

International  
Muon Collider  
Collaboration



# Machine-detector interface studies for a 10 TeV muon collider

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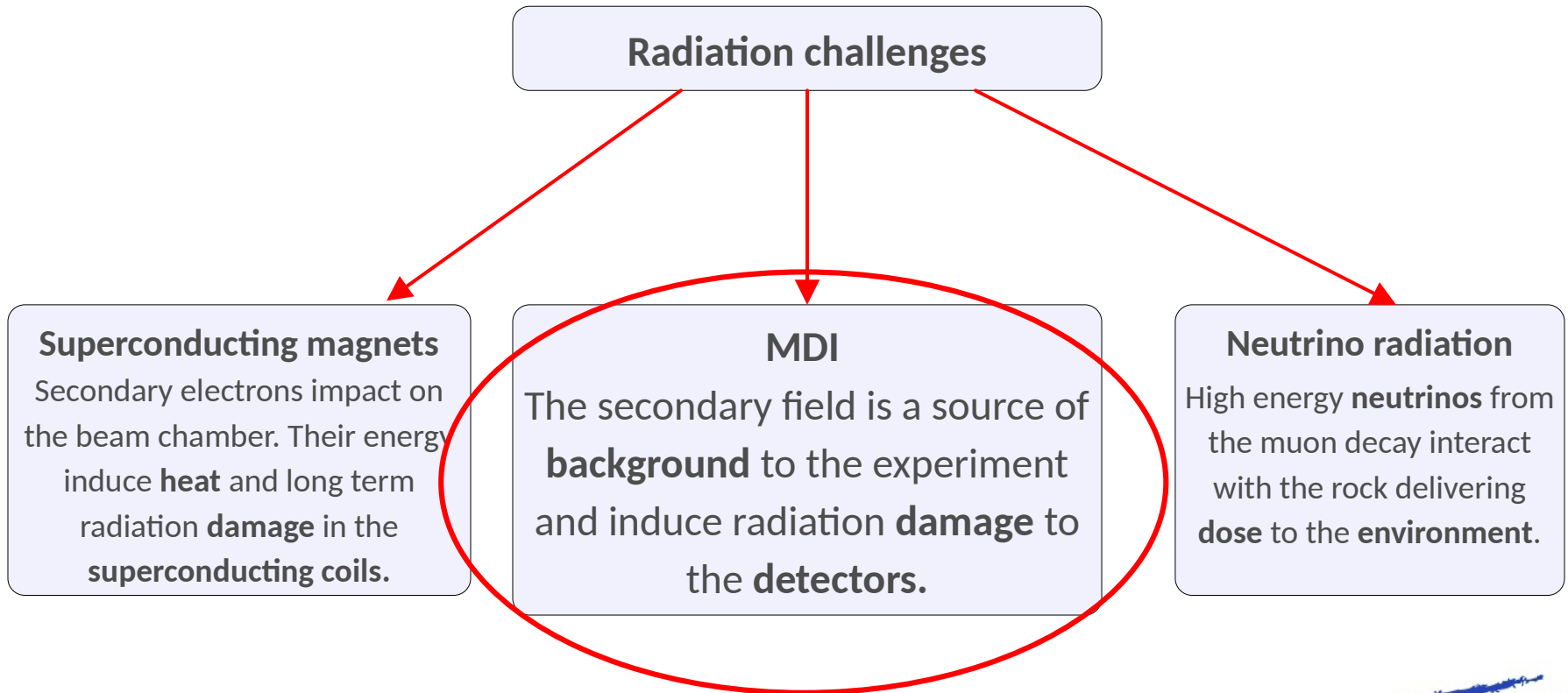
On behalf of the International Muon Collider Collaboration

# Outline

- **Muon collider (MC) radiation challenges**
- **Muon collider (MC) machine detector interface (MDI)**
  - Beam induced background (BIB): halo, muon decay and incoherent pair production by muons
  - Secondary electron losses on the aperture
  - Precedent work in the MAP collaboration
- **BIB for a 10 TeV machine**
  - Muon decay as main source of background and comparison with other machines
  - Incoherent pair production as a non negligible BIB
- **Lattice studies**
  - Lattice design influence on BIB
  - Muon decay in the chromaticity correction section
- **Nozzle design optimization**
- **Conclusions**

# Muon collider: radiation challenges

- **Muons** are **unstable particles**, with a rest lifetime of  $\tau = 2.197 \mu\text{s}$ . They decay spontaneously into electron and positrons (depending on the muon original charge), which are the main contributors to the secondary radiation field.



# Interaction region: MDI

- MDI is a **difficult challenge** for the muon collider. First studies were done by the MAP collaboration (energies up to 6 TeV). So far, no studies were performed for a 10 TeV collider.
- Objectives of the new studies within the IMCC:
  - Devise a conceptual IP design achieving **background levels compatible with detector operation**, both in terms of physics performance and acceptable cumulative radiation damage.
  - The focus energies are 3 TeV and 10 TeV.
- Starting from the **geometry** of the nozzle devised by the **MAP collaboration** [5], first MDI studies for colliders up to 10 TeV have been conducted.

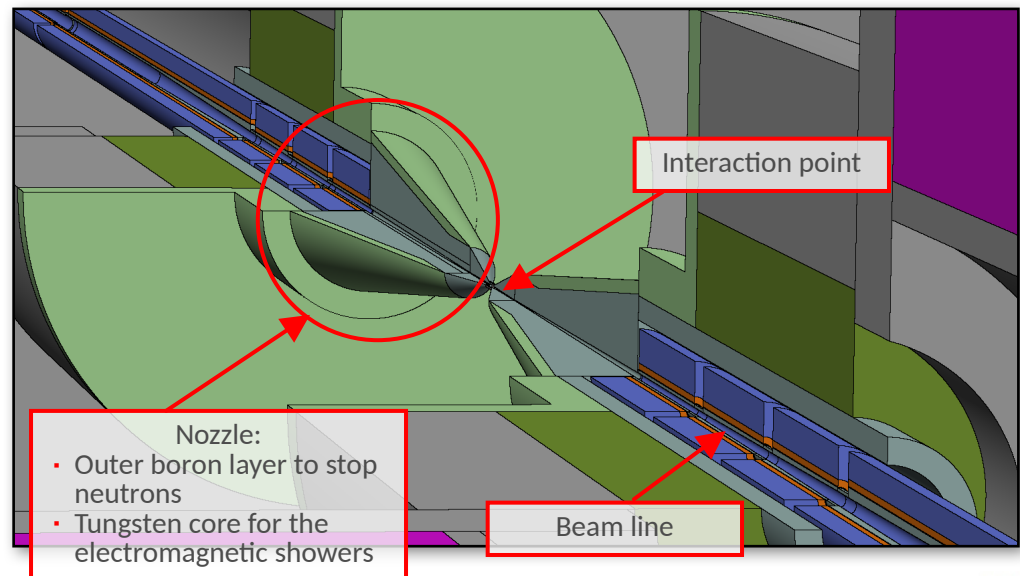
Parameters table

Tentative target parameters  
Scaled from MAP parameters

Comparison:  
CLIC at 3 TeV: 28 MW

Parameter	Unit	3 TeV	10 TeV	14 TeV
L	$10^{34} \text{ cm}^{-2}\text{s}^{-1}$	1.8	20	40
N	$10^{12}$	2.2	1.8	1.8*
$f_r$	Hz	5	5	5
$P_{\text{beam}}$	MW	5.3	14.4	20
C	km	4.5	10	14

## Geometry of the MDI



\*The results presented here are referred to 1 bunch crossing containing  $2E12$  muons

# MDI: radiation sources

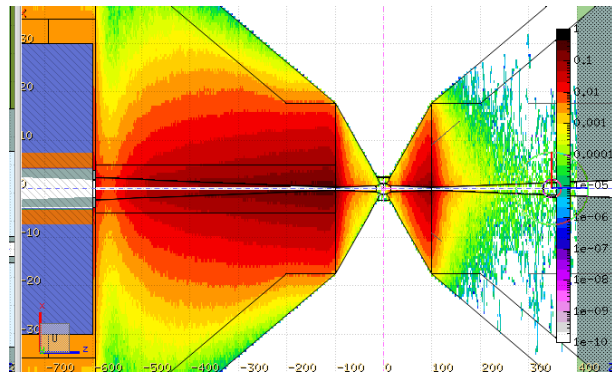
- Main source of detector **background** for all collider energy options.
- Main responsible for heat and radiation effects in the accelerator components.

- Potential contribution to the BIB and damage on accelerator components.
- Levels of acceptable halo losses to be defined. (**halo cleaning**)

- **Muon decay** around the ring
- **Incoherent  $e^-/e^+$  pair production** during bunch crossing in IP
- **Beam-halo losses** at aperture bottlenecks

- Potential problem for the detector **background**.
- Proven not to be an issue for low energy colliders, providing a **solenoid field** of  $\sim 1$  s T. [5].
- Under study in the 10 TeV collider.

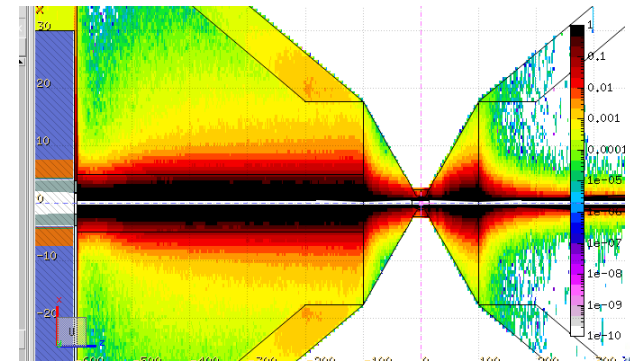
Neutron fluence



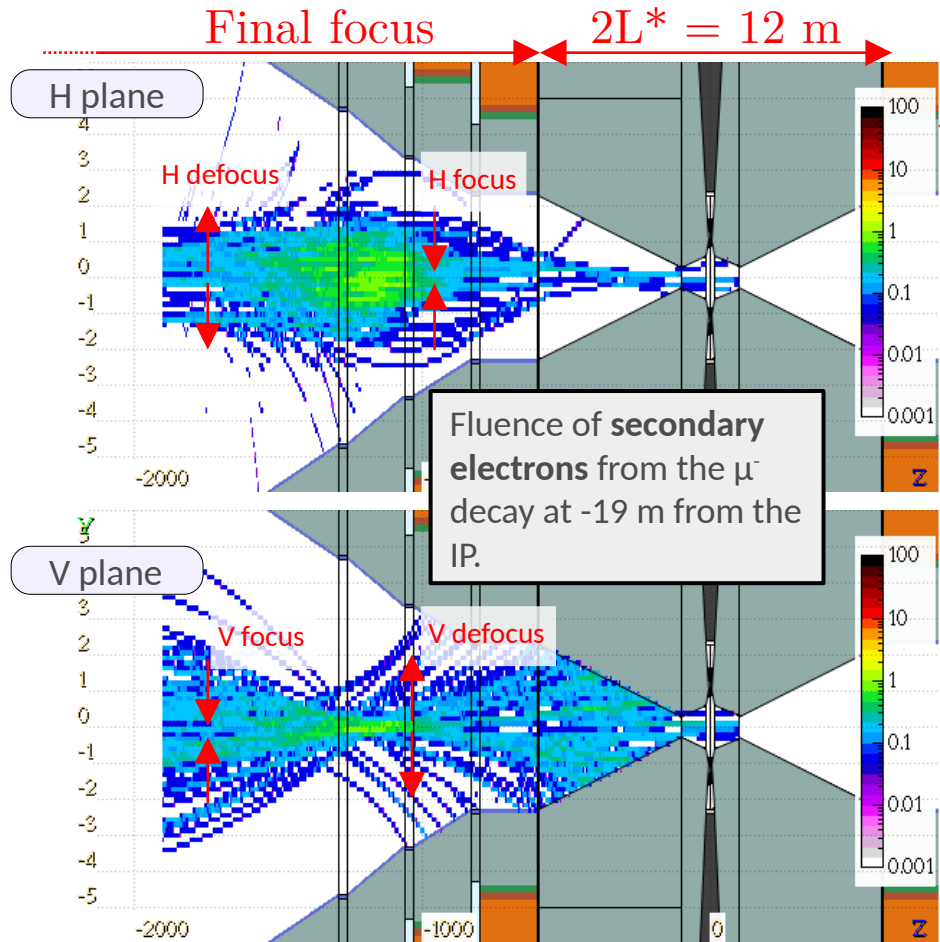
Effects

Secondaries will **interact** with the **machine components** and with the **detectors**. In figure, a thick **nozzle shielding** **protects** the **detector area** by the strong fluences arising from the muon decay.

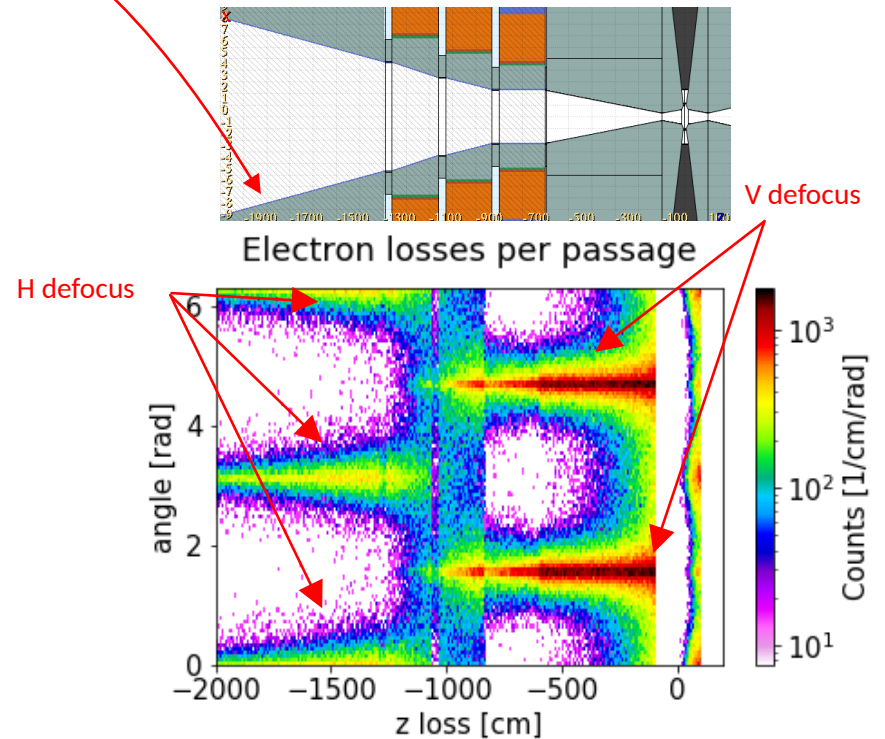
Photon fluence



# $e^+/e^-$ impact on aperture: qualitative view

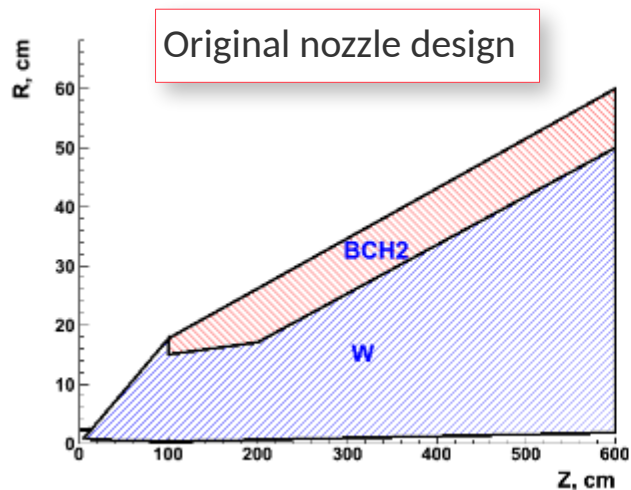


- Final focusing fields induce peaks in the azimuthal distribution of the  $e^-/e^+$  impact position.
- (but!) The azimuthal dependence is diluted to negligible levels by the W nozzle.



# MDI past results (MAP)

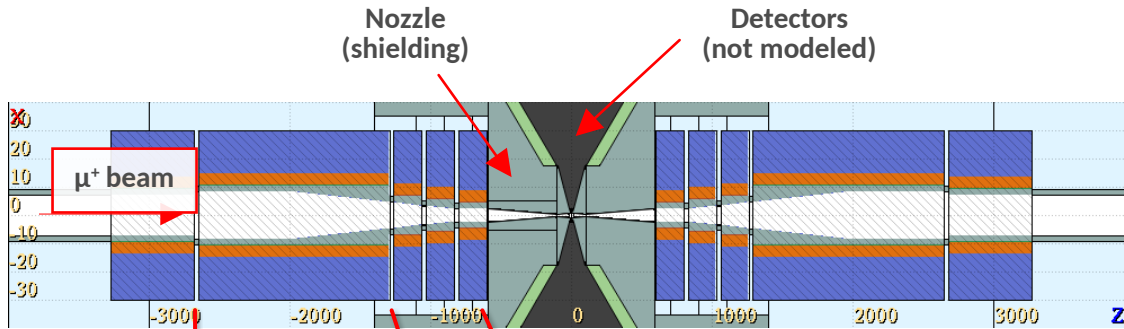
- In the context of the **MAP collaboration**, the muon collider detector background and Machine-detector interface has been thoroughly studied [5-8].
- They observed that most **background** particles are generated in the last **25 m straight section**, except **muons** that can be produced **further away**.
- The MAP collaboration optimized **nozzles** for colliders up to 1.5 TeV (with MARS code).
- Recent **FLUKA** results are in a **good agreement** with the past studies.



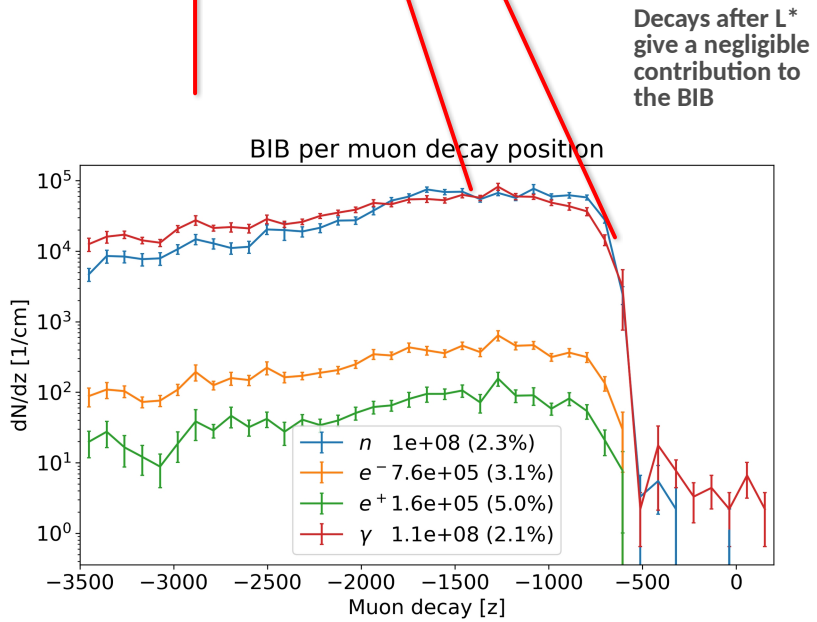
FLUKA/MARS15 results for the BIB of a 1.5 TeV muon collider from [9]

Particle ( $E_{th}$ )	MARS15	FLUKA
Photon (100 keV)	$8.6 \cdot 10^7$	$5 \cdot 10^7$
Neutron (1 meV)	$7.6 \cdot 10^7$	$1.1 \cdot 10^8$
Electron/positron (100 keV)	$7.5 \cdot 10^5$	$8.5 \cdot 10^5$
Ch. Hadron (100 keV)	$3.1 \cdot 10^4$	$1.7 \cdot 10^4$
Muon (100 keV)	$1.5 \cdot 10^3$	$1 \cdot 10^3$

# 10 TeV: BIB from muon decay (Final focusing)



- The 10 TeV geometry **has to be extended** to fully collect the BIB from decay position further away from the IP.
- The total number of particles entering into the detector is **not significantly worse** than the lower energy results!



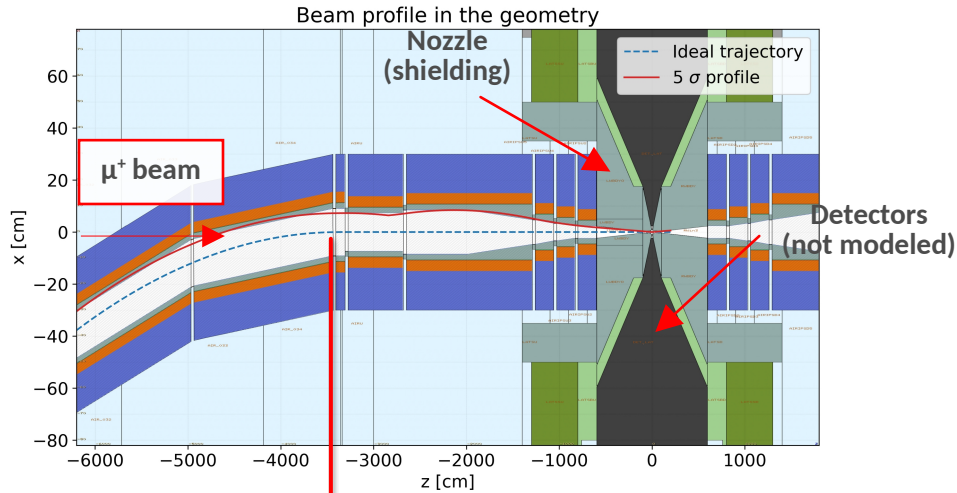
Total particle number: comparison with different collider energies

Collider energy	1.5 TeV	3 TeV	10 TeV
Photons	7.1E+7	9.6E+7	1.07E+8
Neutron	4.7E+7	5.8E+7	1.01E+8
e <sup>+</sup> /e <sup>-</sup>	7.1E+5	9.3E+5	9.6E+5
Ch. hadrons	1.7E+4	2.0E+4	4.3E+4
Muons	3.1E+3	3.3E+3	4.8E+3

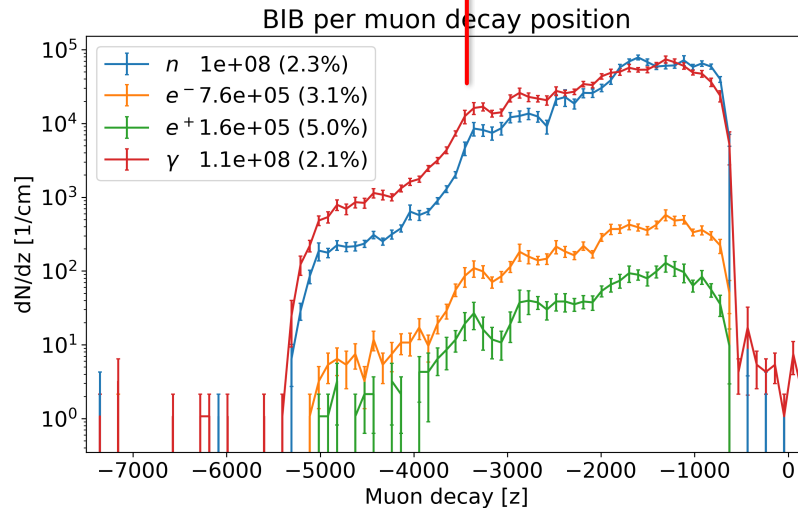
Non optimized [14-15]



# 10 TeV: BIB from muon decay (Final focusing and chromaticity correction)



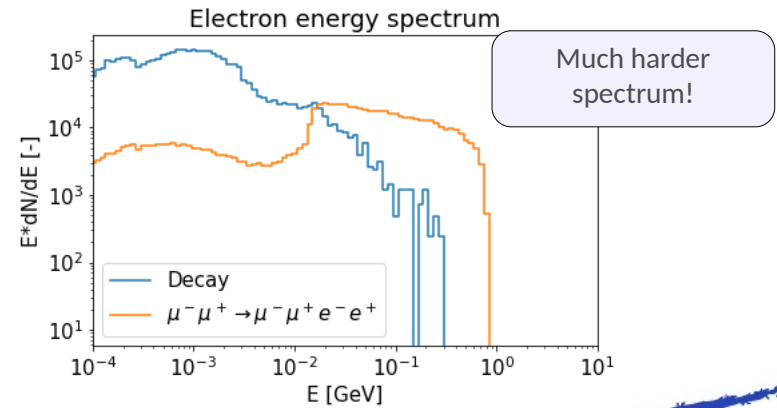
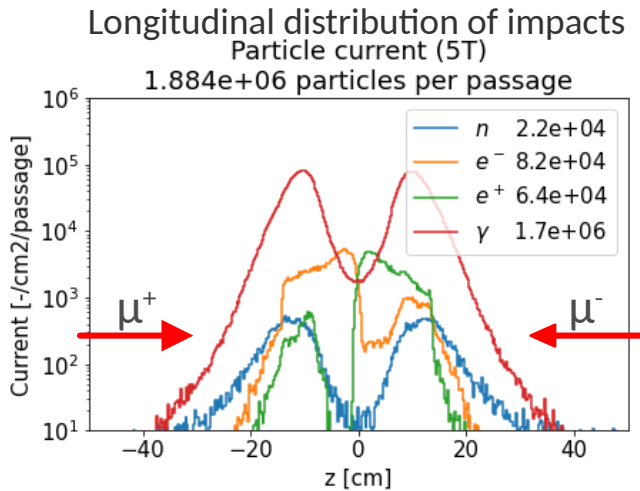
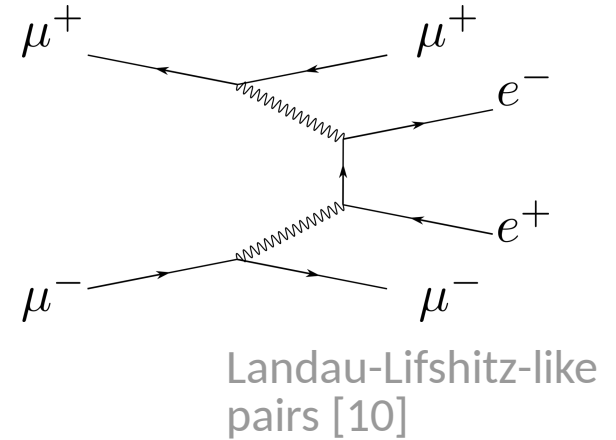
- The total number of particles entering into the detector is **not significantly worse** than the lower energy results!
- Before the final focusing magnet, there is a sharp decrease in the signal for what concerns most of the produced particles



Muon production still not statistically significant!

# 10 TeV: BIB from incoherent pair production

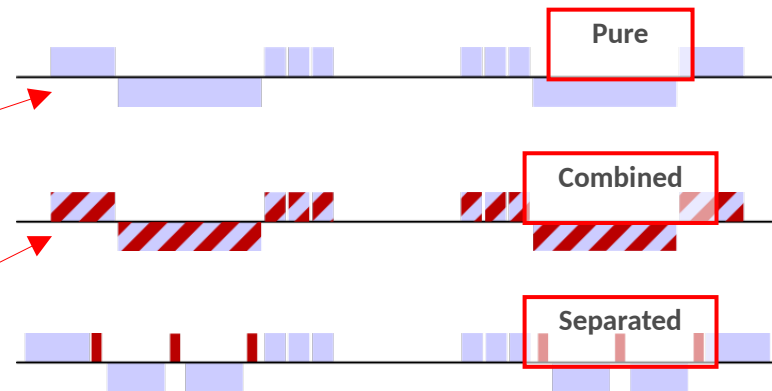
- At very high beam energies, beam-beam effects are not negligible. The most important phenomenon is due to the **incoherent beam-beam pair production  $\mu^+\mu^-\rightarrow\mu^+\mu^-e^+e^-$** .
  - The incoherent pair production  $e^+/e^-$  are provided by D. Schulte and are obtained by a **Guinea-Pig simulation**
- The **total number** of crossing is much **lower** than the muon **decay** case.
- The produced electrons are **energetic** and they **impact** directly on the **detectors**, since are generated in the IP, hence they might be dangerous despite the low total number.



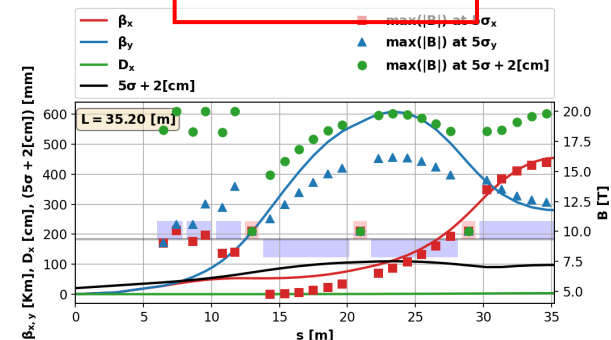
# 10 TeV: possible lattice design choices

- A first attempt to reduce the BIB is conducted working on the **lattice** just before the IP. In principle, having a **dipolar component** in the lattice is **beneficial**, since all the low energy electrons are forced to impact on the magnet sides.

- We considered three possibilities (from K. Skoufaris and C. Carli) for the lattice in the final focusing:
  - Only quadrupoles, with no dipoles and no dipole component (**pure**).
  - Combined function magnets, where there are no dipole magnet, but each quadrupole contains a 2T dipolar component (**combined**).
  - Having both dipoles and quadrupoles in the final triplet, but without exploiting combined function magnets. In this case we “separate” the dipolar component in short 10 T dipole magnets (**separated**).

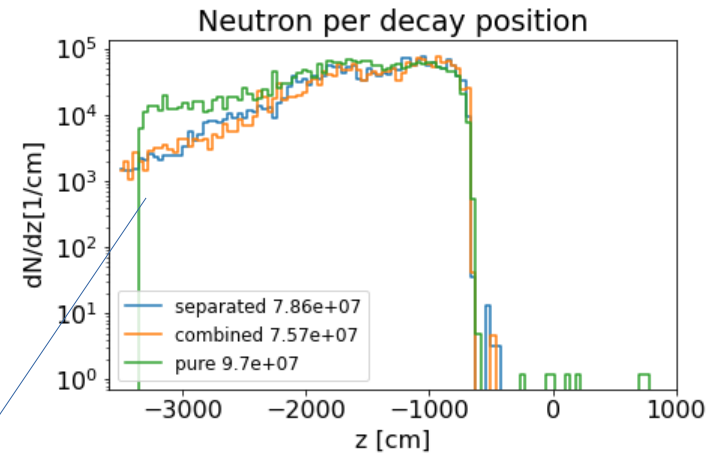
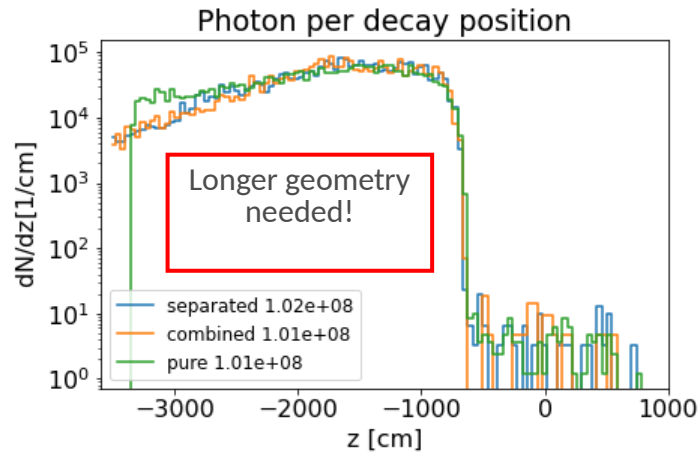


Twiss parameters of the separated lattice



# 10 TeV: lattice effects on BIB

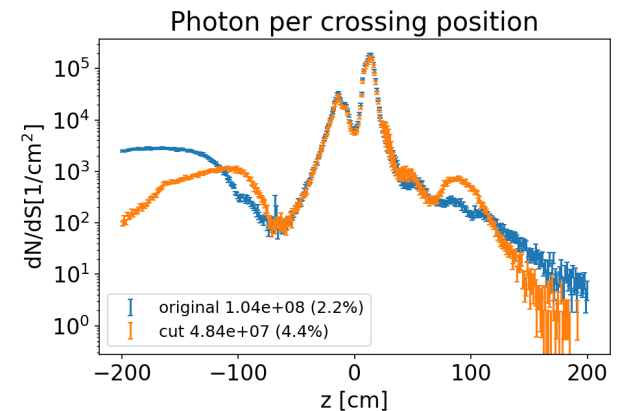
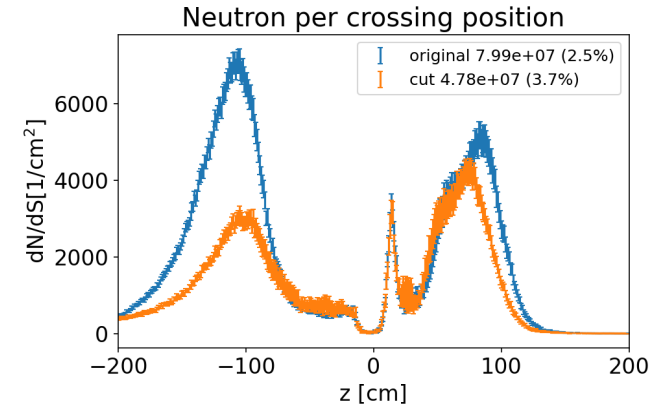
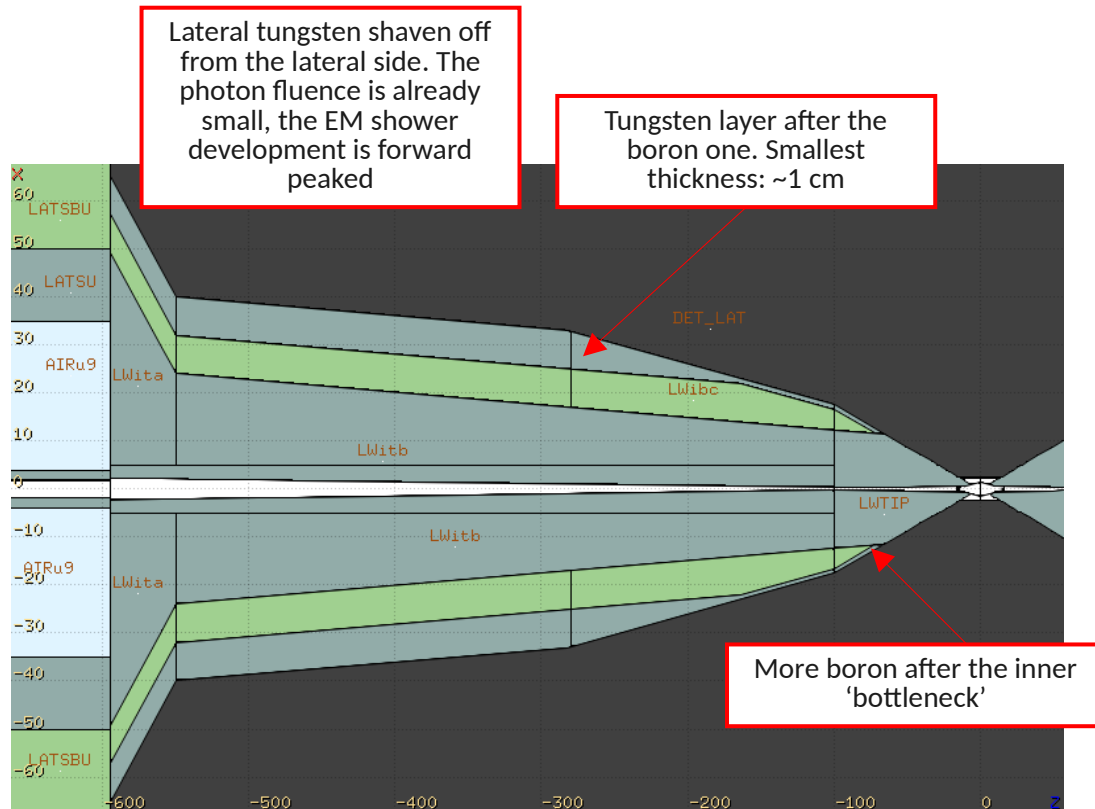
- The **contribution of different decay position** to the BIB for a positive muon beams is reported. As expected, the further away the decay occurs, the less background will arrive to the detector area.
- The overall capability to suppress BIB with **lattice design choices** is limited. Even if we **reduce slightly** the BIB from far away, other optimization means have to be found.



With a dipole component in the final focusing, small reduction far away from IP

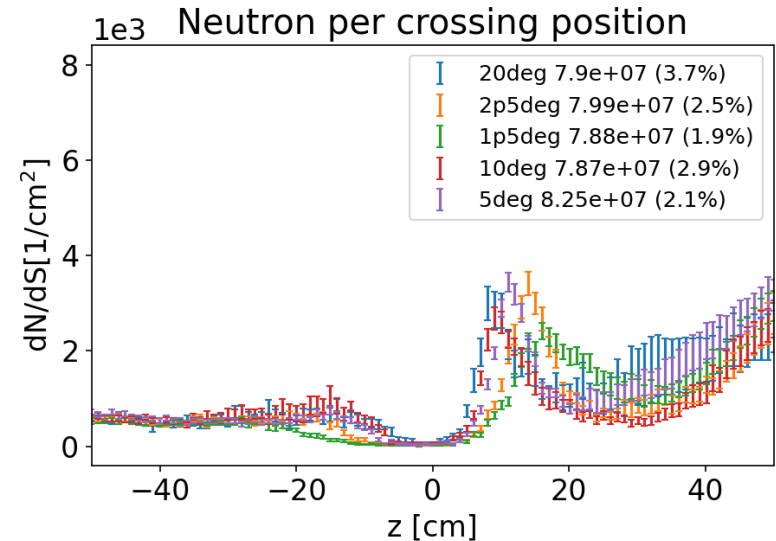
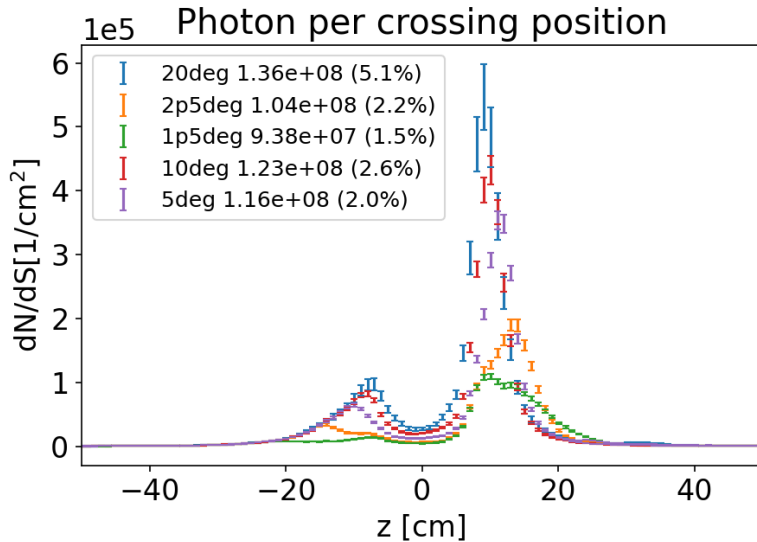
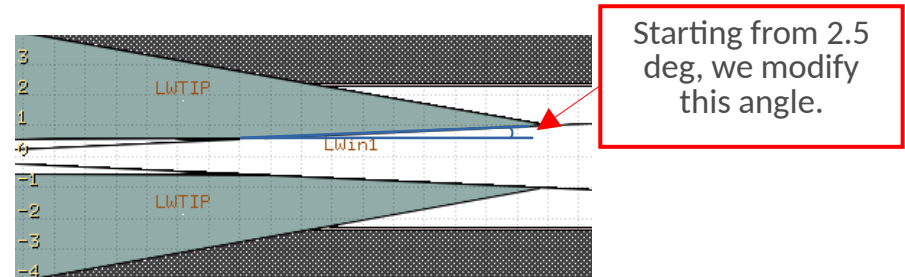
# Current nozzle optimization: nozzle shape

- Considering the particle fluences in the **nozzle**, a tentative nozzle geometry reshaping has been conducted.



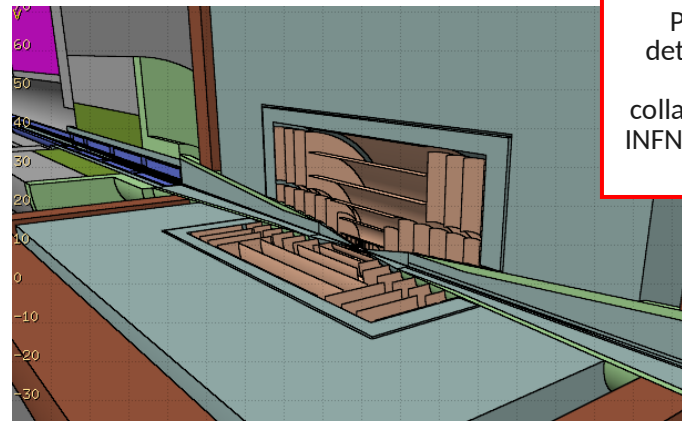
# Current nozzle optimization: angle tip

- Considering the aperture of the **nozzle**, various **angles** have been tested. The scope of the optimization of these parameters, is not to reduce the overall number of particles going into the detectors, but to **reduce their peaks**.
- The results shows a clear advantage to **reduce the tip angle** down to very small values.

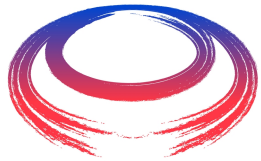


# Conclusions

- **Muons decay** induces an intense secondary radiation field in all component of the muon collider. The MDI design is vital to mitigate the phenomenon.
- The situation with the **high energy option** (10 TeV) is **not significantly worse** in comparison with the 3 TeV collider.
- Different lattices do not significantly alter the BIB from muon decay in close proximity with the final focusing, while changing the nozzle shape alters the background in a more substantial way.
- Contributions from decay further away from the final focus region are not significant.
- At 10 TeV the incoherent pair production from muon is a non negligible source of radiation, while with lower energies this phenomenon is mitigated by the solenoidal magnetic field.
- **Next steps:**
  - Simulate the BIB in the case of longer  $L^*$  (10 m instead of presently 6 m)
  - Continue the optimization of the nozzle design at different energies
  - Detectors response and radiation damage shall be studied



Preliminary  
detector design  
taken in  
collaboration with  
INFN from the CLIC  
layout



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**Thank you for the  
attention!**



# References

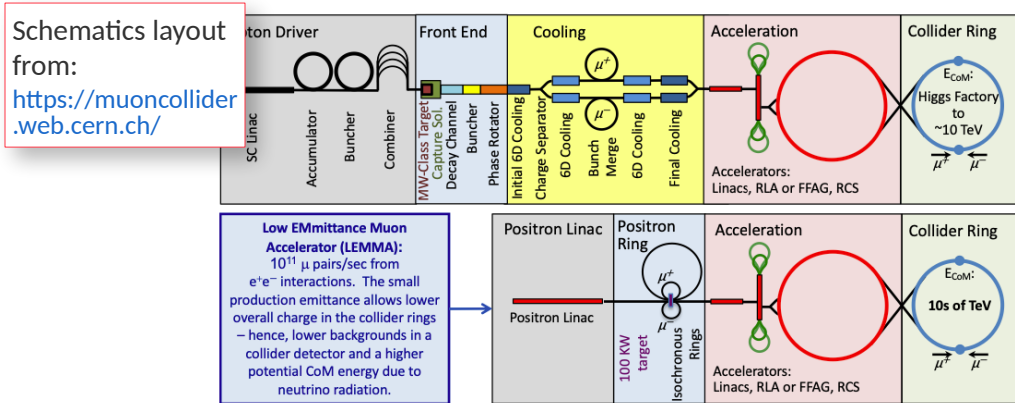
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- [14] [https://indico.cern.ch/event/1134938/contributions/4765158/attachments/2402421/4117427/BIB\\_CCuratolo\\_4mar2022.pdf](https://indico.cern.ch/event/1134938/contributions/4765158/attachments/2402421/4117427/BIB_CCuratolo_4mar2022.pdf)
- [15] [https://indico.fnal.gov/event/51315/contributions/225846/attachments/148314/190521/casarsa\\_BIBcomparison.pdf](https://indico.fnal.gov/event/51315/contributions/225846/attachments/148314/190521/casarsa_BIBcomparison.pdf)

# Muon collider: concept and motivations

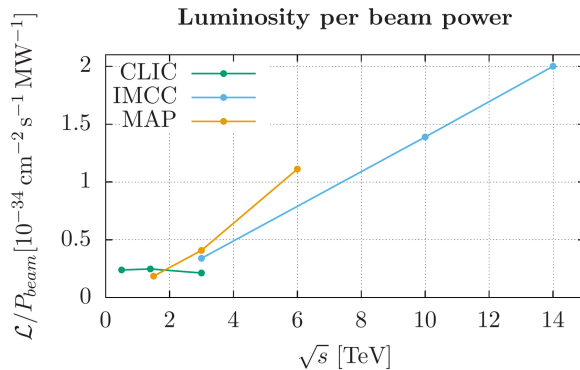
- Among various particles accelerated in colliders, muons have already been under consideration for a long time [1]. Very promising results were achieved in the contest of the **MAP collaboration** [2-3]. The following work is in the context and on behalf of the **International Muon Collider Collaboration (IMCC)**.

**Why?**

- A multi-TeV muon collider could investigate **Higgs properties** with an unprecedented precision. [2]
- With  $\sqrt{s} = 10$  TeV we can explore **new physics at high energies**. [2]



With a muon collider the **luminosity per beam power** increases with the collider energy!



$$\mathcal{L} \propto \gamma \langle B \rangle \sigma_{\delta} \frac{N_0}{\epsilon \epsilon_L} f_r N_0 \gamma$$

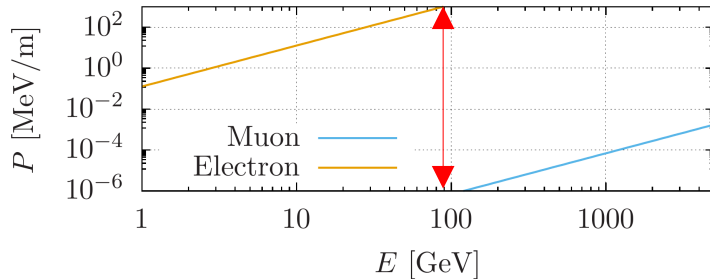
High energy (points to  $\gamma$ )  
 High field in collider ring (points to  $\langle B \rangle$ )  
 Large energy acceptance (points to  $\sigma_{\delta}$ )  
 Dense beam (points to  $\frac{N_0}{\epsilon \epsilon_L}$ )  
 High beam power (points to  $f_r N_0 \gamma$ )

# Muon collider: advantages

## Synchrotron radiation\*

- The muon mass: **105.7 MeV/c<sup>2</sup>**. **Synchrotron radiation (SR) is not a limiting factor** for muon circular colliders.

Energy emitted by SR per unit length

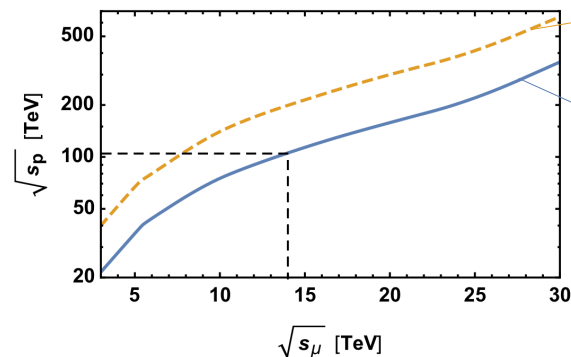


Muons emit  $(m_\mu/m_e)^4 = 1.6 \cdot 10^9$  less synchrotron radiation than electrons

## Lepton collisions

- Muons, as leptons, are elementary particles, and they allow collision where the entire center of mass energy is involved (in proton collision the energy is shared among constituents)
- Same performance of proton colliders, but with much lower center of mass energy! [2]**

Energies at which proton/ $\mu$ -colliders have similar performances



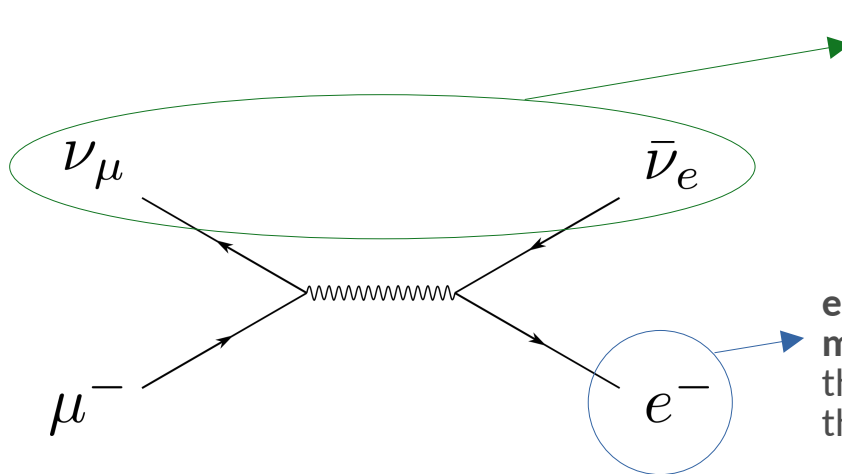
Assuming comparable Feynman amplitudes for muon and proton production processes

Proton production enhanced due to QCD production (factor of 10)

\* of the primary muon beam

# Muon collider: radiation challenges

- Muons are unstable particles**, with a rest lifetime of  $\tau = 2.197 \mu\text{s}$ . They decay spontaneously into electron and positrons (depending on the muon original charge).

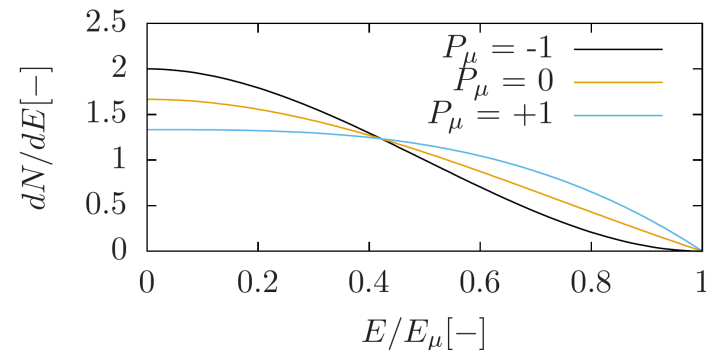


**Neutrinos:** they hardly interact with the accelerator component, therefore little concern for the beam-machine interaction. The only concern is due to **dose delivered to the environment** outside the surface.

**e<sup>-</sup>/e<sup>+</sup>:** they carry around **1/3 of the original muon energy** and they are responsible for the heat load and the **radiation damage** of the **accelerator components**.

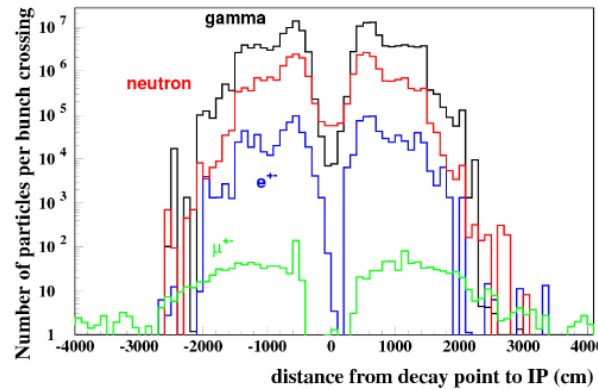
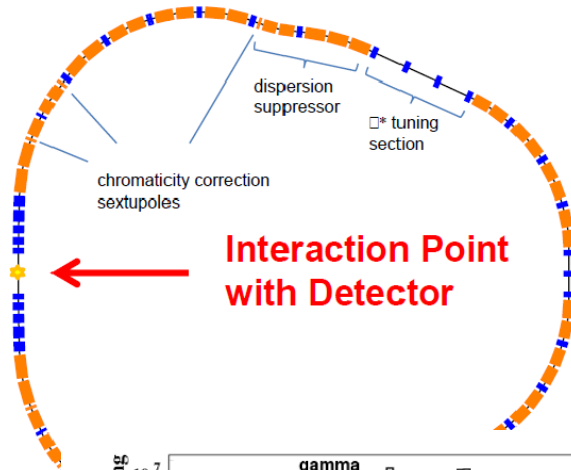
**Original muon:** thanks to the Lorentz boost, it will survive for  $\gamma\tau$ . In any case, the muon production/acceleration/collision must be **extremely fast**.

Electron energy spectrum



# Higgs factory

125 GeV: Higgs' factory [11,12]



# 1.5 TeV spectra

