

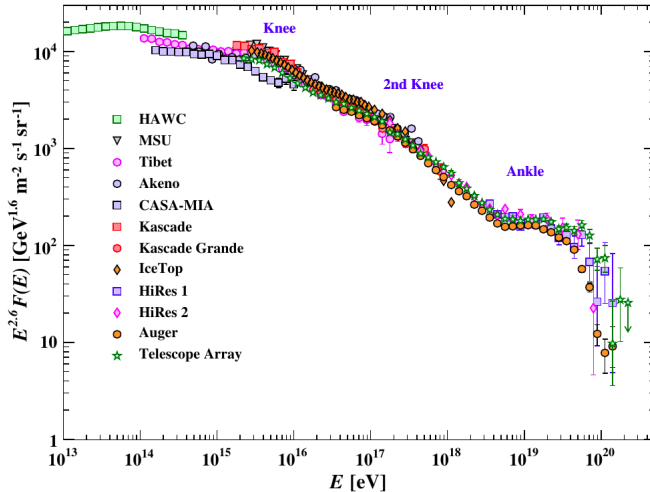
Geometry Calibration of IceCube

using Photon Timing from Downgoing Muons

Matti Jansson

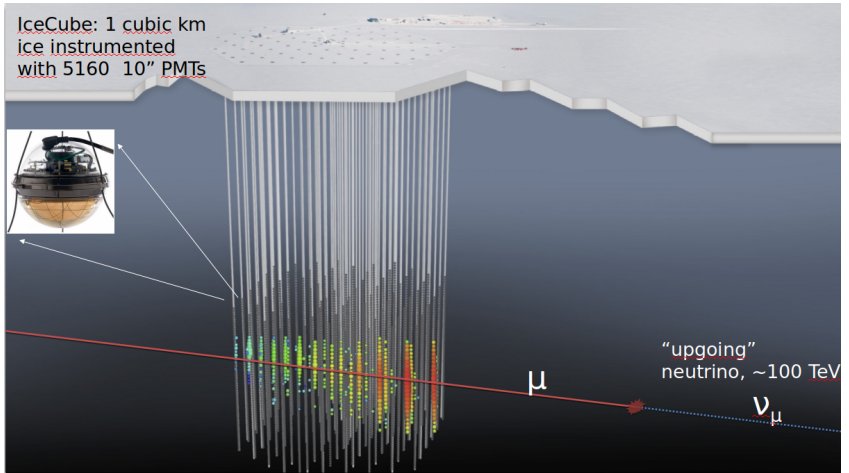
matti.jansson@fysik.su.se

Cosmic Rays

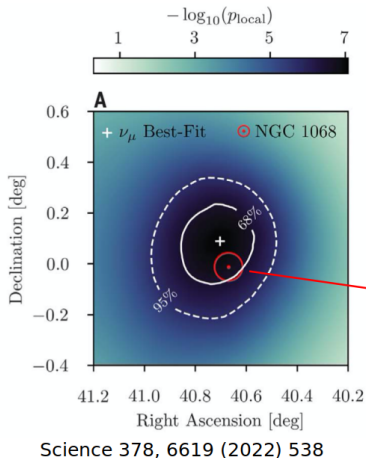


Neutrino Astronomy

IceCube: 1 cubic km
ice instrumented
with 5160 10" PMTs



Neutrino Emission from NGC 1068

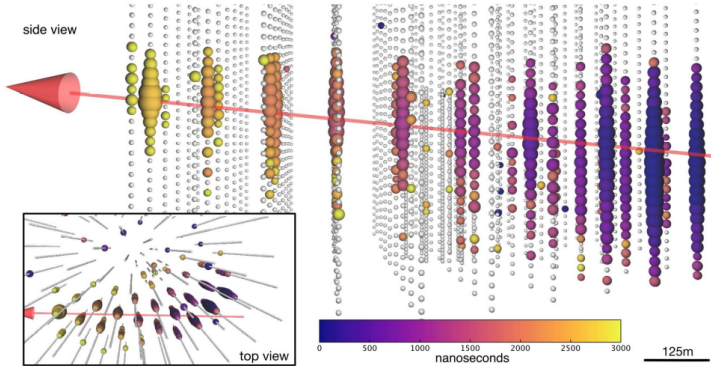


4.2 σ evidence of neutrino emission from the nearby (47 Mly) active galaxy NGC 1068

Reconstructed source location is 0.11° away from the galaxy's center



IceCube Alert Event IC170922A

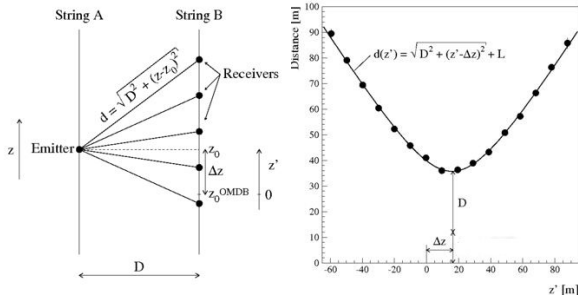


Science 361, 6398, (2018) eaat1378

Geometry Calibration

How do you measure the position of something 2 kilometers under the ice?

Flasher data for z calibration



The method for extracting relative depth from flasher data. On the left is an overview of a single DOM flashing and for DOMs on nearby strings the time delay will give a measured distance d . By doing this to all DOMs on a pair of strings we get the plot on the right where each dot is the average of all distances with the nominal vertical distance of z' and the bars shows the spread. The parameters D , Δz and L are fitted for and Δz is the calibrated vertical offset between the two strings.

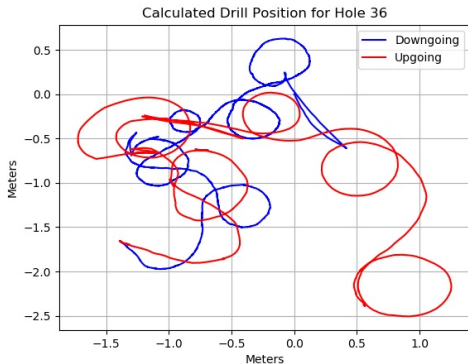
What about x and y ?

Installation



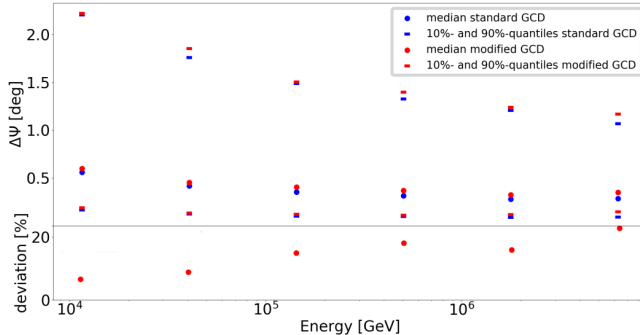
The drill works by flushing 90°C water, melting ice in front. During drilling the depth is measured by a pressure gauge and the orientation is measured by two pendulums and a fluxgate magnetometer. This data can be used to reconstruct the drill hole by integrating along the path of the drill.

Drill Data



But as we were drilling with hot water it did not have to melt the ice in a straight line and the drill data has not been trusted for details, but it does tell us that the DOMs are less than ~ 2 m from nominal.

Impact of x and y Misalignment



$\Delta\Psi$: opening angle between true and reconstructed direction

If the drill data is representative for x and y we have 15% larger opening angle.

What about x and y ?

- The drill data is not self-consistent.
- The flasher method has been unable to pin down x and y .

What about x and y ?

- The drill data is not self-consistent.
- The flasher method has been unable to pin down x and y .
- Maybe we could fit x and y if we had more flashers closer to the DOMs.

What about x and y ?

- The drill data is not self-consistent.
- The flasher method has been unable to pin down x and y .
- Maybe we could fit x and y if we had more flashers closer to the DOMs.
- But we have something that lights up the detector in every spot.

What about x and y ?

- The drill data is not self-consistent.
- The flasher method has been unable to pin down x and y .
- Maybe we could fit x and y if we had more flashers closer to the DOMs.
- But we have something that lights up the detector in every spot.
- We have millions of muon tracks in the detector every day lighting up the detector.

What about x and y ?

- The drill data is not self-consistent.
- The flasher method has been unable to pin down x and y .
- Maybe we could fit x and y if we had more flashers closer to the DOMs.
- But we have something that lights up the detector in every spot.
- We have millions of muon tracks in the detector every day lighting up the detector.
- Our reconstruction gives us the origin of that light.

What about x and y ?

- The drill data is not self-consistent.
- The flasher method has been unable to pin down x and y .
- Maybe we could fit x and y if we had more flashers closer to the DOMs.
- But we have something that lights up the detector in every spot.
- We have millions of muon tracks in the detector every day lighting up the detector.
- Our reconstruction gives us the origin of that light.
- Given the PMT timings and tracks we can estimate the location of our DOMs.

Method

- For a large number of tracks:
 - Remove hits on the string being calibrated.
 - Require at least 30 hits left.
 - Reconstruct the track.
- For track i , corresponding photon times $t_{i0} \dots t_{in_i}$ and noise level b find DOM position $(\hat{x}, \hat{y}, \hat{z})$ that maximizes

$$L(x, y, z) = \text{Log} \left(\prod_i \prod_{t=t_{i0}}^{t_{in_i}} (p(t|i, x, y, z) + b) \right)$$

, where p is the photon time pdf.

Cherenkov cone

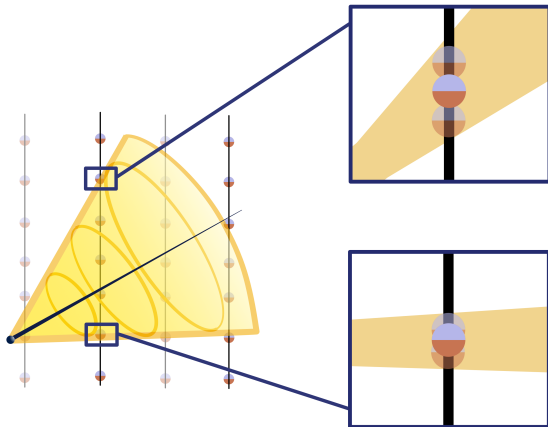
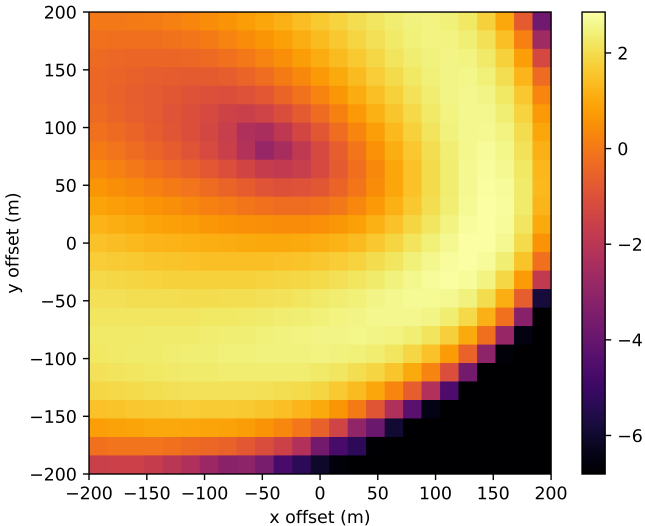
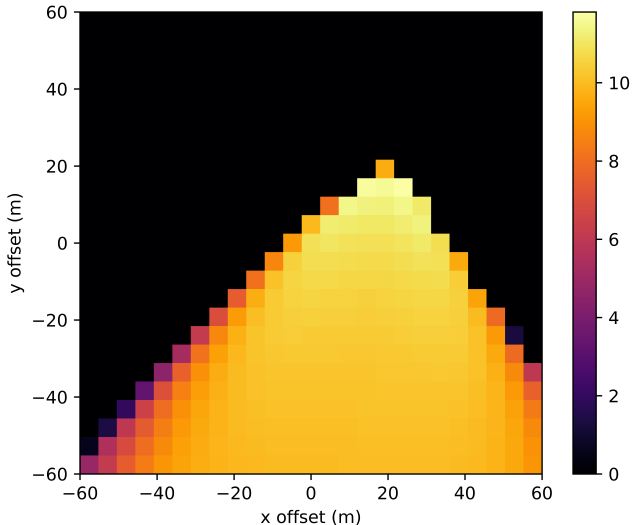


Illustration by K. Deoskar.

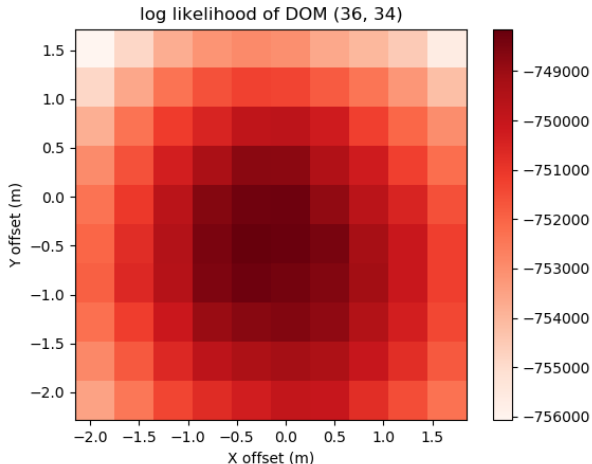
Delta LLH of Single Event (late hit)



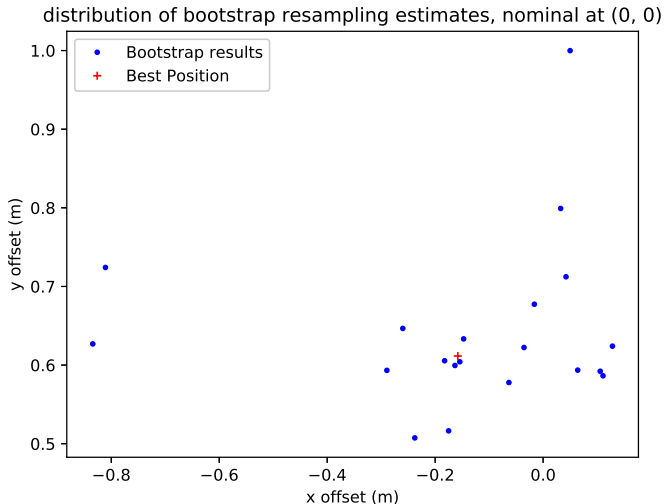
Delta LLH of Single Event (early hit)



Likelihood of many events

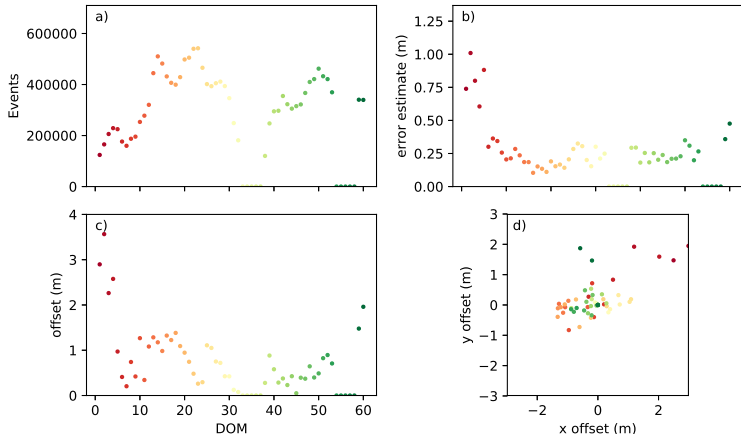


Estimating Error with bootstrap



String 35 results

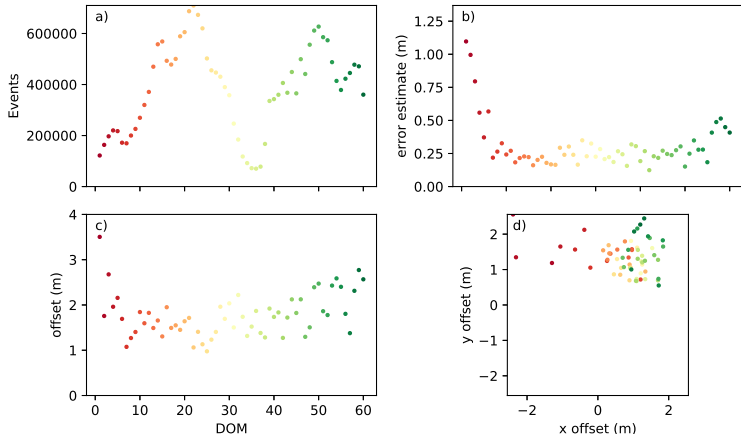
String 35 data



a) number of events per DOM. b) error estimate given by bootstrap.
 c) offset from nominal positions. d) best fit positions for each DOM.

String 36 results

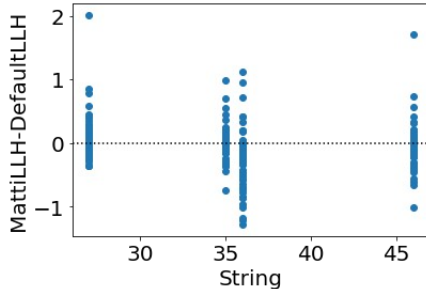
String 36 data



- a) number of events per DOM. b) error estimate given by bootstrap.
 c) offset from nominal positions. d) best fit positions for each DOM.

Confirmation with flashers

Recent improvements in modeling of light propagation in the ice (in particular birefringence) has shown flasher data to be consistent with my results.



Using the new geometry for the central strings we find that for strings where the DOM positions are clustered around nominal there is no change in likelihood but for string 36 which has moved over a meter the new position is a significantly better fit to the flasher data.

Summary

- Angular precision improves sensitivity to cosmic neutrino sources.
- The horizontal geometry is a systematic error that may impact angular resolution up to 15% .
- A muon based calibration is the first method in IceCube with precision to measure meaningful offsets from nominal.
- Using 4 days of data, three strings were consistent with nominal positions and one string is found with a statistically significant shift from nominal.
- The change in position of string 36 has been shown to also improve the description of flasher data.
- Next: Plan to apply both the muon and improved flasher calibration method to other strings, crosschecking each other, in order to provide an xy geometry for IceCube.

Thank you for your attention!

