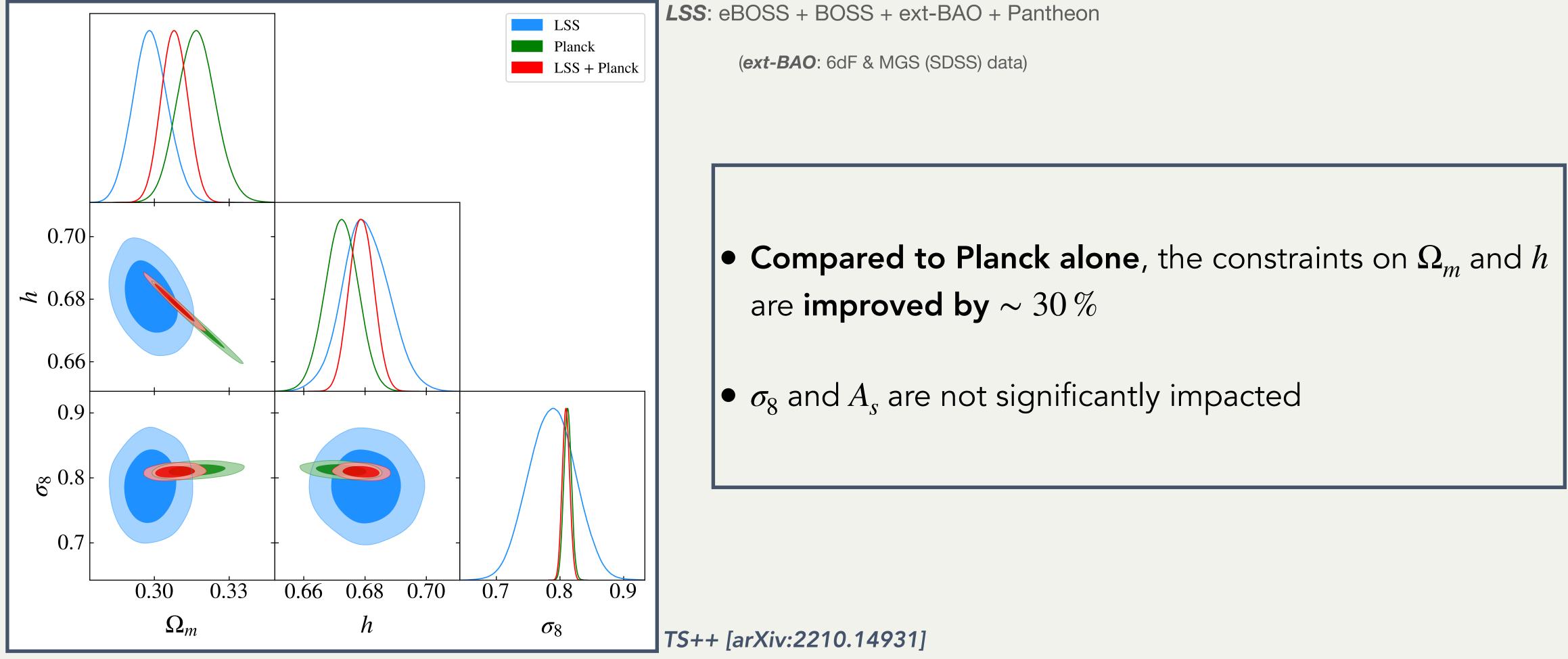
- eBOSS, BOSS and Planck are consistent at  $\lesssim 1.8\sigma$  on all cosmological parameters
- h is ~  $1\sigma$  lower for eBOSS than for BOSS, while  $\sigma_8$  is ~  $1.5\sigma$  higher
- The h and  $\sigma_8$  Planck values are in-between those of BOSS and eBOSS

 $\rightarrow$  there is no tension between Planck and BOSS/eBOSS

TS++ [arXiv:2210.14931]

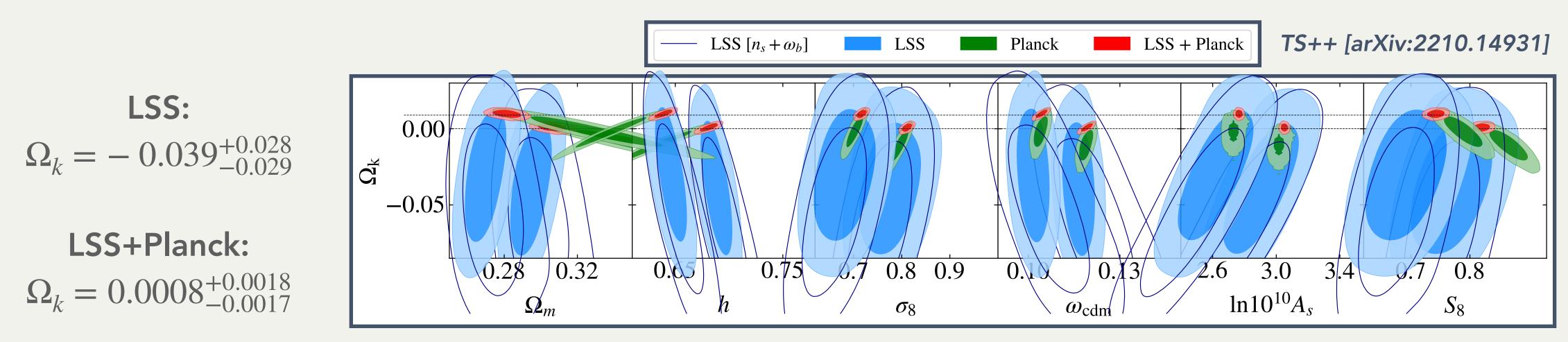


## LSS data combined with Planck



## **Extensions to** $\Lambda$ **CDM:** curvature density fraction $\Omega_k$

- With LSS data only, we find  $\Omega_k$  compatible with zero curvature at  $1.3\sigma$
- conventional BAO/ $f\sigma_8$  analysis
- favored) negative values of  $\Omega_k$



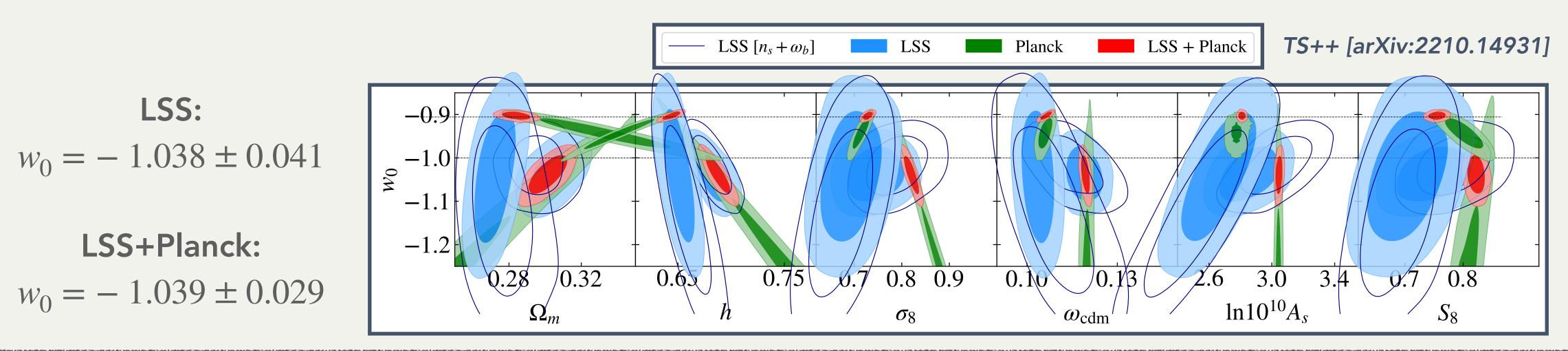
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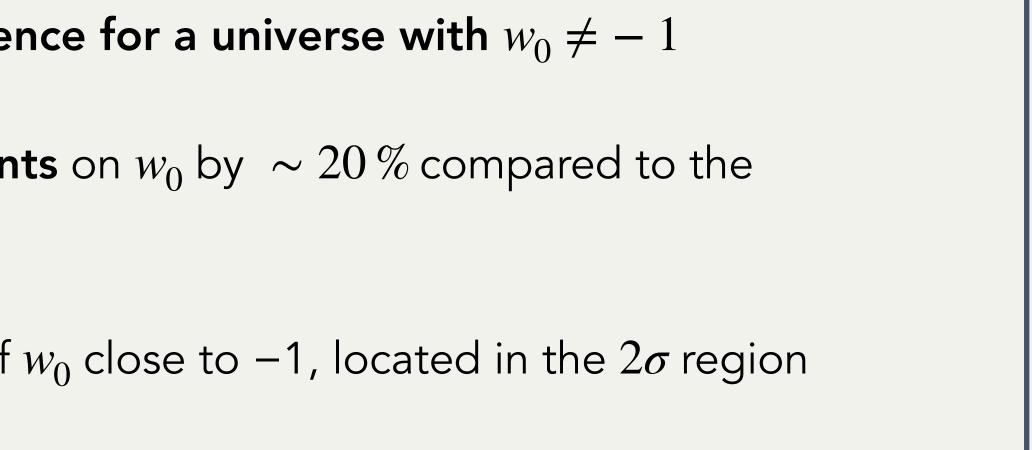
• The EFT analysis significantly improves the constraints on  $\Omega_k$  by  $\sim 50\%$  compared to the

• The combination of LSS and Planck leads to a strong constraint and excludes the (slightly

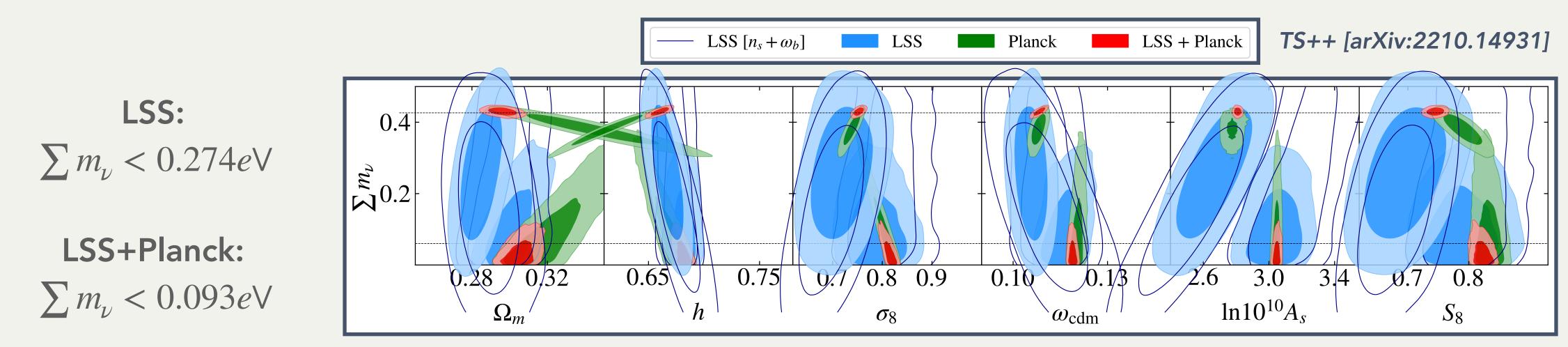
## **Extensions to** $\Lambda$ **CDM:** dark energy equation of state $w_0$

- With the LSS data only, we find **no evidence for a universe with**  $w_0 \neq -1$
- The EFT analysis **improves the constraints** on  $w_0$  by ~ 20 % compared to the conventional BAO/ $f\sigma_8$  analysis
- The addition of LSS data select values of  $w_0$  close to -1, located in the  $2\sigma$  region reconstructed from Planck data





- The LSS constraint derived in this work is only  $\sim 10\%$  weaker than the Planck constraint  $\sum m_{\nu} = 0.241 eV$
- The EFT analysis **significantly improves the** conventional BAO/ $f\sigma_8$  analysis (  $\sum m_{
  u} = 4.84$



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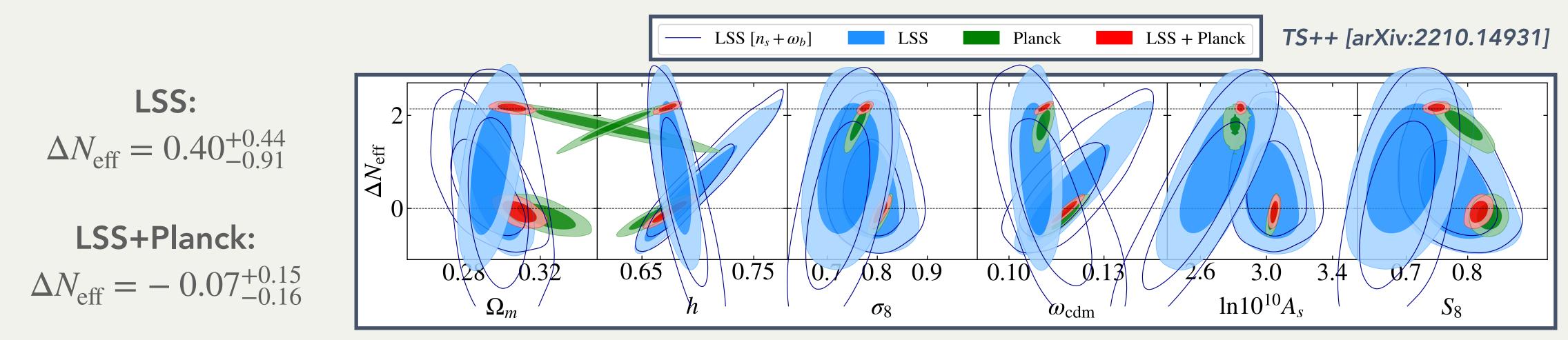
constraints on 
$$\sum m_{\nu}$$
 (by a factor of  $\sim 18$ ) over the eV)

Palangue-Delabrouille++ [arXiv:1911.09073]

• This analysis disfavors the inverse hierarchy at  $\sim 2.2\sigma$  & is competitive to the Lyman- $\alpha$  constraints

## **Extensions to** $\Lambda$ **CDM:** effective number of relativistic species $N_{\rm eff}$

- The value of  $\Delta N_{\rm eff}$  is compatible with the standard model
- Unlike EFTofLSS, the conventional BAO/ $f\sigma_8$  analysis is unable to constrain this parameter
- The addition of the LSS data **improves** the results of Planck alone by  $\sim 25~\%$





- The EFTofLSS is a novel method that provides an accurate description of LSS data at a controlled precision
- Constraints from LSS data are **competitive with CMB data**
- EFTofLSS allows to highlight that there is no tension between current BOSS/eBOSS data and Planck data (but not in tension with weak lensing neither)
- Data are consistent with  $\Lambda$ CDM at  $\lesssim 1.3\sigma \rightarrow$  Strong constraints on canonical extensions to  $\Lambda$ CDM e.g. LSS+Planck:  $\sum m_{\nu} < 0.093 eV$
- EFTofLSS provides interesting constraints on non-trivial extensions of the  $\Lambda$ CDM model:  $\rightarrow$  see [TS et al. '22, arXiv:2203.07440] for **Decaying Cold Dark Matter**  $\rightarrow$  see [TS et al. '22, arXiv:2208.05930] for Early Dark Energy

### Conclusion











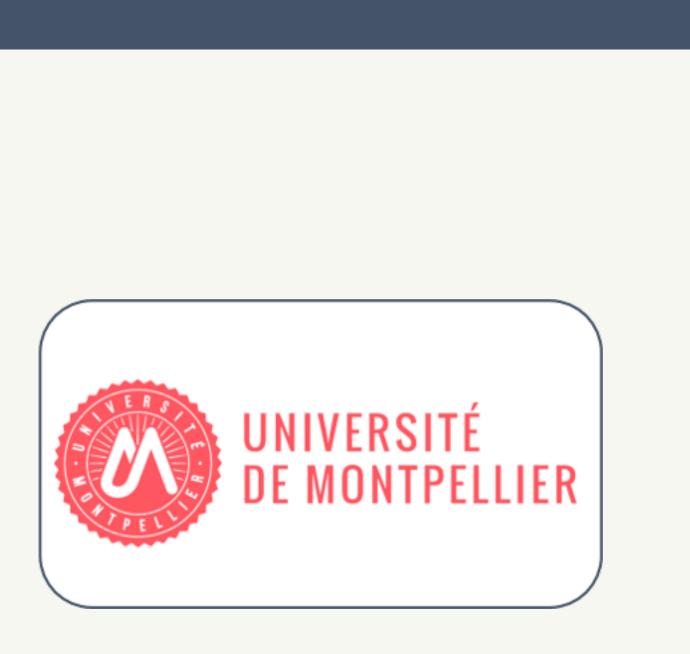
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# Thanks for your attention

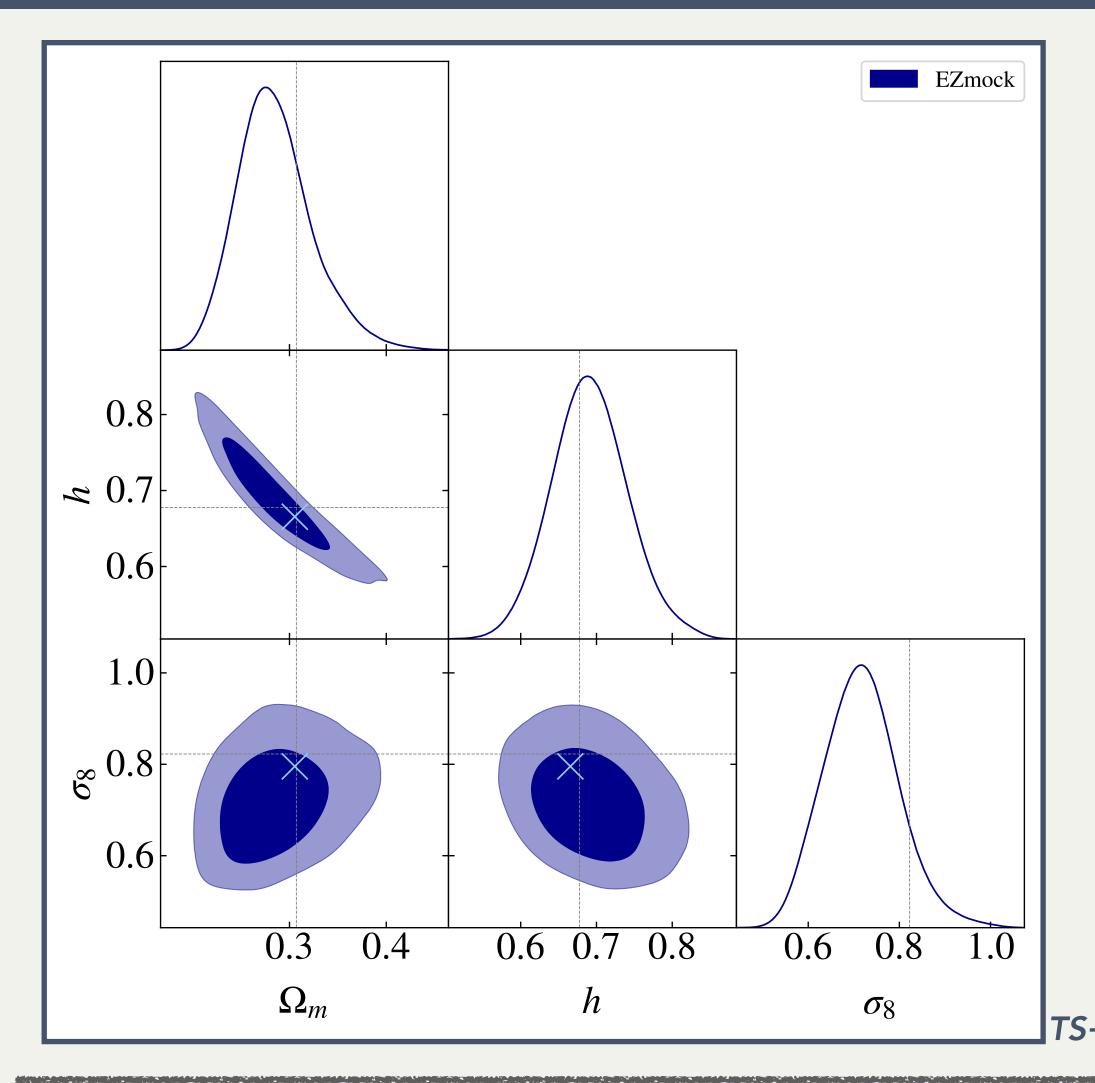
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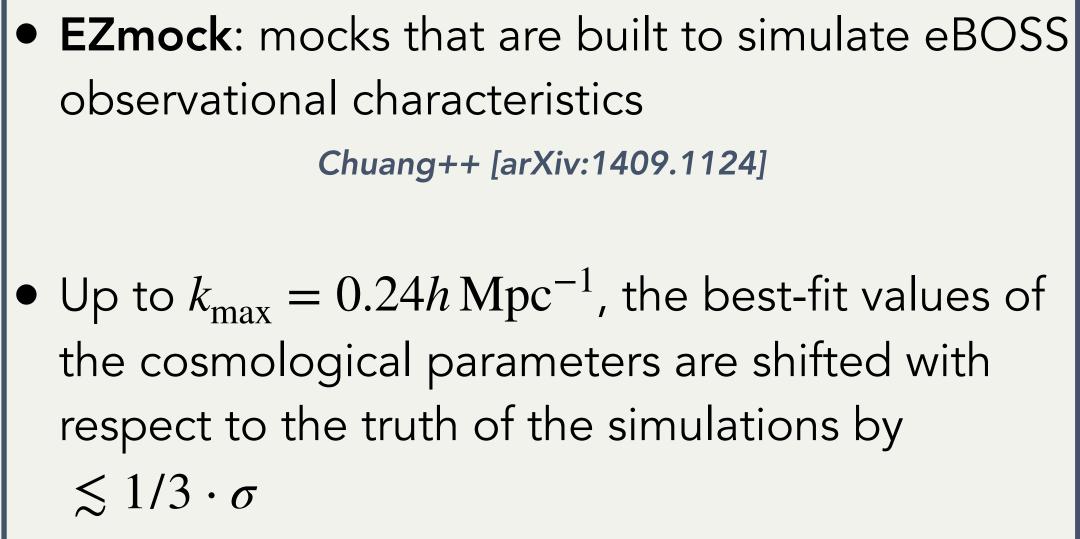


#### Based on arXiv:2210.14931 **TS**, Pierre Zhang and Vivian Poulin

### Determination of the cut-off scale $k_{max}$ of the one-loop prediction The EZmock



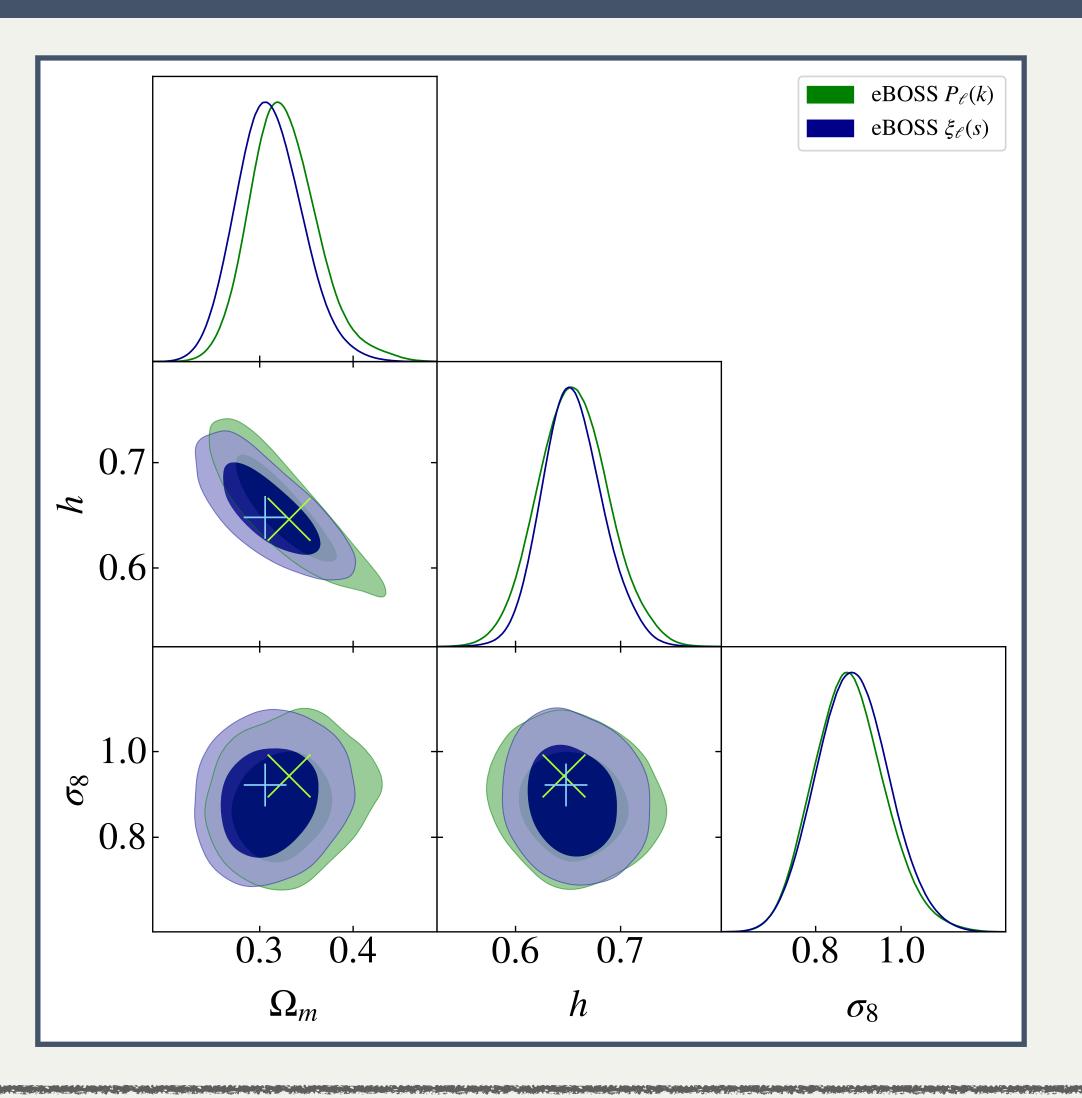
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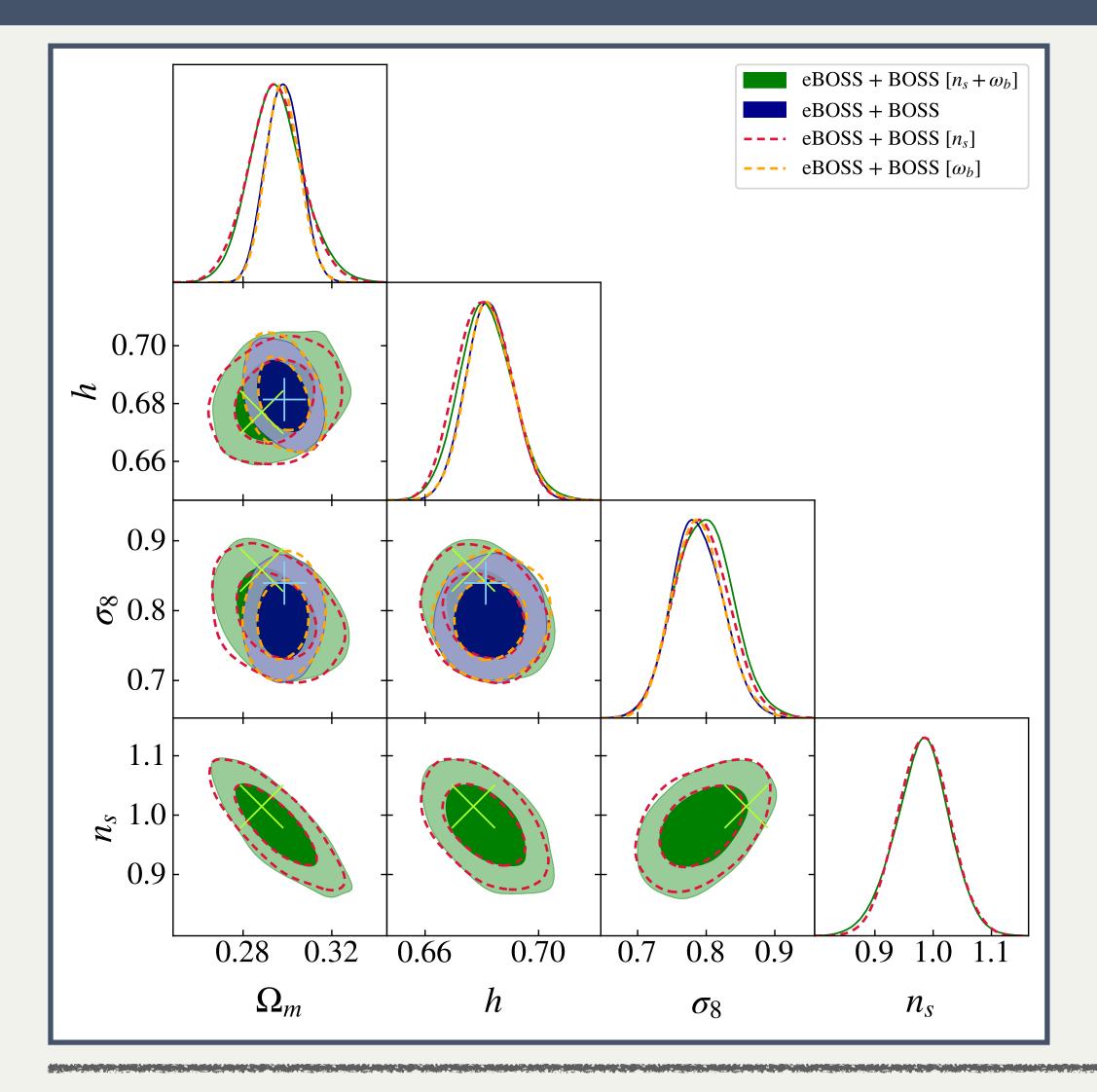
TS++ [arXiv:2210.14931]



## eboss $P_{\ell}(k)$ vs eboss $\xi_{\ell}(k)$



# Variation of $n_s$ and $\omega_b$



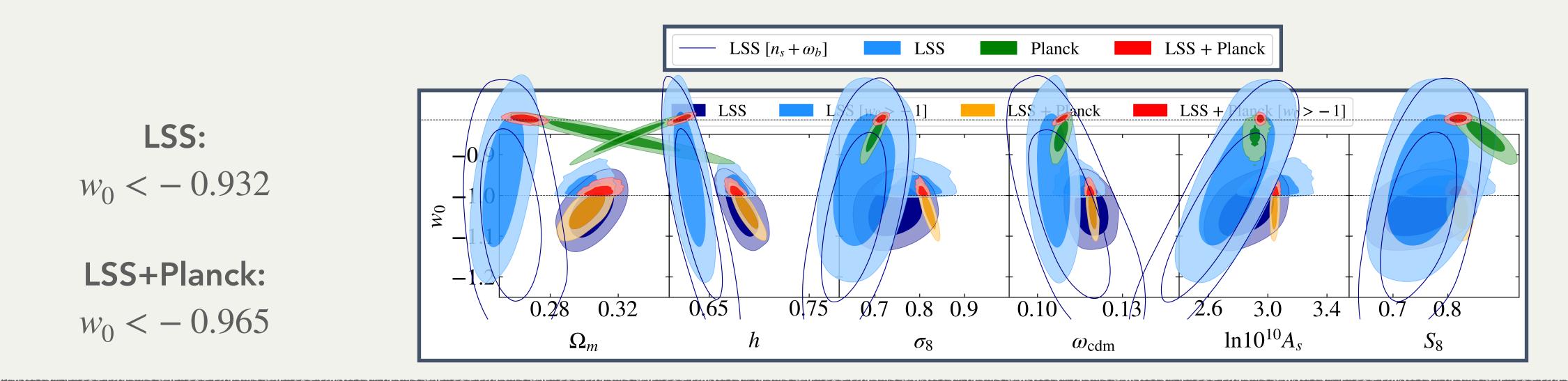
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- We impose a uninformative large flat prior on  $n_{s'}$ while we impose a BBN Gaussian prior on  $\omega_b$
- The variation of  $\omega_b$  within the BBN prior has a negligible impact on the cosmological results: we have a relative shift of  $\lesssim 0.04\sigma$
- The variation of  $n_s$  within a uninformative large flat prior leads to a relative shift  $\leq 0.4\sigma$



# Dark energy equation of state $w_0 \ge -1$

- negligible way, while it remains globally stable for the LSS + Planck
- For these analyses,  $\Delta \chi^2 = 0$  with respect to  $\Lambda$ CDM, since we obtain best-fit values of  $w_0 = -1$



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• One can see that this new prior shifts the 2D posteriors inferred from the LSS data in a non-