Progress in the Standard Model precision calculations

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Increasing precision allows us to change paradigms

- ► Tycho de Brahe (~ 1601) Mars orbits, Rudolphine tables; → Johannes Kepler (~ 1609) - planets laws;
- O. Lummer et al. (1900) blackbody radiation
 → M. Planck (1900) quanta;
 EXPERIMENT → THEORY
- Gravitational waves;

 $\mathsf{THEORY} \to \mathsf{EXPERIMENT}$

Particle Physics

(i) muon discovery, J/Ψ (ii) $(g-2)_e$, $(g-2)_{\mu}$ (iii) V-A, parity; Note the 100th Birthday Anniversary of Prof. Chen Ning Yang, link

► (i)
$$\tau^{\pm}$$
 (tau lepton);
(ii) Tevatron - top quark discovery;
(iii) H^0 (scalar Higgs-Englert boson)
SM corrections matters! LEP, SLAC, LHC (see backup slides)

- \rightarrow M. Veltman (1977) ρ -parametr $\sim m_t^2, \ln(m_H^2);$
- Neutrinos (masses, mixing angles, CP phase(s));

Future Colliders (FC); THEORY ↔ EXPERIMENT

Precision, true inception



Abri Blanchard bone (\sim 30 000 BC),

Alexander Marschack, 'Cognitive Aspects of Upper Paleolithic Engraving', Current Anthropology (1972) Vol.13, 3/4

Interpretation: Chantal Jegues-Wolkiewicz \longrightarrow probably the first lunar calendar

Similarly Lascaux caves' paintintings, \sim 17 000 BC.

Two ways for discoveries (in both cases precision is crucial):

- 1. within the known theory (anomalies¹)
- 2. new processes and (rare) phenomena;

¹'I have always suspected that, one day, (...) they [JG: experimentalists] would like to see what would happen, just for the fun of it, if they falsely report that there exists a certain bump, or an oscillation in a certain curve, and see how the theorists predict it. I know these men so well that the moment I thought of that possibility I have honestly always been concerned that some day they will do just that. Then you can imagine how absurd the theoretical physicists would sound, making all these complicated calculations to demonstrate the existence of such a bump, while these fellows are laughing up their sleeves.' – R.P. Feynman)

I have chosen to discuss:

(i) Z-pole precision physics;
(ii) Neutrino precision physics ²;

 \longrightarrow K. Grzanka, continued

²Massive neutrinos (RHNs) are in the (non-minimal?) SM.

NOT COVERED

Future: W, t, H

▶ $e^+e^- \rightarrow W^+W^-$ at 161 GeV: $\delta m_W^{exp} = 0.5 \div 1$ MeV. Challenge to get the same TH error: NNLO $e^+e^- \rightarrow 4f$.

► $e^+e^- \rightarrow t\bar{t}$ at 350 GeV: $\delta m_t^{exp} = 17$ MeV Big challenge for theory, today > 100 MeV, future projection \leq 50 MeV: \sim 10 MeV unc. from mass def.; \sim 15 MeV from α_s unc. to threshold mass def.; \sim 30 MeV - h. orders resummation

► $e^+e^- \rightarrow HZ$ at 240 GeV: Kinematic constraint fits with $Z \rightarrow ll$ and $H \rightarrow bb$, ..., $m_H = 125.35$ GeV ±150 MeV [link CMS], $\Gamma_H = 4.1^{5.1}_{4.0}$ MeV, $\Gamma_H < 13$ MeV at 95 % C.L., 1901.00174 $\delta m_H^{exp} = 10$ MeV; Theory errors subdominant.

Monte Carlo generators (not discussed!) 'QED challenges at FCC-ee precision measurements',

S. Jadach and M. Skrzypek, Eur.Phys.J.C 79 (2019) 9, 756 1903.09895

Workshop: Precision calculations for future e+e colliders: targets and tools,



CERN 2022, talk by V. Sotnikov

LHC, HL-LHC: QCD, EW-QCD, ...

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NEEDS FOR PRECISION: THE FUTURE

Motivation For Precision Studies: Z,W,H,t and flavour electroweak factories



https://arxiv.org/abs/2203.06520 [The Future Circular Collider: a Summary for the US 2021 Snowmass Process]

Phase	Run duration (years)	Center-of-mass Energies	Integrated Luminosity	Event Statistics
		(GeV)	(ab ⁻¹)	
FCC-ee-Z	4	88-94	150	$5 \cdot 10^{12}$ Z decays
FCC-ee-W	2	157-163	10	10 ⁸ WW events
FCC-ee-H	3	240	5	10^6 ZH events 25k WW $\rightarrow H$
FCC-ee-tt	5	340-365	0.2 ÷1.5	$\begin{array}{c} 10^6 \ t\overline{t} \text{ even ts} \\ 200 \text{ k ZH} \\ 50 \text{ k WW} \rightarrow H \end{array}$

Jorgen D'Hondt, "Strategies and plans for particle physics in Europe",

Epiphany 2021, https://indico.cern.ch/event/934666



Tera-Z Physics:

BREATHTAKING

EXPERIMENTAL

PRECISION

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Eur. Phys. J. Plus (2022) 137:92

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Table 3 Measurement of selected precision measurements at FCC-ee, compared with present precision. Statistical errors are indicated in boed phase. The systematic uncertainties are initial estimates, aim is to improve down to statistical errors. This set of measurements, together with those of the Higgs properties, achieves indirect sensitivity to new physics up to a scale Λ of 70 TeV in a description with dim 6 operators, and possibly much higher in specific new physics (non-decoupling) models

Observable	Present value \pm error	FCC-ee stat.	FCC-ee syst.	Comment and leading exp. error
m _Z (keV)	91186700 ± 2200	4	100	From Z line shape scan
				Beam energy calibration
Γ_Z (keV)	2495200 ± 2300	4	25	From Z line shape scan
				Beam energy calibration
$\sin^2\theta_W^{\text{eff}}(\times 10^6)$	231480 ± 160	2	2.4	from $A_{FB}^{\mu\mu}$ at Z peak
				Beam energy calibration
$1/\alpha_{\text{QED}}(m_Z^2)(\times 10^3)$	128952 ± 14	3	Small	From $A_{FB}^{\mu\mu}$ off peak
				QED&EW errors dominate
R^Z_ℓ (×10 ³)	20767 ± 25	0.06	0.2-1	Ratio of hadrons to leptons
-				Acceptance for leptons
$\alpha_{s}(m_{Z}^{2}) \ (\times 10^{4})$	1196 ± 30	0.1	0.4-1.6	From R^{Z}_{ℓ} above
$\sigma_{\rm had}^0 \; (\times 10^3) \; ({\rm nb})$	41541 ± 37	0.1	4	Peak hadronic cross section
				Luminosity measurement
$N_{\nu}(\times 10^3)$	2996 ± 7	0.005	1	Z peak cross sections
				Luminosity measurement
$R_{b} (\times 10^{6})$	216290 ± 660	0.3	< 60	Ratio of bb to hadrons

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LEP/SLC HAS PAVED THE WAY

Experimental measurements at Z-pole: after unfolding (\rightarrow Z. Was talk)



QED unfolding

Altogether $17 \cdot 10^6$ Z-boson decays at LEP Cross section : Z mass and width dnad [dn] σ ALEPH DELPHI L3 OPAL 30 20 measurements (error bars increased by factor 10) 10 ofrom fit OED correcte M 92 94 E_{cm} [GeV] 86 88 90 ~30% QED corrections (ISR)

Z-resonance: QED and EW, Standard Model Theory for the FCC-ee Tera-Z

stage, https://e-publishing.cern.ch/index.php/CYRM/issue/view/89

1. Z-resonance and $\gamma, Z', \ldots \longrightarrow$ Laurent series,

$$\mathcal{M} = \frac{R}{s - s_0} + \sum_{n=0}^{\infty} (s - s_0)^n \ B^{(n)}, \quad s_0 = \bar{M}_Z^2 + i\bar{M}_Z\bar{\Gamma}_Z.$$

2. We want to extract EW Z-vertex couplings and definitions like $\sin^2 \theta_{\text{eff}}^f$, but in reality, we deal with complicated process

$$e^+e^- \rightarrow f^+f^-$$
 + invisible $(n \ \gamma + e^+e^- \text{pairs} + \cdots)$

$$\sigma^{e^+e^- \to f^+f^- + \dots}(s) = \int dx \ \widehat{f(x)} \ \underbrace{\sigma^{e^+e^- \to f^+f^-}(s')}_{\bullet} \ \delta(x - s'/s)$$

 \longrightarrow form factors, QED separation/deconvolution, non-factorizations $_{(\text{backup slides})}$

To determine the structure function/flux function kernels and hard scattering ansatz for data preparation or for unfolding is one of the challenges of the Tera-Z physics.

EWPOs and $N^{x}LO$ SM CORRECTIONS

Rough scheme for extracting the $Z f \bar{f}$ vertex and EW corrections



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EWPOs, Z pole

$$\begin{split} &\sigma_{\rm had}^0 &= &\sigma[e^+e^- \to {\rm hadrons}]_{s=M_{\rm Z}^2}, \\ &\Gamma_Z &= &\sum_f \Gamma[Z \to f\bar{f}], \\ &R_\ell &= &\frac{\Gamma[Z \to {\rm hadrons}]}{\Gamma[Z \to \ell^+\ell^-]}, \quad \ell=e,\mu,\tau, \\ &R_q &= &\frac{\Gamma[Z \to q\bar{q}]}{\Gamma[Z \to {\rm hadrons}]}, \quad q=u,d,s,c,b. \end{split}$$

The remaining EWPOs are cross section asymmetries, measured at the Z pole, e.g., forward-backward asymmetry

$$A_{\rm FB}^f = \frac{\sigma_f \left[\theta < \frac{\pi}{2}\right] - \sigma_f \left[\theta > \frac{\pi}{2}\right]}{\sigma_f \left[\theta < \frac{\pi}{2}\right] + \sigma_f \left[\theta > \frac{\pi}{2}\right]},$$

where θ is the scattering angle between the incoming e^- and the outgoing f.

EWPOs and Form Factors



Note approximate factorization of weak couplings

$$A_{FB} = \frac{\left[\int_0^1 d\cos\theta - \int_{-1}^0 d\cos\theta\right] \frac{d\sigma}{d\cos\theta}}{\sigma_T} \sim \underbrace{\frac{A_e}{2a_e v_e}}_{a_e^2 + v_e^2} \underbrace{\frac{2a_f v_f}{2a_f v_f}}_{a_f^2 + v_f^2} + \text{corrections}$$

$$\begin{split} \mathbf{A}_{f} &= \quad \frac{2\Re e_{a_{f}}^{v_{f}}}{1+\left(\Re e_{a_{f}}^{v_{f}}\right)^{2}} = \frac{1-4|Q_{f}|\mathrm{sin}^{2}\,\theta_{\mathrm{eff}}^{\mathrm{f}}}{1-4|Q_{f}|\mathrm{sin}^{2}\,\theta_{\mathrm{eff}}^{\mathrm{f}} + 8(Q_{f}\mathrm{sin}^{2}\,\theta_{\mathrm{eff}}^{\mathrm{f}})^{2}},\\ \mathrm{sin}^{2}\,\theta_{\mathrm{eff}}^{\mathrm{f}} &= \quad F\left(\Re e_{a_{f}}^{v_{f}}\right) \end{split}$$

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A. Freitas et al., "Theoretical uncertainties for electroweak and Higgs-boson precision measurements at FCC-ee", https://arxiv.org/abs/1906.05379

Quantity	FCC-ee	Current intrinsic error		Projected intrinsic error
				(at start of FCC-ee)
$M_{\rm W}$ [MeV]	0.5–1 [‡]	4	$(\alpha^3, \alpha^2 \alpha_{\rm s})$	1
$\sin^2 \theta_{\rm eff}^{\ell} [10^{-5}]$	0.6	4.5	$(\alpha^3, \alpha^2 \alpha_{\rm s})$	1.5
$\Gamma_{\rm Z}$ [MeV]	0.1	0.4	$(\alpha^3, \alpha^2 \alpha_{\rm s}, \alpha \alpha_{\rm s}^2)$	0.15
$R_b [10^{-5}]$	6	11	$(\alpha^3, \alpha^2 \alpha_s)$	5
$R_l \ [10^{-3}]$	1	6	$(\alpha^3, \alpha^2 \alpha_{ m s})$	1.5

[‡]The pure experimental precision on $M_{\rm W}$ is $\sim 0.5 \,{\rm MeV}$.

Quantity	FCC-ee	future parametric unc.	Main source
M_{W} [MeV]	0.5 - 1	1 (0.6)	$\delta(\Delta \alpha)$
$\sin^2 \theta_{eff}^{\ell} [10^{-5}]$	0.6	2 (1)	$\delta(\Delta \alpha)$
Γ _Z [MeV]	0.1	0.1 (0.06)	$\delta \alpha_s$
$R_b [10^{-5}]$	6	< 1	$\delta \alpha_{s}$
$R_{\ell} [10^{-3}]$	1	1.3 (0.7)	$\delta \alpha_{s}$

Important input parameter errors are $\delta(\Delta \alpha) = 3 \cdot 10^{-5}$, $\delta \alpha_s = 0.00015$. $M_H, m_t, \ldots \longrightarrow K$. Grzanka's talk E.g. the bosonic 2-loop corrections shift the value of Γ_Z by 0.51 MeV when using M_W as input and 0.34 MeV when using G_{μ} as input.

Please recall: $\delta \Gamma_{Z, \text{FCC}-\text{ee}} = 0.1 \text{ MeV}$

Dubovyk et al, https://doi.org/10.1016/j.physletb.2018.06.037

Γ_i [MeV]	$\Gamma_e, \Gamma_\mu, \Gamma_\tau$	$\Gamma_{\nu_e}, \Gamma_{\nu_{\mu}}, \Gamma_{\nu_{\tau}}$	Γ_d, Γ_s	Γ_u, Γ_c	Гь	$\Gamma_{\rm Z}$
Born	81.142	160.096	371.141	292.445	369.56	2420.2
$O(\alpha)$	2.273	6.174	9.717	5.799	3.857	60.22
$O(\alpha \alpha_{\rm S})$	0.288	0.458	1.276	1.156	2.006	9.11
$\mathcal{O}(N_f^2 \alpha^2)$	0.244	0.416	0.698	0.528	0.694	5.13
$\mathcal{O}(N_f \alpha^2)$	0.120	0.185	0.493	0.494	0.144	3.04
$\mathcal{O}(\alpha_{\rm bos}^2)$	0.017	0.019	0.058	0.057	0.167	0.505
$\mathcal{O}(\alpha_{t}\alpha_{s}^{2}, \alpha_{t}\alpha_{s}^{3}, \alpha_{t}^{2}\alpha_{s}, \alpha_{t}^{3})$	0.038	0.059	0.191	0.170	0.190	1.20

- ▶ 2016, estimation, <u>bosonic NNLO $\sim 0 \pm 0.1$ MeV</u> (see backup slides) 2018, exact result: 0.505 MeV
- * Fixed values of M_W

Complicated subject, theoretical, parametric errors.

 Standard Model Theory for the FCC-ee Tera-Z stage, https://e-publishing.cern.ch/index.php/CYRM/issue/view/89
 Theoretical uncertainties for electroweak and Higgs-boson precision measurements at FCC-ee, https://arxiv.org/abs/1906.05379

Errors, a simple observation:

- 1. Lack of knowledge about HO corrections is a real pain; estimates even in the perturbative regime can differ substantially from concrete results.
- 2. Estimations for each next piece of HO take into account **AMOUNT** of the correction
- 3. Real calculation gives a **CONCRETE** number, with an error which is at least 2 digits.

These points are essential when we are at the level of accuracy which approaches experimental precision.

Updates for error estimations

 \blacksquare Theory error estimate is not well defined, ideally $\Delta_{th} \ll \Delta_{exp}$

- Common methods: Count prefactors (α , N_c , N_f , ...)
 - Extrapolation of perturbative series
 - Renormalization scale dependence
 - Renormalization scheme dependence
- Also parametric error from external inputs $(m_t, m_b, \alpha_s, \Delta \alpha_{had}, ...)$

see, Ayres Freitas: https://arxiv.org/abs/1604.00406

E.g.: Intrinsic theory error estimation for Γ_Z , 1804.10236 [1604.00406]

1. Geometric series

$$\delta_{1}: \mathcal{O}(\alpha^{3}) - \mathcal{O}(\alpha_{t}^{3}) \sim \frac{\mathcal{O}(\alpha^{2}) - \mathcal{O}(\alpha_{t}^{2})}{\mathcal{O}(\alpha)} \mathcal{O}(\alpha^{2}) \sim 0.20 \text{ MeV} [0.26 \text{ MeV}]$$

$$\delta_{2}: \mathcal{O}(\alpha^{2}\alpha_{s}) - \mathcal{O}(\alpha_{t}^{2}\alpha_{s}) \sim \frac{\mathcal{O}(\alpha^{2}) - \mathcal{O}(\alpha_{t}^{2})}{\mathcal{O}(\alpha)} \mathcal{O}(\alpha\alpha_{s}) \sim 0.21 \text{ MeV} [0.3 \text{ MeV}]$$

$$\delta_{3}: \mathcal{O}(\alpha\alpha_{s}^{2}) - \mathcal{O}(\alpha_{t}\alpha_{s}^{2}) \sim \frac{\mathcal{O}(\alpha\alpha_{s}) - \mathcal{O}(\alpha_{t}\alpha_{s})}{\mathcal{O}(\alpha)} \mathcal{O}(\alpha\alpha_{s}) \sim 0.23 \text{ MeV}$$

$$\delta_{4}: \mathcal{O}(\alpha\alpha_{s}^{3}) - \mathcal{O}(\alpha_{t}\alpha_{s}^{3}) \sim \frac{\mathcal{O}(\alpha\alpha_{s}) - \mathcal{O}(\alpha_{t}\alpha_{s})}{\mathcal{O}(\alpha)} \mathcal{O}(\alpha\alpha_{s}^{2}) \sim 0.035 \text{ MeV}$$

 $\delta_5: \mathcal{O}(\alpha_{bos}^2) \sim \mathcal{O}(\alpha_{bos})^2 \sim \mathbf{0.1} \text{ MeV } \underline{[\text{Now we know it!}]}$

Total:
$$\delta\Gamma_Z = \sqrt{\sum\limits_{i=1}^5 \delta_i^2} \sim 0.4$$
 MeV [0.5 MeV]

Summary: estimations for higher order EW and QCD corrections

δ_1 :	δ_2 :	δ_3 :	δ_4 :	δ_5 :	$\delta\Gamma_Z$ [MeV]	
$\mathcal{O}(lpha^3)$	${\cal O}(lpha^2 lpha_{ m s})$	${\cal O}(lpha lpha_{ m s}^2)$	${\cal O}(lpha lpha_{ m s}^3)$	$\mathcal{O}(\alpha_{bos}^2)$	$=\sqrt{\sum_{i=1}^5 \delta_i^2}$	
TH1 (estimated error limits from geometric series of perturbation)						
0.26	0.3	0.23	0.035	0.1	0.5	
TH1-new (estimated error limits from geometric series of perturbation)						
0.2	0.21	0.23	0.035	$< 10^{-4}$	0.4	
δ_1' :	δ_2' :	δ_3' :	δ_4 :		$\delta\Gamma_Z$ [MeV]	
$\mathcal{O}(N_f^{\leq 1}\alpha^3)$	$\mathcal{O}(lpha^3 lpha_{ m s})$	$\mathcal{O}(lpha^2 lpha_{ m s}^2)$	${\cal O}(lpha lpha_{ m s}^3)$		$\sqrt{\delta_1'^2 + \delta_2'^2 + \delta_2'^3 + \delta_4^2}$	
TH2 (extrapolation through prefactor scaling)						
0.04	0.1	0.1	0.035	10^{-4}	0.15	

A. Blondel et al., "Theory for the FCC-ee : Report on the 11th FCC-ee Workshop Theory and Experiments",

https://e-publishing.cern.ch/index.php/CYRM/issue/view/110

1. Calculating N^3LO with 10% accuracy (two digits), we can replace intrinsic error estimation

$$\delta\Gamma_Z = \sqrt{\sum_{i=1}^5 \delta_i^2} \sim 0.4 \text{ MeV}$$

by

$$\delta \Gamma_Z = \sqrt{\sum_{i=1}^{5} (\delta_i / 10)^2} \sim 0.04 \text{ MeV.}$$

2. The requirement FCC-ee^{*exper. error*} (Γ_Z) ~ 0.1 MeV can be met.

METHODS AND TOOLS

Direct numerical approach³

Sector decomposition (SD) method:

- FIESTA [2016], [A.V.Smirnov]
- pySecDec [2022], [Expansion by regions with pySecDec],
- The Mellin-Barnes (MB) method:
 - MB [M.Czakon, 2006]

MBnumerics [J.Usovitsch, I.Dubovyk, T.Riemann, 2015] – Minkowskian kinematics

- Differential equations (DEs) method:
 - DiffExp [F. Moriello, 2019; M. Hidding, 2021],
 - AMFlow [X. Liu, Y.-Q. Ma, 2022] AMFlow,
 - SeaSyde [T. Armadillo, R. Bonciani, S. Devoto, N. Rana, A. Vi, 2022]

 $M_H, m_t, \ldots \longrightarrow \mathsf{K}$. Grzanka's talk

³All programs are public

SM and BSM: NEUTRINOS*

* Neutrino physics itself enters the precision era (mass ordering, C-nature, CP phases); more in backup slides.

LEP (W^{\pm}, Z) , LHC (H^0) - shaping the Standard Model



STEVEN WEINBERG 1933-2021 A mind to rank with the greatest

Steven Weinberg, one of the greatest theoretical physicists of all time, passed away on 23 July, aged 88. He revolutionised particle physics, quantum field theory and cosmology with conceptual break throughs that still form the foundation of our understanding of physical reality.

Weinberg is well known for the unified theory of weak and electromagnetic forces, which earned him the Nobel Prize in Physics in 1979, jointly awarded with Sheldon Glashow and Abdus Salam, and led to the prediction of the Z and W vector bosons, later discovered at CERN in 1983. His breakthrough was the realisation that some new theoretical ideas, initially believed to play a role in the description of nuclear strong interactions, could instead explain the nature of the weak force. "Then it suddenly occurred to me that this was a perfectly good sort of theory, but I was applying it to the wrong kind of interaction. The right place to apply these ideas was not to the strong interactions, but to the weak and electromagnetic interactions," as he later recalled. With his work. Weinberg had made the next step in the unification of physical laws, after Newton understood that the motion of apples on Earth and planets in the sky are governed by the same gravitational force, and Maxwell understood that electric and magnetic phenomena are the only one model oversection of a single force



Steven Weinberg radically changed the way we look at the universe.

In my life, I have built

physicists, and will certainly continue to serve future generations. Steven Weinberg is among the very few individuals who, during the course of the history

$N_{\rm eff}$: (Good) Things Come in 3s?

The Number of Neutrino Species,

D. Denegri, B. Sadoulet, M. Spiro, Rev.Mod.Phys. 62 (1990) 1



106 Z events

1989:

Initial measurements of Z-boson resonance parameters in e^+e annihilation, SLC Colaboration Phys. Rev. Lett. 63, 724 ALEPH, OPAL, L3, DELPHI, MARKII (SLC): $N_{\nu} = 3.12 \pm 0.19$ CERN, 13.10.1989, Video (~12,000 Z decays) [LEP, 2006] (~17 mln Z decays)

 $N_{\nu} = 2.9840 \pm 0.0082$

Update: [P. Janot and S. Jadach, 2019](only 1σ off from N=3)

 $N_{\nu} = 2.9963 \pm 0.0074$

Theorem: [C. Jarlskog, 1990]

In the Standard Model with n left-handed lepton doublets and N-nright-handed neutrinos, the effective number of neutrinos, N_{ν} , defined by

$$\Gamma(Z \to \nu' s) \equiv N_{\nu} \Gamma_0,$$

where Γ_0 is the standard width for one masseless neutrino, satisfies

$$N_{\nu} \leq n$$
.

Cosmology: $N_{eff} = 3.044$. J. Froustey, C. Pitrou, M. Volpe, JCAP 12 (2020) 015, J. Bennett, G. Buldgen, M. Drewes, Y. Wong, JCAP 03 (2020) 003, JCAP 03 (2021) A01

$$\boldsymbol{\Sigma}_{\boldsymbol{x}}^{(f)} = \underbrace{\left(\boldsymbol{V}_{\text{osc}} \right)_{\alpha i} \boldsymbol{\nu}_{i}^{(m)}}_{\text{SM part}} + \underbrace{\left(\boldsymbol{V}_{lh} \right)_{\alpha j} \boldsymbol{\widetilde{\nu}}_{j}^{(m)}}_{\text{BSM part}}$$

BSM part

7

PMNS data analysis (Nonunitarity), 3+1

New limits on neutrino non-unitary mixings based on prescribed singular values, W. Flieger, JG, K. Porwit, JHEP 03 (2020) 169

- $\begin{array}{l} \bullet \quad (\mathbf{I}): \ m > EW. \\ Ours: |U_{e4}| \in [0, 0.021], \quad |U_{\mu 4}| \in [0.00013, 0.021], \quad |U_{\tau 4}| \in [0.0115, 0.075]. \\ Others: |U_{e4}| \le 0.041, \quad |U_{\mu 4}| \le 0.030, \quad |U_{\tau 4}| \le 0.087 \ [\text{J. de Blas, 2013}] \end{array}$
- (II): $\Delta m^2 \gtrsim 100 eV^2$.

 $Ours: |U_{e4}| \in [0, 0.082], \quad |U_{\mu4}| \in [0.00052, 0.099], \quad |U_{\tau4}| \in [0.0365, 0.44].$

• (III):
$$\Delta m^2 \sim 0.1 - 1 eV^2$$
.

$$\begin{split} Ours: |U_{e4}| &\in [0, 0.130], \quad |U_{\mu4}| \in [0.00052, 0.167], \quad |U_{\tau4}| \in [0.0365, 0.436].\\ Others: |U_{e4}| \in [0.114, 0.167], \quad |U_{\mu4}| \in [0.0911, 0.148], \quad |U_{\tau4}| \leq 0.361.\\ \mbox{[C. Giunti et al., 2017]} & [M. Dantler et al., 2018] \end{split}$$

 \rightarrow In some cases we improved (blue), in some not (red).

BSM and new scales, S. Kanemura, FCC November Week 2020

https://indico.cern.ch/event/923801

Two Possibilities satisfying current data


Effective Theories vs Concrete Models, hep-ph/9909242.

 $\Delta r_{BSM} \stackrel{?}{=} \Delta r_{SM}(m_t, m_H, ...) + "$ subleading effects from h.o.c."

Confronting electroweak precision measurements with New Physics models

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$$(\Delta \rho)_{LRM} \simeq \frac{m_t^2}{\mathbf{M}_{\mathbf{W}_2}^2 - M_{W_1}^2}$$

 $(\Delta \rho)_{SM} \simeq \frac{m_t^2}{M_W^2};$

$$G_F \sim \frac{g^2}{M_W^2}$$

Abstract. Precision experiments, such as those performed at LEP and SLC, offer us an excellent opportunity to constrain extended gauge model parameters. To this end, it is often assumed that in order to obtain more reliable estimates, one should include the sizable one-loop standard model (SM) corrections, which modify the Z^0 couplings as well as other observables. This conviction is based on the belief that the higher order contributions from the "extension sector" will be numerically small. However, the structure of higher order corrections can be quite different when comparing the SM with its extension: thus one should avoid assumptions which do not take account of such facts. This is the case for all models with $\rho_{trow} = M_{W}^2/(M_Z^2 \cos^2 \Theta_W) \neq 1$. As an example, both the manifest left–right symmetric model and the $SU(2)_L \odot U(1)_Y \odot \widetilde U(1)$ model, with an additional Z⁰ boson, are discussed, and special attention to the top contribution to $\Delta \rho$ is given. We conclude that the only sensible way to confront a model with the experimental data is to renormalize it self-consistently. If this is not done, parameters which depend strongly on quantum effects should be left free in fits, though essential physics is lost in this way. We should note that the arguments given here allow us to state that at the level of loop corrections (indirect effects) there is nothing like a "model-independent global analysis" of the data. A theorist's commentary

Why are neutrino masses still Beyond the Standard Model Physics?

We do not know how to write neutrino masses:

• Are ν data described by left-handed Majorana masses?

$$\Delta \mathcal{L} = rac{1}{2} m_L \overline{\nu_L^c} \nu_L$$
 (maybe!)

• Are ν data described by right-handed Majorana masses?

 $\Delta \mathcal{L} = rac{1}{2} m_R \overline{
u_R^c}
u_R$ (not by itself!)

• Are ν data described by Dirac masses?

$$\Delta \mathcal{L} = m_D \overline{\nu_L} \nu_R + ext{H.c. (maybe, but I hope not!)}$$



$Z \to l_1^{\pm} l_2^{\mp} \colon Z \to e\mu, \mu\tau, e\tau$



Fig: Charged lepton flavor violation from massive neutrinos in Z decays J. Illana, T. Riemann, Phys.Rev.D 63 (2001) 053004, hep-ph/0010193 ν_i, ν_j - light, heavy, Dirac, Majorana, in general. In addition: EWPOs and oblique corrections (i.e. S,T,U).

Neutrinos, Corrections

the list is not complete ...

Leptonic flavor changing Z0 decays in $SU(2) \times U(1)$ theories with right-handed neutrinos,

J. G. Korner, A. Pilaftsis and K. Schilcher, Phys. Lett. B 300 (1993), 381, hep-ph/9301290

Mixing renormalization in Majorana neutrino theories, B.A. Kniehl, A. Pilaftsis, Nucl.Phys.B 474 (1996) 286, hep-ph/9601390

Effects of heavy Majorana neutrinos on lepton flavor violating processes, G. Hernández-Tomé et al, Phys.Rev.D 101 (2020) 7, 075020 1912.13327

Improving Electro-Weak Fits with TeV-scale Sterile Neutrinos, E. Akhmedov et al, JHEP 05 (2013) 081, 1302.1872

Loop level constraints on Seesaw neutrino mixing, E. Fernandez-Martinez et al, JHEP 10 (2015) 130, 1508.03051

For RHNs we can compare SM HO terms with BSM effects: $HO = HO_{\rm SM} + HO_{BSM}$ (work in progress, Dirac neutrinos).

Z-pole physics was essential in the past and will remain the butter and bread of precision particle studies if future e^+e^- colliders come to reality.

- 1. Challenges at Z-pole:
 - 1.1 3-loop EW and mixed EW-QCD, leading 4-loop corrections for $Z \rightarrow 2f$ vertices
 - 1.2 QED interference effects, non-factorizable corrections
 - 1.3 Adjusting MC generators at NNLO and beyond (Bhabha (!), exclusive NNLO $e^+e^- \to f\bar{f}$).
- 2. Challenge to improve input parameters ($\alpha, \alpha_{\rm s}$, physics at ZH, WW, tt)
- 3. Challenge to optimize/understand paths towards BSM discovery (RHNs, DM, CP effects,...)
- 4. Challenge: SM(BSM)EFT, precision physics for concrete BSM models
- 5. Challenge: Tools (MC generators, multiloop numerical, analytical programs)

 * 'FCC-ee: the challenge for theory', talk at 4th FCC Physics and Experiments Workshop, link

BACKUP

Precision changed history: justice, law, crime, trade, economy, social, ...

Example: To be just was precisely to use balance.



Wisdom 11:20

'By weight, measure and number, God made all things'

 $\underline{Code \ of \ Hammurabi}$ 1772 BC - any taverner using false weights could be served up with the death penalty

Input and calculated/measured parameters



Fig. from the FCC-ee report ' α_{QED} ' by F. Jegerlehner in 1905.05078

Introduction to Precision Electroweak Analysis by J. Welss, 0512342



Science 376 (2022) 6589, 170-176

$$\begin{array}{rcl} {\rm SM}: M_W &=& 80.357 \pm 6 \ {\rm MeV}, \ ({\rm PDG2020}) \\ {\rm Global}: M_W &=& 80.379 \pm 12 \ {\rm MeV}, \ ({\rm PDG2020}) \\ {\rm CDFII}: M_W &=& 80433.5 \pm 9.4 \ {\rm MeV} \end{array}$$

FCC-ee forecast : $M_W = X \pm 0.4 \text{ MeV}!$

Input and calculated/measured parameters

Experimental values:

$$\begin{split} \hat{\alpha} &= 1/137.0359895(61), \ \gamma^* \to e^+ e \\ \hat{G}_F &= 1.16639(1) \times 10^{-5} \,\text{GeV}^{-2} \text{ muon decay} \\ \hat{m}_Z &= 91.1875 \pm 0.0021 \,\text{GeV} \\ \hat{m}_W &= 80.426 \pm 0.034 \,\text{GeV} \\ \hat{s}_{\text{eff}}^2 &= 0.23150 \pm 0.00016, \text{effective} \sin^2 \theta_{\text{W}}, A_{LR} \equiv \frac{(1/2 - \hat{s}_{\text{eff}}^2)^2 - \hat{s}_{\text{eff}}^4}{(1/2 - \hat{s}_{\text{eff}}^2)^2 + \hat{s}_{\text{eff}}^4} \\ \hat{\gamma}_{l+l-} &= 83.984 \pm 0.086 \,\text{MeV} \\ \mathbf{g}(= e/s_W) \, SU(2) \\ \mathbf{g}'(e/c_W) \, U(1)_Y \longrightarrow \\ \mathbf{v} \, \text{VEV}, \\ \mathbf{v} \, \text{VEV}, \end{split} \qquad \begin{cases} \hat{\alpha} = \frac{e^2}{4\pi} \\ \hat{G}_F = \frac{1}{\sqrt{2v^2}} \\ \hat{m}_Z^2 = \frac{e^2v^2}{4s^2c^2} \\ \hat{m}_W^2 = \frac{e^2v^2}{4s^2} \\ \hat{s}_{\text{eff}}^2 = s^2 \\ \hat{\Gamma}_{l+l-} = \frac{v}{96\pi} \frac{e^3}{s^3c^3} \left[(-\frac{1}{2} + 2s^2)^2 + \frac{1}{4} \right] \end{split}$$

Janusz Gluza

Shaping the SM, tree level estimates

In terms of $\hat{\alpha}, \hat{G}_F$ and \hat{m}_Z

$$\hat{m}_W^2 = \pi \sqrt{2} \hat{G}_F^{-1} \hat{\alpha} \left(1 - \sqrt{1 - \frac{4\pi \hat{\alpha}}{\sqrt{2} \hat{G}_F \hat{m}_Z^2}} \right)^{-1}$$

$$\begin{split} \hat{s}_{\text{eff}}^2 \hat{c}_{\text{eff}}^2 &= \frac{\pi \hat{\alpha}}{\sqrt{2} \hat{G}_F \hat{m}_Z^2} \quad \equiv \quad \hat{s}_{\text{eff}}^2 = \frac{1}{2} - \frac{1}{2} \sqrt{1 - \frac{4\pi \hat{\alpha}}{\sqrt{2} \hat{G}_F \hat{m}_Z^2}} \\ \hat{\Gamma}_{l^+ l^-} &= \quad \frac{\sqrt{2} \hat{G}_F \hat{m}_Z^3}{12\pi} \left\{ \left(\frac{1}{2} - \sqrt{1 - \frac{4\pi \hat{\alpha}}{\sqrt{2} \hat{G}_F \hat{m}_Z^2}} \right)^2 + \frac{1}{4} \right\} \end{split}$$

 $\begin{array}{lll} Prediction: \hat{m}_W &=& 80.939 \pm 0.003 \, {\rm GeV} \, 15\sigma \, {\rm away} \\ Prediction: \hat{s}_{\rm eff}^2 &=& 0.21215 \pm 0.00003 \, 120\sigma \, {\rm away} \\ Prediction: \hat{\Gamma}_{l+l^-} &=& 84.843 \pm 0.012 \, {\rm MeV} \, 10\sigma \, {\rm away} \end{array}$

Shaping SM, oblique corrections also not sufficient



$$\tau_{\mu}^{-1} = \frac{\hat{G}_F^2 m_{\mu}^5}{192\pi^3} K(\alpha, m_e, m_{\mu}, m_W)$$

$$\begin{array}{ll} \frac{(\hat{G}_F)^{\rm th}}{\sqrt{2}} & = & \frac{g^2}{8m_W^2} \left[1 + i\Pi_{WW}(q^2) \left(\frac{-i}{q^2 - m_W^2} \right) \right]_{q \to 0} \\ & = & \frac{1}{2v^2} \left[1 - \frac{\Pi_{WW}(0)}{m_W^2} \right]. \end{array}$$

Janusz Gluza

Primary role of SM radiative corrections, F. Jegerlehner, in 1905.05078

$$\sin^2 \Theta_i \, \cos^2 \Theta_i = \frac{\pi \, \alpha}{\sqrt{2} \, G_\mu \, M_Z^2} \, \frac{1}{1 - \Delta r_i} \quad \Delta r_i = \Delta r_i(\alpha, G_\mu, M_Z, m_H, m_{f \neq t}, m_t) \,,$$

$$\Delta r_i = -\frac{c_W^2}{s_W^2} \,\Delta\rho + \Delta r_{i \text{ reminder}} \,,$$
$$\Delta \rho = \frac{3 \,m_t^2 \,\sqrt{2} G_\mu}{16 \,\pi^2}$$

 $\hat{\alpha}(m_Z) = \frac{\hat{\alpha}}{1 - \Delta \alpha(m_Z)} = \frac{e^2}{4\pi} \left[1 + \frac{\Pi_{\gamma\gamma}(m_Z)}{m_Z^2} \right] \sim 128 \text{ (137 at the Thomson limit)}$

Still, well visible disagreement between SM prediction and experiment for EWPOs without subleading SM corrections, and only with the leading corrections $\Delta \alpha(m_Z)$ and $\Delta \rho$.

 $r_{i \text{ reminder}}$ matters! (see also backup slides)

F. Jegerlehner, in 1905.05078

Example: the W and Z mass from
$$\alpha(M_Z)$$
, G_{μ} and $\sin^2 \Theta_{\ell \,\text{eff}}$:
(i) $\sin^2 \Theta_W = 1 - M_W^2/M_Z^2$,
 $\sin^2 \theta_{\ell,\text{eff}}(M_Z) = \left(1 + \frac{\cos^2 \Theta_W}{\sin^2 \Theta_W} \Delta \rho\right) \sin^2 \Theta_W$,
 $\Delta \rho = \frac{3M_t^2 \sqrt{2}G_{\mu}}{16\pi^2}$; $M_t = 173 \pm 0.4 \, GeV$

The iterative solution with input $\sin^2 \theta_{\ell,\text{eff}}(M_Z) \equiv (1 - v_\ell/a_\ell)/4 = 0.23148$ (EXP!) is $\sin^2 \Theta_W = 0.22426$. (ii) $M_W^{\text{exp}} = 80.379 \pm 0.012$; $M_Z^{\text{exp}} = 91.1876 \pm 0.0021 \text{ GeV}$, $\longrightarrow 1 - M_W^2/M_Z^2 = 0.22263$. Predicting then the masses we have

$$M_W = \frac{A_0}{\sin^2 \Theta_W} ; \ A_0 = \sqrt{\frac{\pi \alpha}{\sqrt{2}G_\mu}} ; \ M_Z = \frac{M_W}{\cos \Theta_W}$$

where, including photon VP correction $\alpha^{-1}(M_Z) = 128.953 \pm 0.016$. For the W, Z mass we then get

 $M_W^{\rm the} = 81.1636 \pm 0.0346$; $M_Z^{\rm the} = 92.1484 \pm 0.0264$.

Deviations (errors added in quadrature): $W: 23 \sigma; Z: 36 \sigma$

FCC-ee fits, CDR link



Observable	present	FCC-ee	FCC-ee	Comment and	
	value \pm error	Stat.	Syst.	leading exp. error	
m _Z (keV)	91186700 ± 2200	4	100	From Z line shape scan	
F (h-W)	9405900 - 9900	4	95	Beam energy calibration	
$I_Z (\text{kev})$	2495200 ± 2300	4	25	From Z line snape scan Boom operate collibration	
-:-20eff(106)	021400 100	0	0.4	Beam energy cambration	
$\sin \theta_{W}(\times 10^{\circ})$	231480 ± 100	2	2.4	From AFB at Z peak	
1 (100050 14	0		beam energy calibration	
$1/\alpha_{QED}(m_Z)(\times 10)$	128952 ± 14	3	small	from A _{FB} on peak	
DZ (10 ³)			0.0.1	QED&E w errors dominate	
$R_{\ell}^{-}(\times 10^{-})$	20767 ± 25	0.06	0.2-1	ratio of hadrons to leptons	
				acceptance for leptons	
$\alpha_s(m_{\tilde{Z}})$ (×10 [*])	1196 ± 30	0.1	0.4 - 1.6	from R _ℓ [~] above	
σ_{had}^0 (×10 ³) (nb)	41541 ± 37	0.1	4	peak hadronic cross section	
				luminosity measurement	
$N_{\nu}(\times 10^{3})$	2996 ± 7	0.005	1	Z peak cross sections	
				Luminosity measurement	
$R_b (\times 10^6)$	216290 ± 660	0.3	< 60	ratio of bb to hadrons	
				stat. extrapol. from SLD	
$A_{FB}^{b}, 0 (\times 10^{4})$	992 ± 16	0.02	1-3	b-quark asymmetry at Z pole	
				from jet charge	
$A_{FB}^{pol,\tau}$ (×10 ⁴)	1498 ± 49	0.15	<2	τ polarization asymmetry	
				τ decay physics	
τ lifetime (fs)	290.3 ± 0.5	0.001	0.04	radial alignment	
τ mass (MeV)	1776.86 ± 0.12	0.004	0.04	momentum scale	
τ leptonic ($\mu \nu_{\mu} \nu_{\tau}$) B.R. (%)	17.38 ± 0.04	0.0001	0.003	e/μ/hadron separation	
m _W (MeV)	80350 ± 15	0.25	0.3	From WW threshold scan	
				Beam energy calibration	
Γ _W (MeV)	2085 ± 42	1.2	0.3	From WW threshold scan	
				Beam energy calibration	
$\alpha_{s}(m_{W}^{2})(\times 10^{4})$	1170 ± 420	3	small	from R_{ℓ}^{W}	
$N_{\nu}(\times 10^{3})$	2920 ± 50	0.8	small	ratio of invis. to leptonic	
				in radiative Z returns	
$m_{top} (MeV/c^2)$	172740 ± 500	17	small	From tt threshold scan	
sop c / /				QCD errors dominate	
$\Gamma_{top} (MeV/c^2)$	1410 ± 190	45	small	From tt threshold scan	
				QCD errors dominate	
SM	10100	0.10	emall	From tt threshold scan	
$\lambda_{top} / \lambda_{top}^{SN}$	1.2 ± 0.3	0.10	SHIGH		
$\lambda_{top}/\lambda_{top}^{SM}$	1.2 ± 0.3	0.10	Sinan	OCD errors dominate	

E.g. effective weak mixing angle

The weak mixing angle $s_W^2 \equiv \sin^2 \theta_W$ has three potential different meanings or functions in the model-building:

(i) It describes the ratio of the two gauge couplings,

$$g'/g = c_W/s_W,$$

usually in the $\overline{\text{MS}}$ scheme.

(ii) It describes the ratio of two gauge boson (on-shell) masses,

$$s_W^2 = 1 - \frac{M_W^2}{M_Z^2}.$$

(iii) It describes the ratio of the vector and axial-vector couplings of an (on-shell) Z boson to fermions,

$$\frac{v_f}{a_f} = 1 - 4|Q_f|s_W^2.$$

This definition is called the effective weak mixing angle, denoted as $\sin^2\theta_W^{f,{\rm eff}}.$

How to unfold - rough scheme

We have to describe

$$e^+e^- \longrightarrow (\gamma, Z) \longrightarrow f^+f^-(\gamma),$$

S-matrix Ansatz in the complex energy plane

$$\mathcal{A}^{e^+e^- \to b\bar{b}} = \underbrace{\frac{R_Z}{s - s_Z}}_{\gamma - Z \text{ interference}} + \underbrace{\frac{R_{\gamma}}{s} + S + (s - s_Z)S' + \dots}_{\gamma - Z \text{ interference}},$$

$$s_Z = \overline{M}_Z^2 - i\overline{M}_Z\overline{\Gamma}_Z$$

R, S, S', ... are individually gauge-invariant and UV-finite - unitarity and analyticity of the S-matrix. IR-finite, when soft and collinear real photon emission is added. [Willenbrock, Valencia,1991] [Sirlin,1991] [Stuart,1991] [Riemann, 1991, 1992] [H. Veltman,1994] [Passera, Sirlin, 1998] [Gambino, Grassi, 2000] [Avramik, Czakon, Freitas, 2006].

The term $R_{\gamma}(s)/s$ is part of the background

• The poles of \mathcal{A} have complex residua R_Z and R_γ .

There is only ONE pole in mathematics, while in physics we observe two of them: photon exchange at s = 0, Z exchange at s₀ = s_Z. Mathematicaly, the appearance of the photon pole is result of summing of part of background around Z pole, s₀ = s_Z

[T. Riemann, APPB 2015]

$$\frac{R_{\gamma}(s)}{s} = \frac{\sum_{n=0}^{\infty} R_n (s-s_0)^n}{s} \\
= \frac{\sum_{n=0}^{\infty} R_n (s-s_0)^n}{s_0 - (s_0 - s)} \\
= \sum_{n=0}^{\infty} R_n (s-s_0)^n \frac{1}{s_0} \frac{1}{1 - \frac{s_0 - s}{s_0}} \\
= \sum_{n=0}^{\infty} R_n (s-s_0)^n \frac{1}{s_0} \left[1 + \frac{s_0 - s}{s_0} + \left(\frac{s_0 - s}{s_0} \right)^2 \cdots \right];$$

Beyond Born level, one can write $\begin{aligned} \mathcal{M}_{\gamma}^{(0)}(e^-e^+ \to f^-f^+) &= \frac{4\pi i \alpha_{em}(s)}{s} Q_e Q_f \gamma_{\alpha} \otimes \gamma^{\alpha}, \\ \mathcal{M}_{Z}^{(0)}(e^-e^+ \to f^-f^+) &= 4ie^2 \frac{\chi_Z(s)}{s} \big[M_{vv}^{ef} \gamma_{\alpha} \otimes \gamma^{\alpha} - M_{av}^{ef} \gamma_{\alpha} \gamma_5 \otimes \gamma^{\alpha} \\ &- M_{va}^{ef} \gamma_{\alpha} \times \gamma^{\alpha} \gamma_5 + M_{aa}^{ef} \gamma_{\alpha} \gamma_5 \otimes \gamma^{\alpha} \gamma_5 \big]. \end{aligned}$

In the pole scheme, where \bar{M}_Z is defined as the real part of the pole of the S matrix, one has

$$\chi_Z(s) = \frac{G_F M_Z^2}{\sqrt{2} 8\pi \alpha_{em}} K_Z(s) \simeq \frac{1}{1 + i\frac{\bar{\Gamma}_Z}{M_Z}} \frac{s}{s - \bar{M}_Z^2 + i\bar{M}_Z\bar{\Gamma}_Z} \simeq \frac{s}{s - M_Z^2 + iM_Z\Gamma_Z(s)},$$
$$\Gamma_Z(s) = \frac{s}{M_Z^2} \Gamma_Z$$

Definitions are related:

$$\begin{split} \bar{M}_Z &\approx M_Z - \frac{1}{2} \ \frac{\Gamma_Z^2}{M_Z} \ \approx \ M_Z - 34 \ \text{MeV}, \\ \bar{\Gamma}_Z &\approx \Gamma_Z - \frac{1}{2} \ \frac{\Gamma_Z^3}{M_Z^2} \ \approx \ \Gamma_Z - 0.9 \ \text{MeV}. \end{split}$$

- Known from LEP. One of examples why changing frameworks/assumptions/simplifications of calculations matter (!).
- ► However, at FCC-ee $\delta \Gamma_Z \sim 0.1$ MeV. Non-facotrization effects must be added properly beyond 1-loop.
- Is it necessary for FCC-ee accuracy to implement MC with radiative corrections calculated at the amplitudes level?
- At this precision it is important which parameters are taken as input parameters in schemes.

EW SM theory at loops, an example ($\Delta_{ef} \neq 0$)

$$\begin{cases} \Gamma_{Z}, \Gamma_{partial} \\ A_{FB,peak}^{eff.,Born}, A_{LR,peak}^{eff.,Born} \\ R_{b}, R_{\ell}, \dots \end{cases} \longrightarrow \begin{cases} v_{\ell,\nu,u,d,b}^{eff} \\ a_{\ell,\nu,u,d,b}^{eff} \\ \sin^{2}\theta_{\text{eff}}^{b}, \sin^{2}\theta_{\text{eff}}^{lept} \end{cases}$$

e.g. : improvements needed for subtle corrections $\Delta_{1,2}$ (e.g. boxes, **2L-boxes**)

$$\begin{split} A_{FB,peak}^{eff.,Born} &= \frac{2\Re e\left[\frac{v_e a_e^*}{|a_e|^2}\right] 2\Re e\left[\frac{v_f a_f^*}{|a_f|^2}\right]}{\left(1 + \frac{|v_e|^2}{|a_e|^2}\right) \left(1 + \frac{|v_f|^2}{|a_f|^2}\right)} + \Delta_1 - \Delta_2 \simeq \frac{3}{4} A_e A_f, \\ \Delta_1 &= 2\Re e\left[\Delta_{ef}\right], \ \Delta_2 = |\Delta_{ef}|^2 + 2\Re e\left[\frac{v_e a_e^*}{|a_e|^2} \frac{v_f a_f^*}{|a_f|^2} \Delta_{ef}^*\right], \\ \Delta_{ef} &= 16|Q_e Q_f| s_W^4(\kappa_{ef} - \kappa_e \kappa_f) \end{split}$$

Scheme of construction and the use of EWPO/EWPP at FCC-ee



 $Born \otimes QED$ still valid? Scheme DATA

⊗ QED

 $\mathcal{O}(\alpha^1)^{\mathrm{noQED}}$

General remarks on usefulness of EWPOs

- 1. EWPOs encapsulate experimental data after extraction of well known and controllable QED and QCD effects, in a model-independent manner.
- 2. They provide a convenient bridge between real data and the predictions of the SM (or SM plus New Physics).
- 3. Contrary to raw experimental data (like differential crosssections), EWPOs are well suited for archiving and long term exploitation.
- 4. In particular archived EWPOscan be exploited over long periods of time for comparisons with steadily improving theoretical calculations of the SM predictions, and for validations of the New Physics models beyond the SM.
- 5. They are also useful for comparison and combination of results from different experiments.

MB and SD methods are very much complementary!

 MB works well for hard threshold, on-shell cases, not many internal masses (more IR);
 SD more useful for integrals with many internal masses

 10^{-8} accuracy achieved for any self-energy and vertex Feynman integral with one of the methods - in Minkowskian region.



2-loops \longrightarrow 3-loops



$$m_1 = M_t, m_2 = M_W$$

The integrals contain up to three dimensionless parameters

$$\left\{\frac{M_H^2}{M_Z^2}, \frac{M_W^2}{M_Z^2}, \frac{m_t^2}{M_Z^2}, \frac{(M_Z + i\varepsilon)^2}{M_Z^2}\right\}$$

Neutrino parameters, development



	Normal Ordering (best fit)		Inverted Ordering ($\Delta \chi^2 = 2.6$)		
	bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range	
$\sin^2 \theta_{12}$	$0.304^{+0.013}_{-0.012}$	$0.269 \rightarrow 0.343$	$0.304^{+0.012}_{-0.012}$	$0.269 \rightarrow 0.343$	
$\theta_{12}/^{\circ}$	$33.44_{-0.74}^{+0.77}$	$31.27 \rightarrow 35.86$	$33.45_{-0.74}^{+0.77}$	$31.27 \rightarrow 35.87$	
$\sin^2 \theta_{23}$	$0.573^{+0.018}_{-0.023}$	$0.405 \rightarrow 0.620$	$0.578^{+0.017}_{-0.021}$	$0.410 \rightarrow 0.623$	
$\theta_{23}/^{\circ}$	$49.2^{+1.0}_{-1.3}$	$39.5 \rightarrow 52.0$	$49.5^{+1.0}_{-1.2}$	$39.8 \rightarrow 52.1$	
$\sin^2 \theta_{13}$	$0.02220\substack{+0.00068\\-0.00062}$	$0.02034 \to 0.02430$	$0.02238\substack{+0.00064\\-0.00062}$	$0.02053 \to 0.02434$	
$\left(\theta_{13} \right)^{\circ}$	$8.57^{+0.13}_{-0.12}$	$8.20 \rightarrow 8.97$	$8.60^{+0.12}_{-0.12}$	$8.24 \rightarrow 8.98$	
$\delta_{CP}/^{\circ}$	194^{+52}_{-25}	$105 \to 405$	287^{+27}_{-32}	$192 \to 361$	
$\frac{\Delta m^2_{21}}{10^{-5}~{\rm eV}^2}$	$7.42^{+0.21}_{-0.20}$	$6.82 \rightarrow 8.04$	$7.42^{+0.21}_{-0.20}$	$6.82 \rightarrow 8.04$	
$\frac{\Delta m^2_{3\ell}}{10^{-3}~{\rm eV}^2}$	$+2.515^{+0.028}_{-0.028}$	$+2.431 \rightarrow +2.599$	$-2.498\substack{+0.028\\-0.029}$	$-2.584 \rightarrow -2.413$	

Super-K, Hyper-K, T2K, NOvA, Antares, KM3NeT, Juno, Dune, SNO+, Daya Bay, Double Chooz, RENO, ...



Conclusion: Neutrino Physics stepped in the precision era. Till 2030: mass hierarchy, δ_{CP} (maybe), absolute masses, Majorana-Dirac, L. Wen, EPS2021.

Neutrinos



Dirac neutrinos: Cosmology,...

- Neutrino oscillation experiments are insensitive to the nature of neutrinos.
- Experiments (such as 0νββ decay) looking for lepton number violating signatures can probe the Majorana nature of neutrinos.
- Plethora of $0\nu\beta\beta$ Experiments :

Past - Heidelberg-Moscow, GERDA-I, NEMO-3; Present- EXO-200, KamLAND-Zen; Future- GERDA-II, NEXT, nEXO, PandaX-III, SpuerNEMO...

- So far **NOT** so good : absence of any positive result.
- Caution : Null result in 0νββ expts. do not necessarily prove neutrinos are Dirac particle.
- Nevertheless, its worth to come up with ideas with implication of Dirac neutrinos in Cosmology too

Dirac neutrinos: Cosmology,...

• Dirac Leptogenesis:

[H. Murayama et al., 2002] [T. Dick et al., 1999]

 \rightarrow Traditional Leptogenesis: small neutrio mass and LNV required to produce lepton asymmetry is provided by heavy right-handed Majorana neutrinos.

 \rightarrow Strategy: (a) additional symmetry introduced to forbid Dirac Yukawa coupling and Majorana mass term, (b) A set of heavy, vector-like pairs of fields introduced whose couplings to the standard model fields contain nontrivial, CP-violating phases which plays the role of heavy right-handed Majorana neutrinos.

• Dirac Neutrinos and freeze-in dark matter: [E. Ma, 2021]

[D. Borah et al., 2018]

 \rightarrow Neutrinos to have only Dirac mass and dark matter relic abdunce generated through freeze-in mechanism, both requires coupling $\mathcal{O}(10^{-11}).$ Such a small coupling may have some unified origin.

BSM and RHNs, FCC-ee CDR vol.1

LFV Z-decays: $(10^{-6} \div 10^{-5})$. FCC-ee $\longrightarrow \sim 10^{-9}$ branching fractions. A. Blondel et al. 1411.5230 ESPPU Briefieng Book 1910.11775



Low-scale leptogenesis with flavour and CP symmetries, M. Drewes et al, 2203.08538 Discrete Flavor Symmetries and Lepton Masses and Mixings, G. Chauhan, et al, 2203.08538 (Snowmass contribution)

Resonant Leptogenesis, Collider Signals and Neutrinoless Double Beta Decay from Flavor and CP Symmetries, G. Chauhan, B. Dev, 2203.08538

LNV, Majorana neutrinos, $H^{\pm\pm}$, ...

Proces	Obecne ograniczenie	Oczekiwany limit	Eksperyment
$\mu^+ \rightarrow e^+ \gamma$	$< 4.2 \times 10^{-13}$	5×10^{-14}	MEG II
$\mu^+ \rightarrow e^+ e^- e^+$	$< 1.0 \times 10^{-12}$	10^{-16}	Mu3e
$\mu^- Al \rightarrow e^- Al$	$< 6.1 \times 10^{-13}$	10^{-17}	Mu2e, COMET
$\mu^- \mathrm{Si/C} \rightarrow e^- \mathrm{Si/C}$	-	5×10^{-14}	DeeMe
$\tau \rightarrow e\gamma$	$< 3.3 \times 10^{-8}$	5×10^{-9}	Be lle II, FC
$\tau \rightarrow \mu \gamma$	$< 4.4 \times 10^{-8}$	10^{-9}	Belle II , FC
$\tau \rightarrow eee$	$< 2.7 \times 10^{-8}$	5×10^{-10}	Belle I I, FC
$\tau \rightarrow \mu \mu \mu$	$< 2.1 \times 10^{-8}$	5×10^{-10}	Belle II, FC
τ \rightarrow e had	$< 1.8 \times 10^{-8}$	3×10^{-10}	Belle II, FC
had $\rightarrow \mu e$	$< 4.7 \times 10^{-12}$	10^{-12}	NA 62
$h \rightarrow e \mu$	$< 3.5 \times 10^{-4}$	3×10^{-5}	HL-LHC , FC
$h \rightarrow \tau \mu$	$< 2.5 \times 10^{-3}$	3×10^{-4}	HL-LH C, FC
$h \rightarrow \tau e$	$< 6.1 \times 10^{-3}$	3×10^{-4}	HL-LH C, FC

Matter-antimatter asymmetry! $\eta = \frac{n_B - n_{\bar{B}}}{n_m a} \simeq \frac{n_B}{n_{\gamma}} \simeq 10^{-10}$



$Z \to l_1^\pm l_2^\mp \colon Z \to e\mu, \mu\tau, e\tau$

$$\begin{split} \Gamma(Z \to \bar{\ell}\ell') &= \frac{\alpha}{3} M_Z |F_L^Z(M_Z^2)|^2 \\ F_L^Z(q^2) &= \frac{\alpha_W}{8\pi s_W c_W} \sum_{i,j}^5 \mathbf{B}^*_{\ell \mathbf{i}} \mathbf{B}_{\ell' \mathbf{j}} \big[\delta_{ij} F(x_i;q^2) \\ &+ C^*_{ij} G(x_i, x_j;q^2) + C_{ij} \sqrt{x_i x_j} H(x_i, x_j;q^2) \big], \end{split}$$

F,G,H - combination of the Passarino-Veltman functions, G. Hernández-Tomé et al, Phys.Rev.D 101 (2020) 7, 075020 1912.13327

$$\mathcal{L}_{W^{\pm}} = -\frac{g}{\sqrt{2}} W_{\mu}^{-} \sum_{i=1}^{3} \sum_{j=1}^{5} B_{ij} \,\bar{\ell}_{i} \gamma^{\mu} P_{L} \chi_{j} + \text{h.c.},$$
$$\mathcal{L}_{Z} = -\frac{g}{4c_{W}} Z_{\mu} \sum_{i,j=1}^{5} \bar{\chi}_{i} \gamma^{u} \left(C_{ij} P_{L} - C_{ij}^{*} P_{R} \right) \chi_{j}$$

$Z \to l_1^\pm l_2^\mp \colon Z \to e\mu, \mu\tau, e\tau$

Important for renormalization

$$\sum_{k=1}^{5} B_{ik} B_{jk}^* = \delta_{ij}, \quad \sum_{k=1}^{3} B_{ki}^* B_{kj} = \sum_{k=1}^{5} C_{ik} C_{jk}^* = C_{ij},$$
$$\sum_{k=1}^{5} m_{\chi_k} C_{ik} C_{jk} = \sum_{k=1}^{5} m_{\chi_k} B_{ik} C_{kj}^* = \sum_{k=1}^{5} m_{\chi_k} B_{ik} B_{jk} = 0.$$

Simple model. 3 active neutrino masses irrelevant: $BR(Z \rightarrow ll') \sim 10^{-55}$

$$\mathcal{M} = \begin{pmatrix} 0 & 0 & 0 & 0 & m_1 \\ 0 & 0 & 0 & 0 & m_2 \\ 0 & 0 & 0 & 0 & m_3 \\ 0 & 0 & 0 & 0 & M \\ m_1 & m_2 & m_3 & M & \mu \end{pmatrix}, \quad \begin{array}{c} m_{\chi_{1,2,3}} = 0, \\ m_{\chi_{4,5}} = \frac{1}{2} \left(\sqrt{4M'^2 + \mu^2} \mp \mu \right), \\ M'^2 = m_1^2 + m_2^2 + m_3^2 + M^2, \end{array}$$

 $\mu=0\longrightarrow$ Dirac neutrino.

$$\begin{pmatrix} -\frac{m_2}{\sqrt{m_1^2 + m_2^2}} & -\frac{m_1 m_3}{m \sqrt{m_1^2 + m_2^2}} & -\frac{m_1 M}{m M'} & -i \frac{m_1 m_{\chi_5}}{M' \sqrt{m_{\chi_5}^2 + M'^2}} & \frac{m_1}{\sqrt{m_{\chi_5}^2 + M'^2}} \\ \frac{m_1}{\sqrt{m_1^2 + m_2^2}} & -\frac{m_2 m_3}{m \sqrt{m_1^2 + m_2^2}} & -i \frac{m_2 M}{m M'} & -i \frac{m_3 m_{\chi_5}}{M' \sqrt{m_{\chi_5}^2 + M'^2}} & \frac{m_3}{m_{\chi_5}^2 + M'^2} \\ 0 & \frac{\sqrt{m_1^2 + m_2^2}}{m} & -\frac{m_3 M}{m M'} & -i \frac{m_3 m_{\chi_5}}{M' \sqrt{m_{\chi_5}^2 + M'^2}} & \frac{m_3}{\sqrt{m_{\chi_5}^2 + M'^2}} \\ 0 & 0 & \frac{m_1}{m M'} & -i \frac{m_3 m_{\chi_5}}{M' \sqrt{m_{\chi_5}^2 + M'^2}} & \frac{m_3}{\sqrt{m_{\chi_5}^2 + M'^2}} \\ 0 & 0 & 0 & \frac{m_1}{M'} & -i \frac{m_1 m_3}{M' \sqrt{m_{\chi_5}^2 + M'^2}} & \frac{M'}{\sqrt{m_{\chi_5}^2 + M'^2}} \\ m_{\chi_4} = m_{\chi_4} = M' \\ \begin{pmatrix} m_2 & m_1 m_3 & m_1 M & i m_1 & 1 m_1 \end{pmatrix} \end{pmatrix}$$

$$\begin{pmatrix} -\frac{1}{\sqrt{m_1^2 + m_2^2}} & -\frac{m_2}{m\sqrt{m_1^2 + m_2^2}} & -\frac{1}{mM'} & -\frac{1}{\sqrt{2}} \frac{M'}{M'} & \frac{1}{\sqrt{2}} \frac{M'}{M'} \\ \frac{m_1}{\sqrt{m_1^2 + m_2^2}} & -\frac{m_2m_3}{m\sqrt{m_1^2 + m_2^2}} & -\frac{m_2M}{mM'} & -\frac{i}{\sqrt{2}} \frac{m_2}{M'} & \frac{1}{\sqrt{2}} \frac{m_2}{M'} \\ 0 & \frac{\sqrt{m_1^2 + m_2^2}}{m} & -\frac{m_3M}{mM'} & -\frac{i}{\sqrt{2}} \frac{M_3}{M'} & \frac{1}{\sqrt{2}} \frac{m_3}{M'} \\ 0 & 0 & \frac{m_1}{M'} & -\frac{i}{\sqrt{2}} \frac{M_1}{M'} & \frac{1}{\sqrt{2}} \frac{M_1}{M'} \\ 0 & 0 & 0 & \frac{m_1}{M'} & -\frac{i}{\sqrt{2}} \frac{M_1}{M'} & \frac{1}{\sqrt{2}} \frac{M_1}{M'} \\ 0 & 0 & 0 & \frac{i}{\sqrt{2}} & -\frac{i}{\sqrt{2}} \end{pmatrix}$$
The simplest 3+1 model (toy model)

$$\mathcal{L}_{\nu}^{M} = -\frac{1}{2}\bar{N} \begin{pmatrix} 0 & 0 & 0 & a_{1} \\ 0 & 0 & 0 & a_{2} \\ 0 & 0 & 0 & a_{3} \\ a_{1} & a_{2} & a_{3} & M \end{pmatrix} + H.c.$$

Four physical neutrino $\nu_1, \nu_2, \nu_3, \nu_4$, with ν_1, ν_2 massless and

 $m_3 < m_4$

 $s_{\alpha} \sim \sqrt{m_3/(m_3+m_4)}$

This mass matrix is diagonalized by the following mixing matrix

$$\begin{pmatrix} c_{\beta} & -s_{\beta}s_{\gamma} & -s_{\beta}c_{\gamma} & 0\\ 0 & c_{\gamma} & -s_{\gamma} & 0\\ c_{\alpha}s_{\beta} & c_{\alpha}c_{\beta}s_{\gamma} & c_{\alpha}c_{\beta}c_{\gamma} & -s_{\alpha}\\ s_{\alpha}s_{\beta} & s_{\alpha}c_{\beta}s_{\gamma} & s_{\alpha}c_{\beta}c_{\gamma} & c_{\alpha} \end{pmatrix}$$

[C. Jarlskog, 1990] [C.O. Escobar et al., 1993]

Isolating non-standard contribution, we can write this as

$$N_{\nu} - 2 = \frac{1}{(x+y)^2} \left[x^2 F(y) + y^2 F(x) + 2xy G(x,y) \right],$$

where

$$\begin{split} F(z) &= (1 - 4z^2)^{\frac{3}{2}}, \\ G(x,y) &= \sqrt{1 + (x^2 + y^2)^2 - 2(x^2 - y^2)} \left[1 - \frac{x^2 + y^2}{2} - 3xy - \frac{(x^2 - y^2)^2}{4} \right], \end{split}$$

and

$$x = m_4/M_Z, \ y = m_3/M_Z, \ m_{3,4} \le M_Z/2.$$

Nonstandard Mixing Estimation



We are looking for the line $y = x \tan^2 \alpha$ lying below experimental limits LEP $:N_{\nu} = 2.9840 \pm 0.0082 \rightarrow \sum_{i} |U_{i4}| \equiv \sin \alpha < 0.174$ $i = e, \mu, \tau$ NEW $:N_{\nu} = 2.9963 \pm 0.0074 \rightarrow \sum_{i} |U_{i4}| \equiv \sin \alpha < 0.121$ $i = e, \mu, \tau$

New constraints on heavy neutral leptons coming from oscillation data analysis and precision e^+e physics, W. Flieger, K. Grzanka, PoS ICHEP2020 (2021) 129

SM and SMEFT

"Personal Remarks on SMEFT for Snowmass", M. E. Peskin, EF1-EF4 meeting, Sept. 2020, pdf

1. "... the interpretation depends on the connection to explicit models of BSM physics".

How $\frac{C_i}{\Lambda^2}$, O_H , O_{WW} , ..., are releated to the BSM param eters?

- 2. Linear dependence on SMEFT parameters?
- 3. SMEFT at high Q^2 vs specific models?
- 4. h.o. SMEFT corrections and cancellations with SMEFT tree level
- 5. SMEFT contribution to SM background.

Estimated theoretical uncertainties from missing higher orders and the perturbative orders (QCD/elw.) of the results included in the analysis.

Partial Width	QCD	Electroweak	Total	on-shell Higgs
$H \to b \bar{b}/c \bar{c}$	$\sim 0.2\%$	$\sim 0.5\%$	$\sim 0.5\%$	N^4LO / NLO
$H\to \tau^+\tau^-/\mu^+\mu^-$	—	$\sim 0.5\%$	$\sim 0.5\%$	— / NLO
$H \to gg$	$\sim 3\%$	$\sim 1\%$	$\sim 3\%$	$N^{3}LO / NLO$
$H\to\gamma\gamma$	< 1%	< 1%	$\sim 1\%$	NLO / NLO
$H \to Z \gamma$	< 1%	$\sim 5\%$	$\sim 5\%$	LO / LO
$H \to WW/ZZ \to 4f$	< 0.5%	$\sim 0.5\%$	$\sim 0.5\%$	NLO/NLO

Projected intrinsic and parametric uncertainties for the partial and total Higgs-boson decay width predictions. The last column: the target of FCC-ee precisions.

decay	intrinsic	para. m_q	para. $lpha_{ m s}$	para. $M_{\rm H}$	FCC-ee prec. on g^2_{HXX}
$H \rightarrow b\bar{b}$	$\sim 0.2\%$	0.6%	< 0.1%	-	$\sim 0.8\%$
$H \rightarrow c\bar{c}$	$\sim 0.2\%$	$\sim 1\%$	< 0.1%	-	$\sim 1.4\%$
$H \to \tau^+ \tau^-$	< 0.1%	-	-	-	$\sim 1.1\%$
$H \rightarrow \mu^+ \mu^-$	< 0.1%	-	-	-	$\sim 12\%$
$H \rightarrow gg$	$\sim 1\%$		0.5% (0.3%)	-	$\sim 1.6\%$
$H \rightarrow \gamma \gamma$	< 1%	-	-	-	$\sim 3.0\%$
$H \rightarrow Z\gamma$	$\sim 1\%$	-	-	$\sim 0.1\%$	
$H \rightarrow WW$	$\lesssim 0.3\%$	-	-	$\sim 0.1\%$	$\sim 0.4\%$
$H \rightarrow ZZ$	$\lesssim 0.3\%^\dagger$	-	-	$\sim 0.1\%$	$\sim 0.3\%$
$\Gamma_{\rm tot}$	$\sim 0.3\%$	$\sim 0.4\%$	< 0.1%	< 0.1%	$\sim 1\%$
+ _					

[†] From $e^+e^- \rightarrow HZ$ production



SM precision parameters determination: $lpha(M_Z^2)$, F. Jegerlechner, pdf

1. $\alpha(M_Z^2)$ in precision physics (precision physics limitations)

Uncertainties of hadronic contributions to effective α are a problem for electroweak precision physics: besides top Yukawa y_t and Higgs self-coupling λ

q, G_{μ} , M_Z most precise input parameters \Rightarrow precision predictions 50% non-perturbative $\sin^2 \Theta_f, v_f, a_f, M_W, \Gamma_Z, \Gamma_W, \cdots$

 $\alpha(M_Z), G_\mu, M_Z$ best effective input parameters for VB physics (Z,W) etc.

$$\frac{\frac{\delta a}{g}}{G_{\mu}} \sim 3.6 \times 10^{-9} \\ \frac{\frac{\delta G_{\mu}}{G_{\mu}}}{M_{Z}} \sim 8.6 \times 10^{-6} \\ \frac{\delta M_Z}{M_Z} \sim 2.4 \times 10^{-5} \\ \frac{\delta (M_Z)}{a(M_Z)} \sim 0.9 \div 1.6 \times 10^{-4} \text{ (present : lost 105 in precision!)} \\ \frac{\delta a(M_Z)}{a(M_Z)} \sim 5.3 \times 10^{-5} \text{ (FCC - ee/ILC requirement)}$$

$$\begin{split} \textbf{LEP/SLD:} & \sin^2 \Theta_{\text{eff}} = (1 - v_l/a_l)/4 = 0.23148 \pm \underbrace{0.00017}_{\delta \Delta \alpha}(M_Z) = 0.00020 \qquad \Rightarrow \qquad \delta \sin^2 \Theta_{\text{eff}} = \underbrace{0.00007}_{0.00007} \text{ ; } \delta M_W/M_W \sim 4.3 \times 10^{-5} \\ & \textbf{affects most precision tests and new physics searches!!!} \\ & \frac{\delta M_W}{M_W} \sim 1.5 \times 10^{-4} \text{, } \frac{\delta M_H}{M_H} \sim 1.3 \times 10^{-3} \text{, } \frac{\delta M_l}{M_l} \sim 2.3 \times 10^{-3} \end{split}$$

For pQCD contributions very crucial: precise QCD parameters α_s , m_c , m_b , $m_t \Rightarrow$ Lattice-QCD

SM precision parameters determination: $\alpha(M_Z^2)$

Still an issue in HVP region 1.2 to 2 GeV data; test-ground exclusive vs inclusive R measurements (more than 30 channels!) VEPP-2000 CMD-3, SND (NSK) scan, BaBar, BES III radiative return! still contributes 50% of uncertainty з. 2017 pQCD pQCD 3.0-3.0 2.5 2.5 2.0 2.0 ď. £ excl vs incl clash 1.5 1.5 1.0 1.0 excl (inc. BaBar) incl (BES-II, KEDR) ncl (exc. BES-II) 0.5 0.5 0.0 1 40 1.60 1.80 2.00 2 20 2.40 2.60 1.40 1 60 1 80 2.00 2.20 2.40 2.60 F (GeV) E (GeV)

 illustrating progress by BaBar and NSK exclusive channel data vs new inclusive data by KEDR. Why point at 1.84 GeV so high? Three approaches should be further explored for better error estimate

Note: theory-driven standard analyses (R(s) integral) using pQCD above 1.8 GeV cannot be improved by improved cross-section measurements above 2 GeV !!!

precision in a:	present	direct	1.7×10^{-4}
		Adler	1.2×10^{-4}
	future	Adler QCD 0.2%	5.4×10^{-5}
		Adler QCD 0.1%	3.9×10^{-5}
	future	via $A_{\rm FB}^{\mu\mu}$ off Z	3×10^{-5}

 Adler function method is competitive with Patrick Janot's direct near Z pole determination via forward backward asymmetry in e⁺e⁻ → μ⁺μ⁻

$$A_{\rm FB}^{\mu\mu} = A_{\rm FB,0}^{\mu\mu} + \frac{3 a^2}{4 v^2} \frac{I}{Z + G}$$

where $\gamma - Z$ interference term Z alone

$$\gamma$$
 only

v vector Z coupling

a axial *Z* coupling

 $I \propto \alpha(s) G_{\mu}$ $Z \propto G_{\mu}^{2}$ $G \propto \alpha^{2}(s)$ also depends on $\alpha(s \sim M_{Z}^{2})$ and $\sin^{2} \Theta_{f}(s \sim M_{Z}^{2})$ sensitive to ρ -parameter (strong M_{t} dependence)

 \Box using *v*, *a* as measured at Z-peak

 $e^+e^- \rightarrow \mu^+\mu^-$ and $\alpha^2(s)$

 $\sigma_{\mu\mu}$:

- 1. the photon-exchange term, \mathcal{G} , proportional to $\alpha^2(s)$;
- 2. the Z-exchange term, Z, proportional to G_F^2 (where G_F is the Fermi constant);
- 3. the Z-photon interference term, \mathcal{I} , proportional to $\alpha(s) \times G_F$

The muon forward-backward asymmetry, $A_{\rm FB}^{\mu\mu}$, is maximally dependent on the interference term

$$A_{\rm FB}^{\mu\mu} = A_{\rm FB,0}^{\mu\mu} + \frac{3}{4} \frac{2}{2} \frac{\mathcal{I}}{\mathcal{G} + \mathcal{Z}},$$

varies with $\alpha_{QED}(s)$ as follows:

$$\Delta A_{\rm FB}^{\mu\mu} = \left(A_{\rm FB}^{\mu\mu} - A_{\rm FB,0}^{\mu\mu} \right) \times \frac{\mathcal{Z} - \mathcal{G}}{\mathcal{Z} + \mathcal{G}} \times \frac{\Delta \alpha}{\alpha}.$$

 $e^+e^- \rightarrow \mu^+\mu^-$ and $\alpha^2(s)$



The best accuracy is obtained for one year of running either just below or just above the Z pole, at 87.9 and 94.3 GeV, respectively.