

CERN Retreat 2022

Andrea Caputo

Something about me



Born in Calabria (1992)
Am I the first TH-Fellow from Calabria?



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First postdoc Tel Aviv - Weizmann
(2020-2022)



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My research

[Caputo et al., JCAP 08 \(2022\) 08, 045](#)

[Caputo et al., Phys.Rev.Lett. 128 \(2022\) 22, 221103 \(CCSNe\)](#)

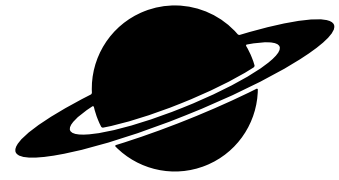
[Caputo et al., Phys.Rev.D 105 \(2022\) 3, 035022 \(CCSNe\)](#)

[Caputo et al., Phys.Rev.Lett. 127 \(2021\) 18, 181102 \(Hypernovae\)](#)

[Caputo et al., Phys.Rev.D 101 \(2020\) 12, 123004 \(Sun\)](#)

O'Hare, [Caputo](#), et al., [Phys.Rev.D 102 \(2020\) 4, 043019 \(Sun\)](#)

Stars and new physics



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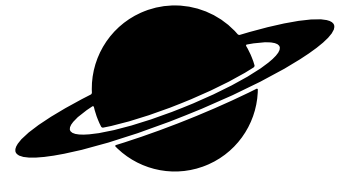
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Dark Matter and telescopes

[Caputo et al., JCAP 03 \(2020\) 001 \(X-rays\)](#)

[Caputo et al., JCAP 05 \(2021\) 046 \(IR\)](#)



Bernal, [Caputo](#), et al., *Phys.Rev.D* 103 (2021) 6, 063523 (optic)

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Dark Matter and other indirect probes

Caputo et al., *Phys.Rev.Lett.* 127 (2021) 1, 011102 (21 cm)

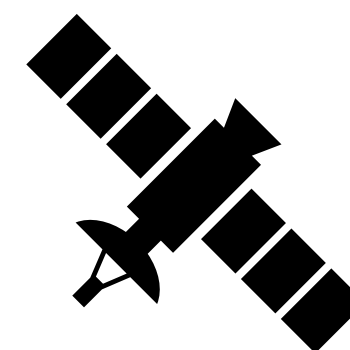
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Bernal, Caputo, et al., *arXiv* 2208.13794 (blazars data)

Caputo et al., *Phys.Rev.D* 100 (2019) 6, 063515 (pulsars)

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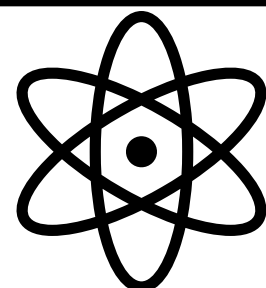


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[Caputo et al., Phys.Rev.D 100 \(2019\) 11, 116007 \(sub-GeV DM\)](#)

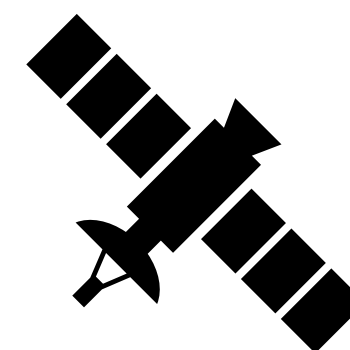
[Caputo et al., Phys.Lett.B 802 \(2020\) 135258 \(sub-GeV DM\)](#)

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[Caputo et al., Phys.Rev.D 104 \(2021\) 9, 095029 \(Dark Photon DM\)](#)

Dark Matter and other indirect probes



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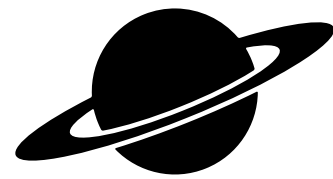
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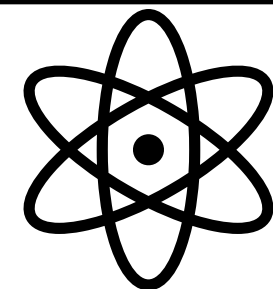
Gravitational Waves

Caputo et al., *JAstrophys.J.* 892 (2020) 2, 90

Toubiana, Sberna, Caputo, et al.,
Phys.Rev.Lett. 126 (2021) 10, 101105

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Dark Matter direct detection



Caputo et al., *Phys.Rev.D* 100 (2019) 11, 116007 (sub-GeV DM)

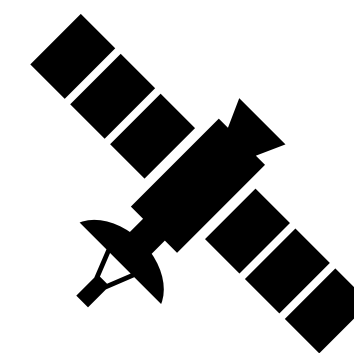
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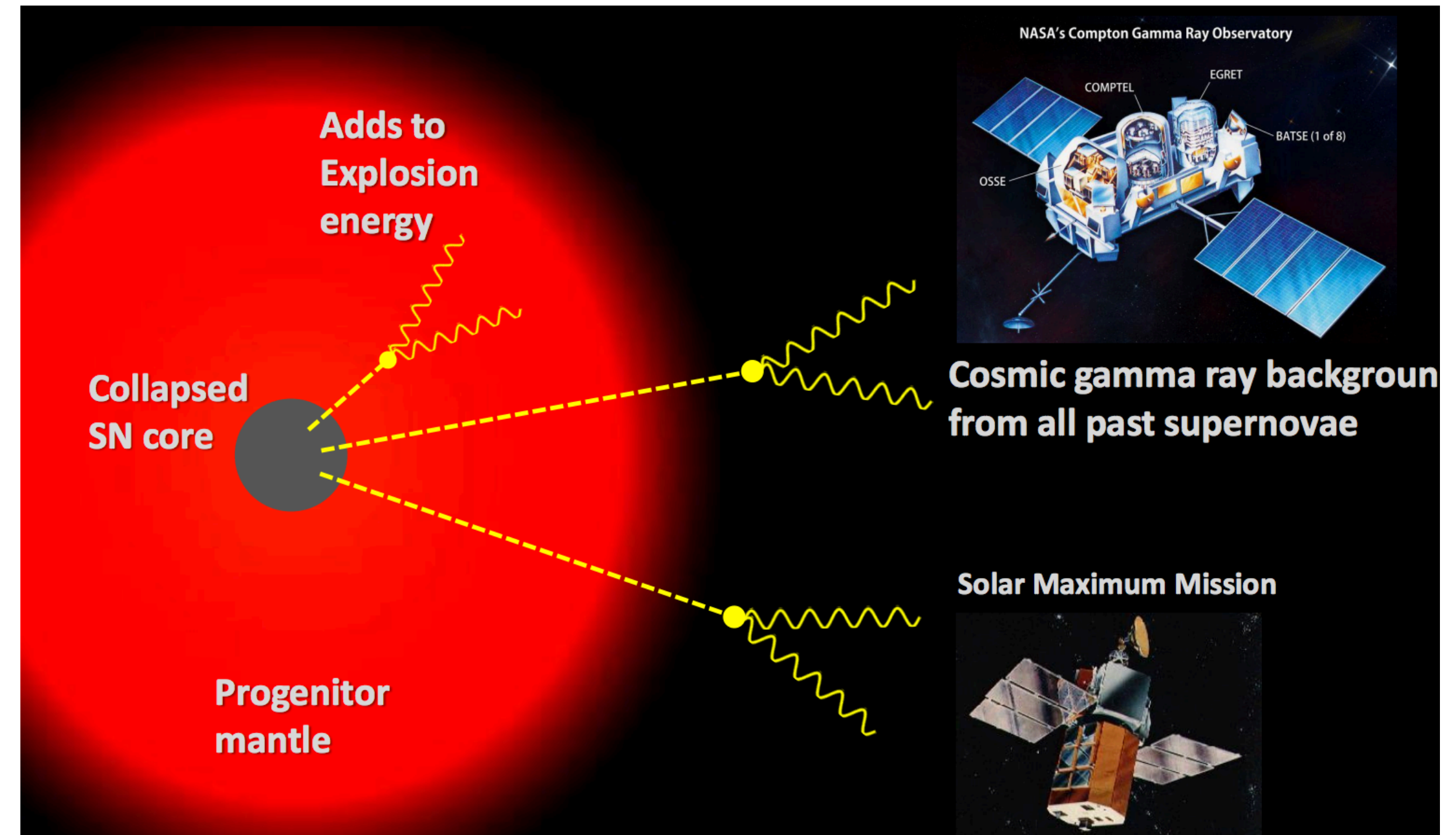
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Few Highlights

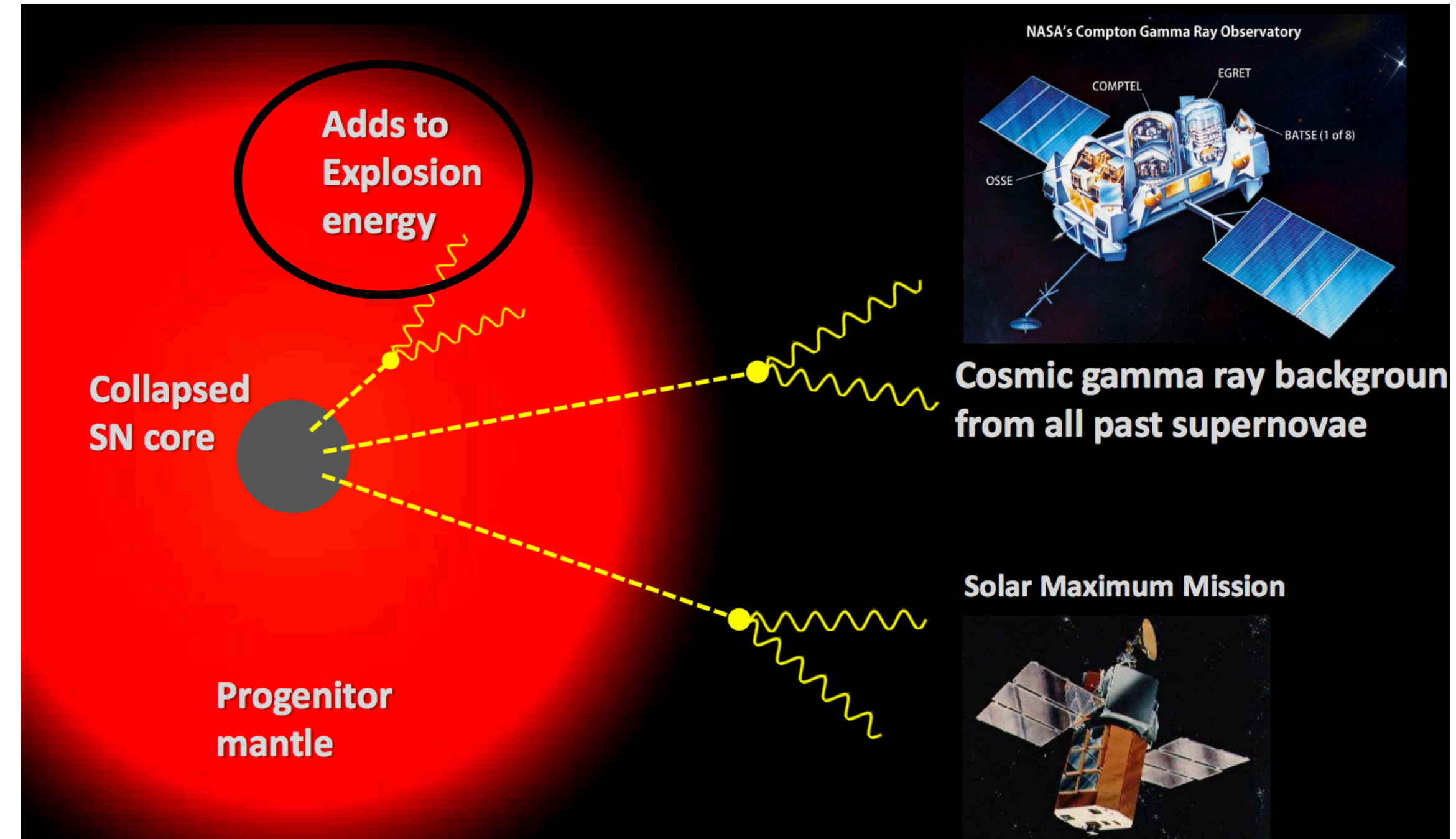
1) Use Supernovae to look for new physics



Few Highlights

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If new particles, let's call them axions, are produced in the SN core and they then **decay in the SN mantle**, they can **dump energy** which should then show up either in **kinetic energy or radiation energy**.



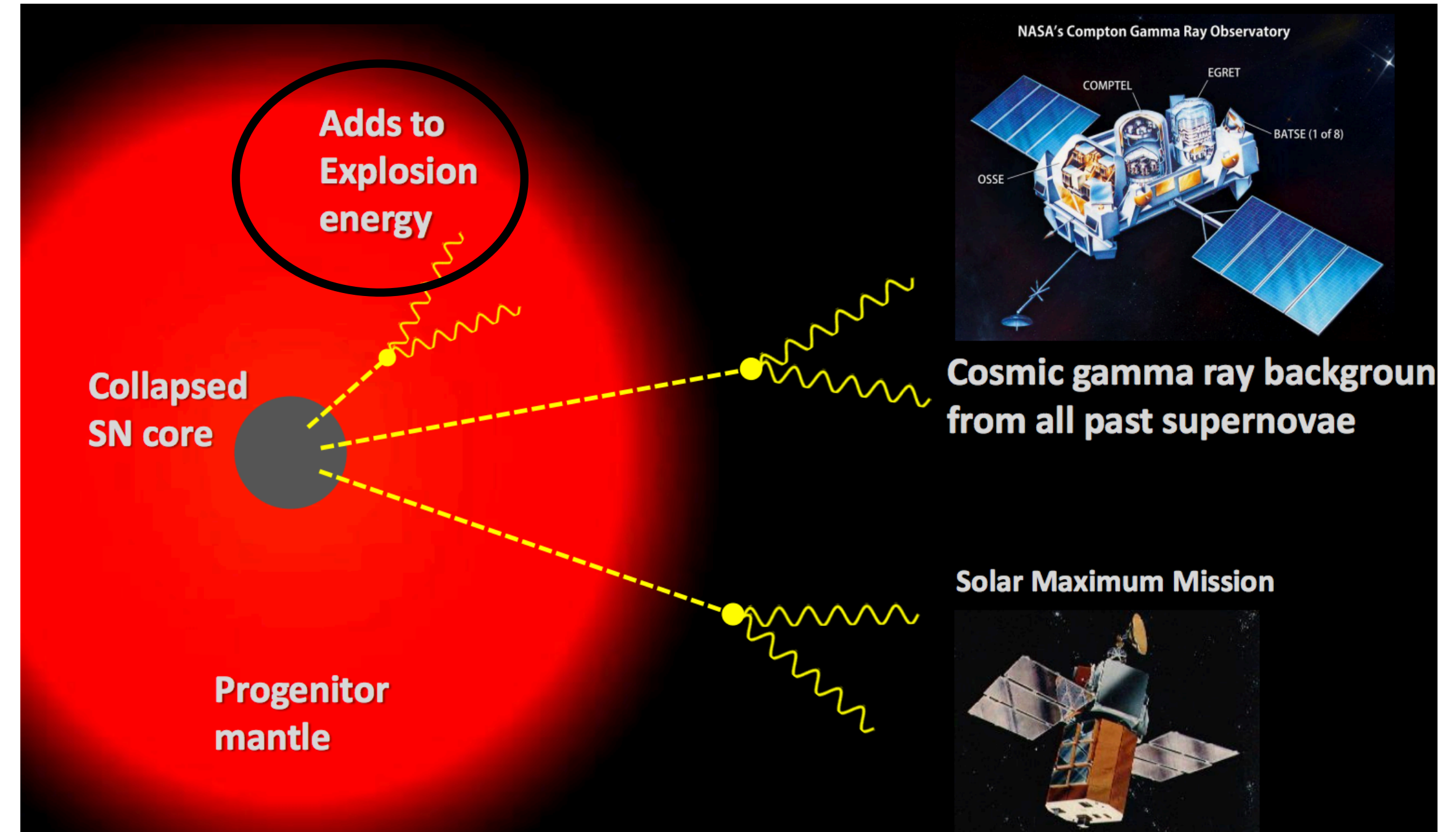
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Example model: axion coupling to photons

$$\mathcal{L}_{a\gamma\gamma} = -\frac{1}{4}g_{a\gamma}a\tilde{F}^{\mu\nu}F_{\mu\nu}$$



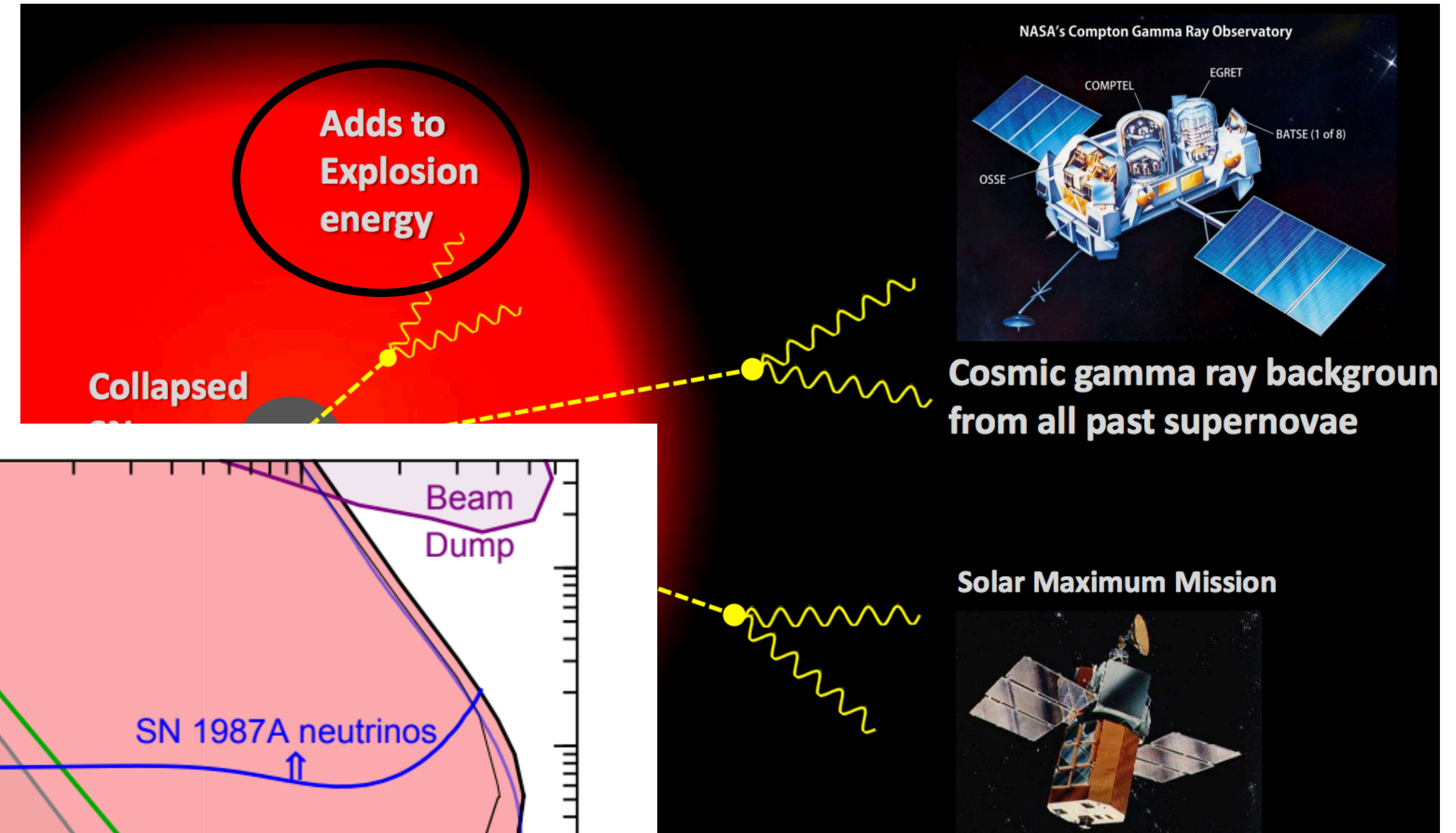
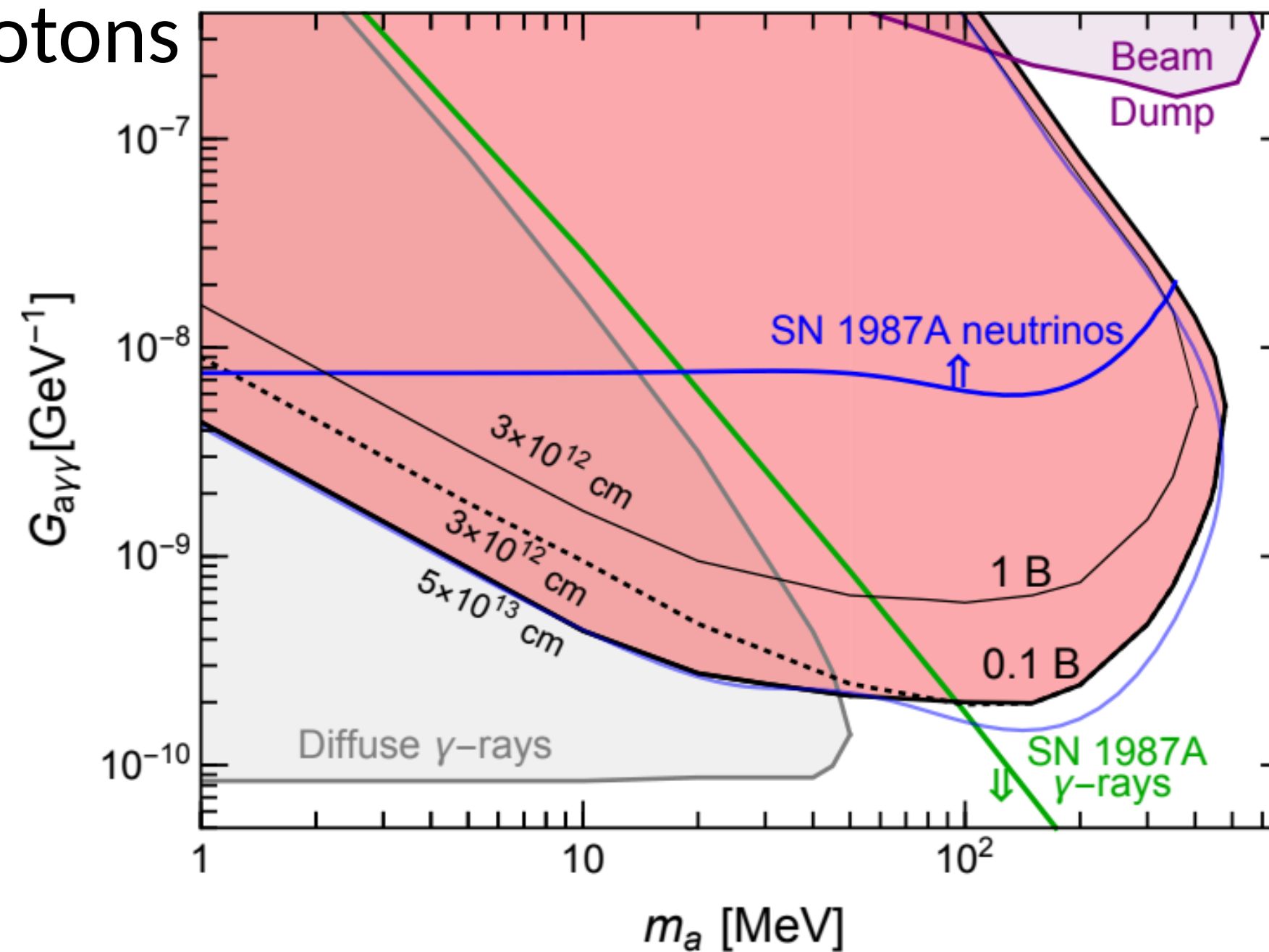
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A.C., H.T. Janka, G. Raffelt, E. Vitagliano, PRL 22

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1) Use Supernovae to look for new physics

2) Dark Matter direct detection

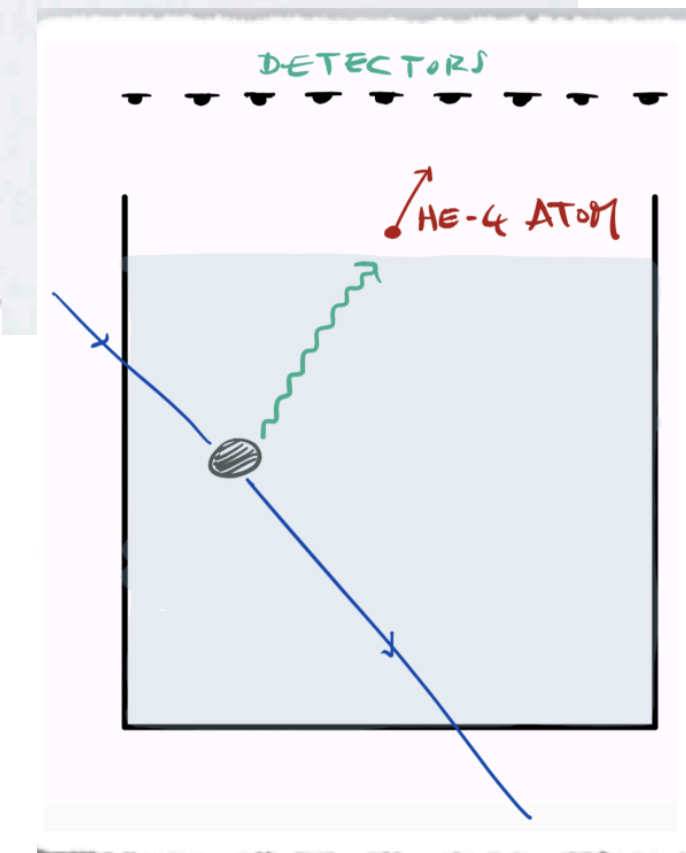
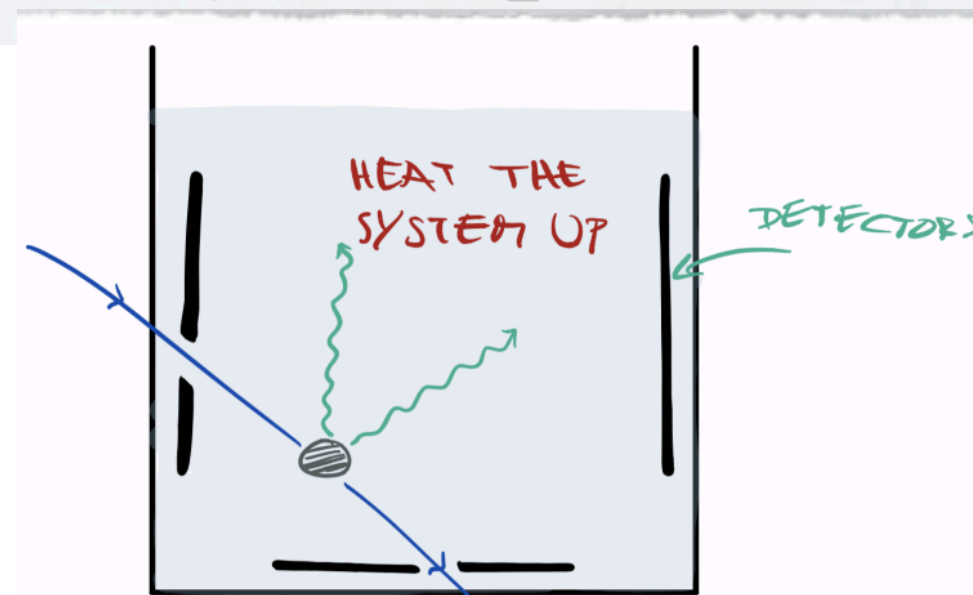
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He-4 dark matter detectors, EFT formalism

$$S_{eff} = \int d^4x \left[\frac{1}{2} \dot{\pi}^2 - \frac{c_s^2}{2} (\vec{\nabla} \pi)^2 - |\partial\chi|^2 - m_\chi^2 |\chi|^2 + \lambda_3 \sqrt{\frac{m_{He}}{\bar{n}}} c_s \dot{\pi} (\vec{\nabla} \pi)^2 + \lambda'_3 \sqrt{\frac{m_{He}}{\bar{n}}} c_s \dot{\pi}^3 \right. \\ \left. - \left(g_1 \sqrt{\frac{m_{He}}{\bar{n}}} c_s \dot{\pi} - \frac{g_1 c_s^2}{2 \bar{n}} (\vec{\nabla} \pi)^2 + \frac{g_2 m_{He} c_s^2}{2 \bar{n}} \dot{\pi}^2 \right) |\chi|^2 \right]$$



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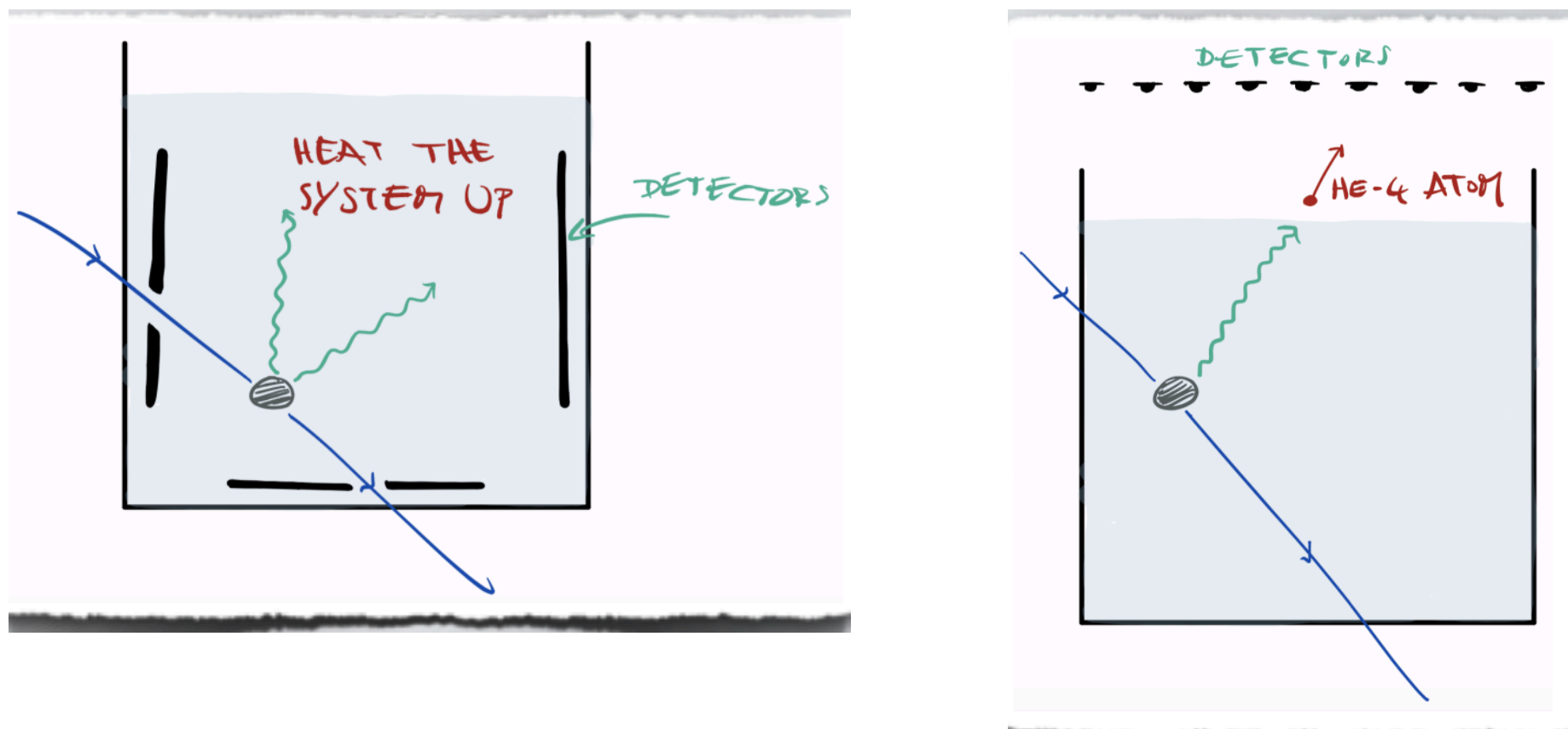
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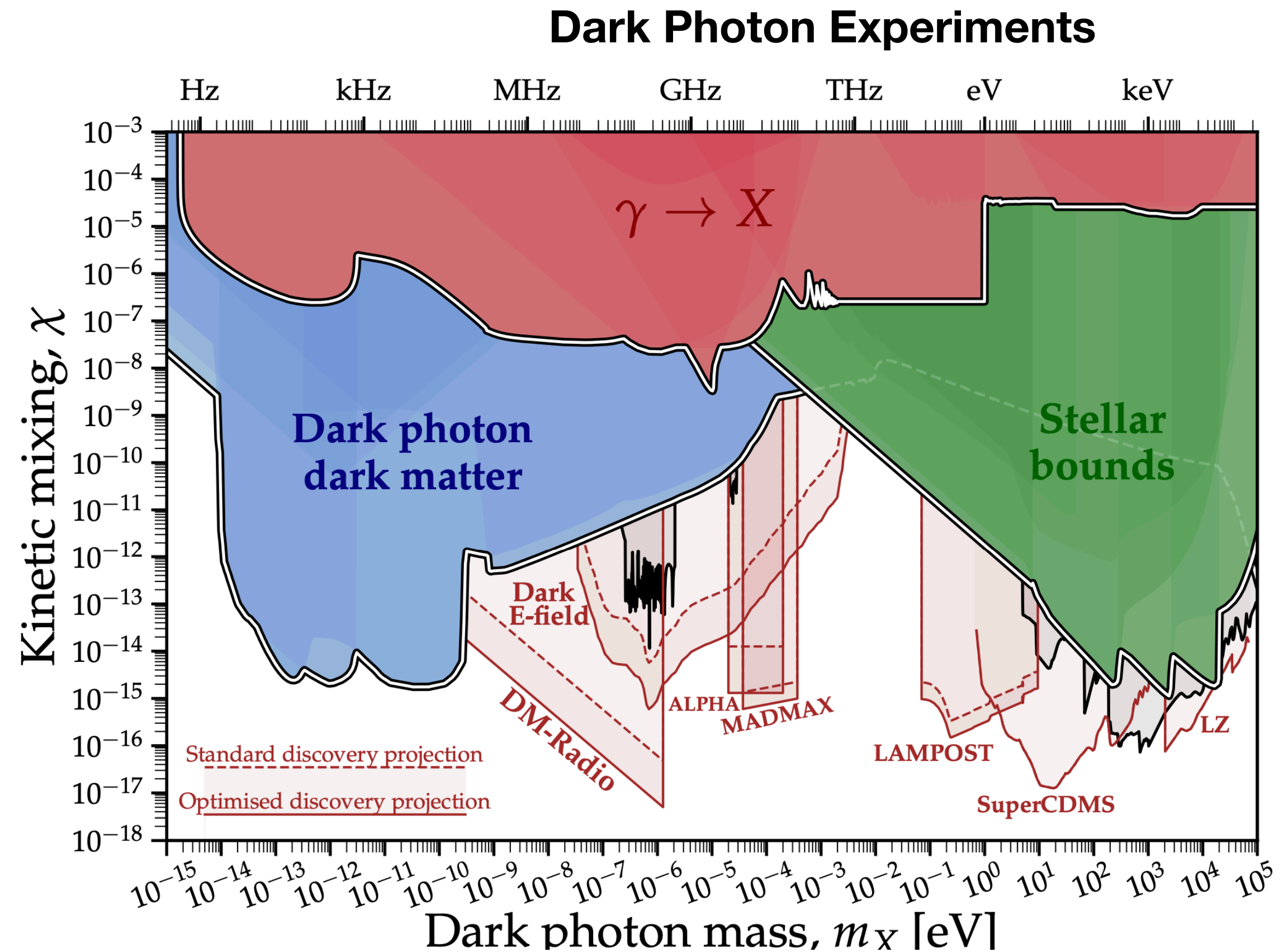
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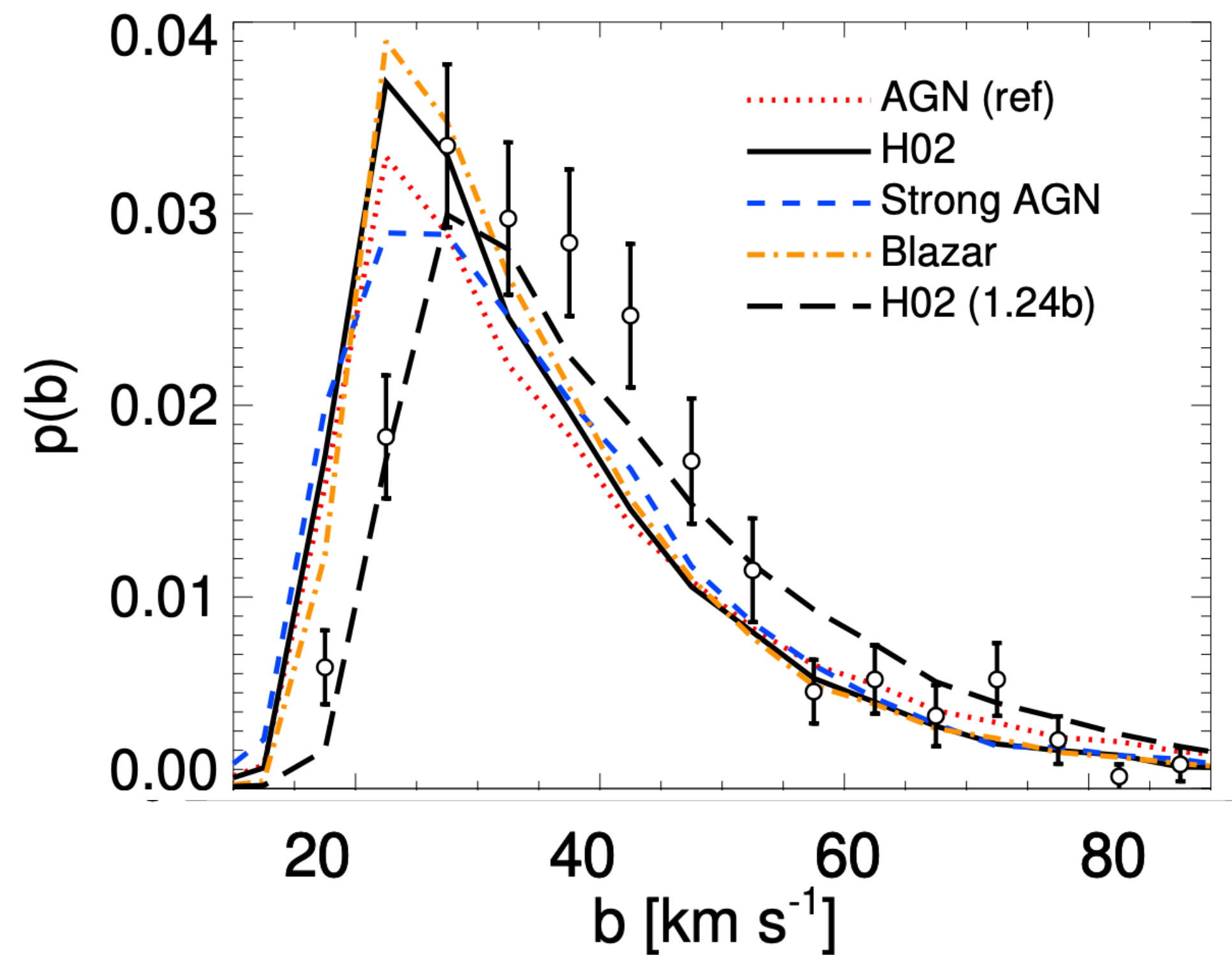
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The **Lyman- α forest** is produced by the absorption of the continuum of a distant source (typically a quasar) caused by the **neutral hydrogen** present along the line-of-sight that connects the observer to the source.

Interestingly, there seems to be a tension between the observed number of lines with **b-parameters** in the range 25–45 km/s and the predictions from simulations

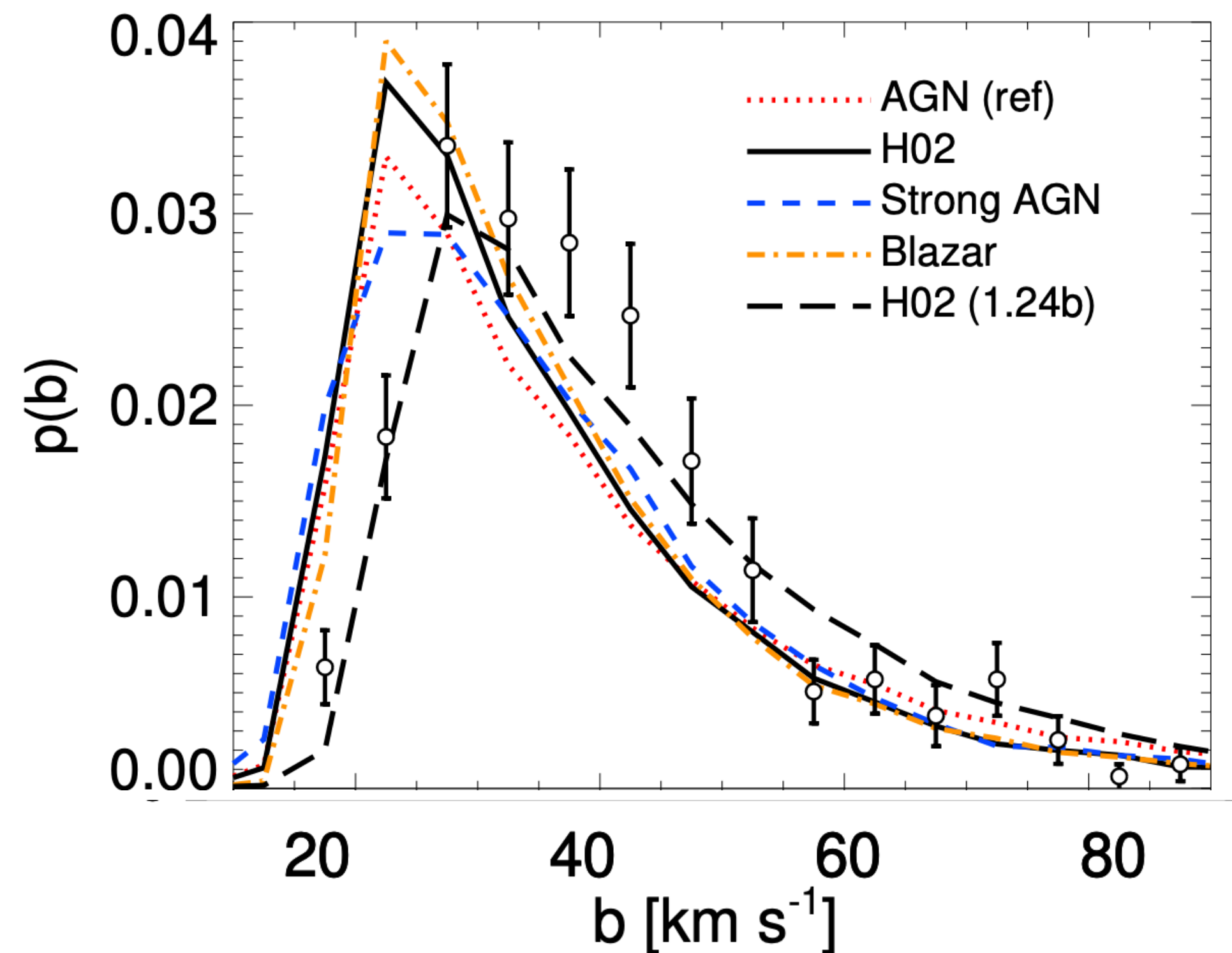


Bolton et al., 2021

$$\Delta\nu = \nu_\alpha (b_{\text{th}}^2 + b_{\text{nth}})^{1/2} / c$$

$$b_{\text{th}} = (2k_B T / m_H)^{1/2}$$

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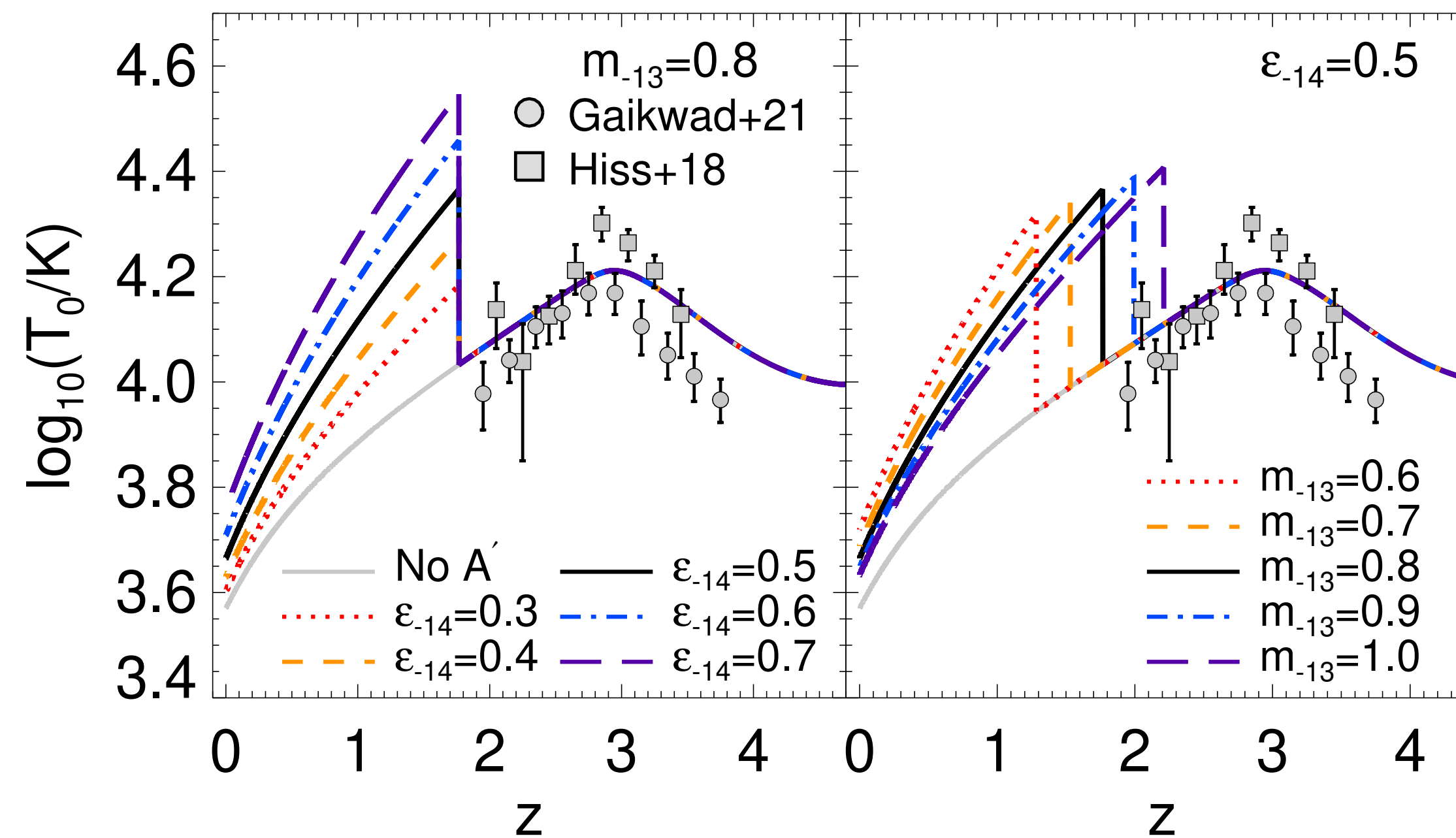
What if **dark matter** is responsible for this?

Dark Photon Dark Matter can be efficiently converted into photons!

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{4}\tilde{F}_{\mu\nu}\tilde{F}^{\mu\nu} + \frac{m^2}{2}\tilde{A}_\mu\tilde{A}^\mu - \frac{e}{(1+\varepsilon^2)^{1/2}}J^\mu\left(A_\mu + \varepsilon\tilde{A}_\mu\right)$$

$$E_{A' \rightarrow \gamma} \sim 2.5 \text{ eV} \left(\frac{\epsilon_{-14}}{0.5} \right)^2 \left(\frac{3}{1 + z_{\text{res}}} \right)^{3/2} \left(\frac{m_{-13}}{0.8} \right)$$

Energy injected per unity baryon via dark photon to photon conversion



**J. Bolton, A.C, H. Liu, M. Viel,
PRL in press**

Physics Magazine Focus

Thermal histories for baryons at the mean background density

Thanks for the attention!

