

# the T-REX project: Micromegas readouts for Rare Event Searches

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# Outline

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T-REX: MM readouts for rare event searches

MM for axion searches: CAST

MM for  $\beta\beta$  searches: NEXT

Radiopurity Measurements

Conclusions

# T-REX

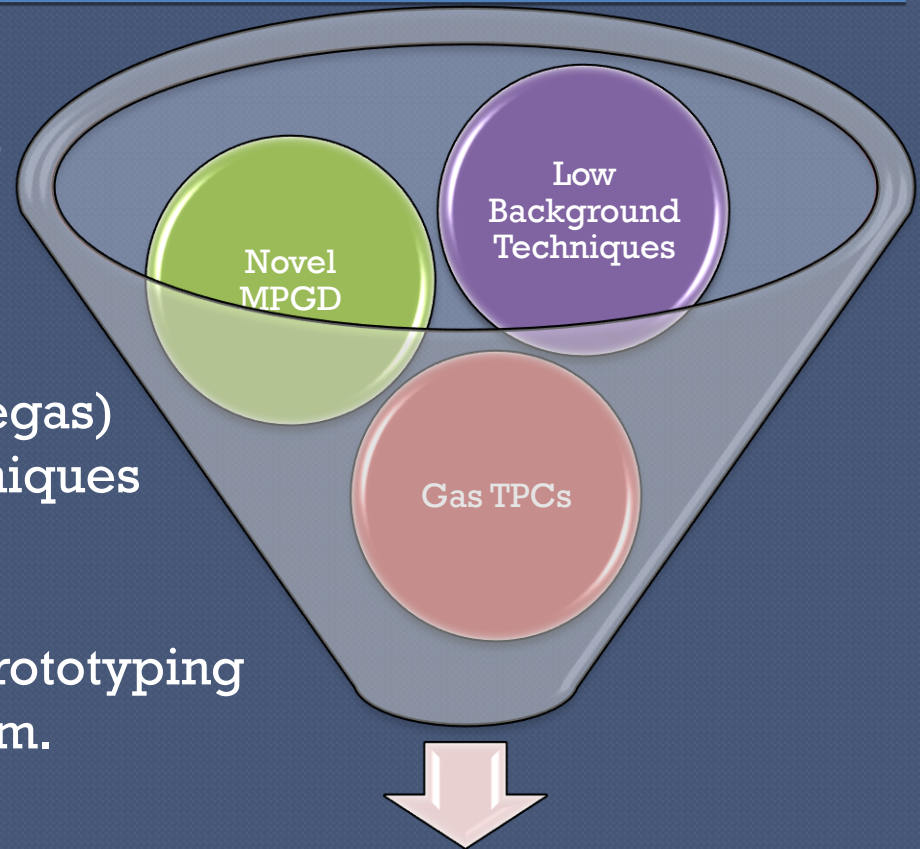
Rare Event Searches:

**topology** can be the key

Gas TPCs can do it

T-REX: merge MPGDs (Micromegas)  
+ low background techniques

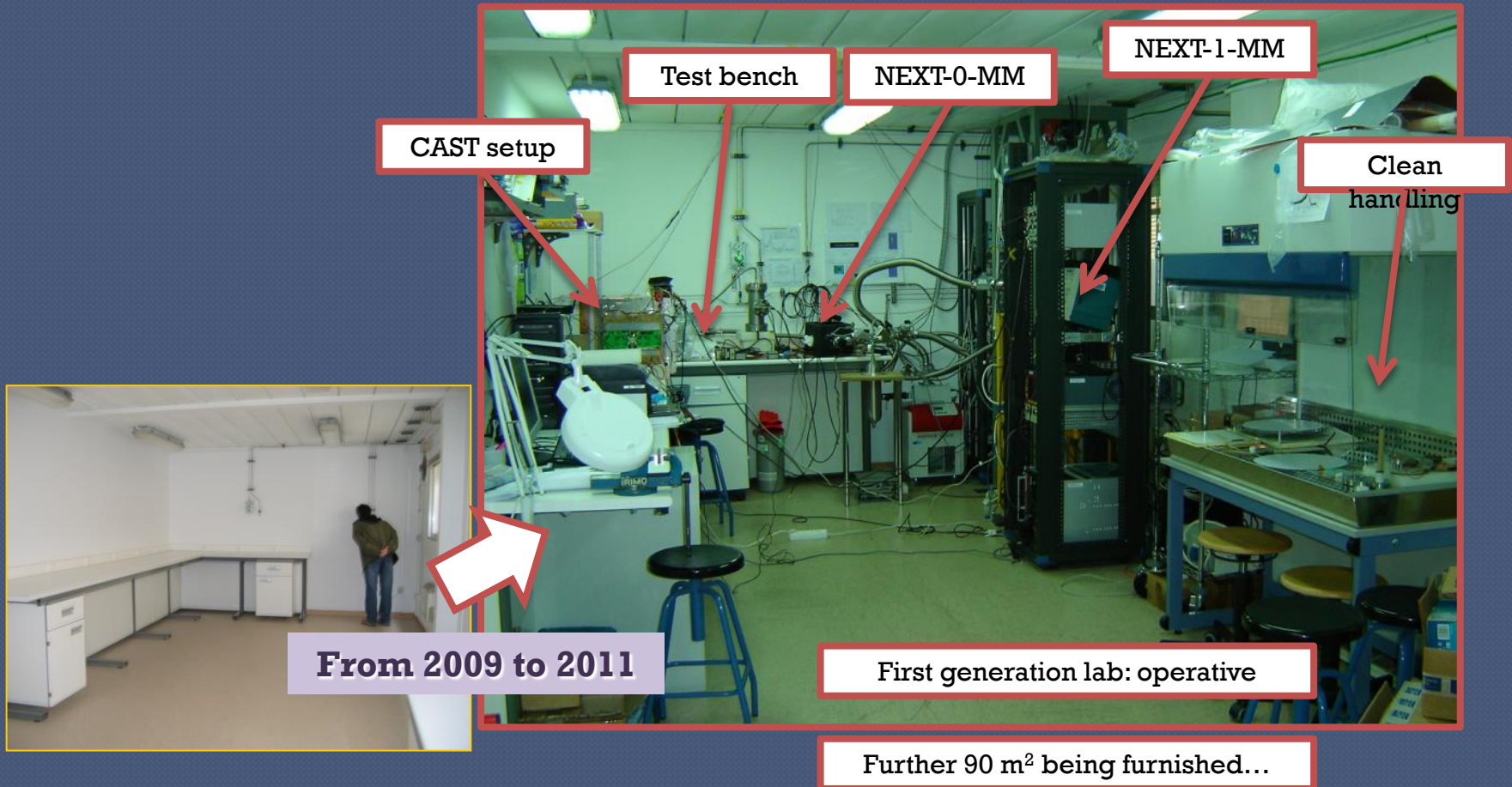
Focus on R&D and small scale prototyping  
ERC St-G funded. IDEAS program.



**New Generation of Rare  
Event Searches**

# T-REX: infrastructure

New gas detector lab at UNIZAR



# Micromegas

Two-region gaseous detector:

Conversion region

Primary ionization  
Charge drift

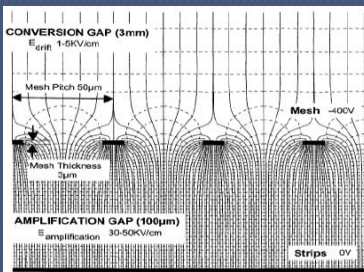
Amplification region

Charge multiplication  
Readout layout

- Strips (1 or 2 D)
- Pixels

*Separated by a Micromesh:*

Very strong and uniform electric field

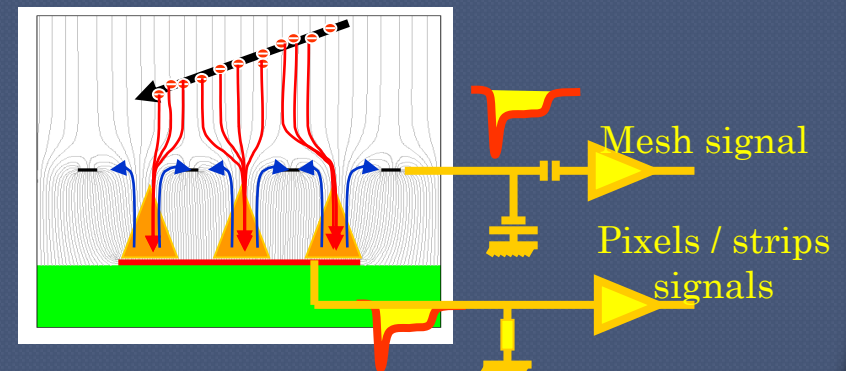
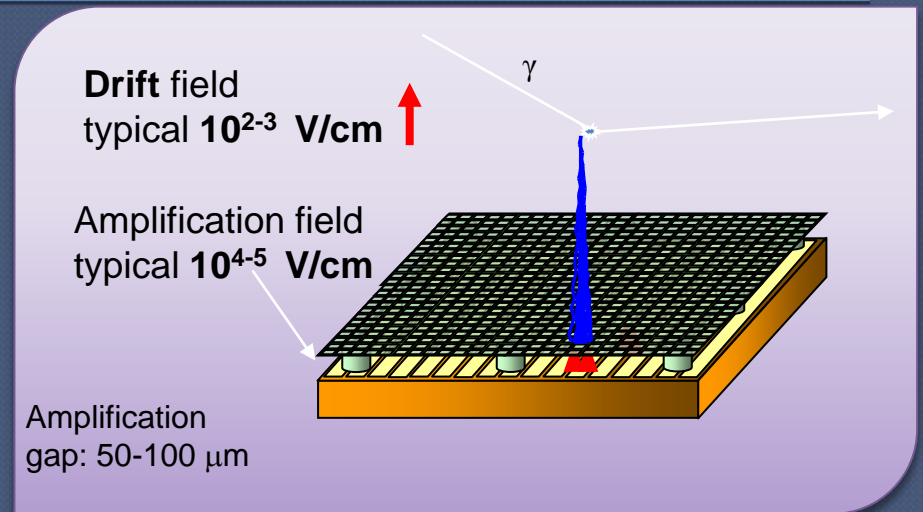


*Giomataris et al. (1996)*

## Advantages:

Simplicity  
Granularity  
Homogeneity  
Large areas

...



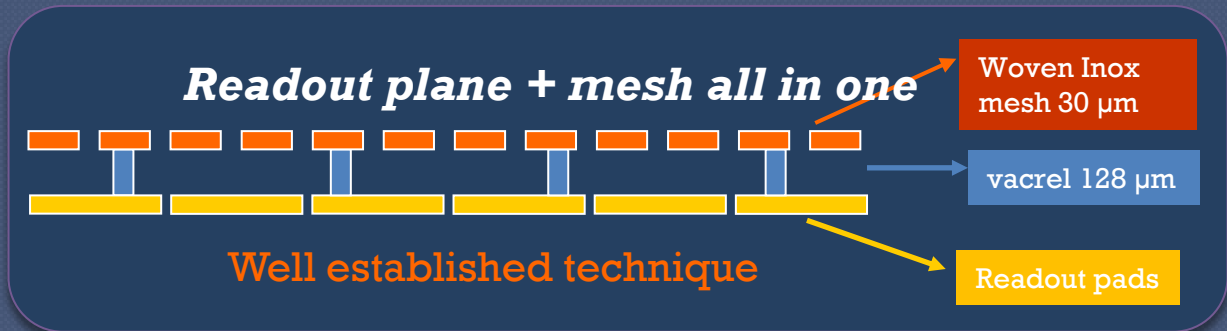
# Micromegas

**Bulk & microbulk** techniques developed for all-in-one fabrication

- Ease of operation
- Large areas

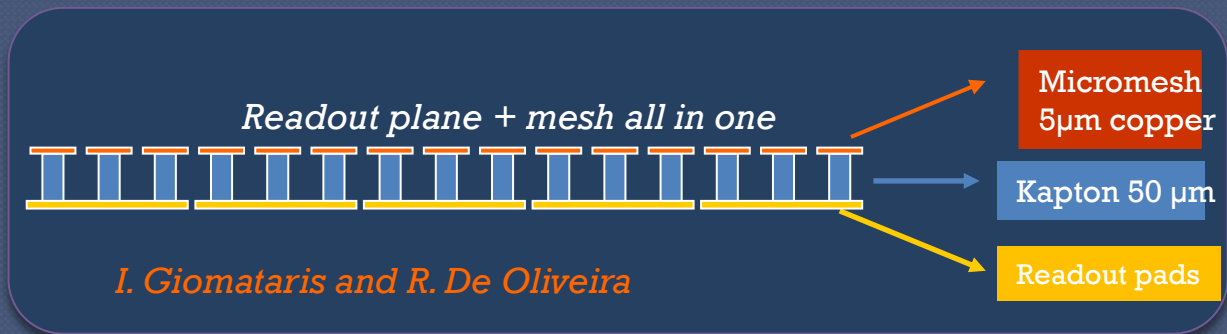
## **BULK**

Robust  
Mature  
Very large areas  
available (2 m<sup>2</sup>)



## **MICROBULK**

Higher  
homogeneity  
Light weight,  
radiopure



See also  
2010 JINST 5 P02001

# $\mu$ M for axion searches: CAST

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S. Aune et al., 2009 NIMA604 15-19

2009 JPCS179 012015

J. Galán et al., 2010 JINST 5 P01009

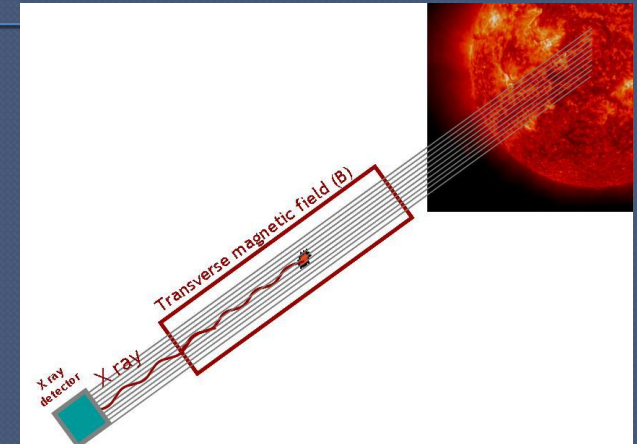
S. Andriamonje et al., 2010 JINST 5 P02001

T. Dafni et al., 2011 NIM A 628 172-176

presented at:

Blois 2010, Vienna Instrumentation 2010, EXRS2010, TPC Symposium 2010, Moriond2010

# CAST experiment @ CERN



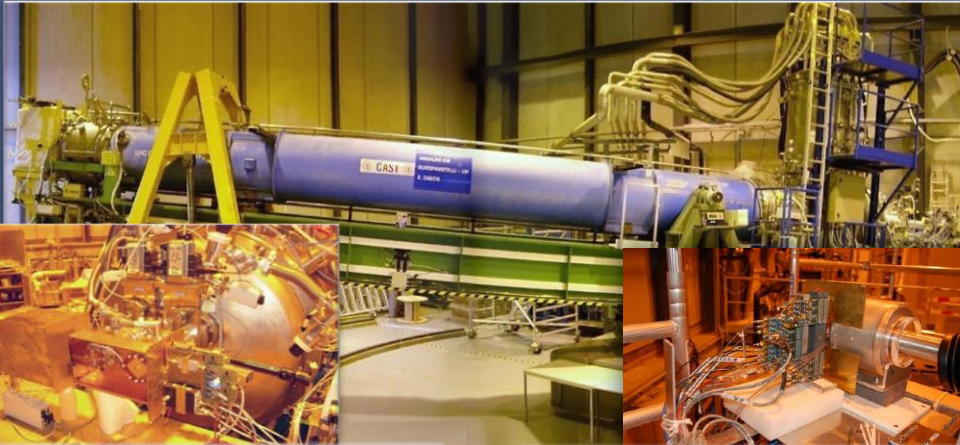
**Axions:** hypothetical particles solving the strong CP problem, and possible dark matter candidates. Produced by the Sun by **photon-to-axion** conversion of the solar plasma photons Detectable via "**helioscope concept**" (Sikivie, 83)

very low background X-ray detectors are necessary  
CAST uses

- 3 microbulk micromegas
- a CCD-X-ray focusing device system



# Microbulks in CAST



2 microbulks at SUNSET

1 microbulk at SUNRISE

Low intrinsic radioactivity:

Light mass, clean materials (copper, plexiglass,...)

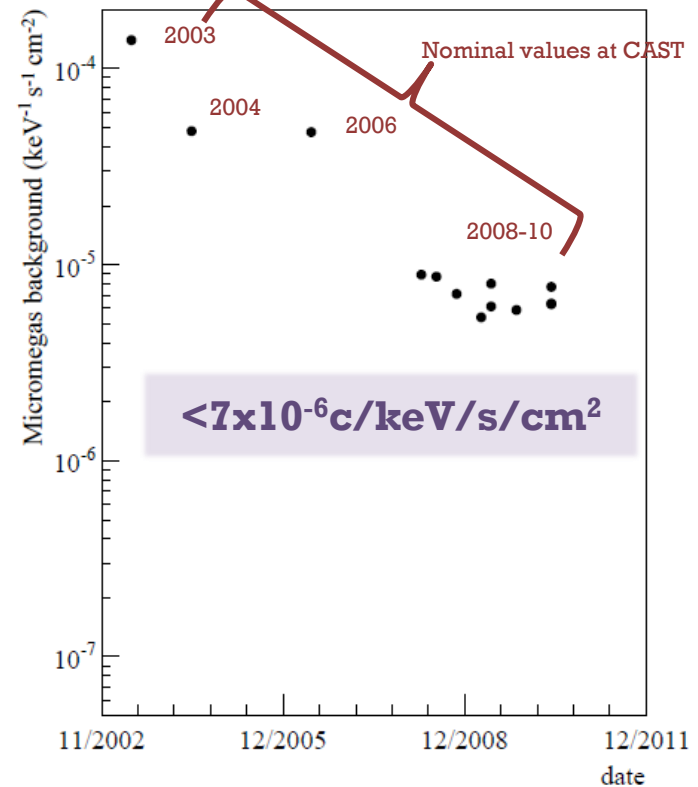
Signal topology → offline analysis

2D readout pattern, Time information

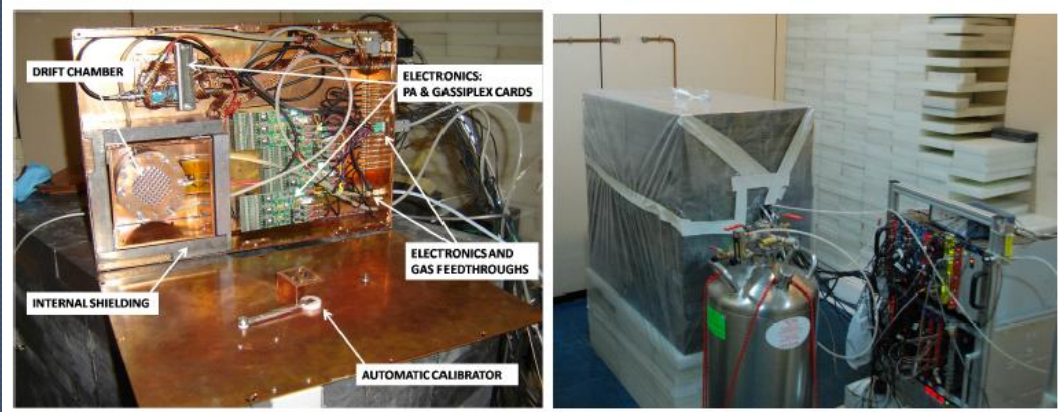
Shielding

archeological lead, inner Cu, N<sub>2</sub> flushing.

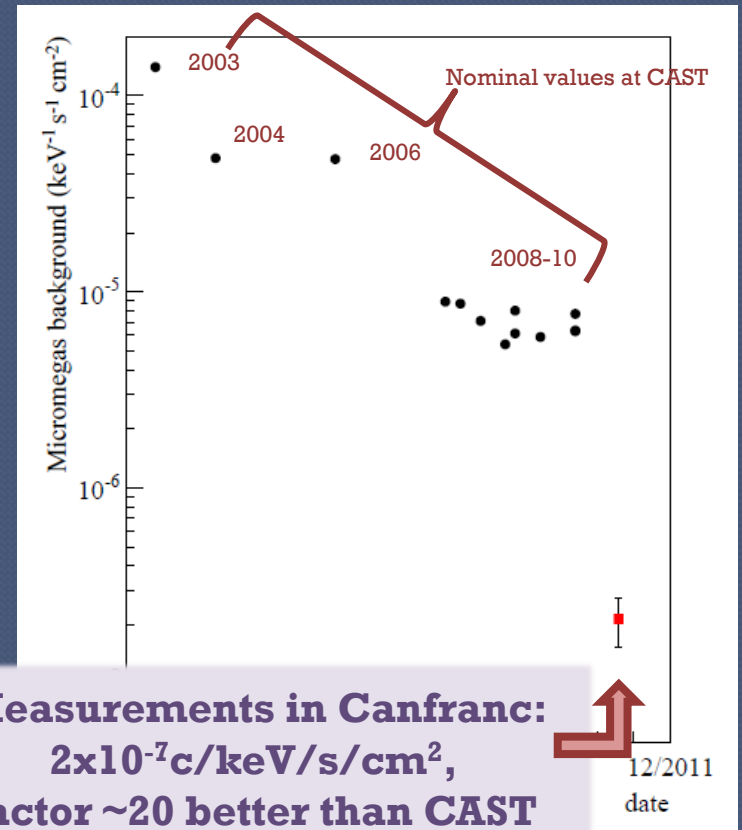
## Background Level history at CAST



# Microbulks in Zgz and Canfranc



## Background Level history at CAST



## Going Underground

(Laboratorio Subterraneo de Canfranc):

<< cosmics

Stable environmental conditions

Better and thicker shielding

Simulation works to build

a background model

**Goal:** to design a new detector  
with improvements implemented

**Measurements in Canfranc:**  
 $2 \times 10^{-7} \text{ c/keV/s/cm}^2$ ,  
factor ~20 better than CAST

# $\mu\text{M}$ for $\beta\beta$ searches: NEXT

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F.J.Iguaz et al. J. Phys.: Conf Series 179 (2009) 012007

T. Dafni et al., NIM A 608 (2009) 259-266

A. Tomas et al., 2009 JINST 4 P11016

S. Cebrian et al., JCAP (2010) 010

Presented in:

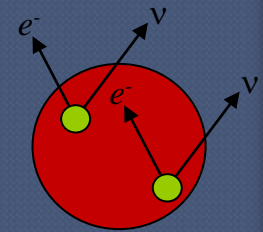
Vienna Instrumentation 2010, IDM2010 (Montpellier), Neutrino2010 (Athens),  
ICHEP2010 (Paris), TPCSymposium 2010(Paris)

# MM readouts for $\beta\beta$

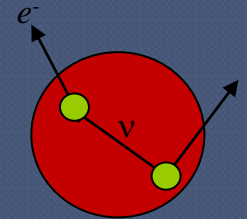
Neutrinoless double beta decay would provide **important information**:

- ✓ on the neutrino nature (Majorana or Dirac)
- ✓ on the neutrino mass scale

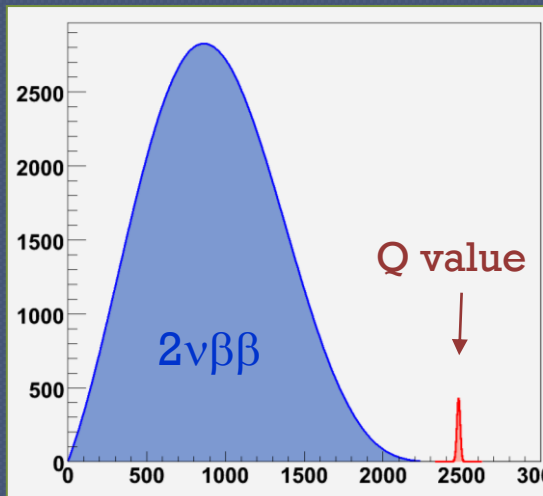
$2\nu\beta\beta$ : Standard process, observed in about 10 isotopes so far



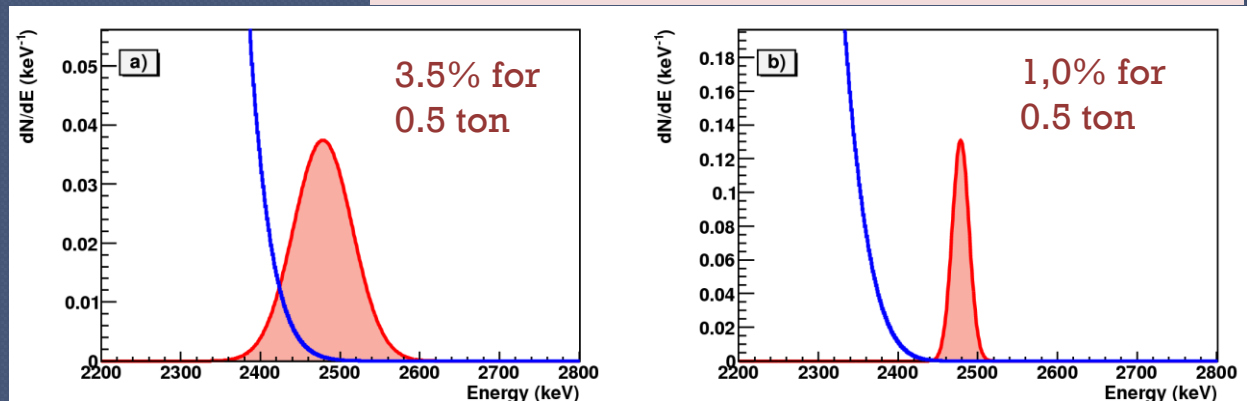
$0\nu\beta\beta$ : Only possible if the neutrino has a mass and is Majorana particle. Yet to be seen



Energy resolution is the only way to distinguish  $0\nu$  from  $2\nu$ .



## The role of Energy Resolution at the ton scale



# NEXT

**A high-pressure, 100kg gaseous Xe TPC to look for the  $0\nu\beta\beta$  decay of  $^{136}\text{Xe}$**

## Project Baseline:

An electroluminescence TPC where the energy is measured with PMTs and the topology is given by SiPM.

## Parallel study:

➤ Equip the detector with Microbulk Micromegas and perform the energy and topology measurement through the charge collection



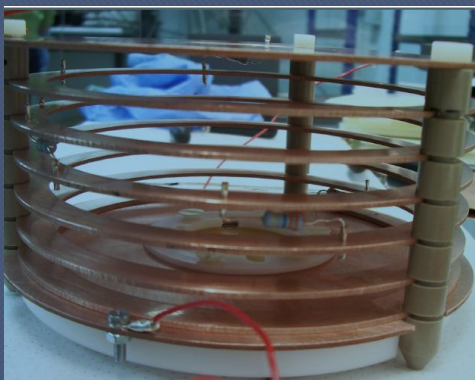
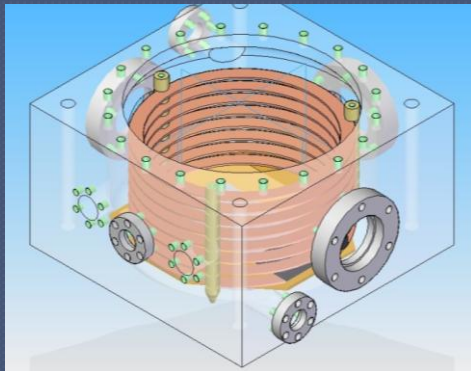
Advances on the R&D related to this option

# NEXT-0-MM Setup



## NEXT-0-MM

Stainless Steel  
UHV specs. bakeable  
Low outgassing materials  
2 liter volume  
Max P 12 bar  
6 cm drift



## Field Cage

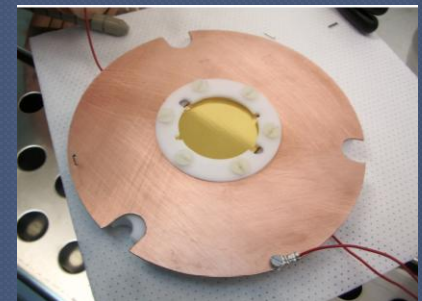
6 copper rings  
3 Peek columns  
6 resistors of 10M $\Omega$

Goal: tests small microbulk readouts in diverse conditions of high pressure Xe

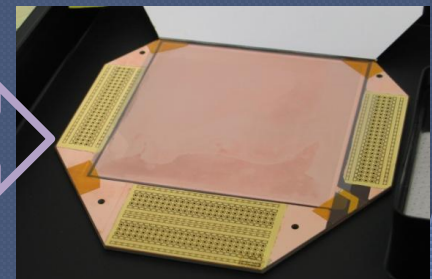
## Microbulk Micromegas

50  $\mu\text{m}$  gap

$\varnothing$  35 mm



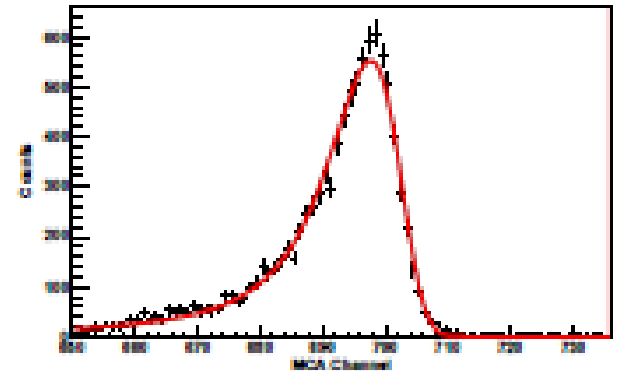
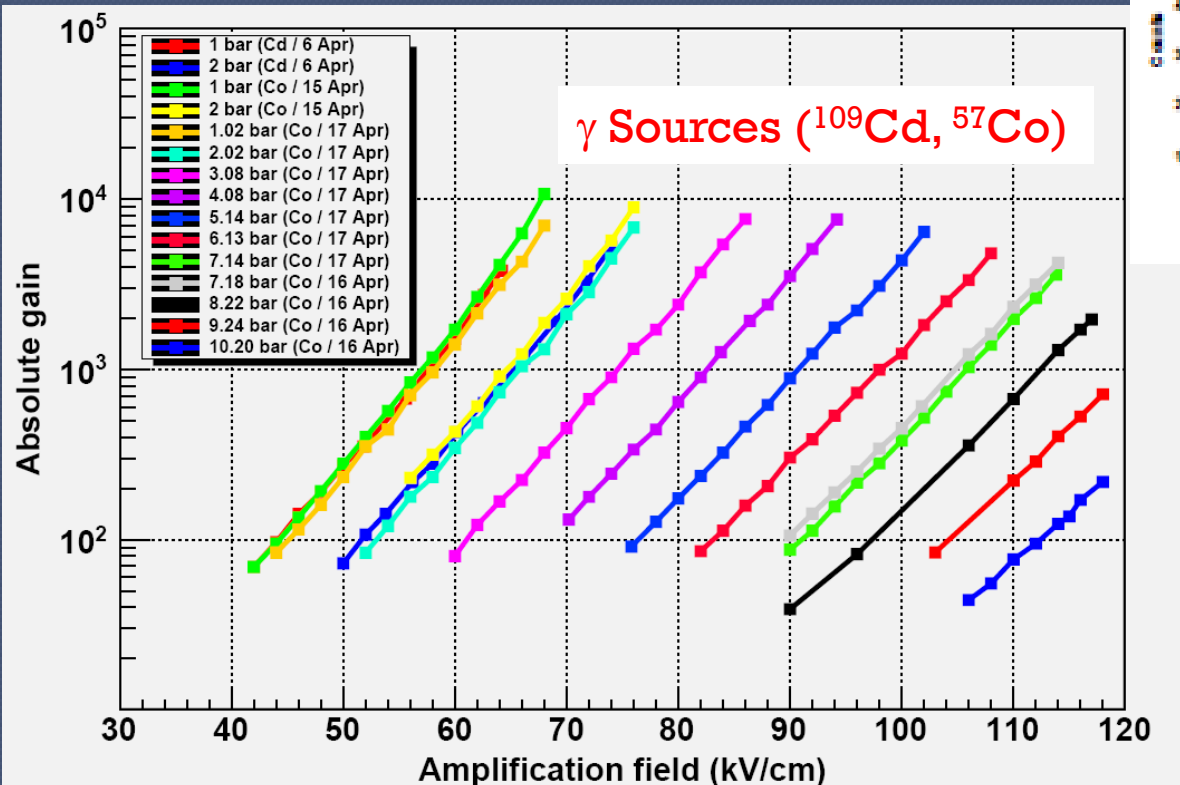
10x10 cm<sup>2</sup>  
Pixelized anode  
12x12 pixels



# NEXT-0-MM

## Tests in Argon-Isobutane

Example of results in Ar-2% Isobutane  
up to 10 bar



5.5 MV  $\alpha$  de  $^{241}\text{Am}$

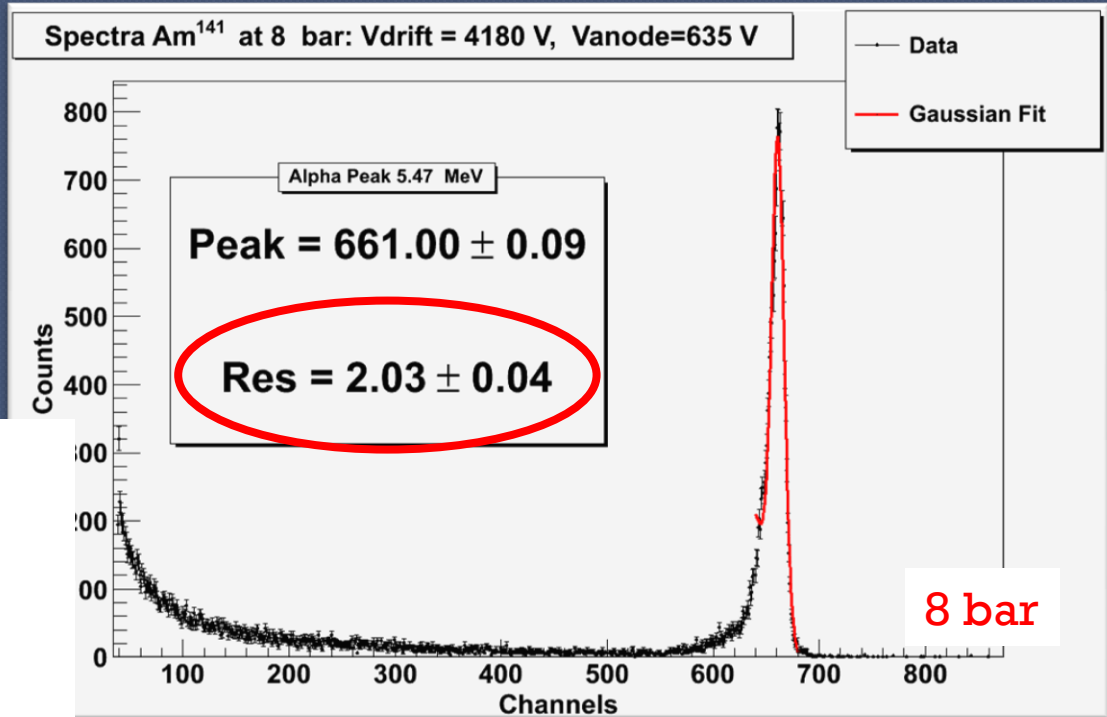
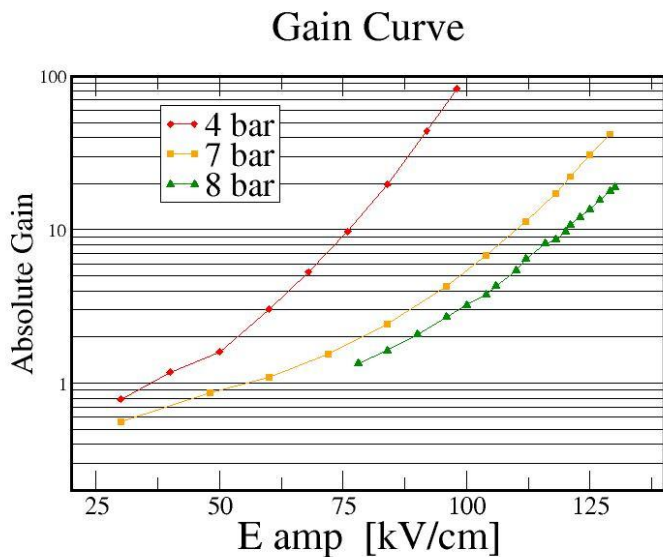
0.7% (FWHM) at 4.75 bar

# NEXT-0-MM

## Tests with $\alpha$ in Pure Argon

Using an  $^{241}\text{Am}$  source

Tests performed in high pressures, reaching 8 bar.

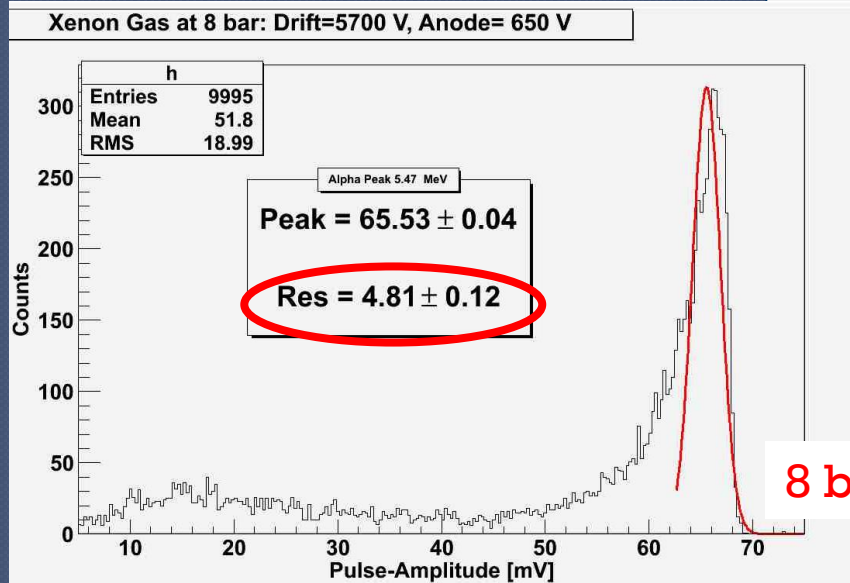
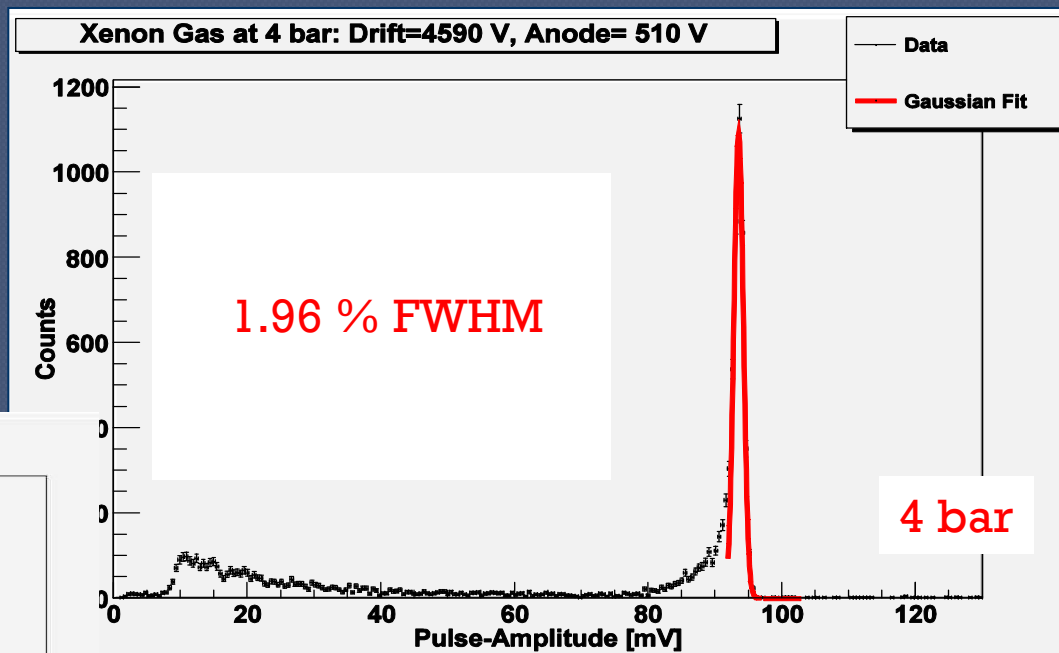




# NEXT-0-MM Tests with $\alpha$ in Pure Xenon

$^{241}\text{Am}$  source

Examples at 4 and 8 bar

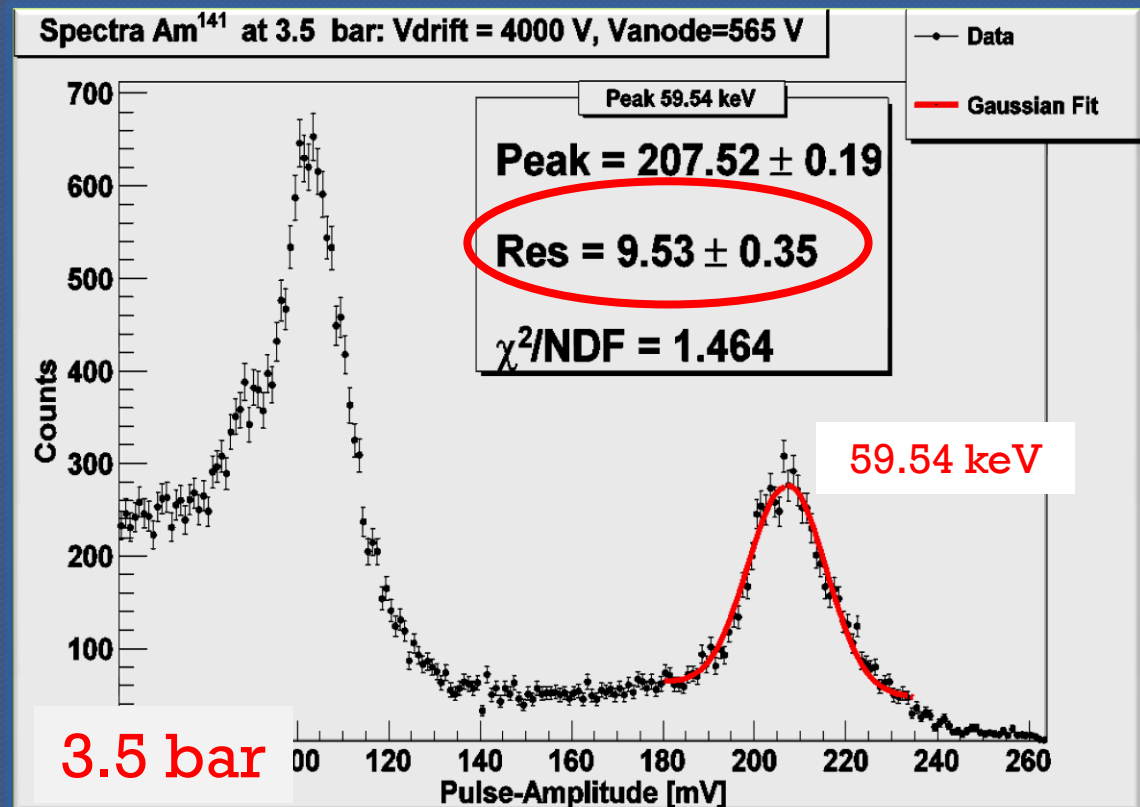


At higher pressures resolution worsens because of Attachment

Measurements with  $\gamma$  in Pure Xenon

$^{241}\text{Am}$  source upside down  
blocks the  $\alpha$  allows the 59.54 keV  $\gamma$

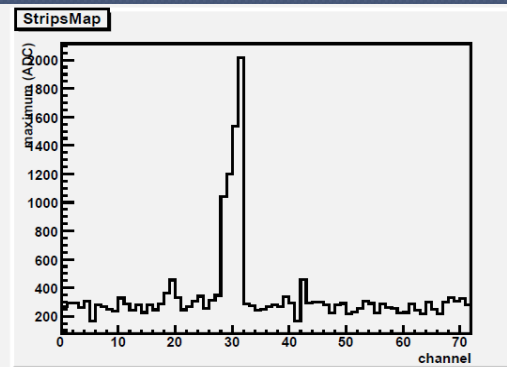
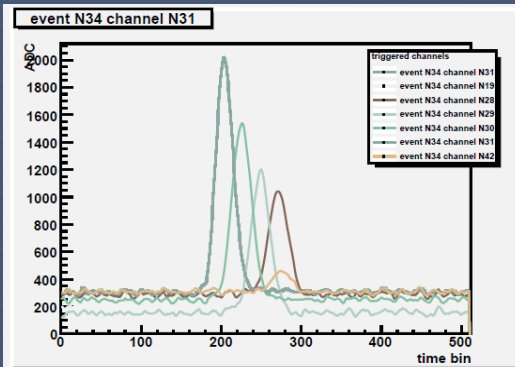
Example taken at 3.5 bar



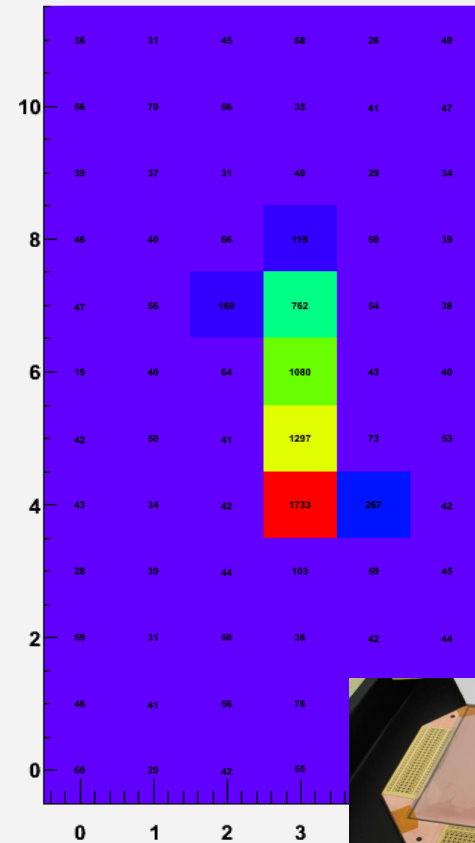
# NEXT-0-MM

## $\alpha$ tracks in Ar-Iso

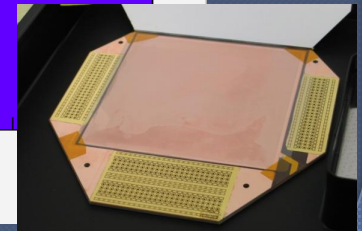
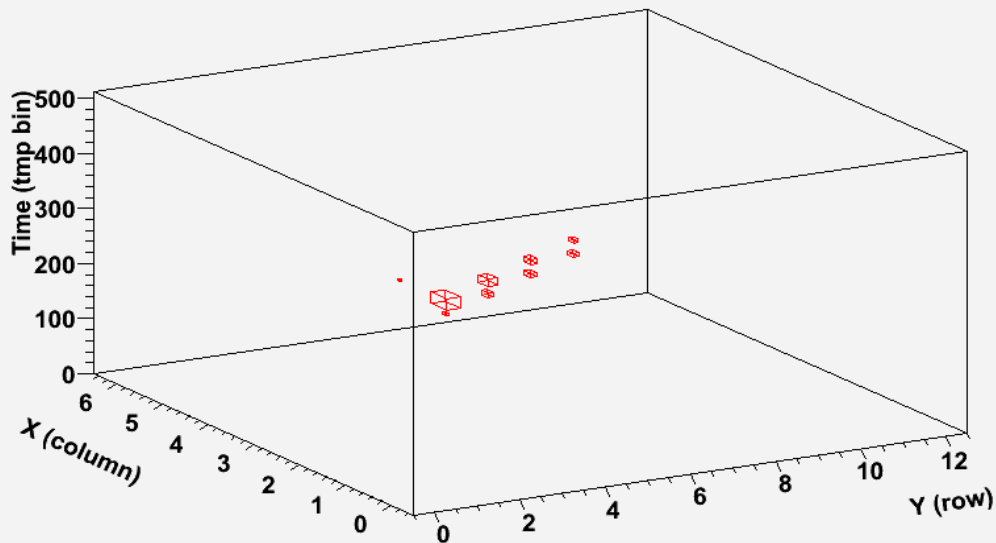
Ar- 2% Isobutane @1bar  
reduced T2K electronics version



Map for event 34



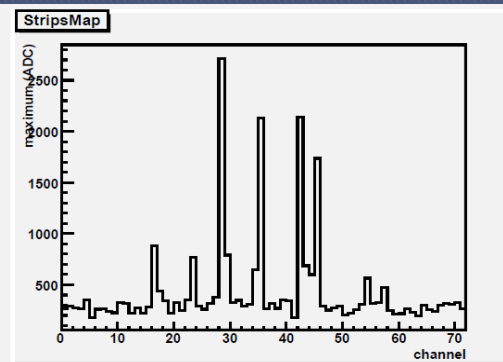
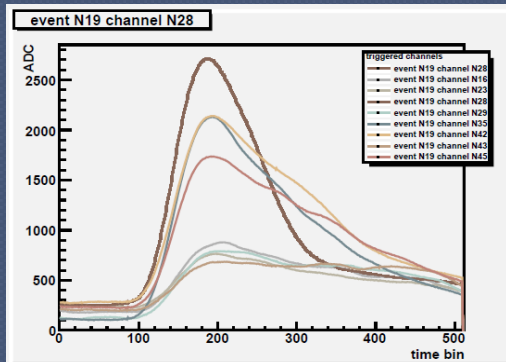
Event 34 reconstructed



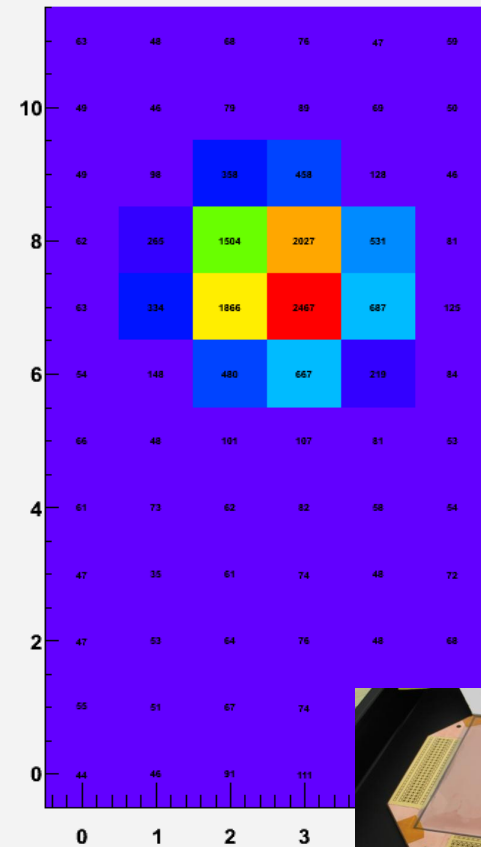
# NEXT-0-MM

## $\alpha$ tracks in pure Ar

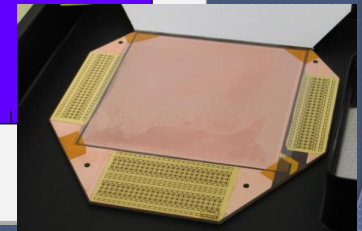
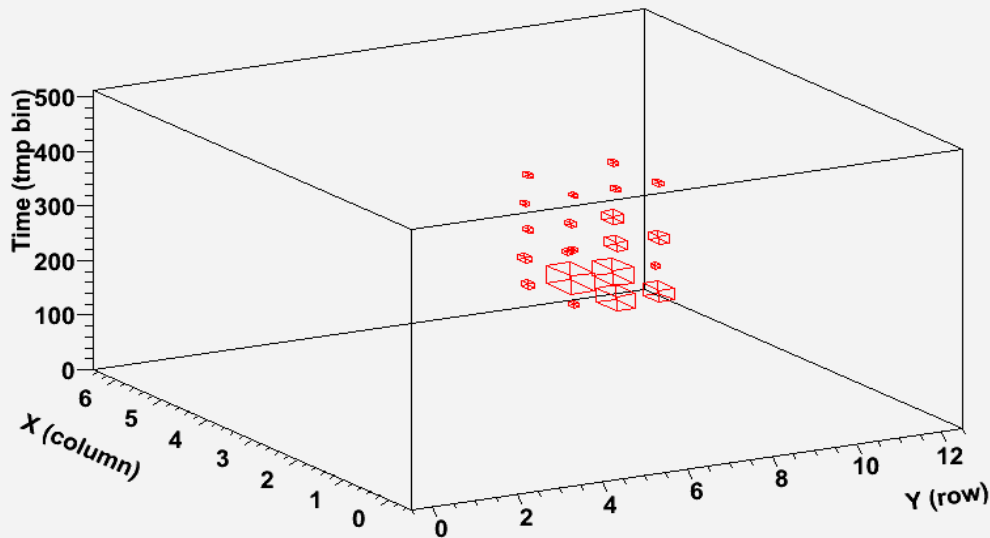
Pure Ar @ 1.23 bar  
reduced T2K electronics version



Map for event 19



Event 19 reconstructed



# NEXT-0-MM Xe-mixtures

Systematic study of **Xe-TMA**  
with Micromegas ongoing.  
Very first results promising:

High gains achievable ( $>10^3$ )

$\gg$  gains than pure Xe at same  
voltage (Penning effect)

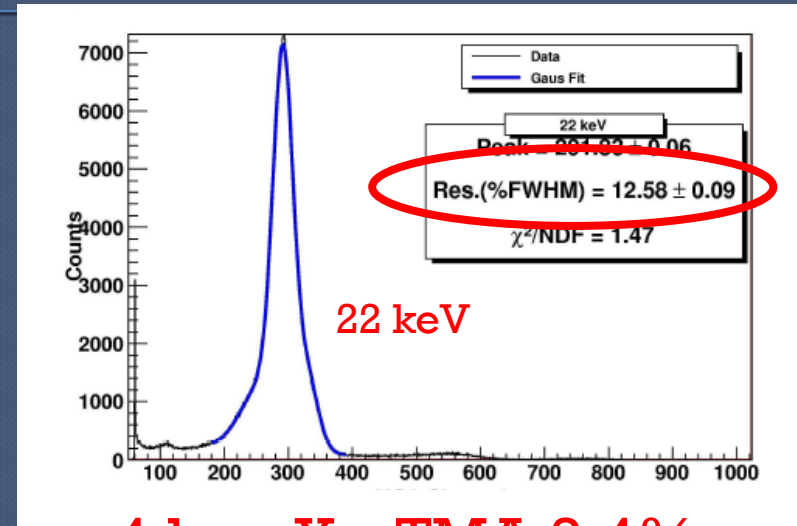
1% resolution seems to be at hand

More work ongoing

Data with **Xe-Ne** also taken:

Higher gains than pure Xe

Energy resolution probably better,  
at least the same (not conclusive)



**4 bar, Xe-TMA 2.4%**

**Extrapolates to**

**1.2 % FWHM at  $Q\beta\beta$**

First data ever taken with MM in this gas, to  
our knowledge

# NEXT-1-MM

Bigger Prototype for  $\sim 1\text{kg}$  of Xe (at 10 bar)

test microbulk readouts in realistic conditions  
 $e^-$  tracks fully contained

- Inner Volume of 74 litres (600mm height,  $\text{\O} 396\text{mm}$ )
- Tested for operation at high pressure (15 bar)
- Steel structure to manipulate the vessel with a crane
- Using the same gas system as for NEXT-0
- Heating and insulation systems



# Equipping NEXT-1-MM

## Field Cage

35 cm drift height  
34 rings

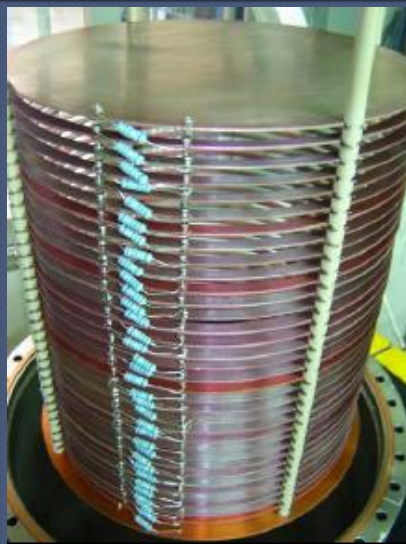
Inner ring  $\varnothing$  28cm

Outer ring  $\varnothing$  30cm

4 PEEK columns

35 resistors

70 PEEK screws



330 M $\Omega$  total resistivity  
For drift fields of  $\sim$  kV/cm  
need to supply 35kV

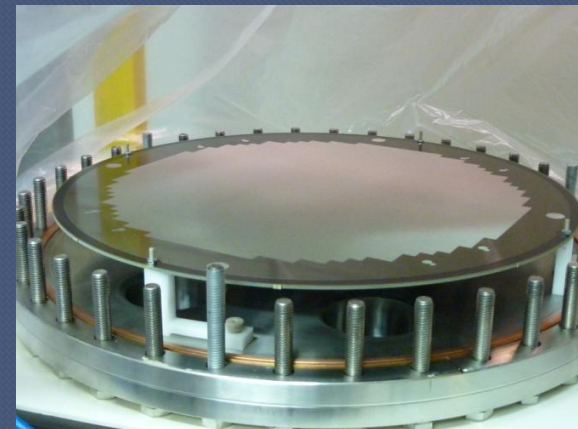
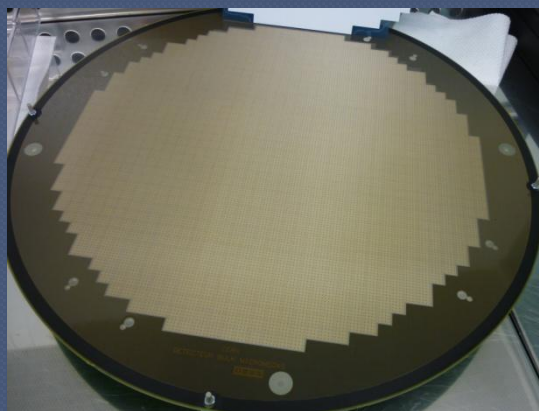
Special HHV FT  
are studied

Cirlex foil between field  
cage and vessel walls



## Bulk Micromegas

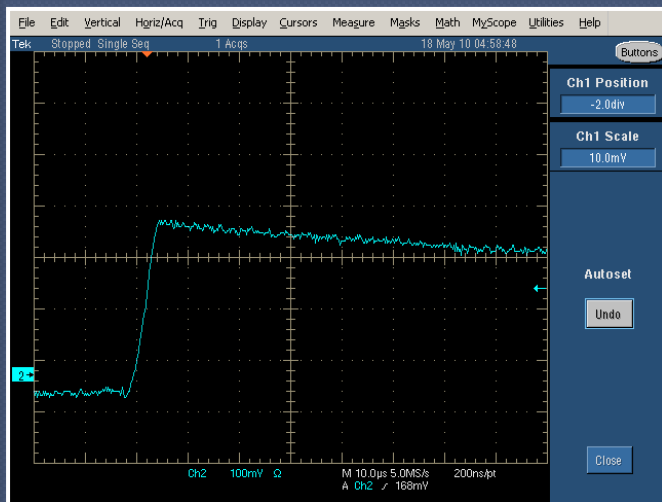
Active region  $\varnothing$   $\sim$ 30cm  
1252 pixels  
independently read



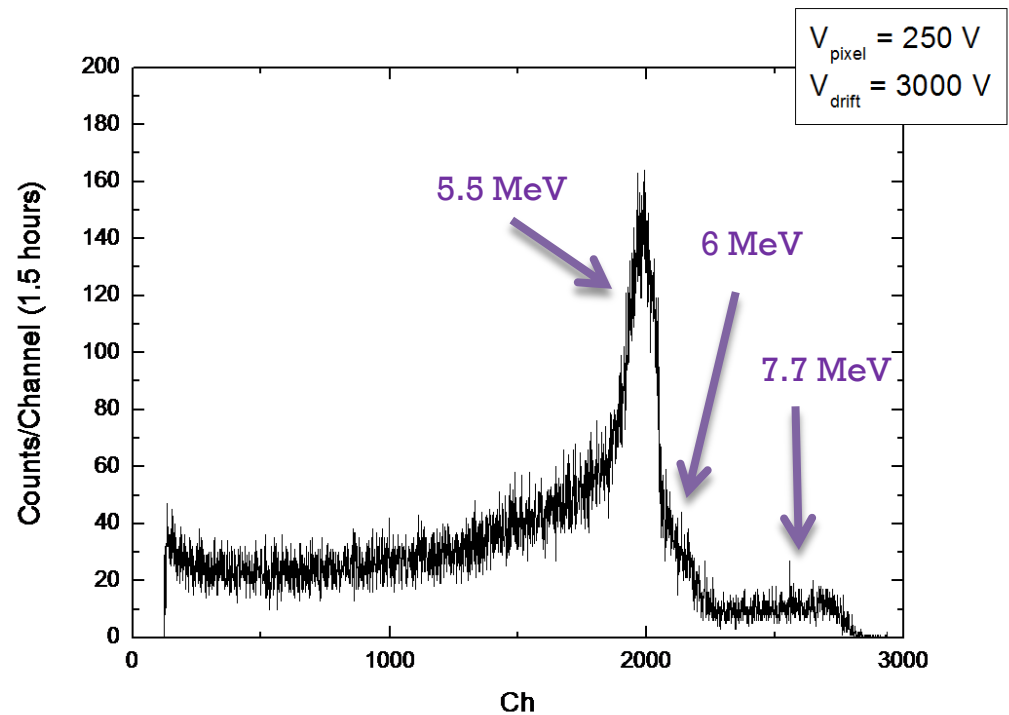
# First Tests

## Reading the mesh

First pulses  
(muons and alpha events)



Rn source diffused  
in Ar - 2%Iso at 1 bar  
All volume active

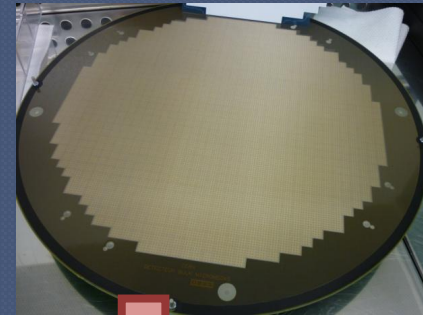




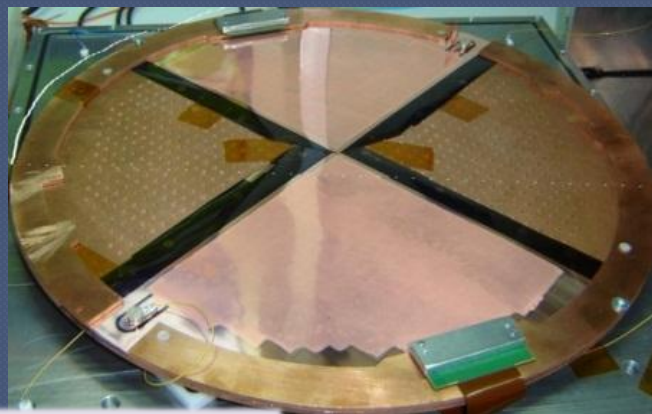
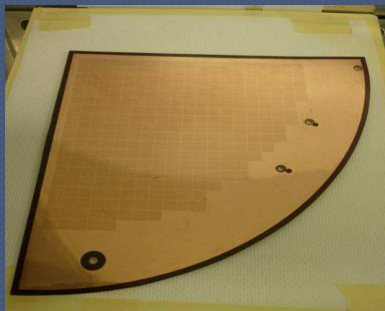
# First Pulses



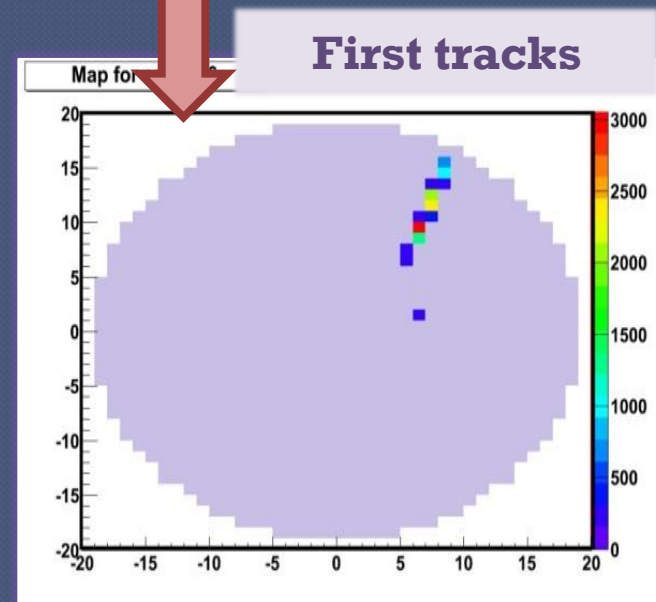
First alpha tracks in Ar  
with AFTER-based DAQ  
with a Bulk readout



Now installing microbulk readout



**Microbulks in 4 sectors**



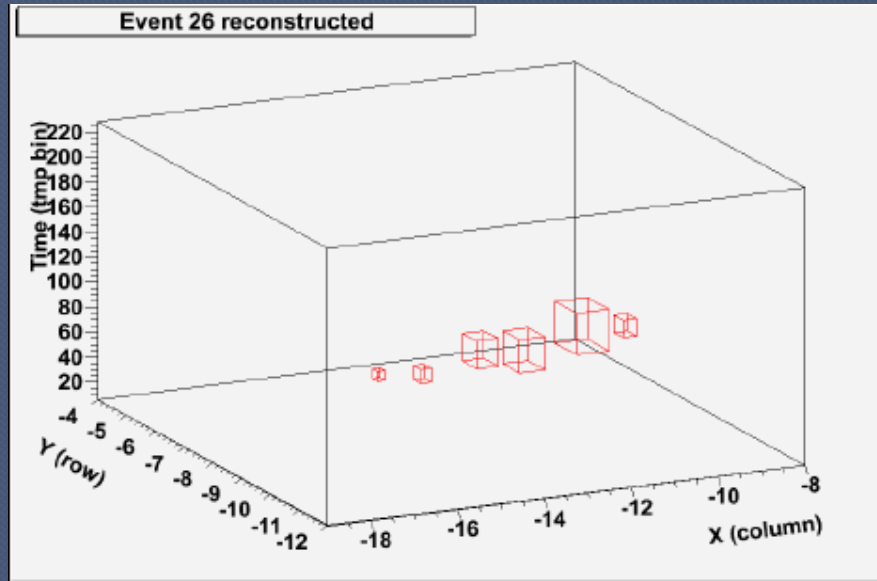
# NEXT-1-MM

## Reading the pixels

Ar-2% Isobutane

Rn source

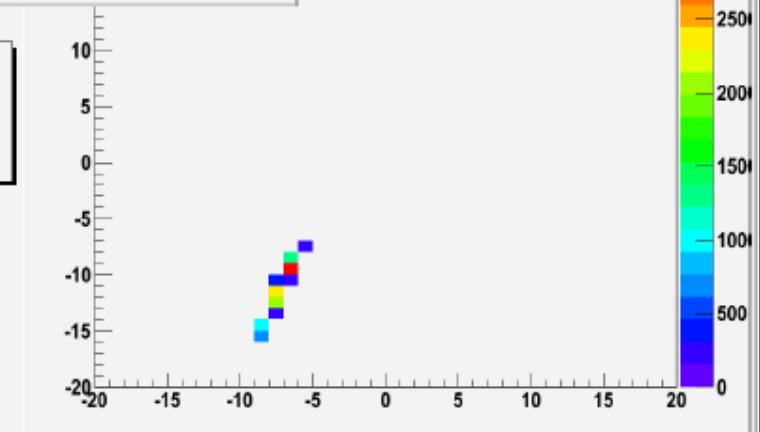
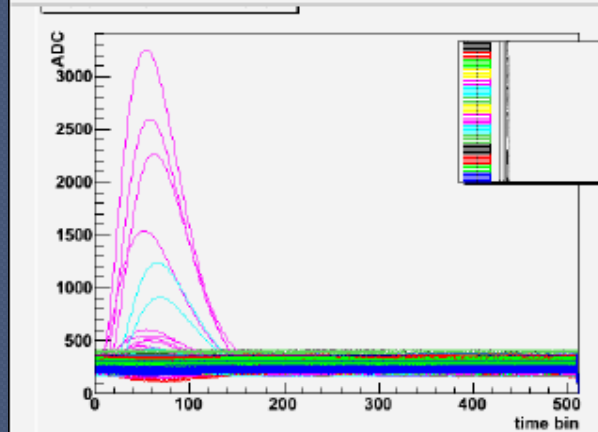
1 bar



NEXT-1-MICROMEGAS

Alpha (Rn) event!

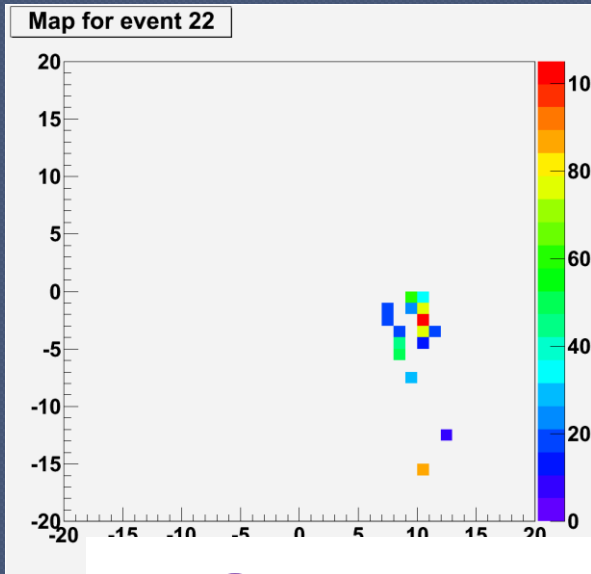
Run68  
Zaragoza 21<sup>th</sup> October 2010



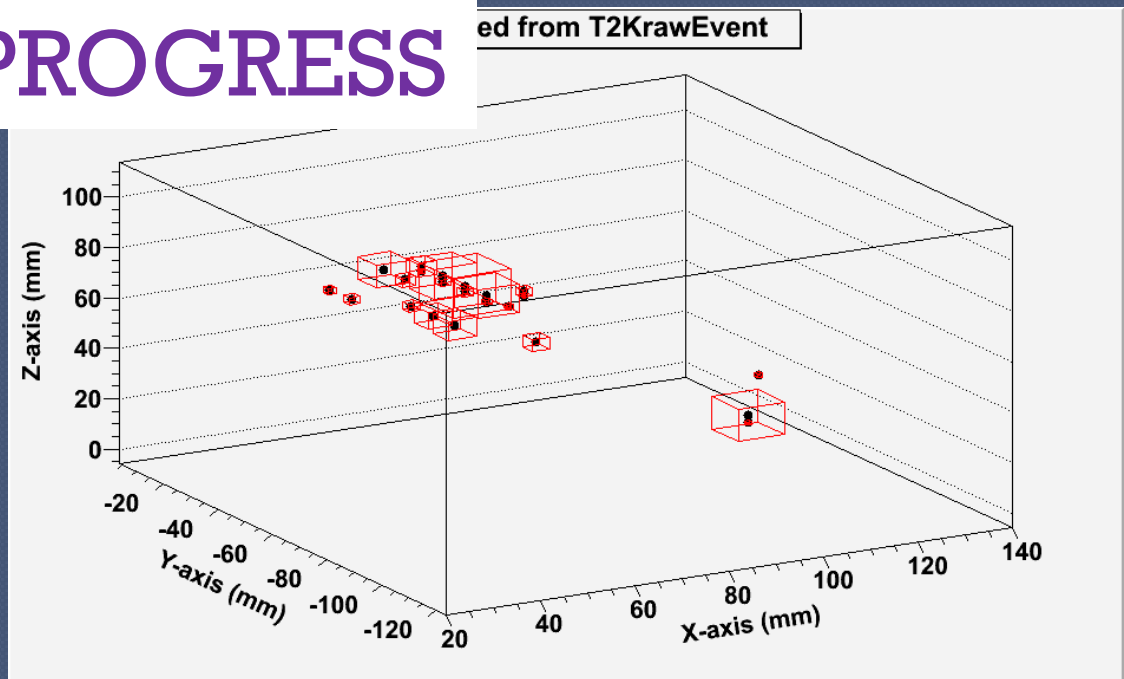
# NEXT-1-MM

## Reading the pixels

First events in Xe !



**WORK IN PROGRESS**



# Radiopurity measurements

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S. Cebrian et al., *Astropart. Phys* (2010) doi:10.1016/j.astropartphys.2010.09.003

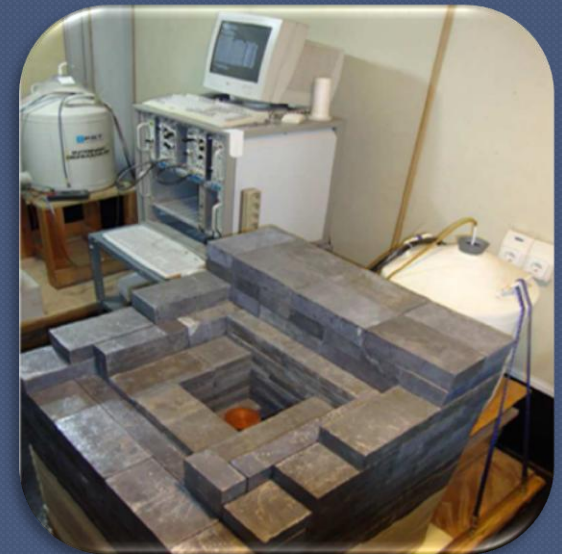
# Radiopurity measurements

Rare event searches dictate the use of radiopure materials  
important parameter for the big-scale experiments  
Necessary to know the quantity, nature and origin of the contamination

Microbulks are mostly Cu & Kapton  
**potentially very radiopure**

Several samples measured  
with HPGe at Canfranc

2 samples of raw material (double clad kapton foil)  
2 samples detached from old CAST detectors



**HPGe detector in Canfranc**

# Radiopurity measurements results

Results (in $\mu\text{Bq}/\text{cm}^2$ )	$^{232}\text{Th}$	$^{235}\text{U}$	$^{238}\text{U}$	$^{40}\text{K}$	$^{60}\text{Co}$
<b>Microbulk mM</b>	<9.3	<13.9	$26.3 \pm 13.9$	$57.3 \pm 24.8$	<3.1*
<b>Kapton-Cu foil</b>	<4.6*	<3.1*	<10.8	<7.7*	<1.6*
<b>Cu-Kapton-Cu foil</b>	<4.6*	<3.1*	<10.8	<7.7*	<1.6*

\*Level obtained from the Minimum Detectable Activity of the detector

- ✓ Very low levels of radioactivity, compatible with the sensitivity of the measurement
- ✓ Contamination probably comes from the treatment of the materials used
- ✓ Next steps: identification of the contaminating steps and find alternatives

S. Cebrian et al., *Astropart. Phys* 34 (2011) 354-359

# Summary

## The T-REX project

Microbulk Micromegas readouts are of **most interest** for such applications.

Examples of development within this scope:

✓ For axions, in CAST:

- Stability of operation

- Very low backgrounds:  $< 1 \times 10^{-5} \text{ c/s/keV/cm}^2$  in CAST, with potentials to go lower , closing in the intrinsic resolution with measurements in Canfranc at  $\sim 2 \times 10^{-7} \text{ c/s/keV/cm}^2$

✓ For double beta decay, in NEXT

- Operation at high pressures

- Energy Resolutions of :

  - 1.96 % (FWHM) at 4 bar of pure Xe for 5.5 MeV  $\alpha$

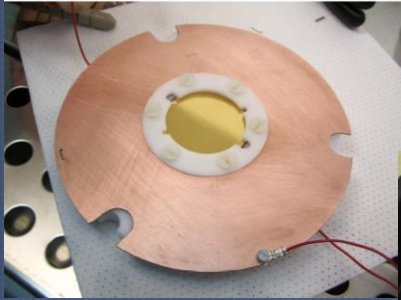
  - 9.5 % (FWHM ) at 3.5 bar of pure Xe for 60keV  $\gamma$

  - 2.4%(FWHM) at 3.4 bar of Xe-2.4%TMA for 22keV  $\gamma$

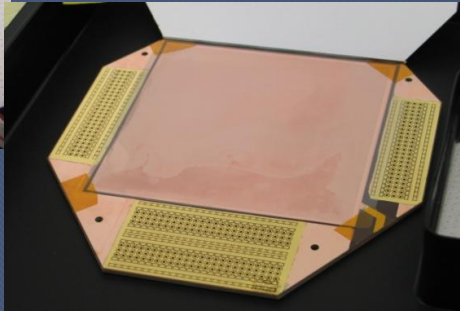
The microbulk planes present **very low** radioactivity levels

Microbulk Micromegas gather all the necessary characteristics

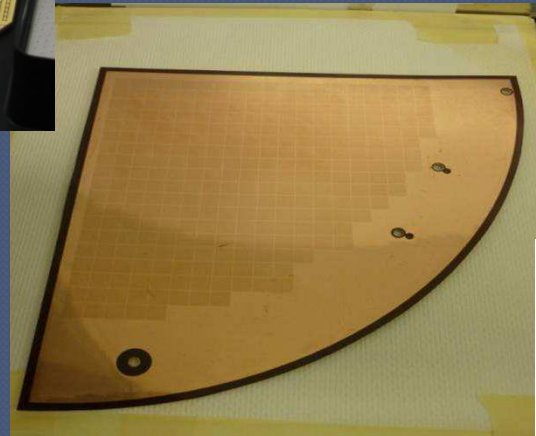
# Microbulk scaling-up



**3 cm Ø**

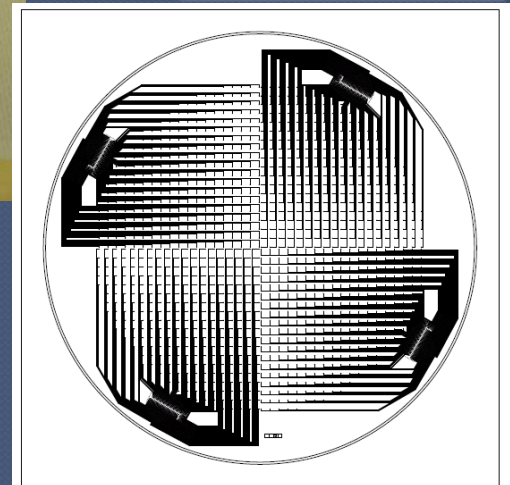


**10 x 10cm<sup>2</sup>  
12 x 12 pixels**



**15 cm radius  
300 pixels  
Largest  
microbulk  
up to now**

**Underway:  
30 cm Ø  
microbulk**





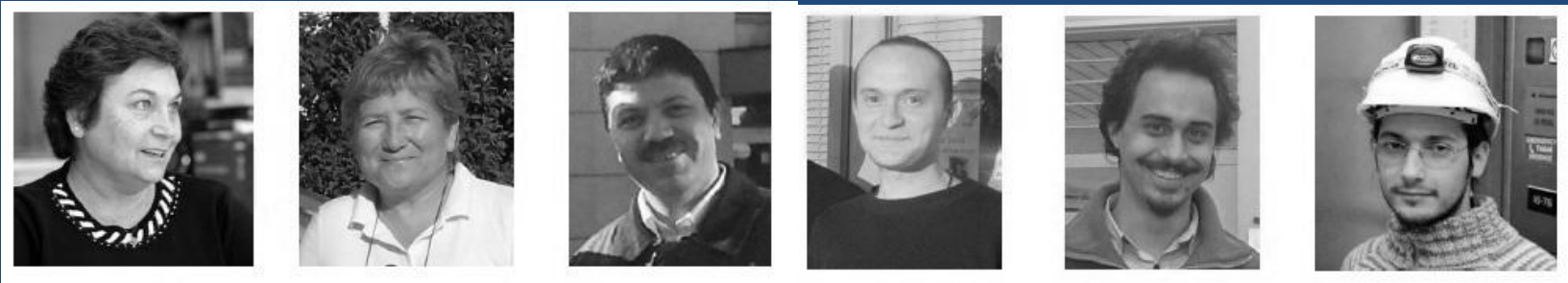
# TREX: infrastructure

~90 m<sup>2</sup> extra space for the project being equipped...



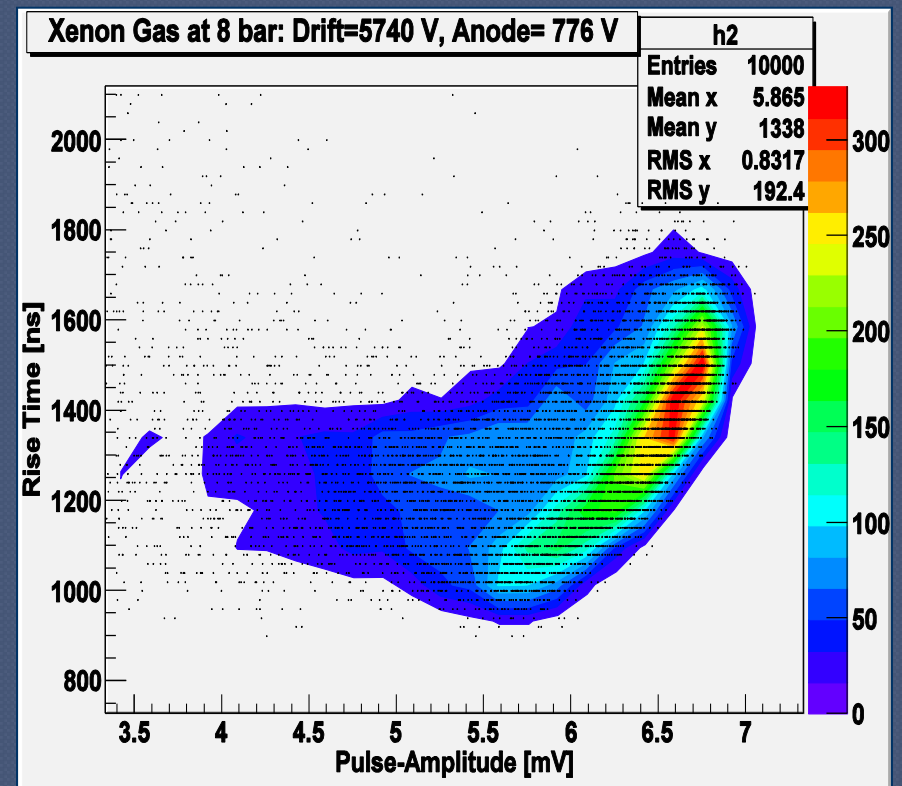
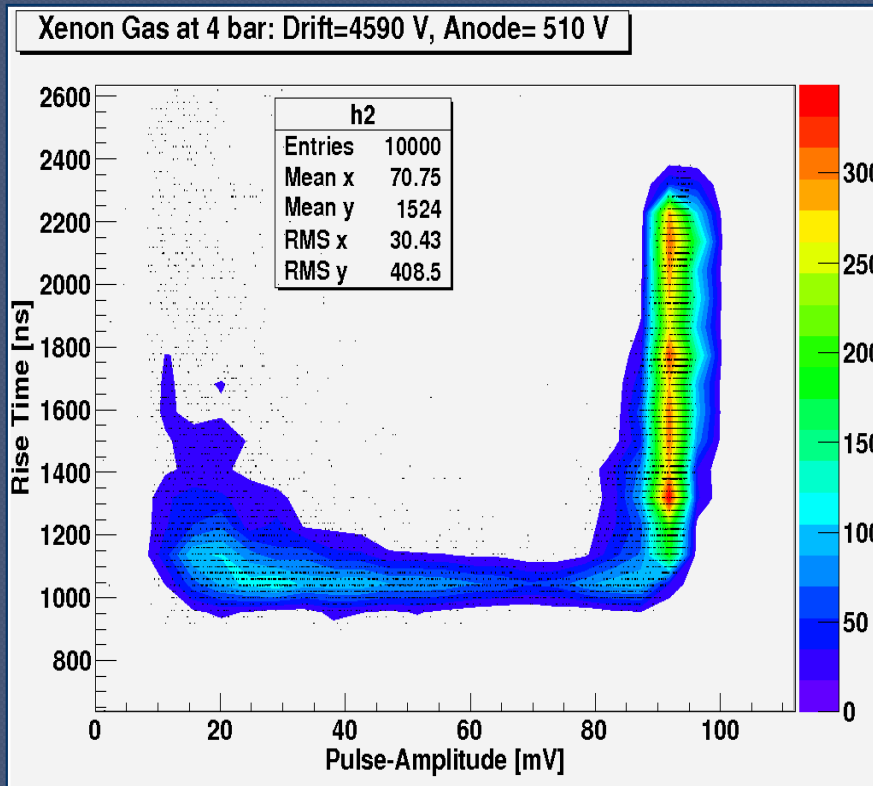
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Thank you for your attention

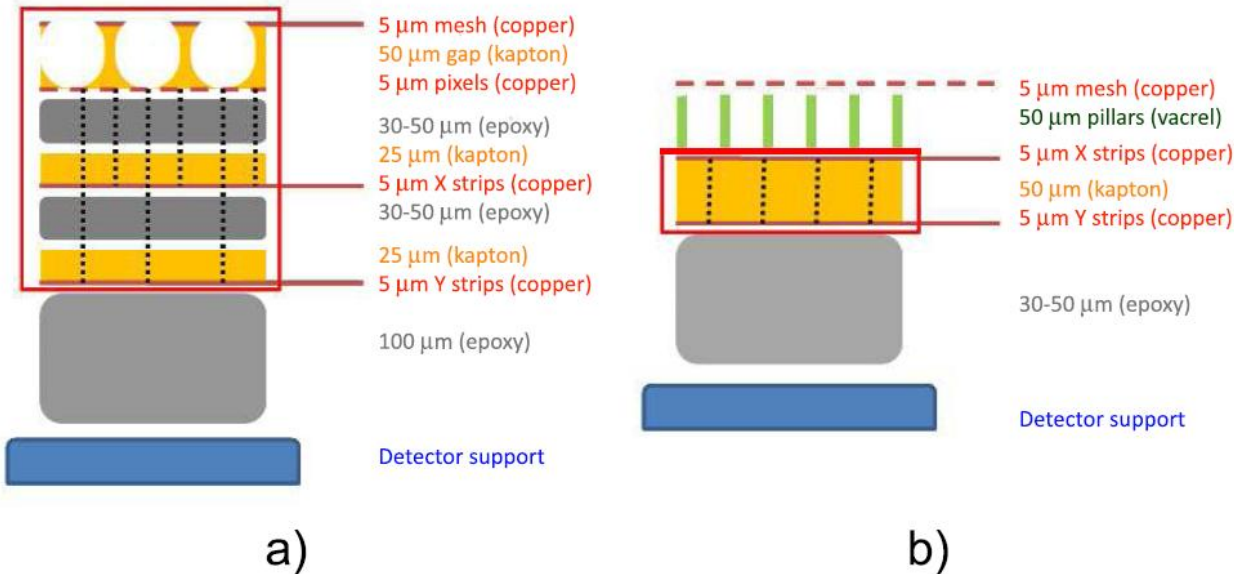


Tests with  $\alpha$  in Pure Xenon

When going to higher pressures the effect of attachment can be clearly seen



# Radiopurity measurements



Results (in $\mu\text{Bq}/\text{cm}^2$ )	$^{232}\text{Th}$	$^{235}\text{U}$	$^{238}\text{U}$	$^{40}\text{K}$	$^{60}\text{Co}$
<b>Microbulk mM</b>	<9.3	<13.9	26.3 $\pm$ 13.9	57.3 $\pm$ 24.8	<3.1*
<b>Kapton-Cu foil</b>	<4.6*	<3.1*	<10.8	<7.7*	<1.6*
<b>Cu-Kapton-Cu foil</b>	<4.6*	<3.1*	<10.8	<7.7*	<1.6*

\*Level obtained from the Minimum Detectable Activity of the detector

S. Cebrian et al., *Astropart. Phys* (2010) doi:10.1016/j.astropartphys.2010.09.003