

Measuring SMBH properties from binary inspirals in LISA

ADRIEN KUNTZ

Conférence PONT - Avignon

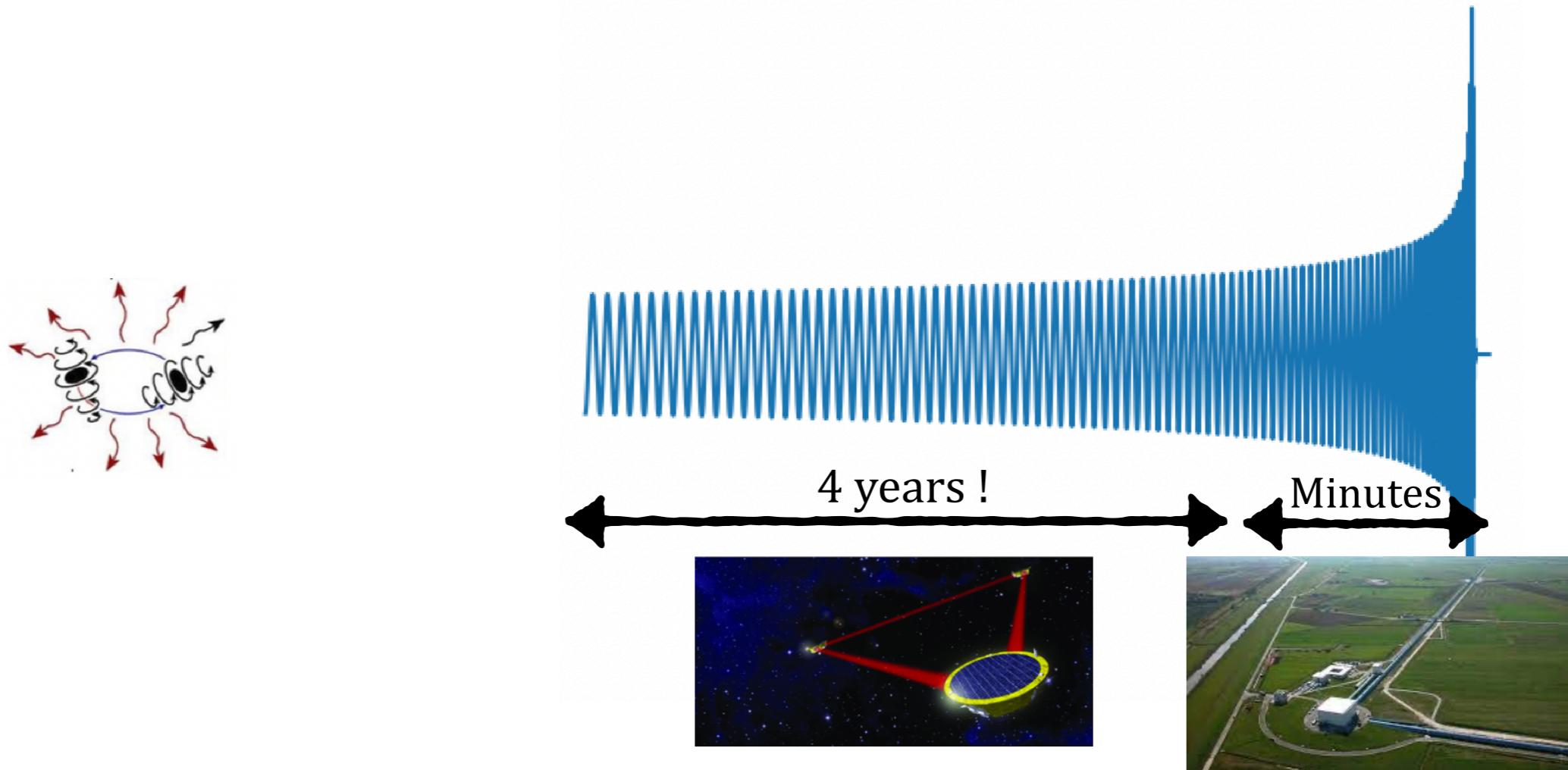


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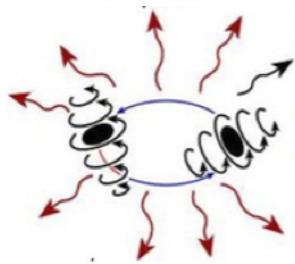


INTRODUCTION

BINARY INSPIRALS IN LISA

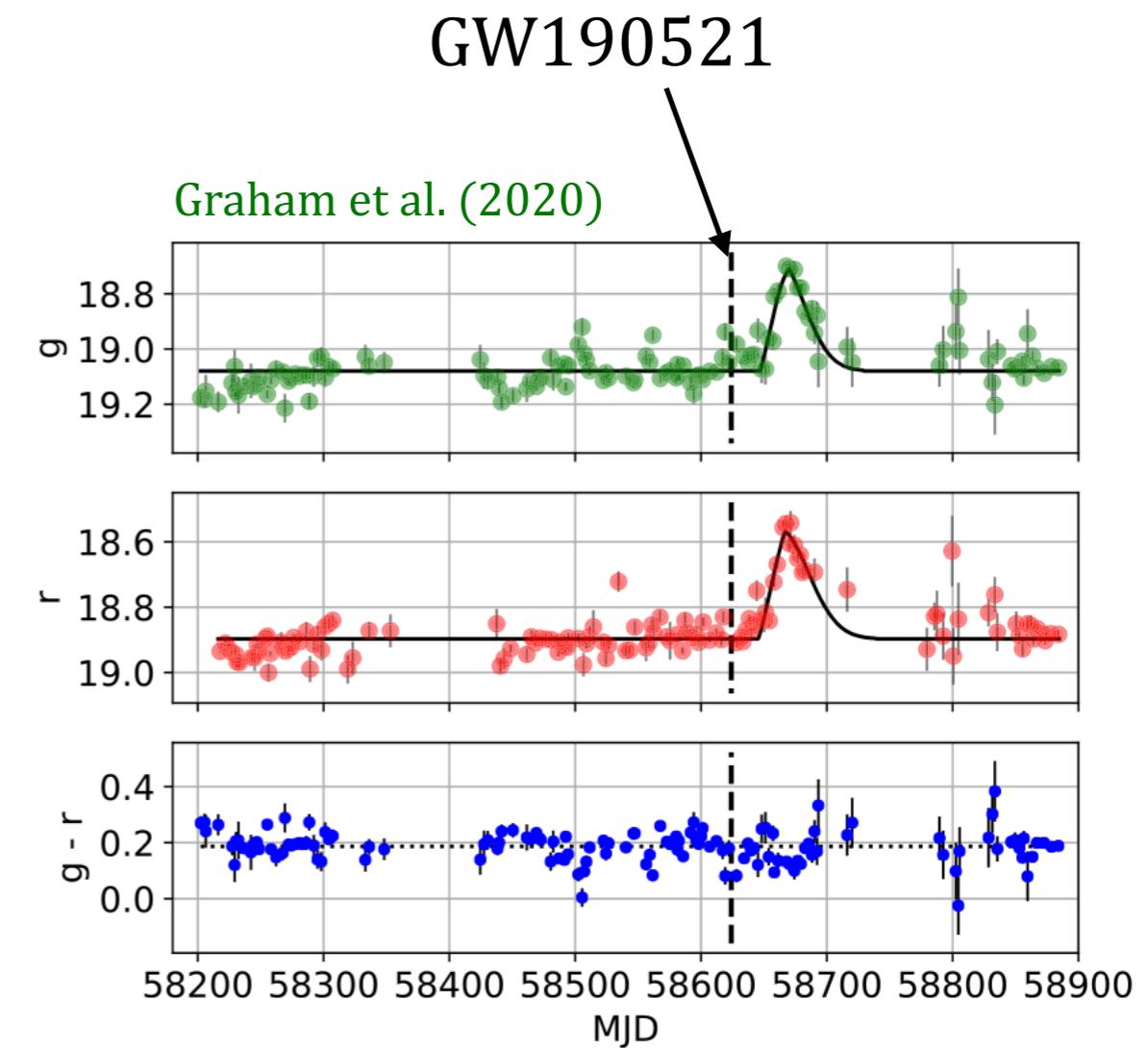
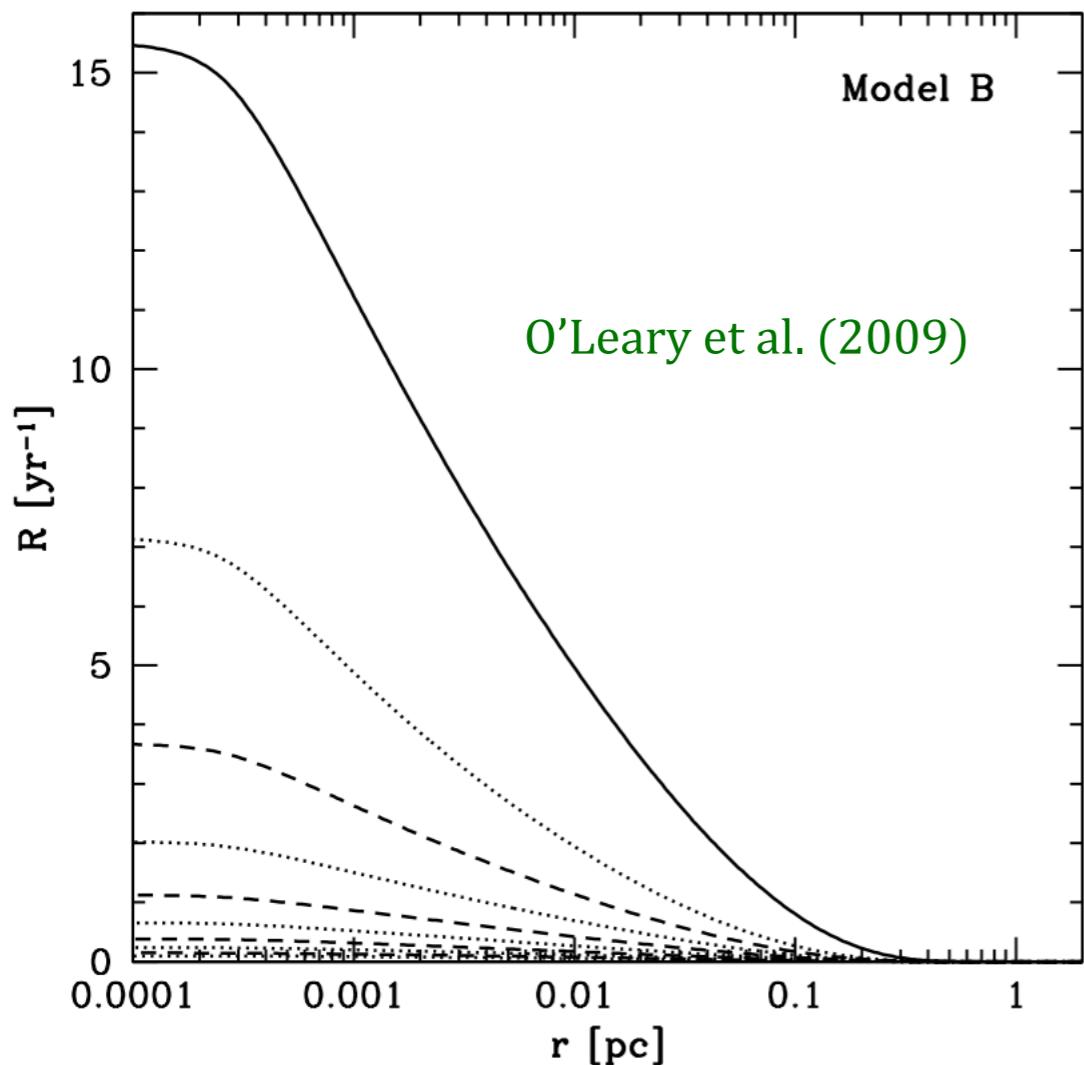


Are vacuum templates precise enough for LISA?



INTRODUCTION

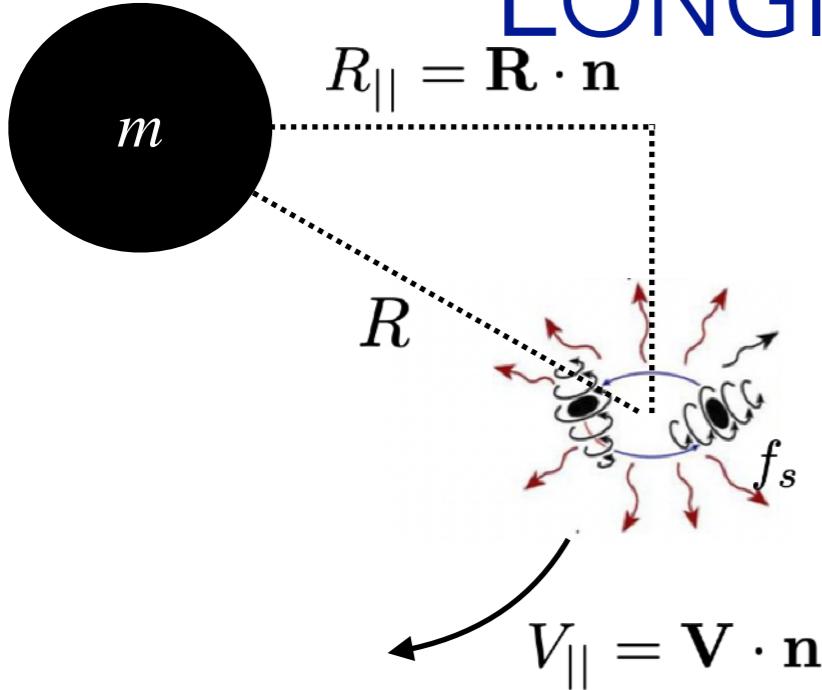
AGN: A FABRIC OF BBH MERGERS



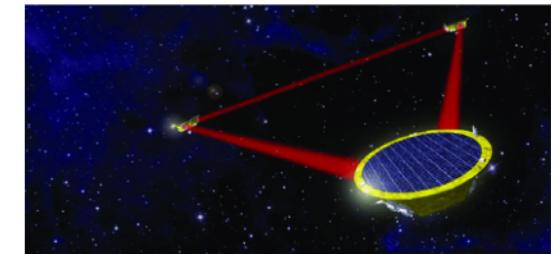
$$a \sim 700R_{\text{sch}}$$

Can the SMBH influence the waveform of the binary ?

LONGITUDINAL DOPPLER SHIFT

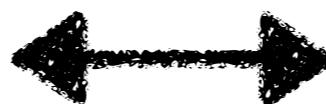


$$f_r = \frac{f_s}{1 + V_{||}}$$



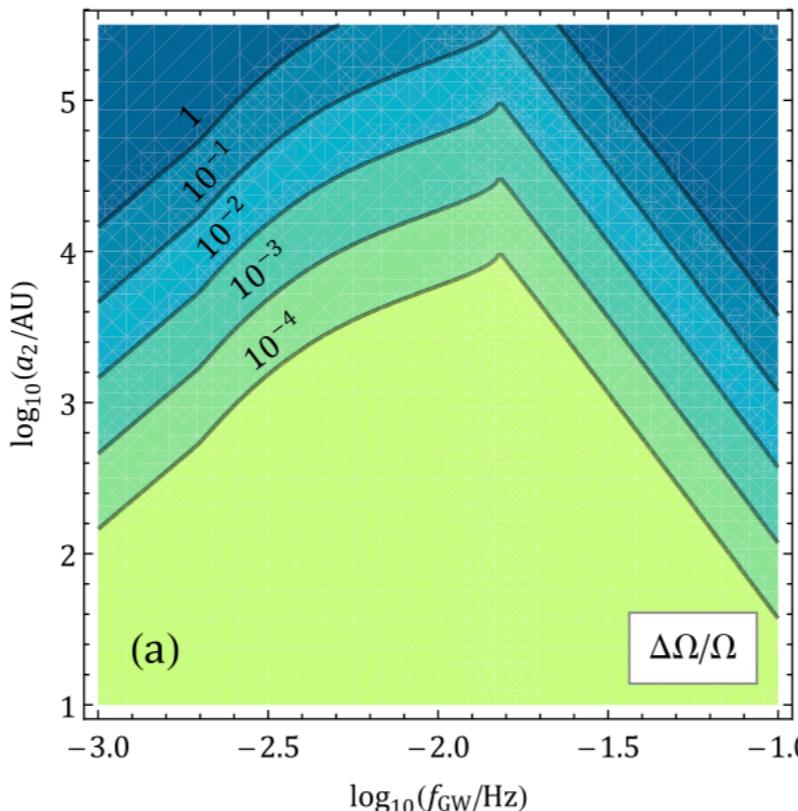
Temporal space

$$h_O(t) = h_S(t - R_{||}(t))$$



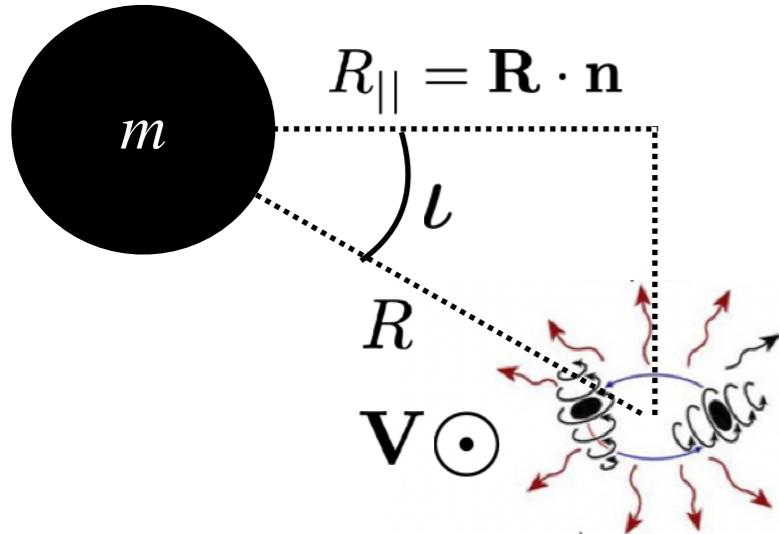
Frequency space

$$\phi_O(f) = \phi_S(f) + 2\pi f R_{||}(f)$$



Randall Xianyu (2019)

DEGENERESCENCE

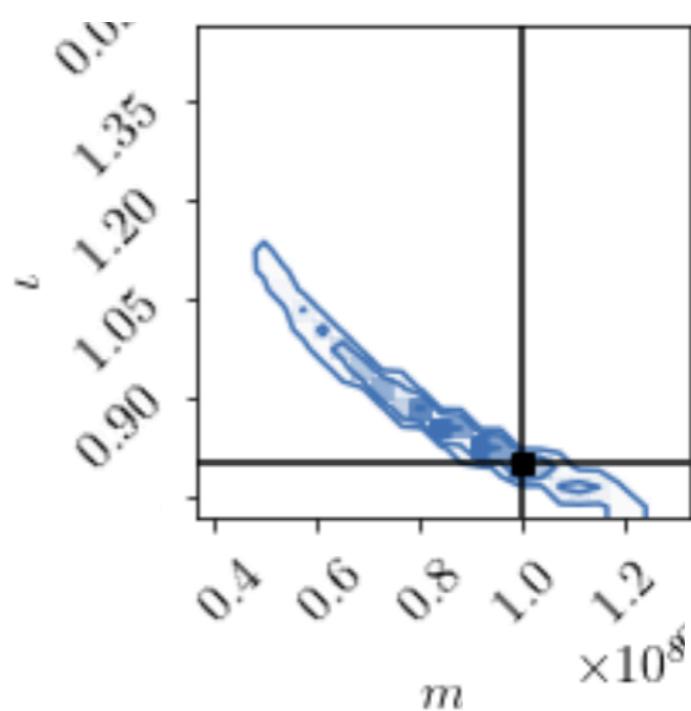


Observer →

$$R_{||} \simeq R \cos i \sin \left(\frac{2\pi t}{P} + \varphi \right)$$

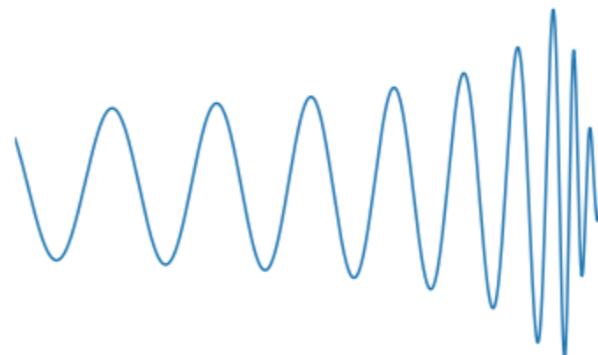
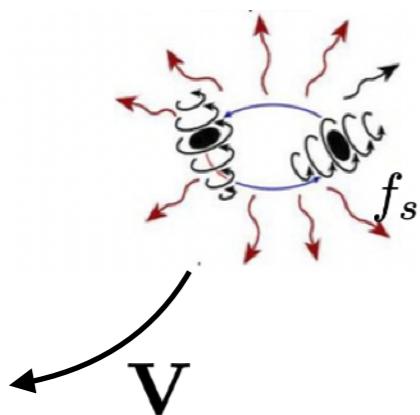
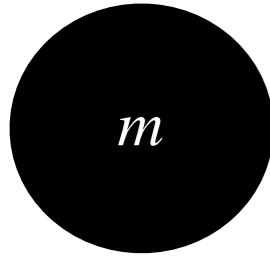
Measure P and R → $Gm = \frac{4\pi^2 R^3}{P^2}$

Mass-inclination degeneracy !

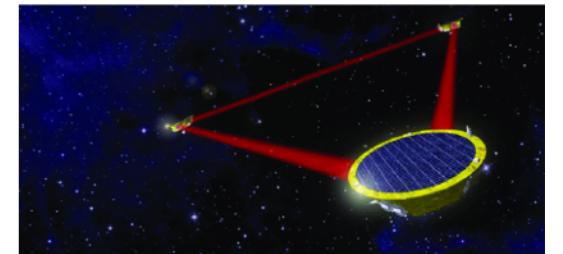


TRANSVERSE DOPPLER

AK, K. Leyde (2022)

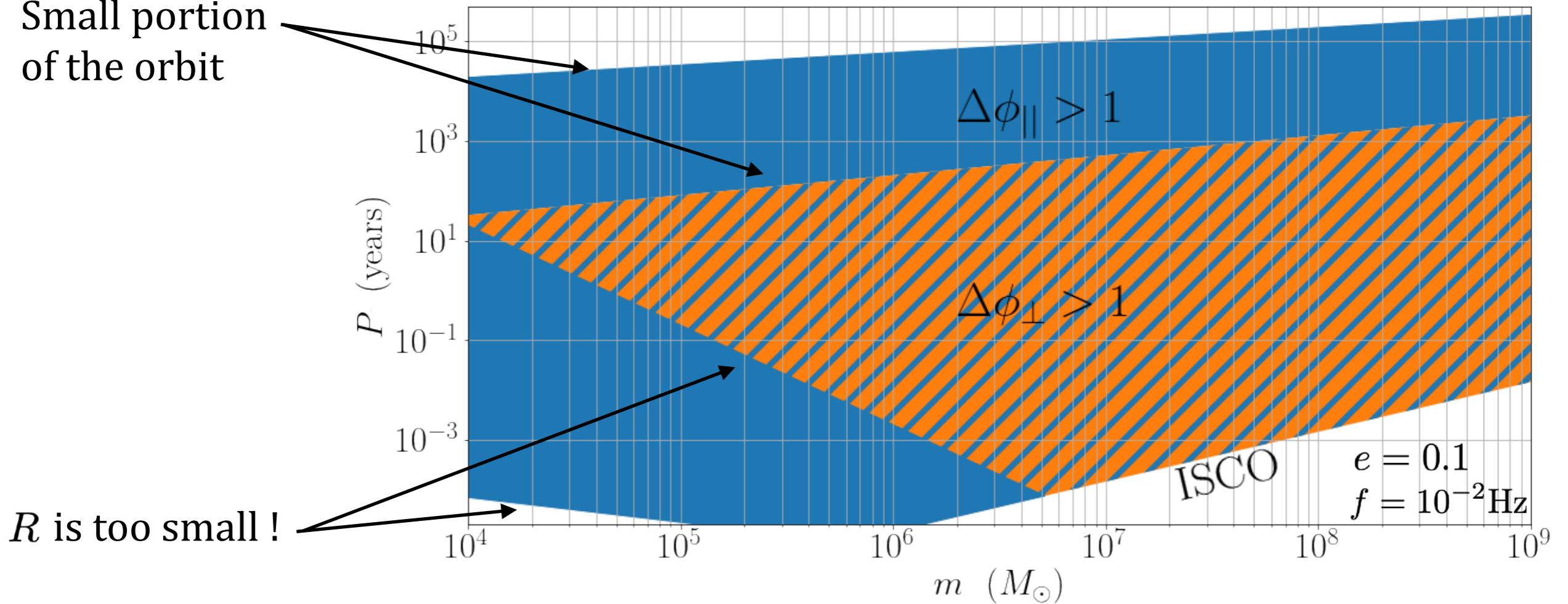


$$f_r = \frac{f_s \sqrt{-g_{\mu\nu} V^\mu V^\nu}}{1 + V_{||}}$$



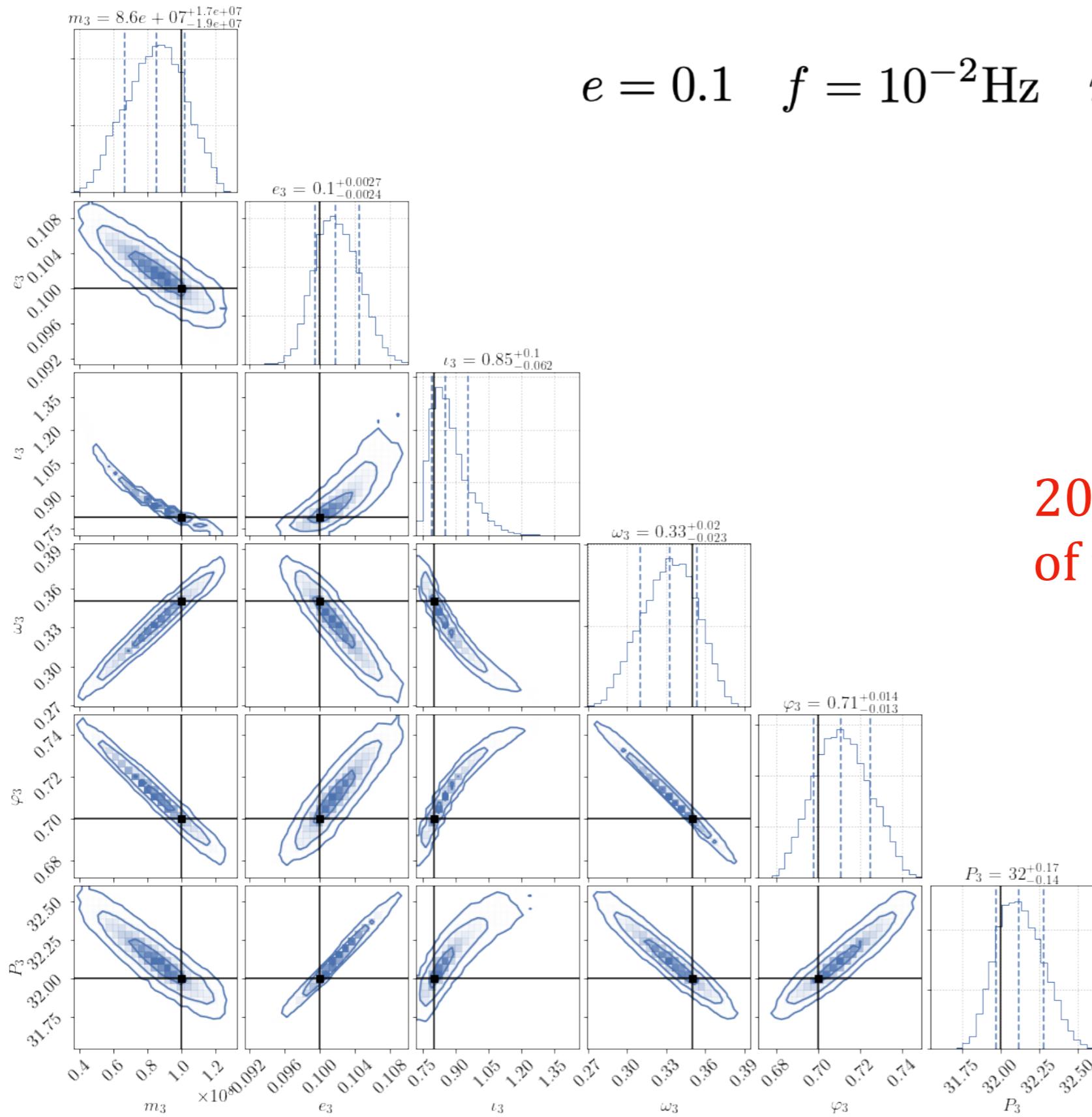
$$\sqrt{-g_{\mu\nu} V^\mu V^\nu} \simeq 1 - \frac{V^2}{2} - \frac{Gm}{R} \quad \rightarrow \quad \Delta\phi_{\perp} \simeq 4\pi f (GmR)^{1/2} e \sin\left(\frac{2\pi t}{P} + \varphi\right)$$

Small portion
of the orbit



MCMC

AK, K. Leyde (2022)



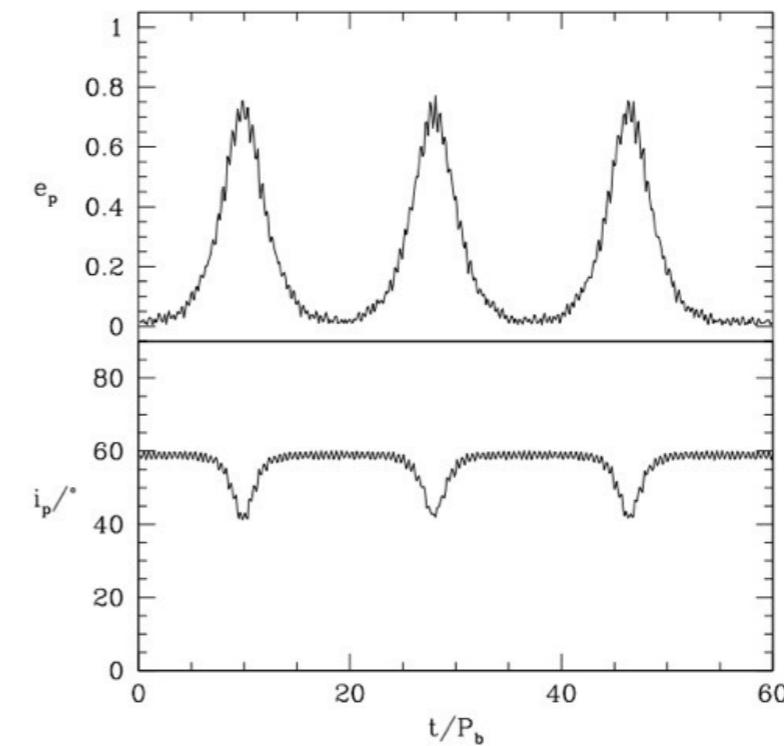
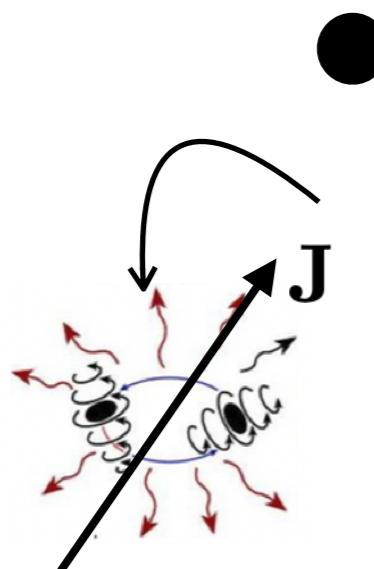
$$e = 0.1 \quad f = 10^{-2}\text{Hz} \quad m = 10^8 M_{\odot} \quad P = 32\text{yr}$$



20% determination
of the SMBH mass

CONCLUSIONS

- New way of measuring SMBH mass
- Other 3-body effects: spin-orbit coupling, Kozai-Lidov oscillations...



MCMC

AK, K. Leyde (2022)

