

Dynamical R-parity breaking: from Cosmology to early LHC

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Based on arxiv: 1010.3968 and 1011.2214

OUTLINE

- Motivation
- Light sterile neutrino(s) & Scale- TeV
- New R-parity breaking interactions
- LHC signatures

MSSM and R-parity

- A fundamental question of Supersymmetry: is R-parity good symmetry or not?
- Conserved R-p implies stable LSP.
- MSSM with R-p implies massless neutrino, or needs higher dimensional operators $\Delta L = 2$.
- **MSSM as a complete theory must break R-p.**
- Gravitino: still possible DM, strong correlations between indirect detection and neutrino mass.

(Bajc, Enkhbat, Ghosh, Senjanovic, YZ 1002.3631)

Origin of R-parity violation

- Take the issue of R-parity breaking seriously.
- Study common predictions of supersymmetric theories with spontaneous R-parity breaking.

Breaking R-parity means: $R - \text{parity} \propto (-1)^{B-L}$

- Break lepton number: Neutrino physics
- Tie the R-p breaking scale to soft mass scale: LHC

Gauging the B-L

- Spontaneously Break R-parity? MSSM $\tilde{\nu}$: No way. Also break B-L **global symmetry**, Goldstone particle (Majoron) -- ruled out by LEP.
(Aulakh, Mohapatra, 82')
- Use a SM singlet -- explicit breaking $\mu_i L_i H_u$ plus a light J -- no prediction on the scale.
- **Gauging B-L** -- possible TeV scale realizations:

$$SU(2)_L \times U(1)_Y \times U(1)_{B-L}$$

$$SU(2)_L \times SU(2)_R \times U(1)_{B-L}$$

The minimal model

- Three RH neutrinos for anomaly cancelation. $U(1)_{B-L}$
- Break B-L: minimal scenario - use RH sneutrino

$$V \sim g_{BL}^4 \tilde{\nu}^{c4} - M_{SUSY}^2 \tilde{\nu}^{c2}$$

- $\langle \tilde{\nu}^c \rangle \approx M_{SUSY} / g_{BL}^2$ also break R-parity - SM singlet.
(Mohapatra, 86', Barger, et al 0812.3661)

- $M_{Z'} \approx M_{SUSY} / g_{BL}$ scale tied to M_{SUSY}

- If low-energy SUSY, theoretical upper bound:

$$M_{Z'} \sim \text{TeV}$$

- Tree level, only one massive RH neutrino, “eaten” by \tilde{Z}' , **two others are light.**

$$\begin{bmatrix} 0 & M_{Z'} \\ M_{Z'} & m_{1/2} \end{bmatrix} \sim \text{TeV}$$

- Loop corrections safe - if $Y_\nu = 0$ break B-L only in one “flavor”.

- HD operators: $\frac{N_1^{c\dagger} N_1^{c\dagger} N_a^c N_b^c}{M^2} \left(\frac{H_u H_d}{M^2} \right)^p \left(\frac{H_u^\dagger H_d^\dagger}{M^2} \right)^q$

- Upper bound on N_a mass $\leq \text{MeV}$. $M \gtrsim 10M_{BL}$

- Decaying slowly, lifetime \sim universe.

- **Must be eV** in order not to dominate the universe.

(Ghosh, Senjanovic, YZ 1010.3968)

- Tree level, only one massive RH neutrino, “eaten” by \tilde{Z}' , two others are light.

$$\begin{bmatrix} 0 & M_{Z'} \\ M_{Z'} & m_{1/2} \end{bmatrix} \sim \text{TeV} \begin{bmatrix} m_{\nu_1} & 0 & 0 & m_D^1 & m_D^2 \\ 0 & m_{\nu_2} & 0 & m_D^3 & m_D^4 \\ 0 & 0 & 0 & m_D^5 & m_D^6 \\ m_D^1 & m_D^3 & m_D^5 & 0 & 0 \\ m_D^2 & m_D^4 & m_D^6 & 0 & 0 \end{bmatrix} \sim \text{eV}$$

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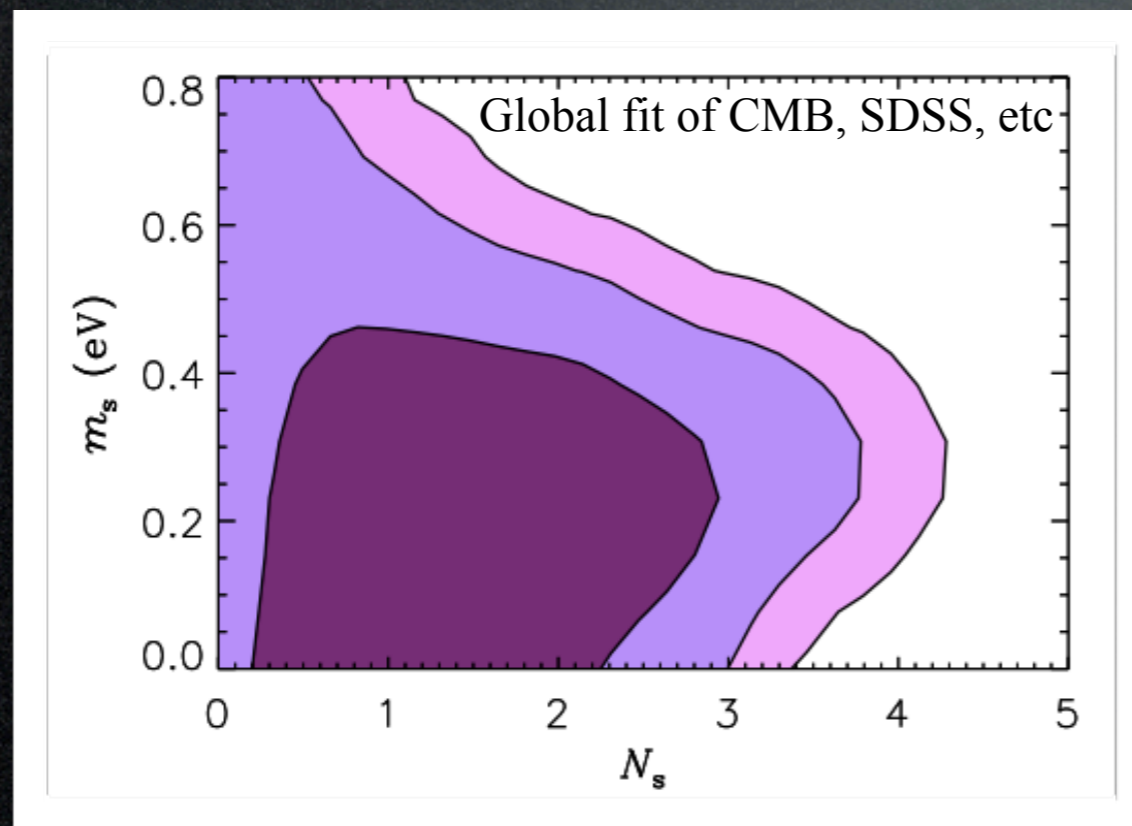
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(Ghosh, Senjanovic, YZ 1010.3968)

Cosmology prefers more neutrinos

- WMAP: $N_{\text{eff}} \approx 4.34^{+0.86}_{-0.88}$ LSS: $N_{\text{eff}} \approx 4.78^{+1.86}_{-1.79}$
- BBN: ${}^4\text{He}, D$ abundances also prefer $N_{\text{eff}} > 3.046$
(Hamann, et al 1006.5276)
- A natural candidate: light RH neutrinos (3+2) -- thermalized through t-channel Z' : $N_i e \rightarrow N_i e$



Phenomenological upper bound:

$$M_{Z'} \lesssim 10\text{TeV}$$

R-parity in left-right theories

- **Minimal version:** $U(1)_Y \times U(1)_{B-L} \rightarrow SU(2)_R \times U(1)_{B-L}$

$$V \sim g_{BL}^4 (\tilde{L}^\dagger \tilde{L} - \tilde{L}^{\dagger c} \tilde{L}^c)^2 + \tilde{L}^\dagger m_{\tilde{l}}^2 \tilde{L} - \tilde{L}^{\dagger c} m_{\tilde{l}^c}^2 \tilde{L}^c$$

(Fileviez Perez, Spinner 0811.3424)

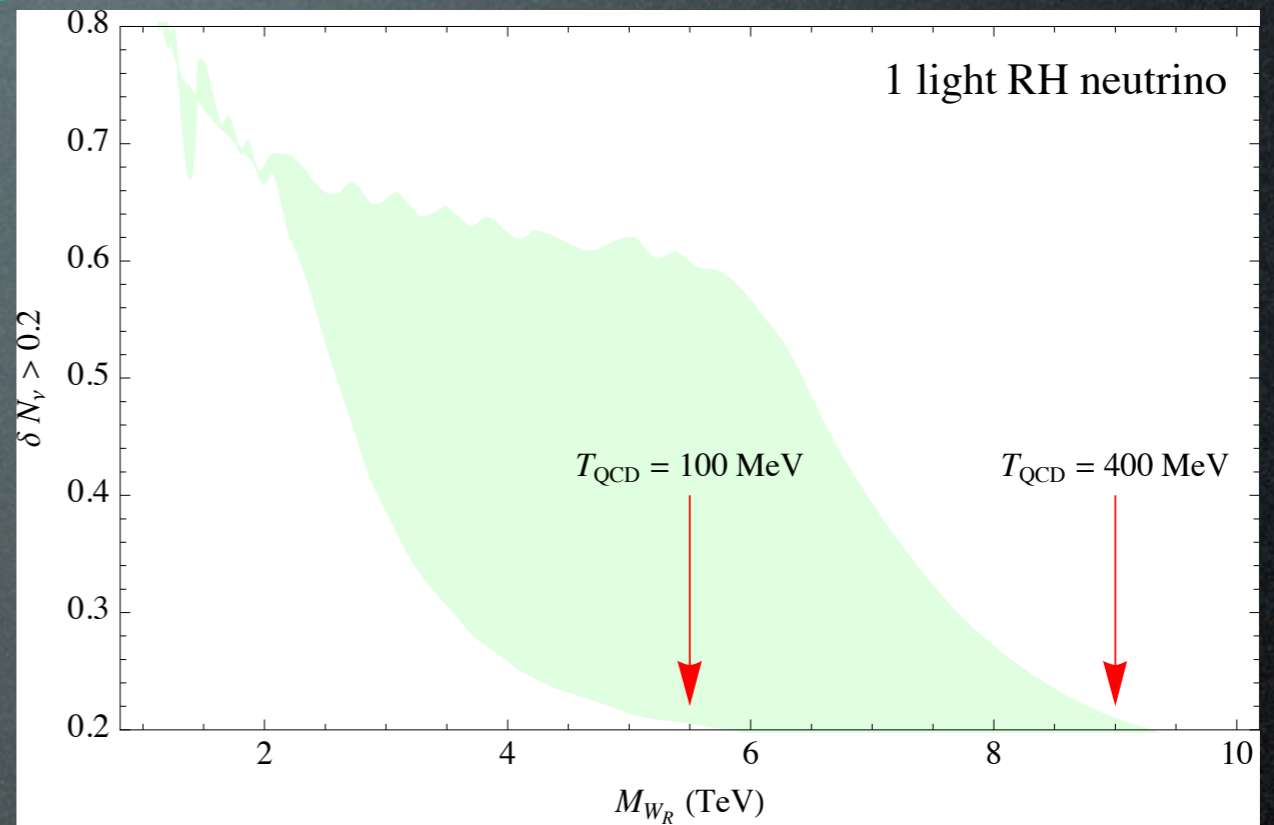
- Asymmetric soft masses -- “parity” symmetry broken at high scale.
- R-p breaking scale: $M_{W_R} \approx M_{SUSY} / g_R$
- **Seesaw version**, add Higgs triplets $\Delta, \bar{\Delta}, \Delta^c, \bar{\Delta}^c$.
- $\langle \Delta^c \rangle$ could **lift one** light RH neutrino mass

Setting the scale

More predictive:

$$M_{W_R} \lesssim 5 - 9 \text{ TeV}$$

- **3+1** scenario:
- Lifting one RH neutrino mass with $\langle \Delta^c \rangle$: **even better** fit to cosmological observations.
- More predictive: **Scale** & bring new signatures from the heavy neutrino to LHC.



(Chen, Ghosh, Mohapatra, YZ 1011.2214)

Seesaw version of susylr

- SUSY limit, tree-level: no symmetry breaking.

$$W \sim \mu_\Delta \text{Tr}(\Delta \bar{\Delta}) + \mu_\Delta \text{Tr}(\Delta^c \bar{\Delta}^c) + f L \Delta^c L$$

- Not D-flat direction: $\langle \Delta^c \rangle = \begin{bmatrix} 0 & 0 \\ v_R & 0 \end{bmatrix}$ $\langle \bar{\Delta}^c \rangle = \begin{bmatrix} 0 & \bar{v}_R \\ 0 & 0 \end{bmatrix}$
(Kuchimanchi, Mohapatra, 93')

- **HD operators:** R-parity conserved by the vev.
(Aulakh, Melfo, Senjanovic, 97')

- **Including soft terms:** R-parity violating direction.

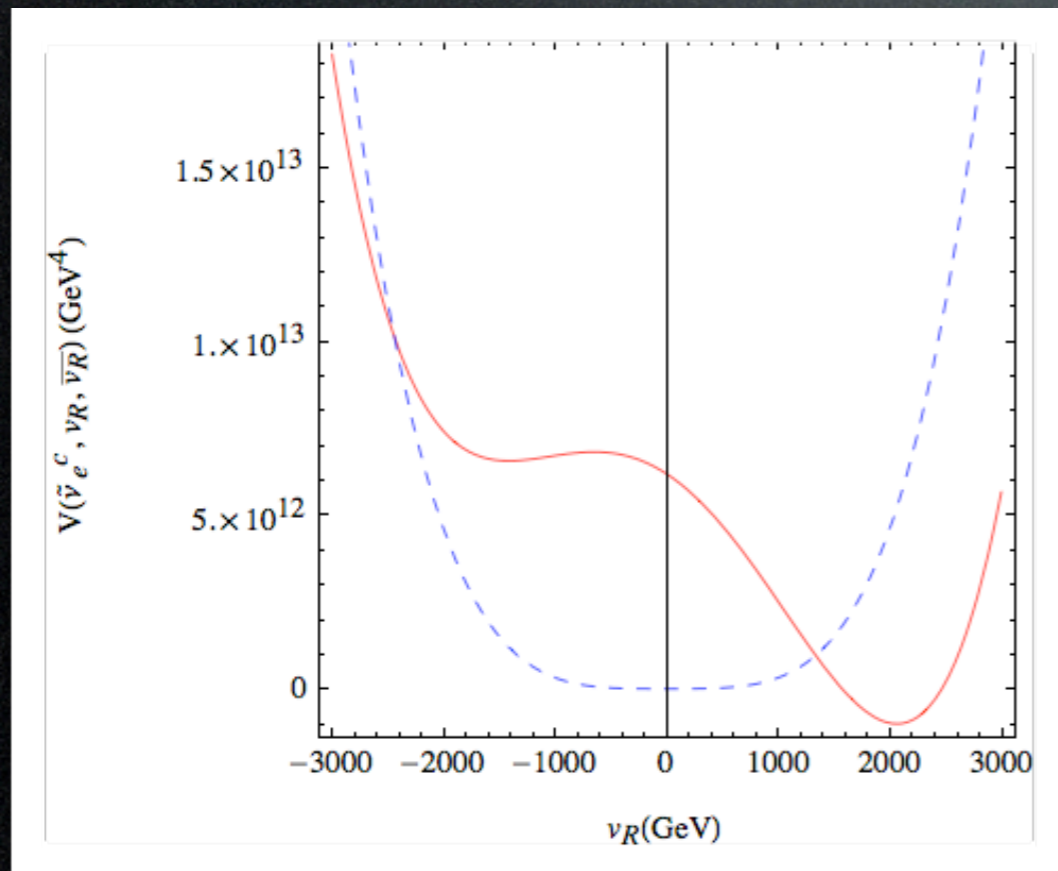
$$V = \text{soft masses} + \left[-2|A_f|v_R - 2|f|\mu_\Delta \bar{v}_R \right] \langle \tilde{\nu}_e^c \rangle^2 + \text{F, D - terms}$$

- If $\langle \Delta^c \rangle$ nonzero, develop tachyonic mass for $\tilde{\nu}_e^c$
(Ji, Mohapatra, Nussinov, YZ, 0808.1904)

Dynamical R-parity breaking

- Deeper vacuum in R-parity breaking direction.
- Amount of R-pV of the same scale as gauge symmetry breaking: $\langle \tilde{\nu}_e^c \rangle \sim v_R \sim \bar{v}_R$.
- Still nearly D-flat: $\langle \tilde{\nu}_e^c \rangle^2 - 2v_R^2 + 2\bar{v}_R^2 \approx 0$.

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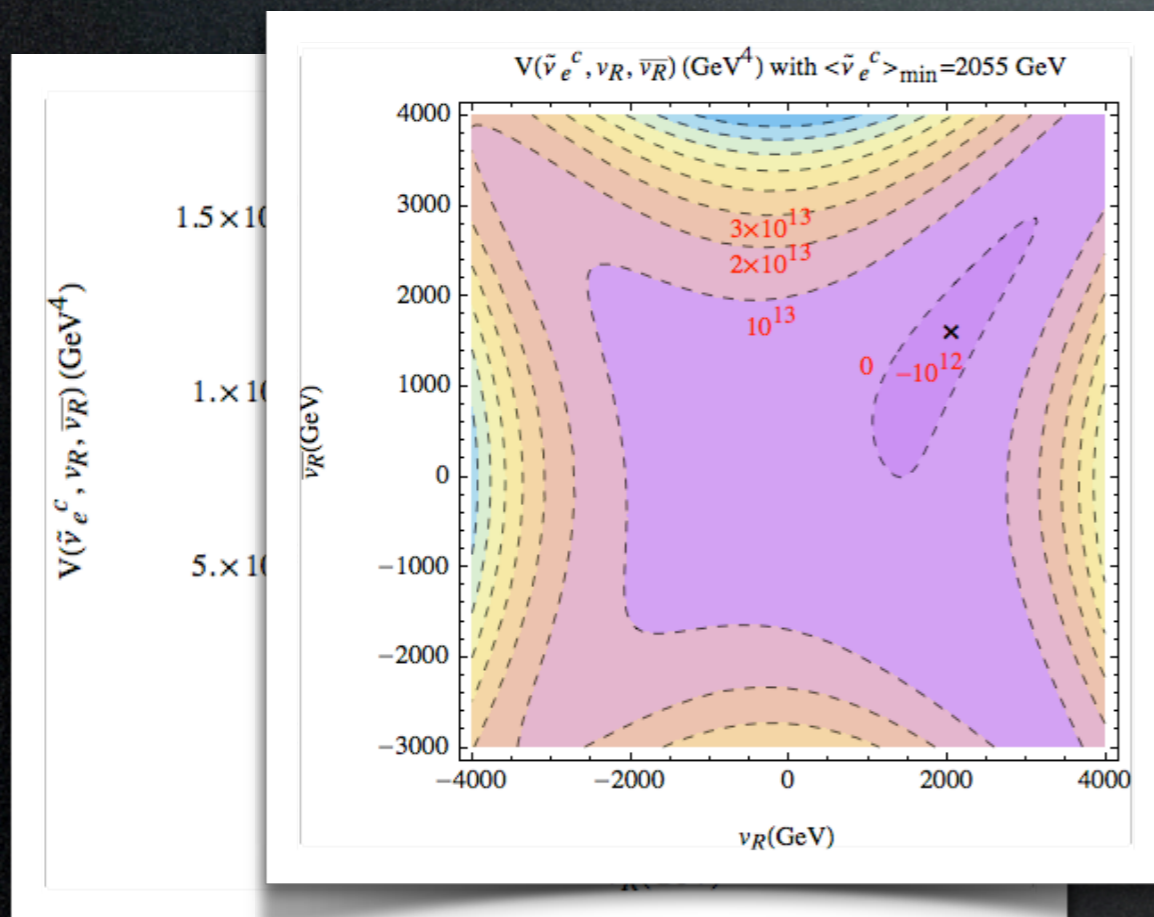
Scale again tied to M_{SUSY}

$$M_{W_R} \approx g M_{\text{SUSY}} / f^2$$

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Generic R-parity breaking

- Generate usual MSSM R-pV terms, but small.. and?
- MSSM: no large lepton-gaugino mixings.

$$\begin{array}{c} e \quad \widetilde{W}^- \\ \widetilde{W}^+ \left[\begin{array}{cc} m_e & 0 \\ g\langle\tilde{\nu}\rangle & m_{1/2} \end{array} \right] \end{array} \xleftrightarrow{SU(2)_L} \begin{array}{c} \nu \quad \widetilde{Z} \\ \widetilde{Z} \left[\begin{array}{cc} 0 & g_Z\langle\tilde{\nu}\rangle \\ g_Z\langle\tilde{\nu}\rangle & m_{1/2} \end{array} \right] \end{array}$$

- SUSYLR: large mixings possible.

$$\begin{array}{c} e \quad \widetilde{W}_R^- \\ \widetilde{W}_R^+ \left[\begin{array}{cc} m_e & g_R\langle\tilde{\nu}^c\rangle \\ 0 & m_{1/2} \end{array} \right] \end{array} \xleftrightarrow{SU(2)_R} \begin{array}{c} \nu^c \quad \widetilde{Z}' \\ \widetilde{Z}' \left[\begin{array}{cc} 0 & g'_Z\langle\tilde{\nu}^c\rangle \\ g'_Z\langle\tilde{\nu}^c\rangle & m_{1/2} \end{array} \right] \end{array}$$

- Physical electron $\hat{e}^c = \theta_{ee}e^c + \theta_{eW}\widetilde{W}_R^+$

Chirality preserving

$$\hat{\nu}^c = \theta_{NN}\nu^c + \theta_{NZ}\widetilde{Z}', \quad \theta_{eW}, \theta_{NZ} \sim O(1)$$

Large RPV interactions

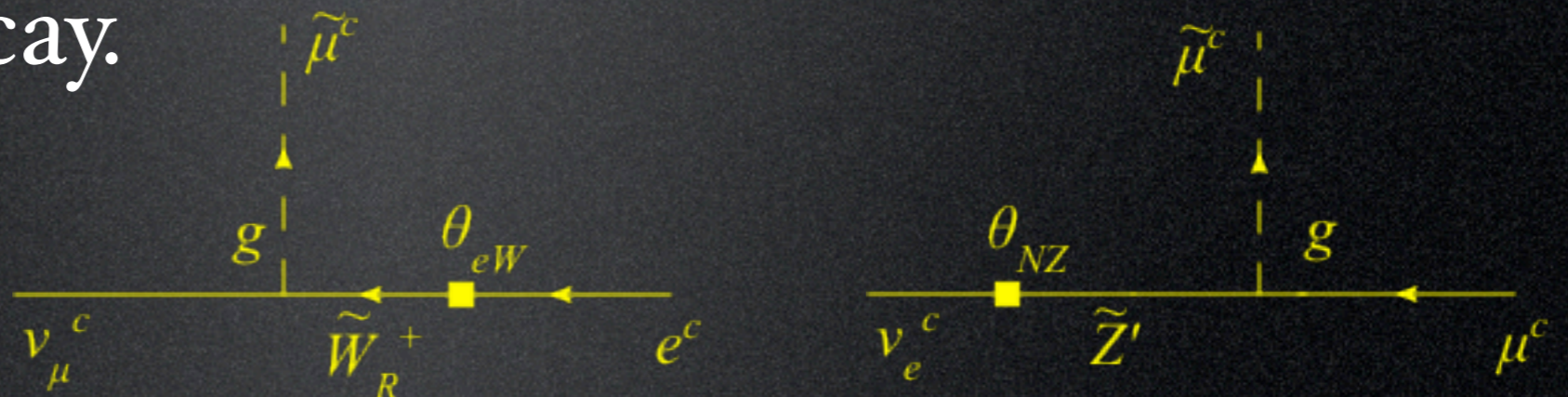
(Chen, Ghosh, Mohapatra, YZ 1011.2214)

- Integrate out extra heavy gauginos.

$$\sqrt{2}g\theta_{eW} \left[\hat{e}^c \nu_\ell^c \tilde{\ell}^{c\dagger} + \hat{e}^c u^c \tilde{d}^{c\dagger} \right] + \sqrt{2}g_{Z'}\theta_{NZ} \left[\hat{\nu}_e^c \ell^c \tilde{\ell}^{c\dagger} + \hat{\nu}_e^c q^c \tilde{q}^{c\dagger} \right] + \text{h.c.}$$

$(\ell = \mu, \tau)$

- As large as gauge coupling.
- No susy counterparts -- susy breaking terms as trigger of R-p violation.
- Facilitate **single production** of smuon/stau from $\nu_{e,\mu}^c$ decay.

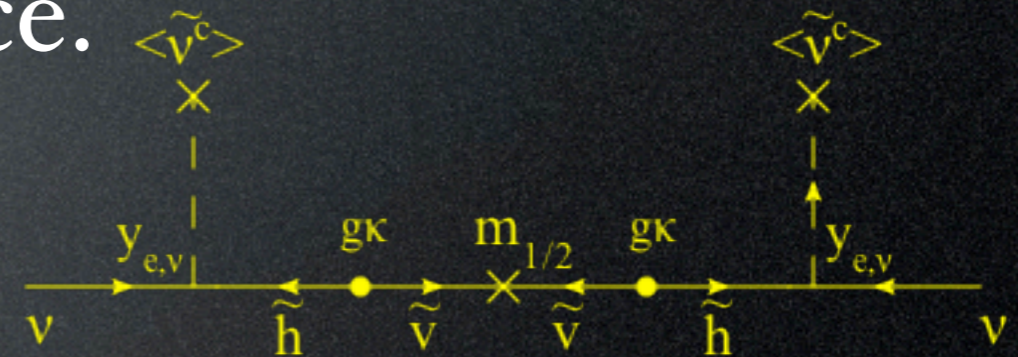


Flavor alignment of R-p V

- **Neutrino Mass:** Yukawa couplings in LR

$$W = Y_\ell L \Phi_1 L^c + Y_\nu L \Phi_2 L^c, \quad \langle \Phi_1 \rangle = \begin{bmatrix} v_1 & 0 \\ 0 & 0 \end{bmatrix}, \quad \langle \Phi_2 \rangle = \begin{bmatrix} 0 & 0 \\ 0 & v_2 \end{bmatrix}$$

- $\langle \tilde{\nu}_\ell^c \rangle$ induces large LNV: LHu unless only couples to $Y_e \sim 10^{-6}$ -- a natural choice.



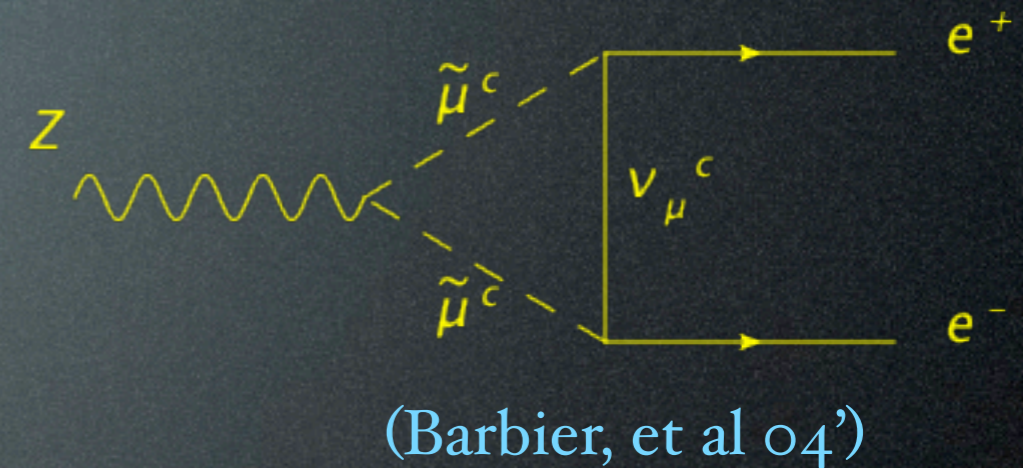
- **Lepton flavor violation:** safe only if R-pV in single (electron) flavor.

- Only suffer the usual SUSY flavor problem.

pre-LHC constraints

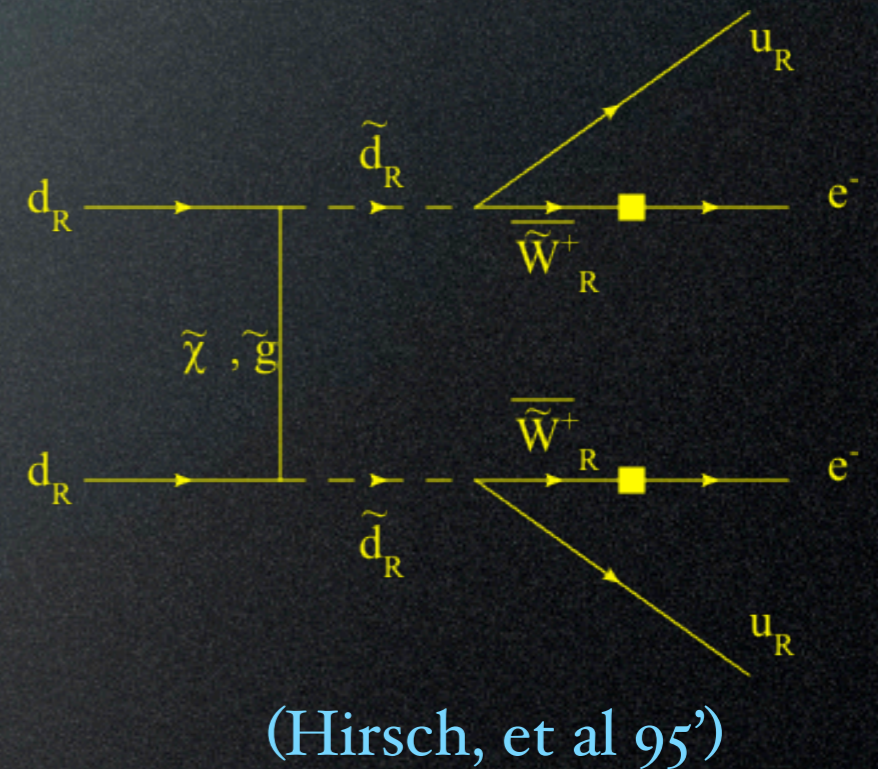
- LEP: loop correction Z to ee.

$$g\theta_{eW} \lesssim 0.5 \left(\frac{m_{\tilde{\mu}^c, \tilde{d}^c}}{100\text{GeV}} \right)$$

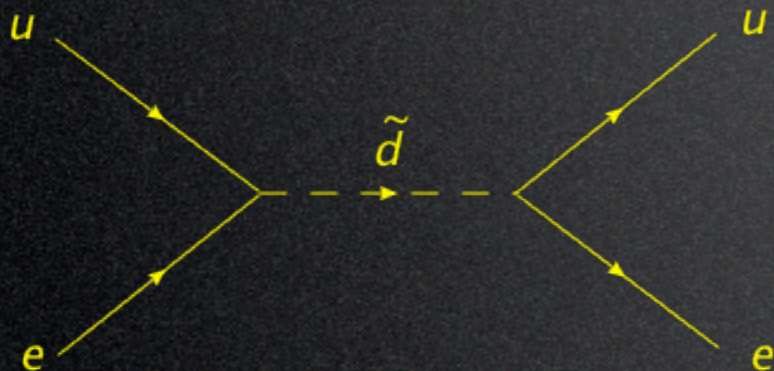


- neutrino-less double beta decay.

$$g\theta_{eW} \lesssim 0.39 \times 10^{-3} \left(\frac{m_{\tilde{d}^c}}{100\text{GeV}} \right)^2 \left(\frac{m_{\tilde{g}}}{100\text{GeV}} \right)^{1/2}$$



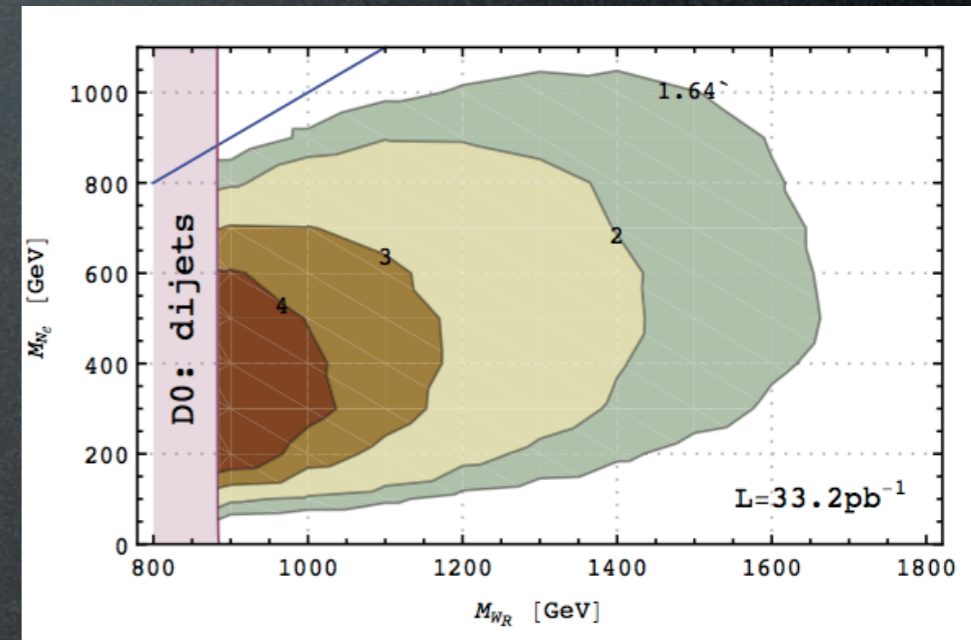
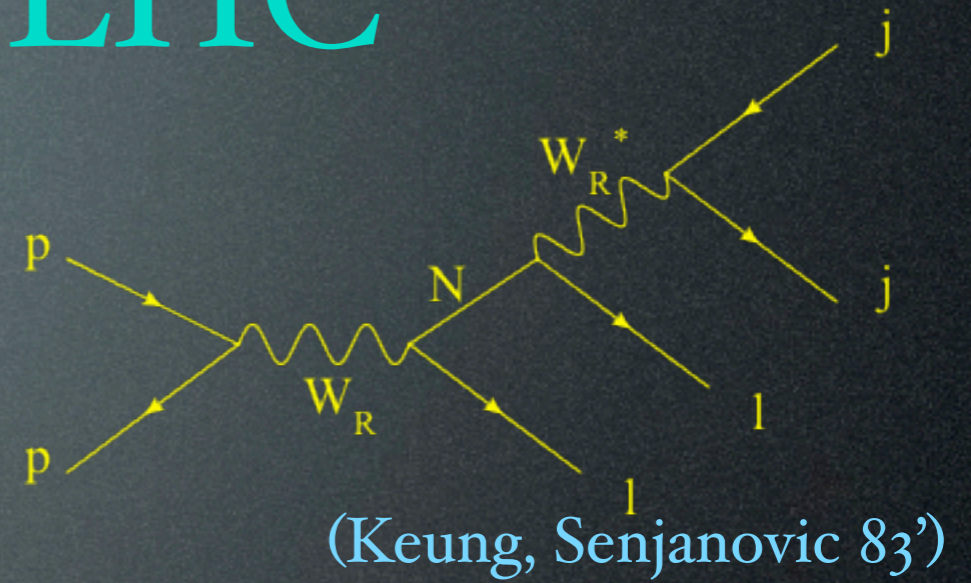
- Hera single production.



$$m_{\tilde{\mu}^c} \gtrsim 100 \text{ GeV}$$

$$m_{\tilde{q}}, m_{\tilde{g}} \gtrsim 1 \text{ TeV}$$

Signature at LHC

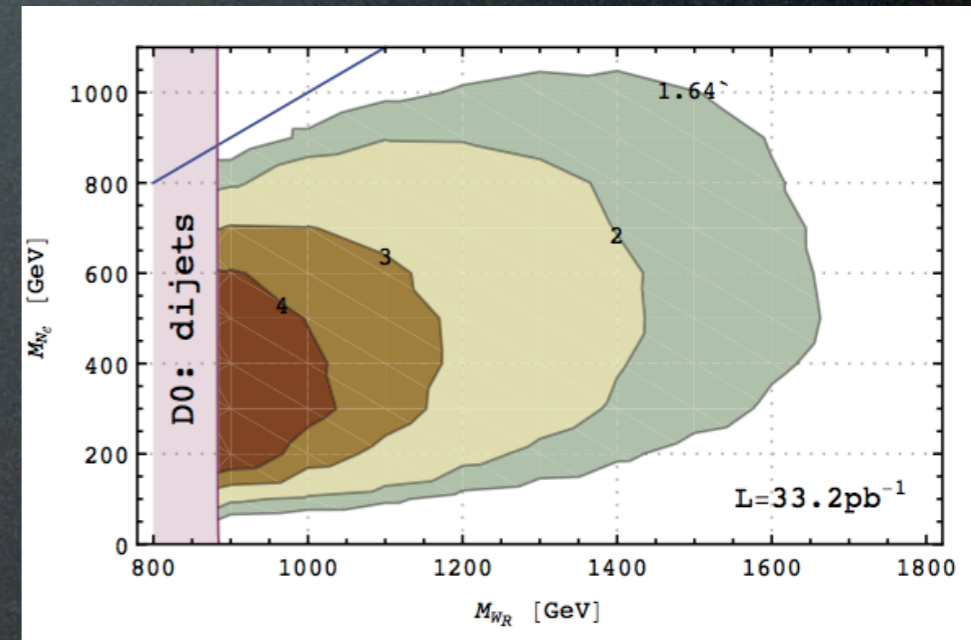
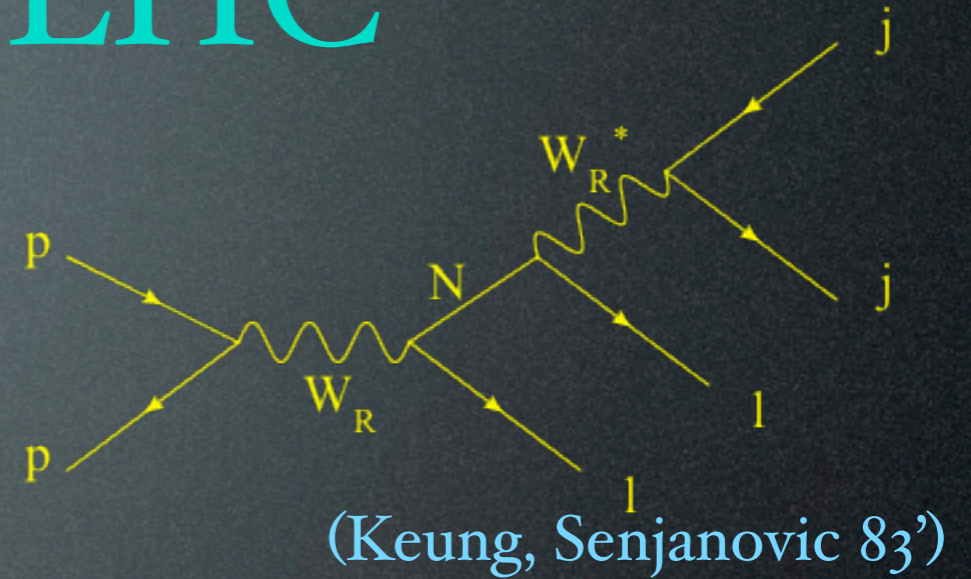
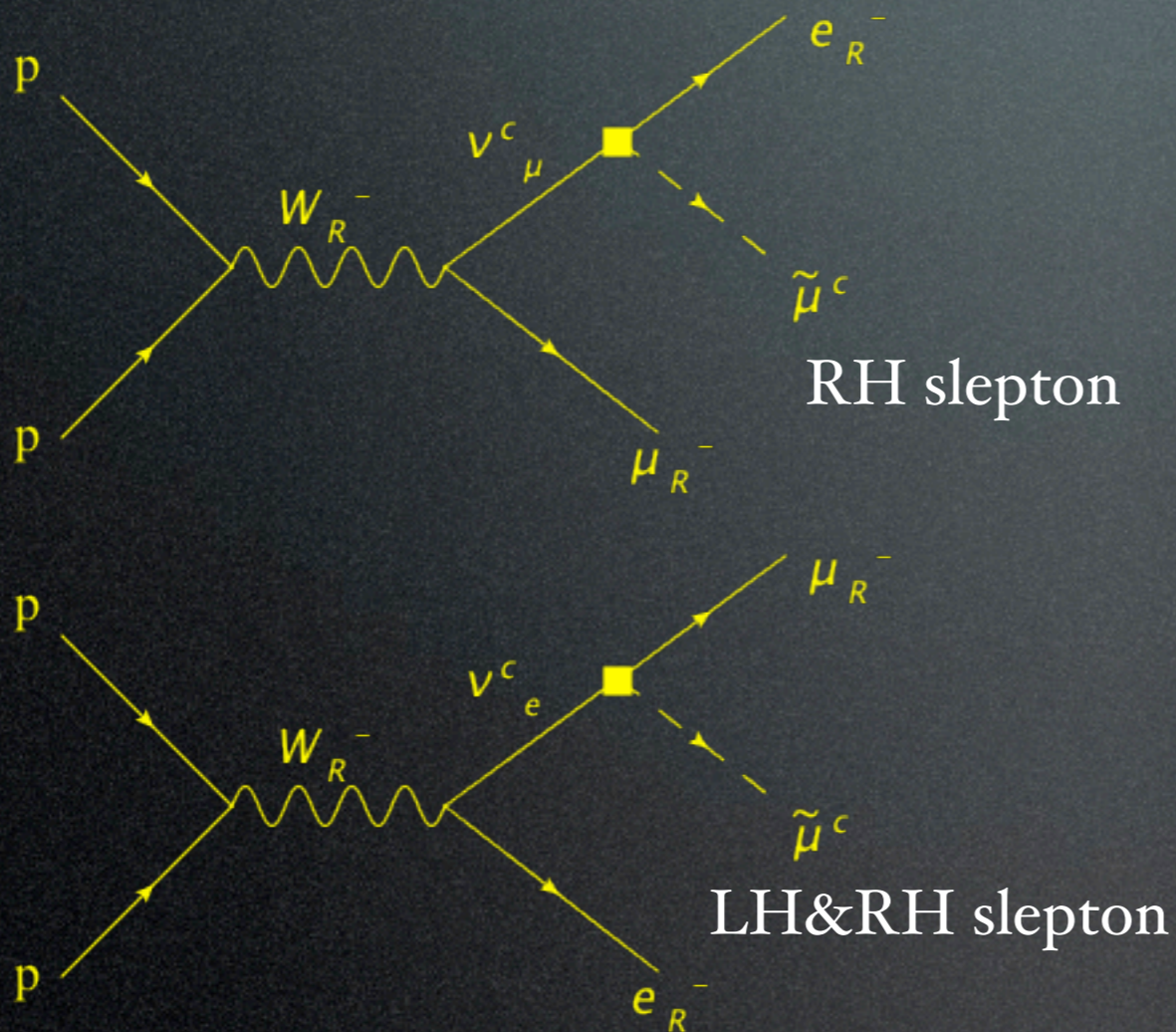


(Nemevsek, Nesti, Senjanovic, YZ 1010.3968)

See also Fabrizio Nesti' talk

Signature at LHC

$$pp \rightarrow e\mu\tilde{\mu}$$



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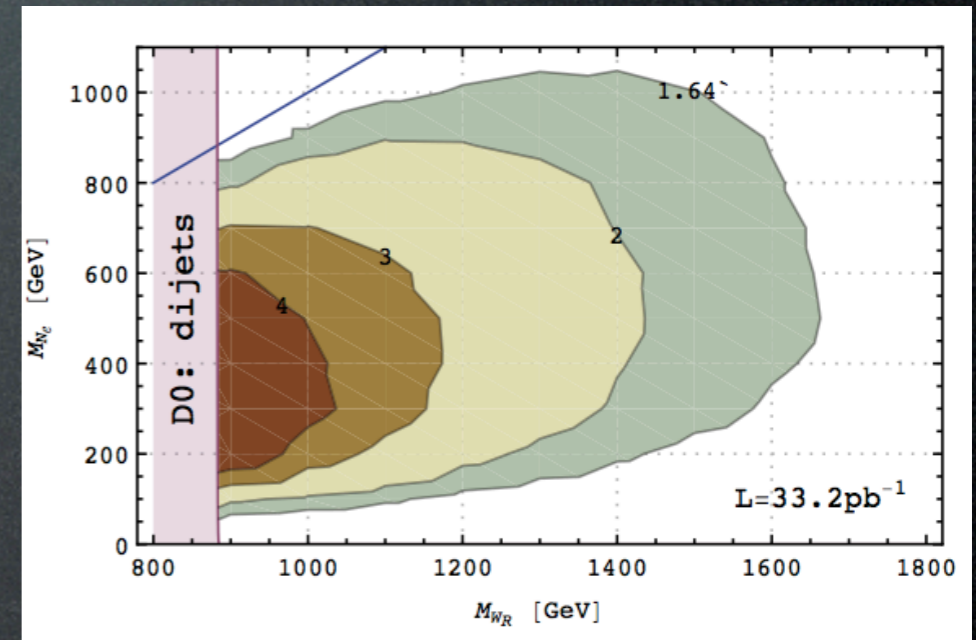
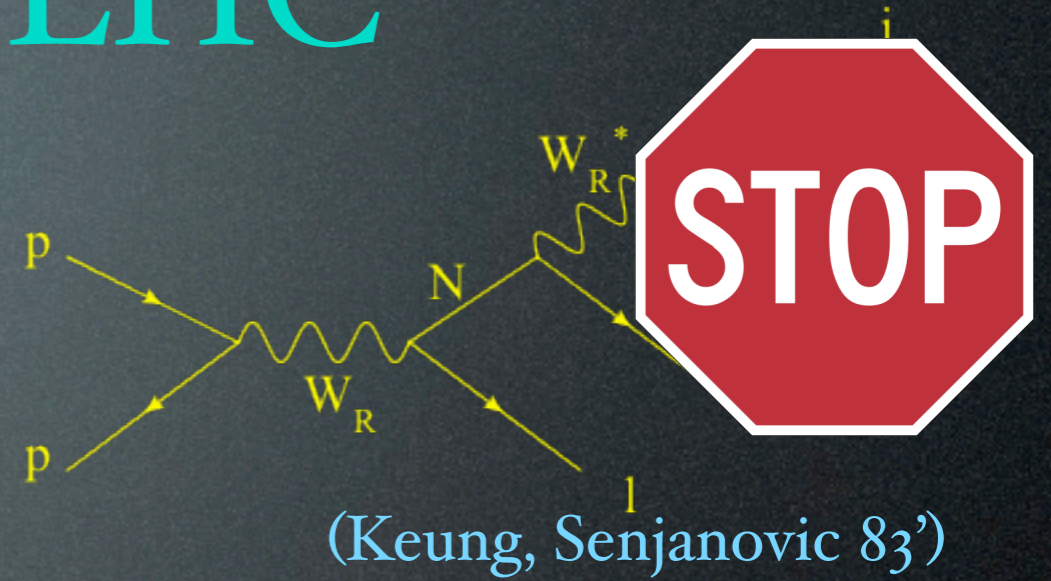
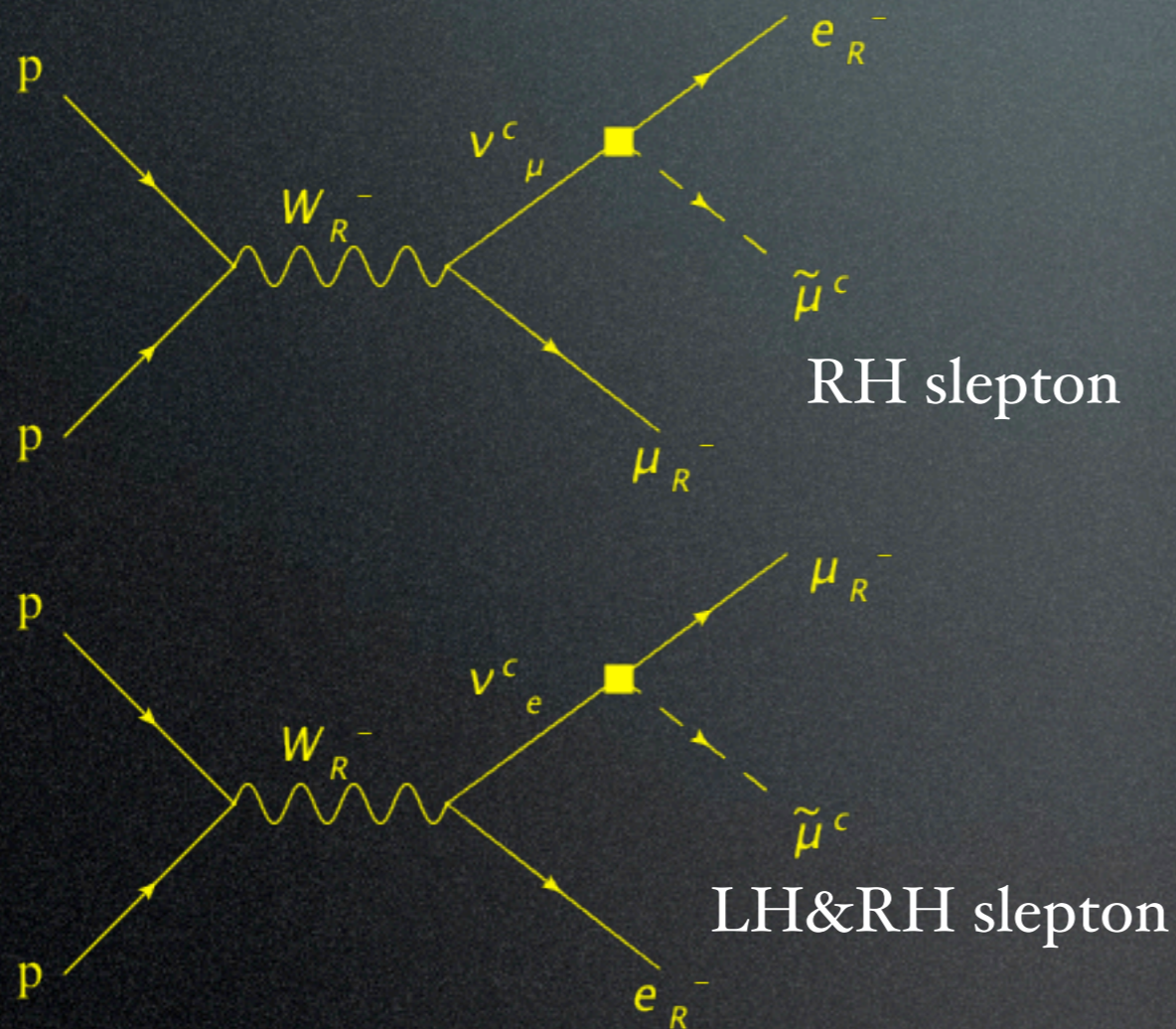
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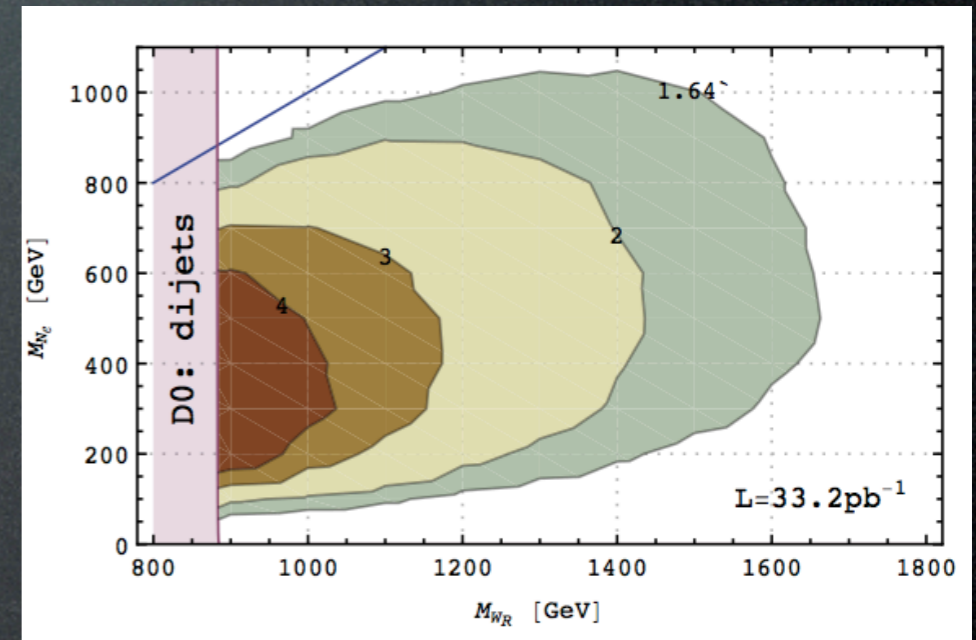
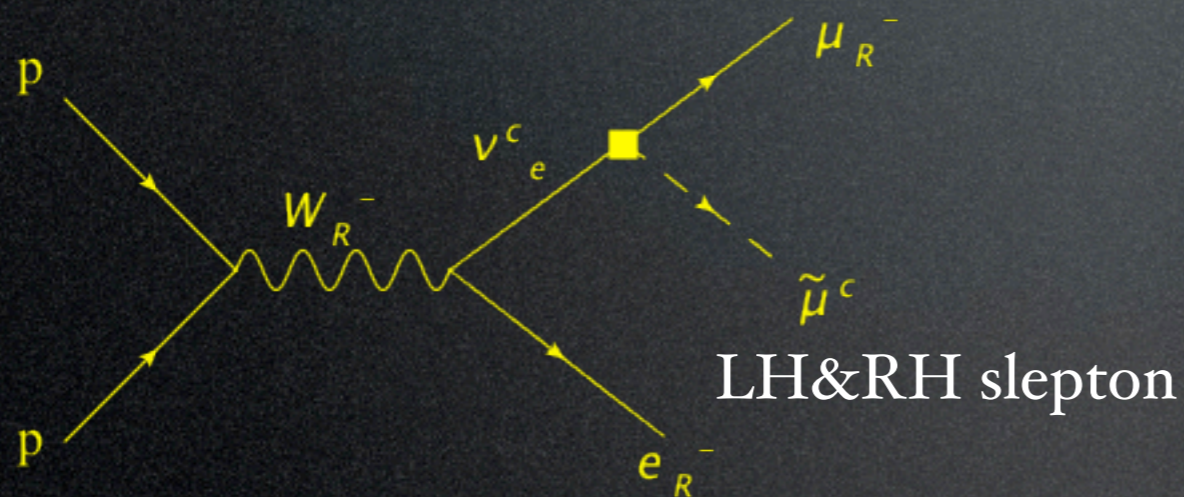
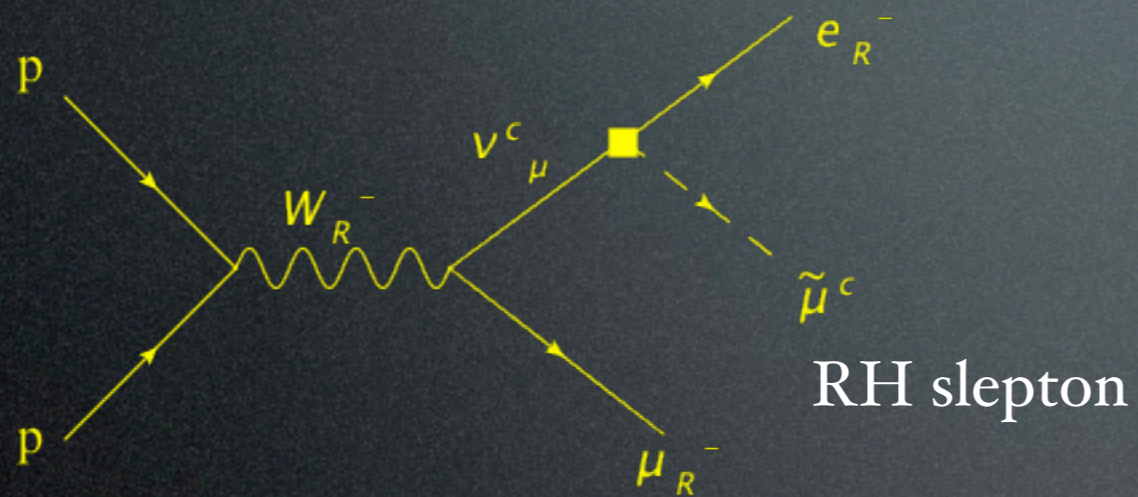
(Ghosh, Senjanovic, YZ 1010.3968)

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Signature at LHC

- Single production of slepton

$$pp \rightarrow e\mu\tilde{\mu}$$



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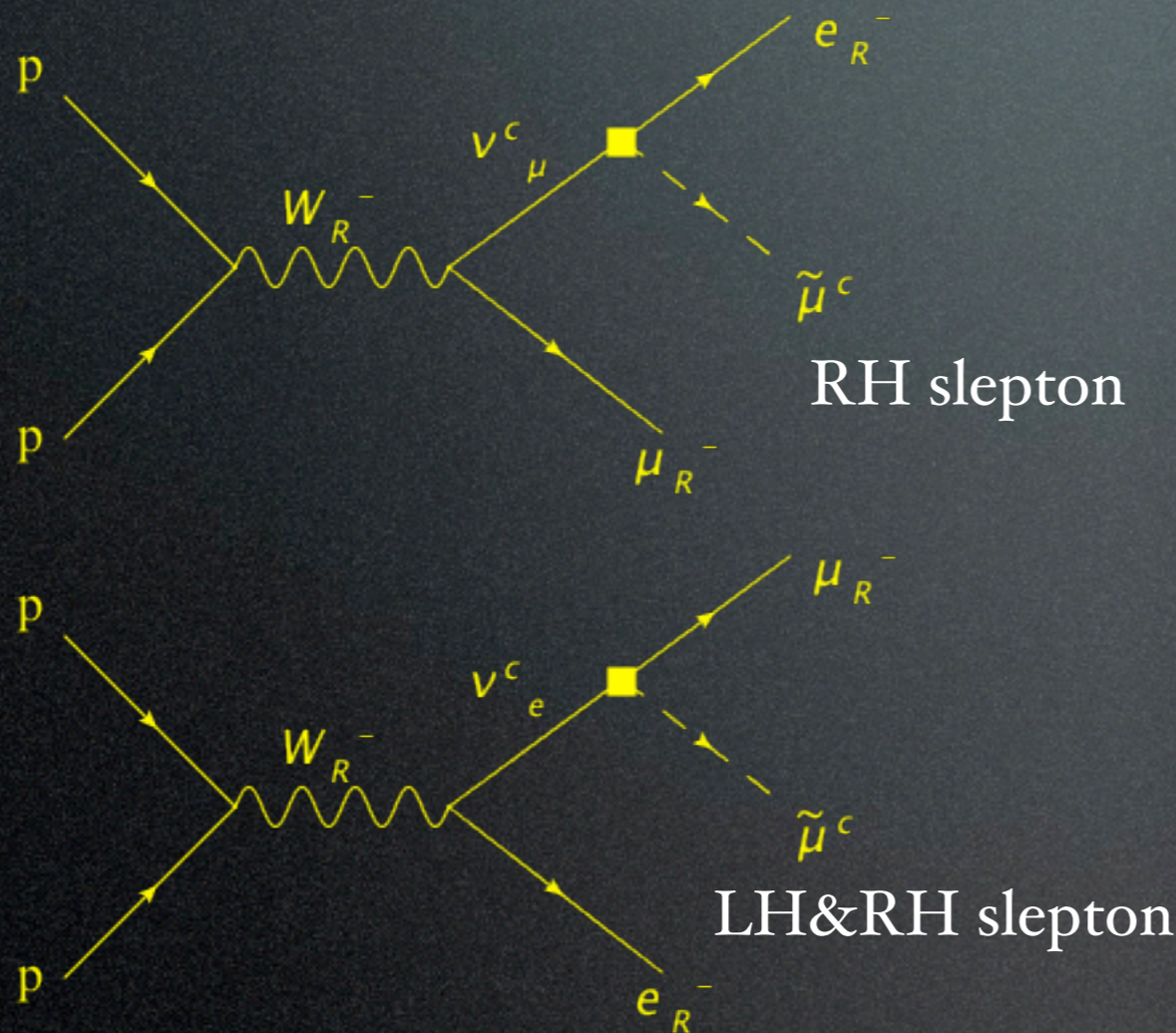
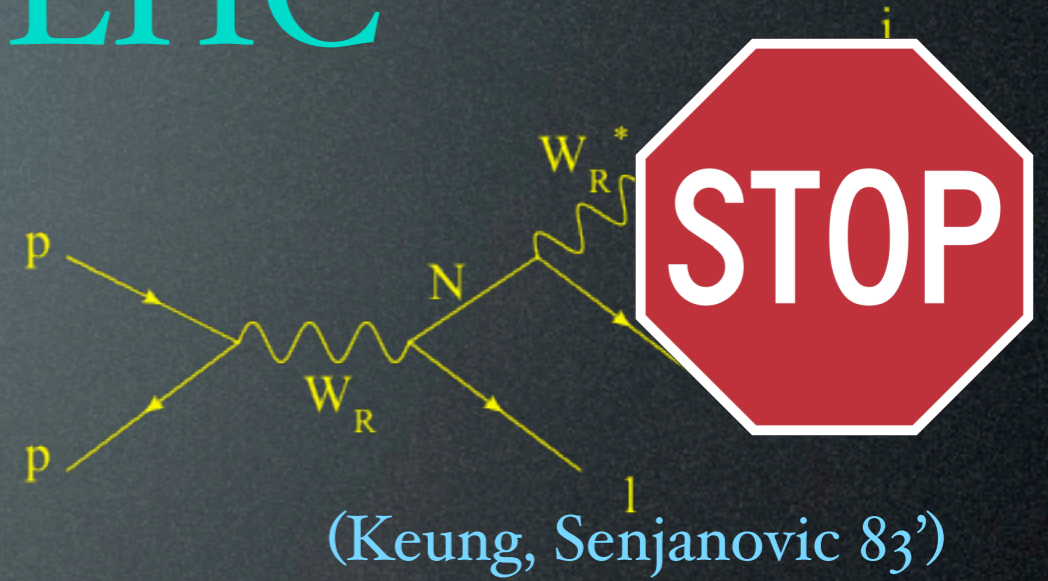
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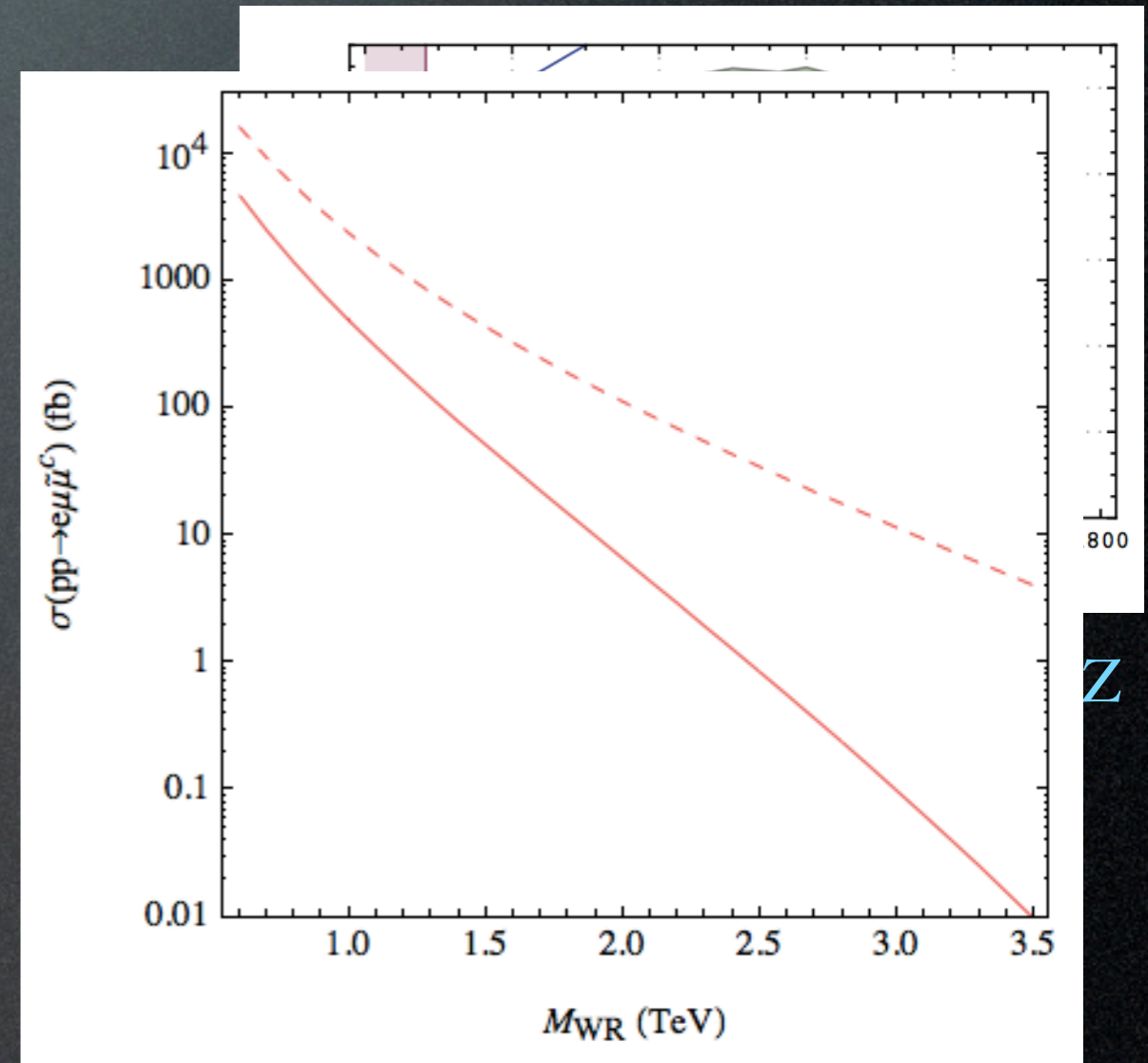
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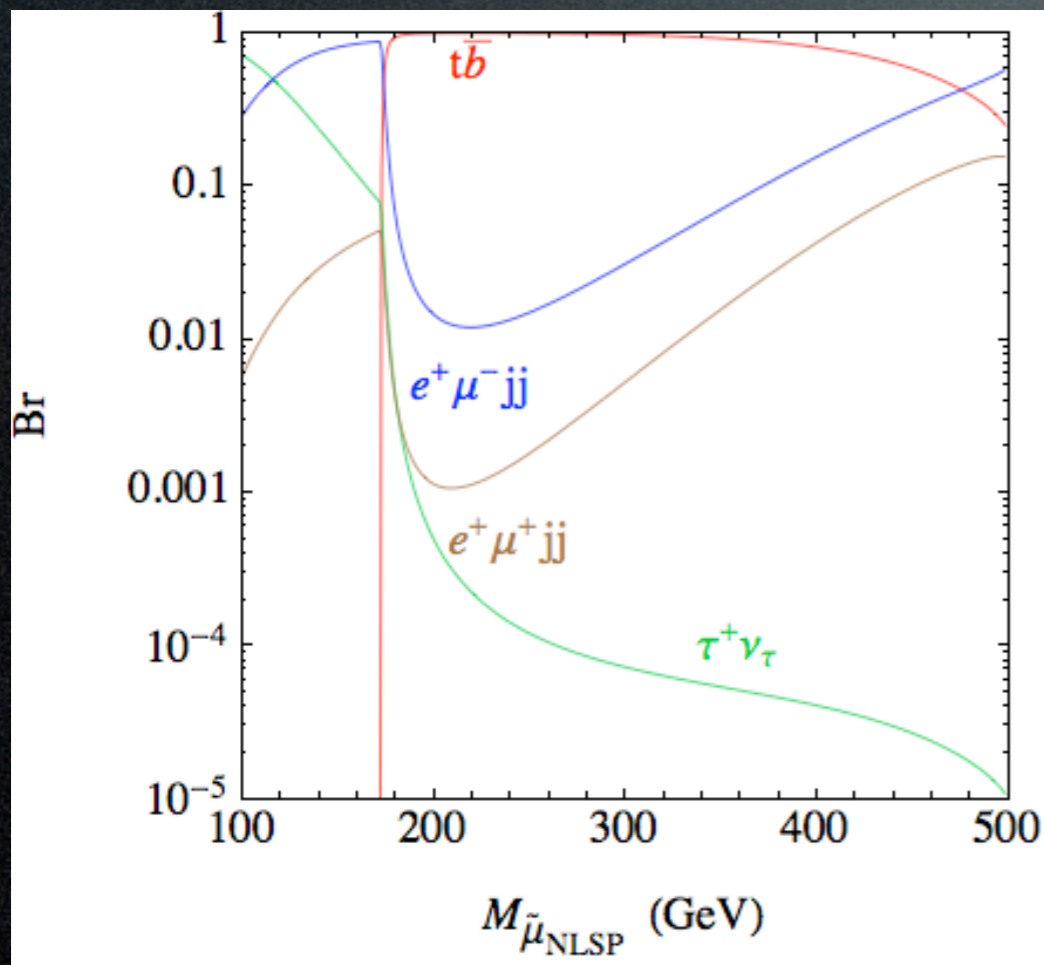
(Ghosh, Senjanovic, YZ 1010.3968)

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Slepton LSP decays

Large smuon mixing $\tilde{\mu}_{\text{LSP}} = \alpha\tilde{\mu}_L + \beta\tilde{\mu}_R, \quad \alpha, \beta \sim O(1)$



decays through:

$$\lambda' \bar{t} b \tilde{\mu}^c, \quad \lambda' \approx y_t y_\nu \langle \tilde{\nu}_e^c \rangle / \mu$$

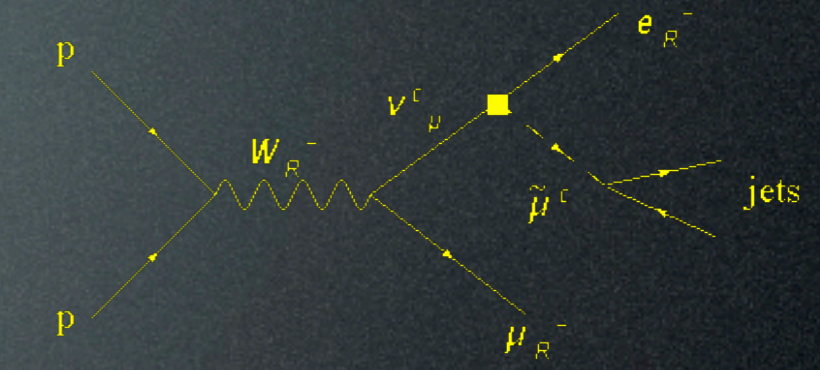
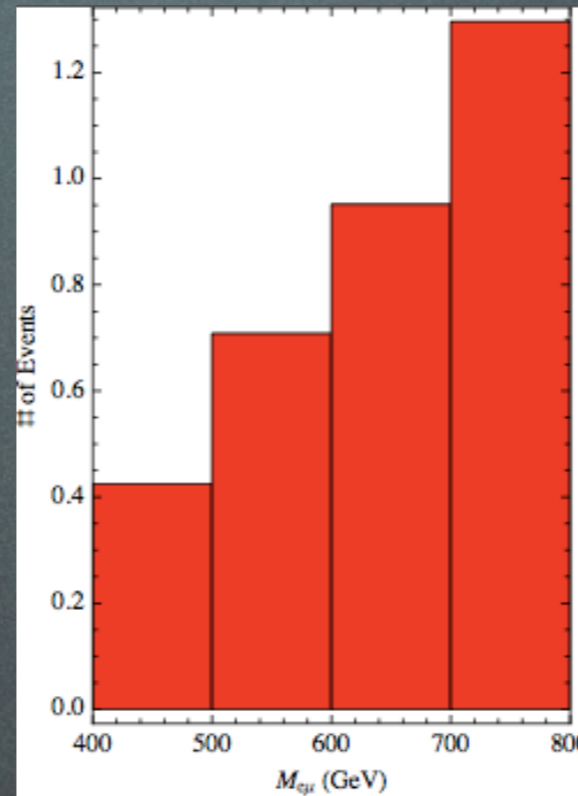
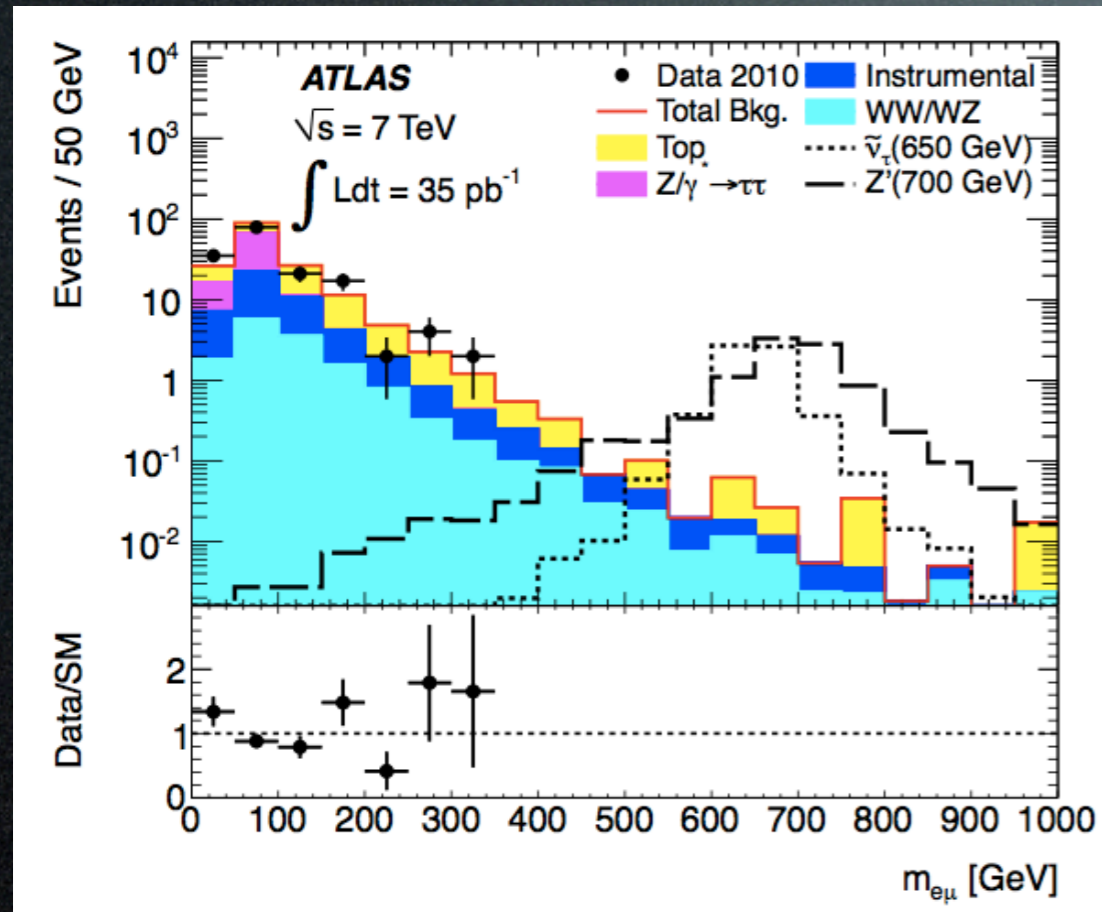
Signatures:

$$pp \rightarrow e^+ \mu^- \bar{t} b, e^+ \mu^+ t \bar{b} \\ \rightarrow e\mu + jets$$

Opposite sign e mu data now available from ATLAS !

(ATLAS collaboration, 1103,5559)

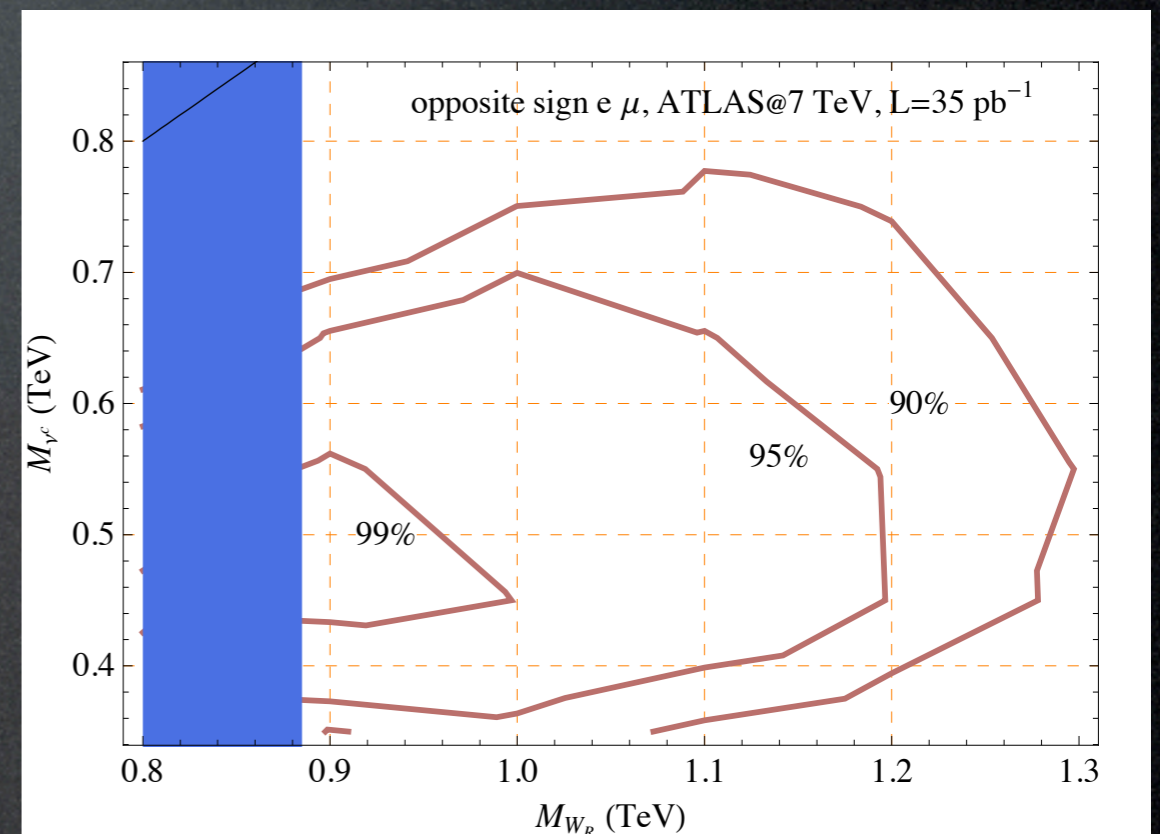
Early LHC limits on dR-p scale



Events in high invariant mass region $> 400 \text{ GeV}$

(YZ, to appear 2011)

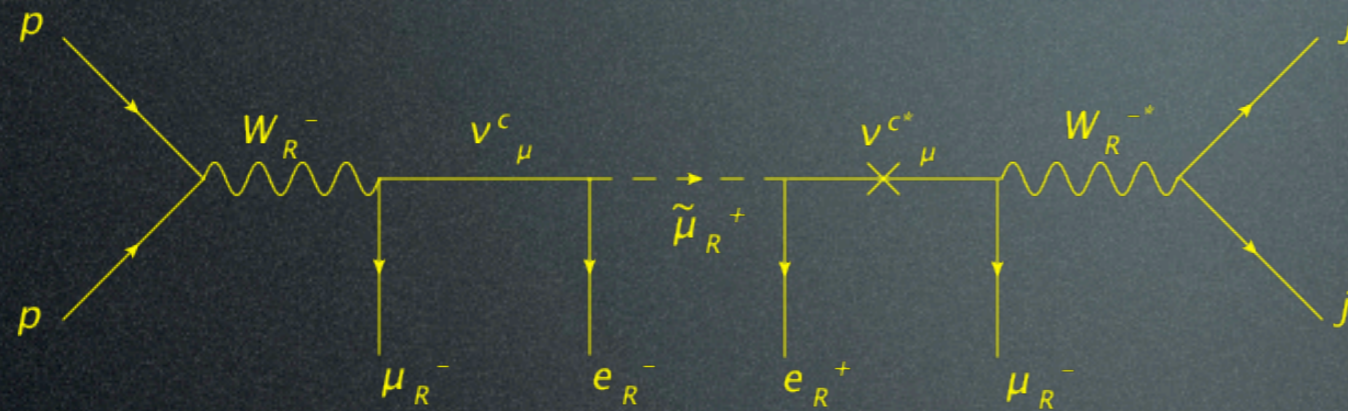
- ATLAS, 7 TeV, 35/pb
- Limit: $M_{W_R} > 1.2 \text{ TeV} @ 95\% \text{ C.L.}$
- Can reach 2-3 TeV for $L \sim \text{fb}^{-1}$



Mainly RH smuon LSP

$$\tilde{\mu}_{\text{NLSP}} \approx \tilde{\mu}^c$$

- dominantly single-produced if:



$$M_{\nu_e^c} > M_{W_R}$$

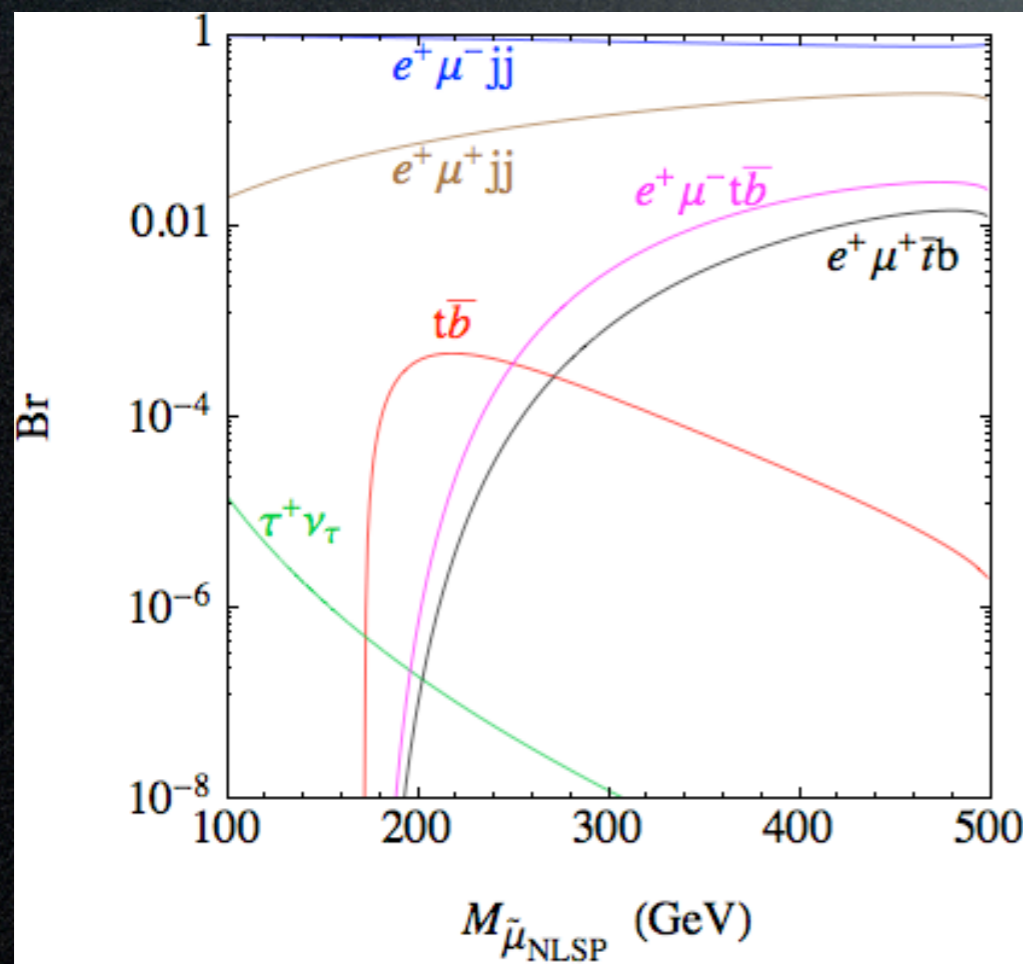
$$M_{\nu_\mu^c} < M_{W_R}$$

Signatures:

$$pp \rightarrow e^\pm e^\mp \mu^\pm \mu^\pm jj$$

- Multi-lepton final states.
- lepton number violating, no \cancel{E}_T .
- All heavy states reconstructable.

Almost background free



Conclusions

We study a class of susy models with R-parity broken spontaneously with extra gauge symmetries.

- Interesting predictions on neutrino mass.
- Cosmology implies the scale could be near TeV.
- A new class of R-parity breaking interactions.
- Distinct signatures at the LHC.

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Thanks! Hvala!