

# Neutrino Radiation for a realistic Collider



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## Introduction Neutrino Radiation Issue



Muons decay (say in

 Radiation due to neutrino beam reaching the earth surface

- Narrow radiation "cone" for a short piece of the machine
- Showers from neutrinos interacting close to earth surface generate dose seen at surface
- Matter in front ("shielding") does not help but makes situation even worse
- Strong increase of maximum dose with muon energy
  - Cross sections about proportional to energy
  - Typical energy per interaction of neutrino with matter proportional to muon energy
  - Opening of radiation cone inversely proportional to muon energy







### **Analytical estimates** absorbed dose per decay



(see, e.g., arXiv:physics/9908017)

straigt section length Ds

 $d = L_s D \mathcal{J}_H = L_s \left( \mathcal{J}_H - \mathcal{J}_H \right)$ 

- Consider say decay  $M^{-} \rightarrow e^{-} + \overline{N}_{e} + N_{m}$
- In muon rest system: anisotropic neutrino distribution with known energy distribution
- Lorentz boost to obtain neutrino direction and energies in lab system (Lorentz factor g)
- Sum over

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- Both neutrino types
- cross sections of possible interactions times fraction of energy deposited per interaction in shower

• Gives  

$$\Delta D = \left( \underbrace{0.410 \cdot 10^{-39} \frac{\text{m}^2}{\text{TeV}} \frac{E_{\mu r}^2}{4}}_{1.10 \cdot 10^{-28} \text{Gy m}^2} \frac{1}{m_u} \frac{1}{L_s^2} \frac{\langle E_v^2 \rangle}{\langle E_v^{\prime 2} \rangle} \frac{1}{N} \frac{dN}{d\Omega} = \left( 1.10 \cdot 10^{-28} \text{ Gy m}^2 \right) \frac{4\gamma^4}{\pi L_s^2} \frac{1}{\left( 1 + \gamma^2 \left( \frac{d}{L_s} \right)^2 \right)^2} \frac{1}{N} \frac{dN}{d\Omega} = \left( 1.10 \cdot 10^{-28} \text{ Gy m}^2 \right) \frac{4\gamma^4}{\pi L_s^2} \frac{1}{\left( 1 + \gamma^2 \left( \frac{d}{L_s} \right)^2 \right)^2} \frac{1}{N} \frac{dN}{d\Omega} = \left( 1.10 \cdot 10^{-28} \text{ Gy m}^2 \right) \frac{4\gamma^4}{\pi L_s^2} \frac{1}{\left( 1 + \gamma^2 \left( \frac{d}{L_s} \right)^2 \right)^2} \frac{1}{N} \frac{dN}{d\Omega} = \left( 1.10 \cdot 10^{-28} \text{ Gy m}^2 \right) \frac{4\gamma^4}{\pi L_s^2} \frac{1}{\left( 1 + \gamma^2 \left( \frac{d}{L_s} \right)^2 \right)^2} \frac{1}{N} \frac{1}{N} \frac{dN}{d\Omega} = \left( 1.10 \cdot 10^{-28} \text{ Gy m}^2 \right) \frac{1}{N} \frac$$

with  $\langle E_n^2 \rangle / \langle E_n^{i2} \rangle = 4g^2 / (1 + g^2 J^2)^2$  the ratio between the square of the neutrino energy in the lab and the rest system and  $N^{-1} dN/dW = (g^2/p)/(1+g^2 J^2)^2$  the angular distribution of neutrinos

• The decay  $\mathcal{M}^+ \rightarrow e^+ + \mathcal{N}_e + \overline{\mathcal{N}}_m$  gives (almost) identical final result despite different cross sections and fractions



### Analytical estimates approximation by Gaussian



 Absorbed dose reasonably well described by Gaussian

$$DD \gg (1.10 \times 10^{-28} \text{ Gy m}^2) \frac{4g^4}{\rho L_s^2} \frac{1}{(1 + g^2 (d/L_s)^2)^2}$$

$$\approx \left(1.10 \cdot 10^{-28} \text{ Gy m}^2\right) \frac{4g^4}{\rho L_s^2} e^{-3g^2 \left(\frac{d}{L_s}\right)^2}$$

- Rms opening angle of neutrino radiation cone of  $J_{rms} = 1/(\sqrt{6}g)$
- Assuming Gaussian for the beam divergence
  - Folding of divergence from muon decay process and beam divergence simple
- Almost suitable as source term to estimate doses generated by neutrino interactions





## FLUKA simulations of doses due to neutrino interactions

CERN

- Motivation: improvement and check of analytical derivations
  - Possible widening of neutrino radiation cone due to lateral extension of shower
  - Effective dose instead of absorbed dose

• Ansatz  

$$D\hat{H} \approx D\hat{H} e^{-d^{2}/2S_{DH}^{2}}$$

$$= w_{R,eff} \cdot (1.104 \cdot 10^{-28} \text{ Gy m}^{2}) \frac{4g^{4}}{\rho(L_{s}^{2} + 6g^{2}S_{s}^{2})} \exp\left(-\frac{d^{2}}{2(L_{s}^{2}/6g^{2} + S_{s}^{2})}\right)$$

$$S_{DH} \approx \sqrt{L_{s}^{2}/(6g^{2}) + S_{s}^{2}}$$

$$D\hat{H} S_{DH}^{2} \approx w_{R,eff} \cdot (1.104 \cdot 10^{-28}) 2g^{2}/(3\rho)$$

#### FLUKA studies by G. Lerner et al.

- with w<sub>R,eff</sub> to convert absorbed dose D to dose equivalent H = w<sub>R,eff</sub>D
- and S<sub>s</sub> the rms radial extension of the shower (assumed as well to be Gaussian?!)
- Result of fitting "by hand" to FLUKA results for 5 TeV muons (green lines in plot)
  - $w_{R,eff} = 1.3 \text{ Sv/Gy}$  and  $S_s = 0.12 \text{ m}$
  - Results may be different for other energies
  - Shower extension s, neglected for further studies (slightly pessimistic)





## **Extrapolation to a collider lattice**



length Ds





### Numerical evaluations Simple cases



- Divergence of muon beam neglected, peak dose rate
  - Straight section Ds , using  $C = 2\rho g E_{rm} / (c e \overline{B})$  and  $L_s \gg 2R_e h$  with  $R_e \gg 6.38 \times 10^6$  m the earth radius  $\frac{dH_s}{dt} = \left(6.85 \cdot 10^{-22} \frac{\text{Sv}}{\text{T}}\right) f_r N_m \left(\frac{E}{5 \text{TeV}}\right)^3 \frac{\text{Ds} \overline{B}}{h}$  for beam energies around  $E \gg 5 \text{TeV}$
  - Bending magnet integration w.r.t s using  $\mathcal{J}_{H} \hat{\mathcal{J}}(s) = \left(\frac{eB}{gE_{rm}}/c\right)s$

$$\frac{dH_B}{dt} = \left(6.85 \cdot 10^{-22} \frac{\mathrm{Sv}}{\mathrm{T}}\right) f_r N_m \left(\frac{E}{5\mathrm{TeV}}\right)^3 \frac{\overline{B}}{h} \int ds \, \exp\left(-3\left(\frac{eBc}{E_{rm}}s\right)^2\right) = \left(2.47 \cdot 10^{-22} \,\mathrm{Sv} \,\mathrm{m}\right) f_r N_m \left(\frac{E}{5\mathrm{TeV}}\right)^3 \frac{\overline{B}}{hB}$$

 Integrated peak equivalent dose per muon beam for one year operation (5000 h = 18 10<sup>6</sup> s) without mitigation measures

Contribution only from of the two muon beams

Mitigation mandatory!!

- $N_m = 1.8 \times 10^{12}$  muons per bunch,
- $f_r = 5$ Hz repetition,
- $E \gg 5 TeV$  beam energy,
- $C = 10000 \, \text{km}$  circumference and
- $\overline{B} = 10.42 \mathrm{T}$  average field

h (m)	L <sub>s</sub> (km)	H <sub>s</sub> (mSv) for ∆s = 0.3 m	H <sub>B</sub> (mSv) for B = 8 T	
100	35.7	3.5	0.52	
200	50.5	1.75	0.26	
500	79.7	0.70	0.105	
784	100	0.45	0.067	





## Numerical evaluations equivalent dose from arc cell at 100 km



- Integrals evaluated for present (work in progress by K. Skoufaris) 10 TeV collider arc half cell
- In collider mid-plane as function of  $\mathcal{J}_H$  (i.e.,  $\mathcal{J}_V = 0$ ) for one year (5000 h operation)





## Mitigation by "Wobbling"



- Wobbling of machine in vertical direction part of MAP proposal?
  - Time-dependent mechanical deformation of ring around arc (including chromatic compensation, matching section and FMC arc cells
  - High precision movement system
  - Impact on optics?
- For 10 TeV com collider with 10 km circumference and say 4.8 km arcs



- Combination of pieces of parabola two pieces with opposite curvature one period
- Say 8 periods 660 m long periods generating angles between -1 mrad and + 1 mrad
- Magnetic field (average) bending in vertical ±0.11 T
- Excursion (maximum total) ±150 mm
- Replaces vertical Gaussian angle distribution with rms opening of ≈0.0086 mrad by about rectangular distribution within ±1 mrad
- => About two order of magnitude reduction of peak dose rates



# **Summary and Outlook**



- Dose generated at the earth surface due to showers generated by neutrino interactions is a serious issue
  - Muon decays generate narrow cone swept by bendings in bending plane
  - In particular high doses in direction of (short) straights sections
  - Numerical estimates based on "source term" from FLUKA simulations (=> Comparison with MARS results to be done?)
- Time-dependent vertical deformation of the whole arc proposed as mitigation measure
  - For a 10 TeV com collider angle variations (linear up and down) in a range ±1 mrad would allow to gain about two orders of magnitude
  - To be combination with other mitigation measures
    - Collider installed deep underground with suitable positioning
    - Appropriate orientation towards uncritical areas (possibly owned by facility)
- Feasibility to be studied
  - High precision movement system
  - Impact on beam dynamics in particular in chromatic compensation and matching sections (vertical dispersion, orbit stability ...)