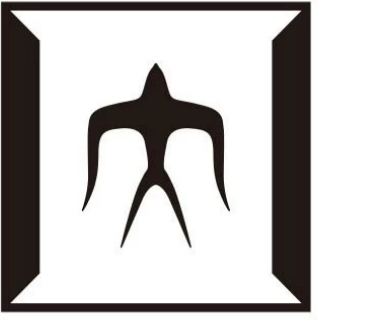


Design and evaluation of a MeV gamma-ray camera aboard a 50-kg class small satellite

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Abstract: Waseda University and Tokyo Tech are developing a **50-kg class small satellite**, tentatively named **INSPIRE**, scheduled for launch in FY2026. The satellite's primary detector is a **BOX-type Compton Camera (CC-BOX)**. This camera covers a dynamic range spanning two orders of magnitude: it observes low energy (30 keV – 200 keV) in **Pinhole mode** and high energy (200 keV – 3 MeV) in **Compton mode**. The CC-BOX comprises a pixelized **Ce:GAGG** scintillator array and an **MPPC array** operating at 40V. The absorber features a four-layer, 20mm thick **depth-of-interaction (DOI)** structure. Additionally, Ce:GAGG elements are positioned on the sides of the detector to enhance its sensitivity than traditional Compton Cameras. For the observation of low-energy gamma rays, the scatter has 5mm square holes to facilitate pinhole imaging. Furthermore, **BGO shields** are positioned on the side and bottom of the detector for efficient event rejection. In this poster, we begin by detailing the design of the CC-BOX, its data processing flow and power specifications. Next, we assessed the detector's anticipated performance through simulations, focusing on detection sensitivity, imaging resolution, energy resolution, and anticipated spectra when observing the Crab Nebula. Finally, we share the findings from a series of environmental tests in preparation for the FY2026 launch.

Introduction

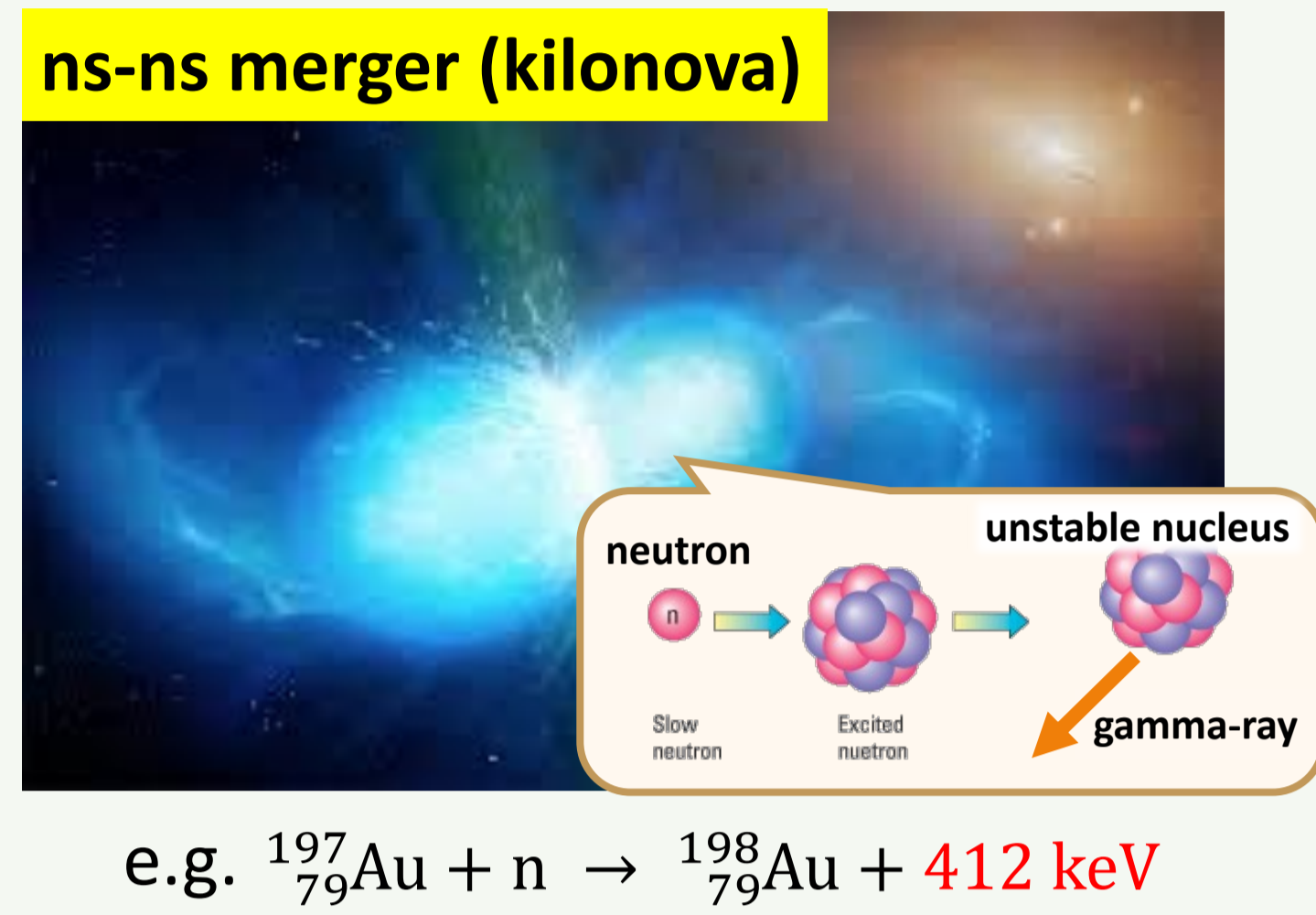
MeV Gamma-ray Astronomy

- ⇒ Nuclear gamma-rays are abundant
- ✓ The origin of rare metals (Pt, Au ++)
- ✓ Blackhole, Gravity waves, kilonova ...

But ...

- ✓ It is necessary to observe in space
- ✓ Background (CXB, Albedo) > Signal

MeV Gamma-ray is difficult to detect



Our Motivation

- ✓ **50-kg class satellite:** Low cost (~2M USD), Short-term development (~5yr)
- ✓ Performance comparable to COMPTEL

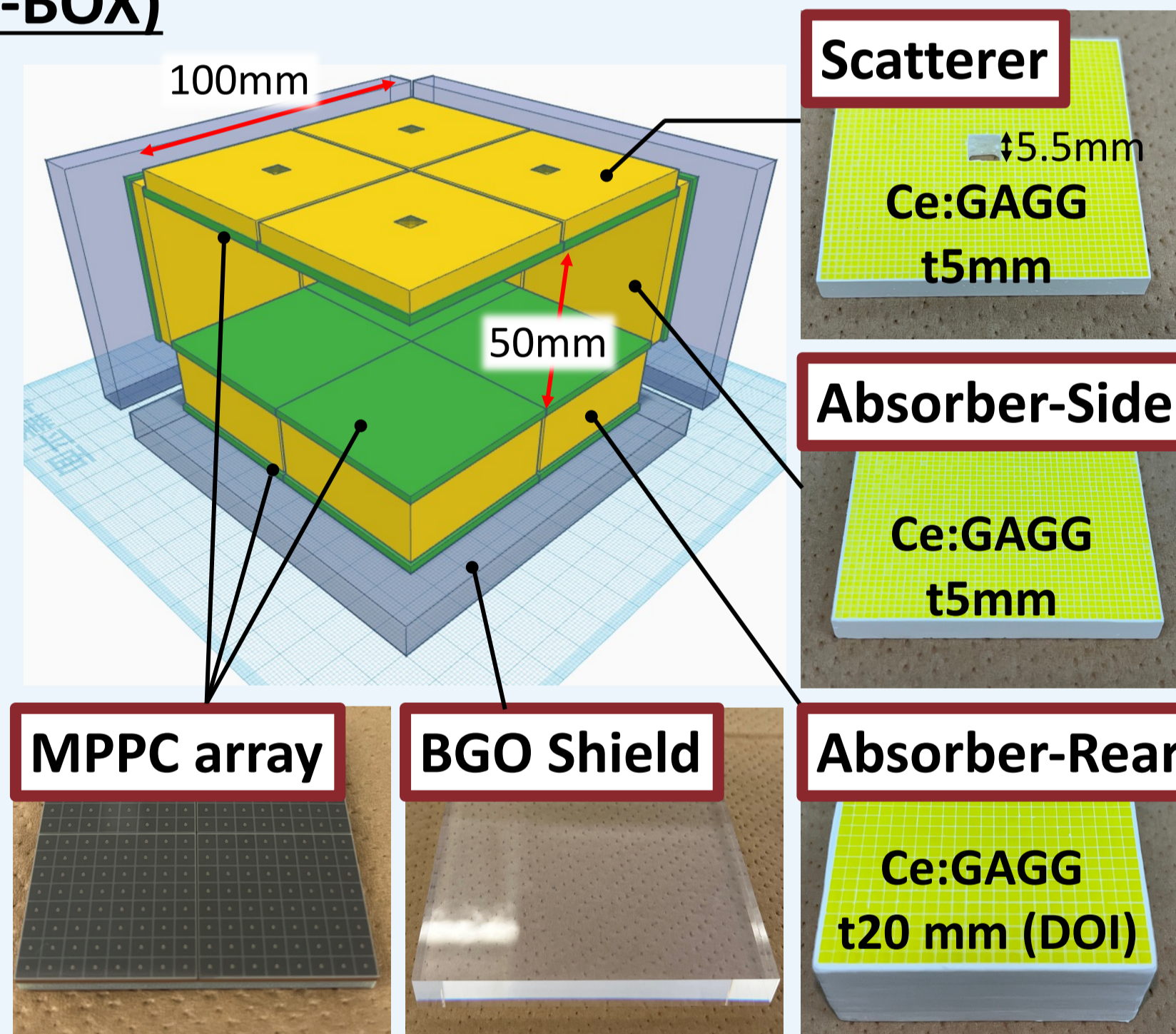
Detector - INSPIRE -

BOX-type Compton Camera (CC-BOX)

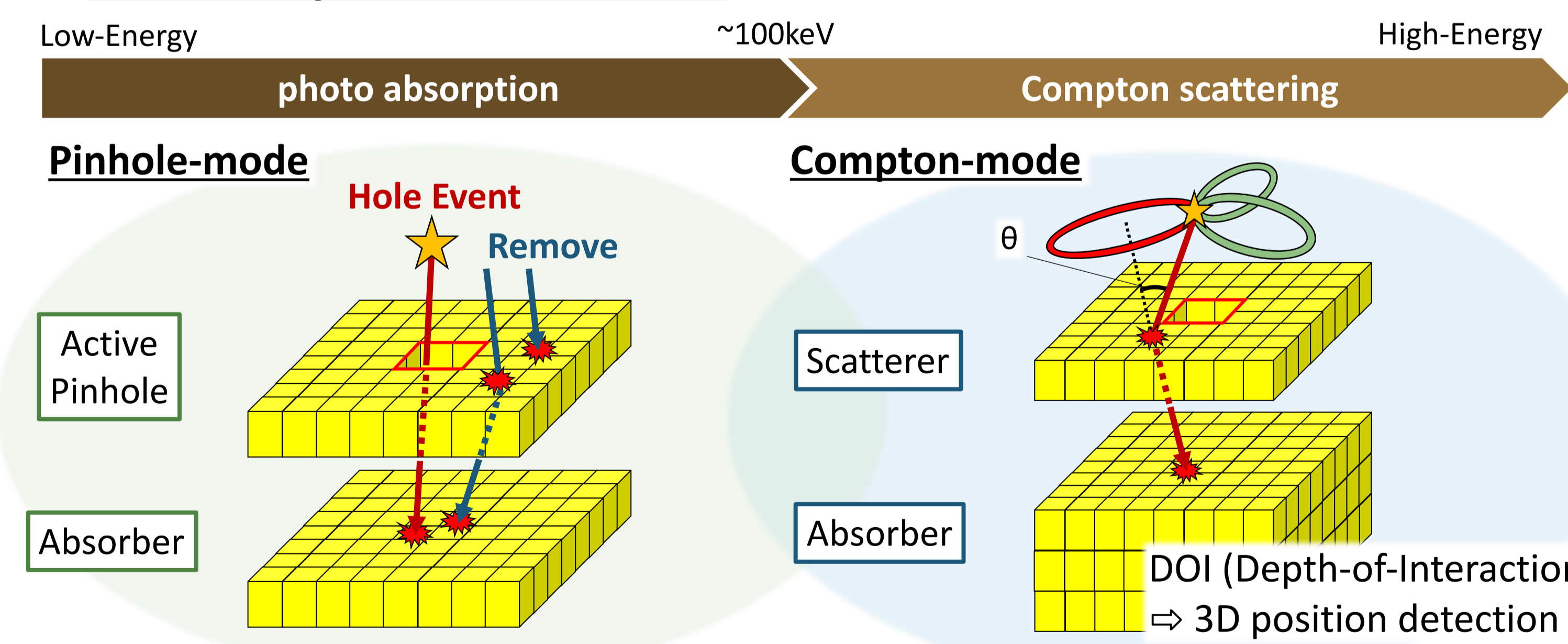
- Scintillator: **Ce:GAGG** array
- veto shield: **BGO** plate
- photon detector: **MPPC** array
- Spatial Resolution**
Scatterer & Side: 1mm Rear: 2mm
- Energy Resolution**
Ce:GAGG: 9%(FWHM) @662keV
BGO: 16%(FWHM) @662keV

characteristic

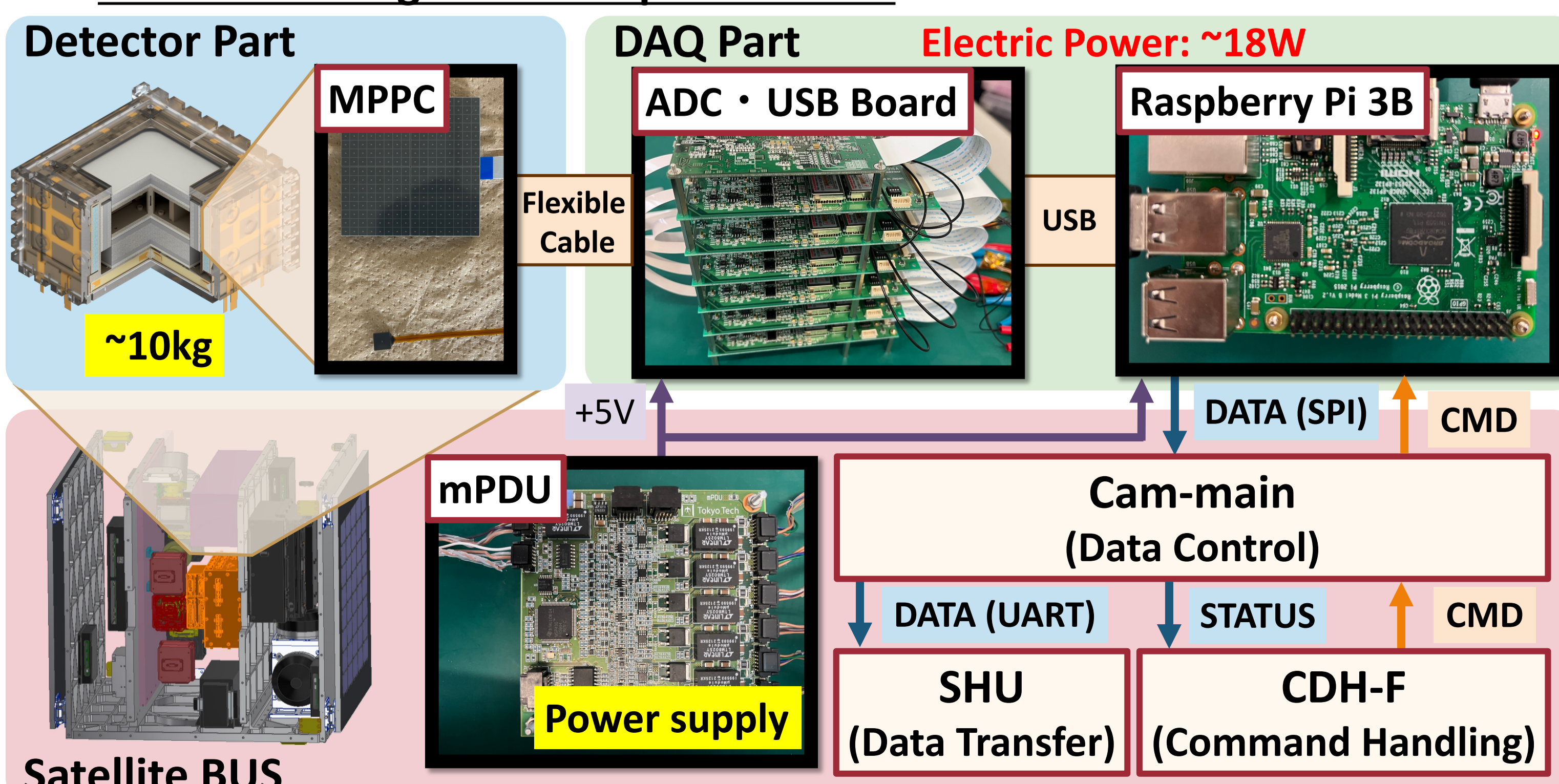
- Active Pinhole: Low-Energy detection
- Rear DOI: Sensitive to high-Energy
- Absorber-side: Higher sensitivity
- BGO shield: BKG/escape event removal



Gamma-ray Detection Method



Data Processing & Power specifications



Performance - Geant4 simulation -

Sensitivity

Continuum Sensitivity

$$S(E) = \frac{f}{\eta(E)} \sqrt{\frac{b(E)}{A \Delta E T}}$$

Line Sensitivity

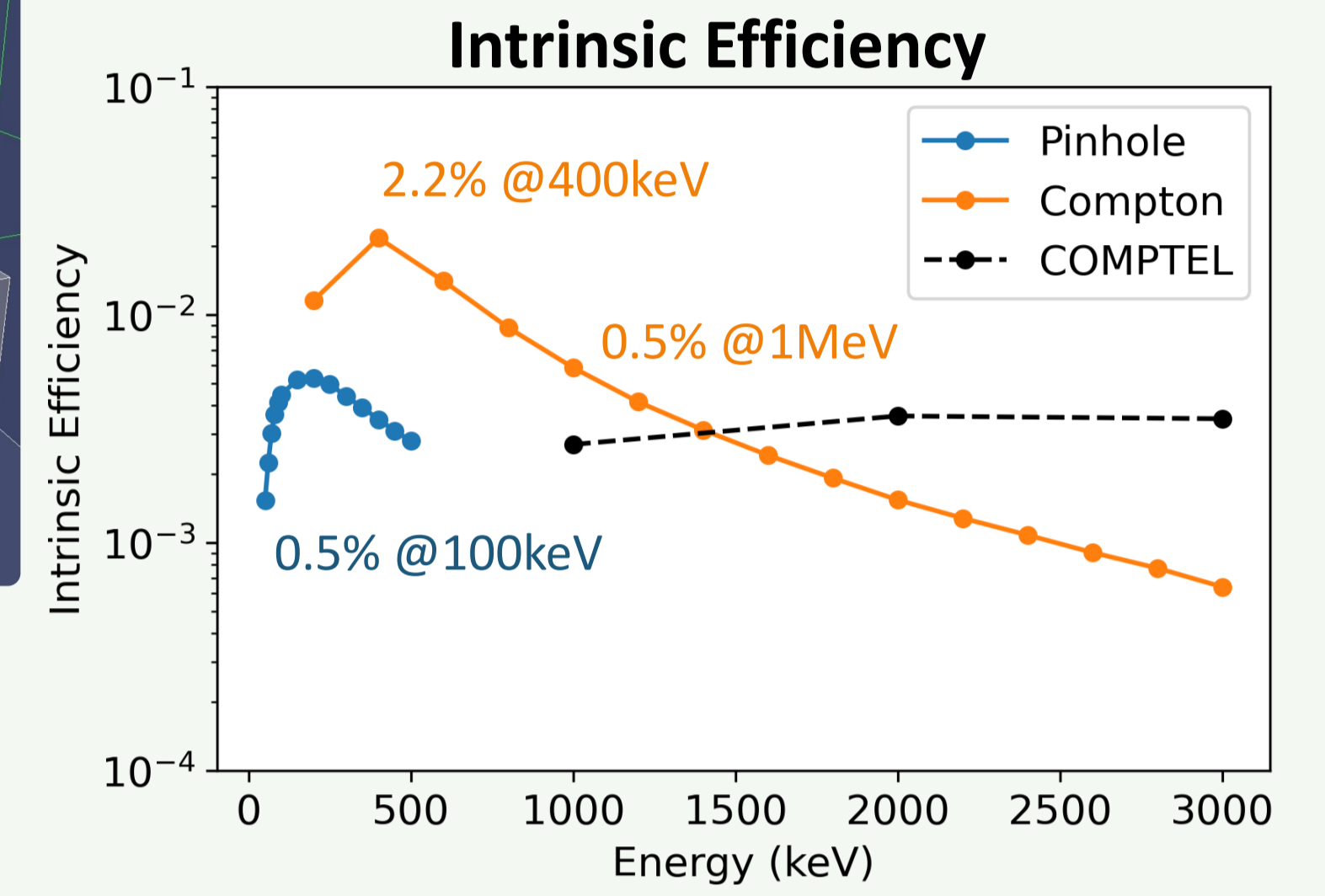
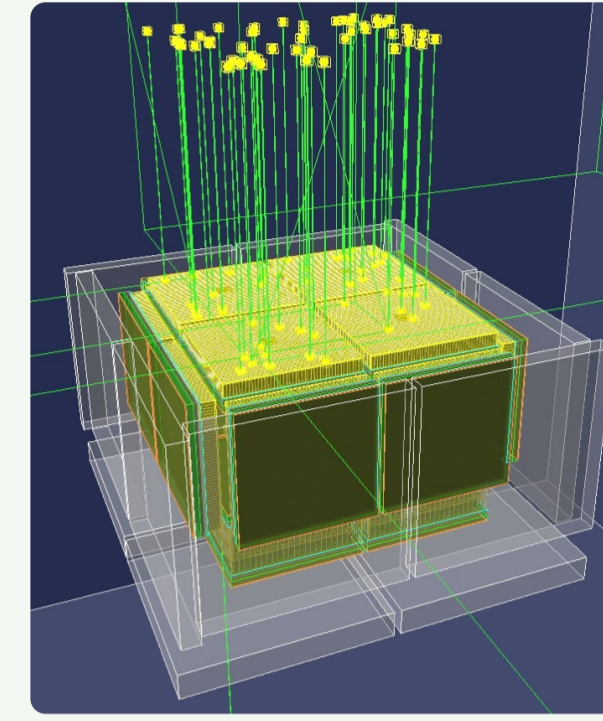
$$S(E) = \frac{f}{\eta(E)} \sqrt{\frac{2b(E)\delta E}{AT}}$$

f: Lower limit of detection σ

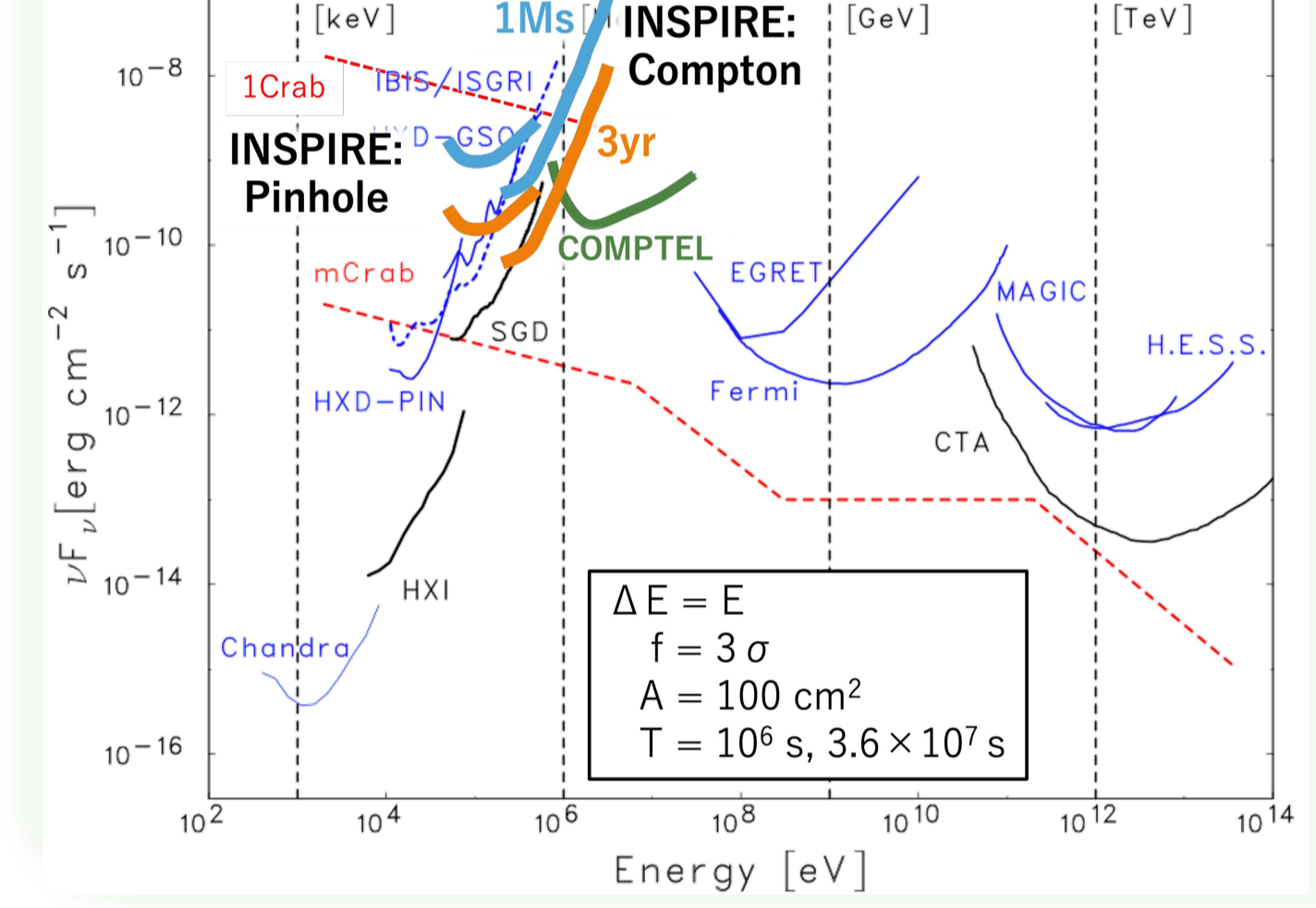
$\eta(E)$: Intrinsic Efficiency

b(E): Detected background (CXB, Albedo)

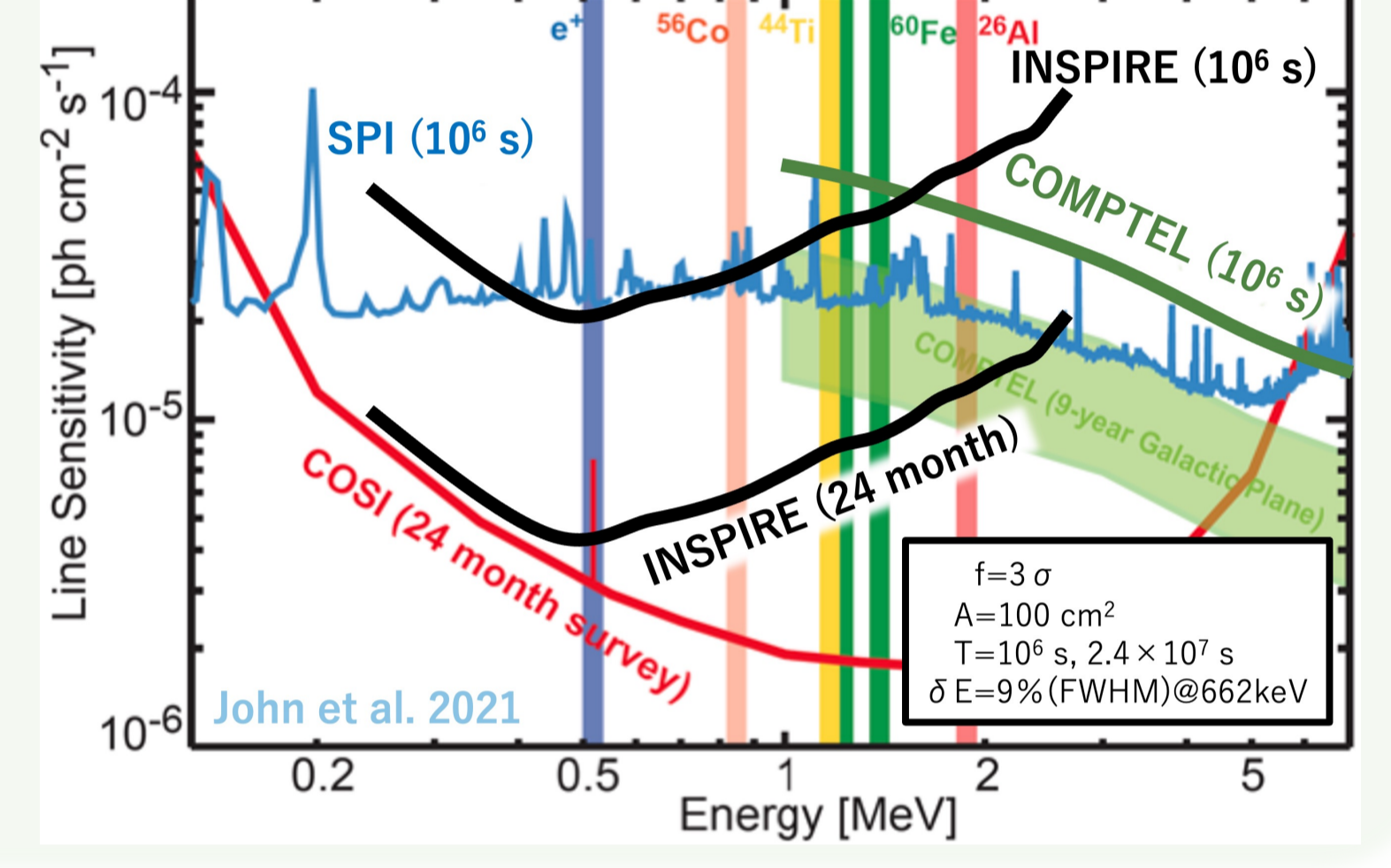
δE : Energy resolution (FWHM)



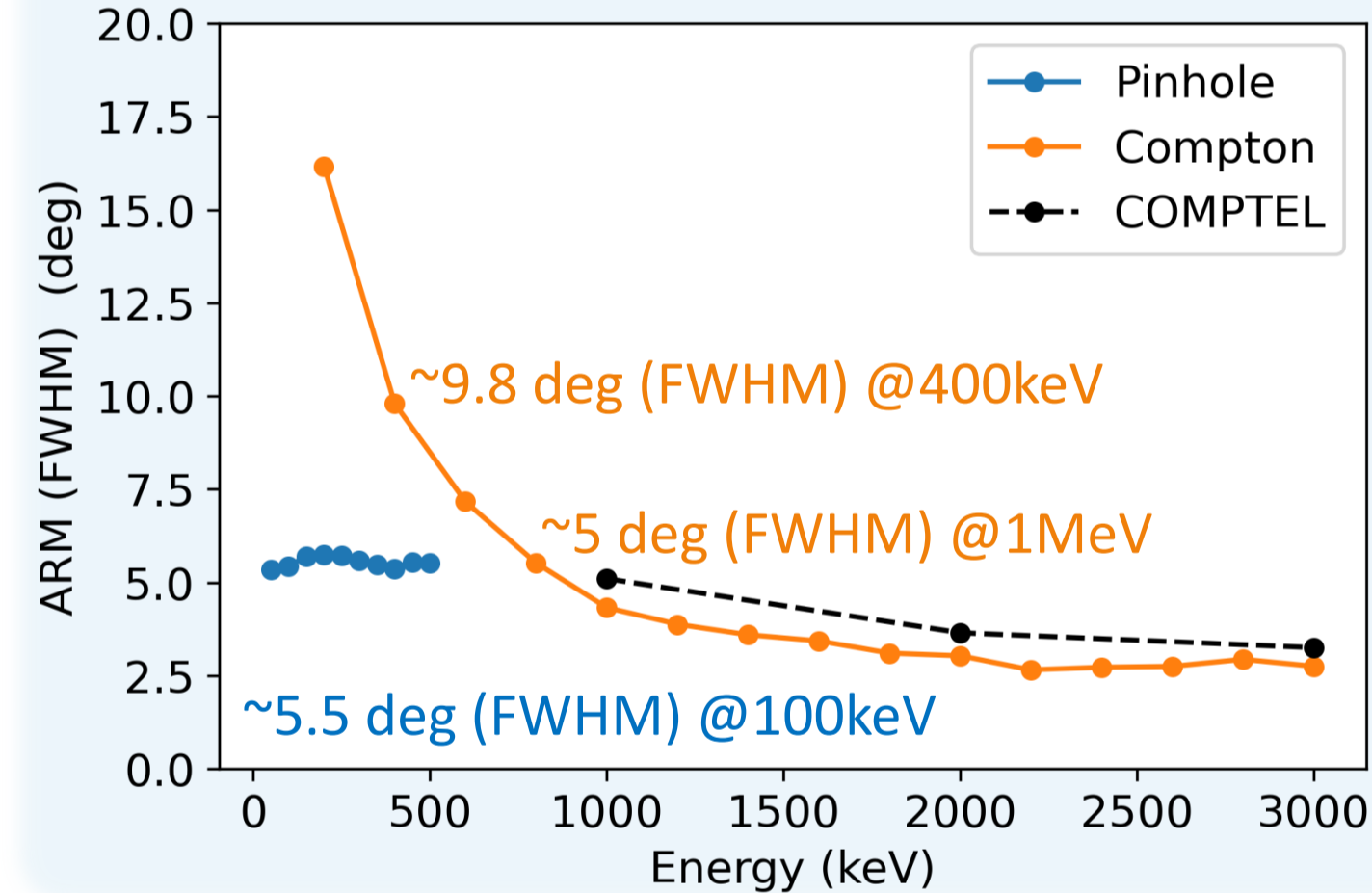
Continuum Sensitivity



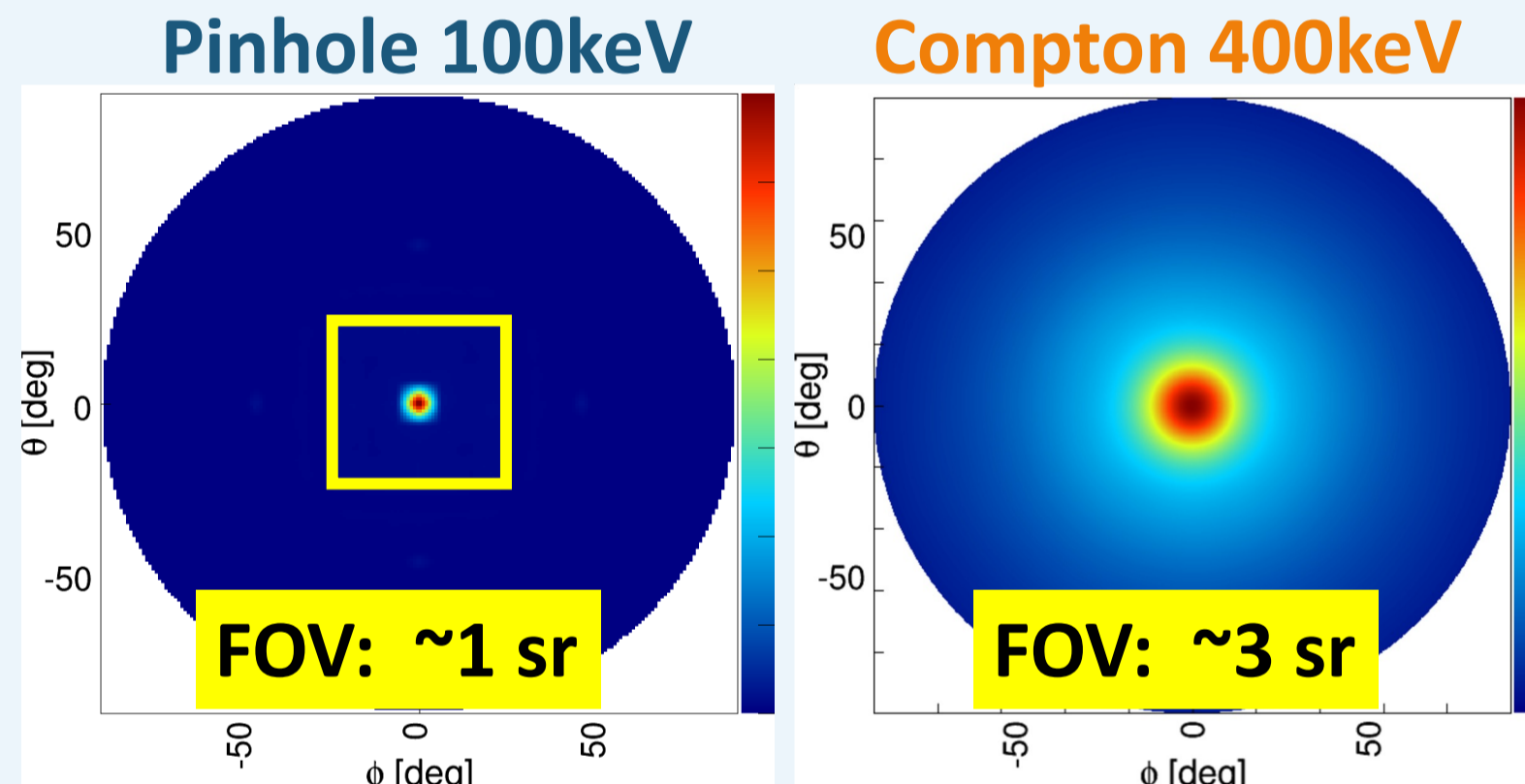
Line Sensitivity



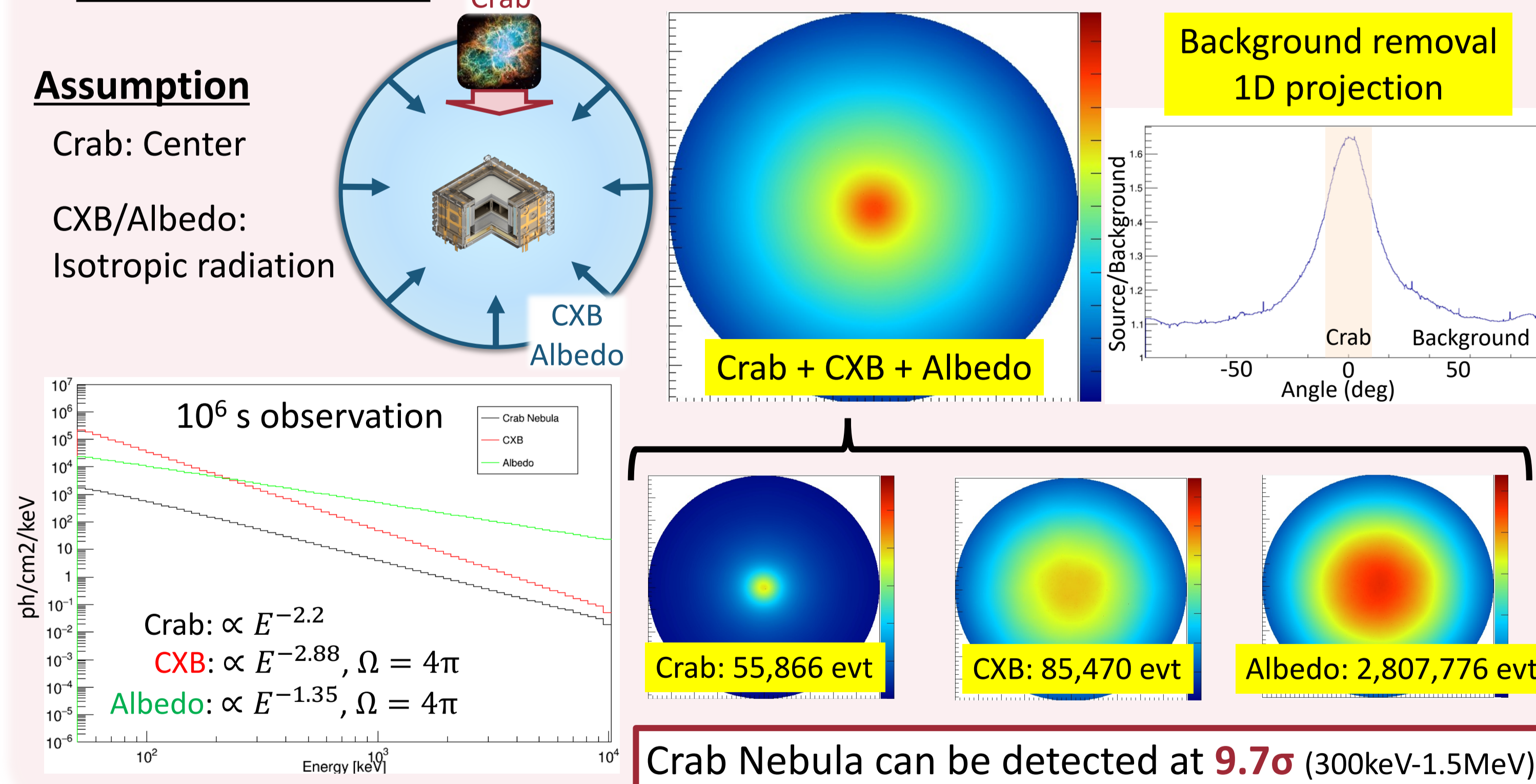
Angular Resolution Measure



Example Image

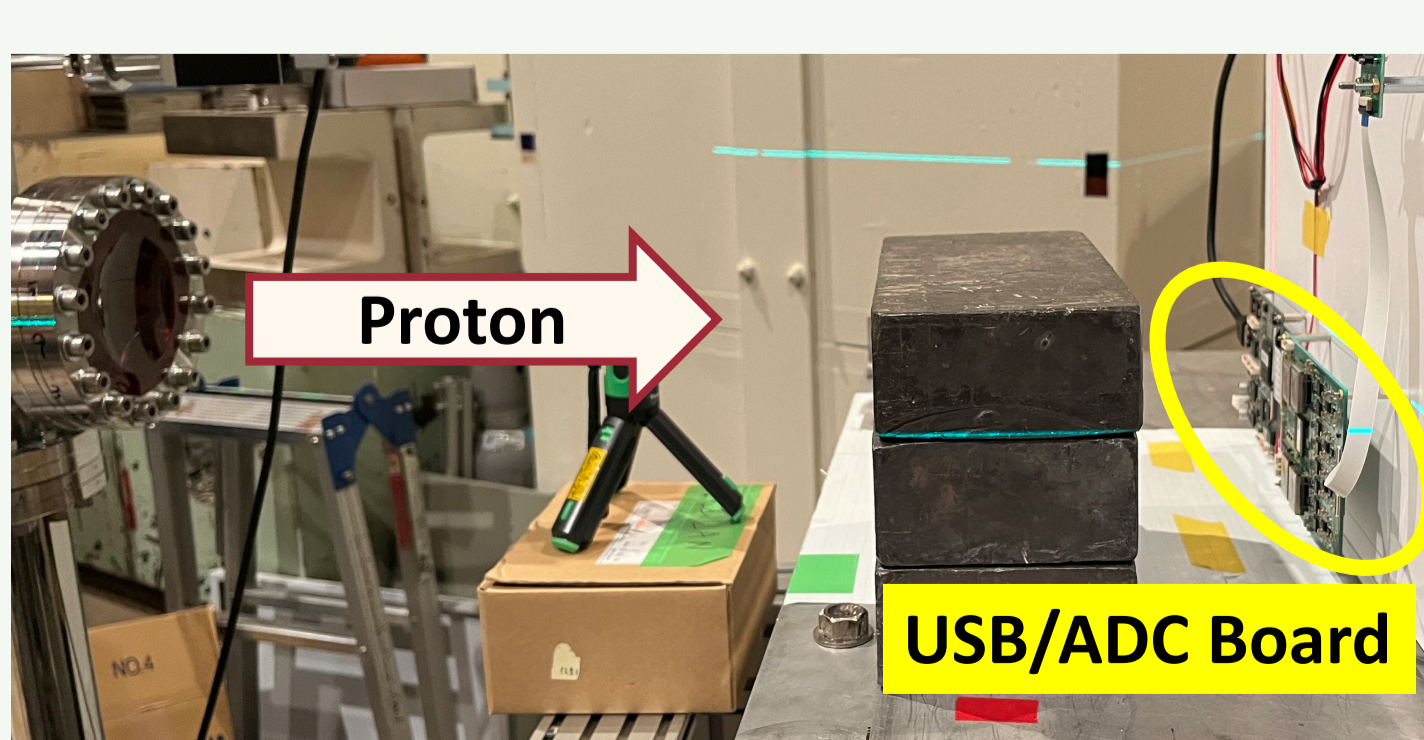


Crab Observation



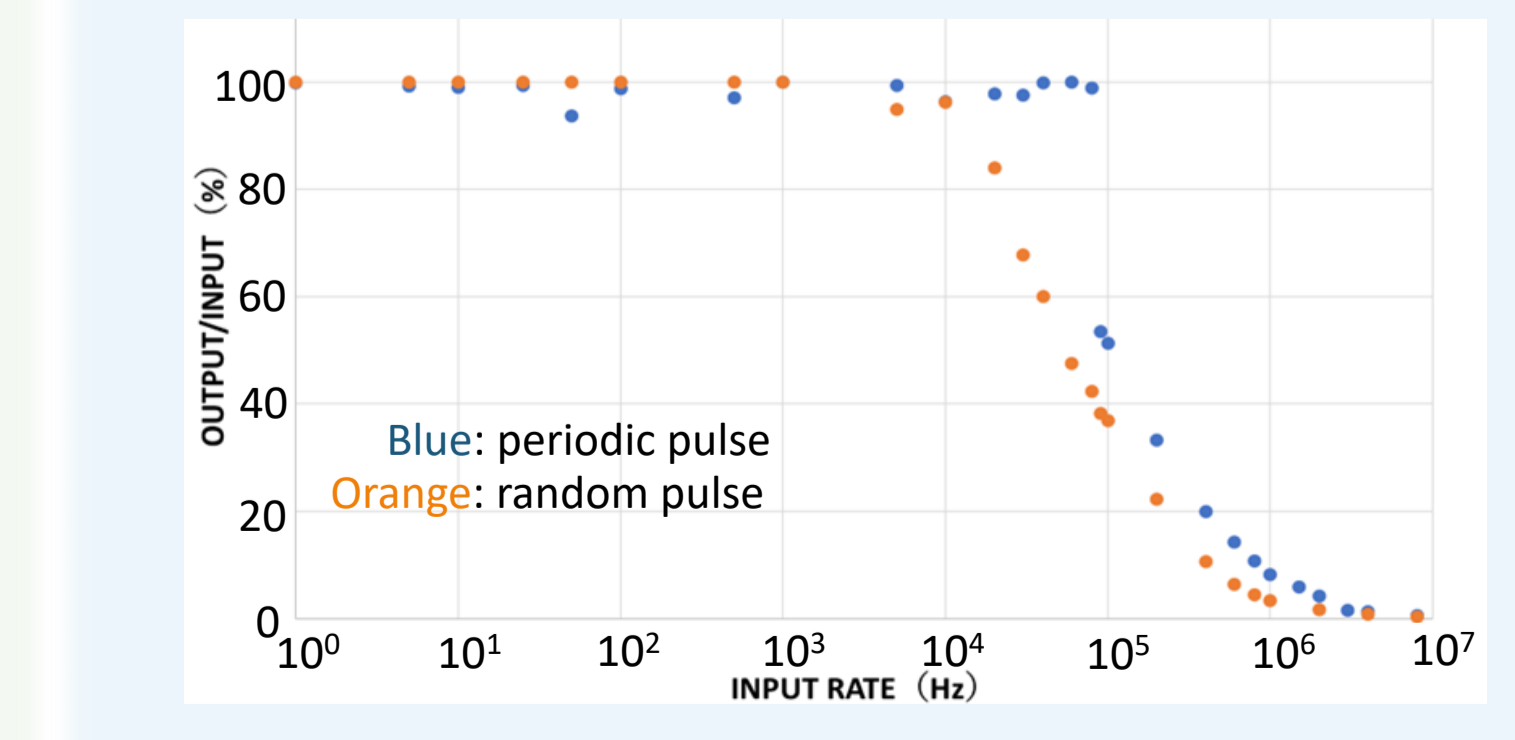
Environmental tests for Launch

Radiation test



- ✓ ~10 years of proton and gamma-ray
- ⇒ SEL / SEU: ~1 / yr
- ⇒ Operation was confirmed after reboot
- ✓ Ce:GAGG doesn't radiate in orbit

Communication test



- ✓ High-rate data acquisition: ~20kHz
- ✓ MPPC signal readout using Raspberry Pi
- Complete engineering model
- ⇒ Vibration test, thermal vacuum test ..