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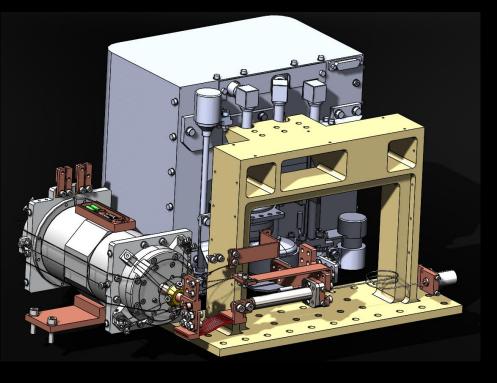


#### Forecast

- Cryogenic temperatures to observe far Universe
- We developed an **ADR** to cool detectors from 250 mK to **100 mK**
- Coupled to a 4He/3He sorption fridge through a 3He gas-gap Heat Switch

• ADR will be used in a **pathfinder telescope** for CMB-S4

- Cooling power at 100 mK: 2 μW for 48 hours
- 4 times more observation than cycling time





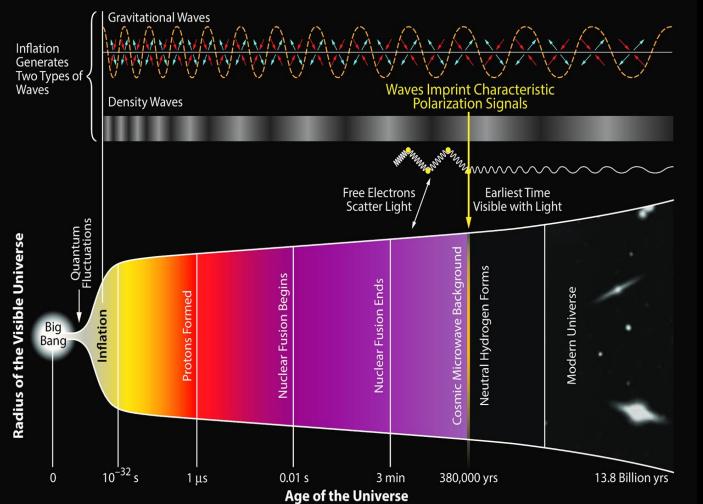


# Outline

- 1. Background
  - Observation of the Universe
  - 100 mK refrigeration
- 2. Material & Methods
  - Cryogenic chain presentation
  - ADR specificities
- 3. Results
  - Parasitic heat losses
  - Pill's characteristics
  - Parasitic magnetic field
- 4. Conclusion

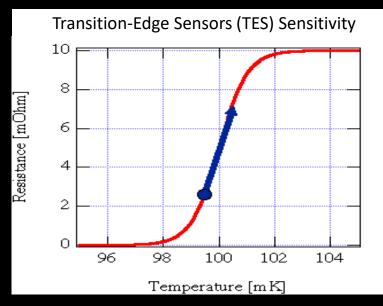
# Background: Observing our Universe





#### History of the Universe

- Looking at CMB frequencies
- Searching for a specific **polarization pattern**
- Could validate cosmic inflation

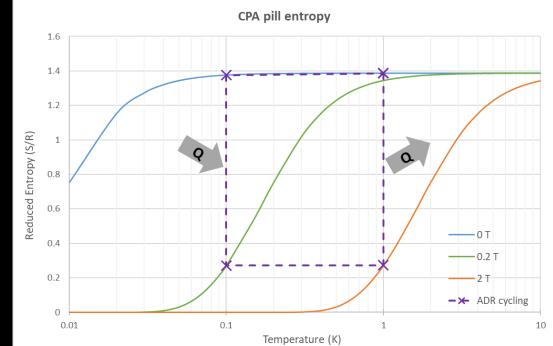


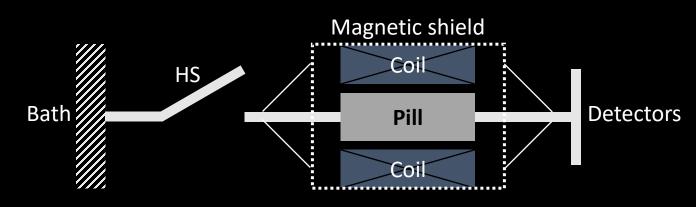
- Detectors must be cooled to 100 mK for 48 hours to increase their sensitivity
- Installed in dry places like the South Pole

# Background : 100 mK Refrigeration



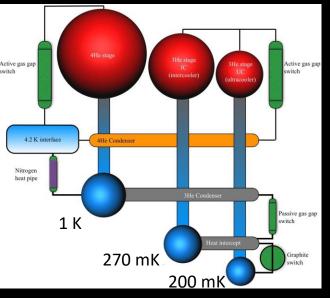
- Cosmic Microwave Background Stage 4 (CMB-S4) is the largest ground-based CMB experiment
- Existing cryostat: 250 mK sorption fridge
- Pathfinder: compact 100 mK ADR
- <u>Same telescope</u>, new detectors
- Adiabatic Demagnetization Refrigerator (ADR):
  - ✓ Paramagnetic pill
  - ✓ Strong magnetic field
  - ✓ Heat dissipated to a thermal bath
  - ✓ Cold finger **thermally isolatable**



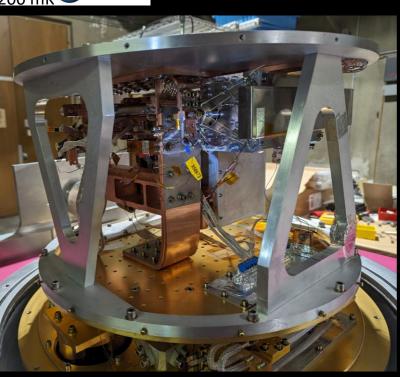


# Material and Methods

- Cryomech PT-415 cooling 2 radiative shields
- 3-stages (4He/3He/3He) sorption fridge
- ADR connected to its <u>biggest 3He</u> evaporator
- ADR specificities:
  - ✓ 226 g **CPA** pill
  - ✓ 16 **doubly-intercepted** Kevlar lines
  - ✓ 3He HS with **deported pump**
  - ✓ Superconducting coil reach **1.3 T** w/ 6 A
  - ✓ Ferromagnetic magnetic shield
  - ✓ Recycled **during** sorption fridge cycle
  - ✓ 12 hours total cycling for 48 hours observations



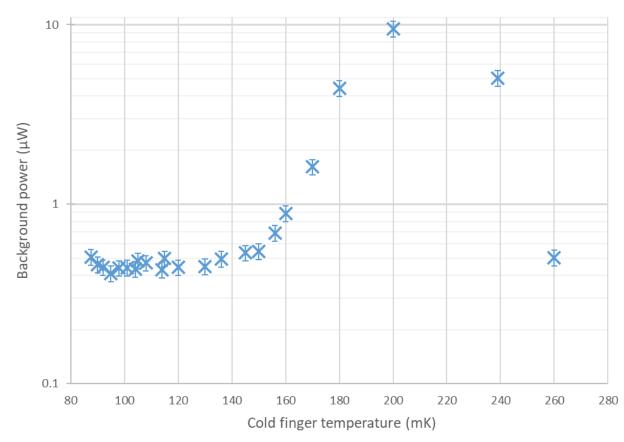




#### Results: Parasitic Heat Losses

- Background power = Radiation + Conduction
- Radiation: minimized thanks to Carbon-loaded Stycast coating
- Conduction: estimated to be
  **0.3 μW** (HS + Kevlar lines)

- Around  $0.45~\mu W$  at 100 mK
- Increase above 150 mK due to liquid 3He in the HS

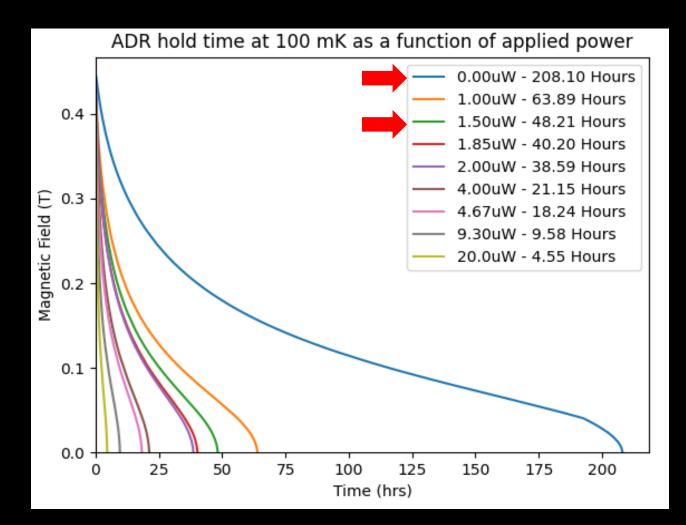






## **Results: Pill's Characterization**

- Several hold time runs at 100 mK with ≠ applied power
- Reference hold time t<sub>0</sub> = 208 hours (4He stage ran out before ADR)
- Enthalpy  $H = t_0 \cdot P_0 = t_1 \cdot (P_1 + P_0)$  $P_0$ : background power
- Pill's enthalpy at 100 mK: 0.34 J
- Cooling power for 48 hours:  $1.5 + 0.45 \approx 2 \mu W$

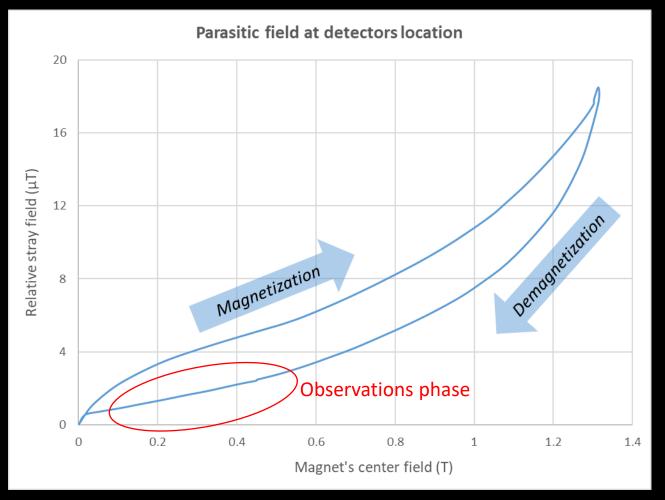




## **Results: Parasitic Magnetic Field**

- ADR needs **1.3 T** to be recycled
- Detectors and amplification circuit extremely sensitive to magnetic field
- Thermally treated ferromagnetic magnetic shield

- **Relative** parasitic field (detectors location)
  - 18 μT maximum during ADR cycling
  - ✓ Below 2.5  $\mu$ T during observations





# Conclusion

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- ADR backed by a 4He/3He/3He sorption fridge cools CMB detectors to 100 mK
- Ratio cycling/observations: 4
- Cooling power at 100 mK: **2 μW for 48 hours** (0.34 J)
- Background power at 100mK: 0.45 μW
- Parasitic field during observations: **2.5**  $\mu$ **T**
- ADR will transform an existing telescope into a CMB-S4 pathfinder with <u>new detectors</u>
- Larger scale telescope may use a **Dilution Fridge**
- More results in the paper!