Background sources and synchrotron radiation issues at JLAB



Compton Polarimeter Backgrounds



- Three primary sources of background
 - Bremsstrahlung
 - Synchrotron radiation
 - Beam halo interaction with beamline elements
- One special source of background (just for some experiments)
 - Neutrons from heavy targets (PREX)



Compton Operation Mode



Photon detector rates

Laser locks and unlocks regularly to allow measurement of backgrounds

- ightarrow Backgrounds highly dependent on beam quality
- → Sometimes extensive tuning is required to achieve good backgrounds *dominant background from beam interaction with apertures in beamline*





Example: stored cavity power droops at high e-beam currents

 \rightarrow Source unknown: synch light or beam scraping heating and distorting mirrors?



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Hall A/C Compton polarimeters use narrow apertures to help protect cavity mirrors from →Large beam related backgrounds →Direct beam strikes

Large beam size, halo will result huge backgrounds from scraping on narrow apertures \rightarrow ion chambers, machine protection system shuts off beam

This system has drawbacks \rightarrow very small halos can still result in significant backgrounds

→ Halo may be small enough to run, but there still may be a lot of junk in your detectors





Beam Halo and Backgrounds



Yves Roblin and Arne Freyberger JLAB-TN-06-048

Sensitivity to beam halo makes us very sensitive to beam tune \rightarrow it is sometimes not possible to achieve ideal tune for Compton and required parameters for experiment

Put model of JLab beam halo (11 GeV) in GEANT model \rightarrow Halo forced to zero at edge of (1 inch) beam pipe



Examples – Hall C Photon Detector



Energy spectrum from Hall A

→Beam energy = 8.8 GeV,
beam current = 20-50 µA
→Laser power ~ 2 kW

Low energy Bremsstrahlung contribution suppressed by discriminator threshold





Synchrotron Radiation at Higher Energies



Photon detector sees synchrotron radiation, primarily from dipoles 2 and 3

Before JLab 12 GeV Upgrade, synchrotron was mitigated with minimal shielding before detector \rightarrow At 6 GeV, this was becoming less effective

At higher energies, shielding alone not sufficient \rightarrow need to mitigate source of synchrotron

J. Benesch et al, Phys. Rev. ST Accel. Beams 18 (2015) 11, 112401





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Dipole Shims for Synchrotron Mitigation

-2

0

Shims added to dipoles in chicane to extend effective length – reduce synchrotron contribution

Basic Dipole

— 3 mm Pb

---- 5 mm Pb

R3

R7 As Built

2

vertical offset (mm)

6 mm aperture

107

10⁶1

10⁵⊦

 10^{3}

-6



10 mm aperture

Initial Higher Energy Running (8.5 GeV)



Initial Higher Energy Running (10.6 GeV)



Acc0/NAcc0, Run=2751, 10mm Aperture



Remotely Adjustable Collimator ("JAWS")

Synchrotron background primarily in vertical direction → Added remotely

adjustable collimator to optimize signal/background



Photon Detector with JAWS (8.8 GeV)



¹⁴ Jefferson Lab

Photon Detector with JAWS (8.8 GeV)



¹⁵ Jefferson Lab

Photon Detector with JAWS (8.8 GeV)



Acc0/NAcc0, Run=2960, 15mm Aperture



Neutron Backgrounds

PREX experiment in Hall A used lead target to measure weak charge radius \rightarrow Many thermal neutrons in hall when beam on target

Used Gd2SiO5:Ce (GSO) for photon detection \rightarrow ¹⁵⁷Gd has large thermal neutron capture cross section



Allison Zec, Ph.D. thesis



Summary

- Dominant background at JLab Compton polarimeters due to beam interaction with material in beamline
 - Most likely small apertures in Fabry-Perot cavity
 - This background can be controlled with beam tuning but sometimes requires much effort and time
- Bremsstrahlung not a significant issue vacuum in Compton area generally at the few 10⁻⁸ level
- Synchrotron was not a major issue before 12 GeV Upgrade
 - Becomes a significantly worse at JLab's highest energies
 - Can be controlled with shielding and appropriate collimation (JAWS)
- In some unique circumstances, neutrons can be a problem
 - PREX experiment (lead target) and GSO detector

