



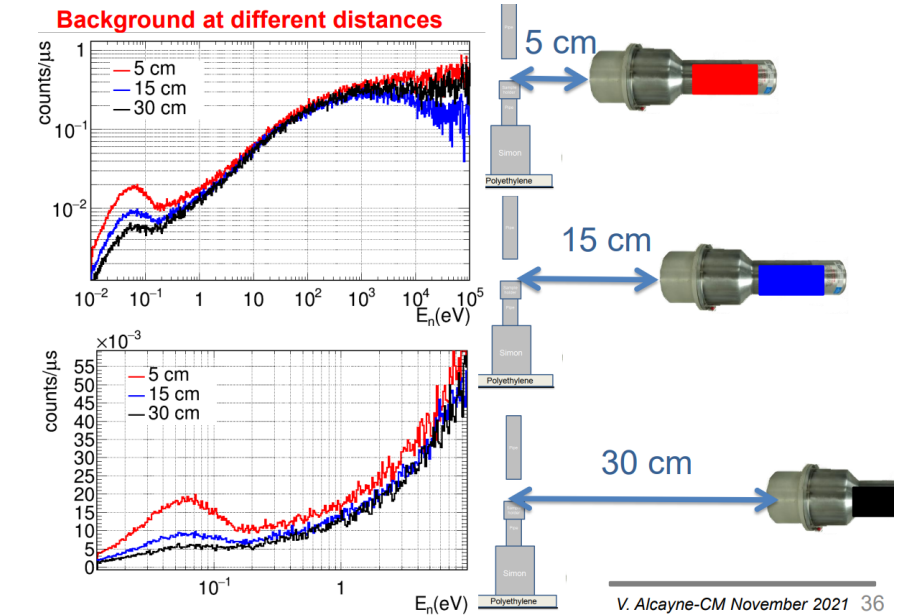
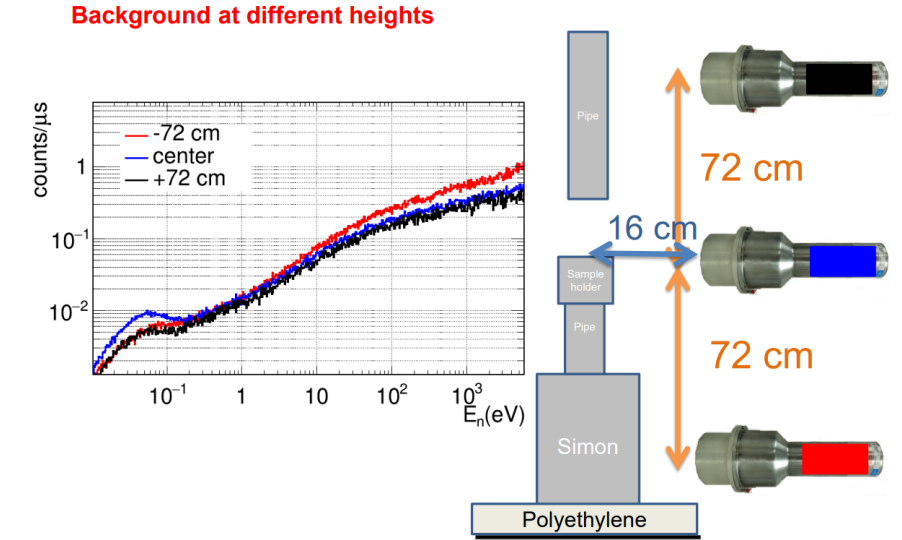
# Signal to Background study for capture measurements in EAR2

V. Alcayne, M. Bacak, A. Casanovas, A. Cintorra, J. Lerendegui-Marco

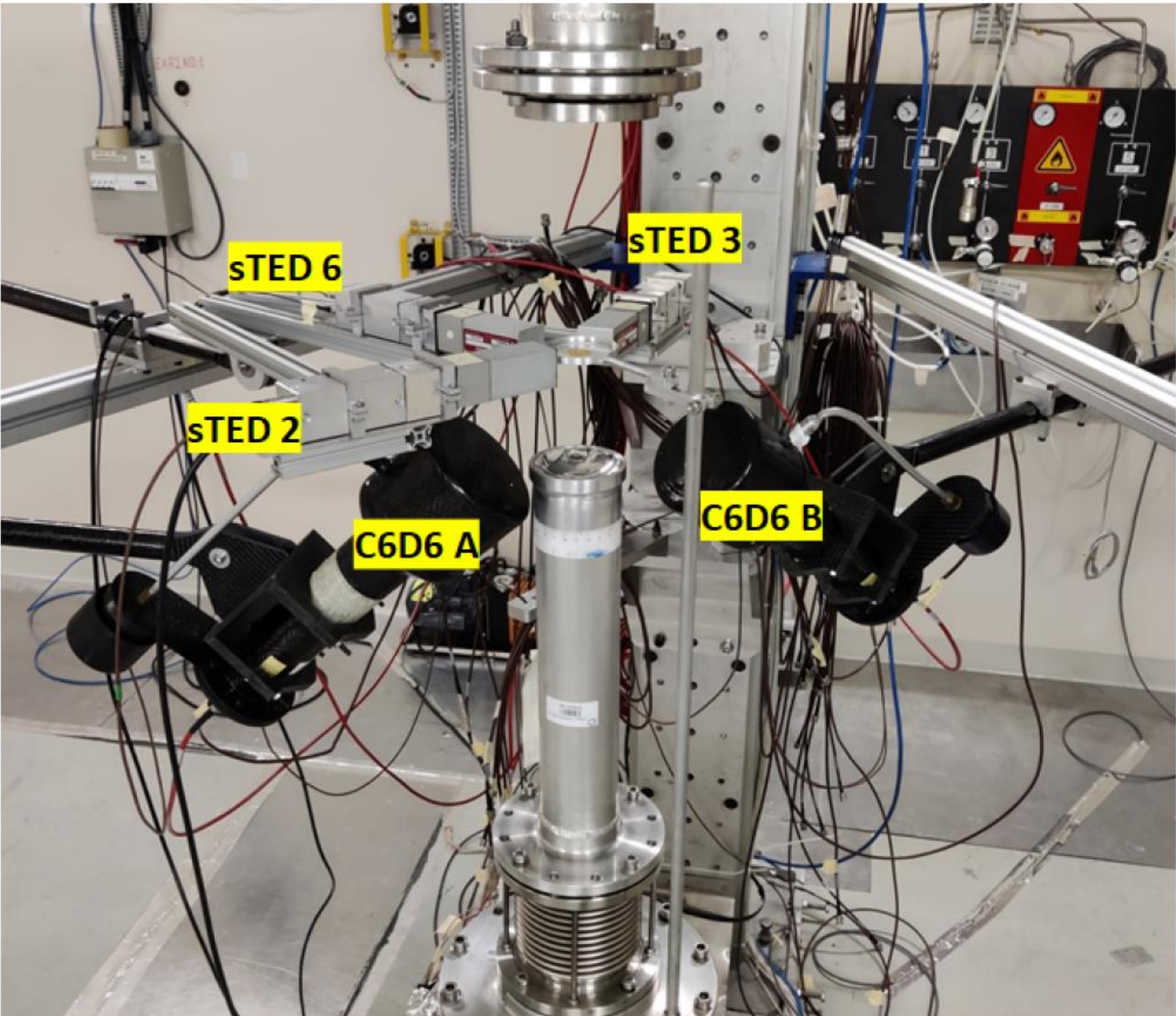
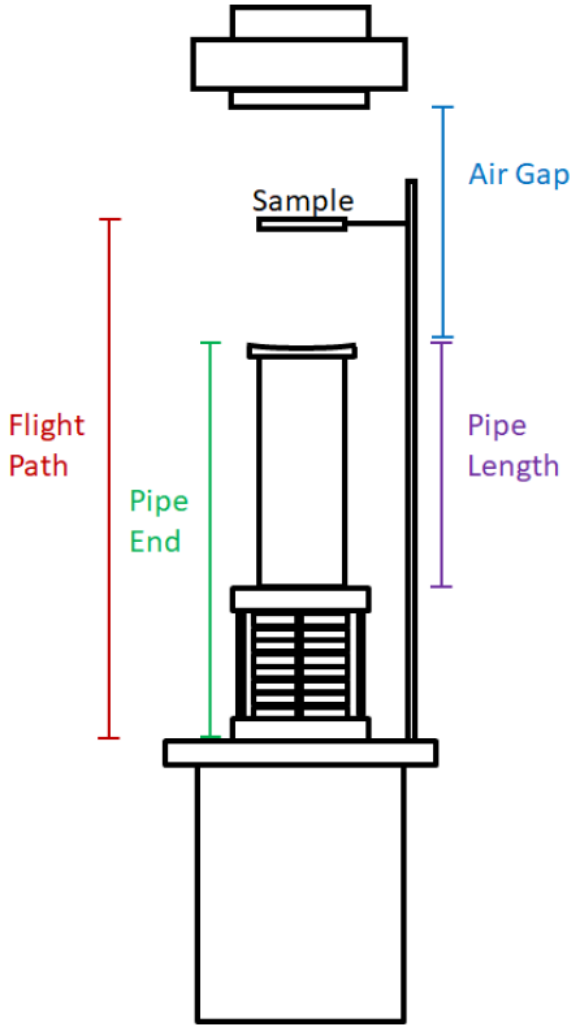
14.12.2022 – n\_TOF Collaboration Meeting

# Introduction & Motivation

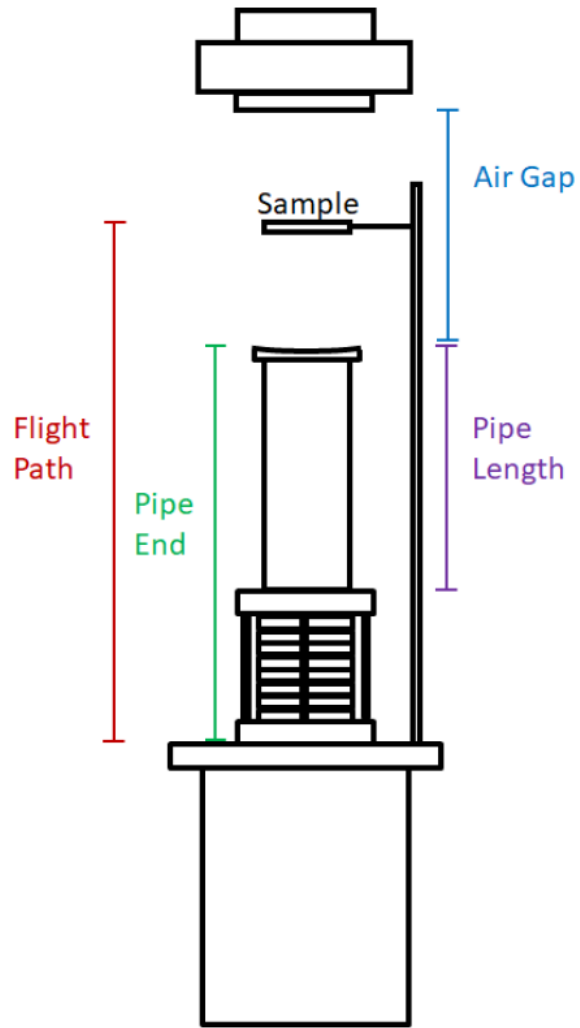
- Triggered by and continuing investigations from:
  - Victor (2021) – [background study of the effect of different distances with B/L6D6](#)
  - César (2022) – tests with  $\text{LaBr}_3$  and  $\text{LaCl}_3$  at different distances and impact on the peak2valley ratio of resonances
- **“What’s the ideal capture position in EAR2 wrt SBR?”**
  - Sample positioning (flight path / BIF)
  - Detector positioning (distance to beam center)
- **Setup:**
  - Detectors at different distances and flight paths
  - Samples at different flight paths



# Exp. Setup



# Exp. Setup and campaign

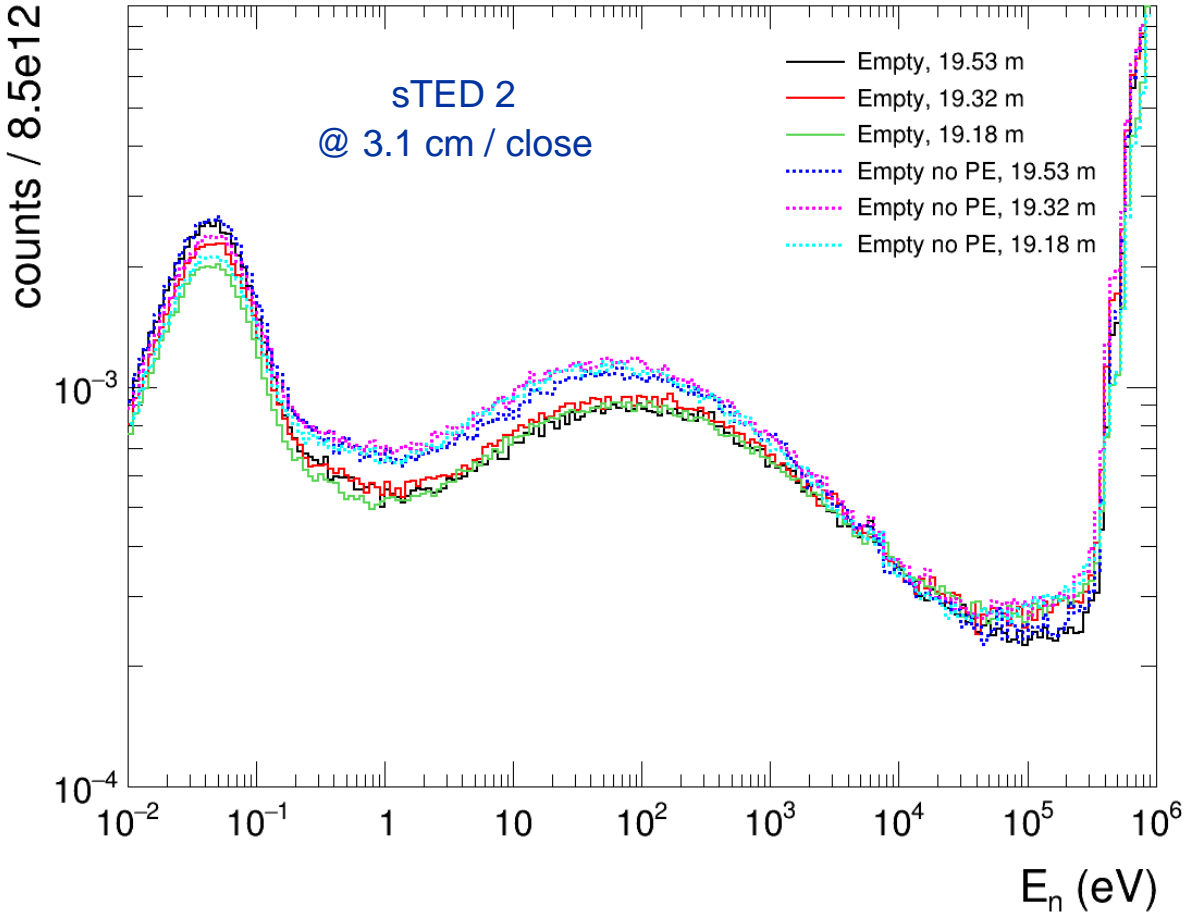
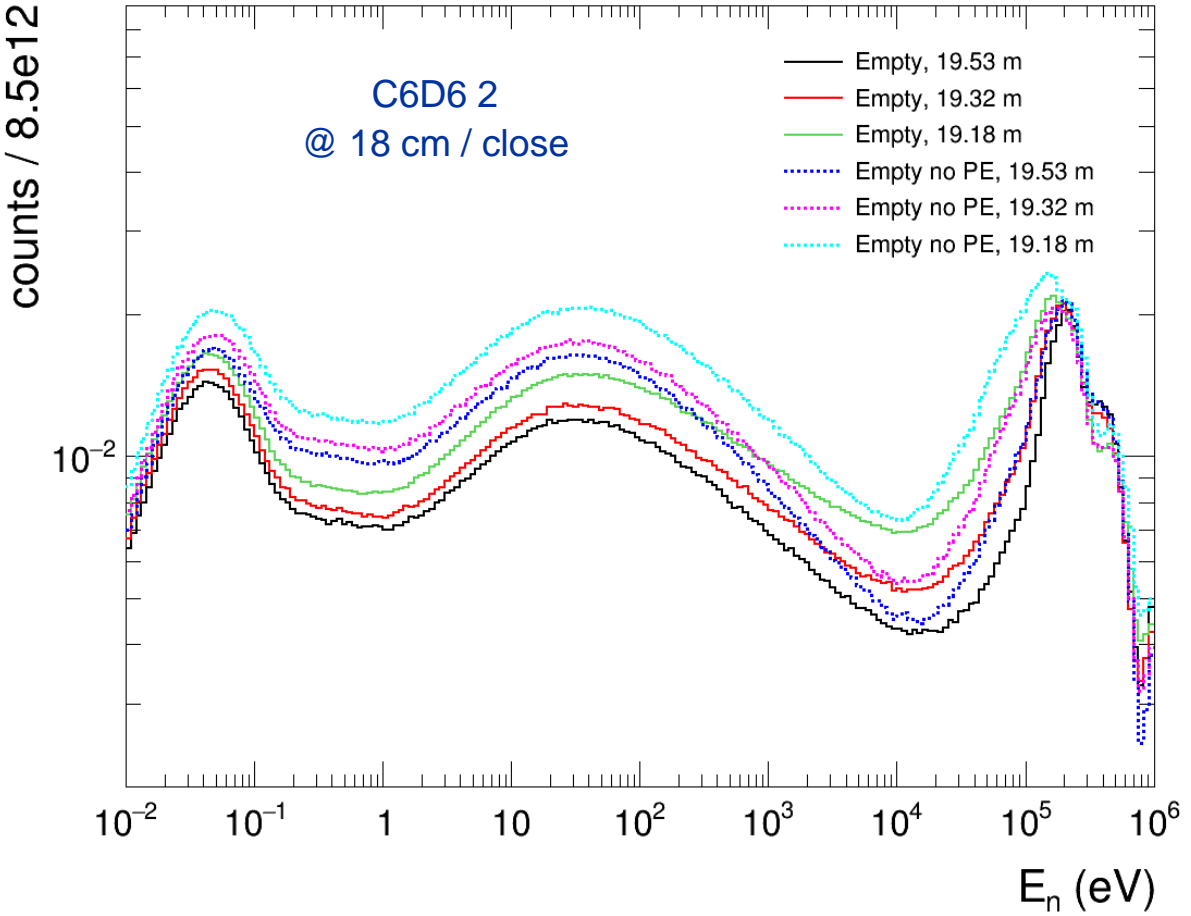


	Flight Path	Pipe End	Pipe Length	Air Gap	$L_0$
<b>FP1</b>	65.5 cm	56.5 cm	38.0 cm	26.5 cm	19.525 m
<b>FP2</b>	44.5 cm	35.0 cm	17.0 cm	33.0 cm	19.315 m
<b>FP3</b>	30.5 cm	20.5 cm	17.0 cm	48.0 cm	19.175 m

Detector	Distance	ID	Source	Mass	Thickness
C6D6 A	24.0 cm	689	$^{197}\text{Au}$ (20 mm)	645 mg	100 $\mu\text{m}$
C6D6 B	18.0 cm	1189	$^{197}\text{Au}$ (10 mm)	-	102 $\mu\text{m}$
sTED 2	3.1 cm	106	C (nat) (20 mm)	2.646 g	5 mm
sTED 3	4.0 cm	1209	Fe (20 mm)	5.143 mg	2.1 mm
sTED 6	5.0 cm	184	Pb (nat) (20 mm)	7.281 mg	2.1 mm

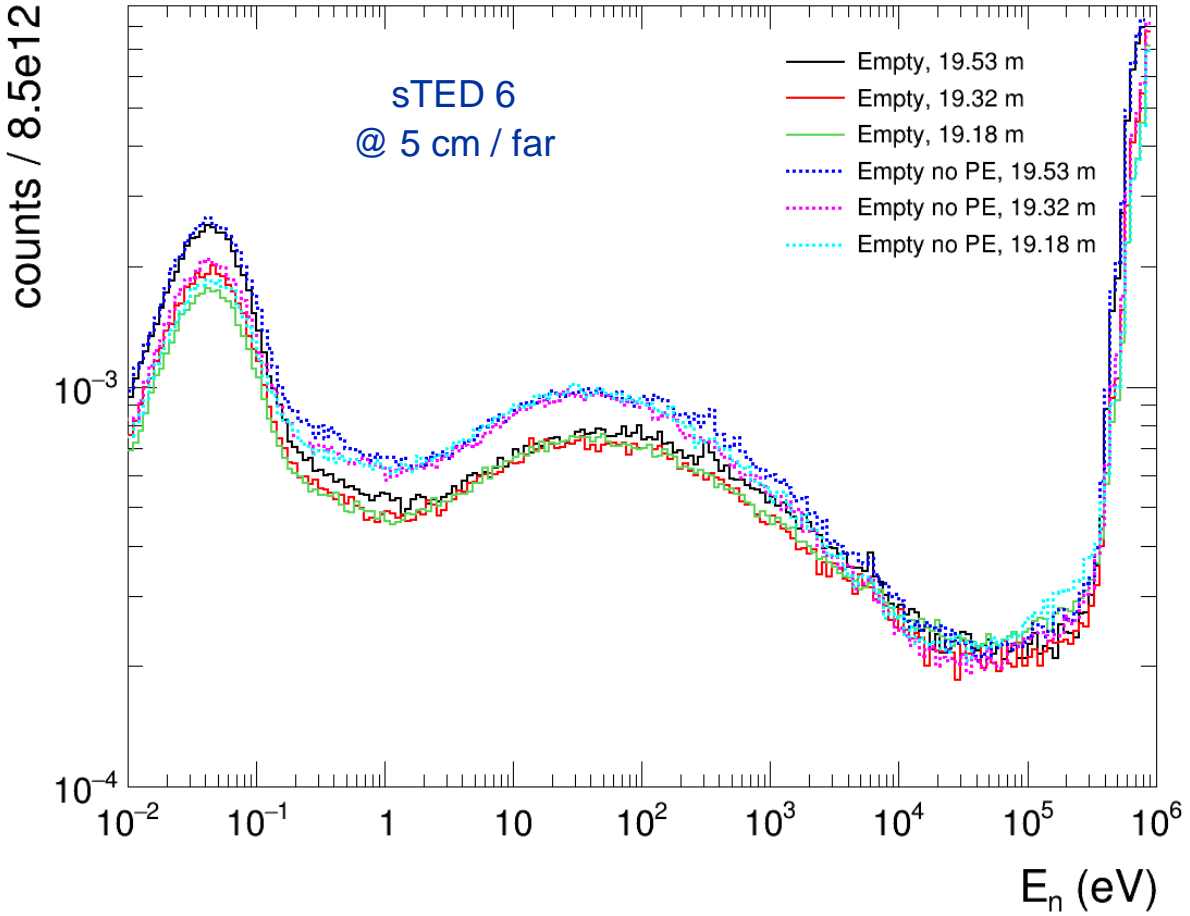
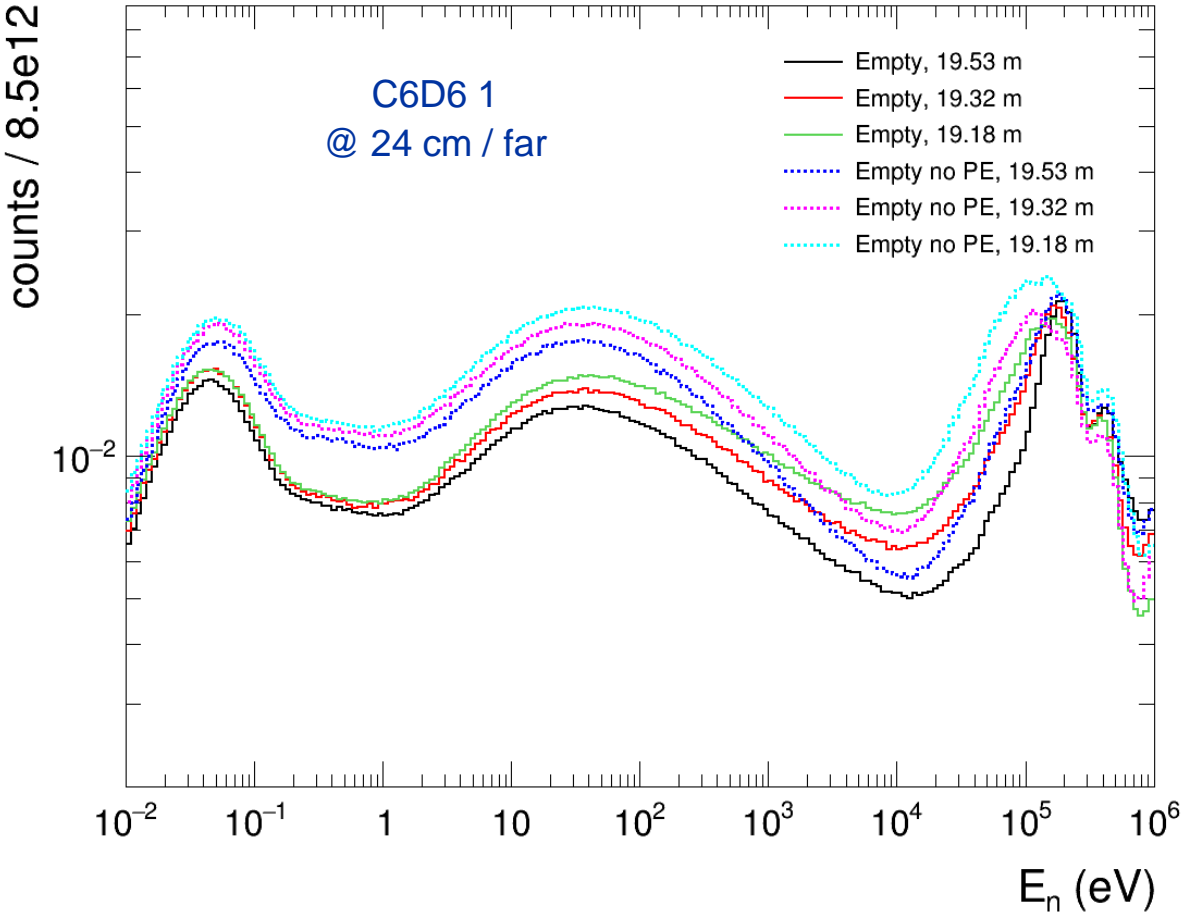
+ empty (with and without PE floor)

# Empties PE-noPE (no normalization required)



noPE → more background irrespective of detector size/position  
Background dependence on flight path for C6D6 (seems stable in STED position range)

# Empties PE-noPE (no normalization required) II

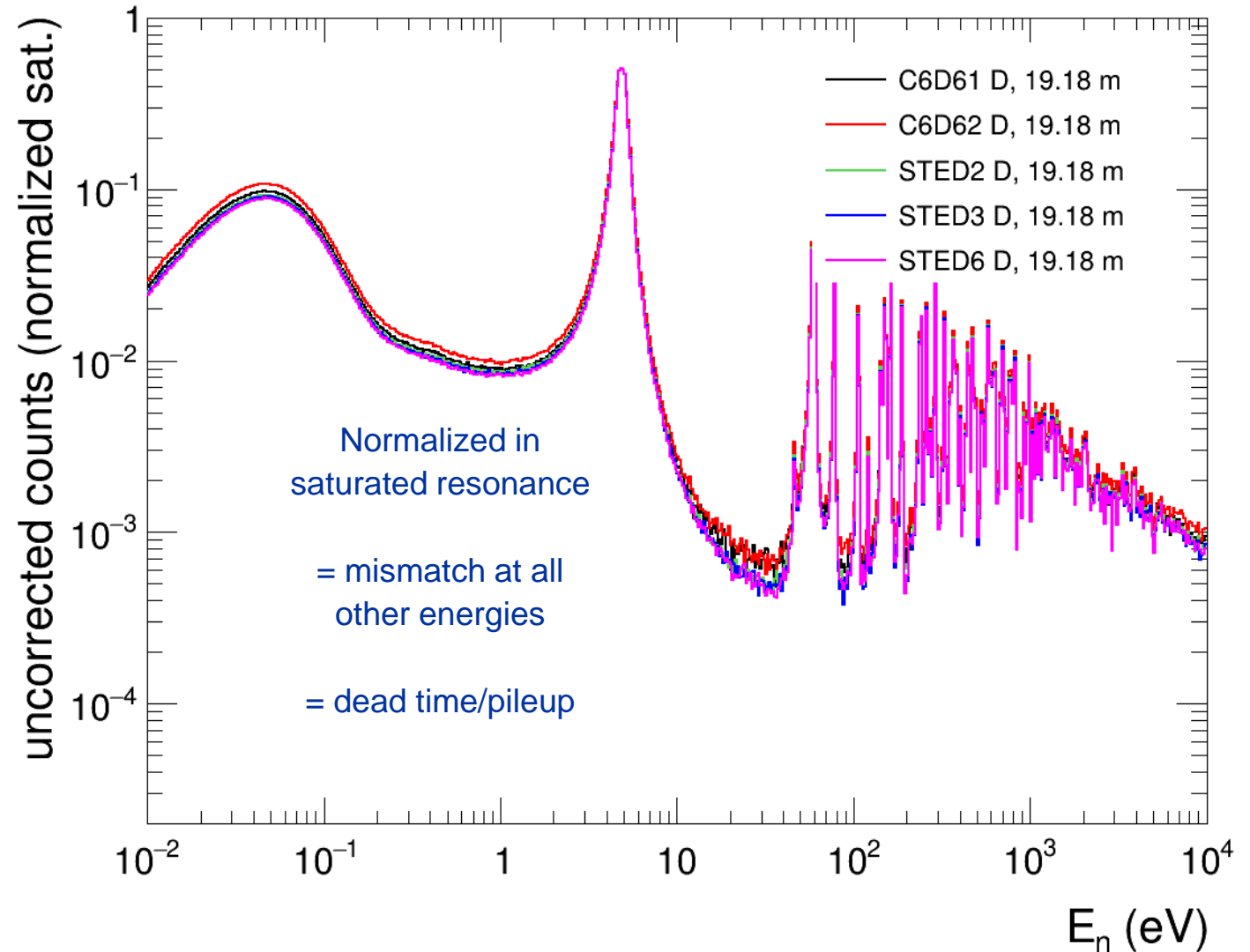


noPE → more background irrespective of detector size/position  
 Background dependence on flight path for C6D6 (seems stable in STED position range)

# Important: normalization with Au to correct for BIF

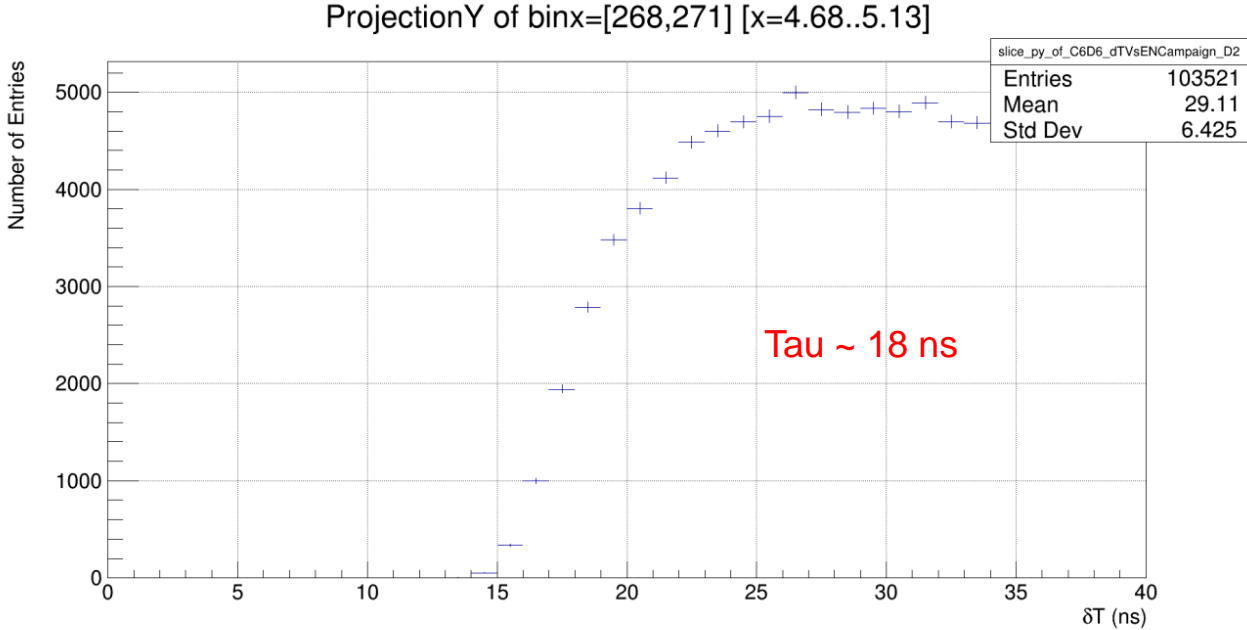
- **Problem: Count rates of up to 10 counts / us:**
  - Important corrections for the normalization otherwise direct comparison is invalid
- **Dead time model:**

$$CF(E_n) = \frac{1}{1 - \tau \cdot CR(TOF)}$$

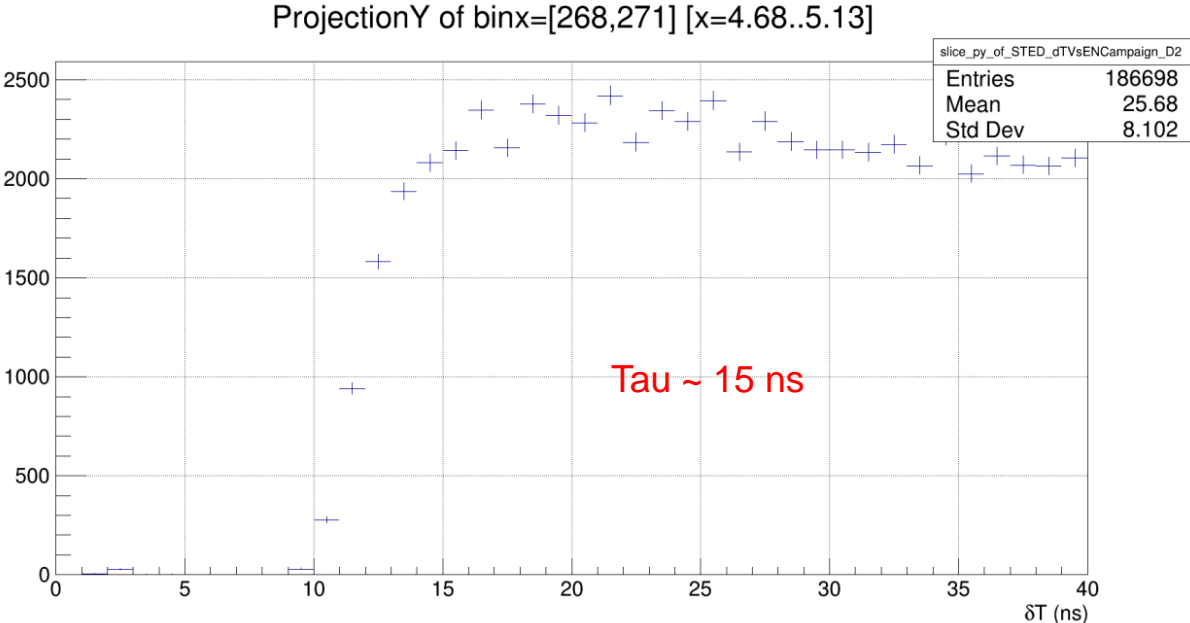


# Pile-Up/Deadtime

$$CF(E_n) = \frac{1}{1 - \tau \cdot CR(TOF)}$$



((b)) C6D6B

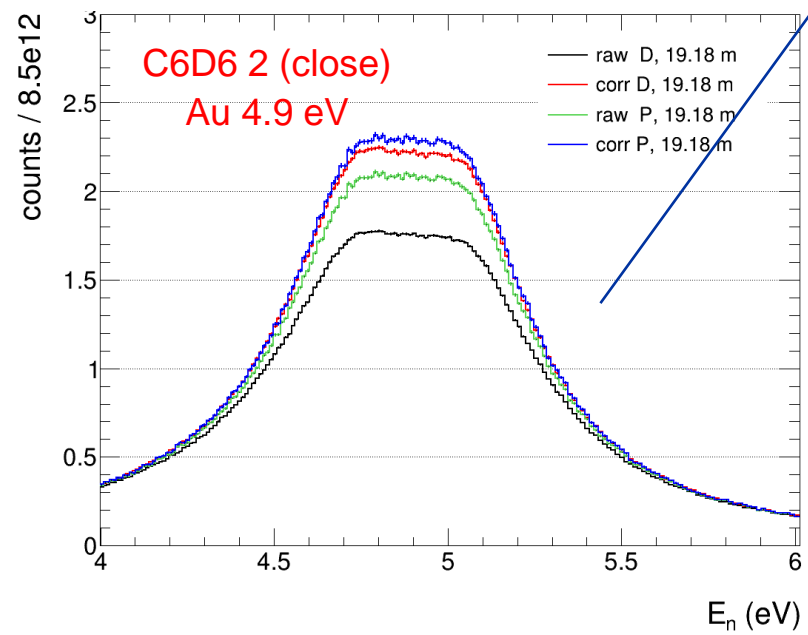
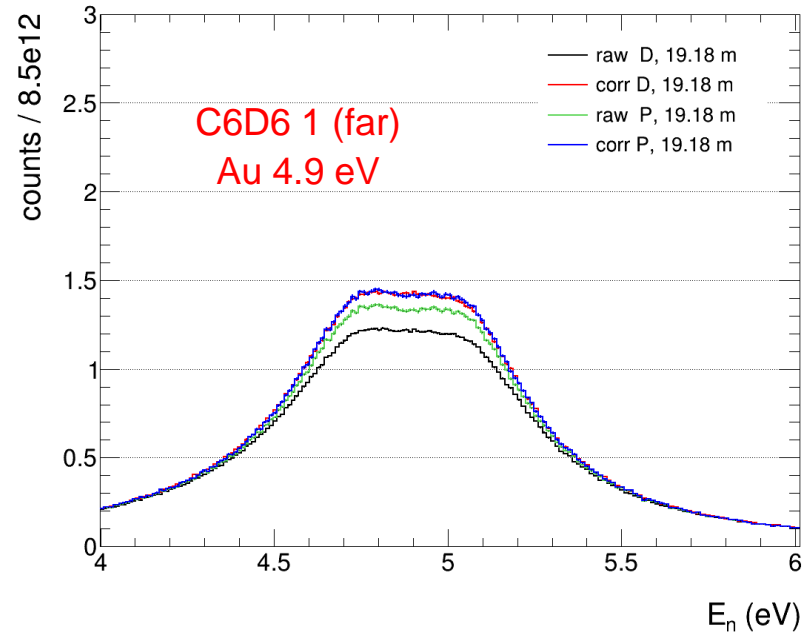
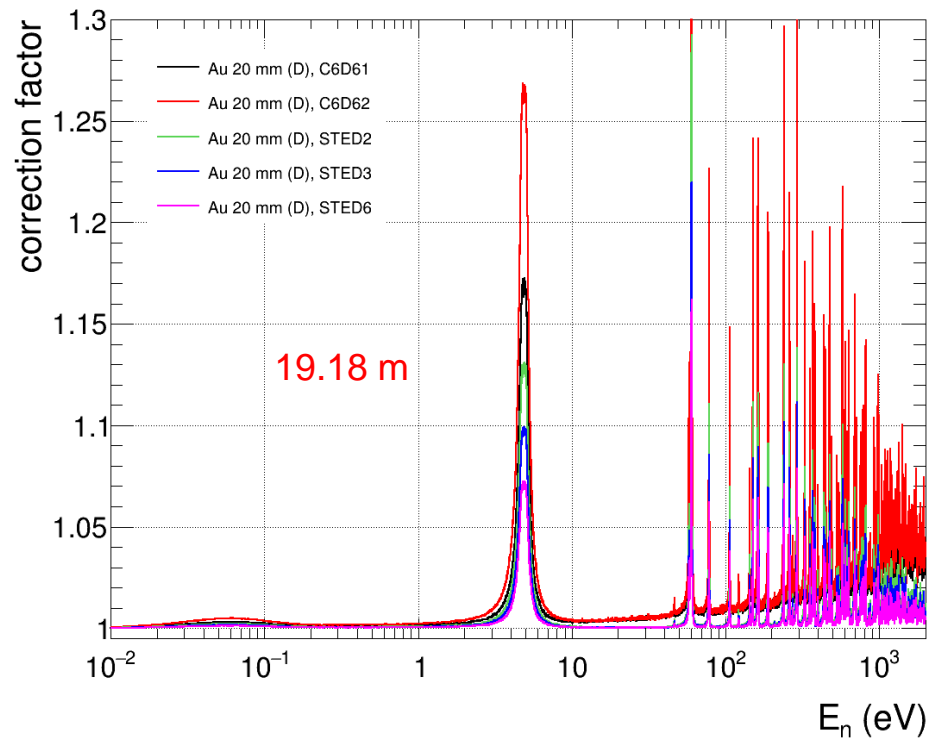


((c)) sTED2

This is an easy-to-implement technique but an approximation potentially not sufficient for a physics measurement but it works well enough for a first comparison

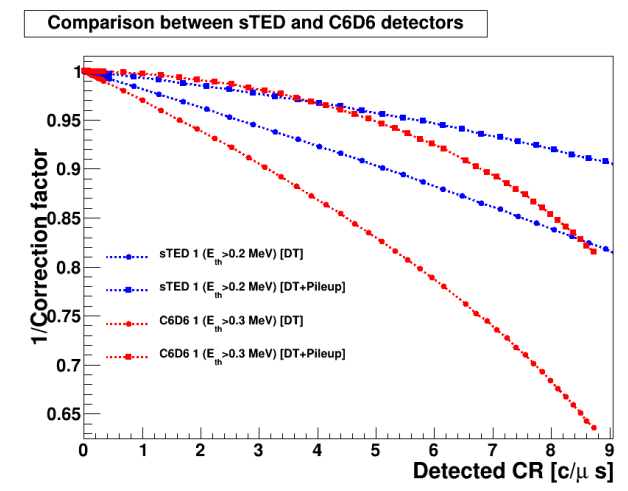


# Pile-Up/Deadtime



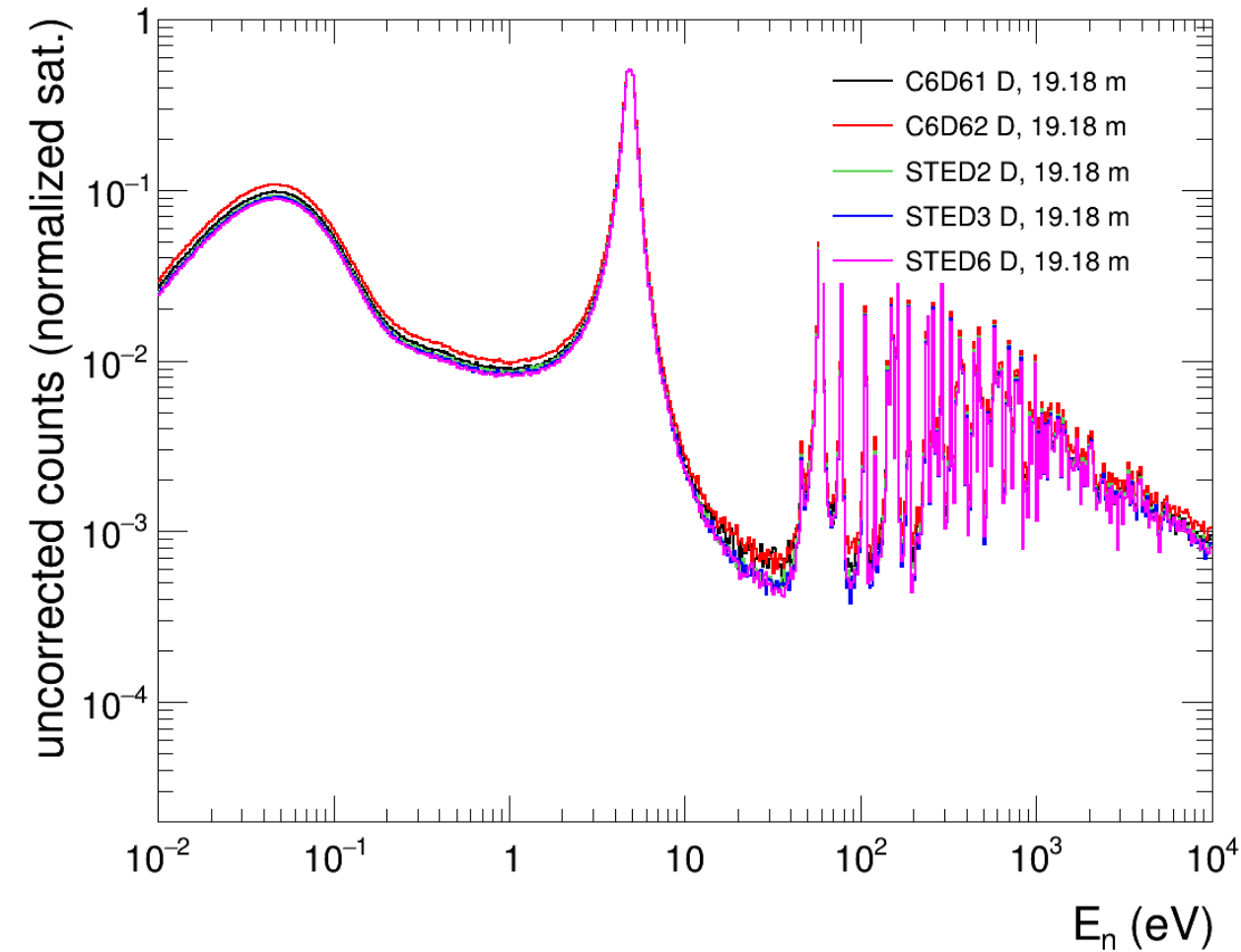
not perfect because not only dead time but also pile-up

See J. Balibrea's <sup>94</sup>Nb talk for a more precise correction

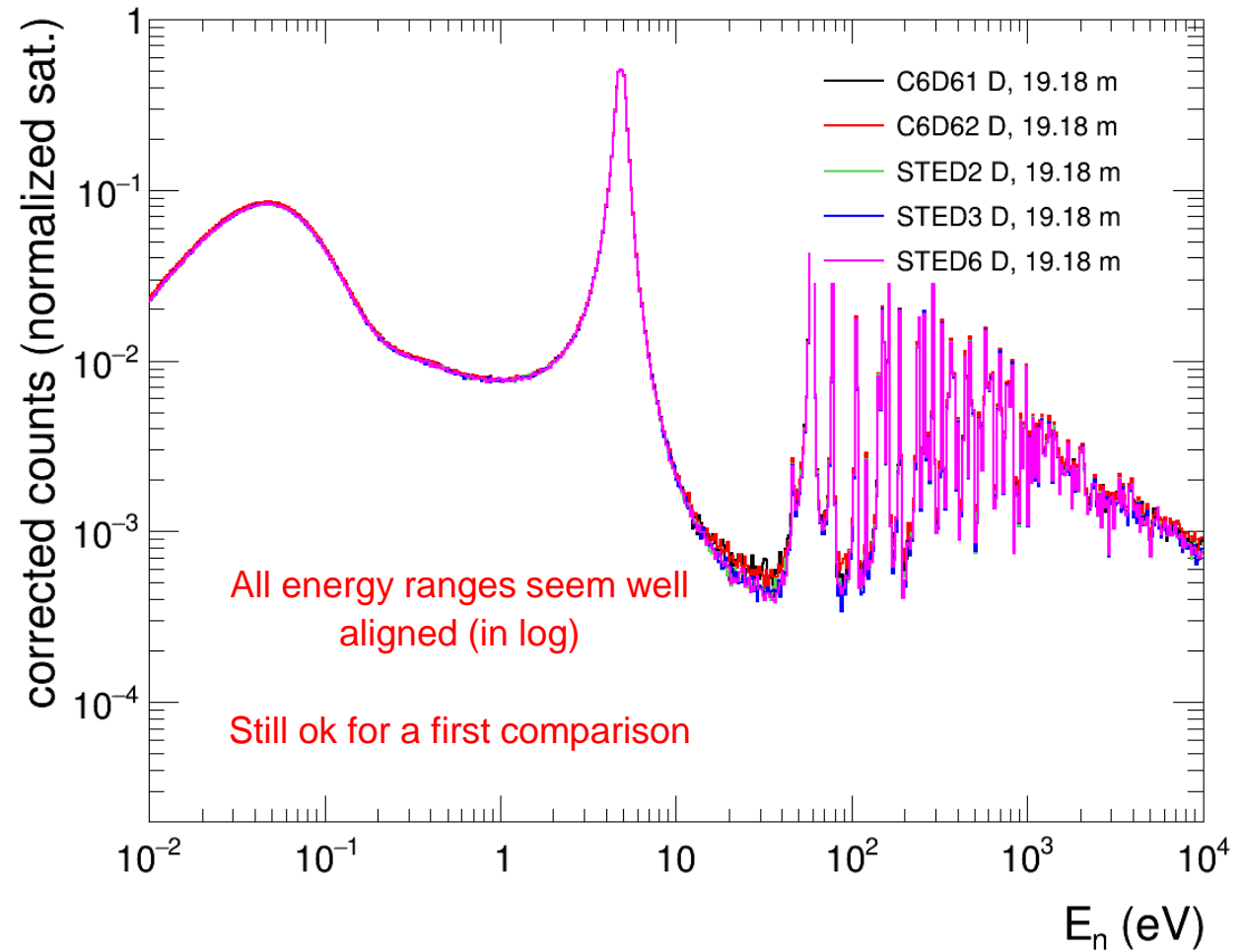


# Au before/after DT correctiong & empty subtraction

Before DT correction



After DT correction

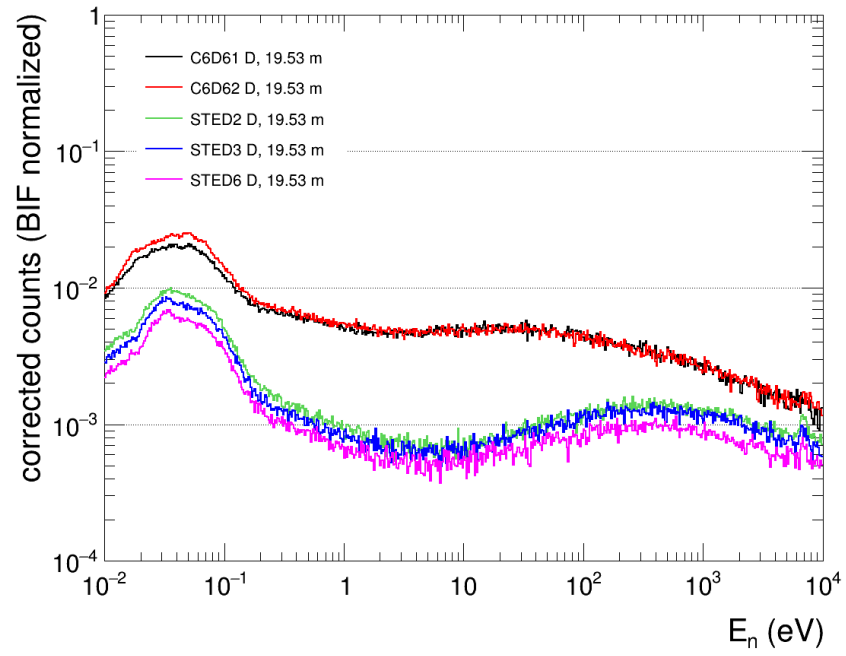
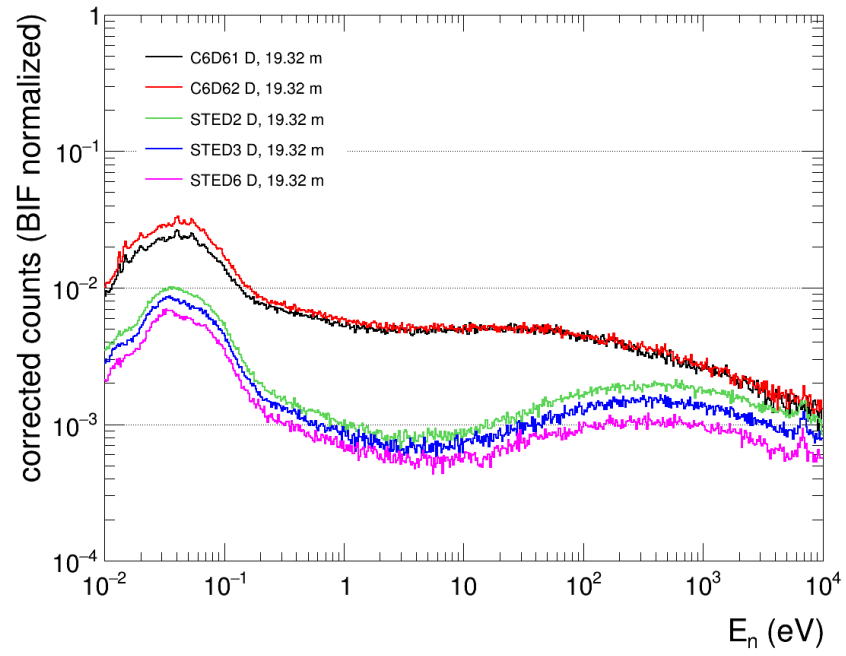
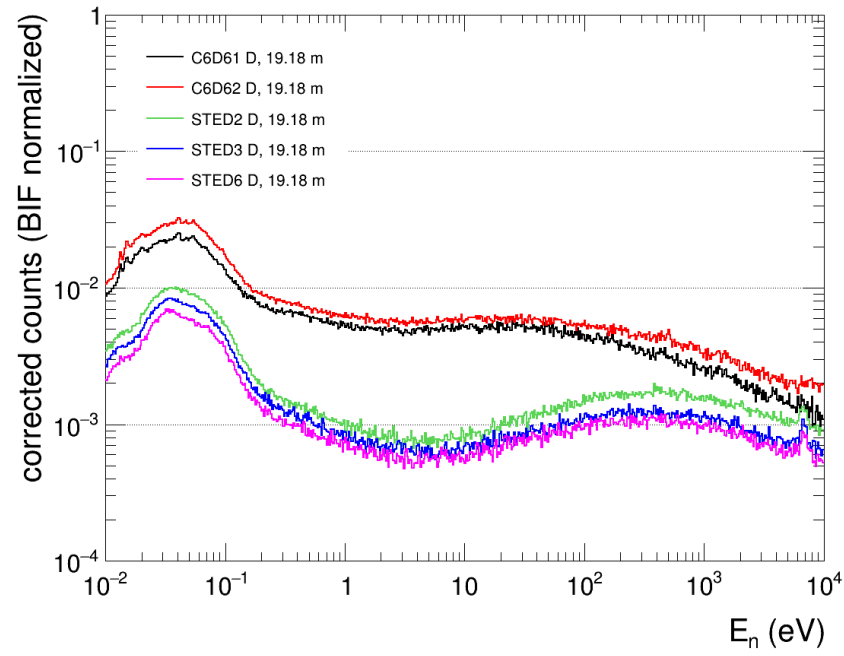


# Extracted normalizations / relative BIFs

	Au integral (per 8.5e12) @ flight path		
	19.53	19.32	19.18
C6D6 1 (A)	2.27	2.63	2.84
C6D6 2 (B)	3.30	3.87	4.44
STED 2	2.09	2.64	2.79
STED 3	1.66	2.01	2.11
STED 6	1.18	1.41	1.53
	Ratio to max flight path (relative BIF)		
		1.16	1.25
		1.17	1.35
		1.26	1.33
		1.22	1.27
		1.19	1.30
<b>Mean</b>		<b>1.20</b>	<b>1.30</b>

Normalize the 19.32 m and 19.18 m flight paths to the higher BIF

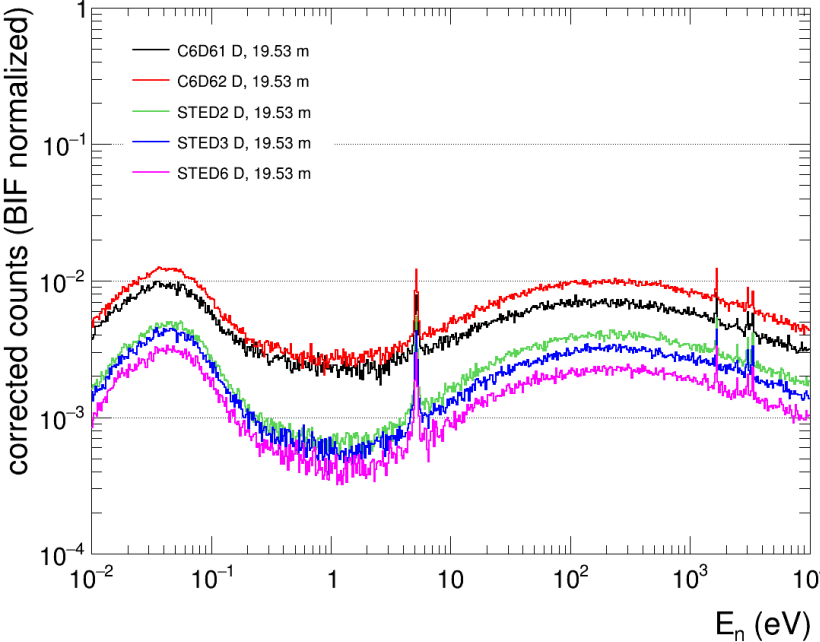
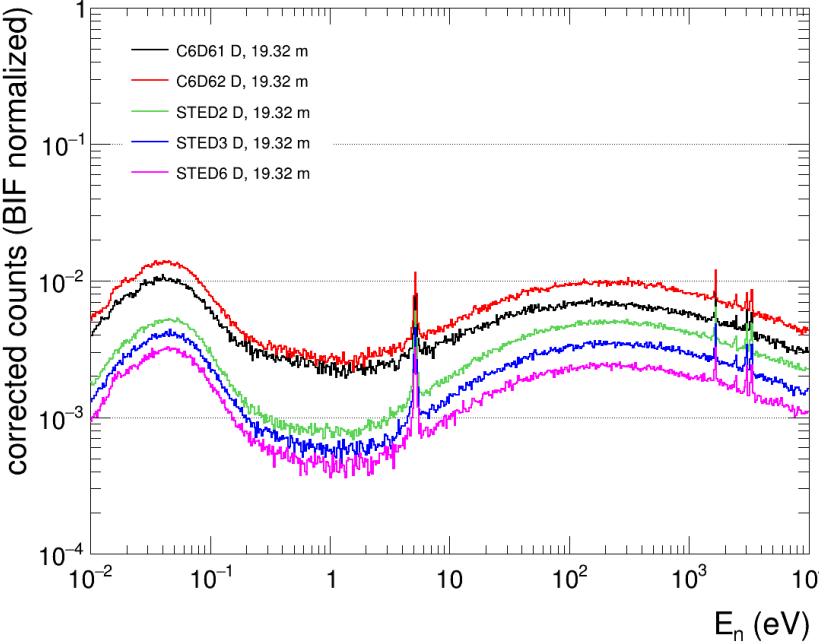
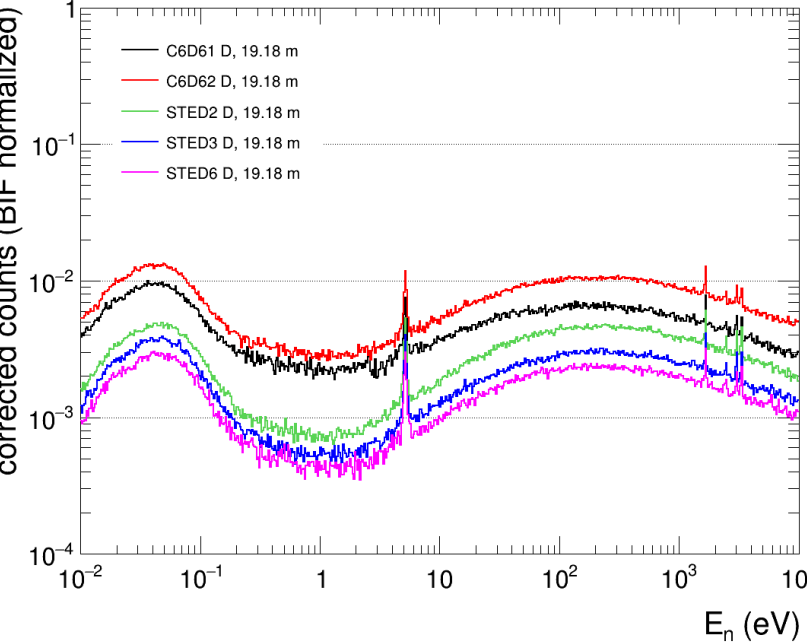
# Cnat BIF normalized – neutron scattering



Detectors relative to each other change but overall trend seems stable.

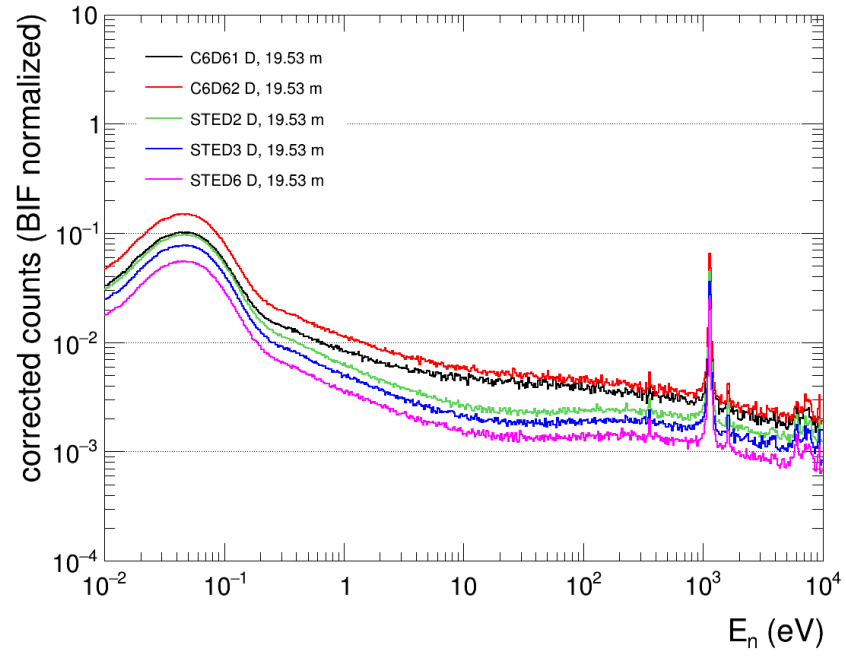
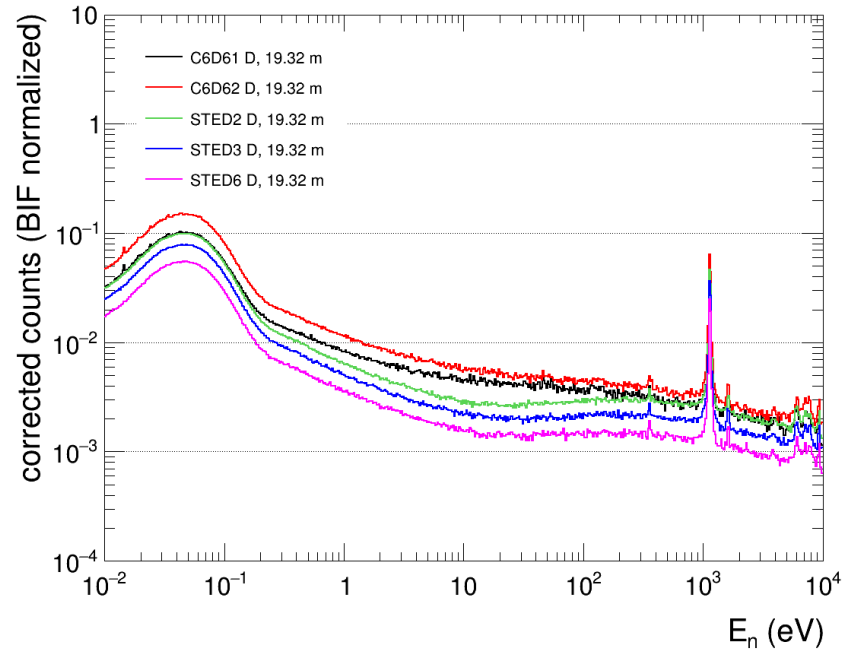
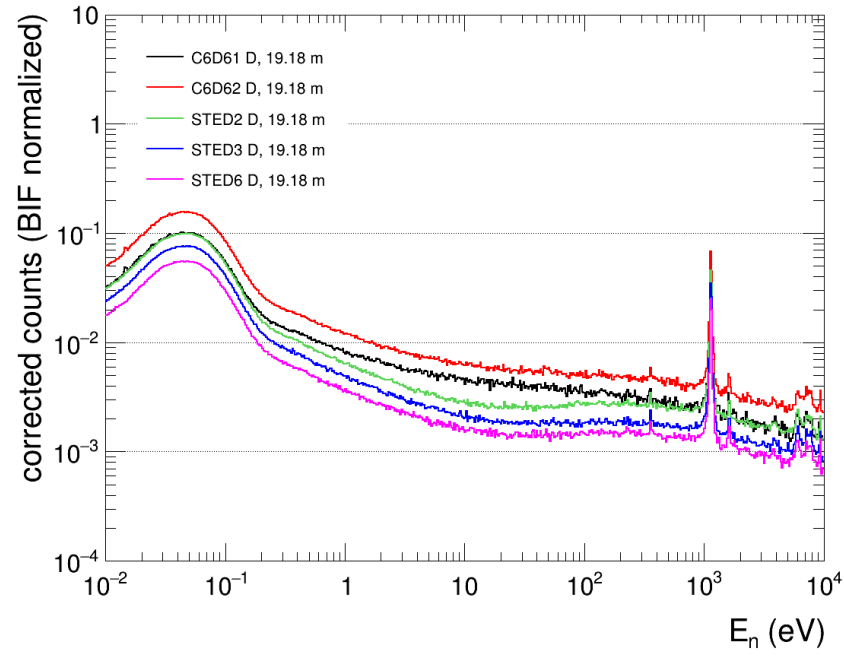
Shape at thermal looks weird – no idea (bragg edges of the carbon crystal?)

# Pb BIF normalized – gamma scattering



Slightly less counts at longer flight paths  
(usual Ag contaminant)

# Fe BIF normalized – neutron and gamma scattering



Slightly less counts at longer flight paths

# Summary & outlook

- **Three weeks of data taking:**
  - 2 setup swaps to different flight paths → thanks to everybody who helped making this as efficient and reproducible as possible
  - Thanks to Oscar for the needed modifications of the sample holder to make it flexible
  - Thanks to our summer student Alan Cintorra (via Carlos) for his interest, help to set up the experiments and a nice portion of data analysis
- **In depth analysis is pending**
  - Master student anyone?
- **First results might indicate:**
  - Potential to increase BIF (30%?) by reducing the flight path without paying a price in additional background (see Fe, Pb (Cnat))
- **The investigation can still be complimented with a finer and wider grid of detectors**



Thanks!

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