

 \exists ongoing Accelarating Expansion of the Universe

 \blacktriangle Standard Interpretation: Universe dominated by Dark Energy permeating all of space

 \blacktriangle in G.R. framework:

Einstein's equal with a positive cosmological constant of the order:

 $\Lambda \approx 10^{-120}$ (in M_{Planck}^4 units)

 \blacktriangle A rather intringuing coincidence:

 $m_{\nu}^4 \lesssim 10^{-116}$ (in M_{Planck}^4 units)

possible link between the two scales?

 \blacktriangle Simple Effective Field Theory description: with a scalar field, ϕ acquiring Potential Energy $V(\phi)$ with positive vacuum energy Λ :

 \blacktriangle de Sitter vacua \blacktriangle

...with some additional requirements: $\phi \rightarrow$ inflation suitable for inflation

The inference from the previous observations and remarks is that ^a variety of fundamental open questions involving ^a vast range of scales are intertwined !

Thence, it would be desirable to contemplate an effective theory with UV completion where Planck-scale Physics are naturally integrated

Currently, the most successful and robust candidate towards a UV completion is

String Theory

Focus of this talk:

EFT from type II-B/F-theory compactified on a Calabi-Yau (CY) Manifold

However

 \downarrow

 \triangle Compactifications characterised by large numbers of massless scalar fields (moduli)

 \star Two basic classes of moduli \star

Recall that a CY is a compact Kähler manifold which admits a Ricci-flat metric g with (closed) $(1,1)$ -Kähler form: $J = g_{i\bar{i}}dz^i \wedge d\bar{z}^{\bar{j}}$, $dJ = 0$

A CY can be deformed in two ways:

- 1. Variation of the Kähler structure $\delta g_{i\bar{j}}$ (mixed type), gives $h^{1,1}$ parameters ^a, the Kähler moduli T^k , $k = 1, 2, ..., h^{1,1}$.
- 2. Pure type metric variations g_{ij} , $g_{\overline{i}\overline{j}}$ giving rise to $h^{2,1}$ complex structure (CS) parameters z^a , $a = 1, 2, ..., h^{2,1}$, associated with:

$$
\Omega_{ijk}g^{k\bar{l}}\delta g_{\bar{l}\bar{m}}\,dz^i\wedge dz^j\wedge d\bar{z}^{\bar{m}}
$$

where Ω is a holomorphic 3-form.

 a_h ^{r,s} dim. of Dolbeault cohomology $H^{r,s} = \frac{\{\omega^{r,s}|\bar{\partial}\omega^{r,s}=0\}}{\{\alpha^{r,s}|\alpha^{r,s}=\bar{\partial}\beta^{r,s-1}\}}$

In addition:

∃ moduli and other fields associated with Type II-B closed string spectrum from L - and R -moving open strings with NS and R b.c. Λ (NS₊, NS₊) : Graviton, dilaton and Kalb-Ramond (KR)-field $g_{\mu\nu}, \phi, B_{\mu\nu} \rightarrow B_2, \quad (\text{def} : e^{\phi} = g_s)$ Λ (R_−, R_−) : Scalar, 2- and 4-index fields (p-form potentials) $\mathbf{C_0}, C_{\mu\nu}, C_{\kappa\lambda\mu\nu} \to C_p, \ p = 0, 2, 4$ 1. A C_0 , $\phi \rightarrow combined\ to\ axion-dilaton\ modulus:$

$$
S = C_0 + i e^{-\phi} \equiv C_0 + \frac{i}{g_s}
$$

2. \triangle Field strengths/magnetic fluxes: $F_p := d C_{p-1}, H_3 := d B_2, \Rightarrow G_3 := F_3 - S H_3$

$\downarrow \downarrow$

 \triangle we conclude that: \triangle $\#$ CY of **Compactifications** and $\#$ fluxes \Rightarrow **Enormous** number of String Vacua

\downarrow

String Landscape

▲ Long standing Question ▲

▲ Are there any de Sitter vacua in the Landscape?

... even if the answer is Yes... we know that they are...

 \Rightarrow Certainly Scarce \Leftarrow

Hence

 $\mathcal A$ Reasonable sequence of $\mathcal T$ asks in the context of type IIB theory:

A Provide masses to moduli fields \Rightarrow Stabilisation

 \triangle The quest for a *de Sitter* vacuum in String Theory (if possible... based only on perturbative corrections)

A Cosmological implications such as inflation

Implementation

A A

A Geometry of internal space. Assuming: *i*): a factorised $T^6 = T^2 \times T^2 \times T^2$ -torus. ii): $3 \times D7$ brane-stacks, each one spans 4 compact dimensions while localised at the remaining 2-d.

^N Context: Type II-B effective Supergravity: Basic 'ingredients': Superpotential W and Kähler potential $\mathcal K$

\blacktriangle The Superpotential \mathcal{W} \blacktriangle

 \blacktriangle A Flux-induced superpotential has been constructed (G.V.W. hep-th/9906070) using $G_3 = F_3 - SH_3$ and $(3,0)$ -form $\Omega(z_a)$:

$$
\mathcal{W}_0 = \int \, \mathbf{G_3} \wedge \Omega(z_a) \; \Rightarrow \; \mathcal{W}_0 = \mathcal{W}_0(z_a, S)
$$

does not depend on Kähler moduli T_i . W_0 must satisfy:

Flatness conditions \blacktriangle

 $\mathcal{D}_{z_0}\mathcal{W}=0$, $\mathcal{D}_S\mathcal{W}=0$:

 $\Rightarrow z_a$ and S stabilised \Leftarrow but!

Kähler moduli ∉ \mathcal{W}_0 ⇒ remain unfixed! A

The Kähler potential \blacktriangle

$$
\mathcal{K}_0 = -\sum_{i=1}^3 \ln(-i(T_i - \bar{T}_i)) - \ln(-i(S - \bar{S})) - \ln(i \int \Omega \wedge \bar{\Omega}) \cdot
$$

 \triangle The classical scalar potential identically zero: \triangle

$$
V = e^{\mathcal{K}} \left(\sum_{I,J} \mathcal{D}_I \mathcal{W}_0 \mathcal{K}^{I \bar{J}} \mathcal{D}_{\bar{J}} \mathcal{W}_0 - 3|\mathcal{W}_0|^2 \right) \equiv 0,
$$

due to flatness conditions and the no-scale structure.

⇓

Kähler moduli completely undetermined!

⇓

Task: Engineer the appropriate geometric set up and compute: Kähler moduli-dependent QUANTUM corrections

The Kähler potential K

 and PERTURBATIVE **String Loop Corrections**

Two types of expansions in String Theory: *i*) Inverse string tension $\propto \alpha'$. *ii*) String coupling g_s

$\triangle \alpha'^3$ Corrections 10-d action with α'^3 (see Becker et al hep-th/0204254): $\mathbf{S} \propto \int d^{10} \sqrt{-g} e^{-2\phi} \left(R + 4(\partial \phi)^2 + \alpha'^3 \nabla^2 \phi Q \right)^{\alpha}$

Compactifation \rightarrow redefinition of 4-d dilaton:

$$
e^{-2\phi_4} = e^{-2\phi_{10}}(\mathcal{V} + \xi/2)
$$

= $e^{-\frac{1}{2}\phi_{10}}(\hat{V} + \hat{\xi}/2)$ (Einstein frame)

where the 6d volume $\hat{\mathcal{V}}$ in Einstein frame are:

$$
\hat{\mathcal{V}} = \frac{1}{3!} \kappa_{ijk} \hat{t}^i \hat{t}^j \hat{t}^k
$$

with $\hat{t^i}$ defined through:

$$
t^k = -\text{Im}(T^k) = \hat{t}^k \left(\frac{S - \bar{S}}{2i}\right)^{-1/2} \equiv \hat{t}^k g_s^{1/2}
$$

^a $Q \rightarrow$ generalisation of 6-d Euler integrand $\int d^6x \sqrt{g}Q = \chi$

 ξ is expressed in terms of the Euler characteristic χ of the manifold:

$$
\xi = -\frac{\zeta(3)}{4(2\pi)^3} \chi = \hat{\xi} \left(\frac{S - \bar{S}}{2i} \right)^{-3/2} \equiv \hat{\xi} g_s^{3/2}
$$

 $\hat{\xi}$ is incorporated into the Kähler potential through the shift ^a

$$
\hat{\mathcal{V}} \rightarrow \mathcal{U}_0 = \hat{\mathcal{V}} + \frac{\hat{\xi}}{2} \equiv \hat{\mathcal{V}} + \frac{\xi}{2} \left(\frac{S - \bar{S}}{2i} \right)^{3/2} \equiv \hat{\mathcal{V}} + \frac{\xi}{2} \frac{1}{g_s^{3/2}}
$$

 α'^3 -modified Kähler potential:

$$
\mathcal{K}_0 \to \mathcal{K} = -\log(-i(S - \bar{S})) - 2\log \mathcal{U}_0 + \mathcal{K}_{cs}
$$

(where: $\mathcal{K}_{cs} = -\ln(i \int \Omega \wedge \bar{\Omega})$)

^aξ in the prepotential $F = \frac{i}{3!} k_{abc} \frac{X^a X^b X^c}{X^0} + \xi X_0$, Candelas et al, NPB (91)

▲ Loop Corrections ▲

 \triangle Previous α'^3 corrections at "tree-level" w.r.t. string-loop series.

\triangle Hypothesis: \triangle

Generic type of **one-loop correction** is captured by

$$
u_1 = \left(\frac{S - \bar{S}}{2i}\right)^{-1/2} f(\hat{\mathcal{V}}) \equiv g_s^{-1/2} f(\hat{\mathcal{V}})
$$

These are included by another shift:

$$
\hat{\mathcal{V}} \to \mathcal{U}_0 \to \left[\mathcal{U} \equiv \hat{\mathcal{V}} + \frac{\xi}{2} \frac{1}{g_s^{3/2}} + g_s^{1/2} f(\hat{\mathcal{V}}) \right]
$$

So, the final form of the corresponding Kähler potential is

$$
\mathcal{K} = -\log(-i(S - \bar{S})) - 2\log \mathcal{U} + \mathcal{K}_{cs}
$$

Evidence: recall that type IIB string theory admits $SL(2, \mathbb{Z})$ This implies invariance of the resulting EFT under some subgroup $\Gamma_s \subset SL(2,\mathbb{Z}).$

Motivation to look for a $SL(2, \mathbb{Z})$ completion.

⇓

Consider the non-holomorphic Eisenstein series: $\mathcal{E}_{\frac{3}{2}} \equiv E_{\frac{3}{2}}(S, S)$:

Observation:

first and second terms are associated with α'^3 and loop corrections.

\blacktriangle The form of $f(V)$ \blacktriangle

(Antoniadis, Chen, GKL, 1803.08941, EPJC-2019) Consider the set up :

 \triangle Configuration of Three intersecting D7-brane stacks

A 4-d Einstein-Hilbert $(\mathcal{E}\mathcal{H})$ term \mathcal{R}_4 in the bulk, generated from higher derivative terms in the 10-d string action (Antoniadis et al hep-th/9707013, etc.)

⇓⇓⇓

$$
f(\hat{\mathcal{V}}) = \sigma + \eta \log(\hat{\mathcal{V}})
$$

The coefficients η and σ are expressed in terms of $\xi \propto \chi$ (Antoniadis, Chen, GKL, JHEP-2020):

$$
\sigma = -\eta = \frac{\zeta(2)}{\zeta(3)}\xi,
$$

The following ratio is of particular interest:

$$
\frac{\hat{\xi}}{\hat{\eta}} = -\frac{\zeta(3)}{\zeta(2)} \frac{1}{g_s^2}
$$

with

$$
\hat{\xi} = \xi g_s^{-3/2} \ ; \ \ \hat{\eta} = g_s^{1/2} \eta
$$

Kähler moduli STABILISATION within a concrete Global Model: $(GKL \& Pramod Shukla \quad 2203.03362 ; JHEP-2022)$

<u>Kreuzer-Skarke</u> (KS) in hep-th/0002240 introduced toric methods to construct Calabi-Yau manifolds in terms of Reflexive Polyhedra

...exploring the KS dataset ... \Rightarrow

Explicit CY_3 Manifold

$$
h^{1,1} = 3, h^{2,1} = 115, \ \chi \equiv 2(h^{1,1} - h^{2,1}) = -224
$$

Assuming a basis of smooth divisors D_1, D_2, D_3 , the Kähler form is

$$
J = 2\sum_{k=1}^{3} t^k D_k
$$

and the case under consideration gives intersection polynomial with only one non-zero intersection:

$$
I_3 = 2D_1D_2D_3
$$

The 6d-volume :

$$
\mathcal{V} = 2 t^1 t^2 t^3 = \frac{1}{\sqrt{2}} \sqrt{\tau_1 \tau_2 \tau_3}
$$

 $(t^{i} \rightarrow 2-cycle, \tau_{i} \rightarrow 4-cycle \text{ moduli}, \text{ subject to } \tau_{i} = 2 t^{j} t^{k})$

Kähler potential including α' and loop corrections:

$$
\mathcal{K}(T_i, S, z_a) = -\log\{-i(S - \bar{S})\} - 2\ln\mathcal{U} + K_{cs}(z_a)
$$
 (1)

$$
\mathcal{U}(T_i, S) = \mathcal{V} + \frac{\hat{\xi}}{2} + \mathcal{U}_1 \tag{2}
$$

Assuming generic $\mathcal{U}_1(S, T^i)$ incorporating any loop corrections.

The effective potential

Identities to compute $K_j = \frac{\partial K}{\partial T_j}$, $K_{ij} = \frac{\partial^2 K}{\partial T_i \partial T_j}$, $K_{Sj} = \frac{\partial^2 K}{\partial S \partial T_j}$ etc:

$$
\tau_i = \frac{\partial \mathcal{V}}{\partial t_i} = \frac{1}{2} \kappa_{ijk} t^j t^k, \ A_{ij} t^i t^j = 6 \mathcal{V}, \ A^{ij} \tau_i \tau_j = 3 \mathcal{V}/2
$$

with A_{ij} the second derivatives of V and their inverse:

$$
A_{ij} = \frac{\partial \mathcal{V}}{\partial t_i \partial t_j} = \kappa_{ijk} t^k, \ \ A^{ij} = (\kappa_{ijk} t^k)^{-1}, \ A_{ik} A^{kj} = \delta_i^j \tag{3}
$$

Computing the inverse Kähler metric $K^{\overline{C}B}$ using $K_{A\overline{C}}K^{\overline{C}B}=\delta_A^B$: $K_{S\bar{S}}K^{\bar{S}S} + K_{S\bar{T}_i}K^{\bar{T}_jS} = 1$ $K_{S\bar{S}}K^{\bar{S}T_j} + K_{S\bar{T}_i}K^{\bar{T}_iT_j} = 0$ $K_{T_i\bar{S}}K^{\bar{S}T_j} + K_{T_i\bar{T}_k}K^{\bar{T}_kT_j} = \delta_i^j$. (4)

These lead to a simple analytic form in the basis S, T^i, z^a :

$$
\mathcal{K}^{A\bar{B}} = \begin{pmatrix} \tilde{\mathcal{P}}_1 & k_{\alpha} \tilde{\mathcal{P}}_2 & \mathcal{O} \\ k_{\alpha} \tilde{\mathcal{P}}_2 & k_{\alpha} k_{\beta} \tilde{\mathcal{P}}_3 - k_{\alpha \beta} \tilde{\mathcal{P}}_4 & \mathcal{O} \\ \mathcal{O} & \mathcal{O} & K_{cs}^{i\bar{j}} \end{pmatrix}
$$
(5)

 $\Rightarrow \text{block-diagonal for } K_{cs}^{i\bar{j}} \text{ but } S \text{ and } T_i \text{ } (\mathcal{V} \text{ }) \text{ mix : } \tilde{\mathcal{P}}_I = \tilde{\mathcal{P}}_I(\mathcal{V}, S).$

Master formula for F-term potential (for generic \mathcal{U}_1 loop corrections)

$$
V_{\alpha' + \text{generic}} = e^{\mathcal{K}} \left(\frac{3\mathcal{V}}{2\mathcal{U}^2} \left(1 + \frac{\partial \mathcal{U}_1}{\partial \mathcal{V}} \right)^2 \frac{4\mathcal{V}^2 + \mathcal{V}\hat{\xi} + 4\hat{\xi}^2}{\mathcal{V} - \hat{\xi}} - 3 \right) |W_0|^2
$$

⇓

A For α' and logarithmic corrections: $U_1 = -\hat{\eta} + \hat{\eta} \log \mathcal{V}$:

$$
V_{\alpha'+\log} = 12g_s e^{K_{cs}} |W_0|^2 \hat{\xi} \frac{\mathcal{V}^2 + 7\hat{\xi}\mathcal{V} + \hat{\xi}^2}{\left(\mathcal{V} - \hat{\xi}\right)\left(2\mathcal{V} + \hat{\xi}\right)^4}
$$

$$
-\frac{3\kappa}{2} |W_0|^2 \frac{2\hat{\eta} - \hat{\eta}\log\mathcal{V}}{2\mathcal{V}^3} + \cdots
$$

$$
\frac{logarithmic}{}
$$

Large Volume Limit

$$
V_F~\approx~C\frac{\hat{\xi}-4\hat{\eta}+2\hat{\eta}\log(\mathcal{V})}{\mathcal{V}^3}
$$

Properties

- Minimum exists for $\hat{\eta} < 0$.
- \triangle Stabilisation at large volume (in weak coupling g_s regime):

$$
\mathcal{V}_{\min} = e^{\frac{7}{3} + \frac{\hat{\xi}}{2|\hat{\eta}|}} \sim e^{\frac{1}{g_s^2}}
$$

 \triangle For F-term potential, AdS-minimum

$$
(V_F)_{\min}\propto \frac{\hat{\eta}}{\mathcal{V}^3}<0
$$

New contributions required to uplift to dS vacuum

 \triangle Uplift to dS occurs through D-terms $(Lüst et al hep-th/0609211; Antoniadis, Chen, GKL 1803.08941)$ associated with universal $U(1)$'s of D7-stacks:

$$
V_{\mathcal{D}} = \frac{g_{D7_i}^2}{2} \left(Q_i \partial_{T_i} K + \sum_j q_j |\Phi_j|^2 \right)^2, \frac{1}{g_{D7_i}^2} = \text{Re} T_i + \cdots
$$

Minimising the total potential:

 $V_{\text{eff}} = V_F + V_D$

 \Rightarrow a minimum and a maximum defined by the double-valued Lambert W-function (i.e., solution of $\mathbf{W}e^{\mathbf{W}} = z$):

$$
\mathcal{V}_{\min} = \frac{\hat{n}}{d} \mathbf{W_{0/-1}} \left(\frac{d}{\hat{n}} e^{\frac{7}{3} - \frac{\hat{\xi}}{2\hat{n}}} \right)
$$

This dramatically constrains acceptable string vacua (fluxes etc)

de Sitter

Plot of V_{eff} vs V for $d' = 10^4 d = \{8.65, 8.85, 9.15\}.$

The lower curve corresponds to AdS vacuum. At large volume, the potential vanishes asymptotically

Inflation

Hybrid scenario can be realised with open string states χ at D7-brane intersections playing the role of waterfall fields (Antoniadis, Lacombe, GKL, 2109.03243, JHEP 2022)

Shape of V_{total} in the presence of χ at large $\mathcal V$ regime:

$$
V_{total} = C \frac{\hat{\xi} - 4\hat{\eta} + 2\hat{\eta}\log(V)}{\mathcal{V}^3} + \frac{d}{\mathcal{V}^2} + V_{\chi}
$$

with V_{χ} the waterfall field potential:

$$
V_{\chi} \;\; = \;\; \sim m^2({\cal V}) \chi^2 + \lambda({\cal V}) \chi^4
$$

The volume modulus can play the role of Inflaton field

 $\phi \propto \log \mathcal{V}$

We find that inflation can be realised with most of the 60 efolds collected near the metastable local mimimum $V_{total}(\mathcal{V}_{min})$. \triangle Inflation ends and false vacuum decays to *Global minimum* through a

waterfall field $\chi: V_{\chi} \sim m^2(\mathcal{V}) \chi^2 + \lambda(\mathcal{V}) \chi^4$.

For $m^2 > 0$ minimum in the *χ*-field direction is at the origin

 $m^2 > 0 \rightarrow \langle \chi \rangle = 0$

When the mass of χ becomes tachyonic, a phase transition occurs and the new vacuum is obtained at a non-vanishing $\langle \chi \rangle$:

$$
m^2 < 0 \ \to \ \langle \chi \rangle \neq 0
$$

A configuration to realise the hybrid scenario in our D7 set-up

- A circled cross shows magnetic field on specific $D7$ and T^2 .
- \triangle $\mathcal{A}_{1,3}$ denote Wilson lines
- \triangle $\pm x_2$ brane separations (uplifting tachyons)
- [⇒] only one tachyonic state ^playing the role of waterfall field:

$$
\alpha' m_{22}^2 \approx -\frac{A}{\mathcal{V}^{1/3}} + B\mathcal{V}^{1/3}, \ \ (A, B) \rightarrow positive\ constants
$$

$$
\langle \chi \rangle \neq 0
$$
 for $\mathcal{V} < \mathcal{V}_{\text{critical}} = \left(\frac{A}{B}\right)^{\frac{3}{2}}$

\bigstar IIB/F-theory:

• Stabilisation of Kähler Moduli possible with Perturbative corrections only:

$$
\mathcal{K} = -2\ln\left(\mathcal{V} + \hat{\xi}/2 + \hat{\eta}\ln\mathcal{V}\right) + \cdots
$$

Origin of log-corrections:

Induced Einstein-Hilbert terms from R^4 -couplings in 10-d theory. This $\mathcal{E}\mathcal{H}\text{-term} \quad \exists \text{ in } 4d \text{ only!}$

⇓

 \star *induced* \mathcal{EH} -term ... indispensable element for a: 4d de Sitter Universe **★** Hybrid Inflation with inflaton $\phi \sim \log V$ and waterfall fields open string states attached on D7's.

AA

D7-branes and Logarithmic corrections

Two ingredients needed for log-corrections:

AA

 \mathcal{A}) Intersecting **D7-brane configuration**:

 \mathcal{B}) Higher derivative couplings in curvature

(generated by multigraviton scattering) (see hep-th/9704145; 9707013; 9707018)

Leading correction term in type II-B action: proportional to the fourth power of curvature: $\propto R^4$

After reduction on $\mathcal{M}_4 \times \mathcal{X}_6$, (with \mathcal{M}_4 4-d Minkowski) R^4 induces a novel Einstein-Hilbert term $\mathcal{R}_{(4)} \propto$ by the Euler characteristic χ :

$$
\propto \chi \int_{M_4} (\zeta(2) - \zeta(3)e^{-2\phi}) \mathcal{R}_{(4)},
$$

induced $\mathcal{E}H$ term

 $\blacktriangle\blacktriangle$ this $\mathcal{E}\mathcal{H}$ term possible in 4-dimensions only!

A New $\mathcal{E}\mathcal{H}$ -term localised at points with $\chi \neq 0$ A KK-exchange between graviton vertex and a $D7$ -brane Momentum D7 $k_I \, \zeta$ Space $\sum_{\geq R_1^{-1}} k_3$ $y=0$ $y=y_{p}$ Graviton vertex k_3 \overline{M} Worldsheet

Details of the Specific CY manifold

The analysis of the divisor topologies using cohomCalg shows that divisors are of ^K³ and SD types and can be represented by the following Hodge diamonds:

 $K3 \equiv$ 1 $\begin{matrix} 0 & 0 \\ 0 & 0 \end{matrix}$ $\begin{array}{ccccccccc}\n1 & & 20 & & 1 \end{array},$ 0 0 1 $SD \equiv 27$ 1 $\begin{matrix} 0 & 0 \\ 0 & 0 \end{matrix}$ 27 184 27 0 0 1

.

Cancellation of all D7-charges

Introduce N_a D7-branes wrapped around divisors D_a and orientifold images D_a' (0811.2936)

$$
\sum_{k} N_{k} ([D_{k} + D'_{k}]) = 8[O7]
$$

D7-branes and O7-planes also give rise to D3-tadpoles which receive contributions also from background 3-form fluxes

Assuming simple case:

D7-tadpoles are cancelled by placing $4\,D7 + D7'$ -branes on top of O7-plane:

$$
N_{D3} + \frac{1}{2} N_{\text{flux}} + N_{\text{gauge}} = \frac{1}{4} (O3 + \chi(O7))
$$

Example

Specific brane setting involving 2 stacks of D7-branes wrapping the divisors D_1, D_6 in the basis,

$$
8[O7] = 4 ([D_1 + D'_1]) + 4 ([D_6 + D'_6])
$$

D3 tadpole condition

$$
N_{D3} + \frac{1}{2} N_{flux} + N_{gauge} = 12
$$

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