

MUON Collider Collaboration: Final Cooling



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Overview

Tracking codes: ICOOL an RF-Track

Discussion on the technical challenges of absorbers

- 1. Heat dissipation of windows**
- 2. Thermodynamic of hydrogen absorbers**

Why final cooling?

Table 1.1: Parameter list defined by the international muon collider collaboration.

Parameter	Symbol	Units	3 TeV	10 TeV	14 TeV
Luminosity	\mathcal{L}	$10^{34} \text{ cm}^{-2}\text{s}^{-1}$	1.8	20	40
Muons/bunch	N	10^{12}	2.2	1.8	1.8
Repetition rate	f_r	Hz	5	5	5
Beam power	P_{beam}	MW	5.3	14.4	20
Collider circumference	C	km	4.5	10	14
Average bending field	$\langle B \rangle$	T	7	10.5	10.5
Norm. long. emittance	$\epsilon_{\text{L,N}}$	MeV m	7.5	7.5	7.5
Energy spread	σ_E/E	%	0.1	0.1	0.1
Bunch length	σ_z	mm	5	1.5	1.07
Interaction point beta	β	mm	5	1.5	1.07
Norm. trans. emittance	$\epsilon_{\perp,\text{N}}$	μm	25	25	25
Interaction point beam size	σ_{\perp}	μm	3.00	0.90	0.63

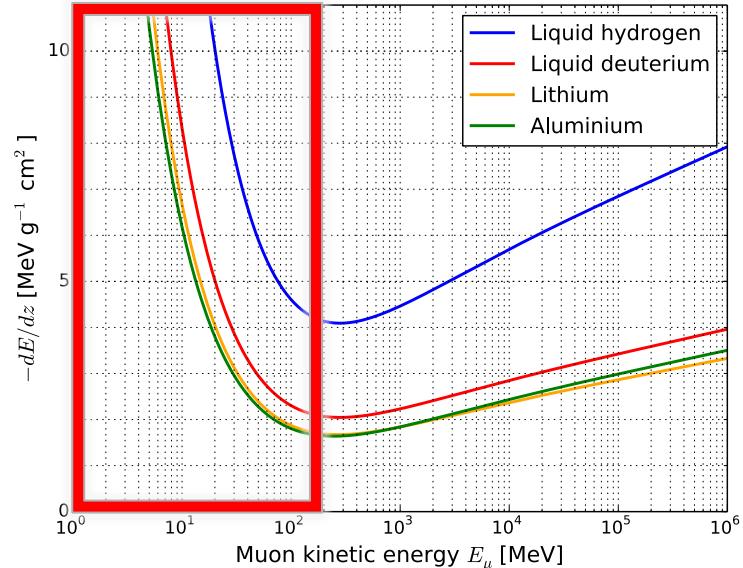
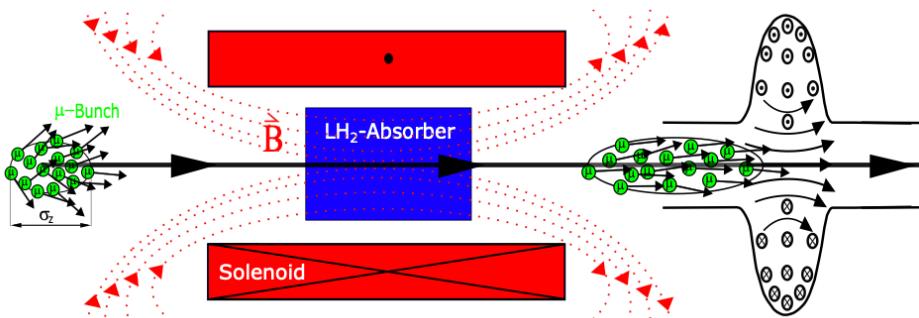
Emittance:

- In past, $\epsilon_{\text{trans}} = 55 \mu\text{m}$ [1]
- Goal is to reach $25 \mu\text{m}$

Ionization cooling

$$\frac{d\epsilon_{\perp,N}}{ds} = -\frac{\epsilon_{\perp,N}}{E\beta^2} \left\langle \frac{dE}{ds} \right\rangle + \frac{\beta_\perp(13.6[\text{MeV}])^2}{2\beta^3 mc^2 L_R}$$

$$\frac{d\epsilon_{L,N}}{ds} = -\epsilon_{L,N} \frac{d}{dE} \left\langle \frac{dE}{ds} \right\rangle + \left\langle \frac{d\epsilon_{L,N}}{ds} \right\rangle_{\text{Straggle}}$$



$$\mathcal{L} \propto \frac{\gamma^3}{C} \frac{N^2}{\epsilon_{\perp,N} \epsilon_{L,N}}$$

Tracking ionization cooling

- Initial beam parameters from 6D-cooling: $\epsilon_{\perp} = 300 \mu\text{m}$, and $\epsilon_L = 1.5 \text{ mm}$
- Goal: find a method for efficient **emittance reductions** $\epsilon_{\perp} = 25 \mu\text{m}$
- Have chosen **ICCOOL** for optimizing parameter:
 1. Initial beam energy
 2. Magnetic field strength
 3. Absorber type and length
 4. Matched beam optics

Acknowledgement to Elena Fol

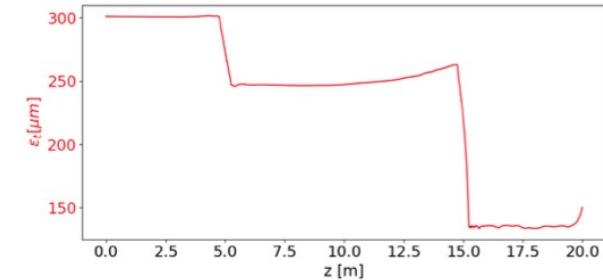
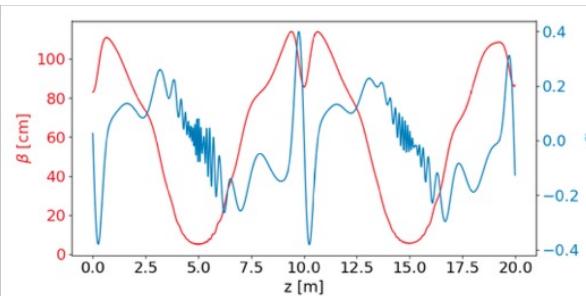
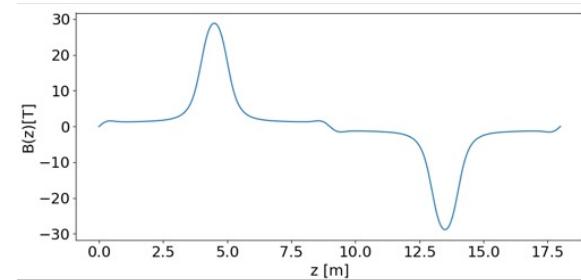
ICOOL example

Acknowledgement to Elena Fol

- 2 cells without RF structure
- Optics matching mitigates emittance blow up
- Varying initial parameters for finding:
 1. Min correlation α in absorbers
 2. Max ϵ_{\perp} decrease, min ϵ_L increase
 3. Max transmission

[2] Varying free parameters

$$\min_p \frac{\Delta\epsilon_{||}}{(\Delta\epsilon_{\perp}\Delta N)} + \bar{\alpha}$$



[2] E. Fol et al, doi:10.18429/JACoW-IPAC2022-WEPOMS047

RF-track example

- **RF-Track by A. Latina (CERN)**
 - Tracking code for of low-energy LINACs
 - Space charge effect included
 - User interface Python/Octave as user interface
- Same parameter optimizer used, but RF-structure included
- Preliminary results for the first 4 cells (RF-track still under development)

Cell configuration						Beam parameters (end of the cell)					
	Length [m]	LH thickness [m]	Drift [m]	1. Freq [MHz]	1. Vmax [MV]	2. Phase	Emittance Tr. [mm mrad]	Emittance Long. [mm]	Bunch length [mm]	Pz [MeV/c]	Pz spread
							300,0	1,56	49,9	135	3,5
1	7,4	0,21	0,9	321	17,9	57	278	2,18	73,5	121,3	3,85
2	5,56	0,17	0,56	266	10,4	46,1	256	2,27	69,3	105,78	3,5
3	5,45	0,11	0,7	249	10,8	81,3	243	3,7	134,5	102,2	5,4
4	4,6	0,11	0,85	201	10,48	76,3	229	6,4	400	95,6	6,7

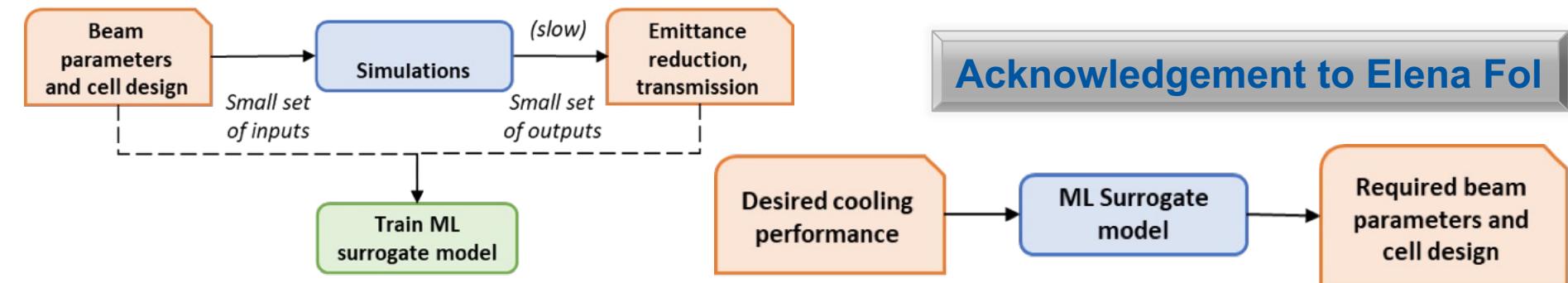
Acknowledgement to Elena Fol

Note: numbers are preliminary,
ionisation cooling in RF-Track
is still under development



Surrogate modeling [4]

- Optimization routines are time-consuming processes
- Initial and final parameter from the optimized simulations will be stored
- Surrogate model allows an automatic mapping between initial condition and simulated results



Technical risks on absorbers

- Due to the **target parameters** the beam size is small at the end of the cooling channel
- High **field magnets** up to **50 T** and $\varepsilon_{\perp} = 25 \mu\text{m} \rightarrow 1 \text{mm}$ beam size in diameter
- Follows: energy **storage** in a **small volume**
- Question: what does this mean for the **absorbers** (last cell)?

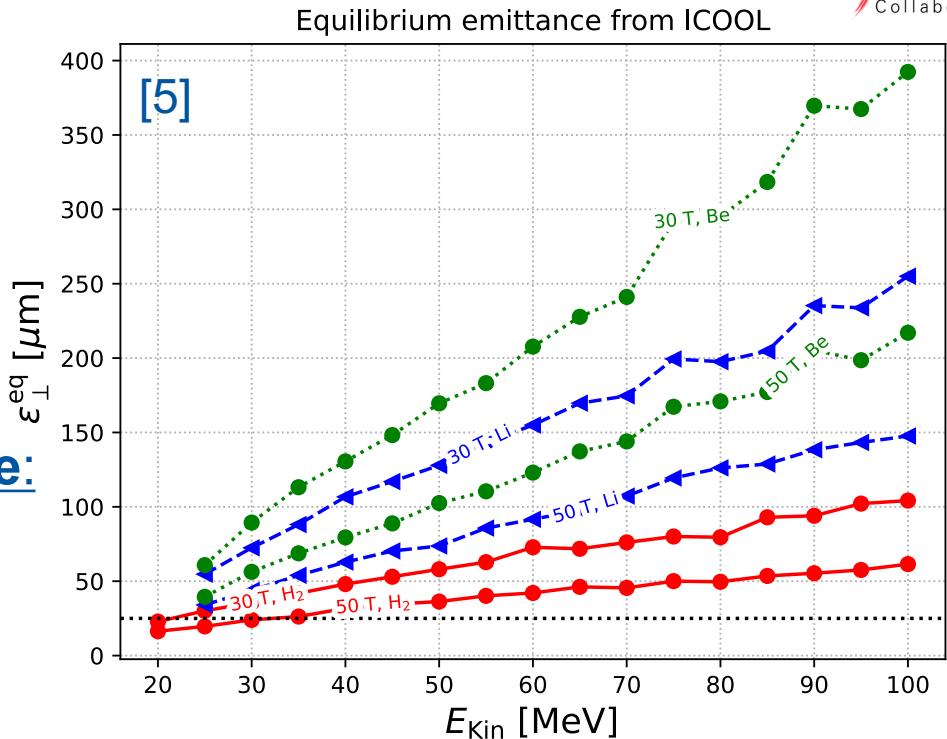
Absorber choice in the last cell



$$\epsilon_{\perp}^{\text{eq}} = \frac{\beta_{\perp}(13.6[\text{MeV}])^2}{2\beta mc^2 L_R \left\langle \frac{dE}{ds} \right\rangle}$$

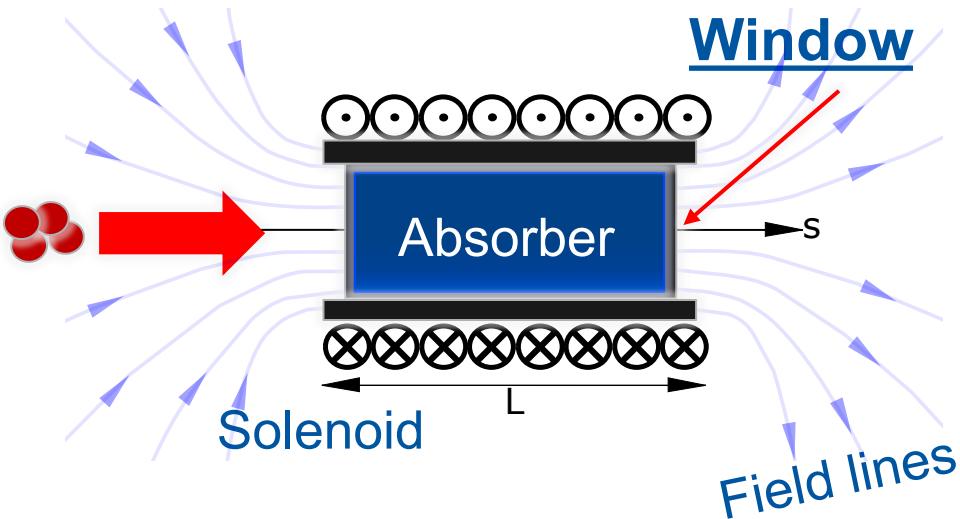
Achieving the target trans emittance:

- **Hydrogen**
- **Strong B-fields**
- **Low beam energy**



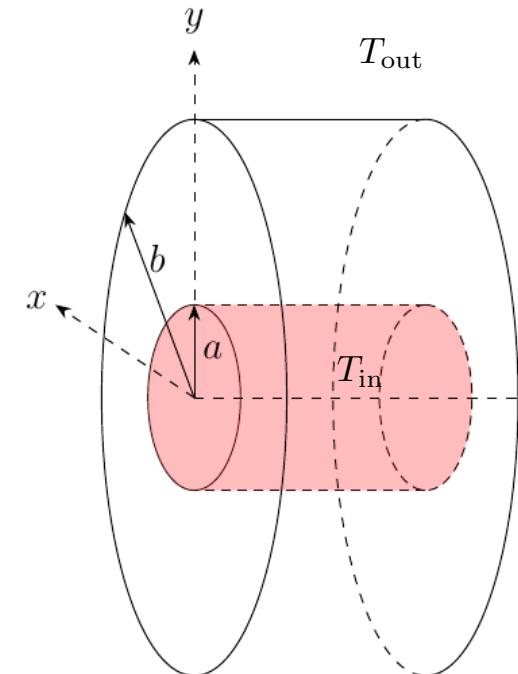
Absorber confinement

- H_2 can be solid, liquid, gaseous, supercritical,...
- Mitigate muons-window interactions
- Requirements: low Z and thin
- Candidates: Be and Si_3N_4



Heat dissipation in the window

Beam channel volume depends on
emittance, field and window thickness

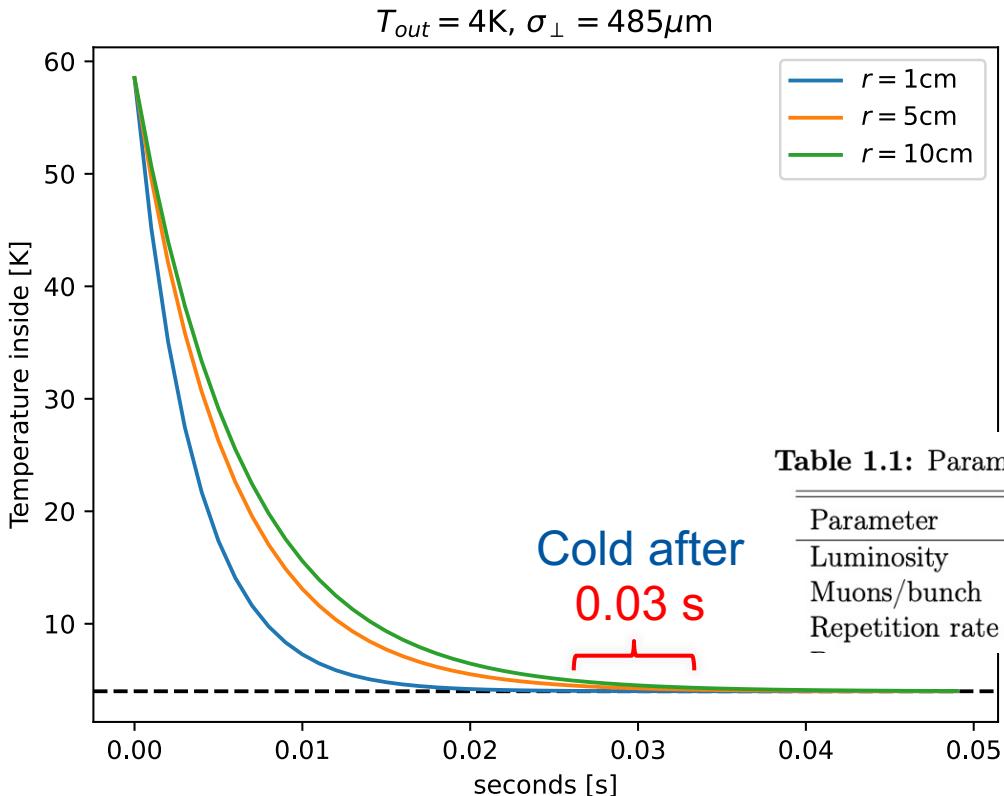


Beam causes an **instantaneous** temperature increase
Heat **dissipation** from inside to the edge

$$\frac{dQ}{dt} = -\lambda \cdot A \cdot \frac{dT}{dr}. \quad \text{Thermal conductivity}$$

$$\int_{T_{in}}^{T_{out}} dT = -\frac{\Delta Q}{\Delta t \cdot \lambda} \frac{1}{2\pi L} \int_a^b \frac{dr}{r}$$

$$\Delta Q = 2\pi L \lambda \frac{T_{in} - T_{out}}{\ln(b/a)} \Delta t$$



Example:

- Beryllium,
 - norm emittance is **25 μm**
 - Solenoid field is **50 T**
 - Window in steady cooled **4 K** bath
- final conditions*

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0.2 seconds



Thermodynamics in liquid H₂

- $\epsilon_{\perp,N}$ reductions = storing energy in the absorber

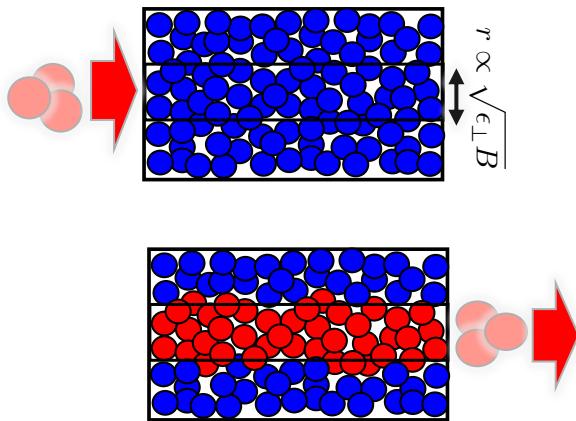
$$\rho = 70 \text{ kg m}^{-3} \quad B = 50 \text{ T}$$

$$A = 0.002 \text{ kg m}^{-3} \quad \epsilon_{\perp,N} = 25 \mu\text{m}$$

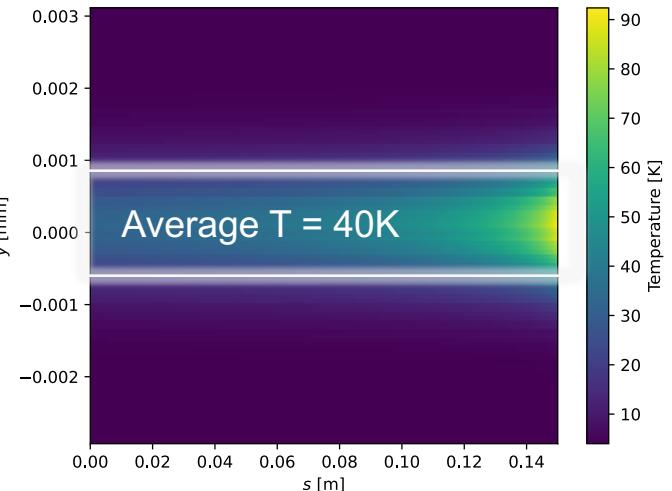
$$n = \frac{\rho V}{A}$$

$$p = \frac{nRT}{V} = \frac{\rho RT}{A}$$

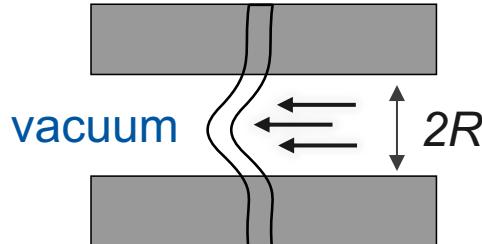
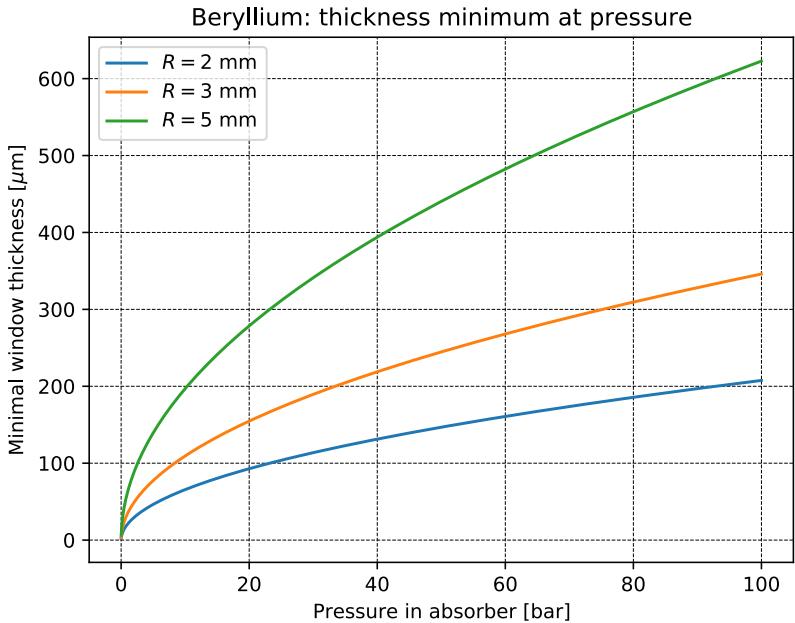
$$p = \frac{70 \text{ kg m}^{-3} \cdot 8 \text{ J mol}^{-1} \text{ K}^{-1} \cdot 40 \text{ K}}{0.002 \text{ kg mol}^{-1}} \approx 100 \text{ bar}$$



20MeV to 5Mev



Minimum window thickness



d ...window thickness [m]

p ...pressure [Pa]

R ...radial size [m]

ν ...Poisson's ratio

σ_{\max} ...Maximum yield strength [N m^{-2}]

$$d = \sqrt{\frac{3}{8} p(1 + \nu) \frac{R^2}{\sigma_{\max}}}$$

- Thick windows influences the beam
 1. Longitudinal emittance
 2. Lifetime

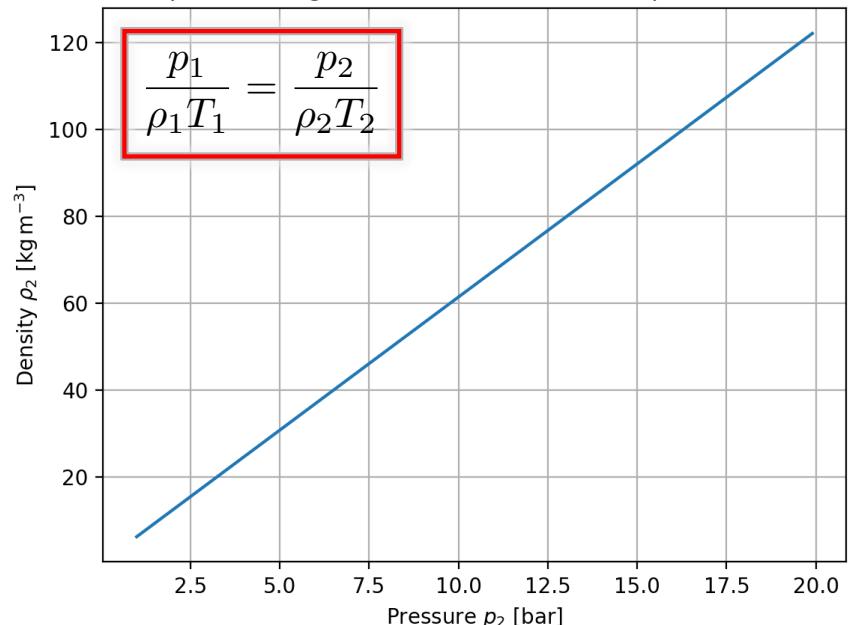
Alternatives to liquid H₂

- Liquid H₂ plays an important part in final cooling
- The low ε_{eq} guarantees high cooling performance for beams with:
 1. High initial beam **momentum**
 2. High initial **transverse emittance**
- But small beam sizes at low energies cause **huge pressures**
- **What about vaporized H₂?**

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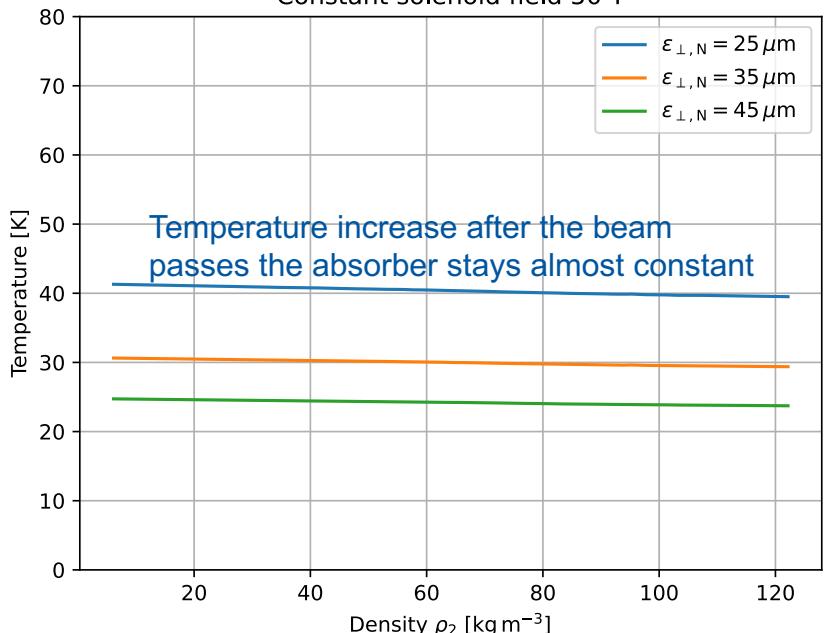
Vaporized H₂ ?

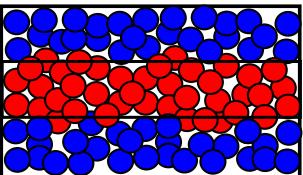
$$\rho_1 = 0.09 \text{ kg m}^{-3}, T_1 = 273 \text{ K}, T_2 = 4 \text{ K}, p_1 = 1 \text{ bar}$$



Initial beam 10MeV → 5MeV

Constant solenoid field 50 T





$$\Delta Q = \left\langle \frac{\partial E}{\partial s} \right\rangle \rho L \quad \Delta Q = c m \Delta T$$

$$\Delta Q = \left\langle \frac{\partial E}{\partial s} \right\rangle \frac{m}{V} L$$

$$\Delta Q = \left\langle \frac{\partial E}{\partial s} \right\rangle \frac{m}{\pi \beta_{\perp} \epsilon_{\perp} L} \quad \Delta T = \frac{\Delta Q}{m c}$$

$$\boxed{\Delta T = \frac{1}{\pi \beta_{\perp} \epsilon_{\perp} c} \left\langle \frac{\partial E}{\partial s} \right\rangle}$$

- The **temperature increase** in gas absorbers depends only on the **beam size ($\epsilon_{\perp} \beta_{\perp}$)**, the **spec. heat capacity (c)** and the **Bethe-Bloch** value for this gas

- It does **NOT** depend on the **gas density!**
- Low gas densities mitigate the pressure increase caused by the beam**
- But low densities cause longer absorbers and **longer solenoids**

Summary

- First simulation with ICOOL and RF-Track:
 - ✓ **Emittance reduction** achieved by optimizing initial cooling condition
 - ✓ ε_{\perp} reduction: **300 μm** to **230 μm** in the first 4 cells in RF-Track
 - ✓ Results trained in a surrogate model for saving simulation time
- Heat **dissipation** of Be windows is fast (**cooled** down after some **10 ms**)
- Liquid H₂ at small beam sizes and low energies leads to a **huge pressure increase** inside the absorber
- **Vaporized** H₂ at low densities leads to **low pressures** but to **longer** absorbers

Outlook

- We focus on engineering works of H₂ absorbers design
- We are going to accelerate the development of RF-Track and compare it with ICOOL to generate an optimized baseline design

Special thanks to

Andrea Latina & Jose Antonio Ferreira Somoza