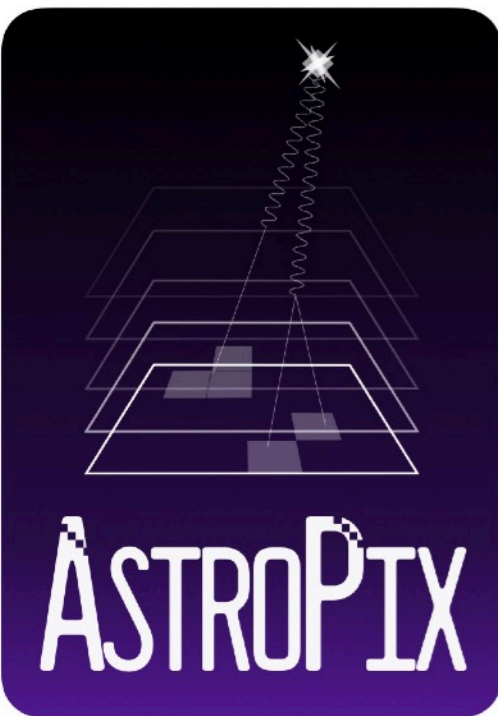


# Performance evaluation of the high-voltage CMOS active pixel sensor AstroPix for gamma-ray space telescopes

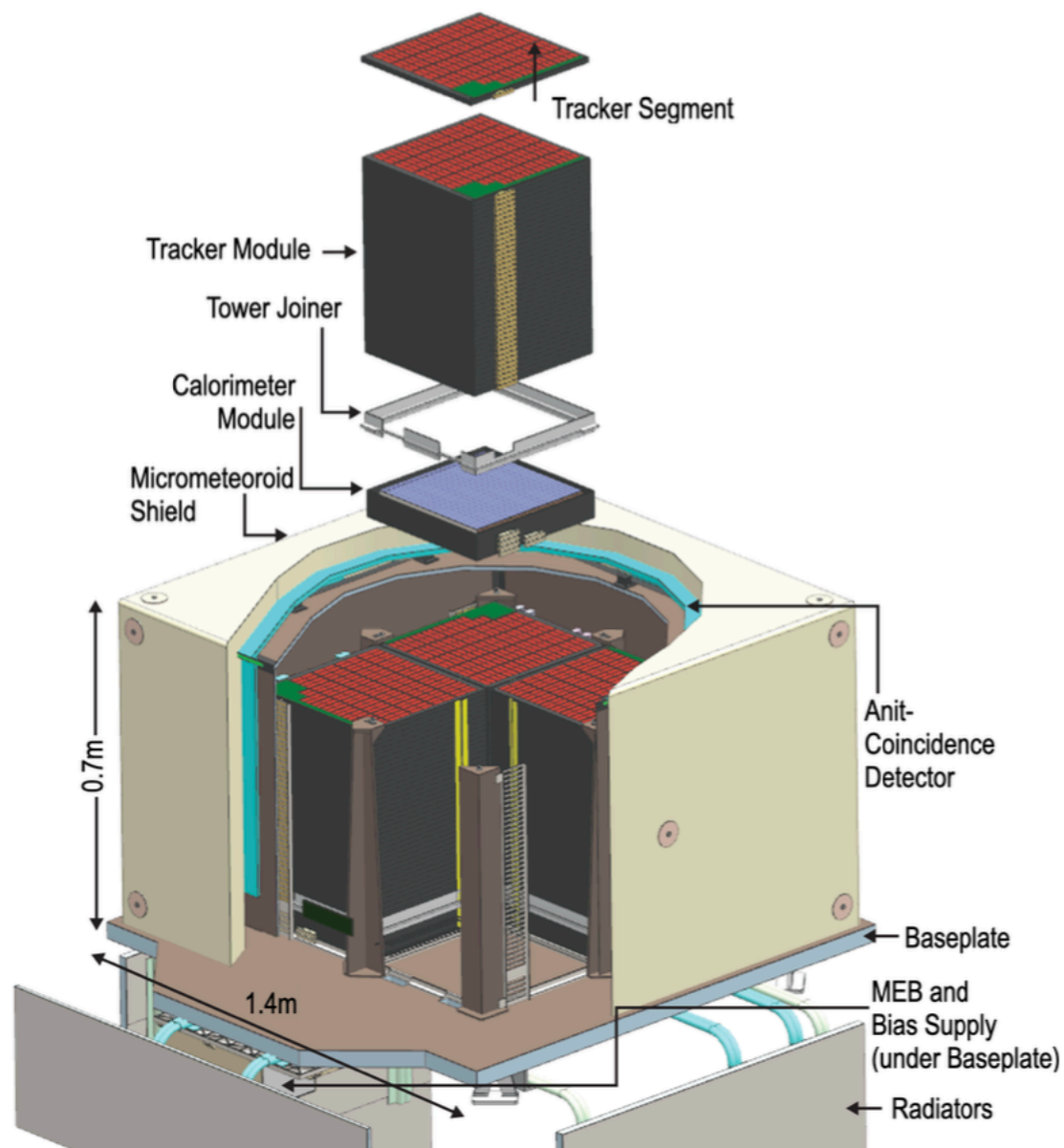
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Regina Caputo, Amanda L. Steinhebel, Daniel Violette,  
Carolyn Kierans, Jeremy S. Perkins (GSFC/NASA),  
Manoj Jadhav, Jessica Metcalfe (ANL), Nicolas  
Striebig, Richard Leys, Ivan Peric (KIT), Michela Negro  
(Louisiana State U.), Taylor Shin (UCSC), Hiroyasu  
Tajima (Nagoya U.), Yasushi Fukazawa, Masaki  
Hashizume, Norito Nakano (Hiroshima U.)**



**13th International “Hiroshima” Symposium on the Development and  
Application of Semiconductor Tracking Detectors (HSTD13)  
Vancouver, Canada  
Dec. 7, 2023**



- AMEGO-X (PI: R.Caputo GSFC/NASA) is a proposed explorer to study extreme astrophysical phenomena in MeV regime

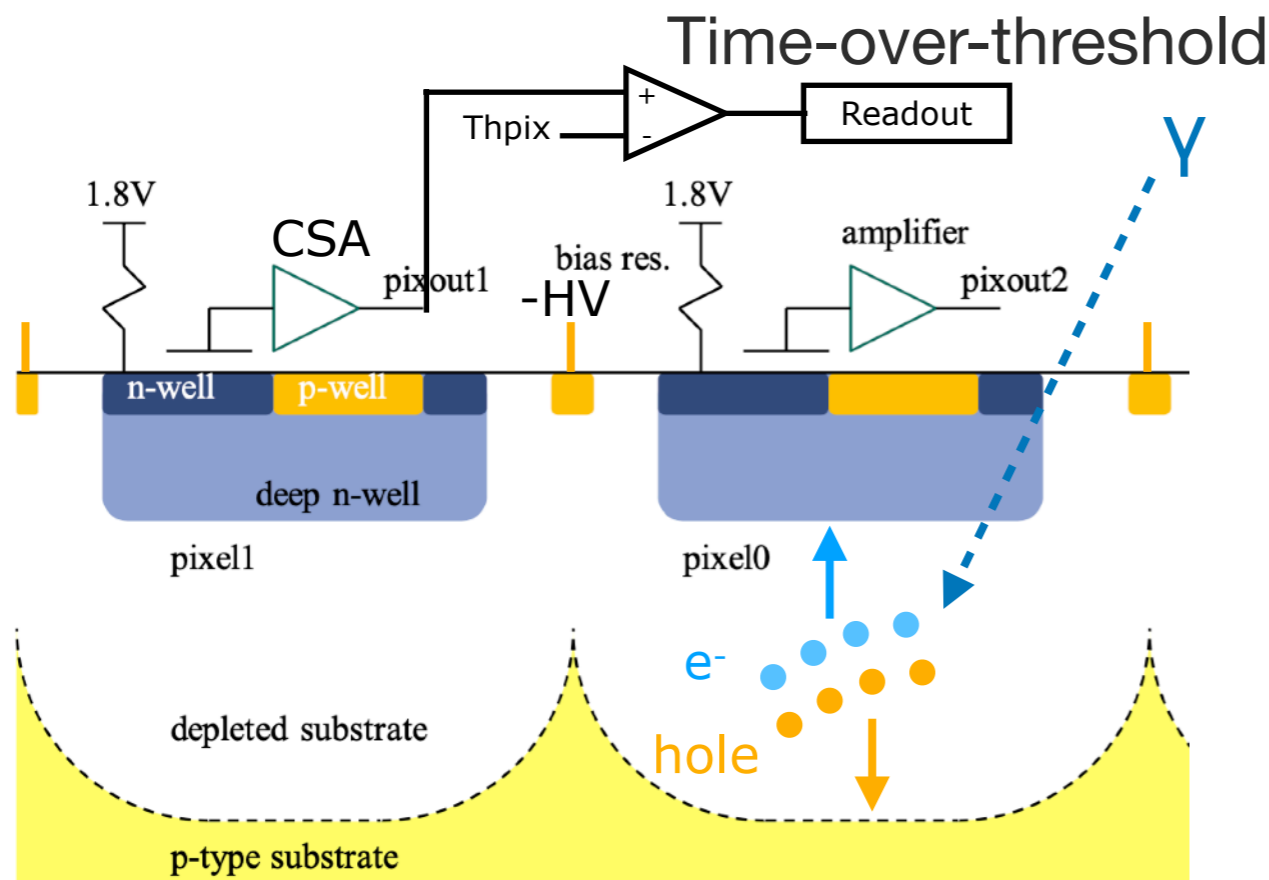


- Shed light on not-well-studied energy regime (25 keV - 1 GeV)
- Huge impact on multi-messenger astronomy thanks to its wide FoV ( $2\pi$  @  $E < 10$  MeV)
- Require a lot of low-noise pixelated silicon sensors ( $\sim 2.4 \times 10^5$  cm<sup>2</sup>) with a wide depletion layer to efficiently perform Compton reconstruction
- Power consumption of silicon sensors must be  $< 1$  mW/cm<sup>2</sup>

→ AstroPix

R. Caputo+ 2022

- AstroPix is a new monolithic HV-CMOS active pixel sensor
  - Full depletion achieved by applying HV
  - Signal processing (CSA → Comparator for Time-over-threshold) is performed on pixel and digitization is done on chip
- Development based on experience from ATLASPix and MuPix  
I.Peric&N.Berger 2018



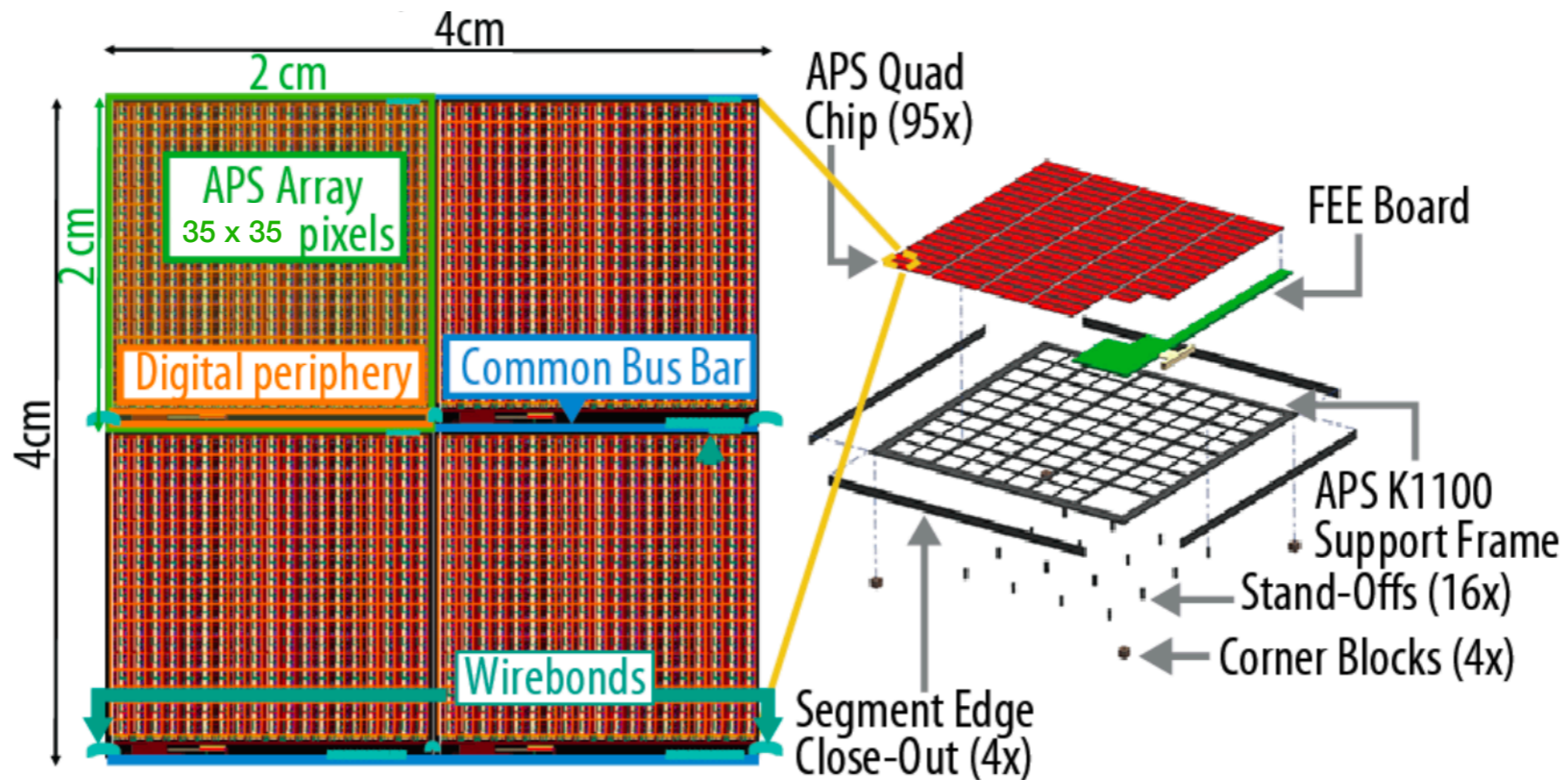
I.Peric with modifications



- AstroPix Team
  - PI: R. Caputo (GSFC/NASA)
  - GSFC, ANL, KIT, UCSC, Hiroshima U, Nagoya U



**Quad Chip** = 4 identical AstroPix array



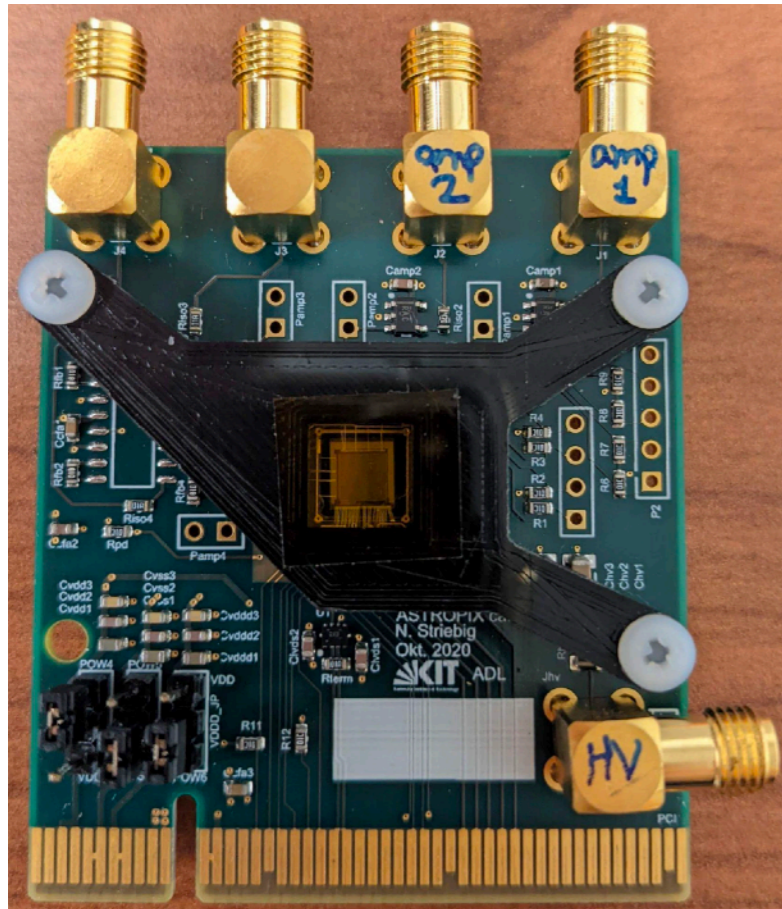
## Requirements

Pixel pitch	$500 \times 500 \mu\text{m}^2$	Dynamic range	25 keV - 700 keV
Thickness	500 $\mu\text{m}$	Energy resolution	< 10% (FWHM) at 60 keV
Power consumption	< 1 mW/cm <sup>2</sup>		



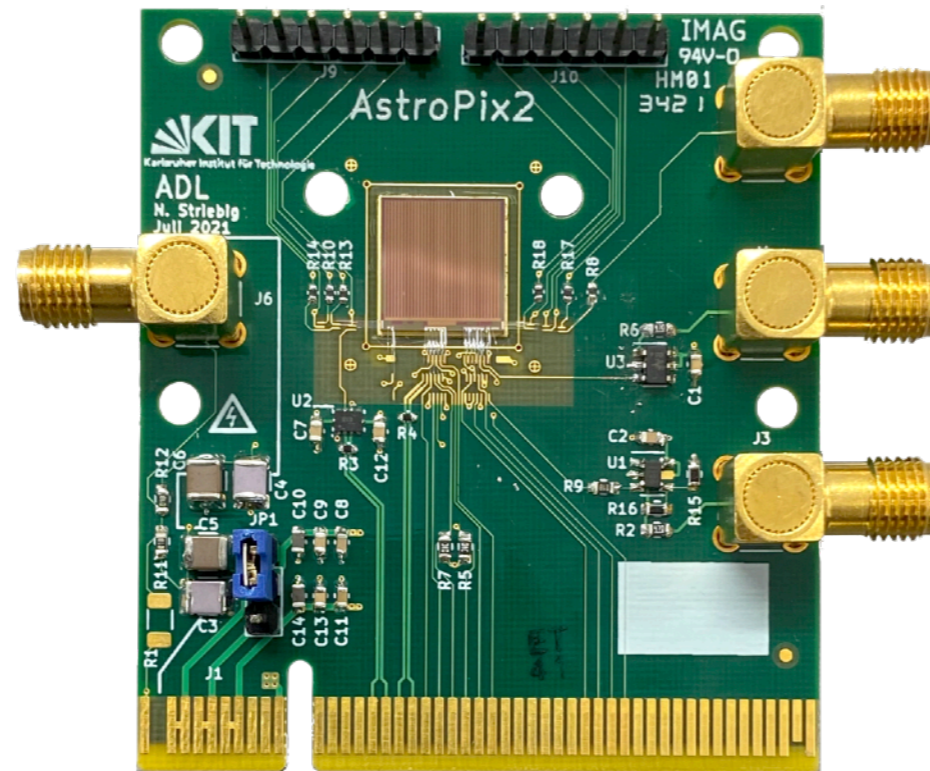
# AstroPix Series

## Version 1



0.5 x 0.5 cm<sup>2</sup> chip  
175 x 175 μm<sup>2</sup> pitch  
18 x 18 pixels  
\*14.7 mW/cm<sup>2</sup>  
\*CSA+comparator only.

## Version 2



1 x 1 cm<sup>2</sup> chip  
250 x 250 μm<sup>2</sup> pitch  
35 x 35 pixels  
\*3.4 mW/cm<sup>2</sup>

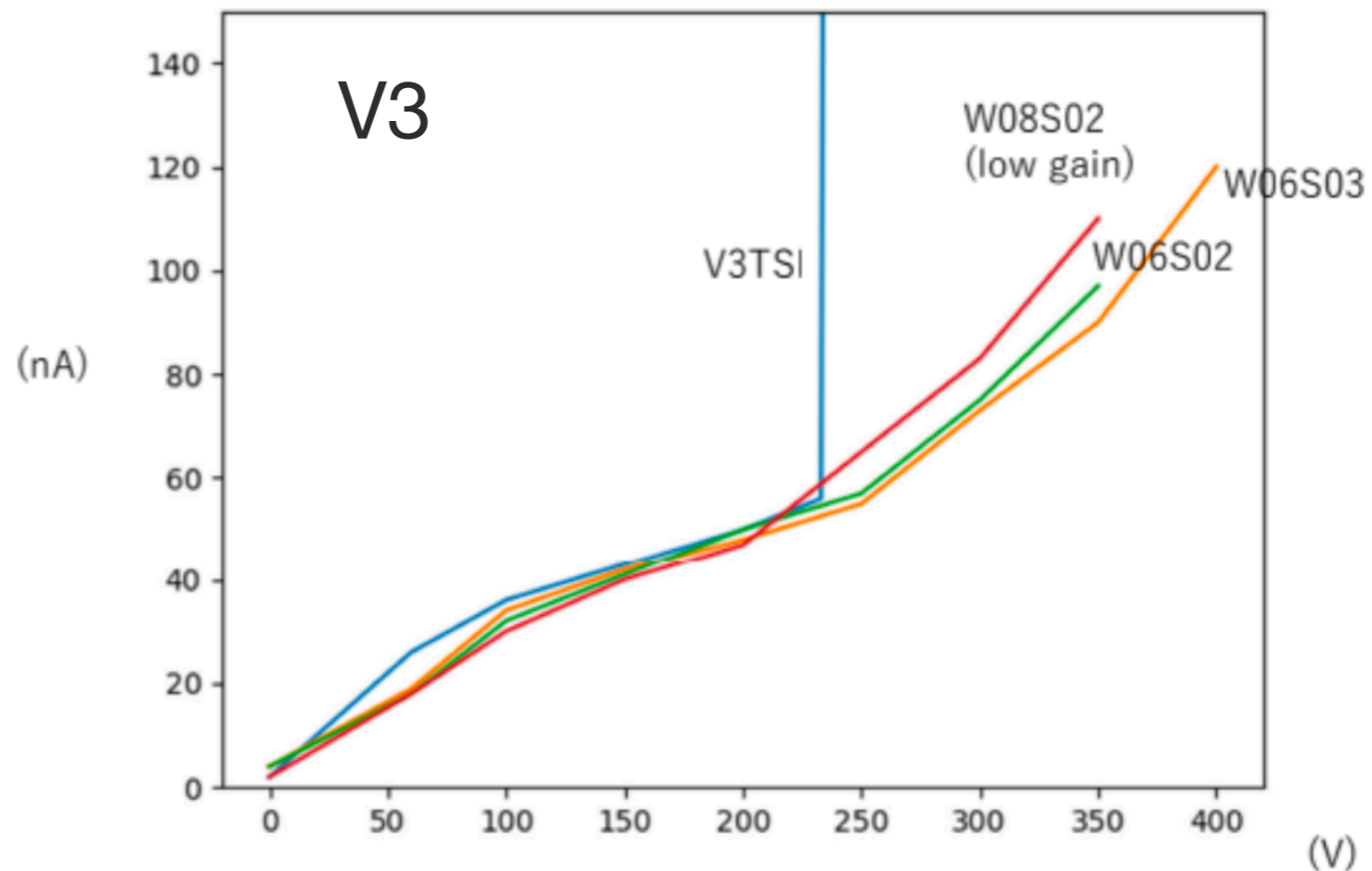
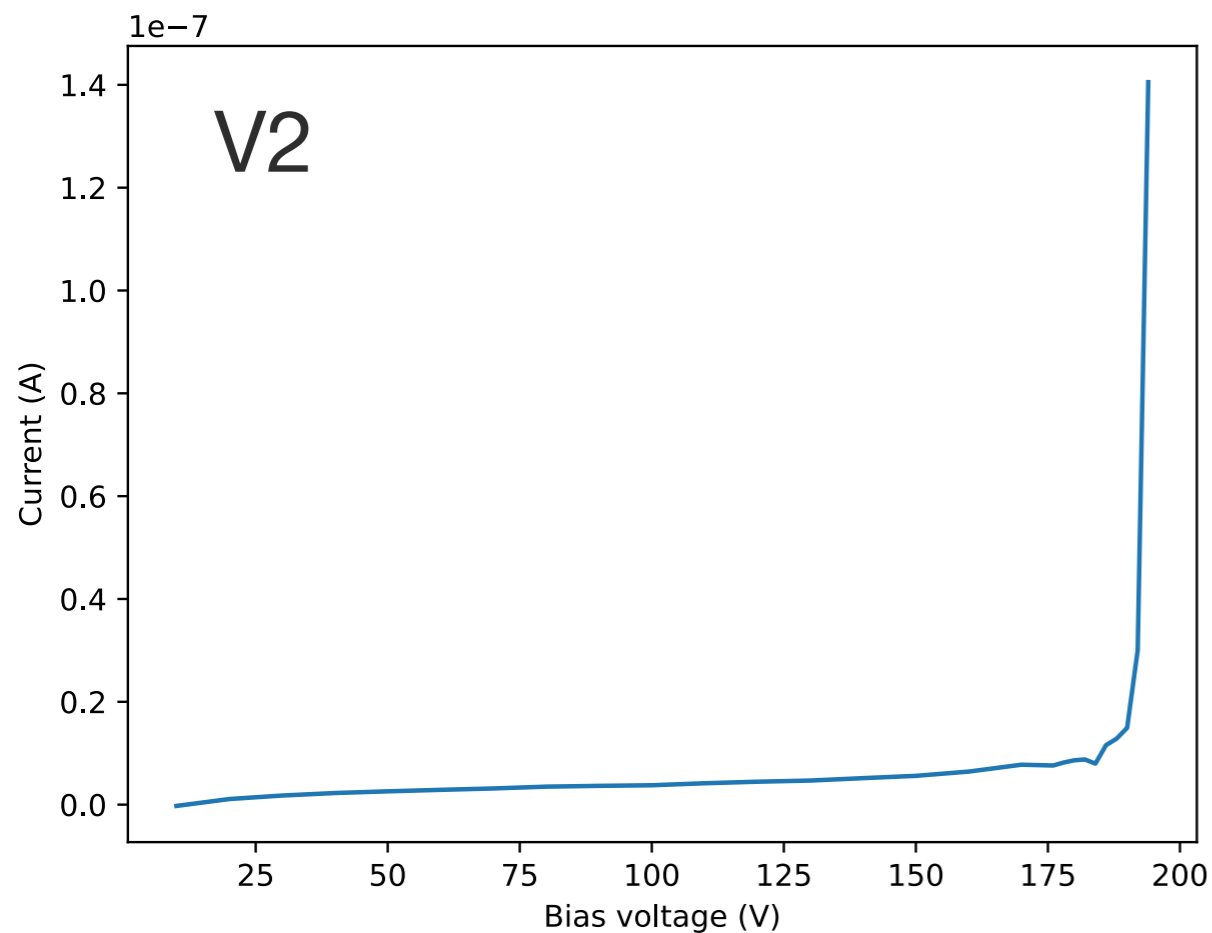
## Version 3



2 x 2 cm<sup>2</sup> chip  
500 x 500 μm<sup>2</sup> pitch  
35 x 35 pixels  
\*1.04 mW/cm<sup>2</sup>  
4.12 mW/cm<sup>2</sup> (incl. digital)

# I-V Curves

- V2: 300  $\Omega\text{cm}$  wafer chip shows breakdown at  $\sim 190$  V
- V3: 300  $\Omega\text{cm}$  wafer chip can go up to  $\sim 400$  V thanks to better clearance to chip-edge  $\leftarrow$  Results in this talk from this wafer
  - 20  $\Omega\text{cm}$  wafer chip breakdown at  $\sim 230$  V
  - 18  $\text{k}\Omega\text{cm}$  wafer chip shows high leakage current in the mA range

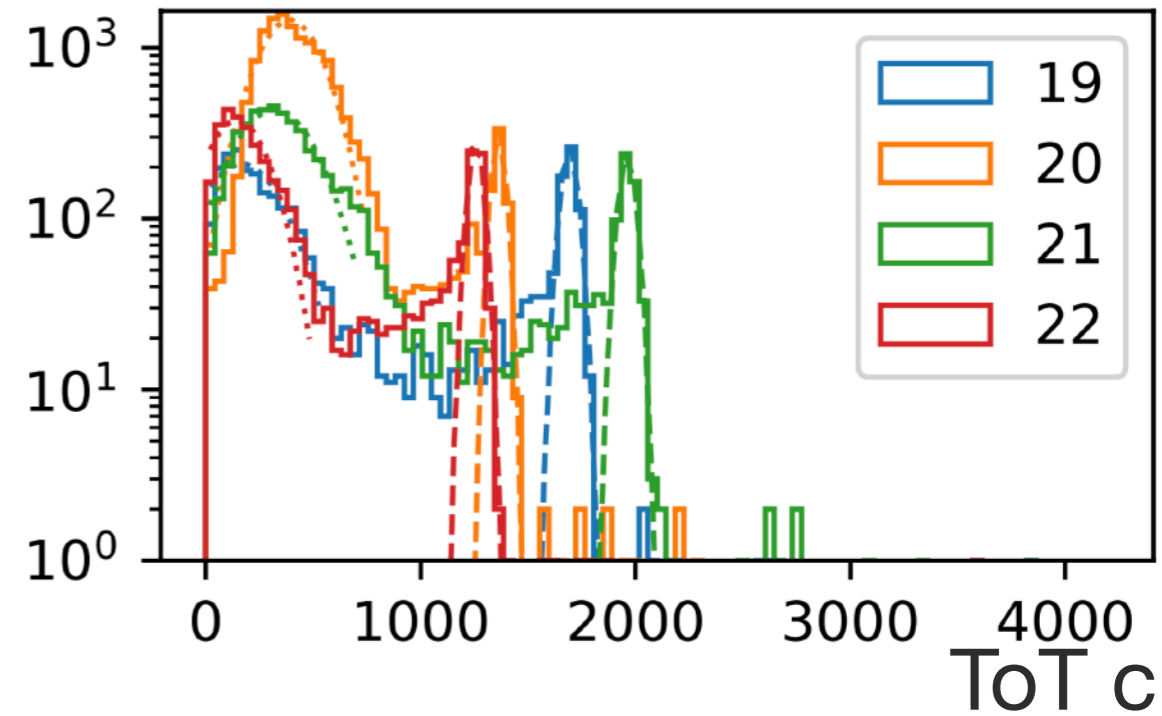
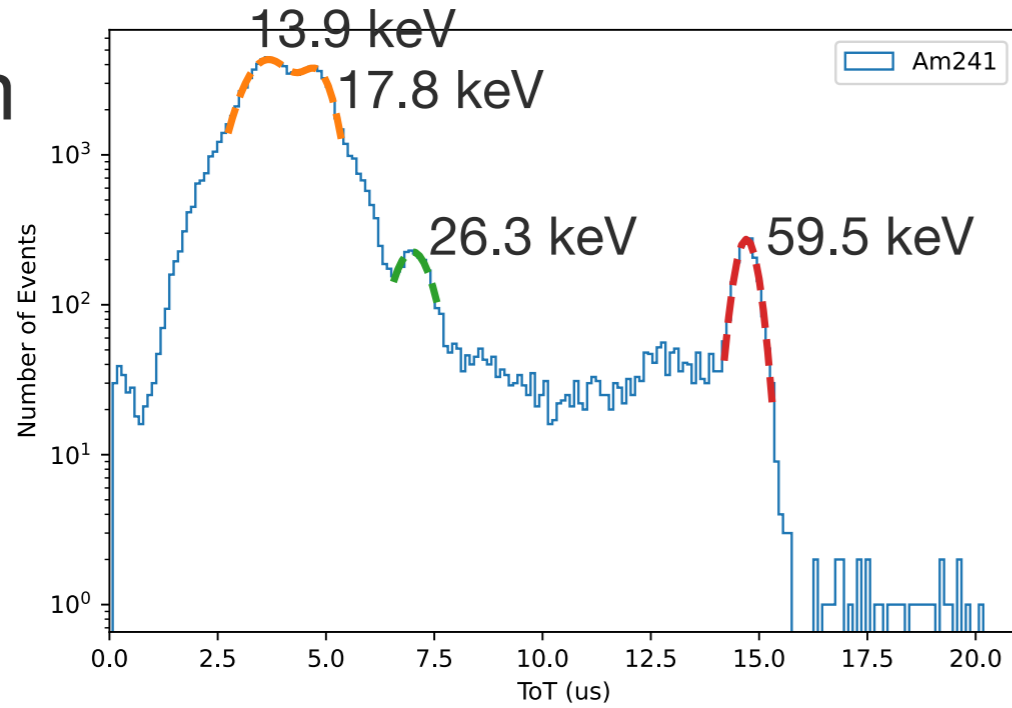


# Energy Spectra

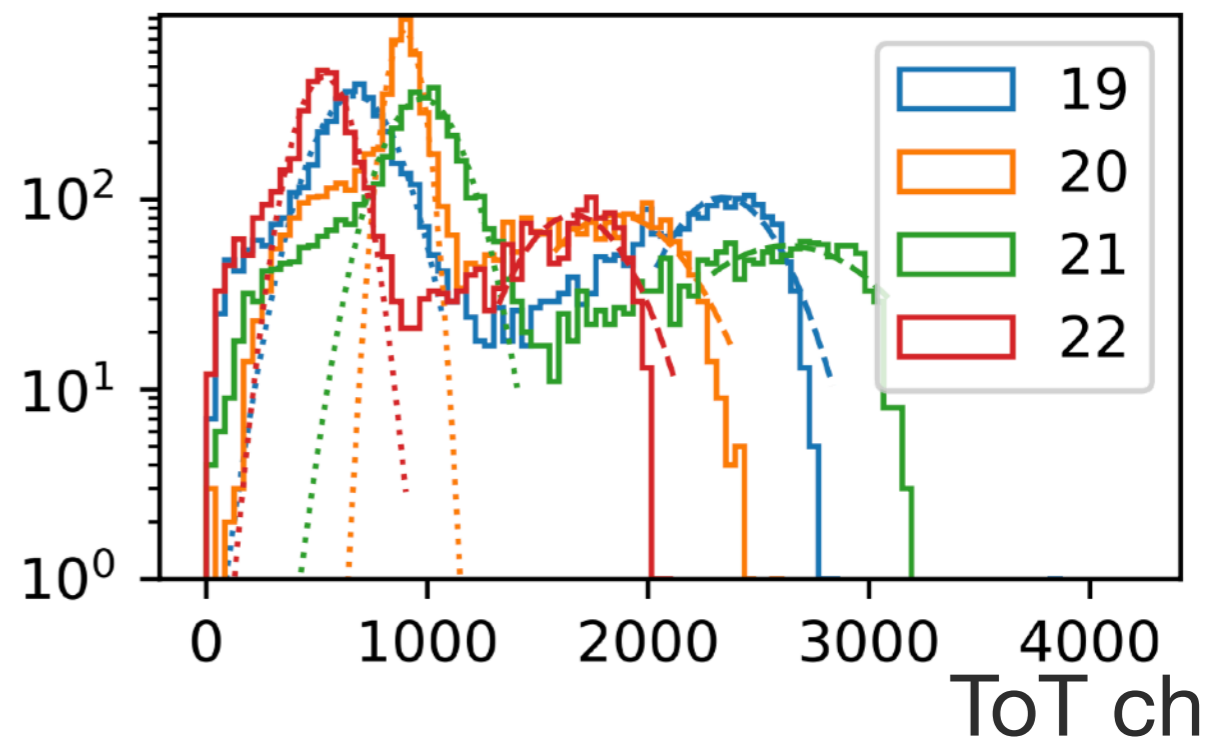
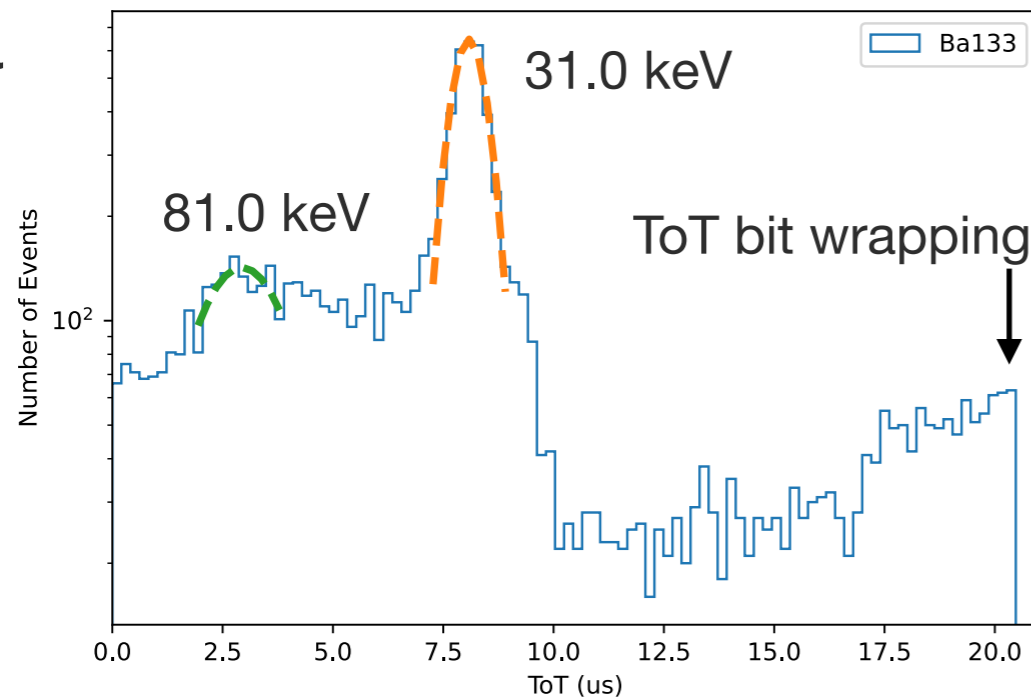
V2

V3

$^{241}\text{Am}$



$^{133}\text{Ba}$

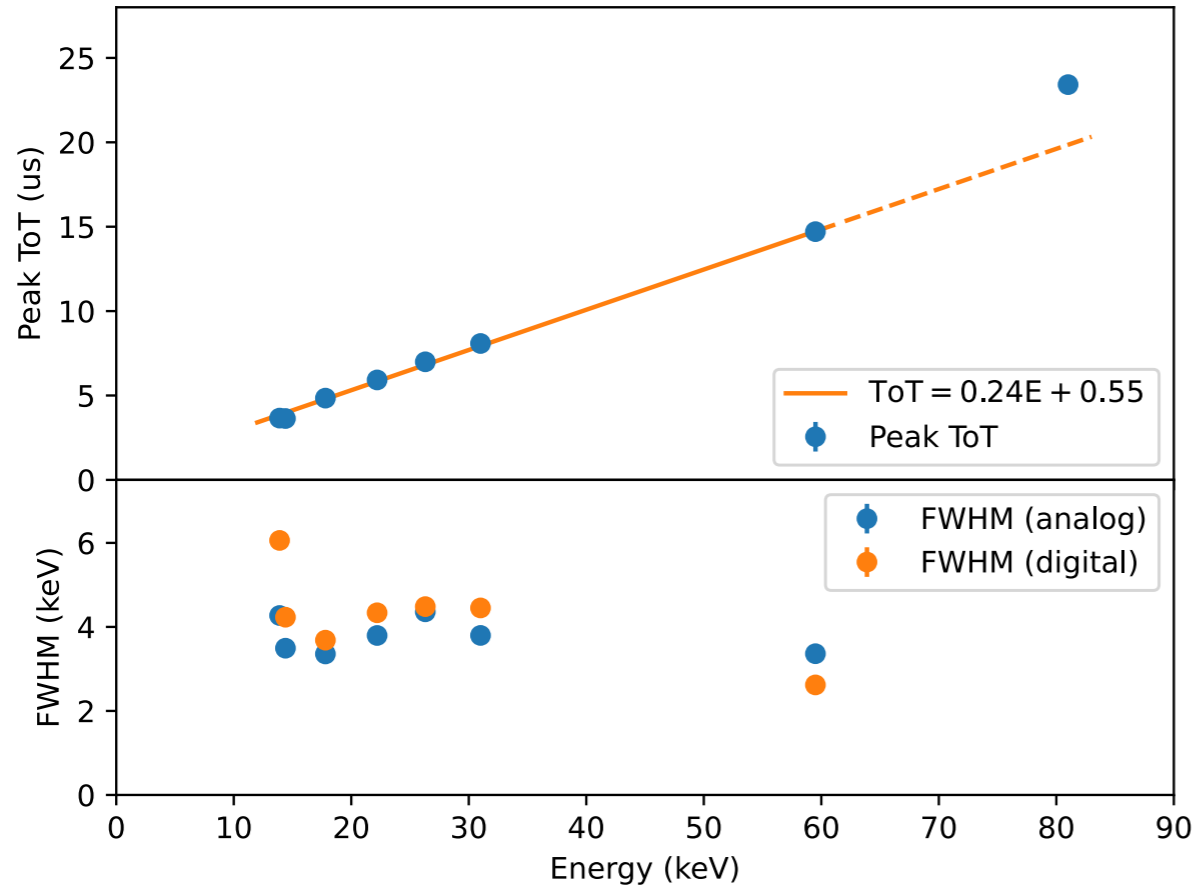




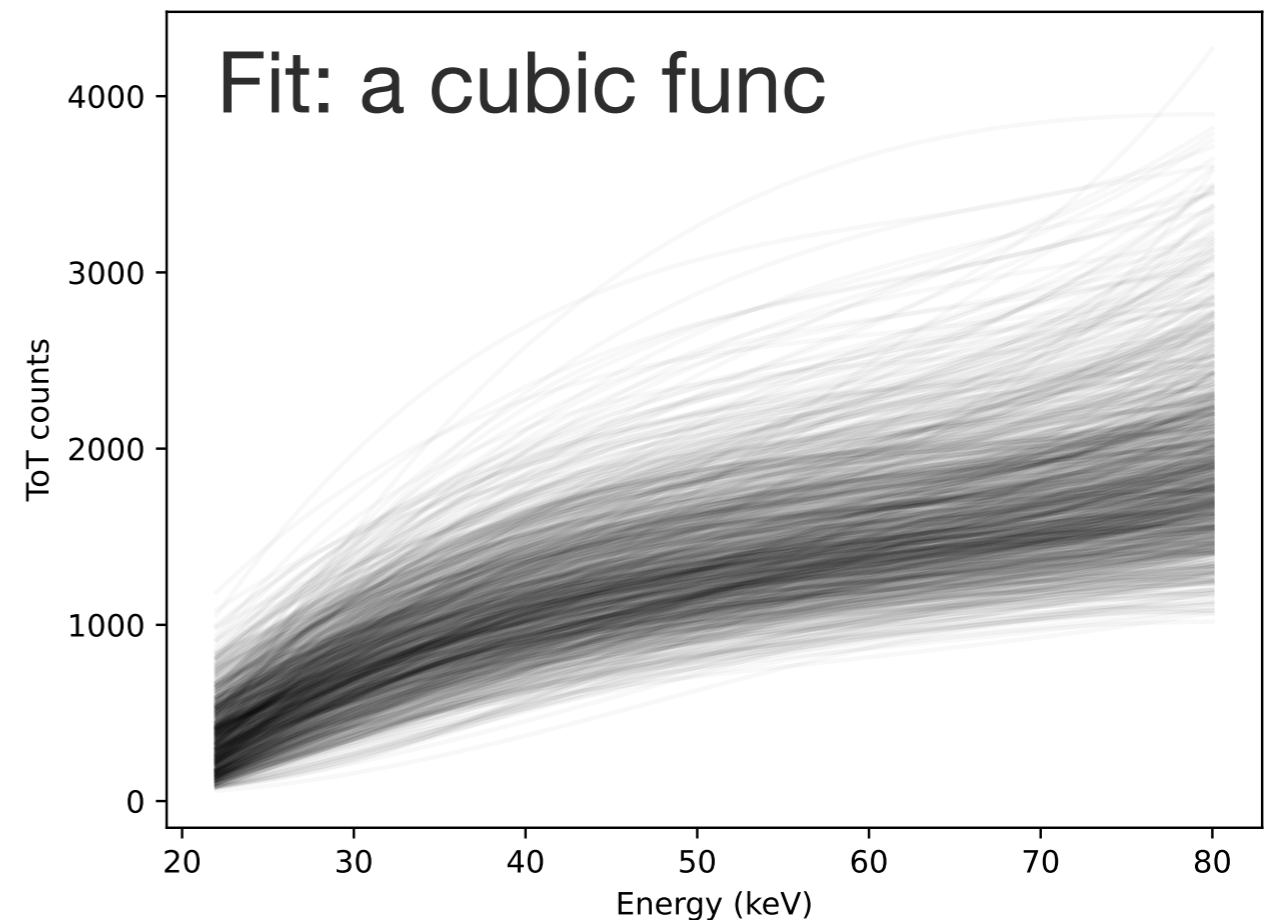
# Energy Calibration

- With updated readout FW/SWs for V3, energy calibration can be performed over the full sensor at once
- Dynamic range: 14 - 80 keV for one pixel of V2 chip  
22 - 80 keV for V3 chip

V2



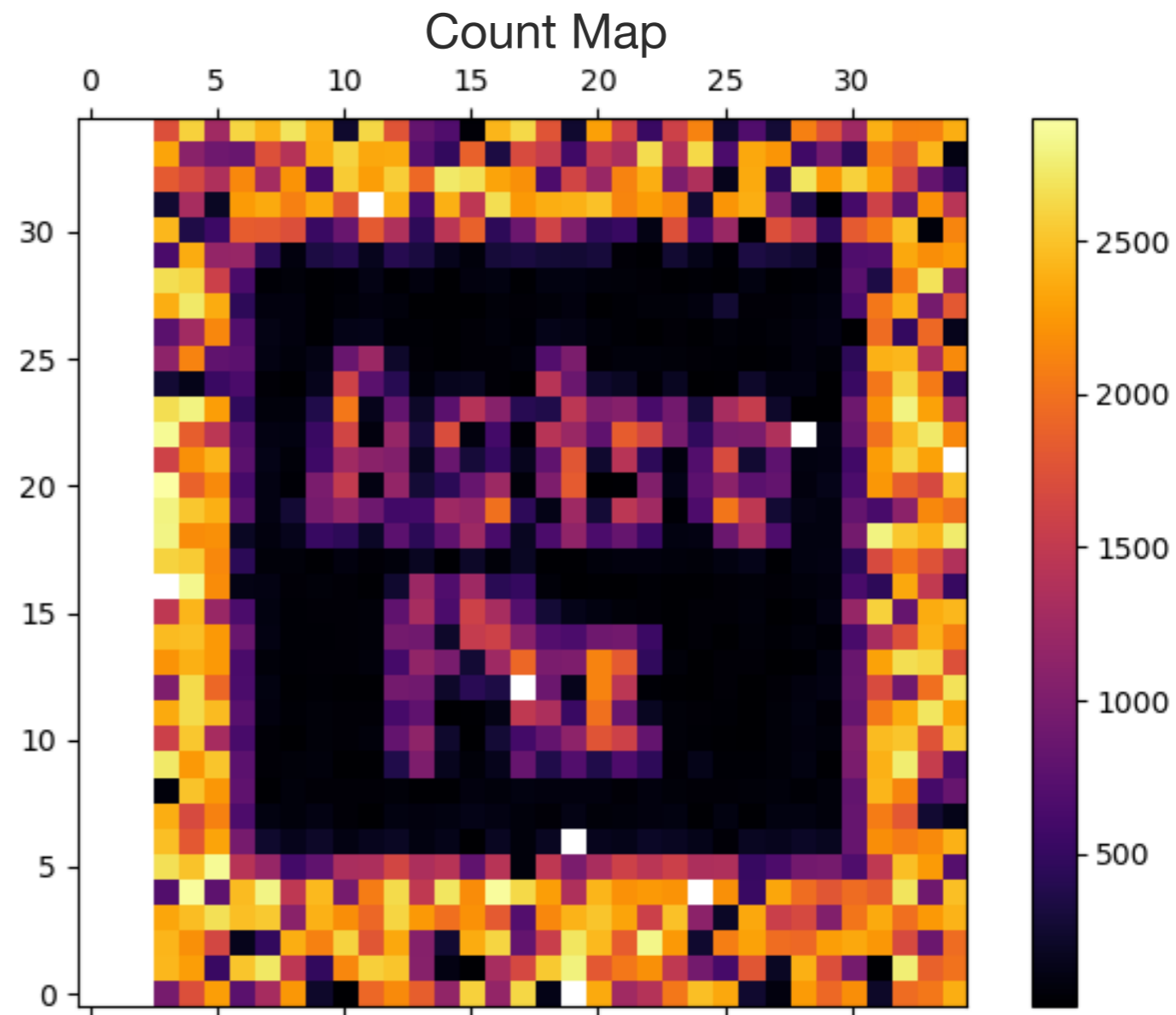
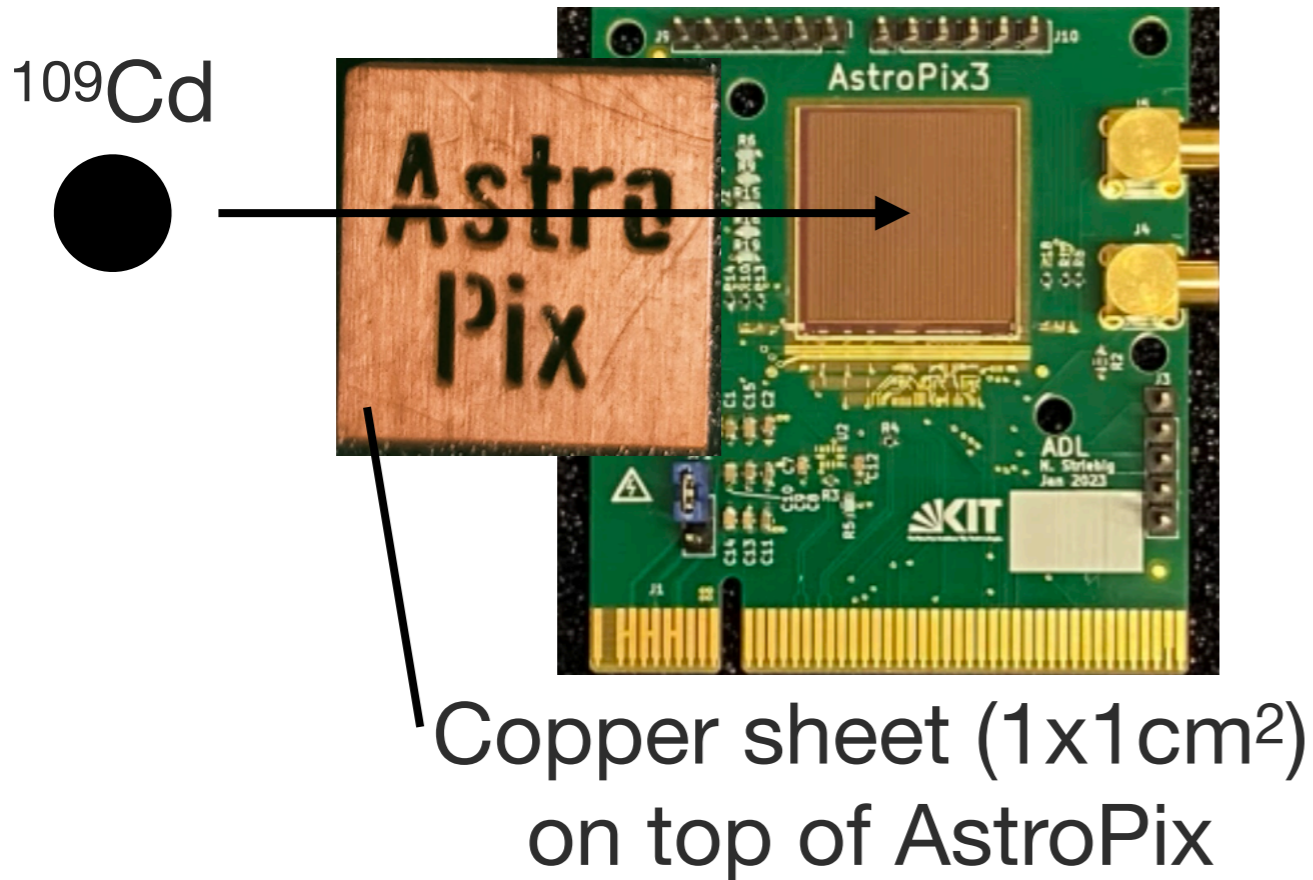
V3





# AstroPix V3: Imaging

- First picture taken by AstroPix
- Only 0.7% of pixels are noisy (masked)



# V3: Depletion Depth Measurements

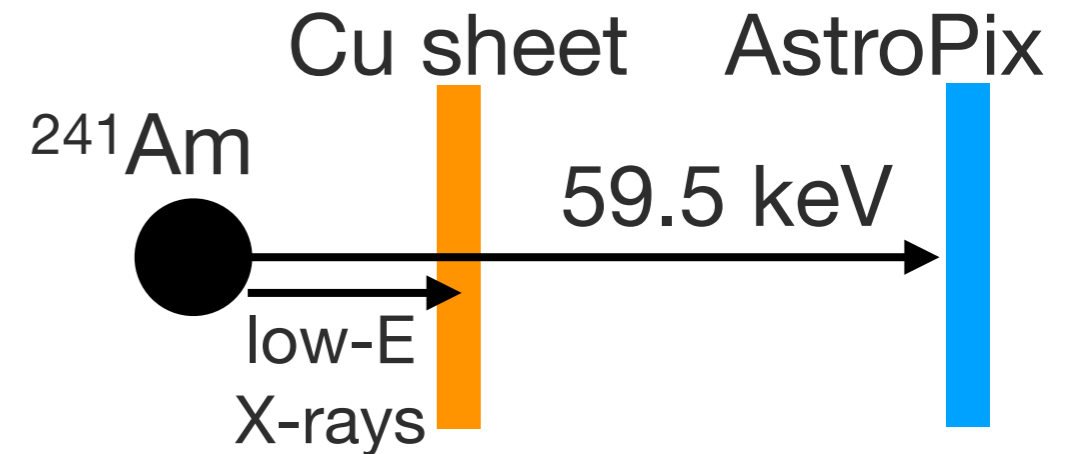
- Estimate from the detection rate of  $^{241}\text{Am}$  59.5 keV events
  - Extract photopeak events from the fitted spectrum
- Compare with a simple model

$$d = \sqrt{2\epsilon\mu\rho(V_{\text{bias}} + V_{\text{built-in}})}$$

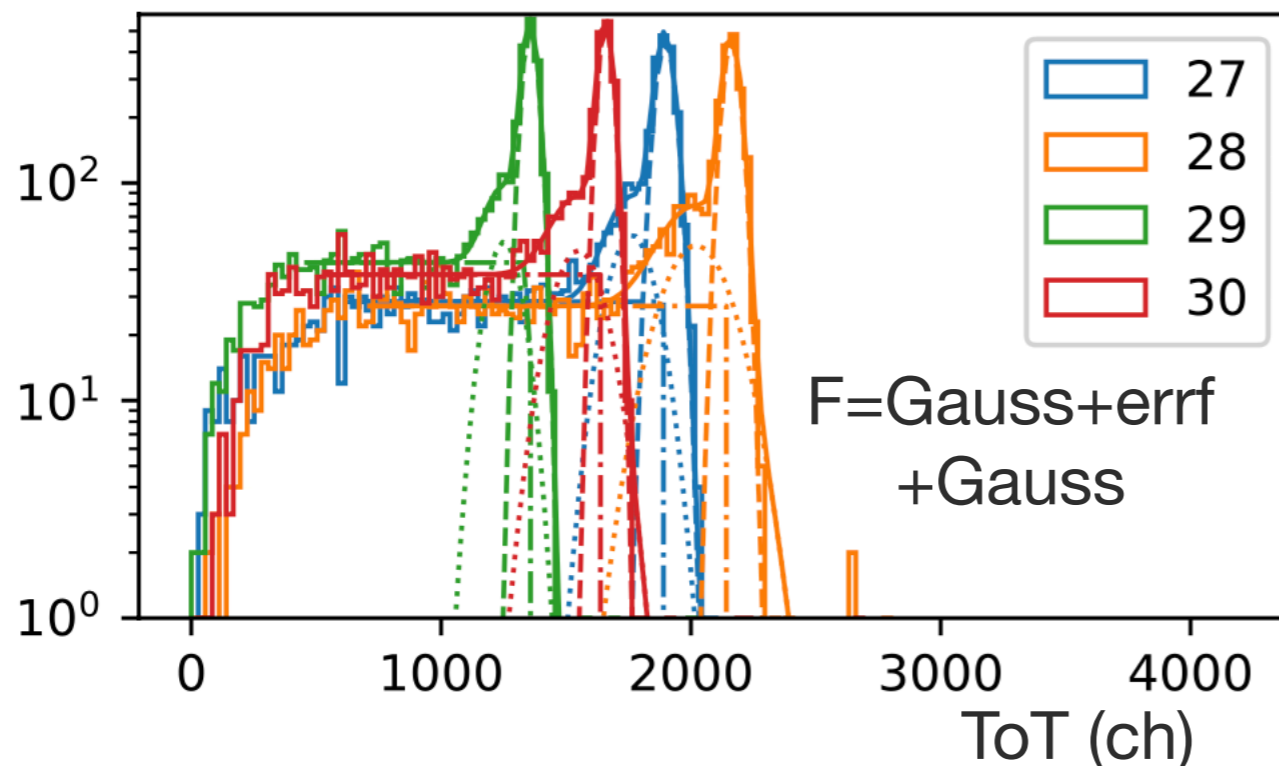
$\epsilon$  : Permittivity       $\mu$  : Hole mobility

$\rho$  : Resistivity 200-400 Ohm cm

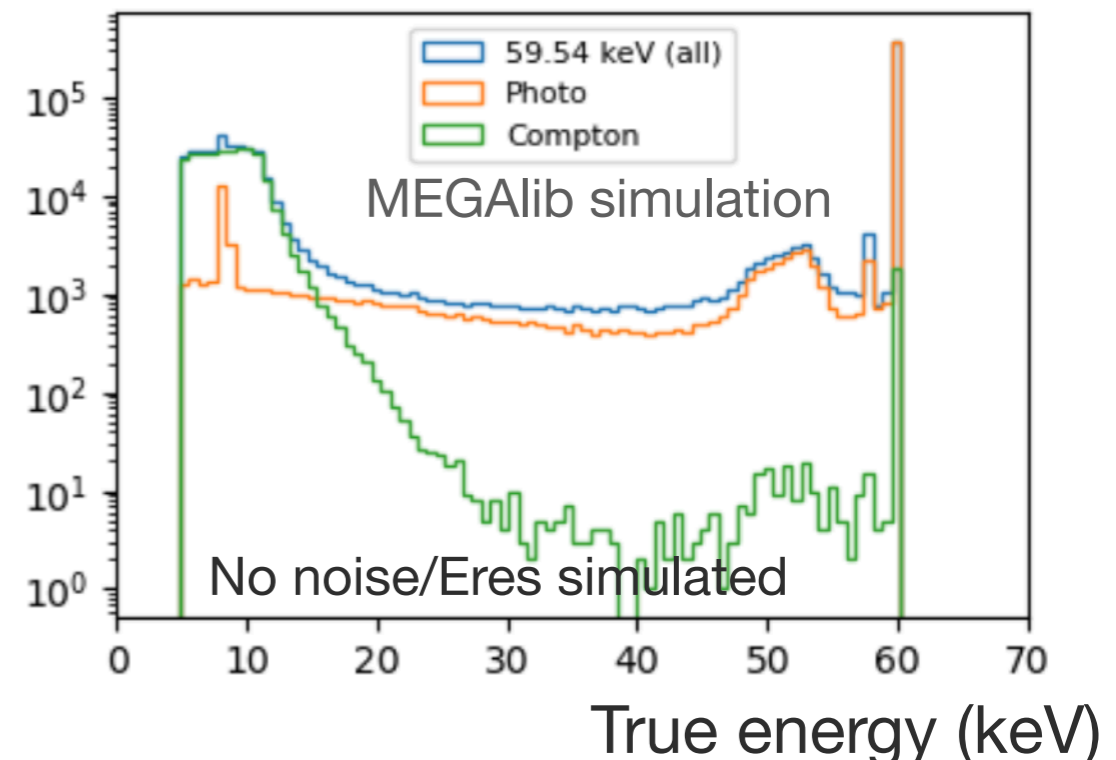
$V_{\text{built-in}}$  : Built-in potencial



Measured  $^{241}\text{Am}$  spectra



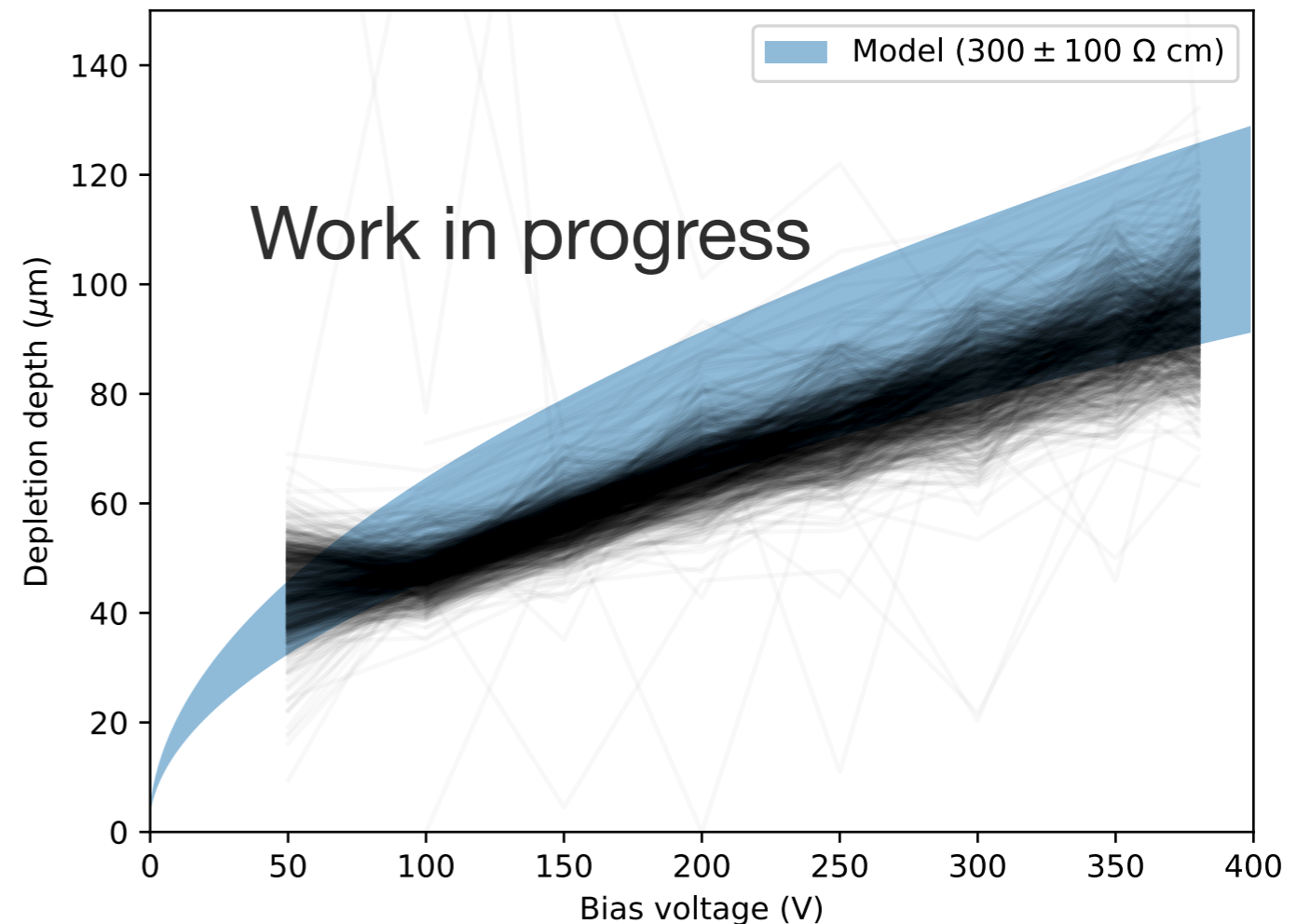
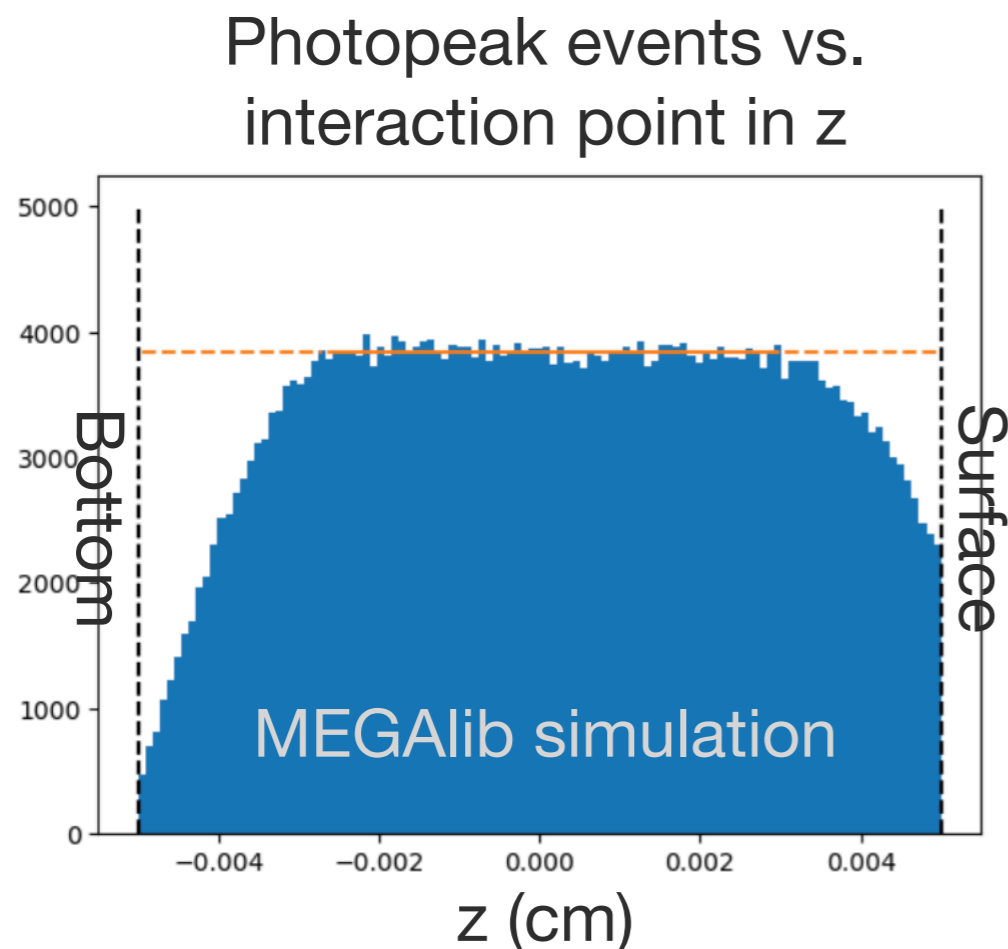
Simulated  $^{241}\text{Am}$  spectra



# V3: Depletion Depth Measurements

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- Sensor volume is corrected for photopeak events
- Measured depths follows the model curve
- Depletion layer develops as expected, but need higher resistivity chips for full depletion

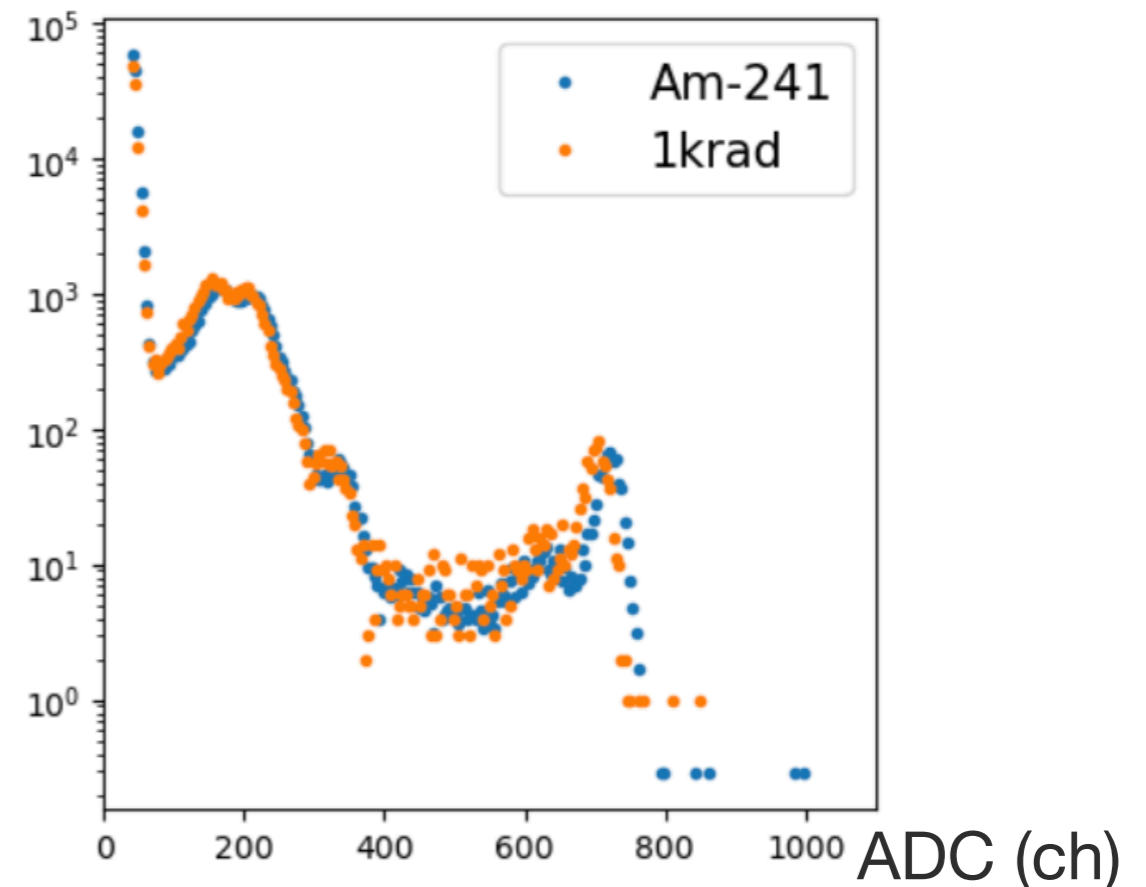




# V2: Radiation Tolerance

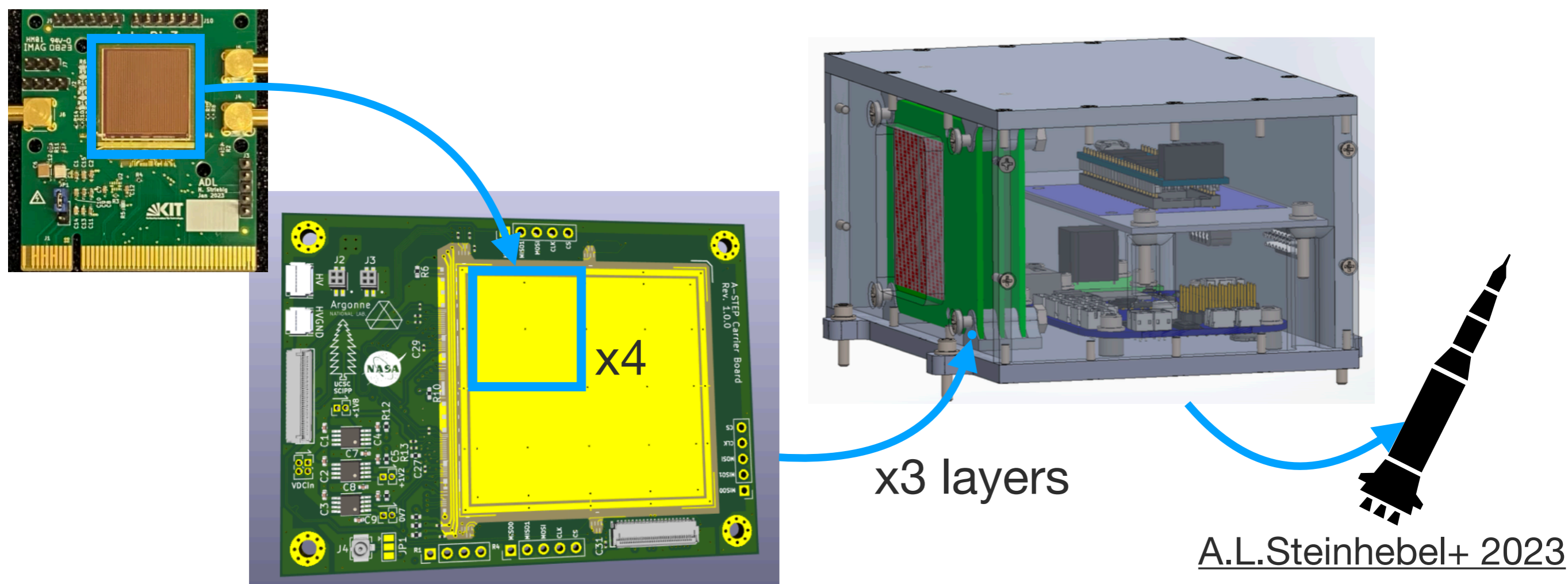
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- Expect  $\sim 1$  krad/yr irradiation on orbit
- One of V2 chips was tested at the high intensity  $^{60}\text{Co}$  facility, Hiroshima U.
- After  $\sim 1$  krad (assumed. TBC) irradiation, the chip works normally, but current draw increases by  $\sim 60\%$  and slight change in gain
- Plan to test with V3 as well



# Quad-Chip

- Minimum component of the AMEGO-X's tracker
- Quad-chip consists of 4 identical V3 chips diced together. Testing will start soon
- Sounding rocket hosted flight (“A-STEP”) is planned in summer 2025 to increase Technical Readiness Level



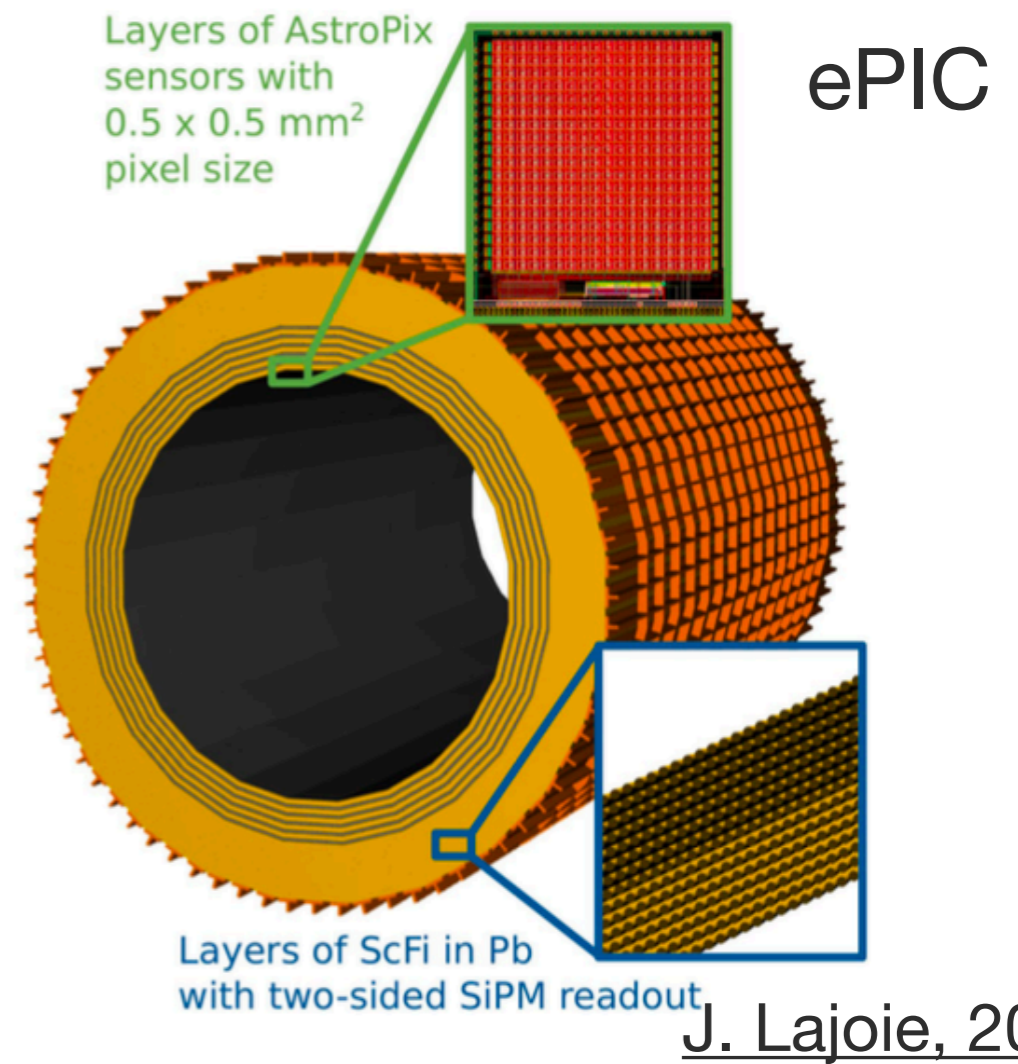
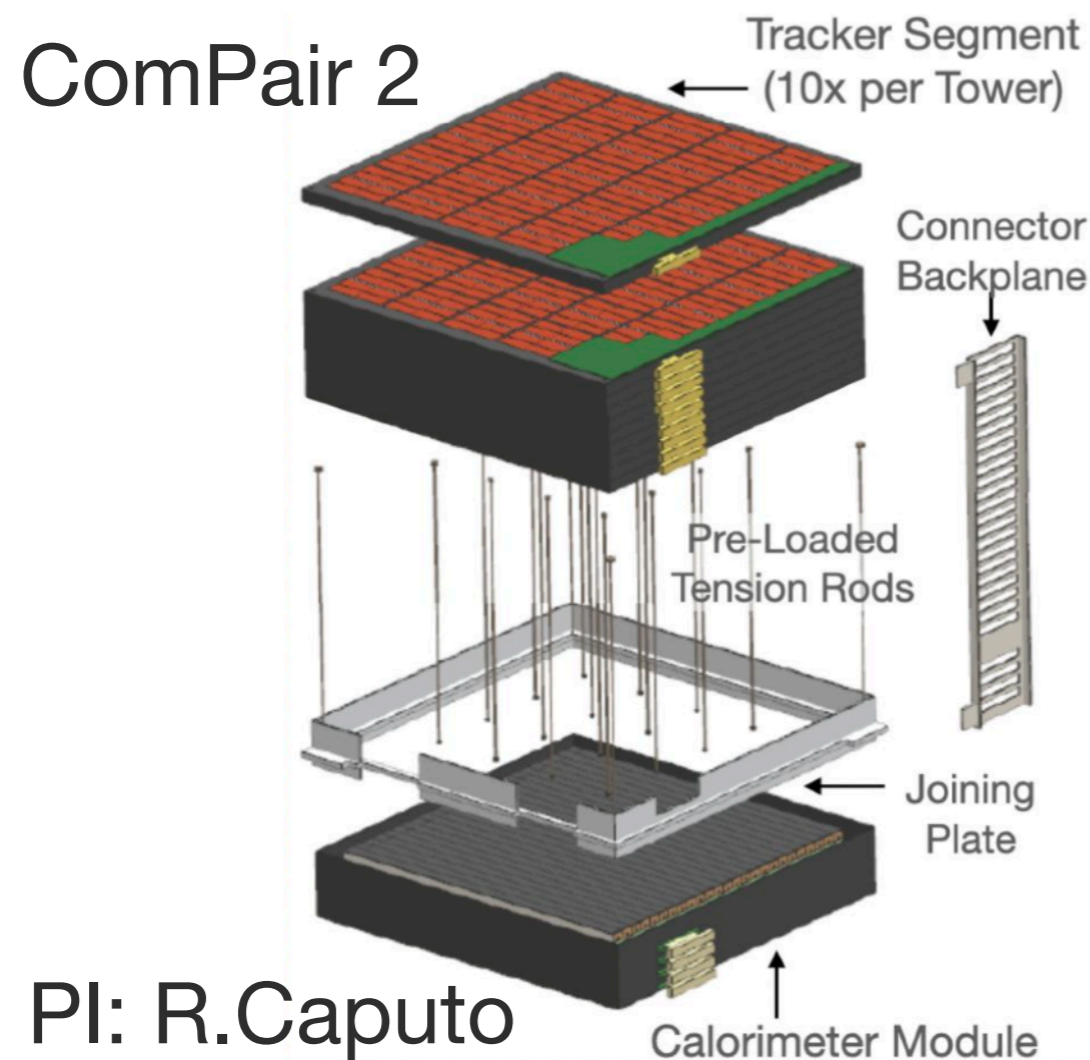
- 500 x 500  $\mu\text{m}^2$  pixel pitch, 1 x 1  $\text{cm}^2$  chip
- Asynchronous timestamp clocks to reduce power consumption
  - Previous: 200 MHz clock  $\rightarrow$  20 MHz clock generated from 2.5 MHz reference clock
- Improved time of arrival resolution: 400 ns  $\rightarrow$  3.125 ns
- Individual tune-DACs for individual pixel calibration/threshold setting
- Will be available to test in early next year

**N.Striebig+ 2023**



# Future Projects

- ComPair 2: Compton-Pair telescope prototype
  - Prototype of AMEGO-X's tracker. Instrument integration in 2026
- ePIC detector in Electron Ion Collider (EIC)
  - Imaging electromagnetic calorimeter. AstroPix delivery in 2029



- AstroPix is a new HV-CMOS active pixel sensor being developed for future gamma-ray telescopes on board
- Pixel pitch in V3 reached to the target size
- Dynamic range is limited to 80 keV. Need to lower gain
- Full energy calibration and full imaging are possible in V3
- Measured depletion depth in V3 shows the depletion layer expands as expected, but only  $\sim 90$   $\mu\text{m}$ . Need high- $\rho$  wafer chips designed to lower leakage current
- V3 quad-chips and V4 chips are coming soon