



5th Croatian Teachers Programme CERN, 3-5.4.2023

Uvod u Fiziku elementarnih čestica

Quarks	<i>u</i> up	<i>c</i> charm	<i>t</i> top	γ photon
	<i>d</i> down	<i>s</i> strange	<i>b</i> bottom	<i>g</i> gluon
Leptons	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	<i>Z</i> Z boson
	<i>e</i> electron	μ muon	τ tau	<i>W</i> W boson

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Standard Model of FUNDAMENTAL PARTICLES AND INTERACTIONS

The Standard Model summarizes the current knowledge in Particle Physics. It is the quantum theory that includes the theory of strong interactions (quantum chromodynamics or QCD) and the unified theory of weak and electromagnetic interactions (electroweak). Gravity is included on this chart because it is one of the fundamental interactions even though not part of the "Standard Model."

FERMIIONS

matter constituents
spin = 1/2, 3/2, 5/2, ...

Leptons spin = 1/2		
Flavor	Mass GeV/c ²	Electric charge
ν_e electron neutrino	<1x10 ⁻⁸	0
e electron	0.000511	-1
ν_μ muon neutrino	<0.0002	0
μ muon	0.106	-1
ν_τ tau neutrino		
T tau		

Spin is the quantum number.
Electric charge of the proton

The energy of an electron in cross section $E = mc^2$, where $c = 1.67 \times 10^{-2}$

Baryons
The

Symbol	Name	Quark content	Electric charge	Mass GeV/c ²	Spin
p	proton	uud	1	0.938	1/2
\bar{p}	anti-proton	$\bar{u}\bar{u}\bar{d}$	-1	0.938	1/2
n	neutron	udd	0	0.940	1/2
Λ	lambda	uds	0	1.116	1/2
Ω^-	omega	sss	-1	1.672	3/2

Matter and Antimatter

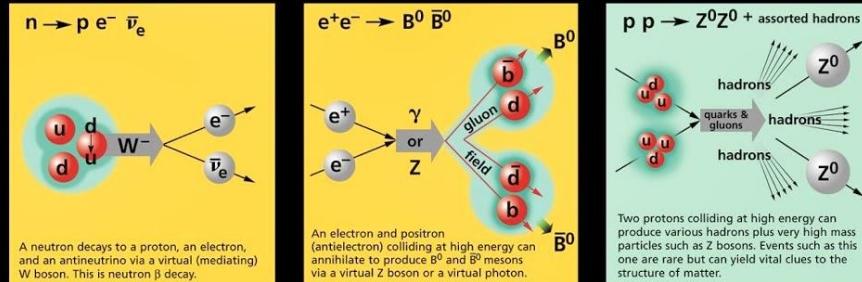
For every particle type there is a corresponding antiparticle type, denoted by a bar over the particle symbol (unless + or - charge is shown). Particle and antiparticle have identical mass and spin but opposite charges. Some electrically neutral bosons (e.g., Z^0 , γ , and $\eta_c = c\bar{c}$, but not $K^0 = d\bar{s}$) are their own antiparticles.

Figures

These diagrams are an artist's conception of physical processes. They are not exact and have no meaningful scale. Green shaded areas represent the cloud of gluons or the gluon field, and red lines the quark paths.

Što znamo o temeljnoj gradnji svemira

Acts on:					
Particles experiencing:		Mass – Energy	Flavor	Electric Charge	Color Charge
Particles mediating:		All	Quarks, Leptons	Electrically charged	Quarks, Gluons
Strength relative to electromagnetism (not yet observed)		10^{-41}	W^+ W^- Z^0	γ	Gluons
for two u quarks at: 10^{-18} m for two protons in nucleus: $3 \times 10^{-17} \text{ m}$		10^{-41}	0.8	1	25
		10^{-41}	10^{-4}	1	60
		10^{-36}	10^{-7}	1	Not applicable to hadrons
					20
See Residual Strong Interaction Note					
Symbol	Name	Quark content	Electric charge	Mass GeV/c ²	Spin
π^+	pion	u <u>\bar{d}</u>	+1	0.140	0
K^-	kaon	s <u>\bar{u}</u>	-1	0.494	0
ρ^+	rho	u <u>\bar{d}</u>	+1	0.770	1
B^0	B-zero	d <u>\bar{u}</u>	0	5.279	0
η_c	eta-c	c <u>\bar{c}</u>	0	2.980	0



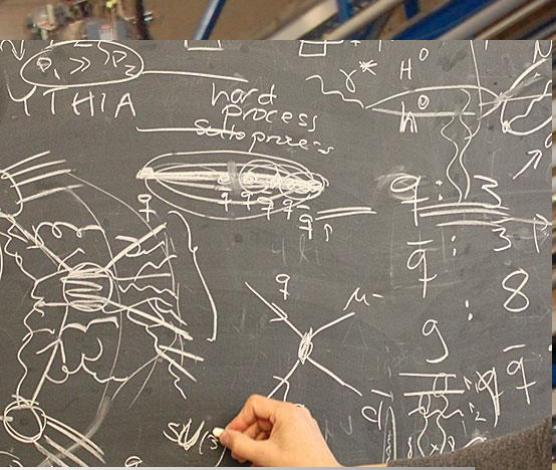
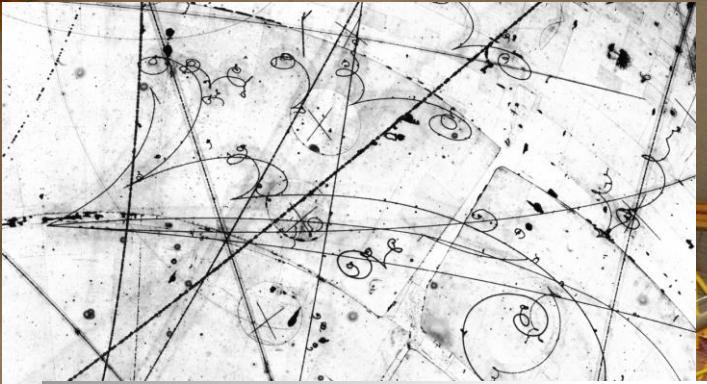
The Particle Adventure
Visit the award-winning web feature [The Particle Adventure](http://ParticleAdventure.org) at <http://ParticleAdventure.org>

This chart has been made possible by the generous support of:

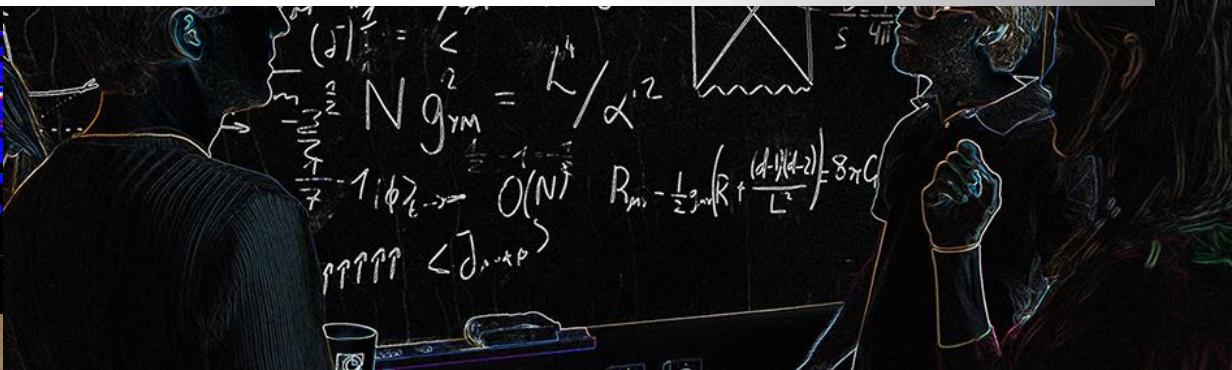
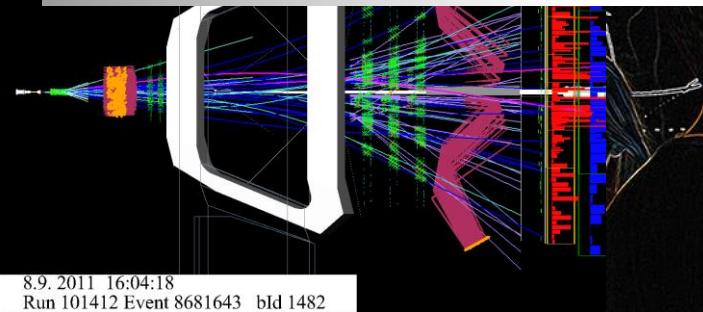
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American Physical Society, Division of Particles and Fields
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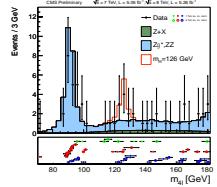
<http://CPEPweb.org>



Preko 100 godina plodonosne suradnje Teorije i Eksperimenta

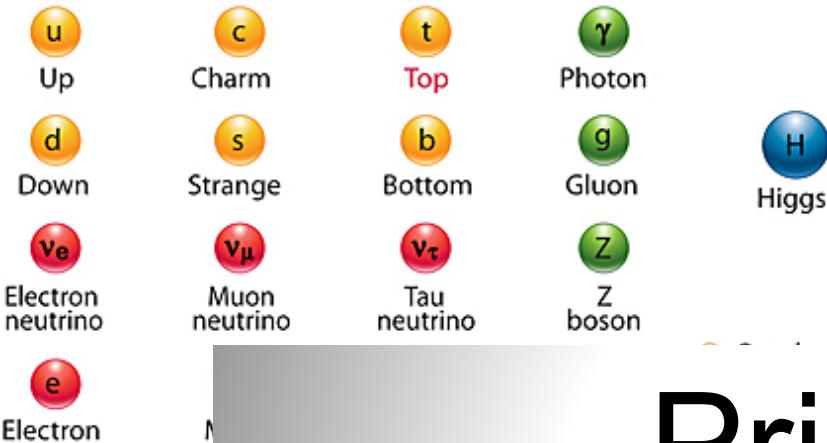


Što su otvorena pitanja?



$$\begin{aligned}\mathcal{L} = & -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\ & + i \bar{F} \partial^\mu F + h.c. \\ & + \chi_1 Y_1 \chi_2 \phi + h.c. \\ & + |\partial_\mu \phi|^2 - V(\phi)\end{aligned}$$

The Standard Model



Pričam

Profesorima fizike

$$\begin{aligned} & -I_f^3 \gamma_5) \bar{\Psi}_f \Psi_f + \\ & -ie(W_\mu^+ A_\nu - W_\nu^+ A_\mu) + ig' c_w (W_\mu^+ Z_\nu - W_\nu^+ Z_\mu|^2 + \\ & -\frac{1}{4} |\partial_\mu Z_\nu - \partial_\nu Z_\mu + ig' c_w (W_\mu^- W_\nu^+ - W_\mu^+ W_\nu^-)|^2 + \\ & -\frac{1}{2} M_\eta^2 \eta^2 - \frac{g M_\eta^2}{8 M_W} \eta^3 - \frac{g'^2 M_\eta^2}{32 M_W} \eta^4 + |M_W W_\mu^+ + \frac{g}{2} \eta W_\mu^+|^2 + \\ & + \frac{1}{2} |\partial_\mu \eta + i M_Z Z_\mu + \frac{ig}{2c_w} \eta Z_\mu|^2 - \sum_f \frac{g}{2} \frac{m_f}{M_W} \bar{\Psi}_f \Psi_f \eta \end{aligned}$$

Zahvaljujem

- U FEČ smo naučili surađivati
- Jedan dio slajdova, ideja, prikaza sam “ukrao” od:
 - Rolf Landua
 - Marie-Helene Schune
 - Ivica Puljak
 - Bill Murray
 - Caroline Collard
 - i puno drugih...
 - Oni lošiji su vjerojatno moji ☺
- Sad ste dio CERN-ove kulture: možete ih dalje koristiti, kopirati, prilagoditi...
 - pptx će biti na raspolaganju

Fizika: eksperimentalna znanost

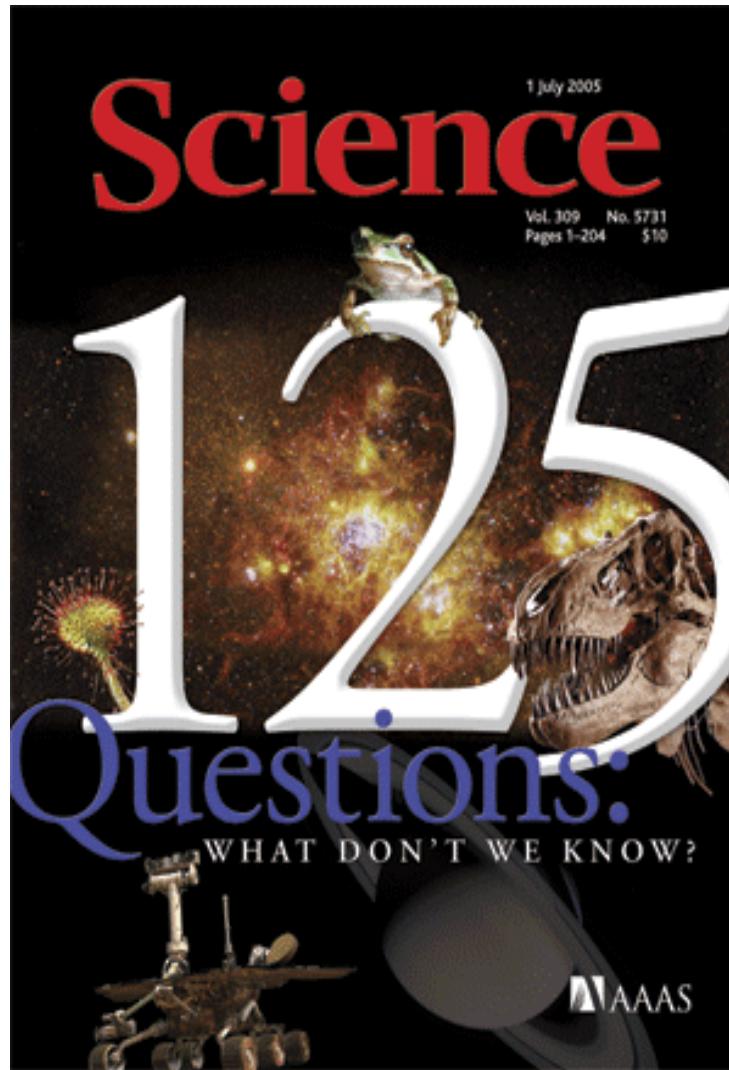


R. P. Feynman

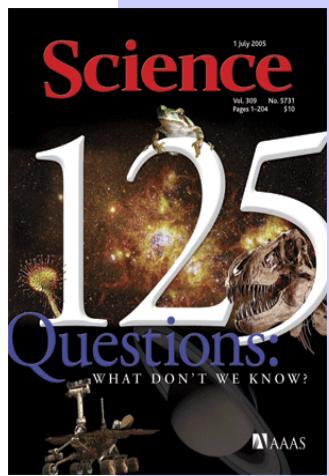
“Physicists learned to realize that whether they like a theory or they don’t like a theory is not the essential question. Rather, it’s whether or not the theory gives predictions that agree with experiments”

Richard Feynman

Velika pitanja!



➤ Pitanje broj 1: **Od čega se sastoji svemir?**



INTRODUCTION AND OPENING ESSAY

What Don't We Know?

D. Kennedy and C. Norman

In Praise of Hard Questions

T. Siegfried

Online Extras

Be sure to check out these online extras related to our 125th Anniversary Issue:

Your Turn

Did we miss your favorite scientific conundrum? Visit our [special online forum](#) to comment on our 125 questions or nominate your own choice.

Anniversary Editorial

Science Editor-in-Chief Donald Kennedy celebrates the magazine's

THE QUESTIONS

The Top 25

Essays by our news staff on 25 big questions facing science over the next quarter-century.

- > [What Is the Universe Made Of?](#)
- > [What is the Biological Basis of Consciousness?](#)
- > [Why Do Humans Have So Few Genes?](#)
- > [To What Extent Are Genetic Variation and Personal Health Linked?](#)
- > [Can the Laws of Physics Be Unified?](#)
- > [How Much Can Human Life Span Be Extended?](#)
- > [What Controls Organ Regeneration?](#)
- > [How Can a Skin Cell Become a Nerve Cell?](#)
- > [How Does a Single Somatic Cell Become a Whole Plant?](#)
- > [How Does Earth's Interior Work?](#)
- > [Are We Alone in the Universe?](#)
- > [How and Where Did Life on Earth Arise?](#)

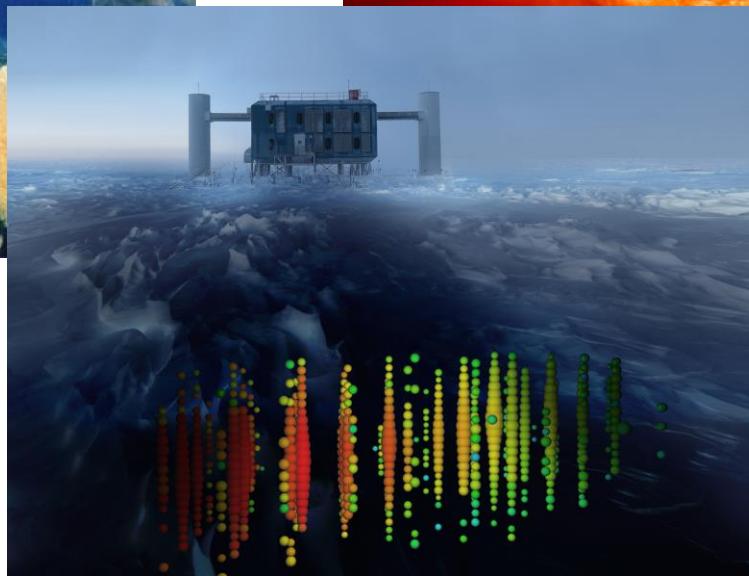
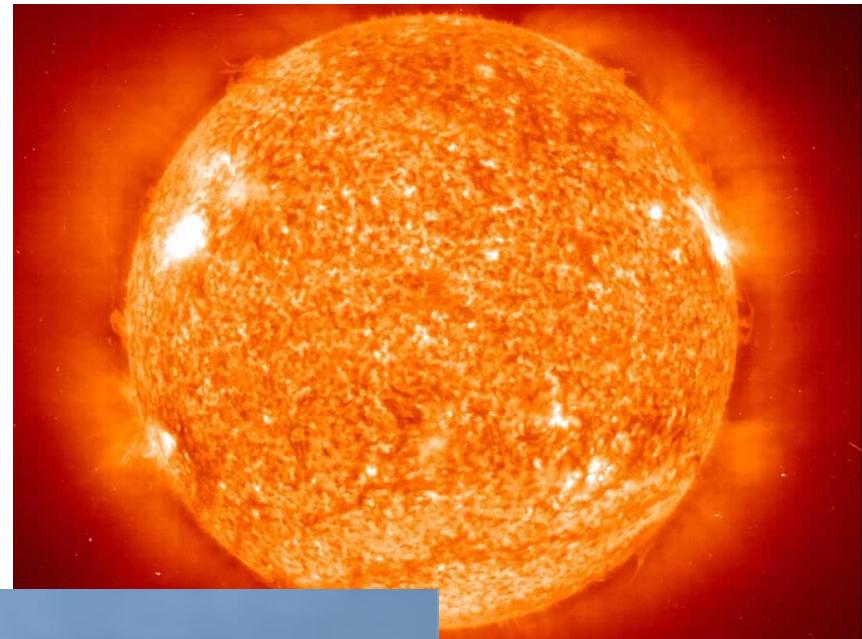
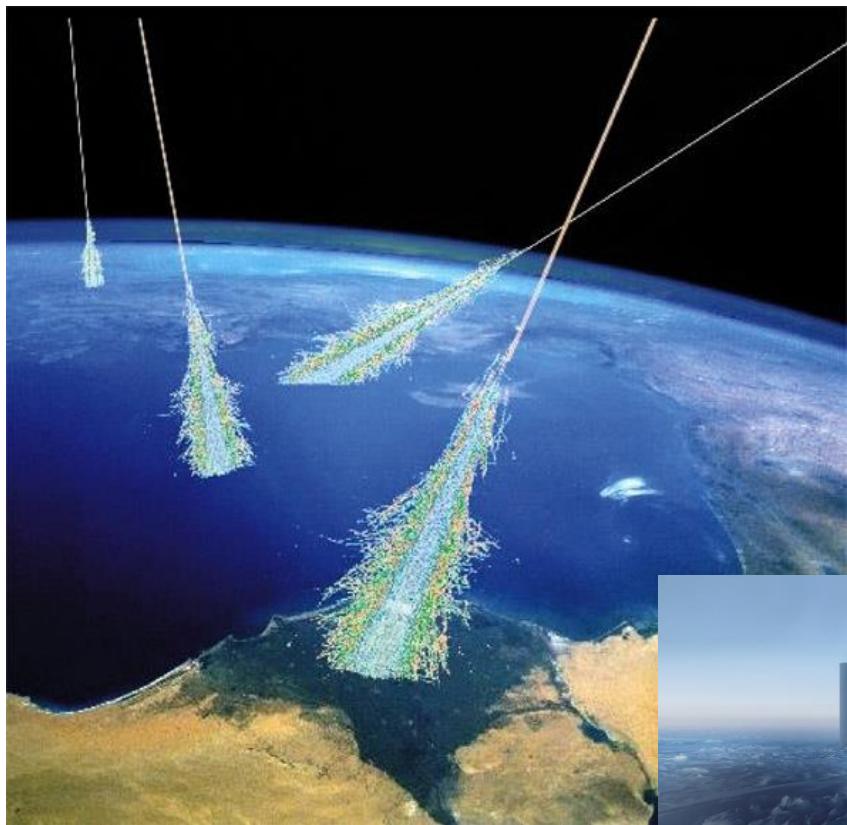
fizika elementarnih čestica: Moderno ime za stara pitanja

Koji su gradivni
elementi svijeta?

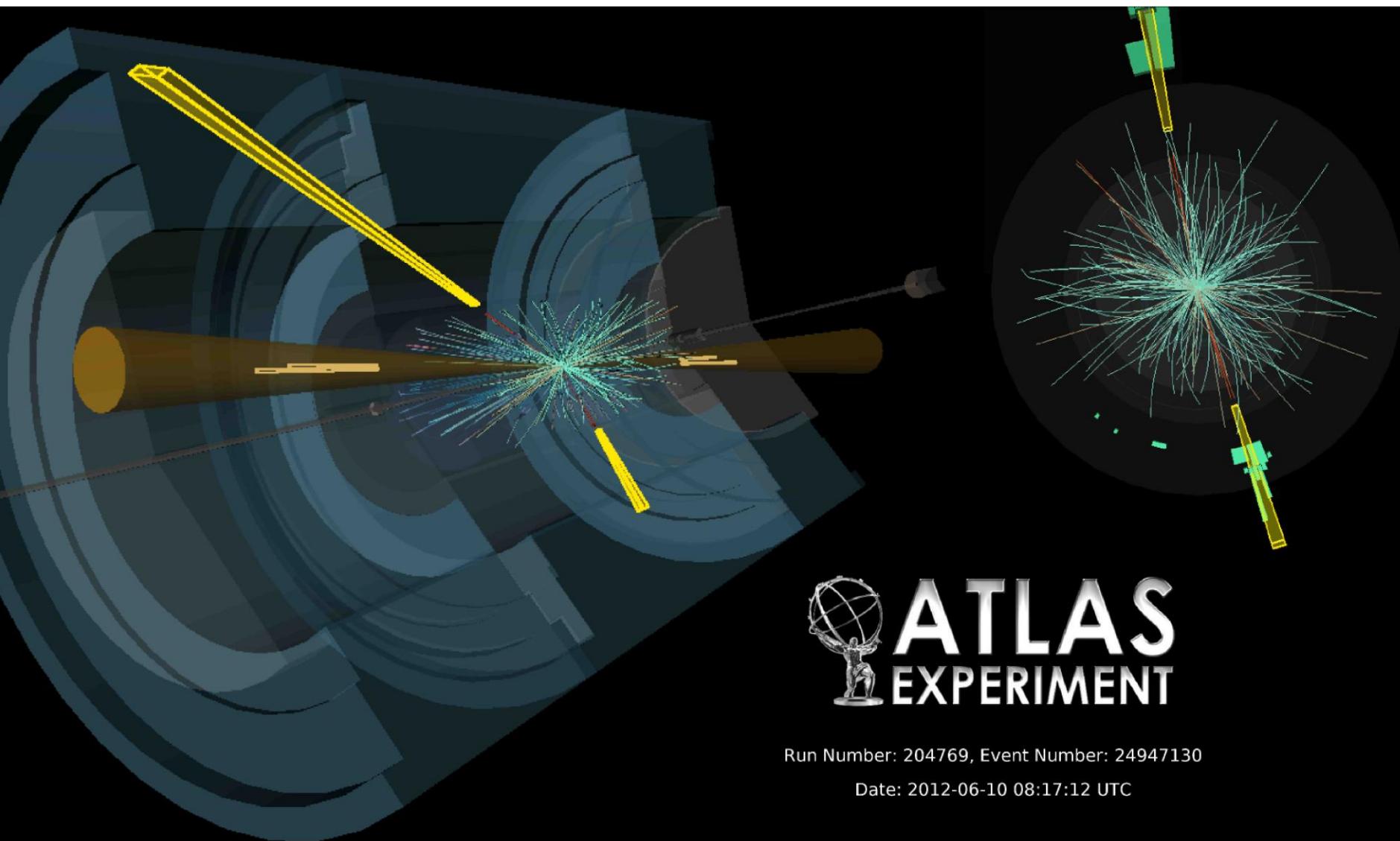
Kako se
ponašaju?



Fenomeni: u prirodi



Fenomeni: u laboratoriju



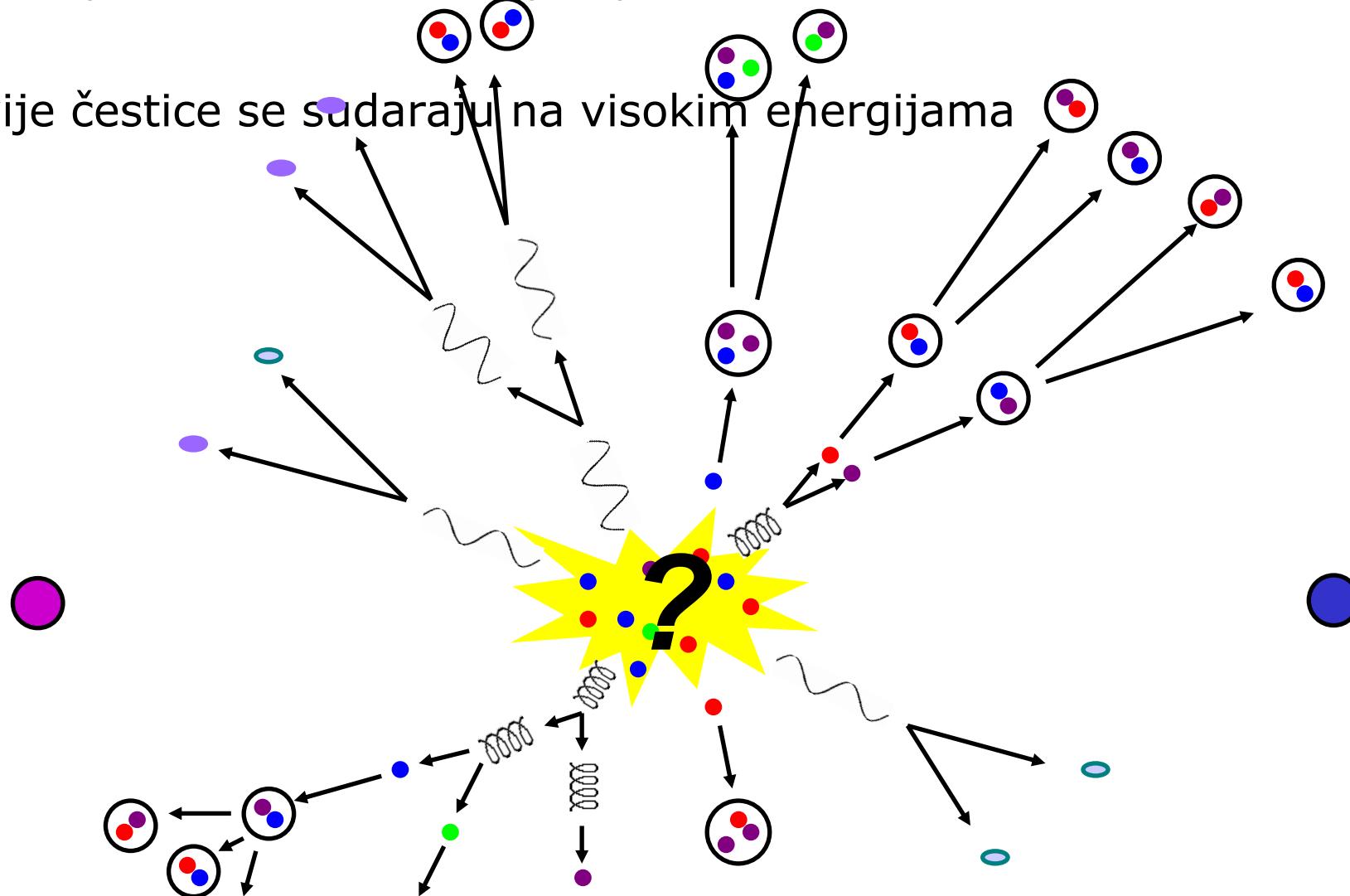
Run Number: 204769, Event Number: 24947130

Date: 2012-06-10 08:17:12 UTC

Fenomeni: sudari čestica

Stvaraju se nove čestice koje mjerimo i proučavamo

Dvije čestice se sudaraju na visokim energijama



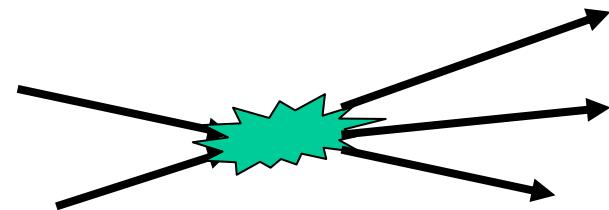
Kako eksperimentalno pristupiti EČ

Sve informacije o elementarnim česticama potječu iz triju izvora:

1. Raspršenja

Čestica na metu ili sudar dviju snopa

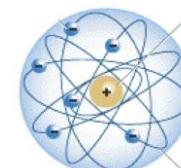
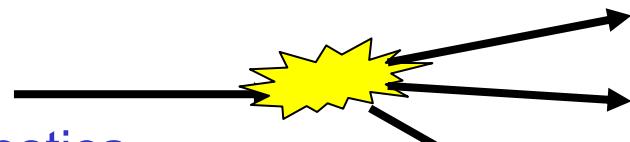
Informacija o interakciji se dobiva iz mjerenja čestica koje proizlaze iz raspršenja (masa, energija, pravac, impuls, spin, ...)



2. Raspadi

Mjerenje ostataka spontanog raspada čestica

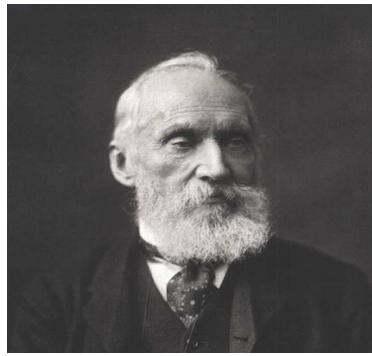
Od svih poznatih čestica: gotovo sve nestabilne!



3. Vezana stanja

Mjerenje svojstava objekata složenih od više čestica (masa, spin, ...)

Fizika na kraju 19. stoljeća



William Thomson
(Lord Kelvin)

A periodic table diagram where the first two columns (H and He) are at the top, followed by a row of noble gases (He, Ne, Ar, Kr, Xe). Below this, the main body of the table is shown with groups of elements color-coded by group number. A separate row below the main table lists the remaining noble gases: La, Ce, Pr, Nd, Pm, Sm, Eu, Gd, Tb, Dy, Ho, Er, Tm, Yb, Ac, Th, Pa, U, Np, Pu, Am, Cm, Bk, Cf, Es, Fm, Md, No. Arrows point from the main table to the separate row of noble gases.

- Cijela fizika opisana kroz:
 - Mehanika (Newton)
 - Termodinamika
 - Elektrodinamika (Maxwell)

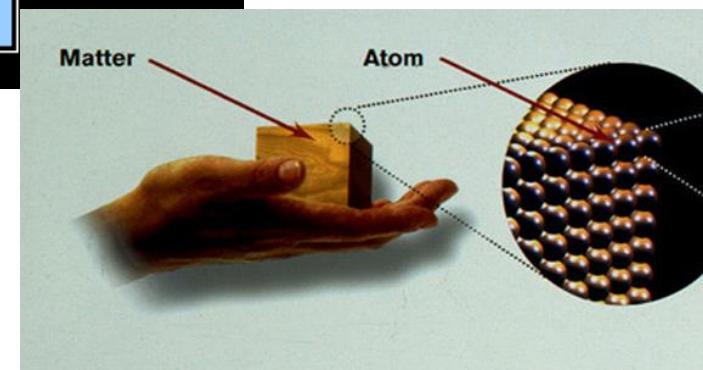
“There is nothing new to be discovered in Physics now. All that remains is more and more precise measurements.”

(Lord Kelvin, 1900)

- Samo nekoliko oblaka na inače vedrom nebnu fiziku:
 - Materija sastavljena od atoma?
 - Spektar zračenja crnog tijela?
 - Michelson-Morley Eksperiment?

Elementarne čestice: do 1897

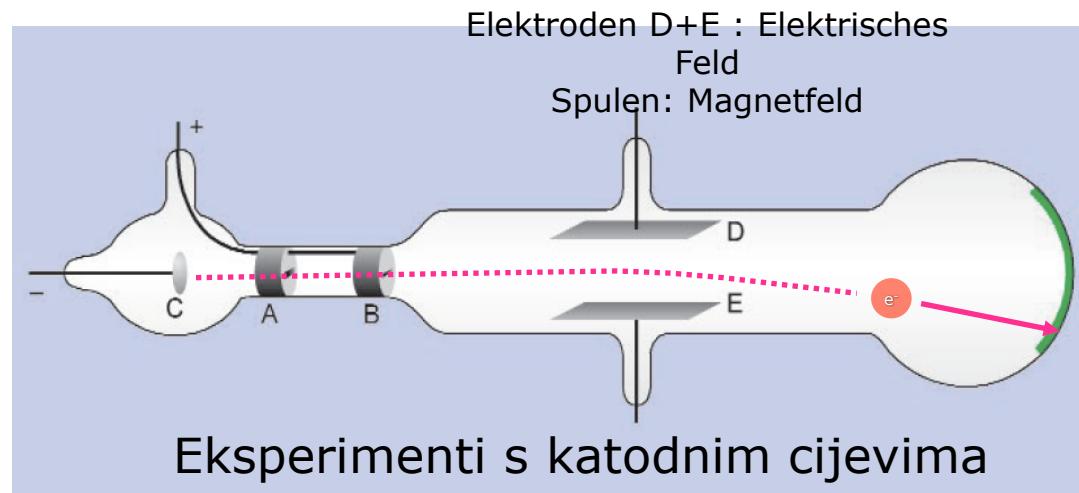
	IA												0					
1	H	IIA											He					
2	Li	Be																
3	Na	Mg	IIIIB	IVB	VB	VIB	VIIB	VIIIB		IB	IIB	III A	IVA	VA	VIA	VIIA		
4	K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Al	Si	P	S		
5	Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe
6	Cs	Ba	*La	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn
7	Fr	Ra	89	104	105	106	107	108	109	110	111	112		114	116	116	118	118
+Ac		104	105	106	107	108	109	109	110	111	112		114	116	116	118	118	
*Lanthanide Series		58	59	60	61	62	63	64	65	66	67	68	69	70	71	Lu		
+Actinide Series		Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb				
+Actinide Series		90	91	92	93	94	95	96	97	98	99	100	101	102	103	Lr		
+Actinide Series		Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No				



1897, prva elementarna čestica: Elektron



J.J. Thomson



1893-1897 Otkriće elektrona (J.J. Thomson)

(Struje u katodnim cijevima se sastoje od čestica)

$$m_e \approx M_H / 1836$$

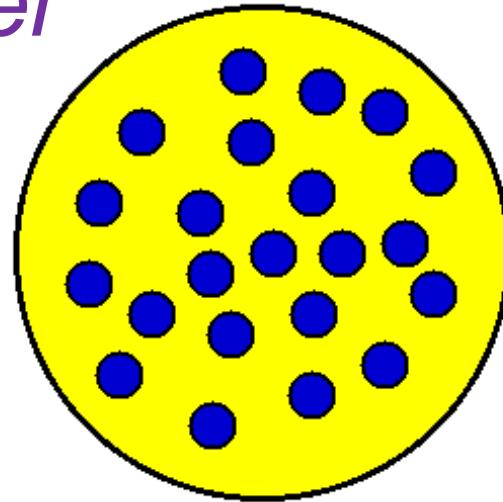
“Could anything at first sight seem more impractical than a body which is so small that its mass is an insignificant fraction of the mass of an atom of hydrogen?” (J.J. Thomson)



ATOMI NISU ELEMENTARNI!

Thomsonov model atoma

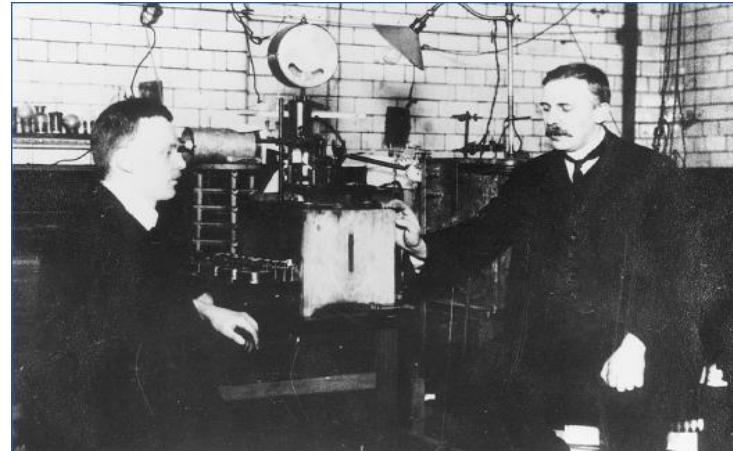
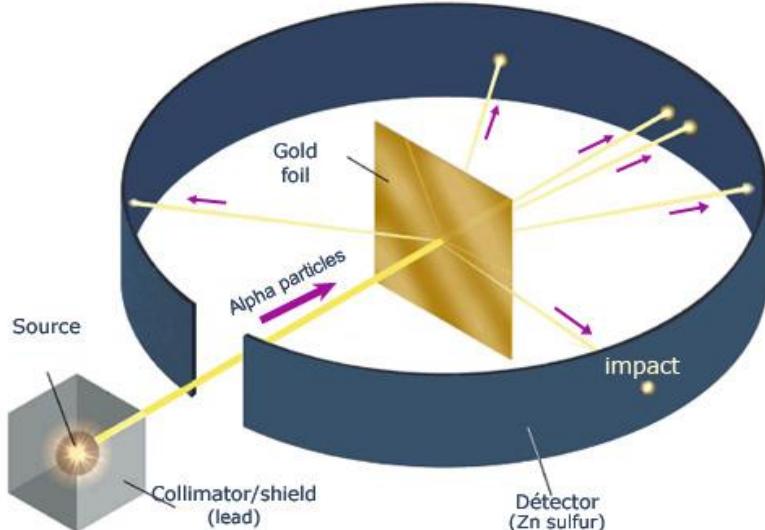
"Plum pudding model"



Thomsonov model atoma:

- Električni nabijena kugla
- Radijus $\sim 10^{-8}$ cm
- Pozitivni naboј: homogeno raspodijeljen u cijelom atomu
- Elektroni s negativnim naboјem raspršeni u kugli

Otkriće atomske jezgre: Rutherfordov eksperiment

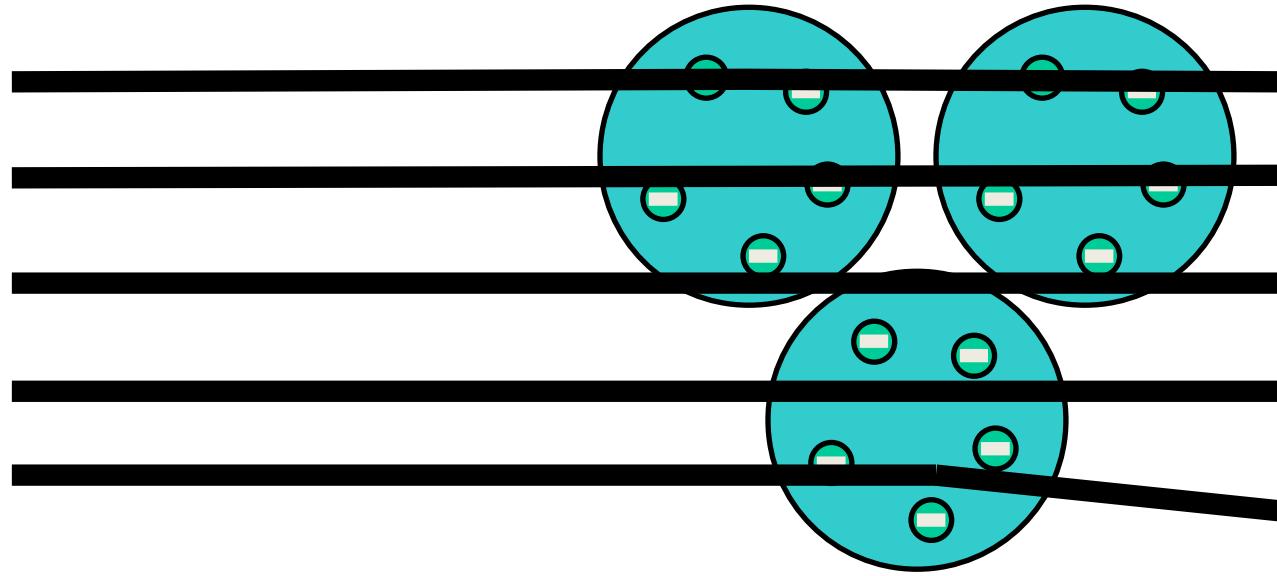


Ernest Rutherford (r) und Hans Geiger (l)
in Manchester

Bitni elementni eksperimenta

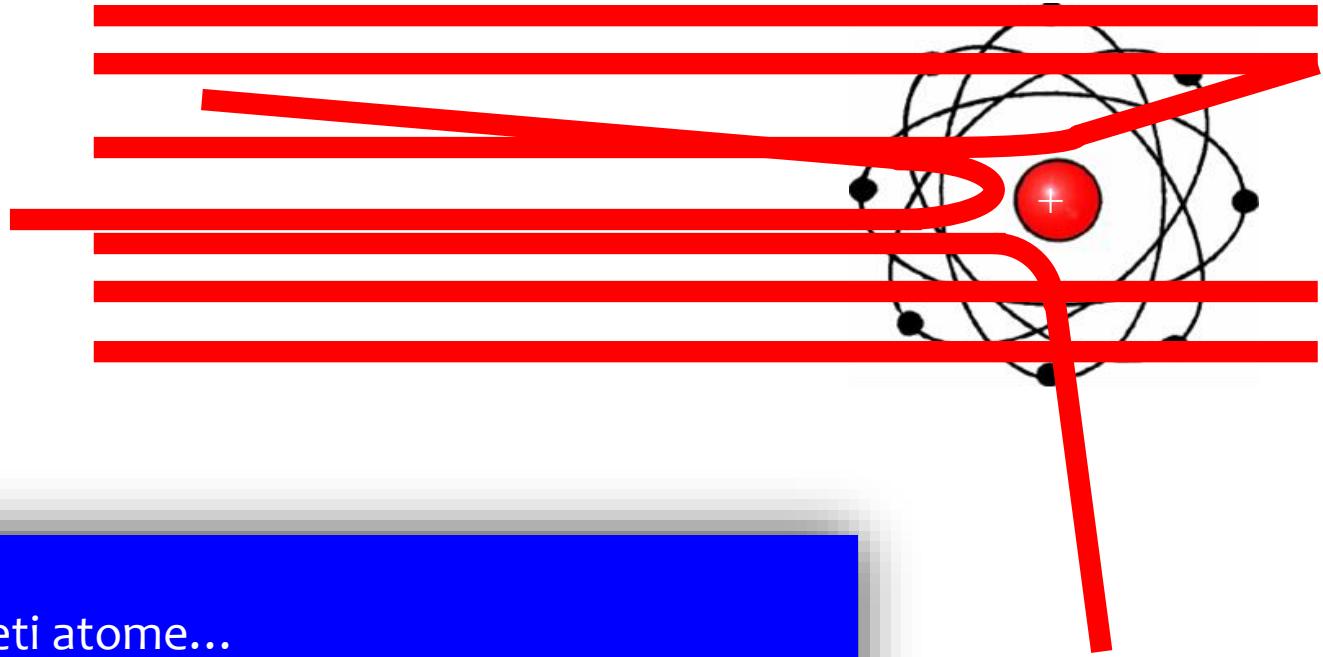
- Snop
- Meta
- Detekcija
- Izbor podataka (u našem žargonu: *triggering*)

Thomsonov model atoma



Elektroni izmiješani s ostatkom atoma.
Atome ne možemo vidjeti... ali možemo tražiti dokaze da su tu
Kakve eksperimentalne rezultate predviđa model?

Rutherfordov model atoma



Rutherford model:

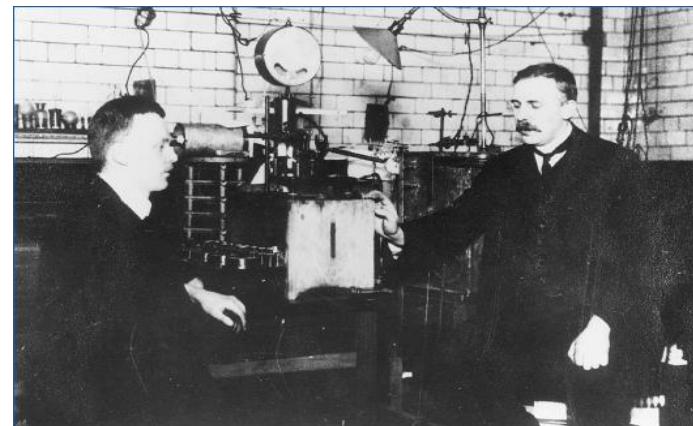
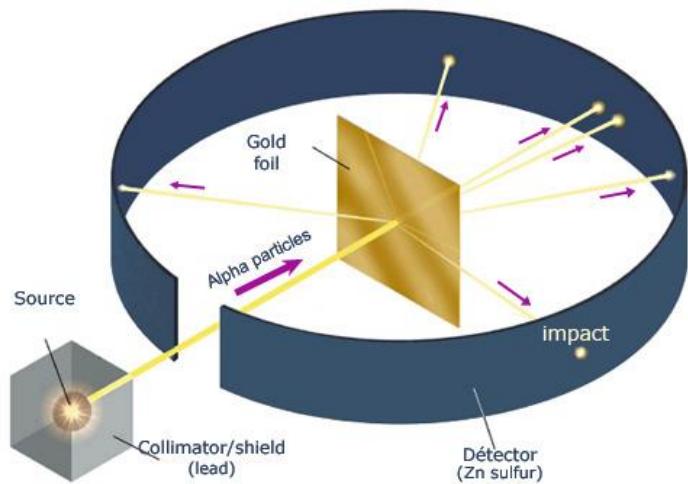
I dalje ne možemo vidjeti atome...

Negativni elektroni oko malog, gustog pozitivnog centra

Kakve eksperimentalne rezultate predviđa ovaj model?

Jezgra atoma je 100 000 puta manja od promjera atom!

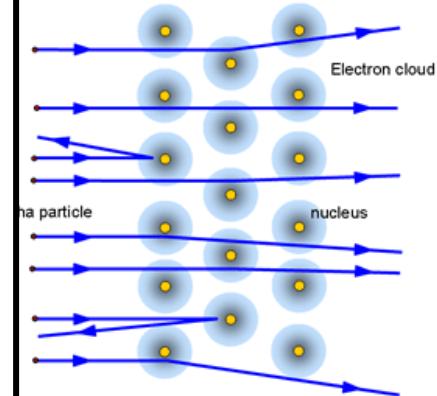
Otkriće atomske jezgre



Ernest Rutherford (r) und Hans Geiger
(l)
in Manchester

It was quite the most incredible event that has ever happened to me in my life. It was almost as incredible as if you fired a 15-inch shell at a piece of tissue paper and it came back and hit you. On consideration, I realized that this scattering backward must be the result of a single collision, and when I made calculations I saw that it was impossible to get anything of that order of magnitude unless you took a system in which the greater part of the mass of the atom was concentrated in a minute nucleus. It was then that I had the idea of an atom with a minute massive centre, carrying a charge.

— Ernest Rutherford^[19]



Gotovo cijela masa atoma koncentrirana u jako maloj jezgri : $< 27 \times 10^{-15} \text{ m}$

2 revolucije u ranom 20. stoljeću

RELATIVNOST

Fizika visokih energija
brzine $\sim c$



Einstein
(1905)

Masa / Energija
 $E = M c^2$

Stvaranje
Novih čestica

KVANTNA FIZIKA

Fizika najmanjih dimenzija



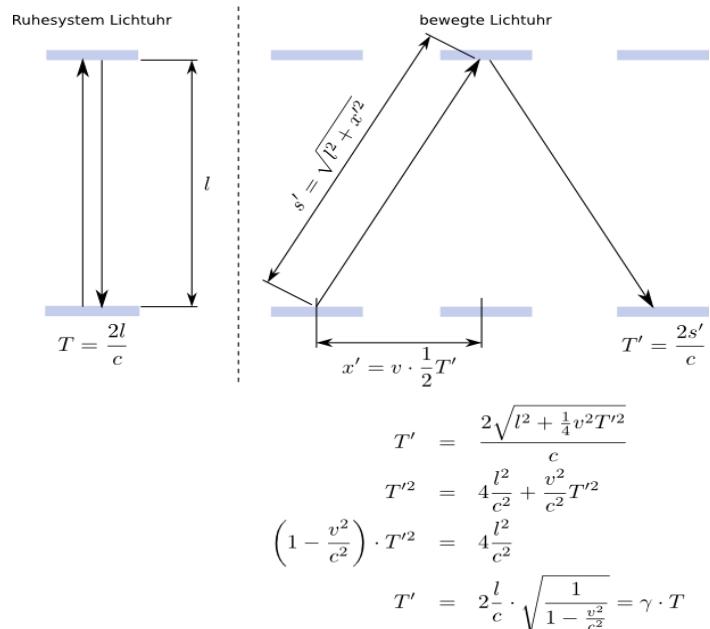
Val / Čestica
 $p = h / \lambda$

Ispitivanje strukture
materije

Bohr
(1913)

Relativnost: principi

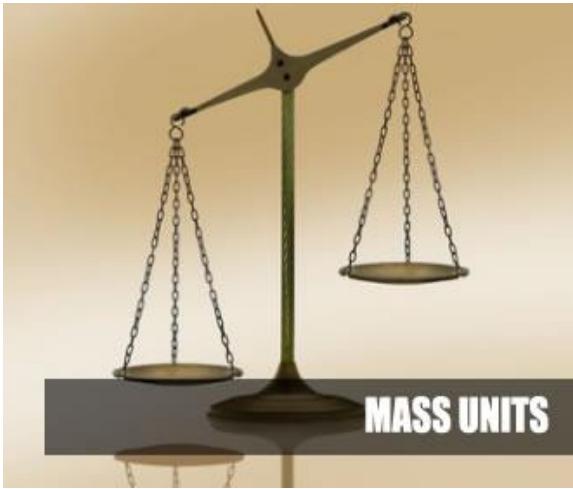
1. Zakoni fizike su isti u svim inercijalnim sustavima
2. Brzina svjetlosti (u vakuumu) je konstantna i ista za sve promatrače



- Nema absolutnog vremena ni prostora
- Dilatacija vremena - duljine

Masa i Energija

- Masa:
 - Spremište energije čestica
 - Ne ovisi o brzini
 - “definirana”
$$M^2 c^4 = E^2 - p^2 c^2$$
 - U miru: $E = M c^2$
 - M je **Lorentz invarianta**: ista u svim sustavima (E i p opčenito nisu)
 - c velik: mala masa= velika količina energije

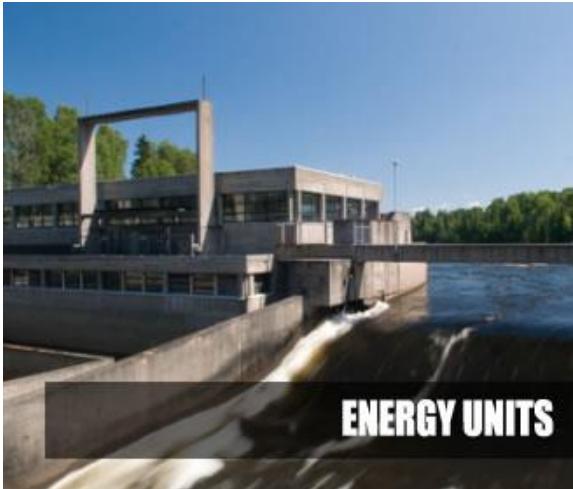


Jedinica za masu (SI): kg

masa elektrona $\approx 10^{-30}$ kg

masa čovjeka ≈ 100 kg

masa Sunca $\approx 10^{30}$ kg



Jedinica za energiju (SI): J

Često se koristi u fizici:

eV (elektron volt)

$1 \text{ eV} = 1.6 \times 10^{-19} \text{ J}$

$1 \text{ GeV} = 10^9 \text{ eV}$

$E = m$

$E = mc^2$

$$E = mc^2$$



$$c = 300\,000 \text{ km/h} = 10^8 \text{ m/s}$$

$$c^2 = 9 \times 10^{16} \text{ m}^2/\text{s}^2 \approx 10^{17} \text{ m}^2/\text{s}^2$$

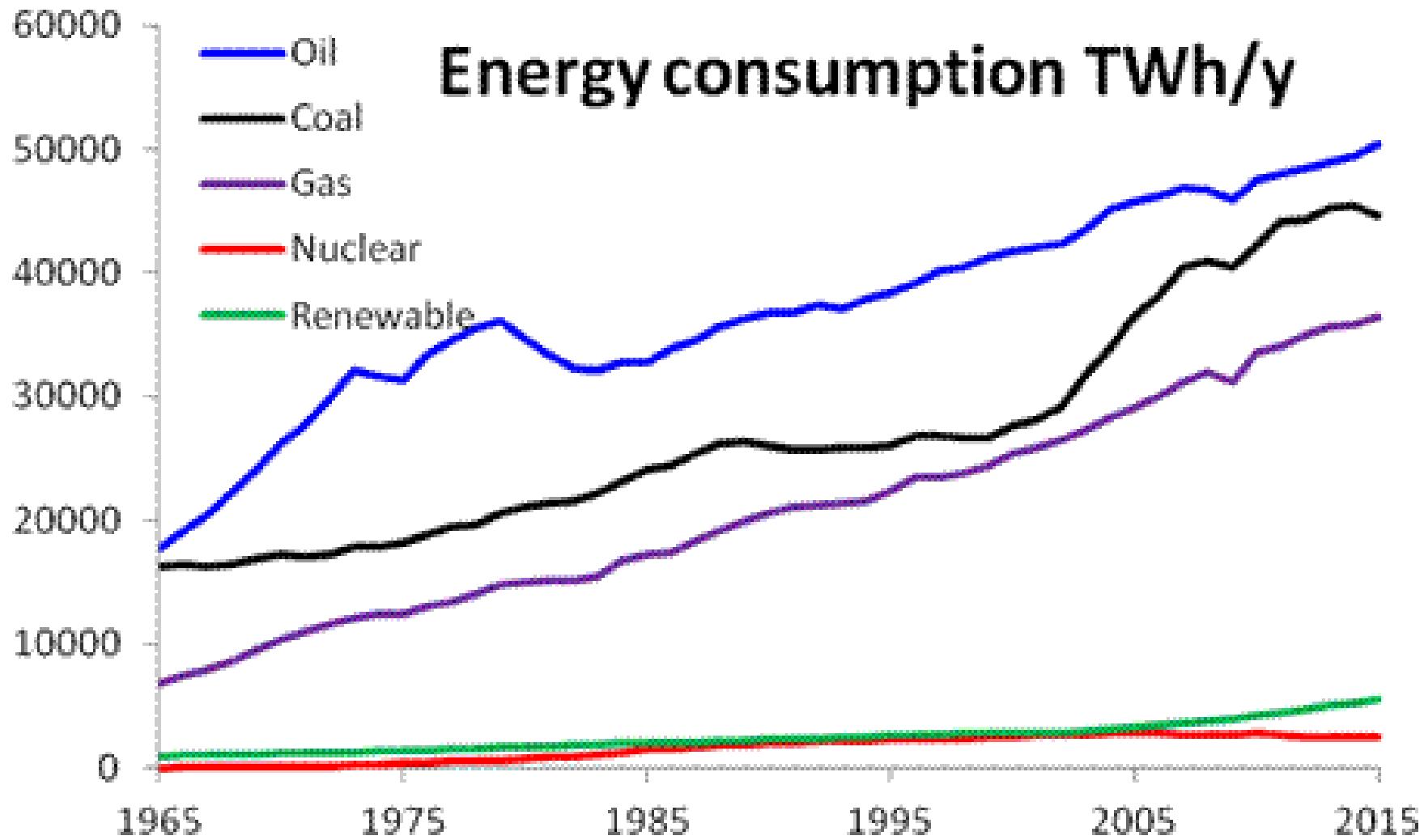


$$m = 100 \text{ kg}$$

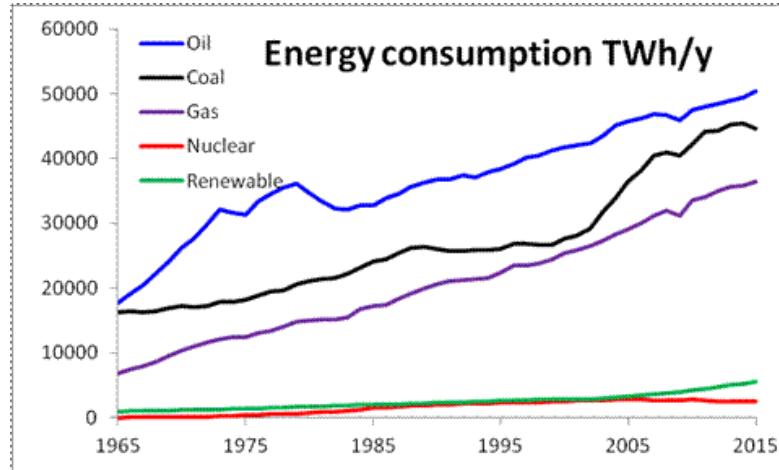


$$E = mc^2 \approx 10^{19} \text{ J}$$

Energy consumption TWh/y



$$2015 : 109136 \text{ TWh} = 3.93 \cdot 10^{20} \text{ J} = \sim 4400 \text{ kg} \cdot c^2$$



=



~2 *



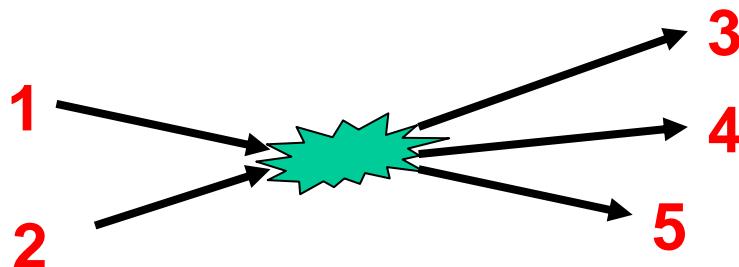
U fizici visokih energija

$$E = mc^2 \rightarrow [m] = \text{GeV}/c^2$$

$$c = 1 \rightarrow E = m \rightarrow [m] = \text{GeV}$$

Očuvane veličine

- Očuvane veličine u svakoj reakciji:
 - Energija, količina gibanja
- Npr. Reakcija $1+2 \rightarrow 3 + 4 + 5$



– Vrijedi

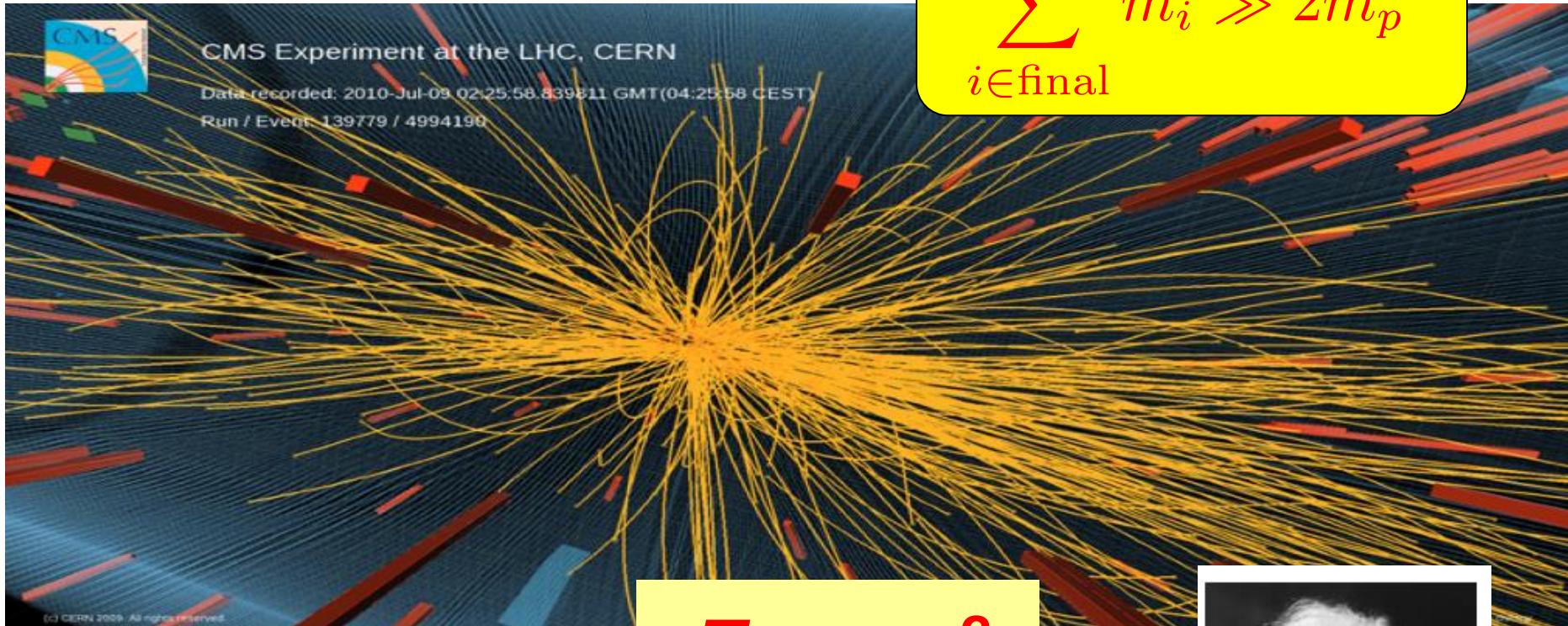
$$E_1 + E_2 = E_3 + E_4 + E_5$$

$$\vec{p}_1 + \vec{p}_2 = \vec{p}_3 + \vec{p}_4 + \vec{p}_5$$

– Ali općenito: $M_1 + M_2 \neq M_3 + M_4 + M_5$
Masa nije očuvana veličina!

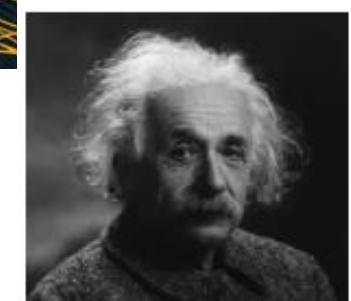
Masa nije očuvana

proton proton sudar na LHC-u



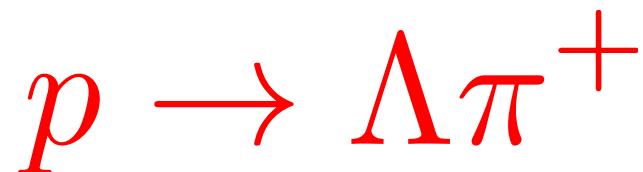
$$\sum_{i \in \text{final}} m_i \gg 2m_p$$

$$E = mc^2$$



Relativnost i raspadi: Test

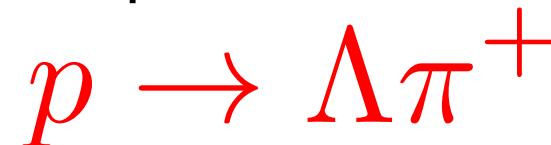
- Mase:
 - Proton: 938 MeV
 - Λ barion: 1116 MeV
 - Pion \square^+ : 140 MeV
- Ako proton miruje, je li dopušten raspad



- a) DA
- b) NE

Relativnost i raspadi: Test

- Zbog očuvanja energije se proton u miru ne može raspasti na način



- $M(p)$ 938 MeV, $M(\Lambda)$: 1116 MeV, $M(\pi^+)$: 140 MeV
- Na koju kinetičku energiju moram ubrzati proton za omogućiti taj raspad:
 - 318 MeV
 - 1256 MeV
 - 2512 MeV
 - Nikad neće biti moguće

Kako identificirati nevidljive čestice

Nije mjerena

$$Z \quad E_Z, \vec{p}_Z$$

$$e^+$$

$$e^-$$

$$E_1, \vec{p}_1$$

Mjerimo

$$E_2, \vec{p}_2$$

Očuvanje

$$E_Z = E_1 + E_2$$

Energije i impusa

$$\vec{p}_Z = \vec{p}_1 + \vec{p}_2$$

Invarijantna masa

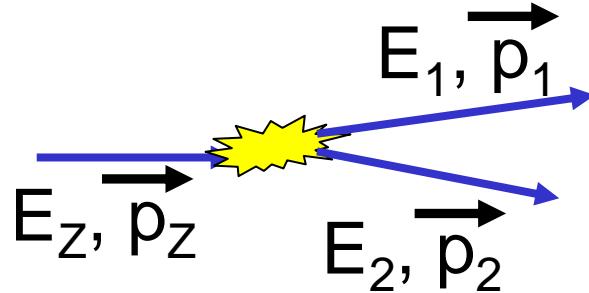
$$M_Z^2 = E_Z^2 - p_Z^2 = (E_1 + E_2)^2 - (\vec{p}_1 + \vec{p}_2)^2$$

Vježba: dodajte nedostajuće c faktore
(dimenzionalna analiza!)

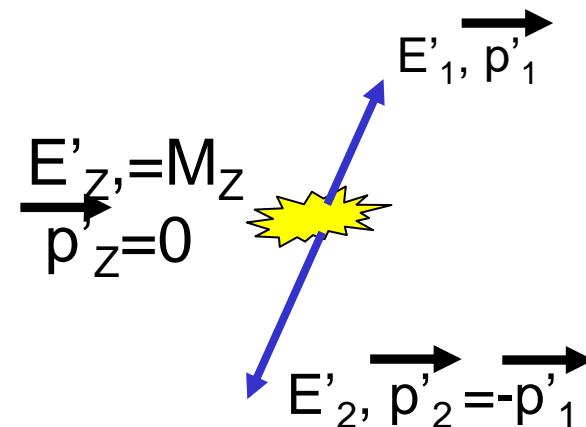
Izbor sustava

Možemo promatrati raspad u više mogućih sustava

Raspad u letu



Raspad u mirovanju



Opčenito: E_i i E'_i nisu jednaki

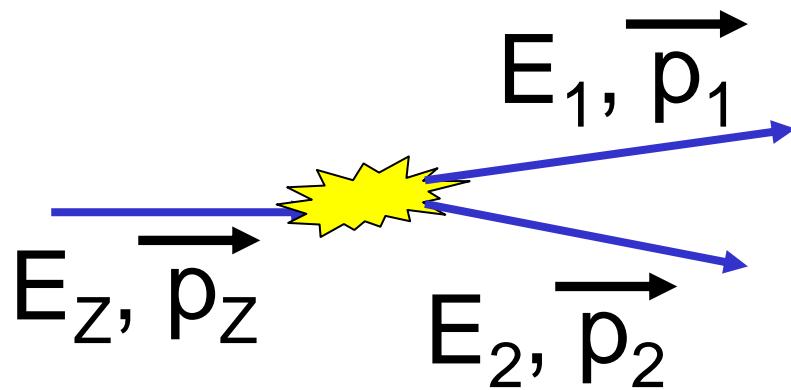
U kojem sustavu računati inv. Masu za dobiti pravu masu čestice?

- 1) U sustavu laboratorija**
- 2) U sustava mirovanja**
- 3) U nekom trećem sustavu**
- 4) U bilo kojem od navedenih**

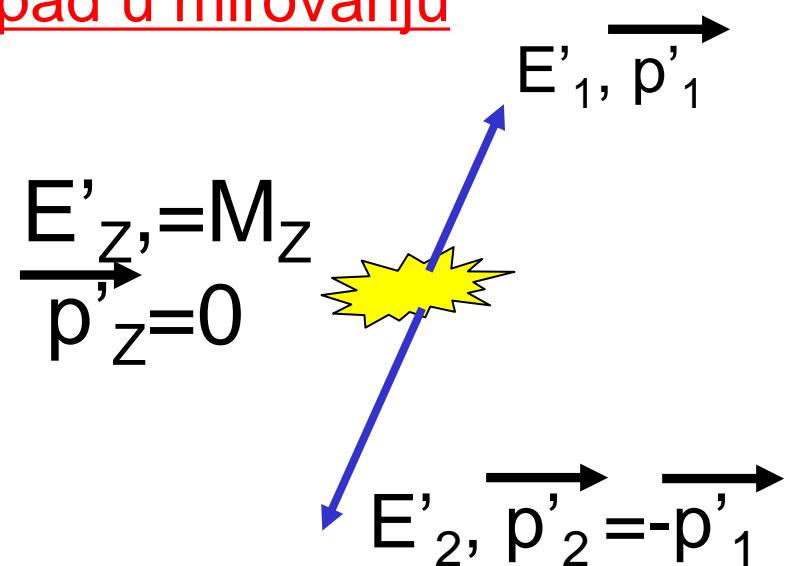
Izbor sustava

Možemo promatrati raspad u više mogućih sustava

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Raspad u mirovanju



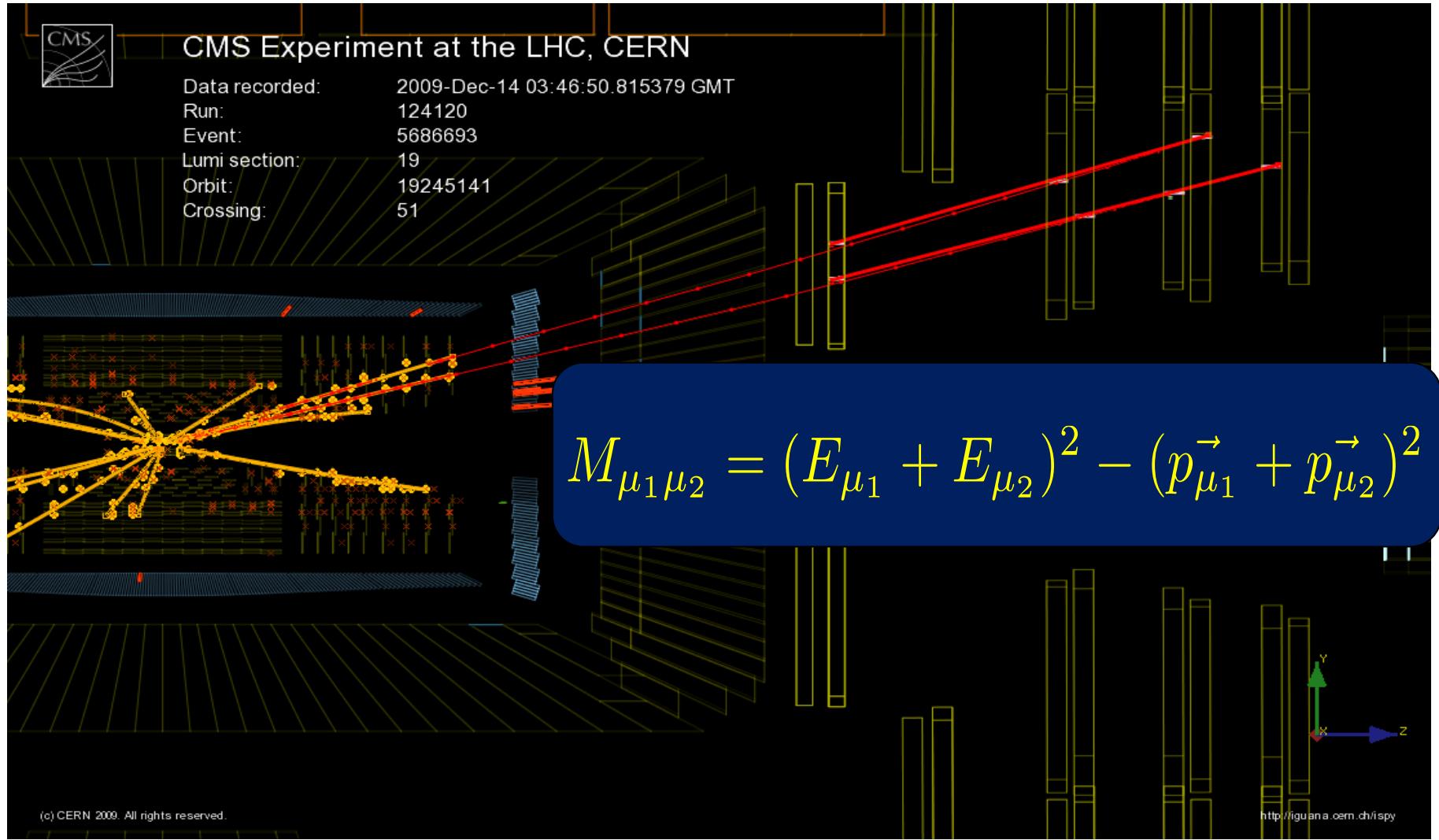
Općenito: E_i i E'_i nisu jednaki

U kojem sustavu računati inv. Masu?

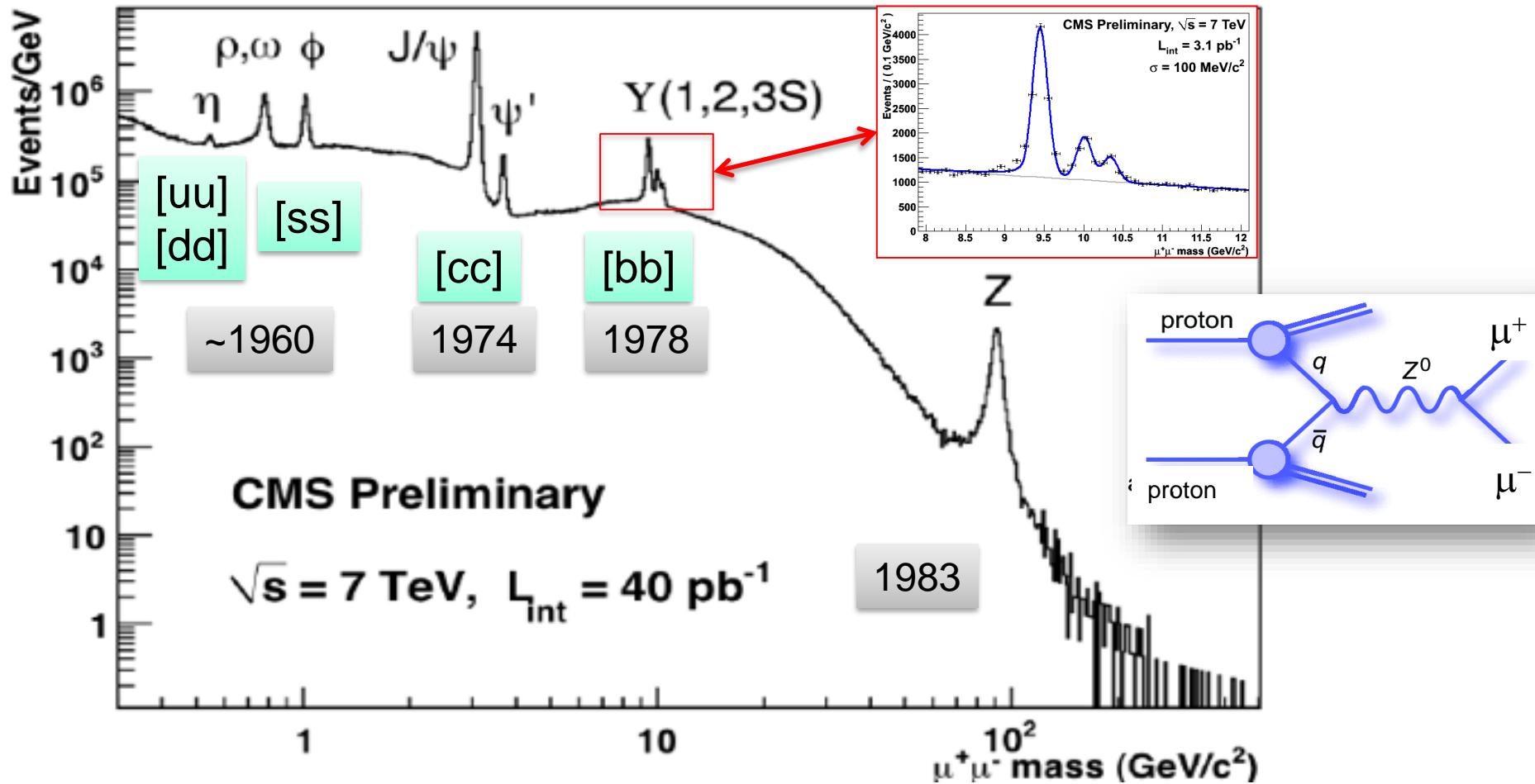
$$\text{Einstein : } (E_1 + E_2)^2 - (\vec{p}_1 + \vec{p}_2)^2 = (E'_1 + E'_2)^2 - (\vec{p}'_1 + \vec{p}'_2)^2$$

NIJE BITNO: INVARIJANTNA MASA NE OVISI O PROMATRAČU

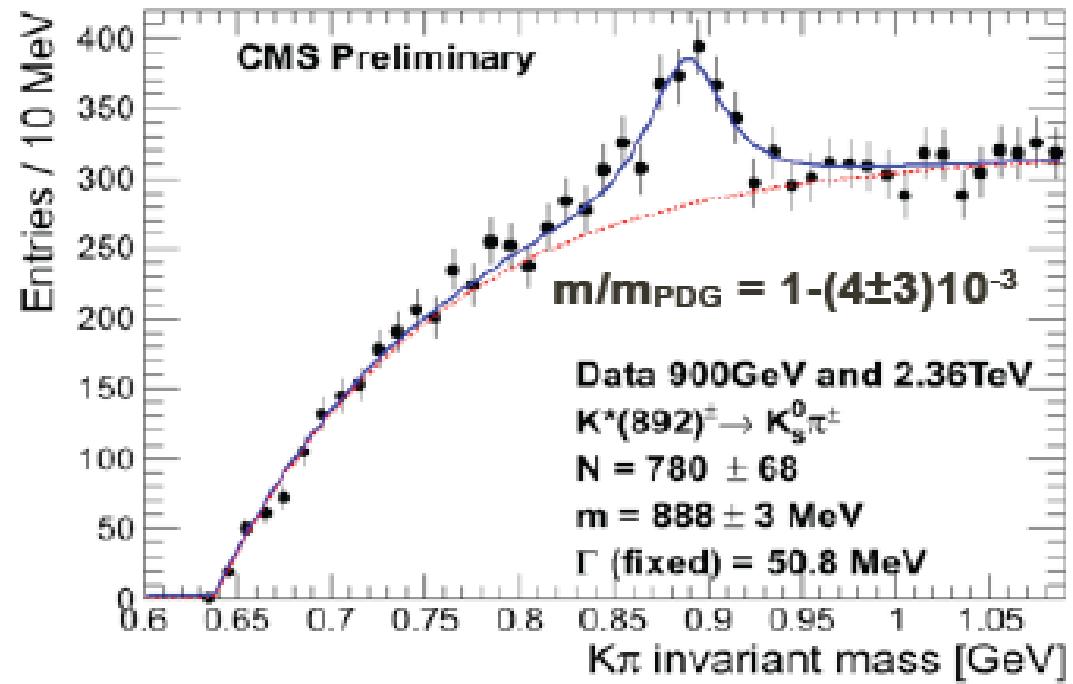
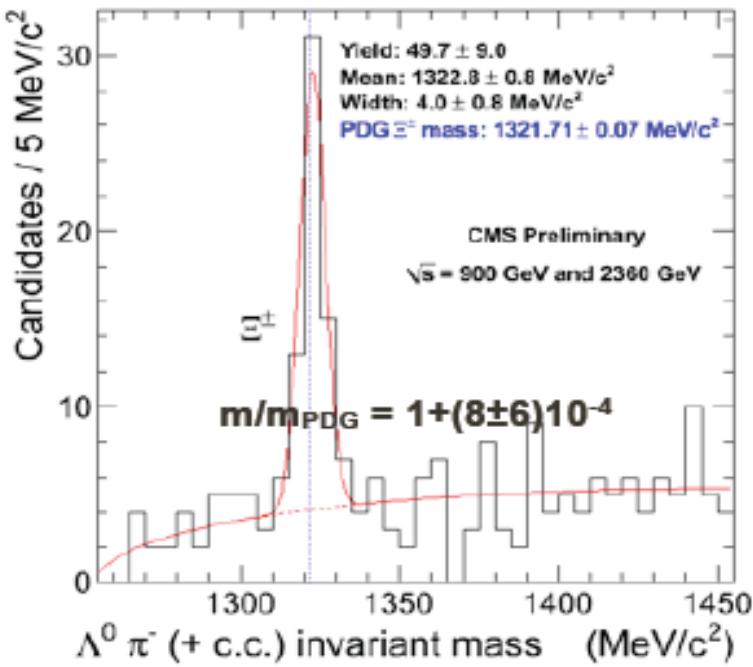
Događaji sa 2 miona



Ponovno otkriće SM na LHC-u



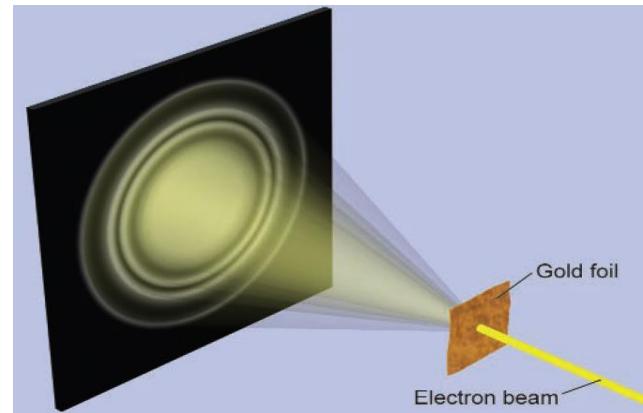
Signal i pozadina – Primjeri iz prvih LHC podataka



Elementarne čestice: kvantni objekti

- Svim česticama pripadaju svojstva čestica:
 - Energija, količina gibanja, masa
 - Za svjetlo (fotone)
E=hν
- Sve imaju i valnu narav:

$$\lambda = h / p$$



Louis de Broglie (1924)

Kvantna neodređenost



Heisenberg (1925)

- Količina gibanja i pozicija ne mogu biti istovremeno precizno mjerene:

$$\Delta x \Delta p \geq \frac{\hbar}{2}$$

- Isto vrijedi za energiju i vrijeme:

$$\Delta E \Delta t \geq \frac{\hbar}{2}$$

Kvantna jednadžba gibanja



Schrödinger
1926

- Schroedinger: ponašanje čestica opisano kroz valnu jednadžbu

$$i\hbar \frac{\partial}{\partial t} \psi = -\frac{\hbar^2 \nabla^2}{2m} \psi + V(r)\psi$$

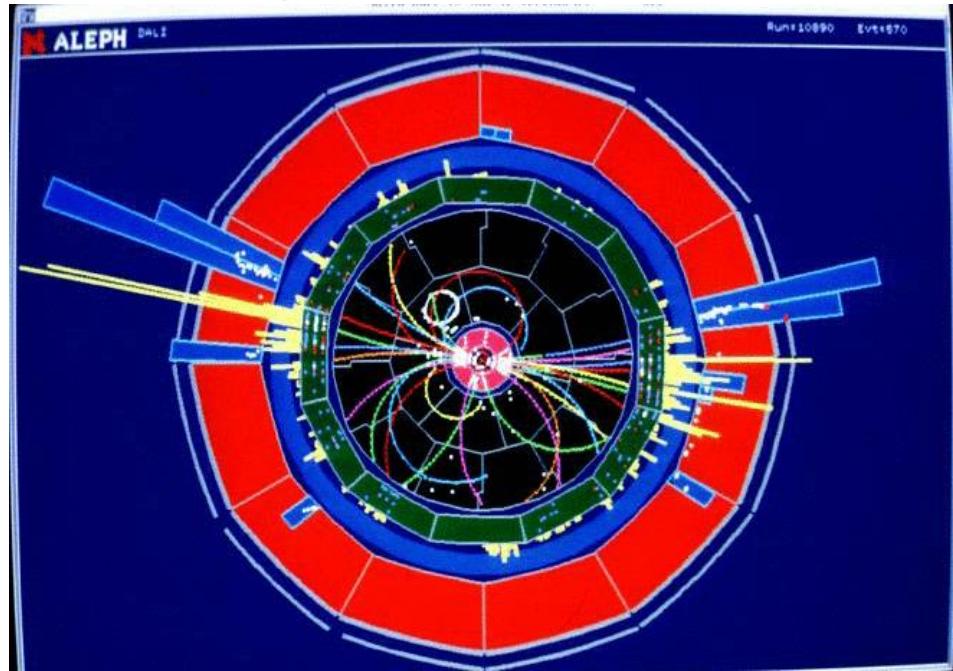
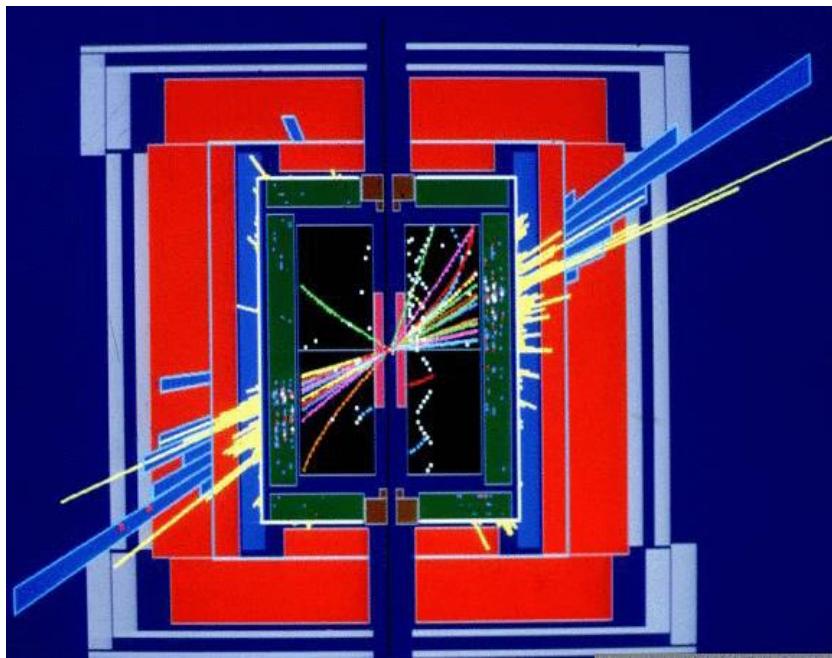
- Novost: valna funkcija nije više direktno mjerljiva nego predstavlja vjerojatnost!
- Nigdje tako dobro vidljivo kao kod elementarnih čestica

Primjer slučajnosti: Z raspad

Eksperiment

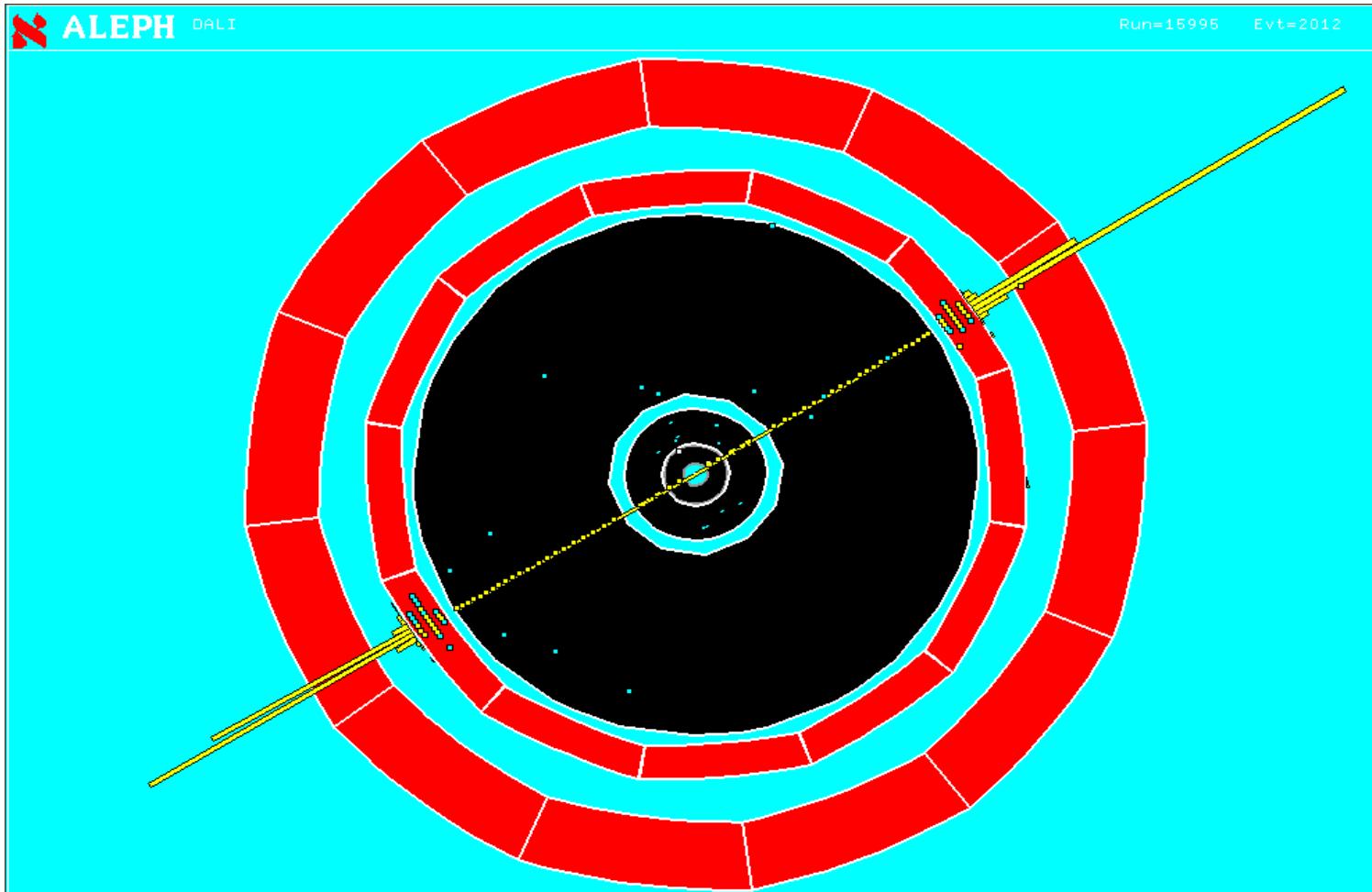
Sudaramo snopove elektrona i pozitrona na energiji od 91 GeV

Produciramo Z bozone i proučavamo njihove raspade...



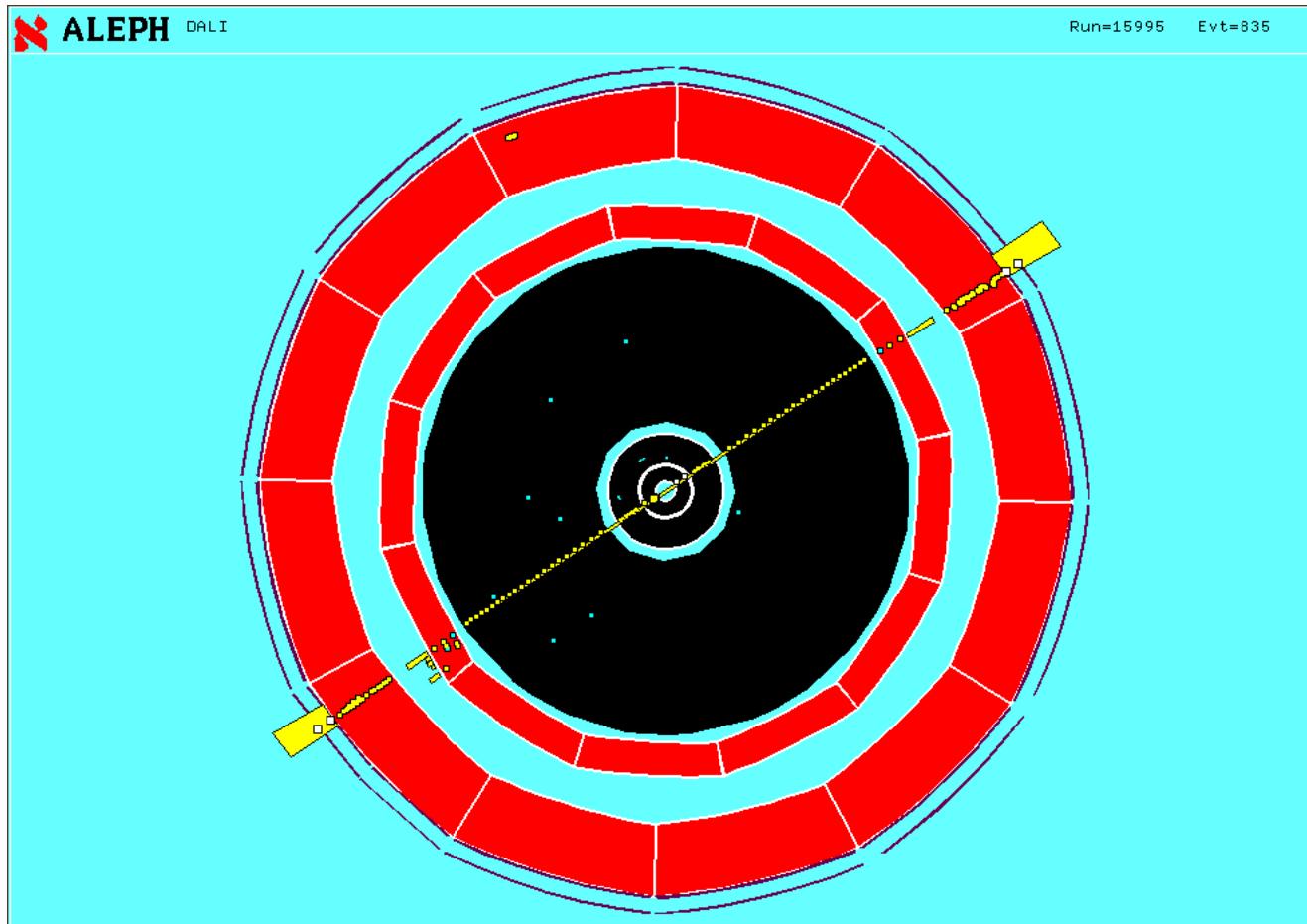
Ajmo to ponoviti...

Još jedan raspad Z bozona



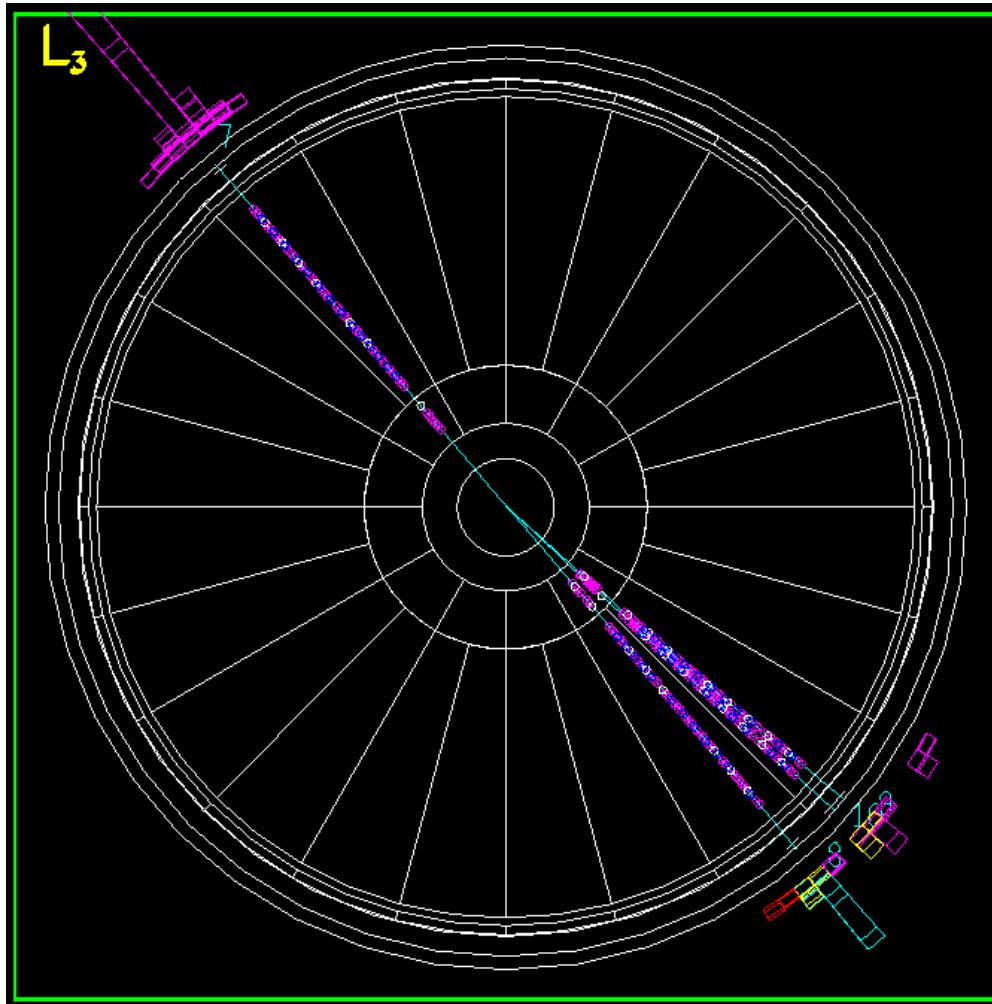
Prilično različito!

I još jednom



Stvarno se radi o ISTOM eksperimentu...

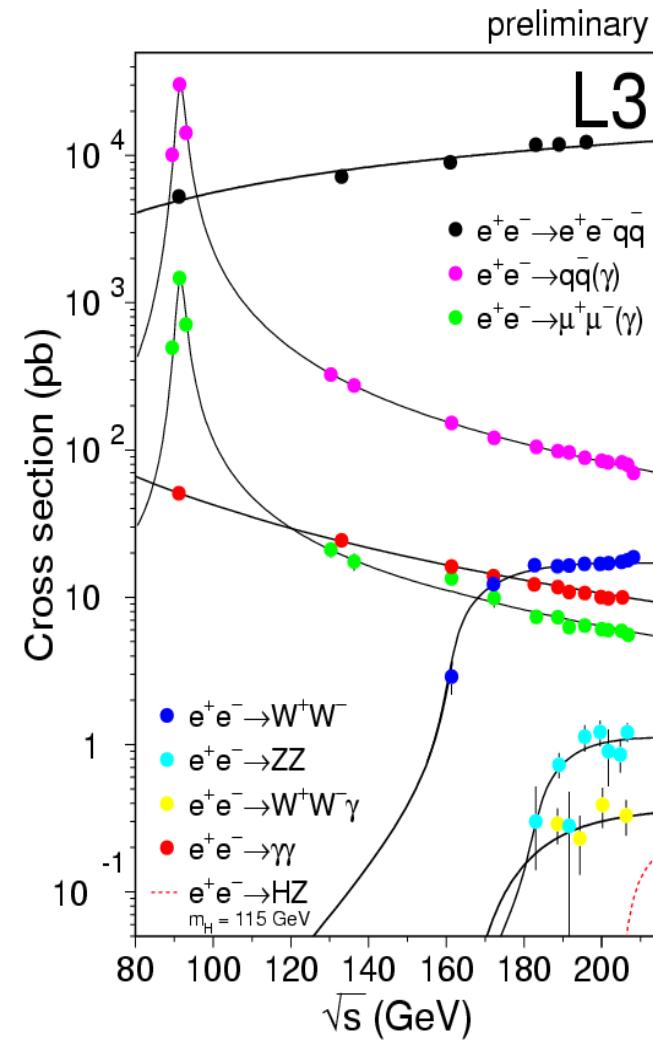
I još jedan

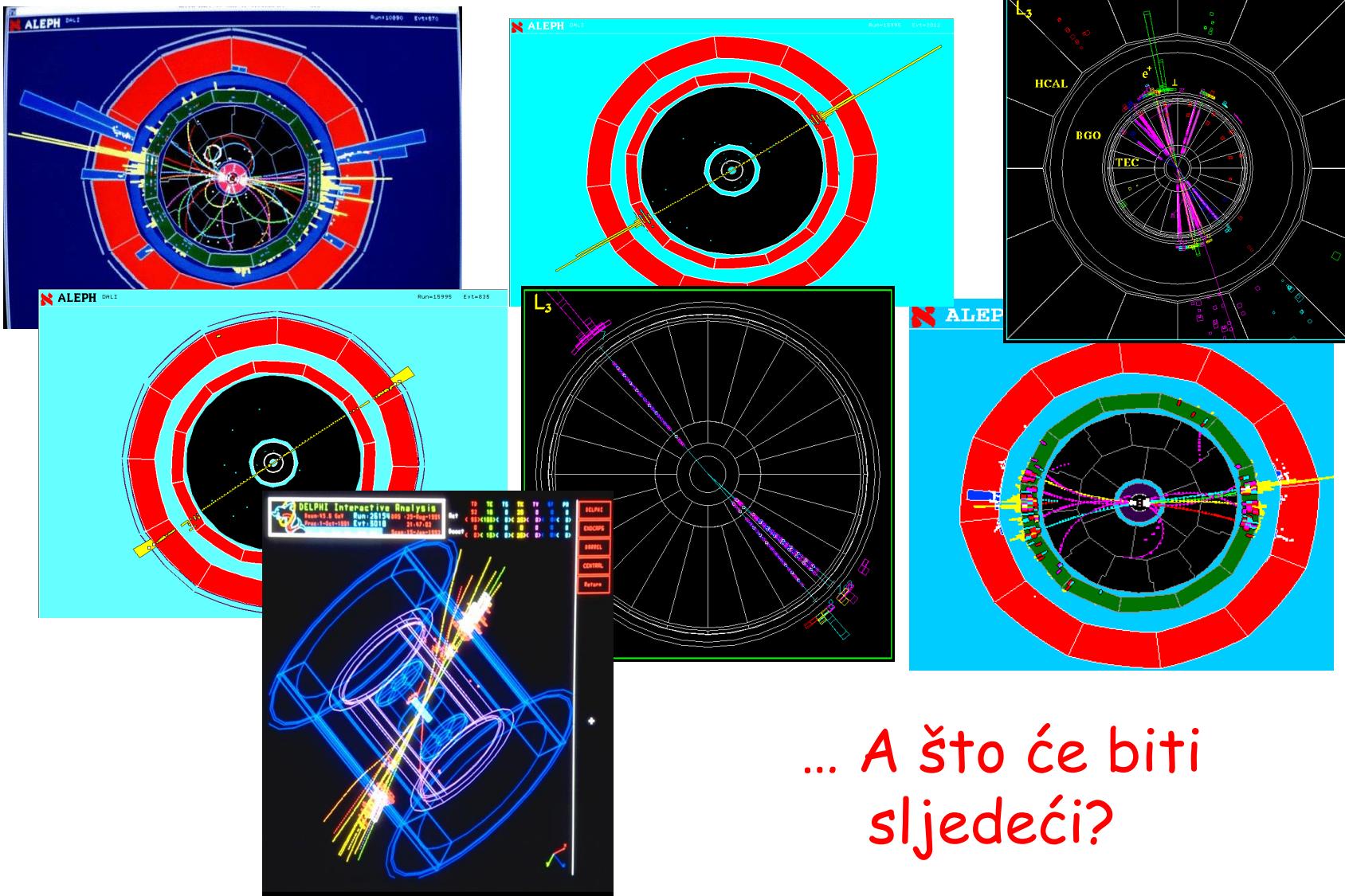


Što nam teorija kaže o Z raspadima

- Predviđa koje su vrste raspada moguće
- Predviđa koliko često (%) će se raspasti u svaki I jako točno!

Ali ne kaže
nam apsolutno
ništa o ...



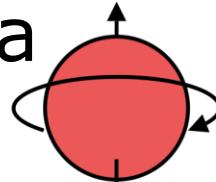


... A što će biti
sljedeći?

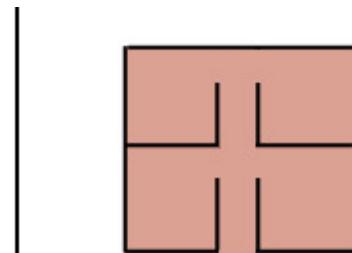
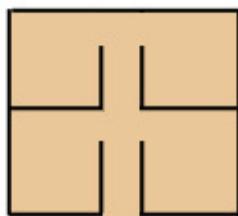
**Ako kvantna teorija vrijedi to nije određeno fizikalnim
zakonima, čisto je **slučajno!****

Kvantno svojstvo: Spin

- Čisto kvantna pojava: intrinzična kutna količina gibanja, SPIN
- Kutna količina gibanja je očuvana veličina
- Kutna količina gibanja može imati samo vrijednosti, u jedinici Planckove konstante:
 - $0, 1, 2, 3, \dots$: Bozoni
 - $\frac{1}{2}, \frac{3}{2}, \frac{5}{2}, \dots$: Fermioni
- Važno svojstvo kvantne teorije:
 - 2 fermiona NE MOGU biti u istom stanju
 - 2 ili više bozona mogu biti u istom stanju



Bozoni i Fermioni



Standard Model of Elementary Particles

three generations of matter (fermions)				
QUARKS	I	II	III	
	mass charge spin u	~2.4 MeV/c ² 2/3 1/2	~1.275 GeV/c ² 2/3 1/2	~172.44 GeV/c ² 2/3 1/2
	down d	-4.8 MeV/c ² -1/3 1/2	~95 MeV/c ² -1/3 1/2	~4.18 GeV/c ² -1/3 1/2
	electron e	~0.511 MeV/c ² -1 1/2	~105.67 MeV/c ² -1 1/2	~1.7768 GeV/c ² -1 1/2
LEPTONS	tau τ	~1.7768 GeV/c ² -1 1/2	muon μ	~0.511 MeV/c ² -1 1/2
	electron neutrino ν _e	<2.2 eV/c ² 0 1/2	muon neutrino ν _μ	<1.7 MeV/c ² 0 1/2
	tau neutrino ν _τ	<15.5 MeV/c ² 0 1/2		
SCALAR BOSONS				
	gluon g	0 0 1	~125.09 GeV/c ² 0 0	
	photon γ	0 0 1	Higgs H	
	Z boson Z	0 1	~91.19 GeV/c ²	
	W boson W	±1 1	~80.39 GeV/c ²	

Fermioni

SPIN

1/2

Bozoni

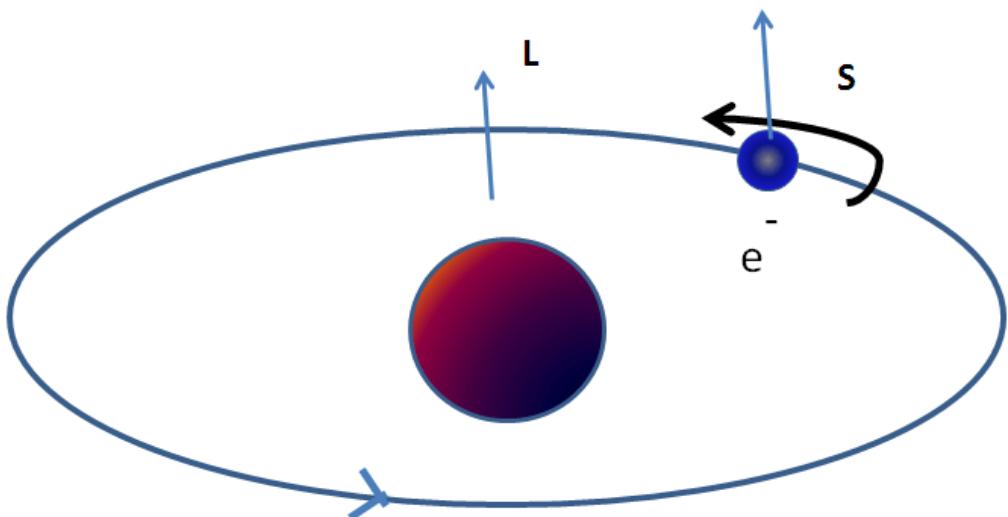
1

0

Kutna količina gibanja

Vrsta kutne količine gibanja	Simbol	Dopuštene vrijednosti
Orbitalni	L	0,1,2,...
Spin (intrizični)	S	0,1/2, 1, 3/2, 2, ...

Ukupna kutna količina gibanja



$$\vec{J} = \vec{L} + \vec{S}$$

Kvantna teorija + Teorija relativnosti

- Schroedingerova jednažba nije relativistička
- Spajanjem dvije teorije dobivamo važne nove pojave:
 - Za svaku česticu postoji i **antičestica**
 - Čestice se mogu stvarati i poništiti:
Broj čestica nije očuvan



BBC FOUR

$$V = \sum P^a V_{p^a}$$

$$V = V_0 - \sum V_A \left\{ 1 - \frac{1}{r_A} \right\}$$

PAUL DIRAC

PHYSICIST

antisym

Čestice i ... Antičestice!



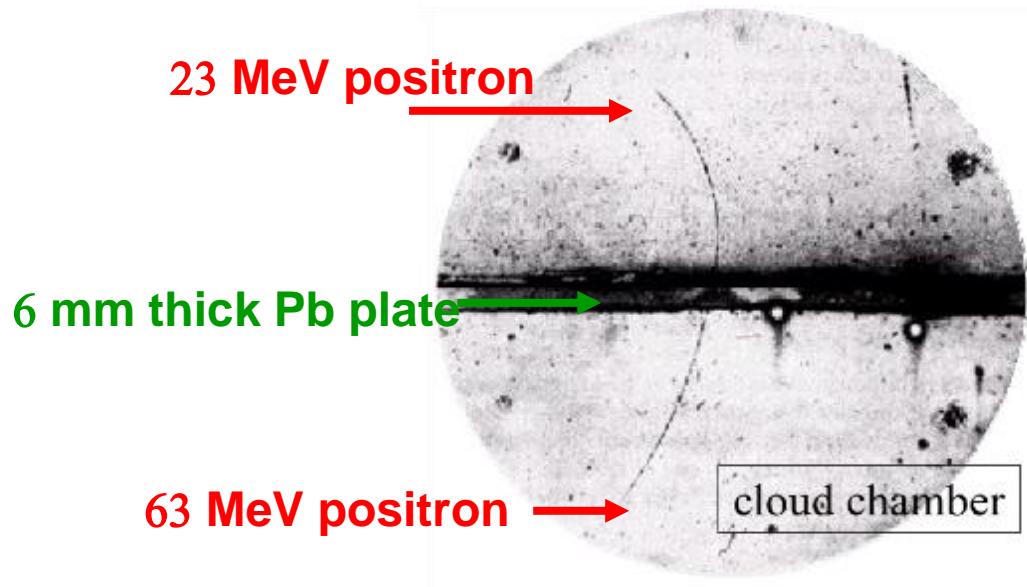
P.A.M. Dirac

Relativnost i kvantna fizika zahtjevaju

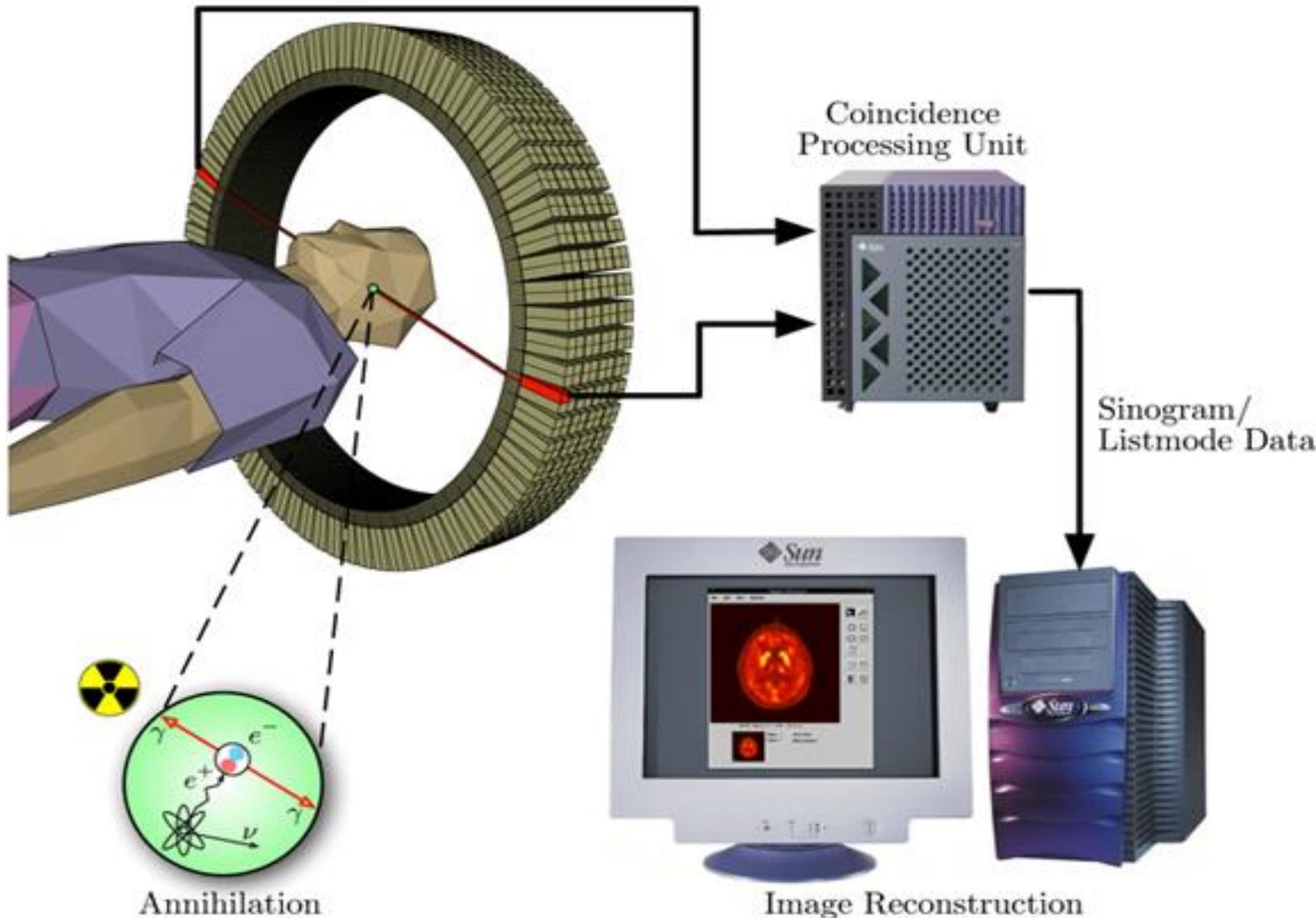
Za svaku česticu postoji i antičestica sa istom
masom i suprotnim kvantnim brojevima

“Otkrio” kao teorijski zahtjev Dirac (1928)

Potvrda: Otkriće pozitrona (anti-elektron): C.D.Anderson (1932)

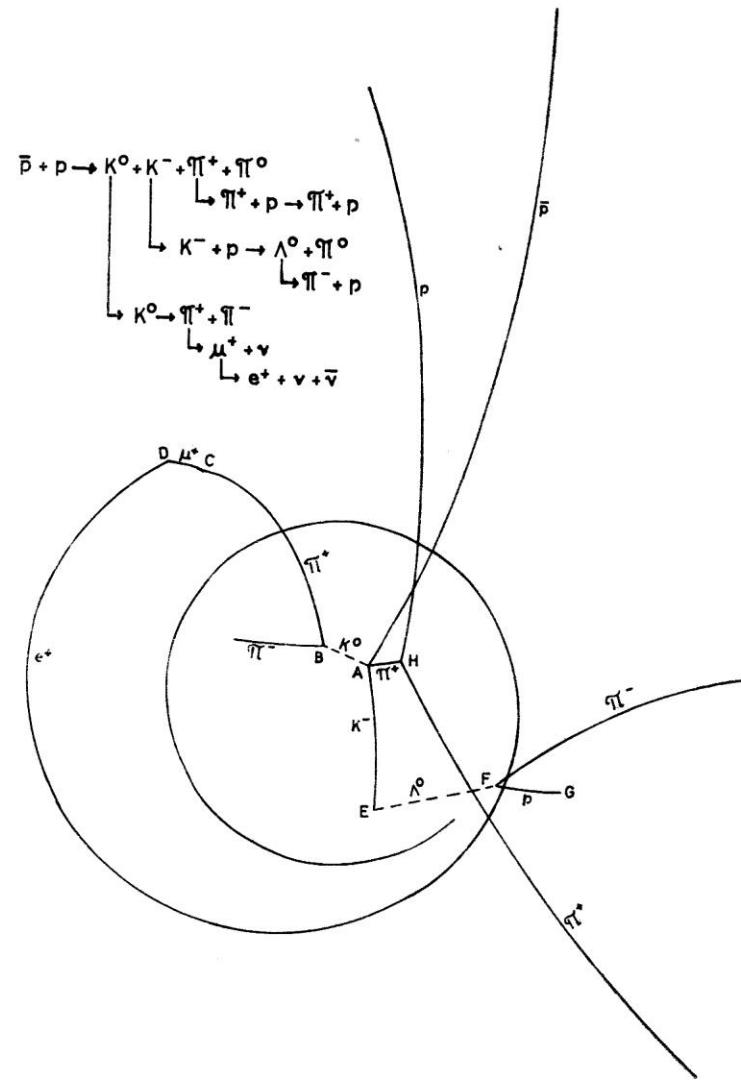
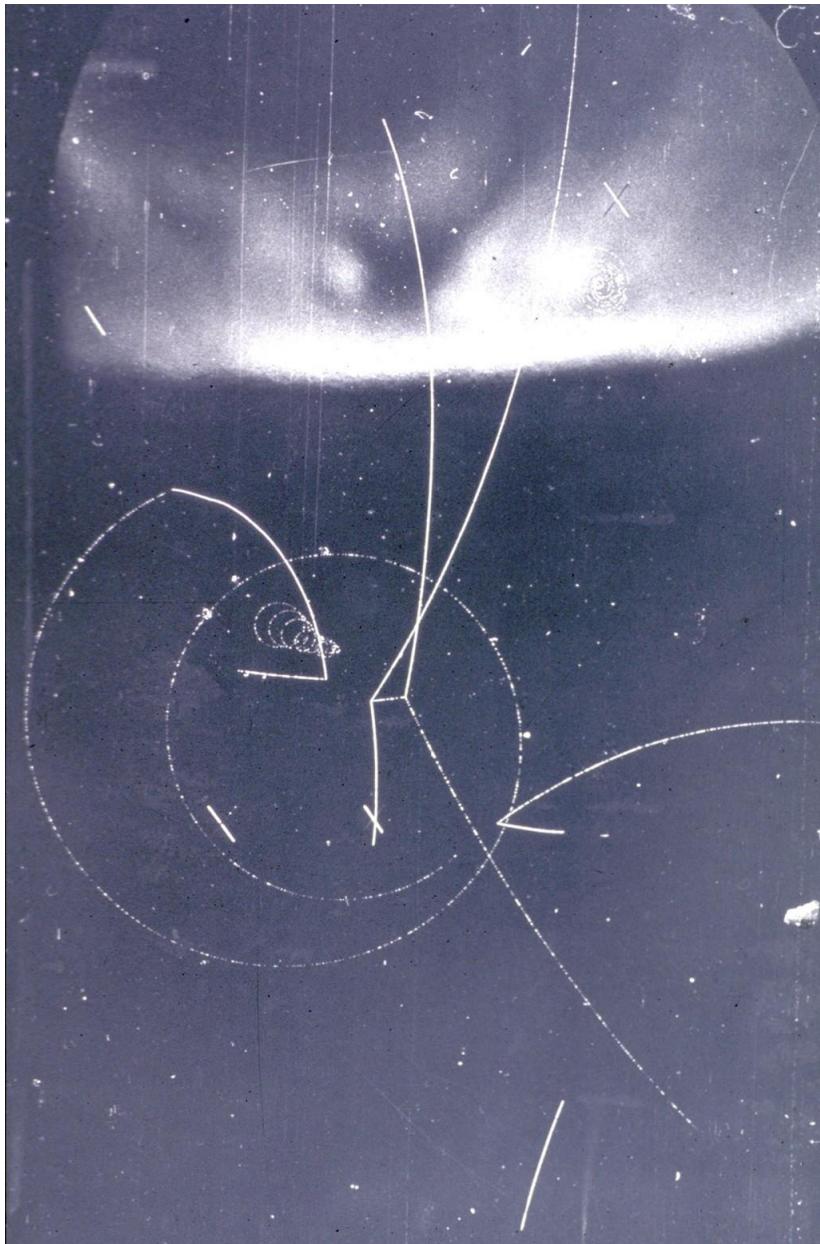


Antimaterija u medicini: PET Tomografija





Example of antiproton annihilation at rest in a liquid hydrogen bubble chamber



1932: Svijet je jednostavan

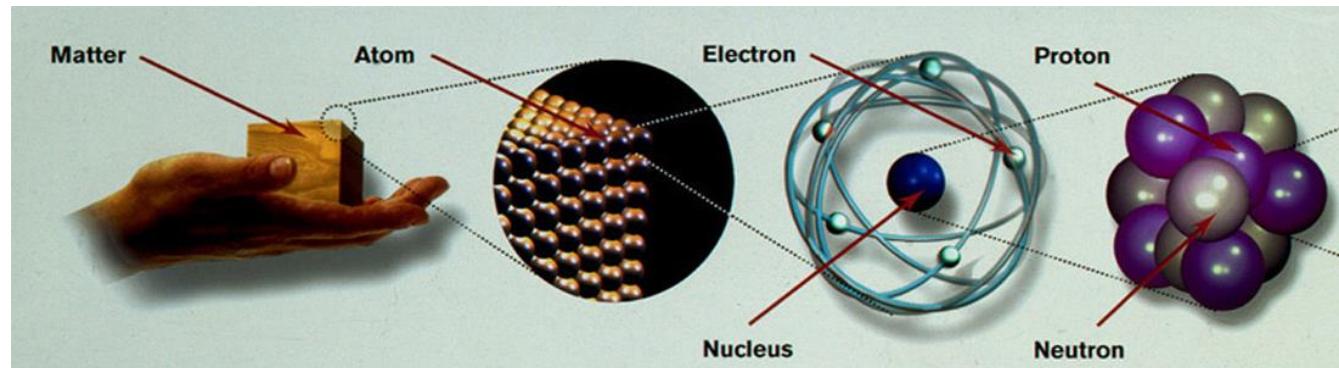
Otkriće neutrona (Chadwick, 1932)

Neutron: Neutralna čestica sa masom
kao proton



Elementarne čestice 1932 g.:

- proton: $Q = + 1 e$
- Neutron: $Q = 0$
- Elektron: $Q = - 1 e$

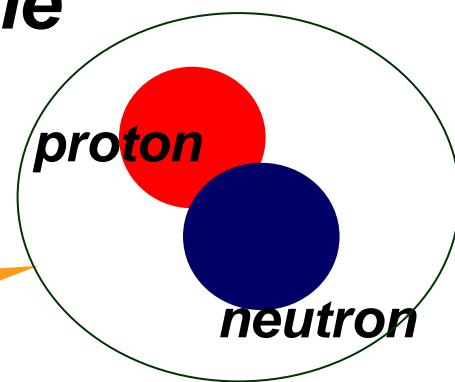
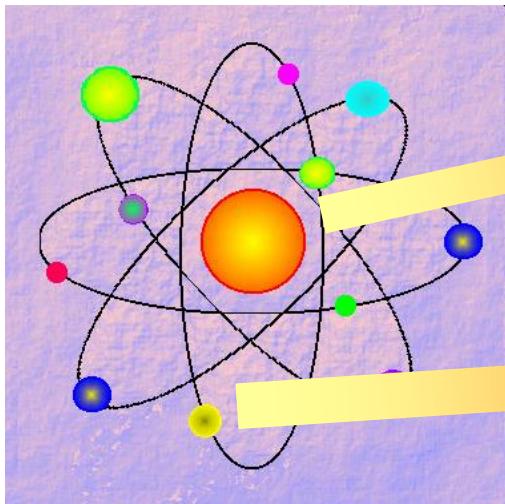


H															He
Li	Be														
Na	Mg														
K	Ca														
Rb	Sr														
Cs	Ba														
Fr	Ra														
Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr
Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe
Lu	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn
Lr	Rf	Db	Sg	Bh	Hs	Mt									
La	Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb		
Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No		

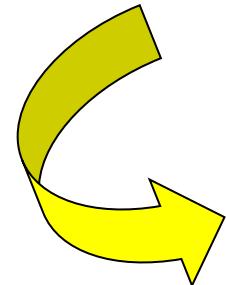
?

Periodic Table

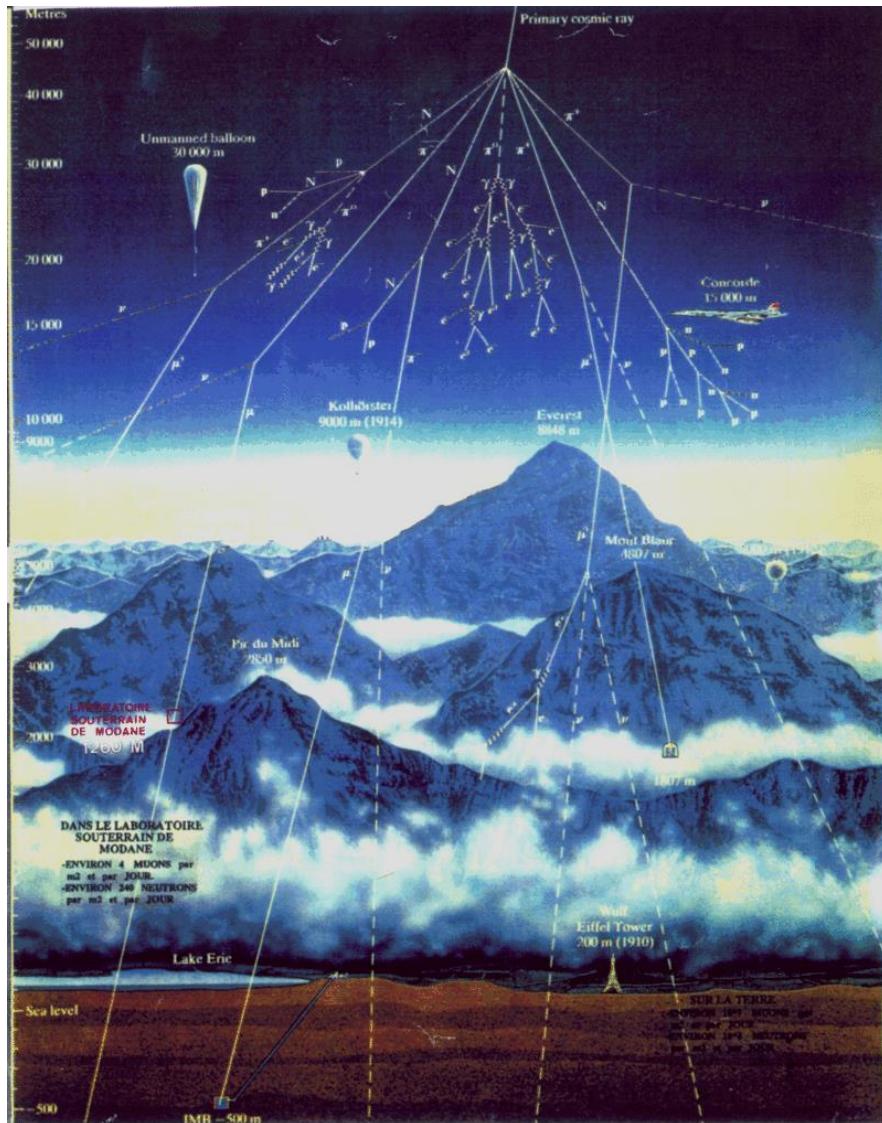
Atom



electron

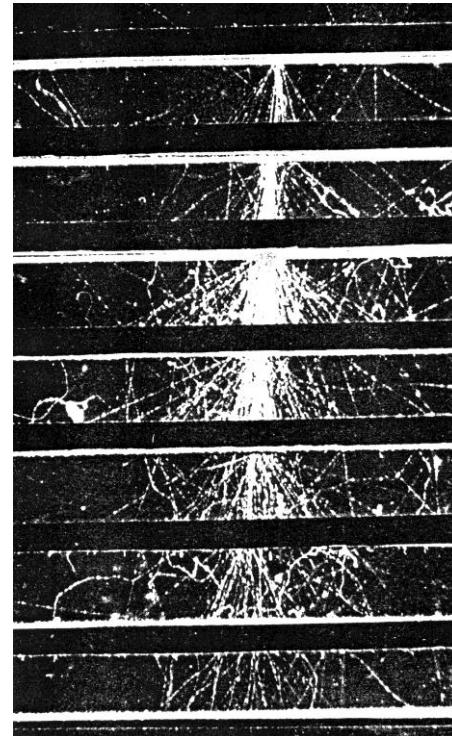


Prvi snopovi: padaju s neba



Kozmičko zračenje

Jedini izvor visoko energetskih čestica do ~1950



Cloud chamber image of an electromagnetic shower.
Pb plates, each 1.27 cm thick

Svijet se zakomplicira...

U kozmičkom zračenju
otkriveni

μ

potpuno neočekivan

(Rabbi: “Who ordered that?”)

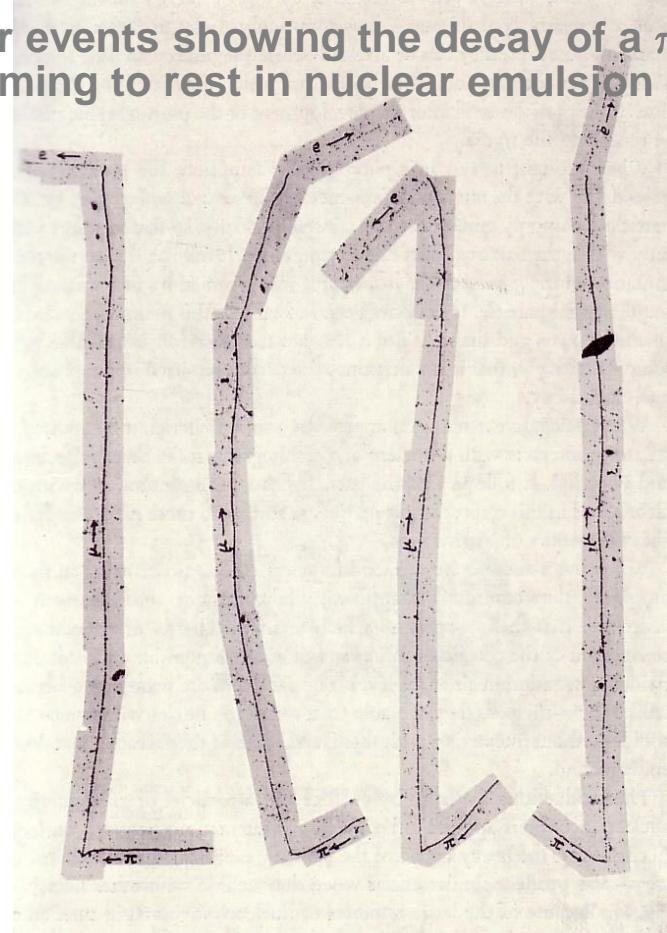
– Teški elektron

π

– Iz sudara zračenja sa
jezgrama

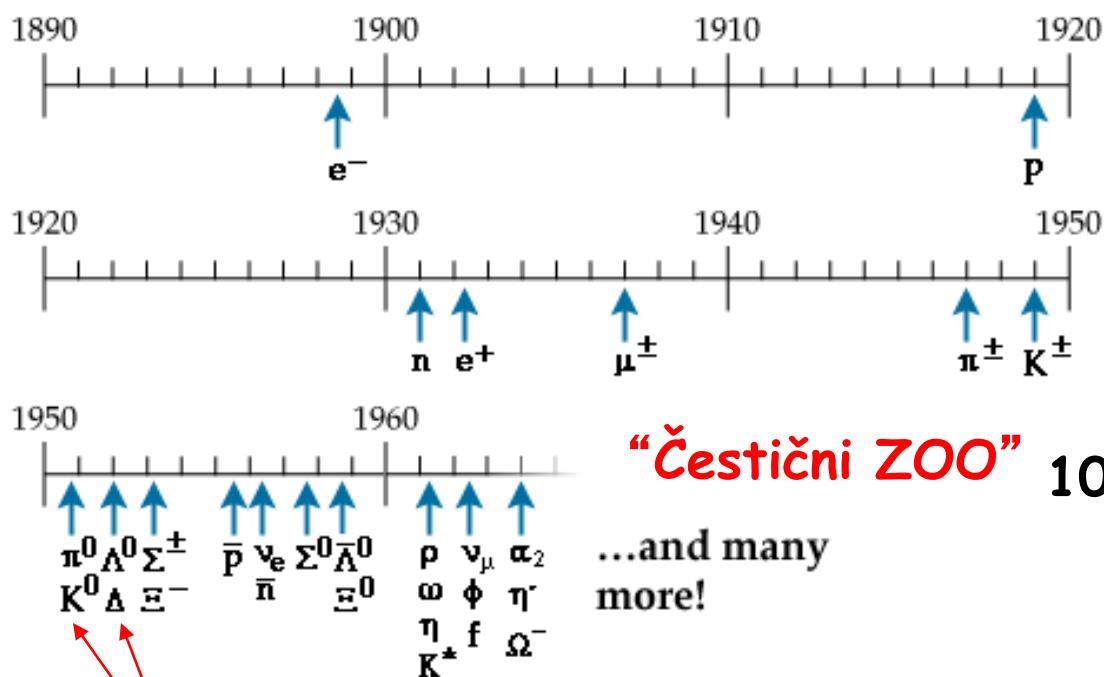
$$\pi^+ \rightarrow \mu^+ \nu_\mu ; \pi^- \rightarrow \mu^- \nu_\mu$$

Four events showing the decay of a π^+ coming to rest in nuclear emulsion

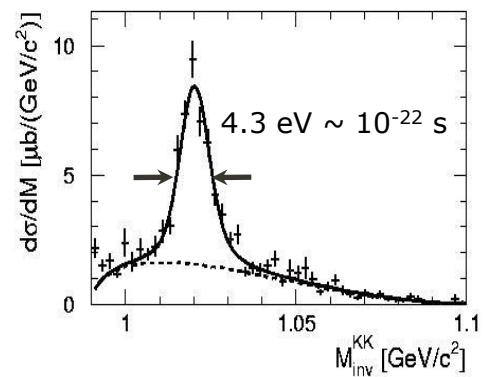


... sve više i više

~ 1950-1960: puno novih čestica otkrivene u kosmičkom zračenju i na prvim akceleratorima (CERN, BNL, ...)



“Strange particles” (Murray Gell-Mann)



“Čestični ZOO”
Ponovo više od
100 elementarnih čestica!

...and many
more!

ali
Jesu li stvarno?

Kvarkovi?

1964 (Gell-Mann, Zweig): Hadroni bi mogli biti građeni od 3 elementarnih čestica:
Gell-Mann ih krsti "quarks"

	<i>u</i>	<i>d</i>	<i>s</i>
Electric charge (units $ e $)	+2/3	-1/3	-1/3
Baryonic number	1/3	1/3	1/3
Strangeness	0	0	-1

i 3 antikvarka (\bar{u} , \bar{d} , \bar{s}) sa suprotnim kvantnim brojevima

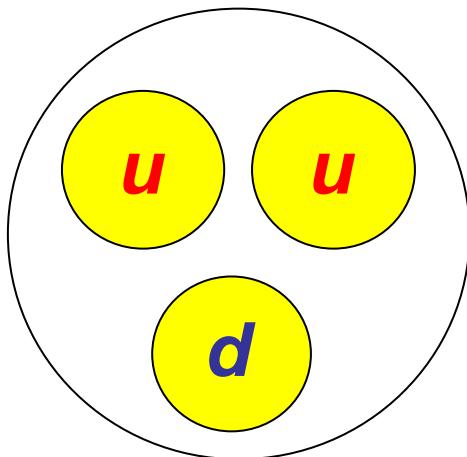
1964 se predlaže ideja

kvarkova

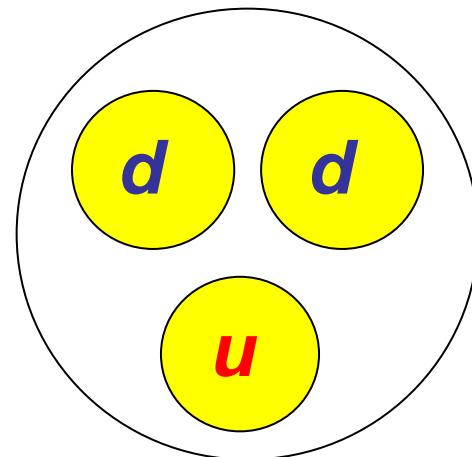


Zweig

Gell-Mann



proton



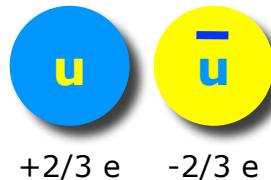
neutron

Elementarne čestica s nabojom manjim od elementarnog naboja (elektrona) koje se pojavljuju u više vrsta ili okusa

Kvarkovi – Moguće kombinacije

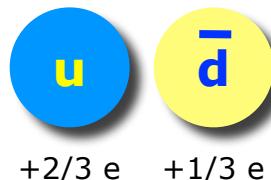
Mezoni

= Kvark-Antikvark

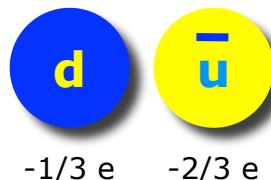


π^0

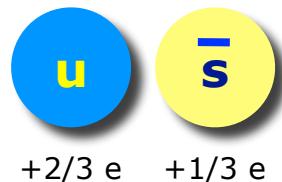
Nekoliko
poznatih
mezona



π^+



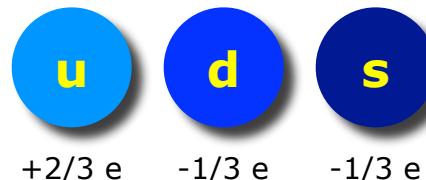
π^-



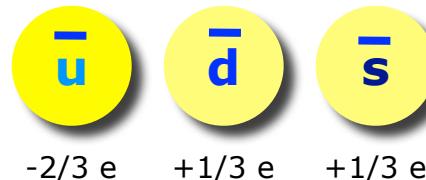
K^+

Barioni:

3 kvarka ili
3 antikvarka



Λ

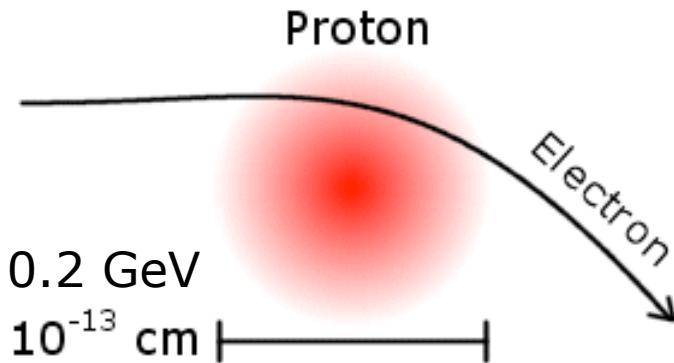


$\bar{\Lambda}$

Kvarkovi postoje!

Otkriće kvarkova

Raspršenje elektrona na protone

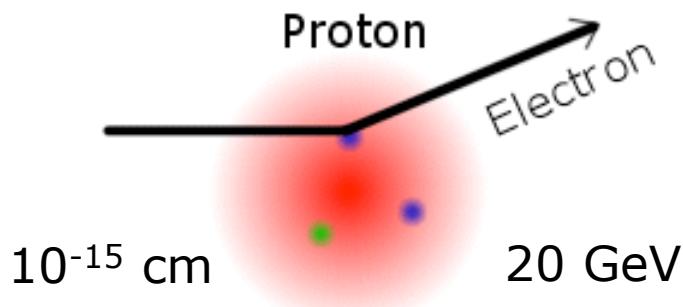


1956 Hofstadter: Mjerenje radiusa protona



Stanford Linear Accelerator Centre

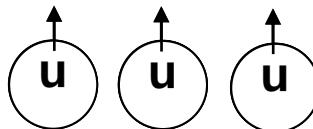
**1967 Friedmann, Kendall, Taylor (SLAC): ~ Rutherford eksperiment s elektronima
--> u protonu se nalaze 3 'tvrda' (točkasta) centra raspršenja**



Mjerenja u skladu s prisutnošću 2 up- 1 down-
Kvarka u protonu

Jaka sila vidi... BOJU!

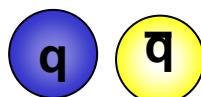
$$\Delta^{++}(u^{\dagger}u^{\dagger}u^{\dagger})$$



A moj princip
isključenja!?



Slaganje boja u nešto bezbojno



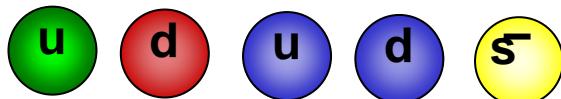
Mezoni



Barioni

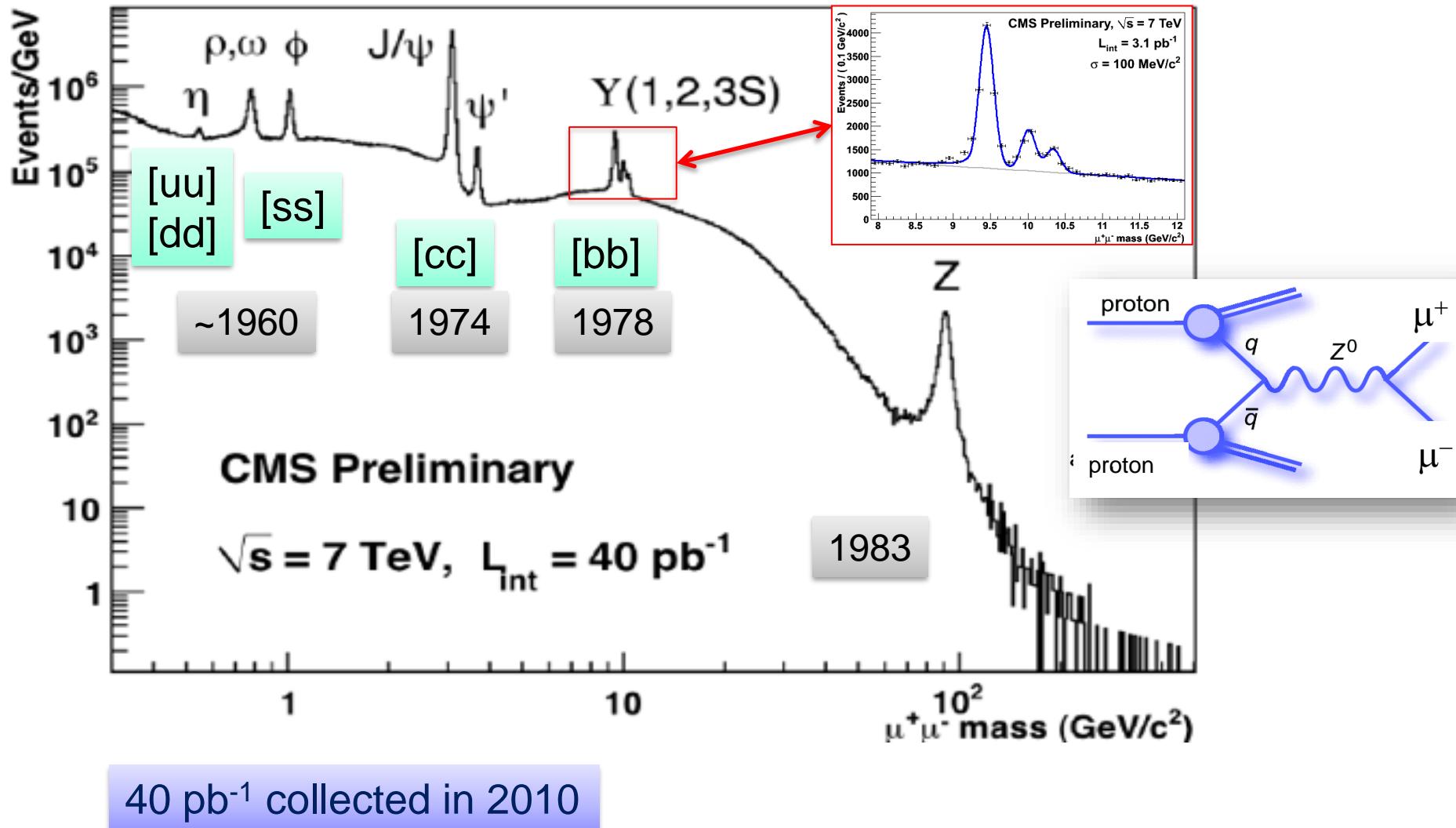


Anti-barioni



Egzotični barion

Nove rezonancije: novi kvarkovi!



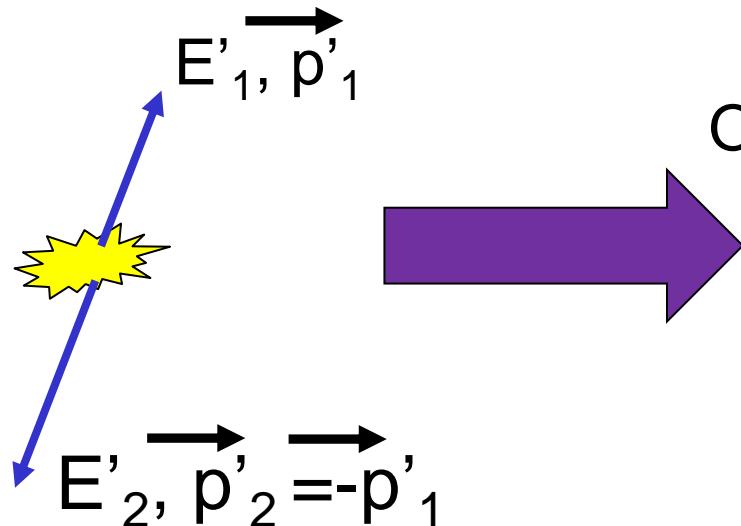
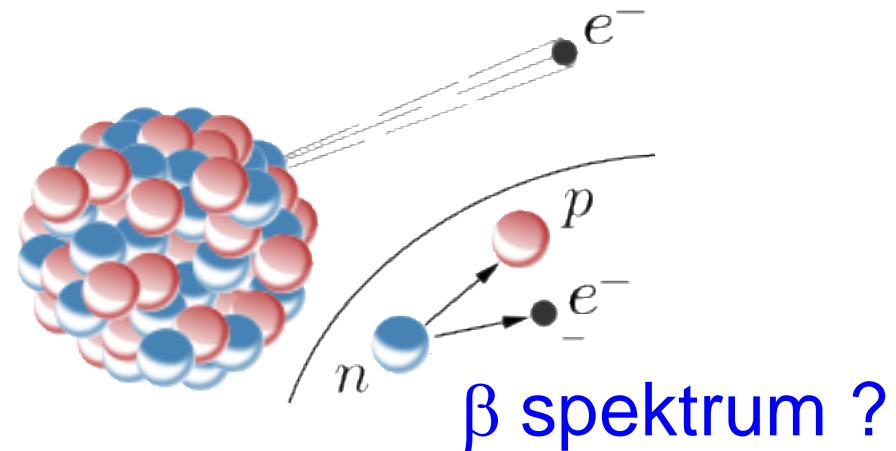
Enigma β raspada

Neke jezgre se raspadaju putem β raspada, npr.



SPIN: 0 1 1/2

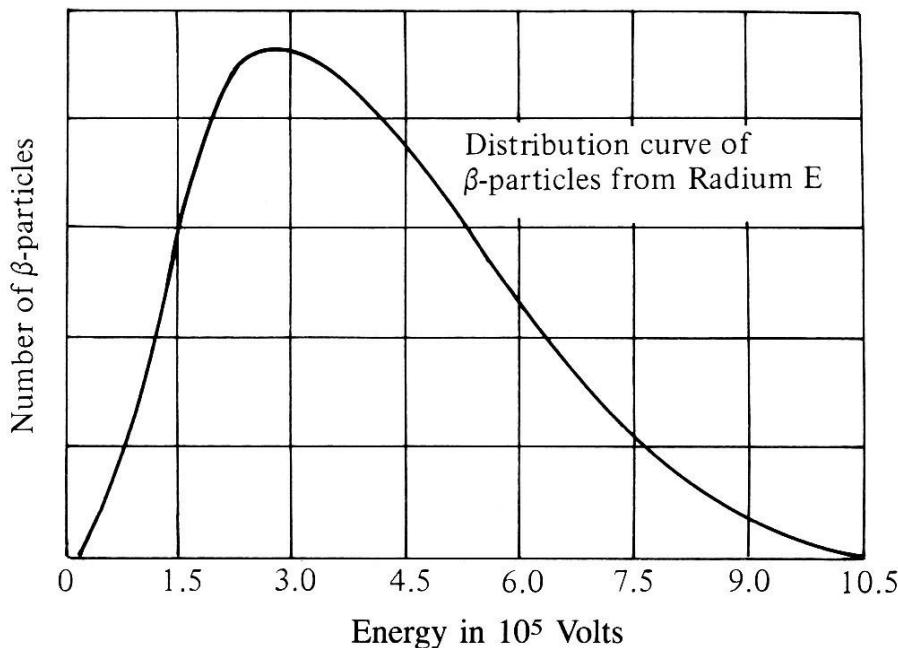
Očuvanje kutne
Količine gibanja????



Problem β raspada: kontinuirani spektar

Problem sa β spektrom
(trebao bi biti monokromatičan)

First measurement by Chadwick (1914)



**Kako se onda zamišljao
 β^- raspad: $n \rightarrow p + e^-$**

Radium E: $^{210}\text{Bi}_{83}$
(a radioactive isotope
produced in the decay chain
of ^{238}U)

Očuvanje energije?

Pauliovo pismo

December 1930: public letter sent by W. Pauli to a physics meeting in Tübingen

Zürich, Dec. 4, 1930

Dear Radioactive Ladies and Gentlemen,

...because of the “wrong” statistics of the N and ${}^6\text{Li}$ nuclei and the continuous β -spectrum, I have hit upon a **desperate remedy to save the law of conservation of energy.** Namely, the possibility that there could exist in the nuclei electrically neutral particles, that I wish to call neutrons, **which have spin $\frac{1}{2}$ and obey the exclusion principle** The mass of the neutrons should be of the same order of magnitude as the electron mass and in any event not larger than 0.01 proton masses. The continuous β -spectrum would then become understandable by the assumption that in β -decay a neutron is emitted in addition to the electron such that the sum of the energies of the neutron and electron is constant.

..... For the moment, however, I do not dare to publish anything on this idea
So, dear Radioactives, examine and judge it. Unfortunately I cannot appear in Tübingen personally, since I am indispensable here in Zürich because of a ball on the night of 6/7 December.

W. Pauli

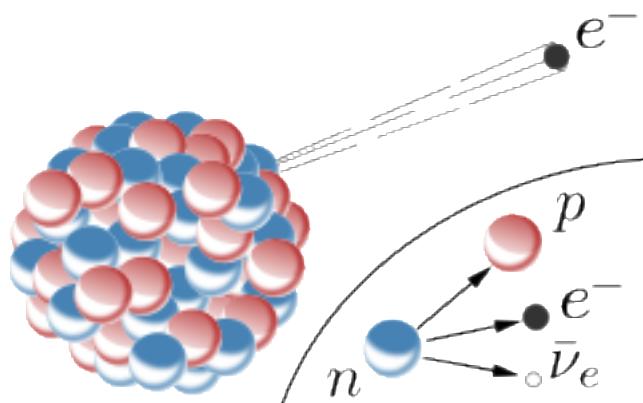
Rješenje problema β raspada: Nova čestica = Neutrino!

Pauli: u β raspadu se emitira još jedna čestica:

- Neutralna
- Nevidljiva (jako slabo međudjeluje)
- Polucijelog spina



SPIN: 0 1 $\frac{1}{2}$ $\frac{1}{2}$: OK!



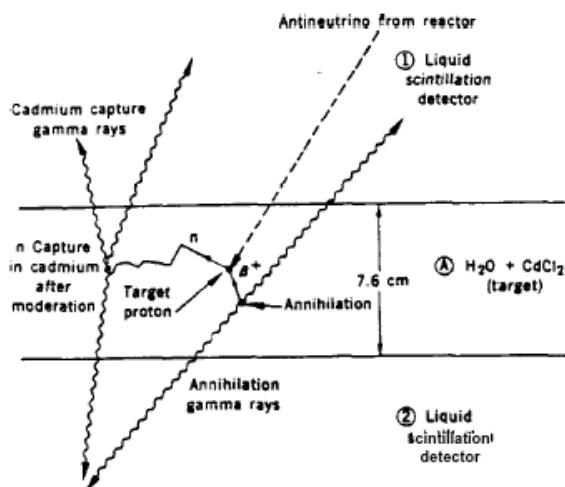
Otkrili smo
nevidljivu česticu
koristeći isključivo
Zakone očuvanja

Neutrini postoje!

ν_e

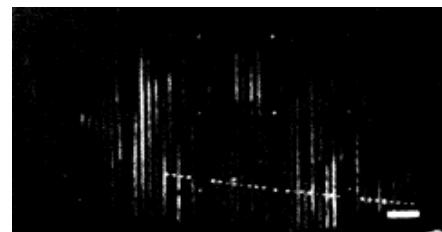
(Reines, Cowan 1953)

$$\bar{\nu} + p \rightarrow e^+ + n$$



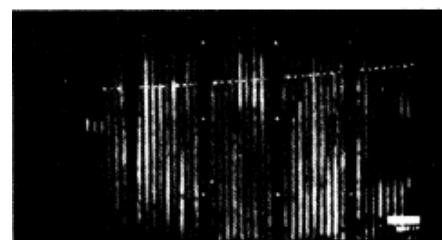
ν_μ

(1962)



ν_τ

(2000)



Three typical single-track events
in the BNL neutrino experiment

Pauli: "I have done a terrible thing. I have proposed a particle that cannot be detected. It is something no theorist should ever do."

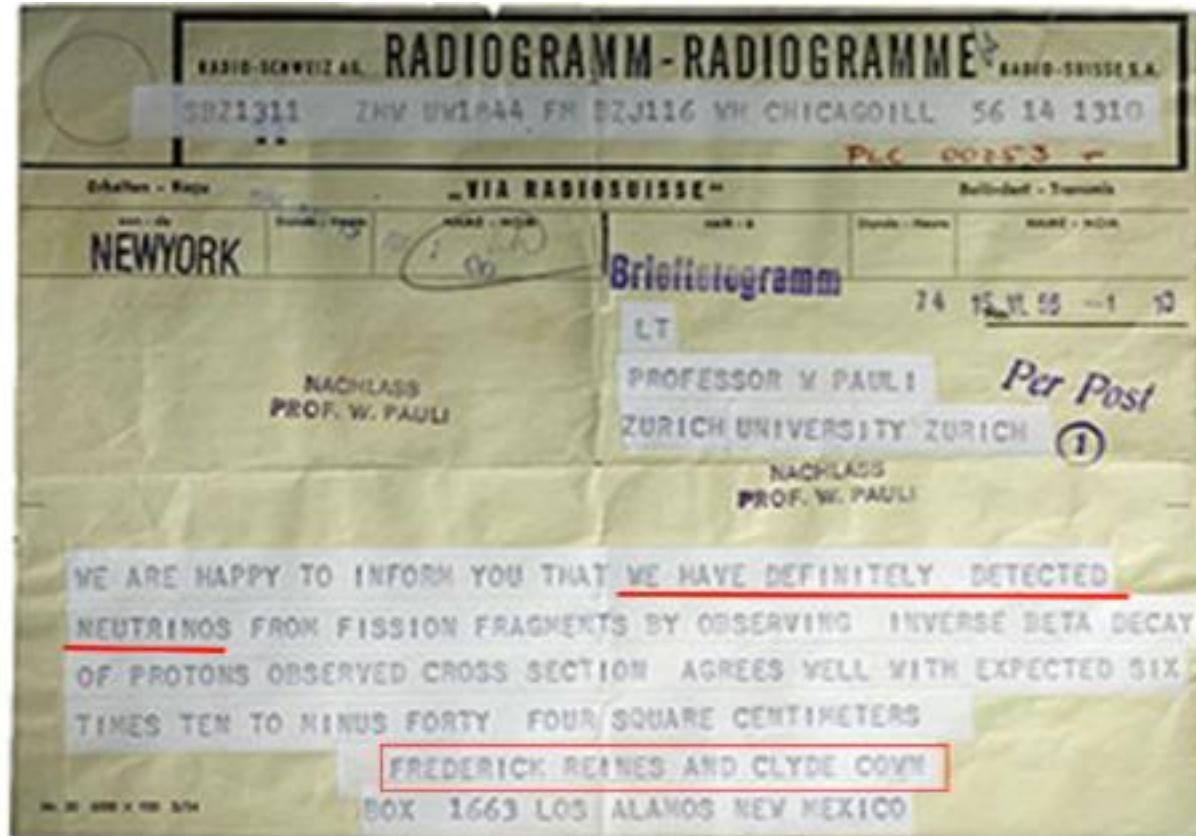


Réacteur de Savannah River (USA)



Reines et Cowan à Pauli (1956)

Pauli :



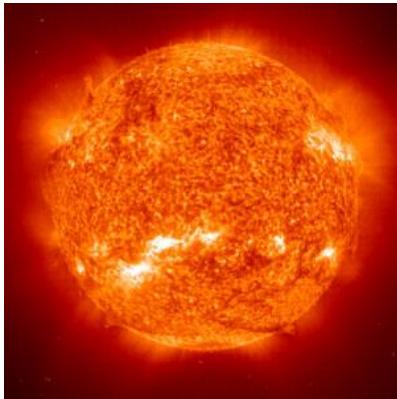
Frederick REINES and Clyde COVAN

Box 1663, LOS ALAMOS, New Mexico.

Thanks for message. Everything comes to
him who knows how to wait.

Pauli:

Neutrini pluštaju sa svih strana



Sunce: 64 milijardi
po cm^2 po sekundi



Supernova iz naše galaksije:
5 milijardi po cm^2 svake
sekunde kroz 10 sekundi



Veliki prasak:
U svakom cm^3 oko 300
neutrina

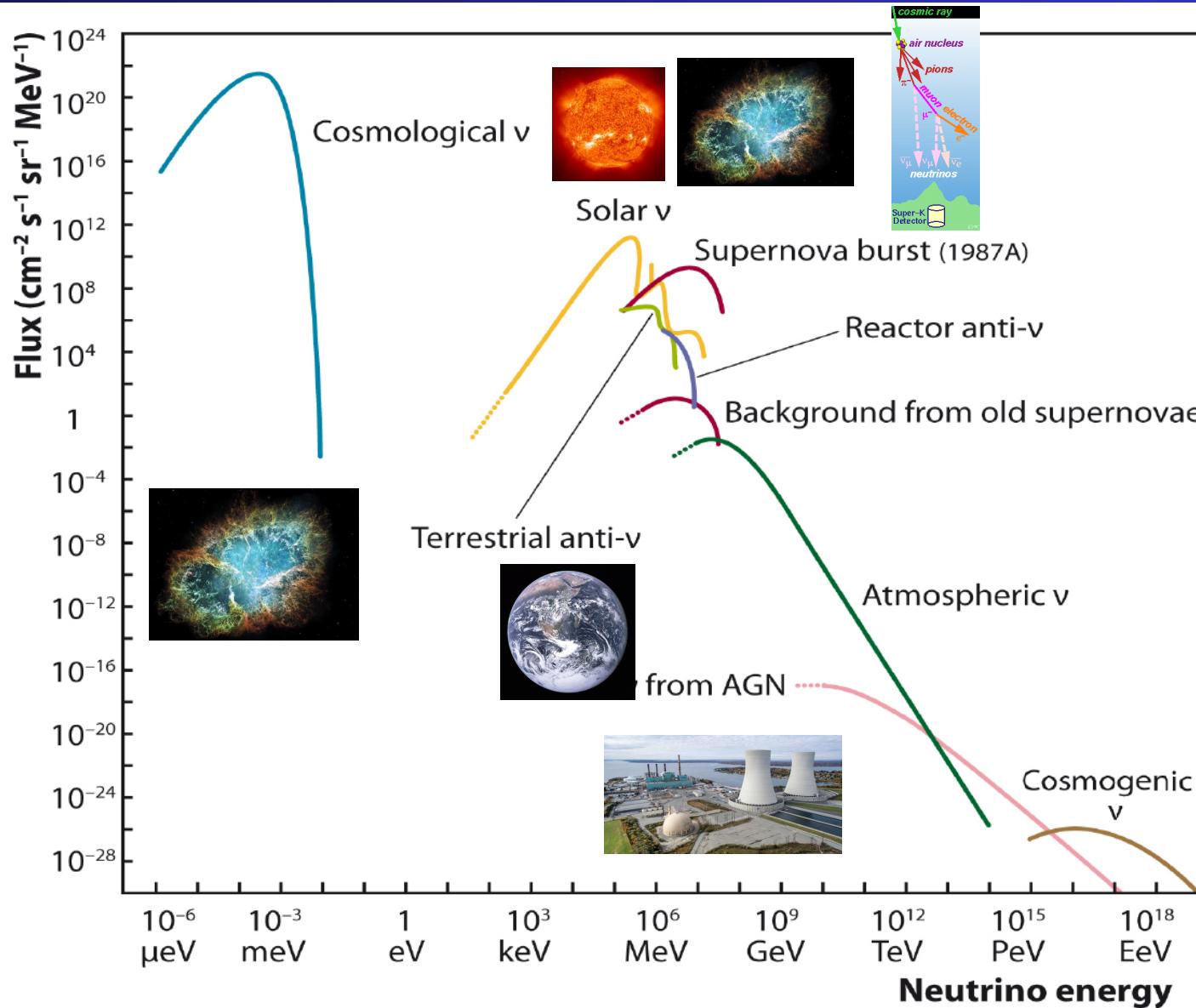


Nuklearke:
Oko 100,000 po cm^2 svake
sekunde na 200km od centrale



Geofizički I atmosferski neutrini:
6 milijuna po cm^2 svake sekunde

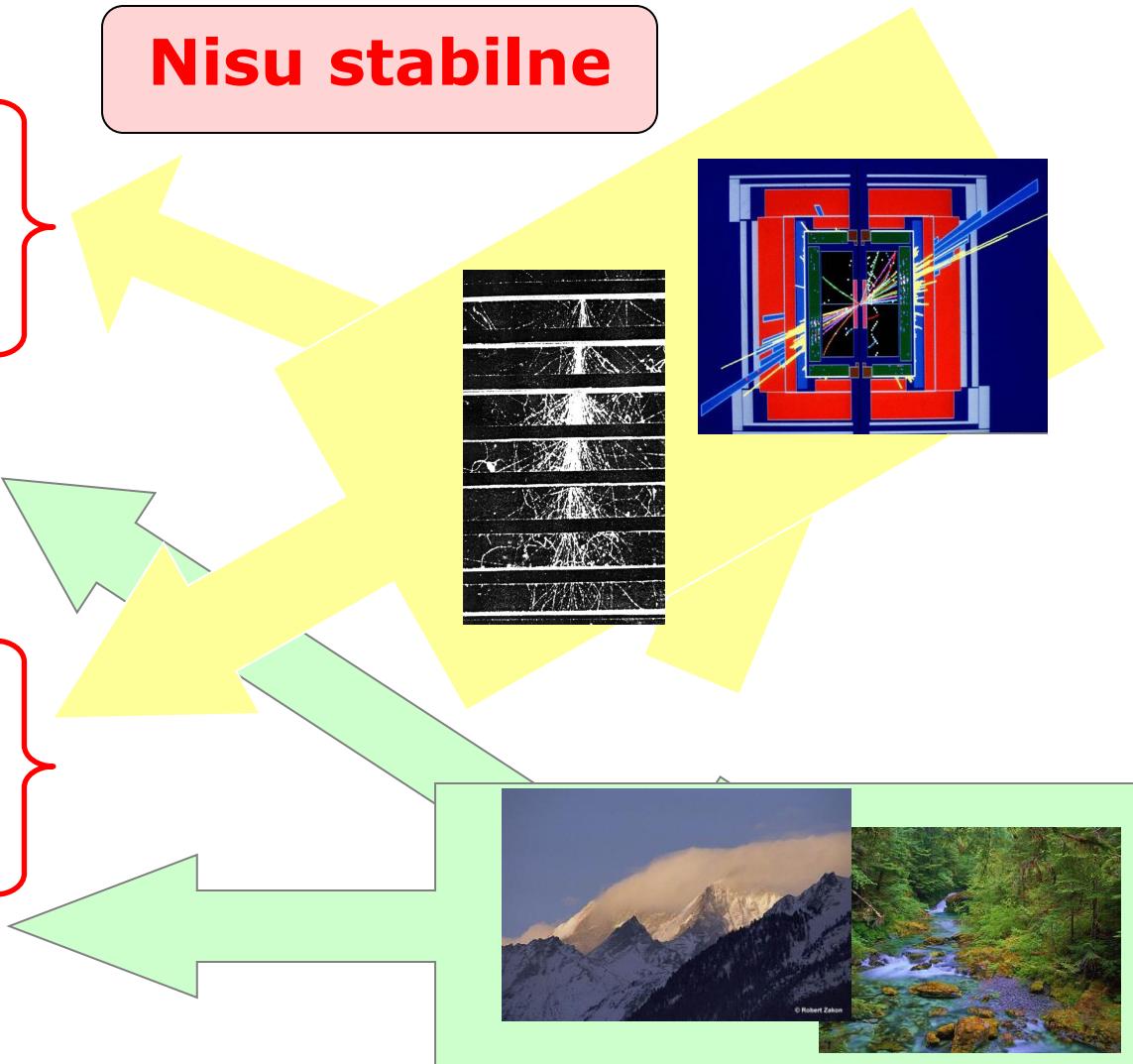
Spektar neutrina



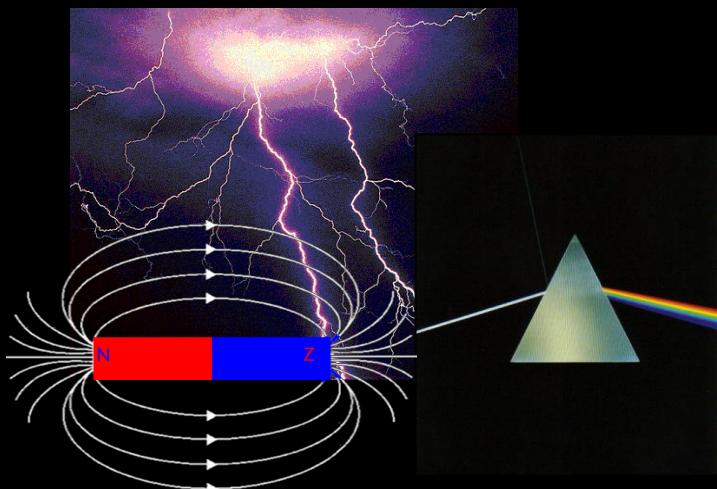
Elementarne čestice, 2019

Leptons		
Electric Charge		
Tau	-1/3	0
Muon	-1	0
Tau Neutrino		
Muon Neutrino		
Electron	-1	0
Electron Neutrino		
Quarks		
Electric Charge		
Bottom	-1/3	2/3
Strange	-1/3	2/3
Up	2/3	2/3
Down	-1/3	2/3
each quark: R, B, G 3 colors		

Nisu stabilne



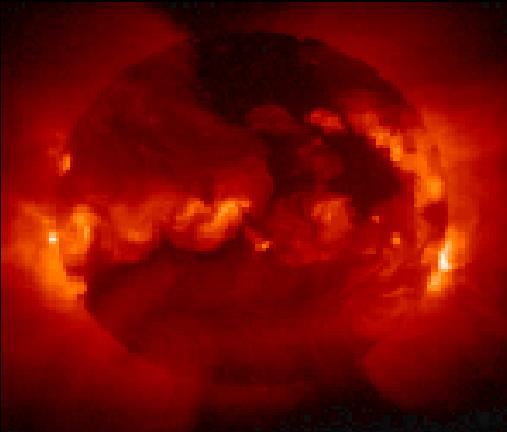
4 temeljne sile



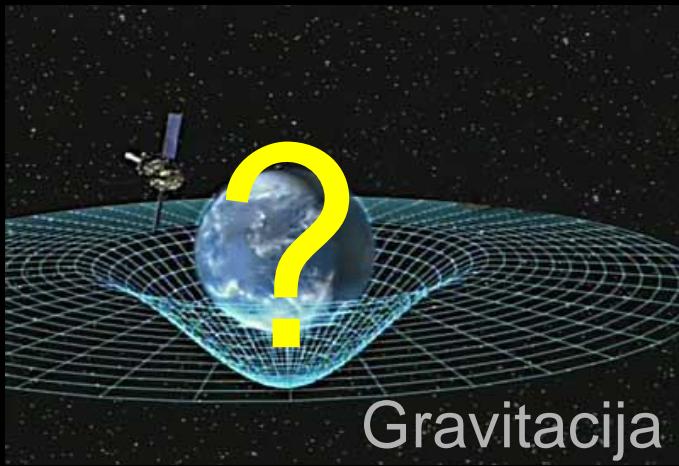
Elektromagnetska sila



Jaka sila



Slaba sila

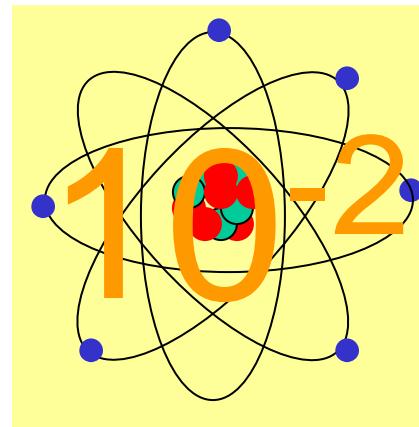


Gravitacija

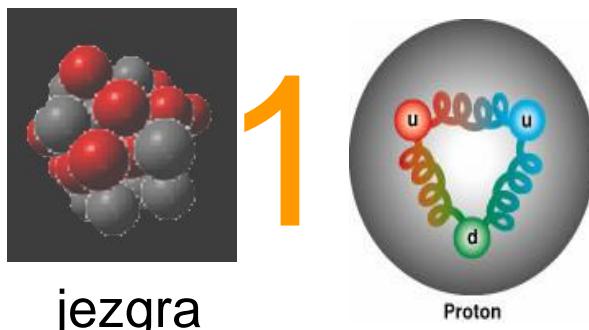
Jačine sila



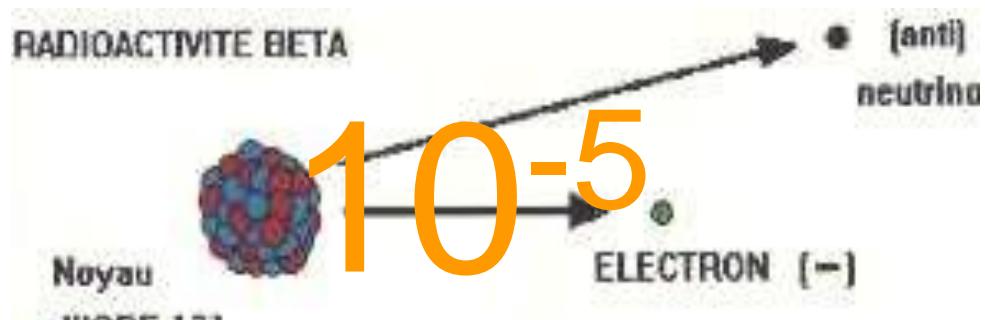
Gravitacijska sila



atom



Jaka nuklearna sila

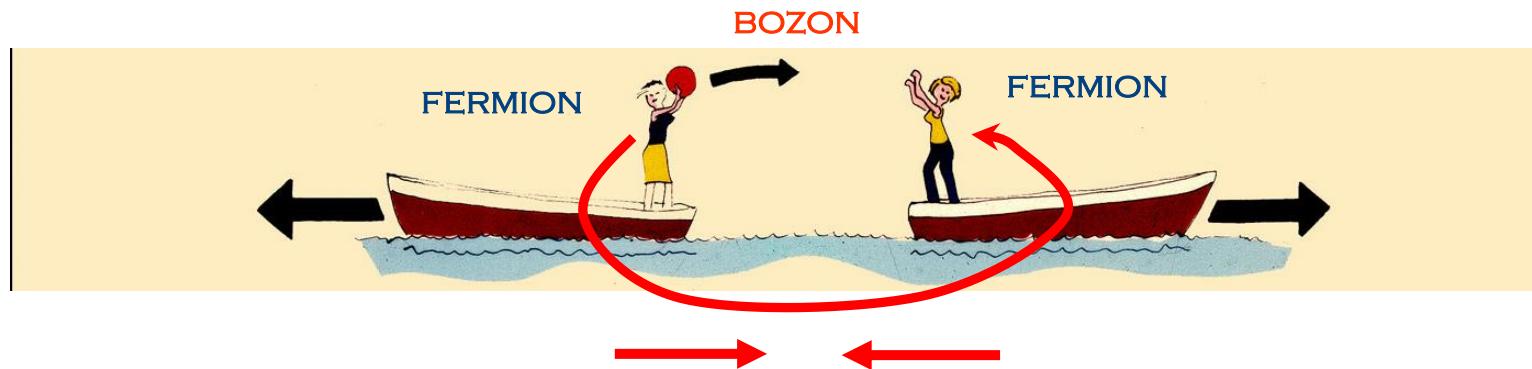


Slaba sila

$$n \rightarrow p + e^- + \nu_e$$
$$d \rightarrow u + e^- + \nu_e$$

Sile: Izmjena čestica

Relativnost: nema djelovanja na daljinu!



Izmjena bozona je izvor sila između fermiona

sila	bozon	primjer
jaka	gluon (bez mase)	atomska jezgra
elektromagnetska	foton (bez mase)	atomske orbite
Slaba	Z^0 $W^+ W^-$ (teške)	radiaktivnost
gravitacija	graviton (bez mase) ??	nebeska tijela

Doseg sila

Stvaranje i izmjena čestice nosioca sile

-> Energija nije očuvana...

za kratko vrijeme!

$$\Delta t \approx \frac{\hbar}{\Delta E} = \frac{\hbar}{mc^2}$$

U vremenu Δt čestica može proći $R = c \Delta t$

$$R \approx \frac{\hbar c}{mc^2}$$

Npr.:

- Nositelj sile nema mase ($m=0$): $R = \infty$
- Nositelj sile mase $m=80$ GeV: $R = 10^{-3}$ fm

Kvantna elektrodinamika

Kvantizirana
elektromagnetska polja:
Foton

Brzina c : masa = 0

Elektromagnetske
interakcije

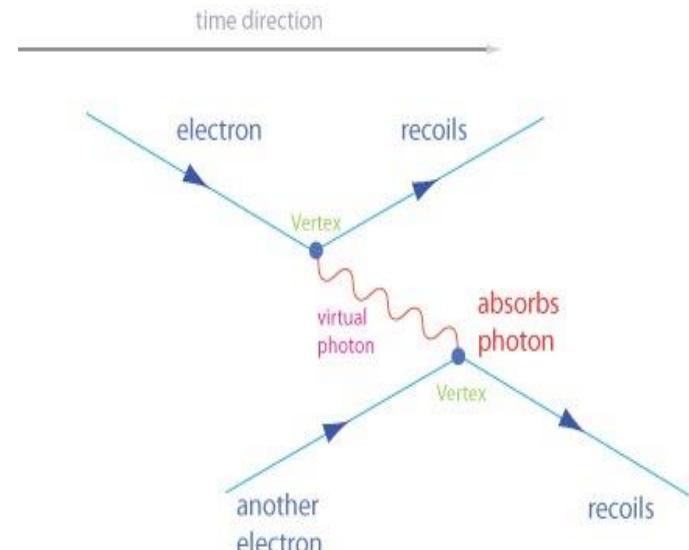
=

Emisija i apsorpcija fotona

Električni naboj očuvan!



R. P. Feynman



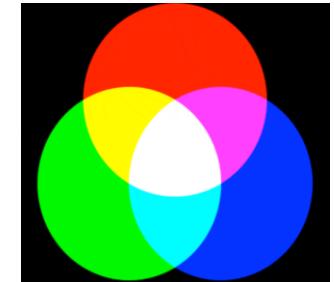
Feynmanovi diagrami

Jaka sila: QCD

Teorija jake interakcije:

Kvantna kromodinamika (**QCD**)

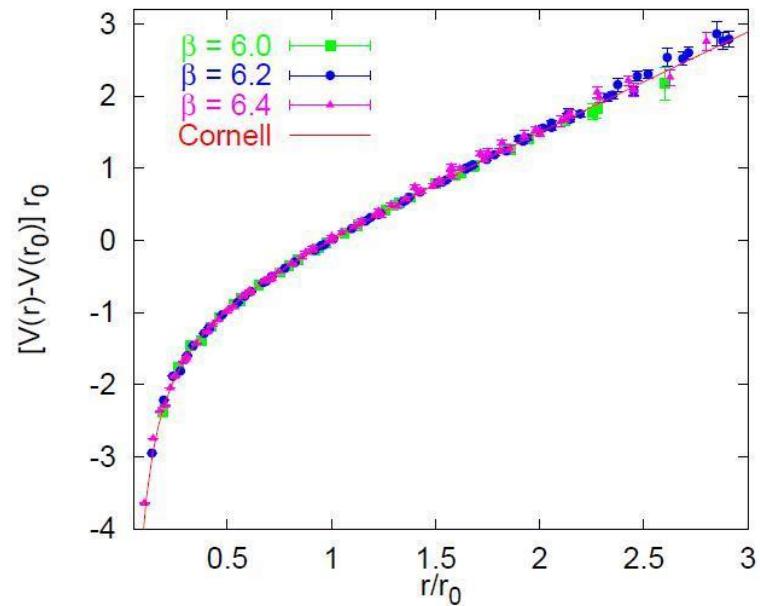
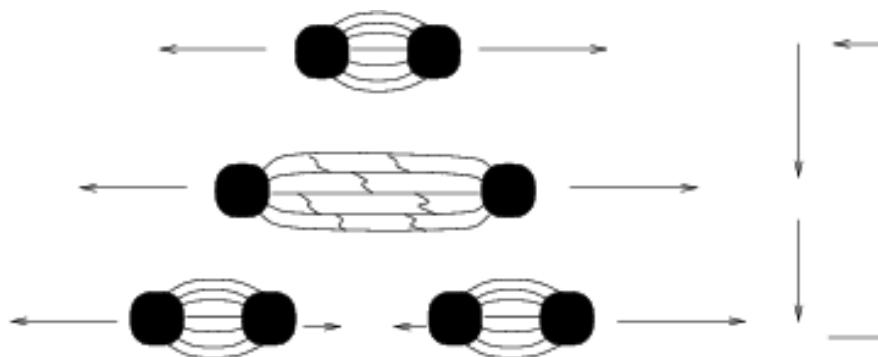
- 3 različita naboja: boje (crveno, zeleno plavo)
- Nosioci jake sile (8 gluona)



SAMO bezbojna stanja mogu postojati: nema slobodnog kvarka

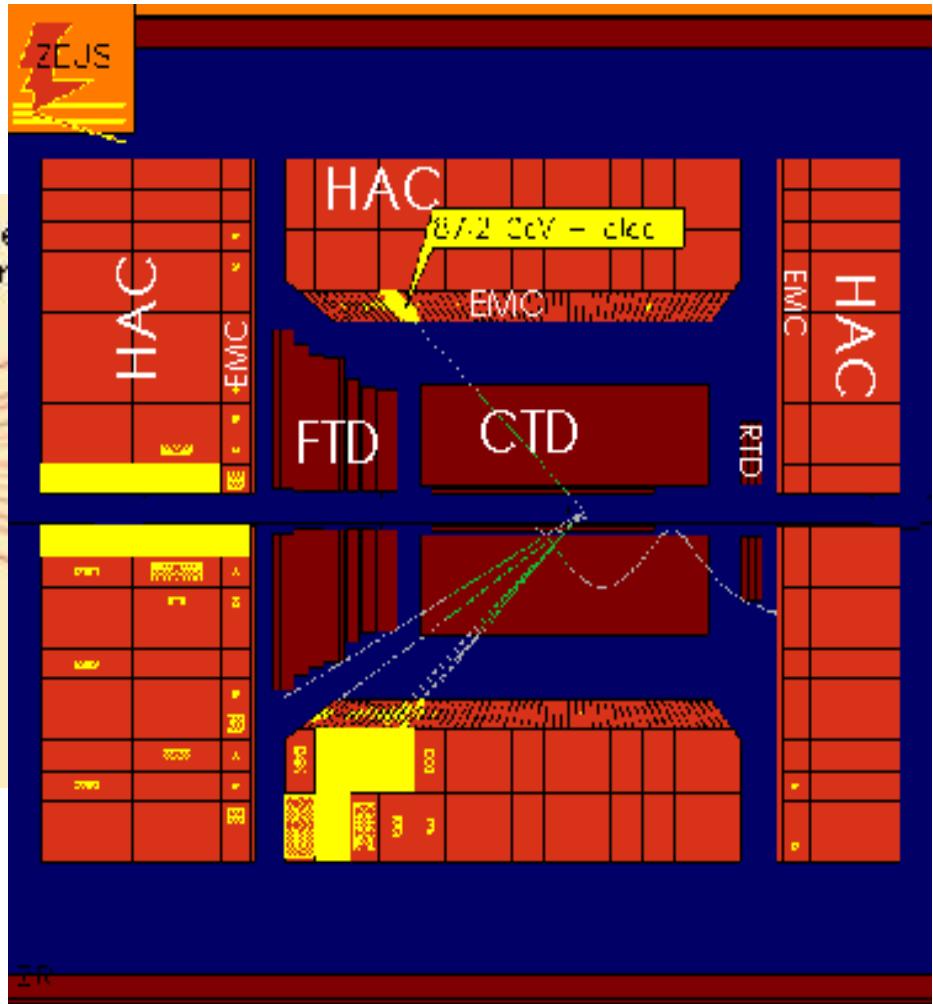
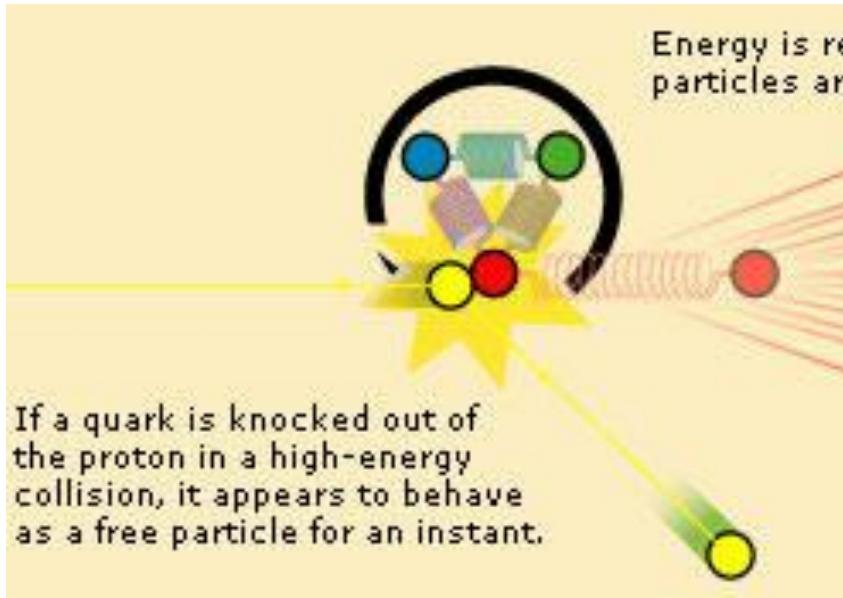
Nema slobodnog kvarka: “Confinement”

Zbog QCD potencijala:
isplativije energički
izvući iz vakuma
par kvark-antikvark
za stoviriti bezbojne
hadrone

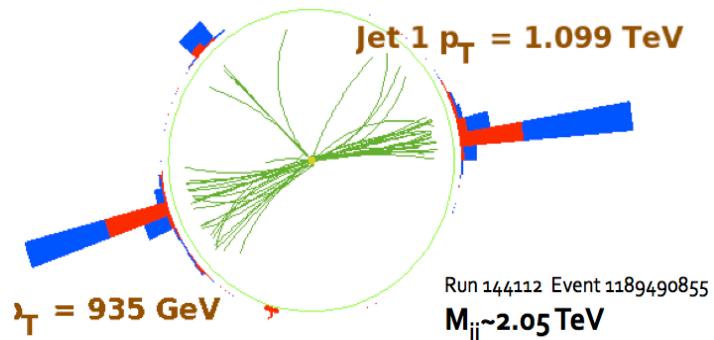
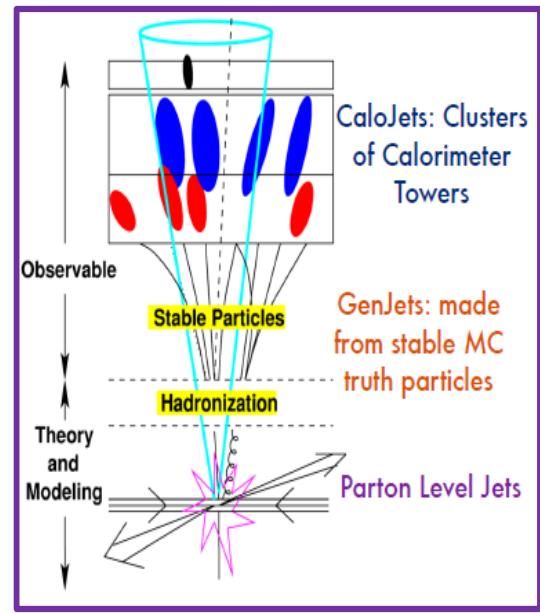
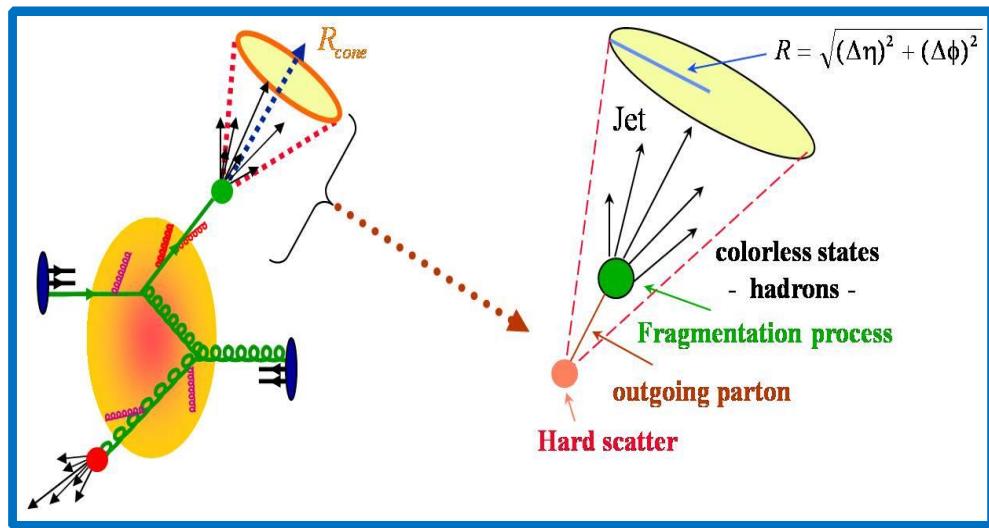


Posljedica:
Hadronski mlaz
“Jet”

Jets (*hadronski mlazovi*)

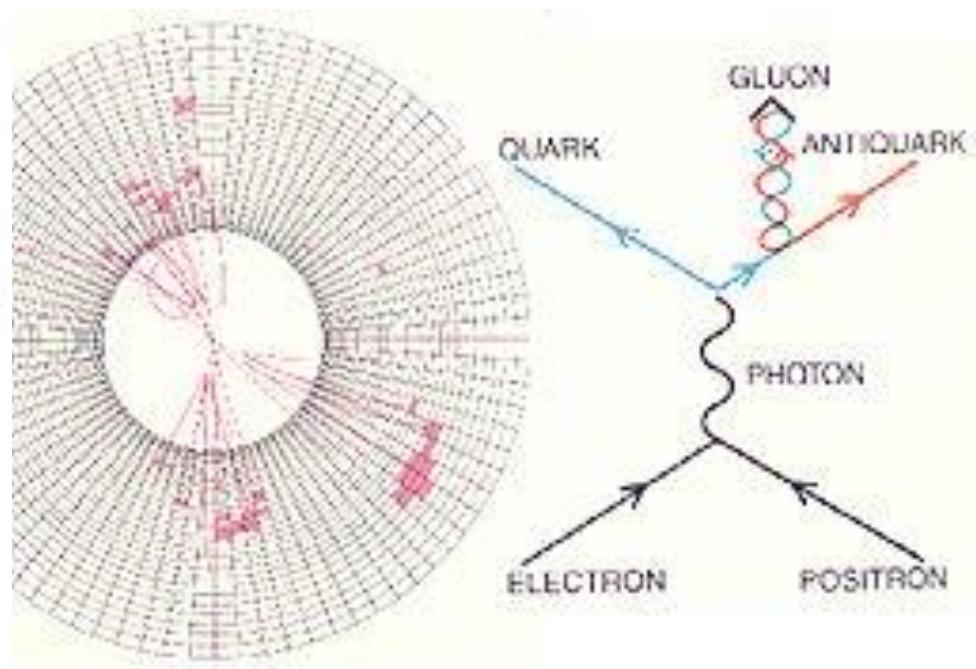
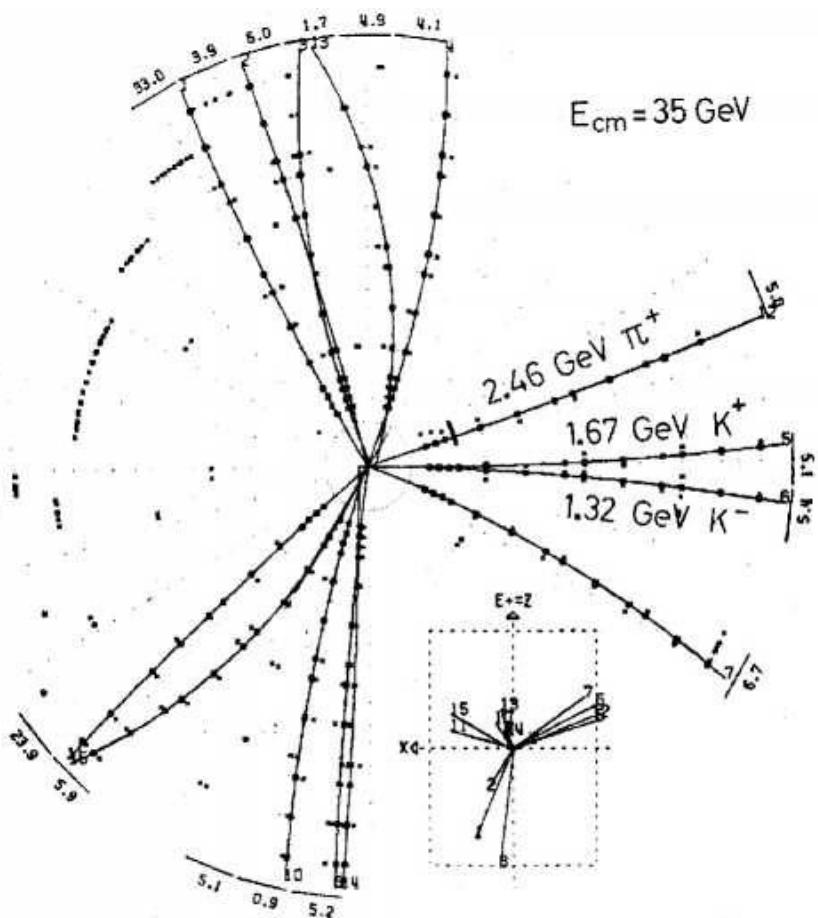


Kvarkovi - Jets



QCD: Gluoni

Otkriće gluona
(DESY, 1979)



PETRA Storage Ring, 1979, DESY (Hamburg)

Očuvano: broj bariona

Svakoj čestici se pripisuje barionski broj:

Čestica	e,μ,τ	Neutrino	Kvark	antikvark
B	0	0	+1/3	-1/3

Barion: +1

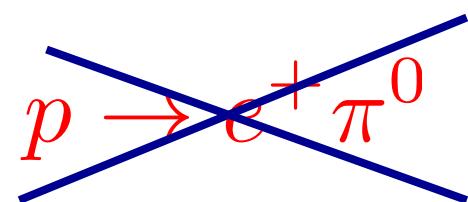
Antibarion: -1

Mezon: 0

Činjenica: barionski broj je očuvan u svim poznatim reakcijama

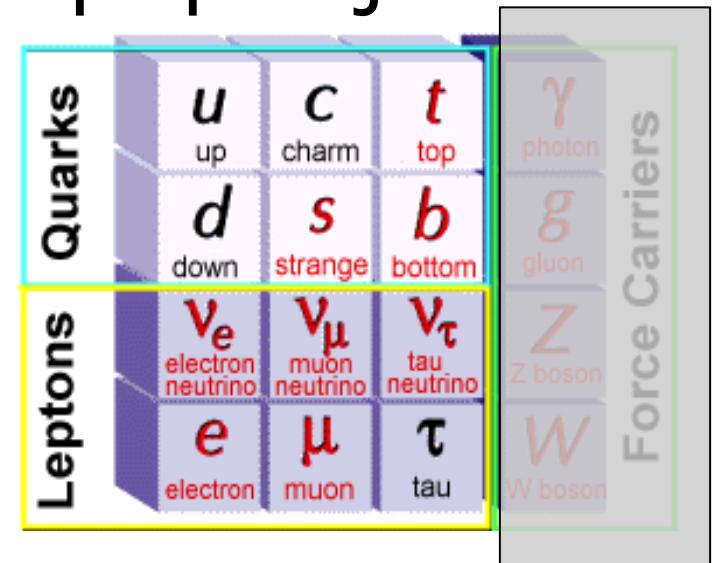
= Zbroj kvarkova i antikvarkova je konstantan

Npr. zabranjuje: Raspad protona

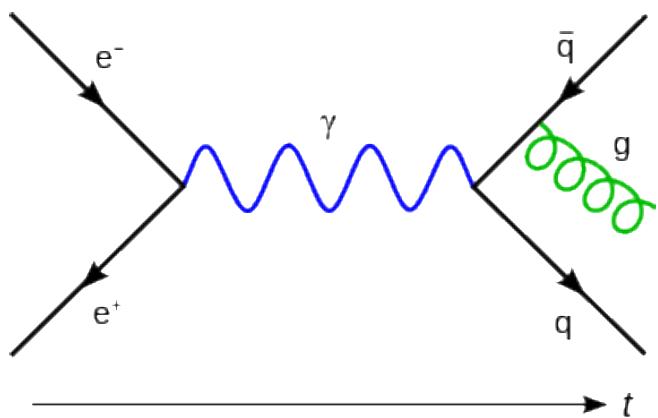
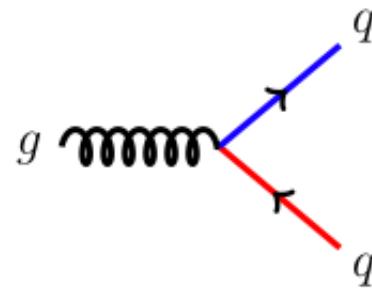
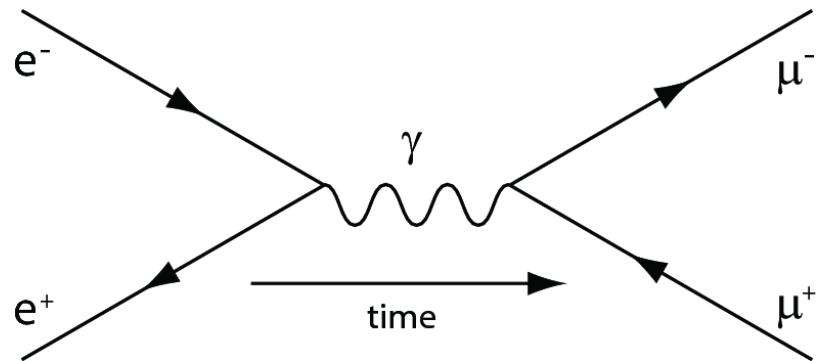


Očuvano (?) svojstvo: "Okus"

- Svakoj vrsti čestica se pripisuje okus:
 - "u,c,t,d,s,b" okus
 - Antičestica: suprotno
- Okus je očuvan u svim jakim i elektromagnetskim reakcijama:
 - Ako se stvara u-kvark, mora i anti-u

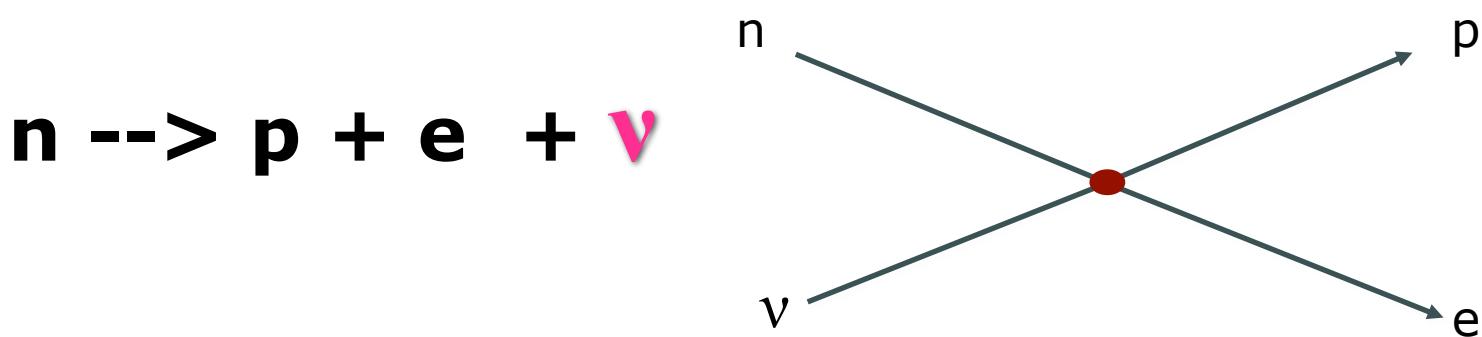


Očuvanje okusa u e.m. i jakim interakcijama



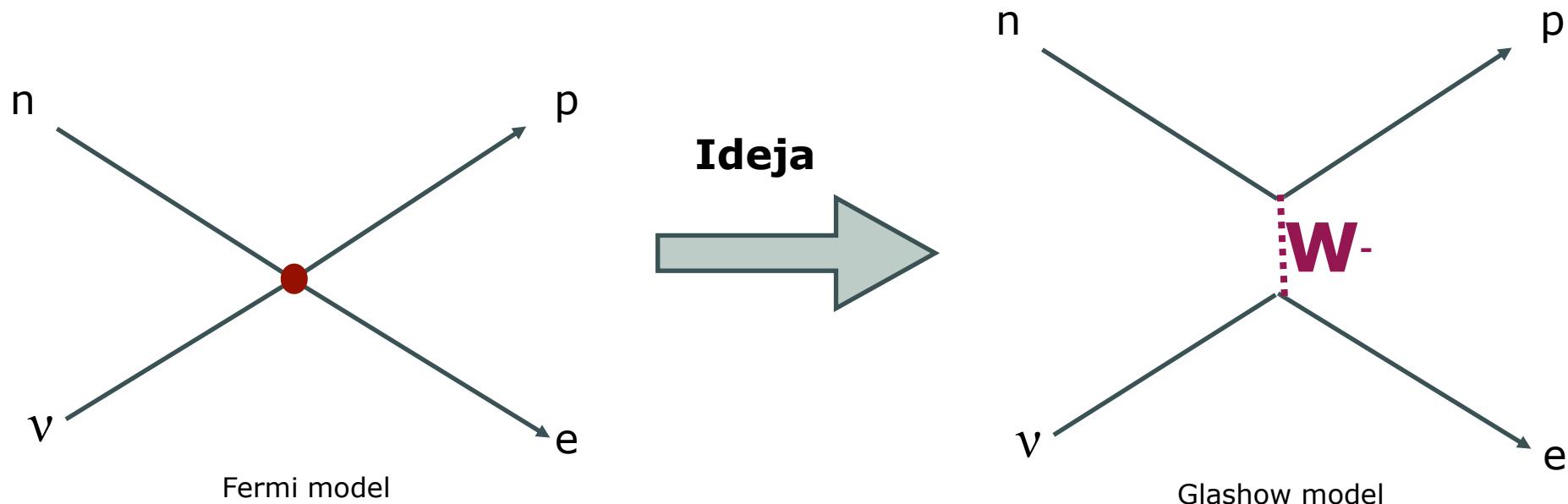
Slaba sila

- Druga vrsta sporijih procesa opažena:
 - Radioaktivni raspadi (β raspadi)
 - Interakcija neutrina
 - “Okus” (vrsta) kvarkova nije očuvan



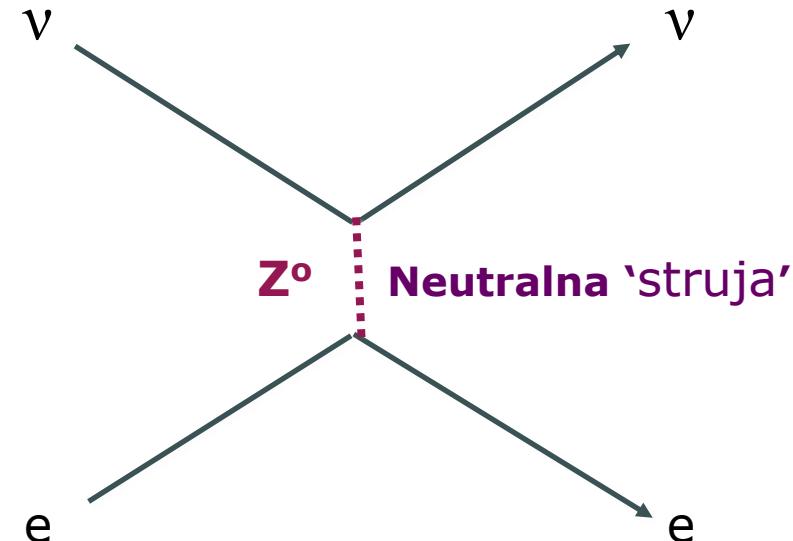
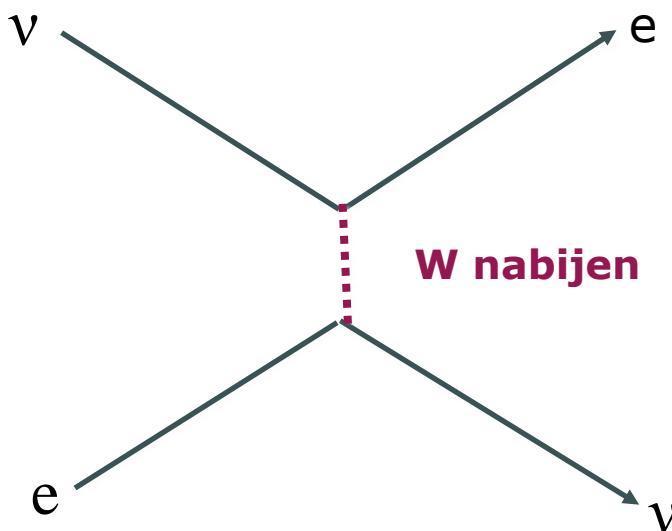
Nosioci slabe sile

Slabe interakcije se odvijaju putem izmjene teškog bozona



Velika masa (80 GeV) obašnjajva krati doseg i malu vjerojatnost
Interakcije za slabu silu

Elektroslaba sila



Glashow, Salam, Weinberg (1968)

Ujedinjenje elektromagnetse i slabe sile

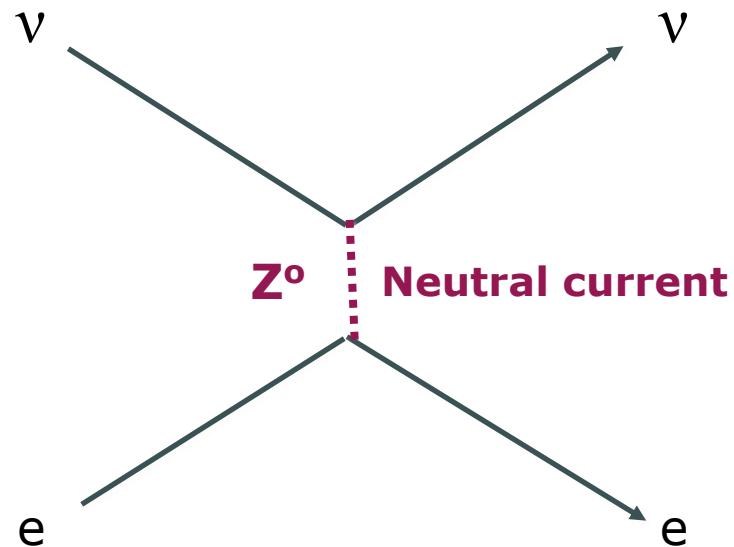
- e.m. i slaba aila su Aspekti iste “elektroslabe sile”
- I kvarkovi i leptoni imaju naboj salbe sile.
- **W i Z Bozoni dobivaju svoju masu od Higgsovog polja (1964)**

Opažanje elektroslabih interakcija

1973

Nuclear Physics B73 (1974) 1–22 North-Holland Publishing Company

Otkriće “neutralnih Struja”, CERN (1973)



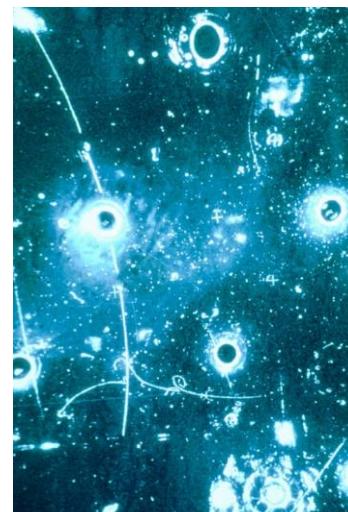
OBSERVATION OF NEUTRINO-LIKE INTERACTIONS WITHOUT MUON OR ELECTRON IN THE GARGAMELLE NEUTRINO EXPERIMENT

F.J. HASERT, S. KABE, W. KRENZ, J. VON KROGH, D. LANSKE,
J. MORFIN, K. SCHULTZE and H. WEERTS
III. Physikalisches Institut der Technischen Hochschule, Aachen, Germany

G. BERTRAND-COREMANS, J. SACTON, W. VAN DONINCK and P. VILAIN*
Interuniversity Institute for High Energies, U.I.B., V.U.B., Brussels, Belgium

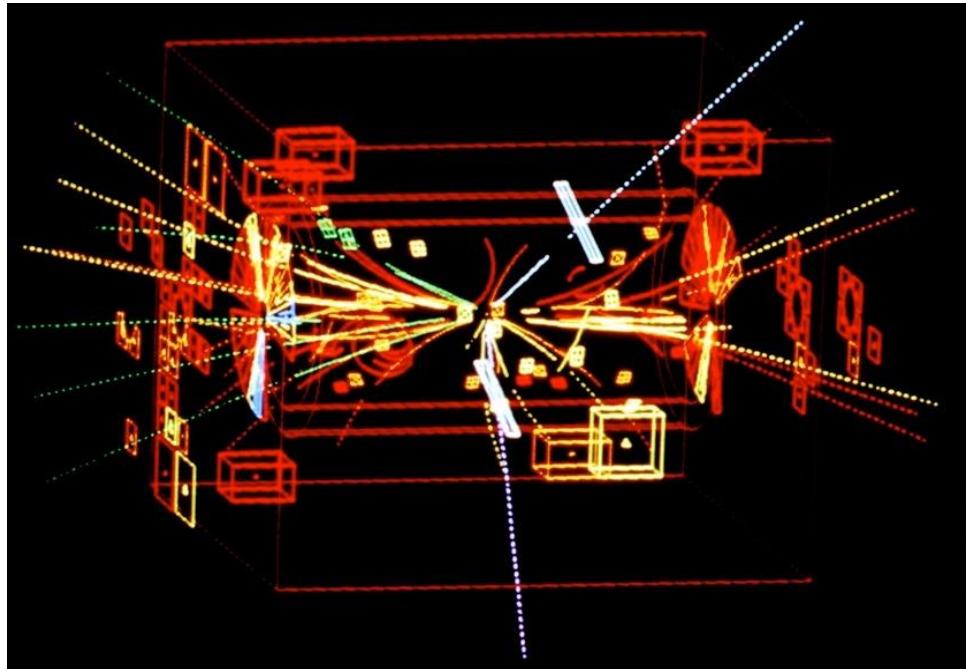
R. BALDI, U. CAMERINI**, D.C. CUNDY, I. DANILCHENKO***, W.F. FRY**
D. HAIDT, S. NATALI†, P. MUSSET, B. OSCULATI, R. PALMER††,
J.B.M. PATTISON, D.H. PERKINS*, A. PULLIA, A. ROUSSET,
W. VENUS†† and H. WACHSMUTH
CERN, Geneva, Switzerland

V. BRISSON, B. DEGRANGE, M. HAGUENAUER, L. KLUBERG,
U. NGUYEN-KHAC and P. PETIAU
Laboratoire de Physique Nucléaire des Hautes Energies, Ecole Polytechnique, Paris France



Otkriće W i Z bozona

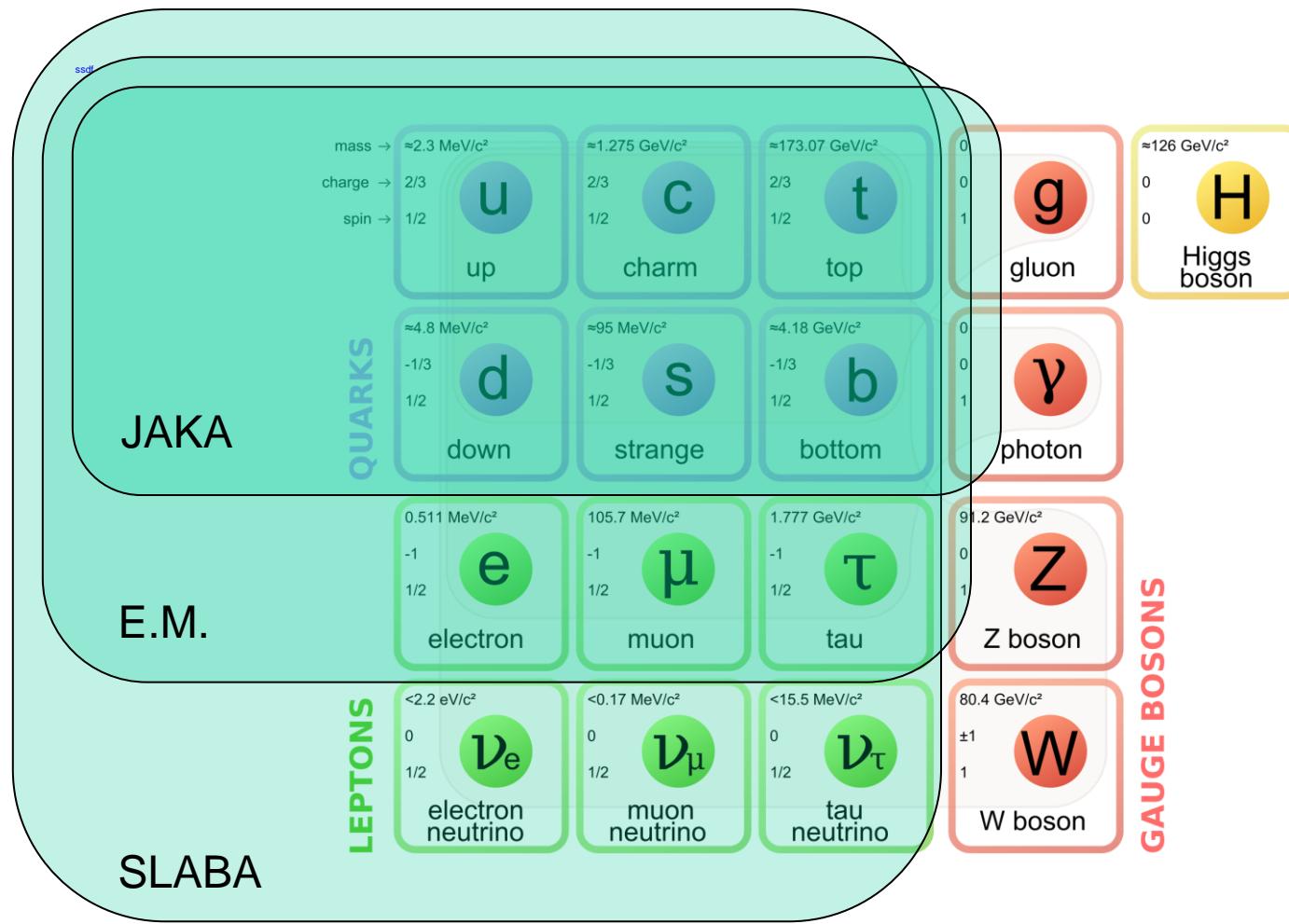
CERN (1983)



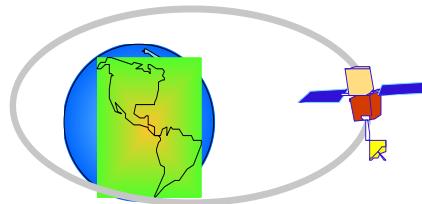
Carlo Rubbia, Simon van der Meer

W kandidat u detektoru UA-1

Čestice Standardnog modela i njihove interakcije



Ujedinjenje sila

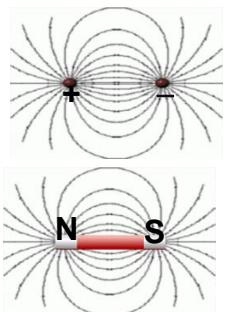


Zemaljska mehanika

Univerzalna gravitacija

Nebeska mehanika

Inercijalna vs. Gravitacijska masa
(I. Newton, 1687.)

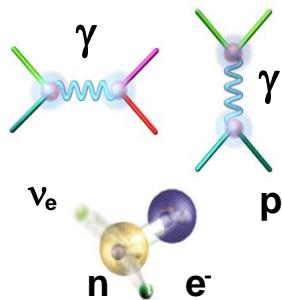


Elektricitet

Elektromagnetizam

Magnetizam

Elektromagnetski valovi (foton)
(J.C. Maxwell, 1860.)



Elektromagnetizam

Elektroslaba

Slaba sila

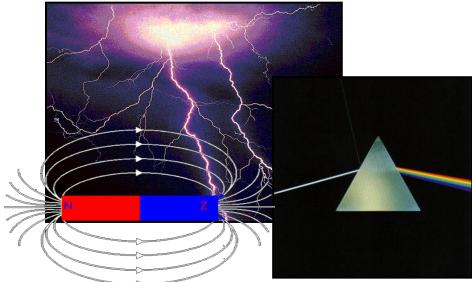
Intermedijalni bozoni W, Z
(1970.-83.)



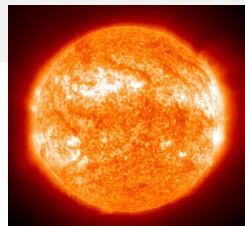
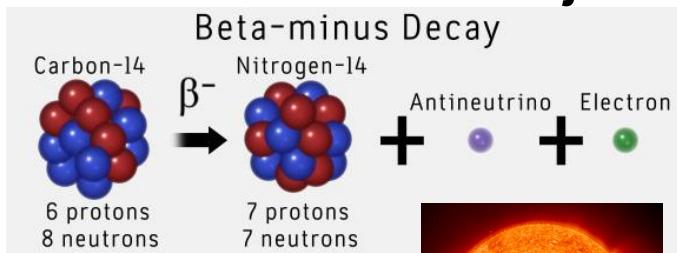
Ispitivanje sve manjih dimeznija
otkriva
dublje pravilnosti
UJEDINJENI OPIS

Elektroslabo ujedinjenje

Elektromagnetizam

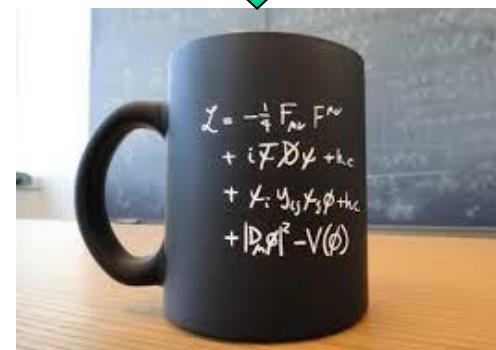
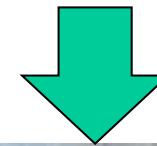


Slabe interakcije



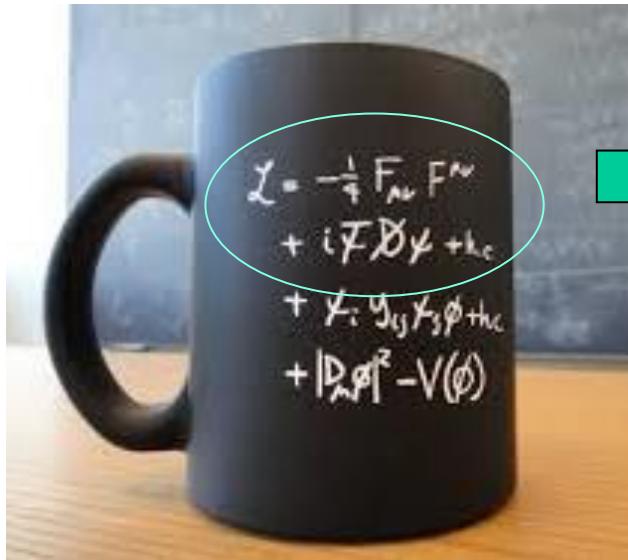
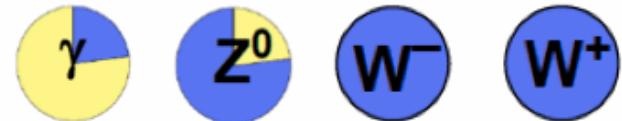
Elektroslaba teorija

- Zajednički izvor za elektromagnetsku i slabu silu
- Elektroslaba teorija temeljena na principu simetrije između 2 sile



Elektroslabi bozoni

- Nosioci elektroslabe sile



Elektroslaba teorija

- svi bozoni bezmaseni
- Sile \sim jednake

Opažanja

e.m. Sila

Foton

$m=0$

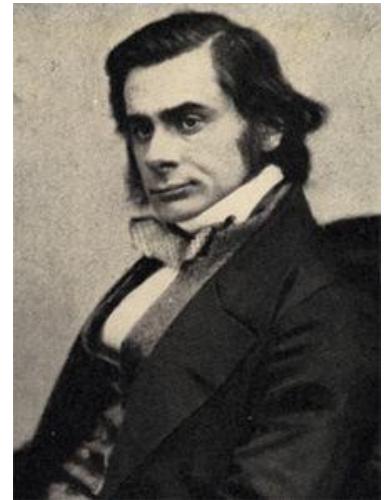
Slaba sila

W & Z bozoni: $m \sim 80-90$ GeV

Slaba sila $<<$ e.m sila

**Različite mase za fotona, W i Z:
lome simetriju i time i SM!**

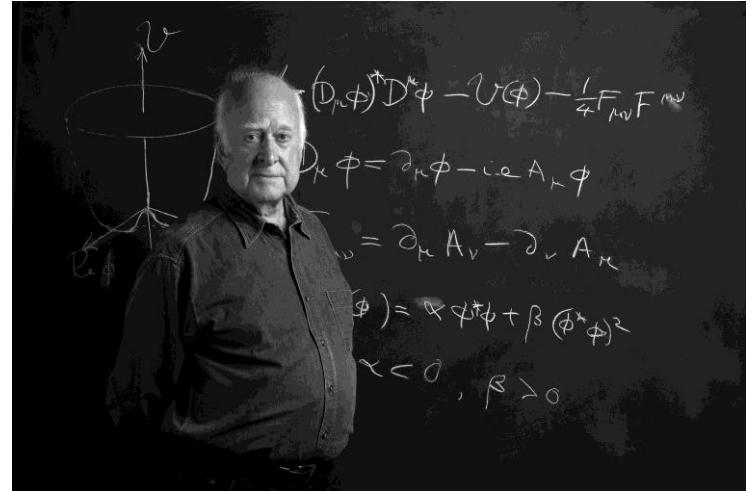
***"The great tragedy of science -
the slaying of a beautiful
hypothesis by an ugly fact."***



Thomas Henry Huxley

Ideja za spasiti SM

- Inovativne ideje za uvesti mase bozona u SM, tzv. "Higgsov mehanizam" dolaze sa više strana u 60im
[Brout-Englert, Higgs](#)
[Guralnik-Hagen-Kibble](#)
- Koristimo "Higgs" kao ime za cijelu skupinu (uključuje i Petera Higgsa)
- Polazimo od ničega: **Vakuum!**
 - Klasična fizika: Vakuum postoji u volumenu iz kojega se odstranila sva materija



U Vakuumu: Higgs!

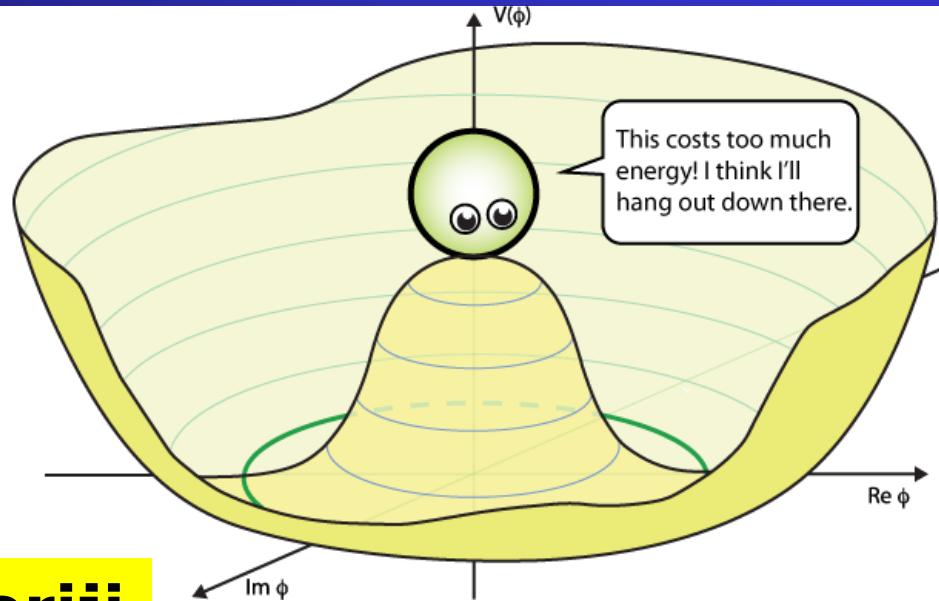
Osnovni postulat
“Higgsovog mehanizma”:

**Postoji novo polje, higgsovo
polje, koje ispunjava cijeli
svemir**

**Njegova je prosjećna vrijednost
postala različita od nule kad se
rani svemir ohladio**

Higgsov mehanizam

Spontani
lom simetrije



Higgsovo polje u teoriji

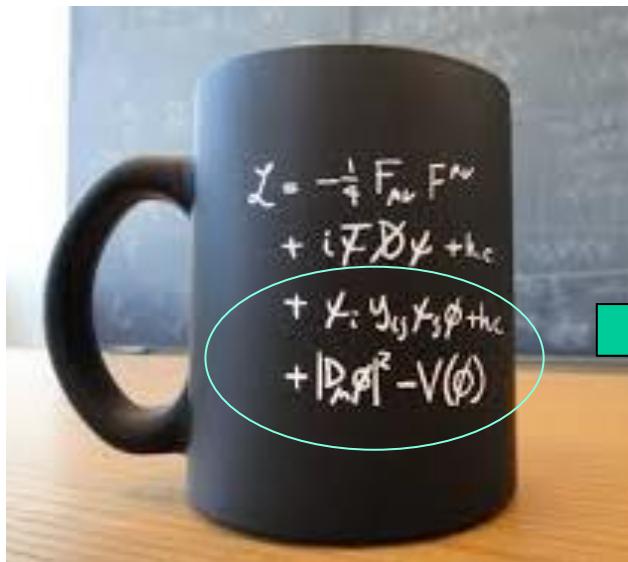
- Ispunjava cijeli prostor
- Nije 0 na minimumu energije
- Lomi simetriju u minimumu
- W i Z bozoni poprimaju masu kroz međudjelovanje s higgsovim poljem

Glashow-Salam-Weinberg Model

Elektroslabe interakcije

+

Higgsov mehanizam

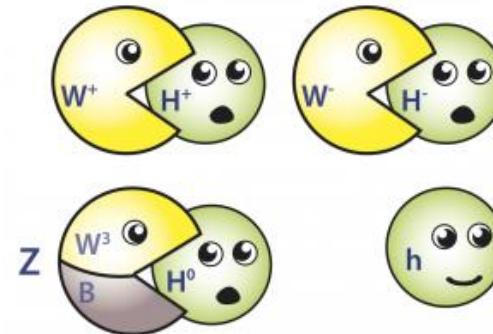


+

Teorija jake interakcije
(QCD)

= Standardni Model

- Simetrija (iako skrivena) spašena
- W i Z bozon dobivaju masu od higgsovog polja



- Fermioni također dobivaju masu kroz međudjelovanje s higgsovim poljem

Higgsov mehanizam

Interakcija s
Higgsovim poljem



Frikcija u
viskoznoj tekučini



Higgsov mehanizam

Interakcija s
Higgsovim poljem

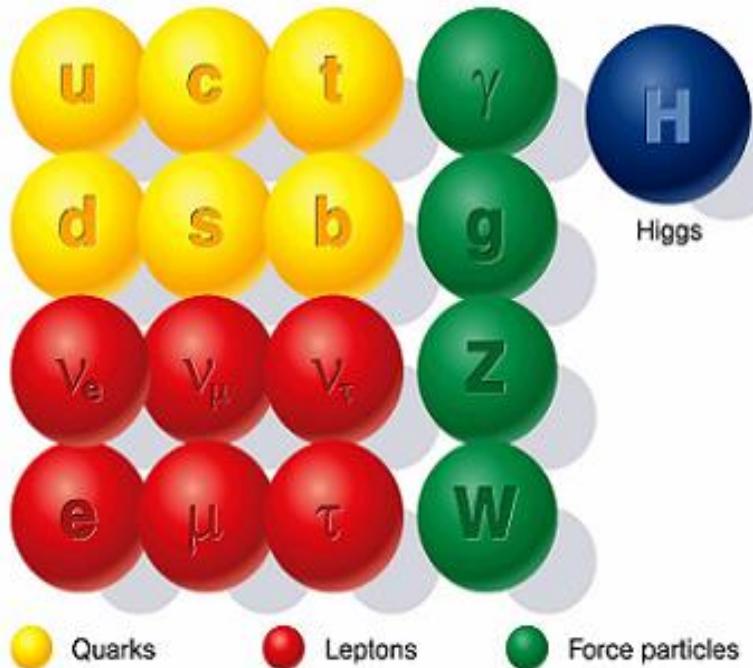


Trenje u viskoznoj
tekućini



Standard Model of Elementary Particles

Standard particles

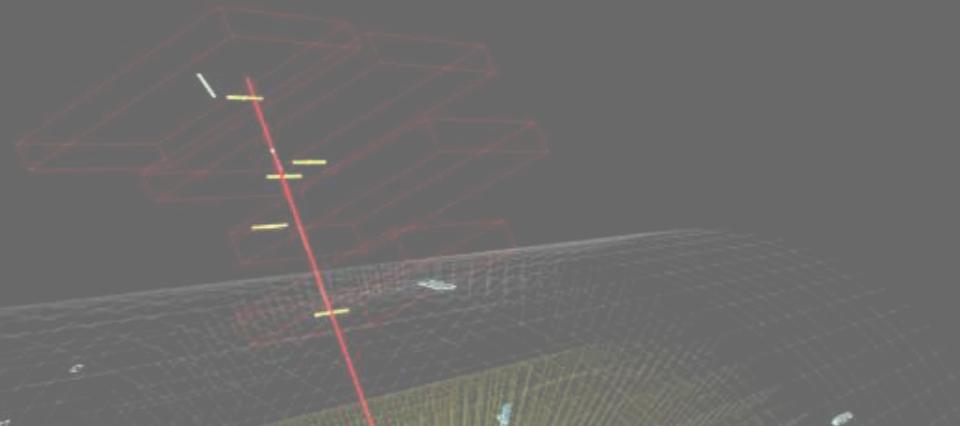


$$\begin{aligned}
 \mathcal{L}_{GWS} = & \sum_f (\bar{\Psi}_f (i\gamma^\mu \partial_\mu - m_f) \Psi_f - e Q_f \bar{\Psi}_f \gamma^\mu \Psi_f A_\mu) + \\
 & \gamma (\bar{a}_L^i \gamma^\mu b_L^i W_\mu^+ + \bar{b}_L^i \gamma^\mu a_L^i W_\mu^-) + \frac{g}{2c_w} \sum_f \bar{\Psi}_f \gamma^\mu (I_f^3 - 2s_w^2 Q_f - I_f^3 \gamma_5) \Psi_f \bar{Z} \\
 & |\partial_\mu A_\nu - \partial_\nu A_\mu - ie(W_\mu^- W_\nu^+ - W_\mu^+ W_\nu^-)|^2 - \frac{1}{2} |\partial_\mu W_\nu^+ - \partial_\nu W_\mu^+ + \\
 & - ie(W_\mu^+ A_\nu - W_\nu^+ A_\mu) + ig' c_w (W_\mu^+ Z_\nu - W_\nu^+ Z_\mu)|^2 + \\
 & - \frac{1}{4} |\partial_\mu Z_\nu - \partial_\nu Z_\mu + ig' c_w (W_\mu^- W_\nu^+ - W_\mu^+ W_\nu^-)|^2 + \\
 & - \frac{1}{2} M_\eta^2 \eta^2 - \frac{g M_\eta^2}{8M_W} \eta^3 - \frac{g'^2 M_\eta^2}{32M_W} \eta^4 + |M_W W_\mu^+ + \frac{g}{2} \eta W_\mu^+|^2 + \\
 & + \frac{1}{2} |\partial_\mu \eta + i M_Z Z_\mu + \frac{ig}{2c_w} \eta Z_\mu|^2 - \sum_f \frac{g}{2} \frac{m_f}{M_W} \bar{\Psi}_f \Psi_f \eta
 \end{aligned}$$

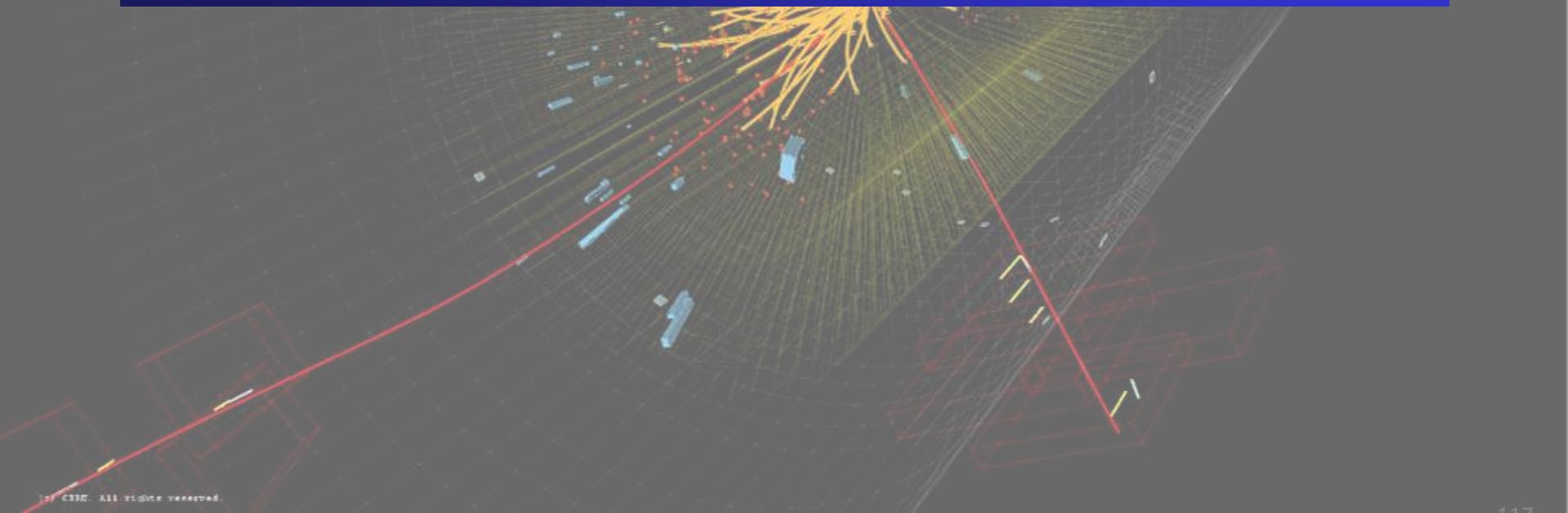


**Jedini sigurno opaženi
Higgs u CMS
eksperimentu do 2012.**

Dr. Peter Higgs



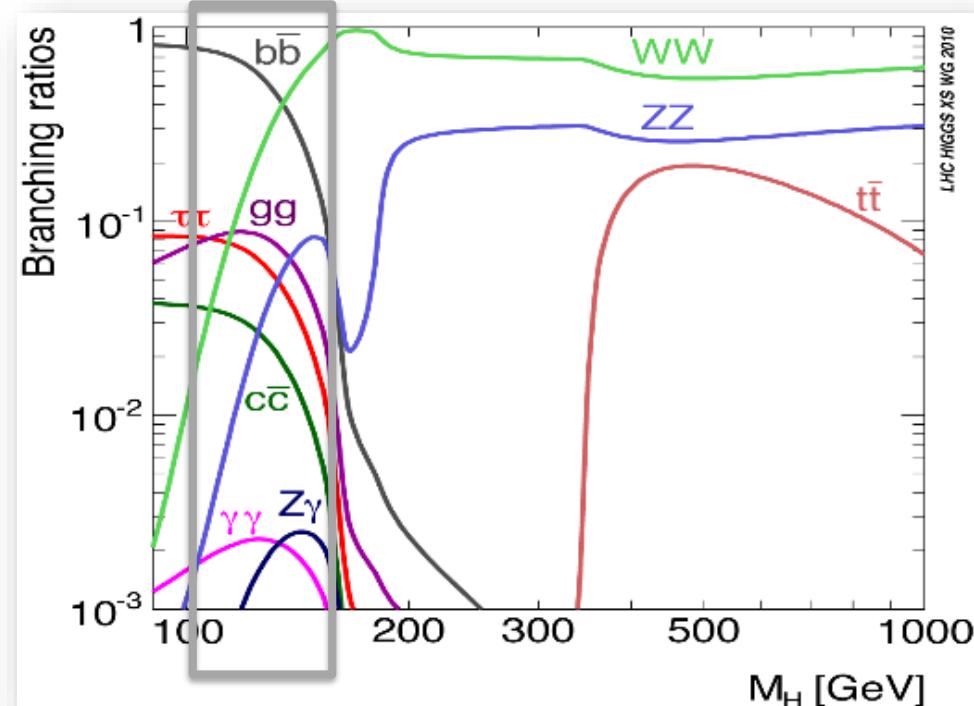
Potraga za Higgsovim bozonom



Kako se raspada Higgsov bozon?

Promatrani kanali raspada

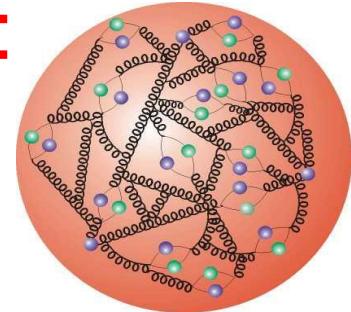
- Visoke mase: WW , ZZ
- Niske mase: bb , $\tau\tau$, WW , ZZ , $\gamma\gamma$
- Područje niskih masa je jako bogato, ali i teško:
Teško identificirati glavne kanale raspada (bb , $\tau\tau$) zbog velike pozadine
- Najprecizniji kanali: $H \rightarrow \gamma\gamma$ & $H \rightarrow ZZ \rightarrow 4l$



Analiza sudara: komplicirana!

Hadronski sudarivač
(sudar ~50 točkastih čestica)

Proton:

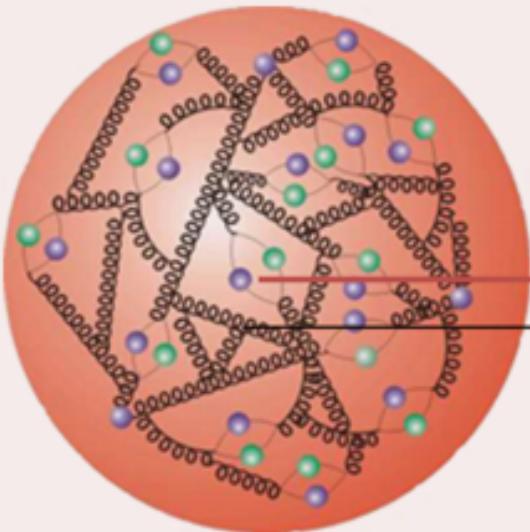


Gdje je Asterix?

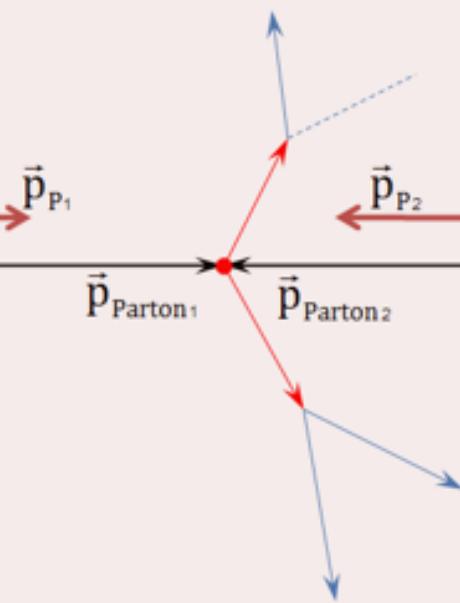
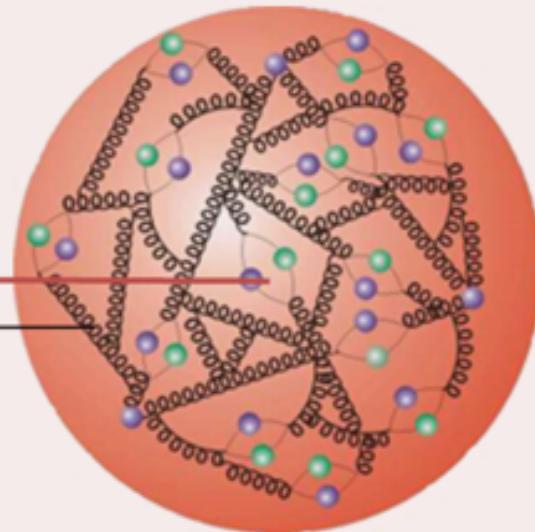


Interactions of constituents of the colliding protons, the so called partons (quarks, gluons)

proton 1



proton 2



\vec{p}_{P_1} ... momentum proton 1

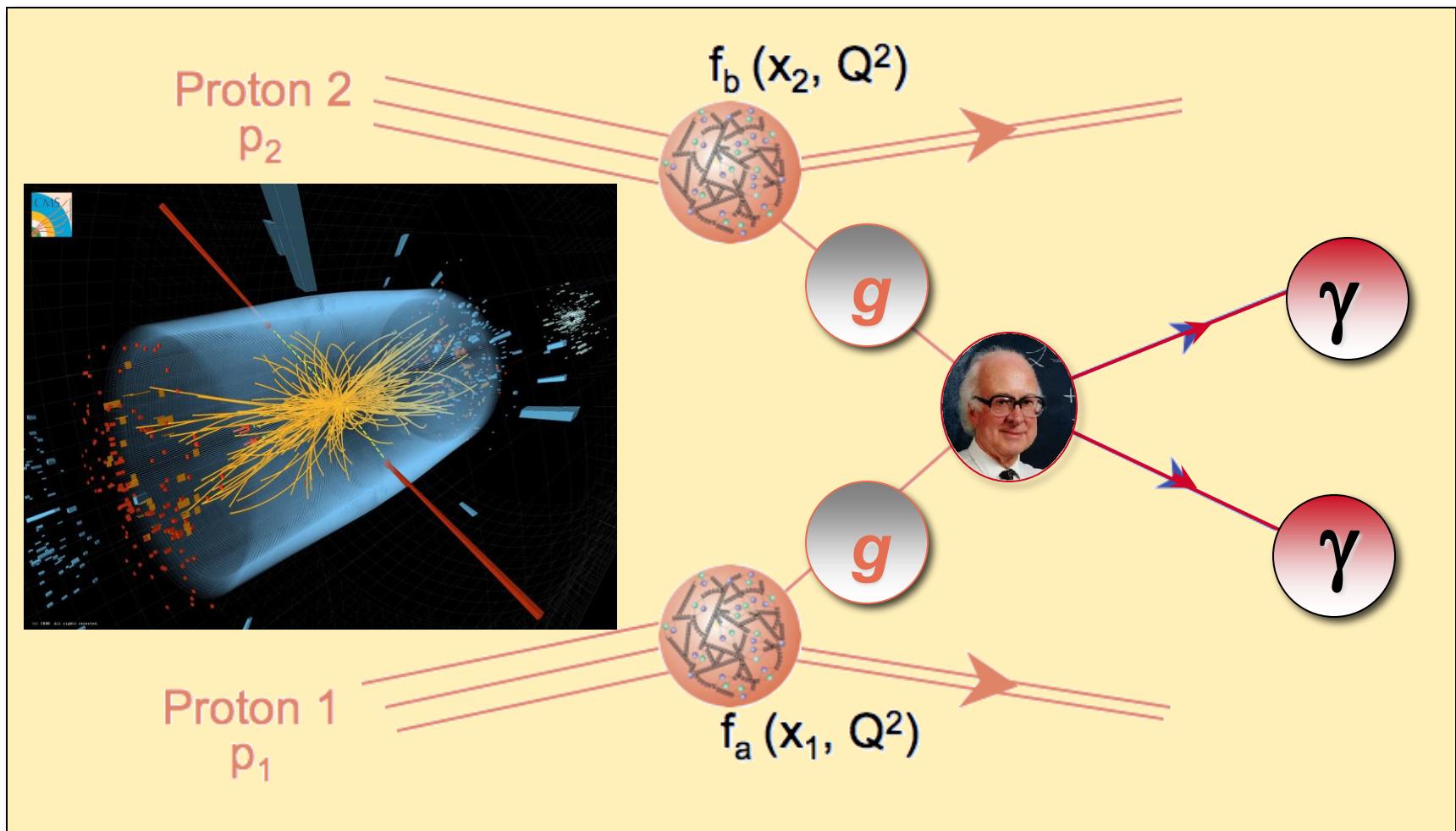
\vec{p}_{P_2} ... momentum proton 2

- interaction vertex

$\vec{p}_{\text{Parton}_1}$... momentum parton 1

$\vec{p}_{\text{Parton}_2}$... momentum parton 2

Osnovni procesi na LHC-u

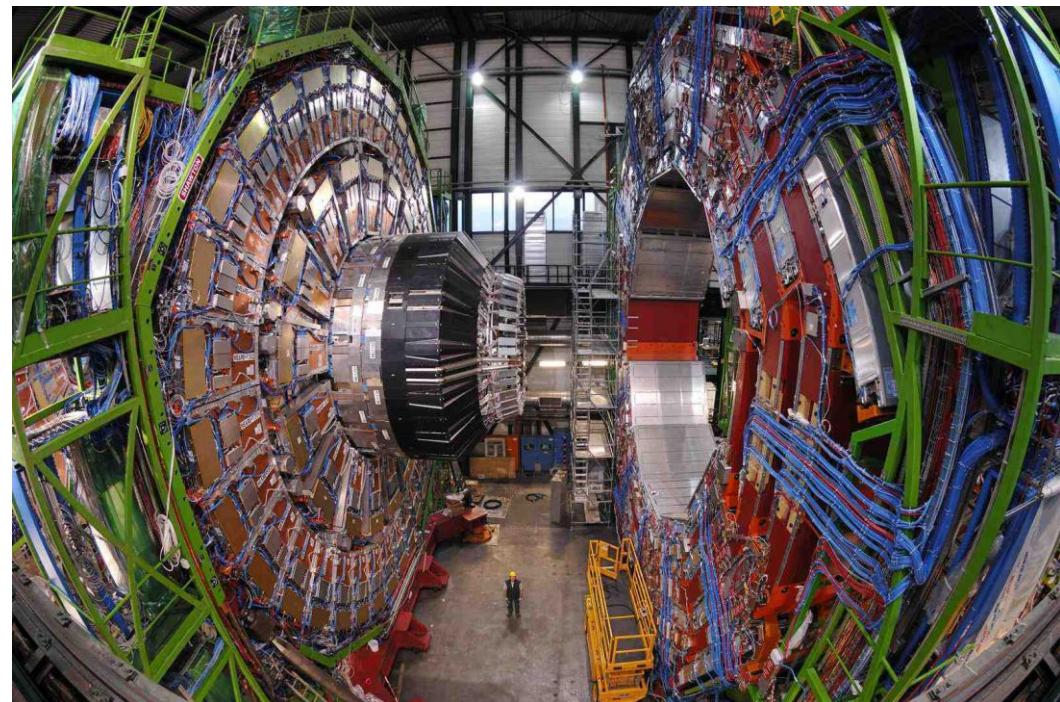


Kako vidimo

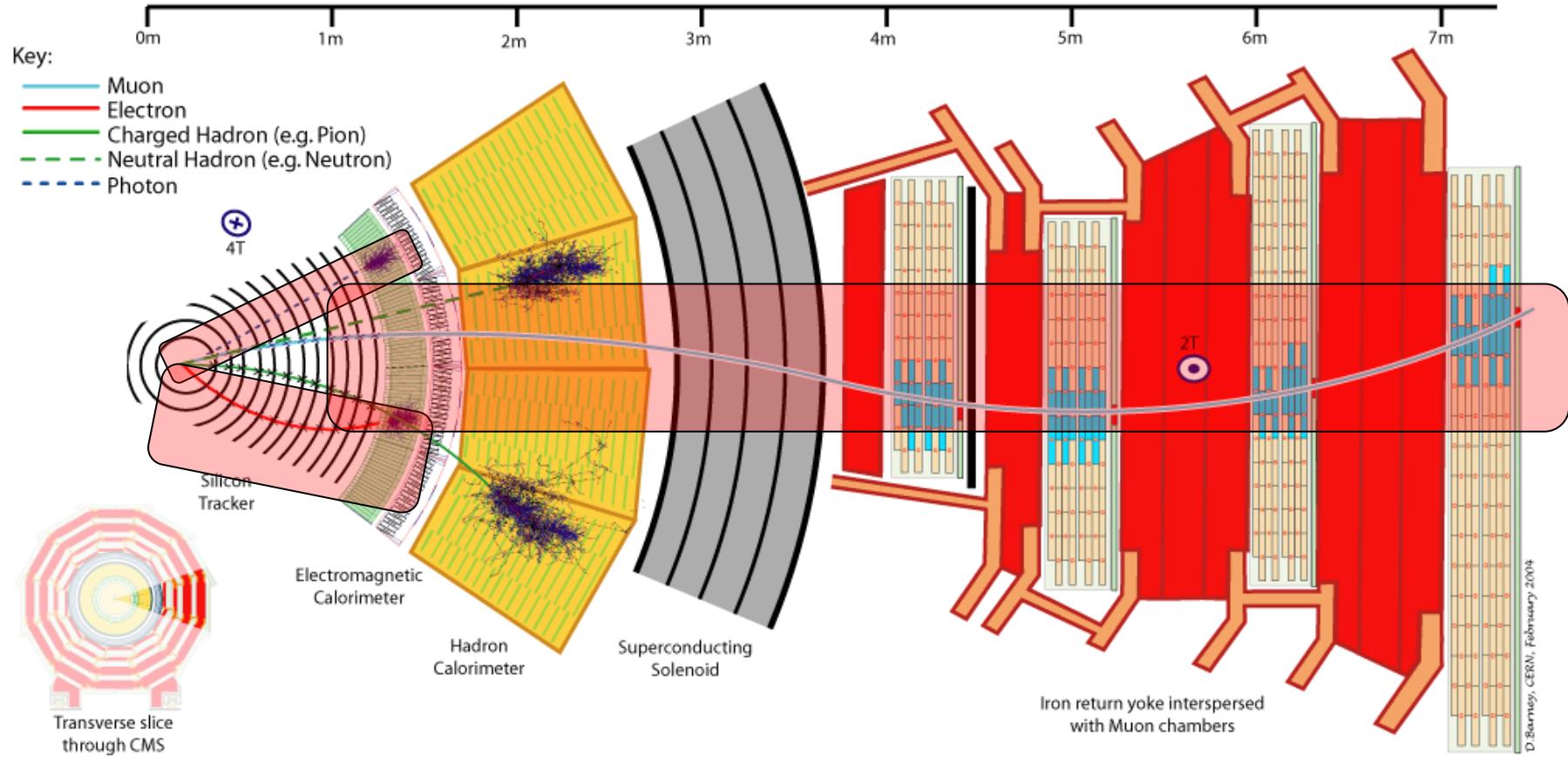
STANDARD MODEL OF ELEMENTARY PARTICLES

QUARKS	UP	CHARM	TOP	GLUON	HIGGS BOSON
	mass 2,3 MeV/c ²	1,275 GeV/c ²	173,07 GeV/c ²	0 0 1	126 GeV/c ² 0 0
	charge 2/3 spin 1/2	2/3 1/2	2/3 1/2	g	H
	u	c	t		
	DOWN	STRANGE	BOTTOM	PHOTON	
	4,8 MeV/c ² -1/3 1/2	95 MeV/c ² -1/3 1/2	4,18 GeV/c ² -1/3 1/2	0 0 1	
	d	s	b	γ	
	ELECTRON	MUON	TAU	Z BOSON	
	0,511 MeV/c ² -1 1/2	105,7 MeV/c ² -1 1/2	1,777 GeV/c ² -1 1/2	91,2 GeV/c ² 0 1	
	e	μ	τ	Z	
LEPTONS	ELECTRON NEUTRINO	MUON NEUTRINO	TAU NEUTRINO	W BOSON	
	<2,2 eV/c ² 0 1/2	<0,17 MeV/c ² 0 1/2	<15,5 MeV/c ² 0 1/2	80,4 GeV/c ² ±1 1	
	ν _e	ν _μ	ν _τ	W	
	Gauge Bosons				

u

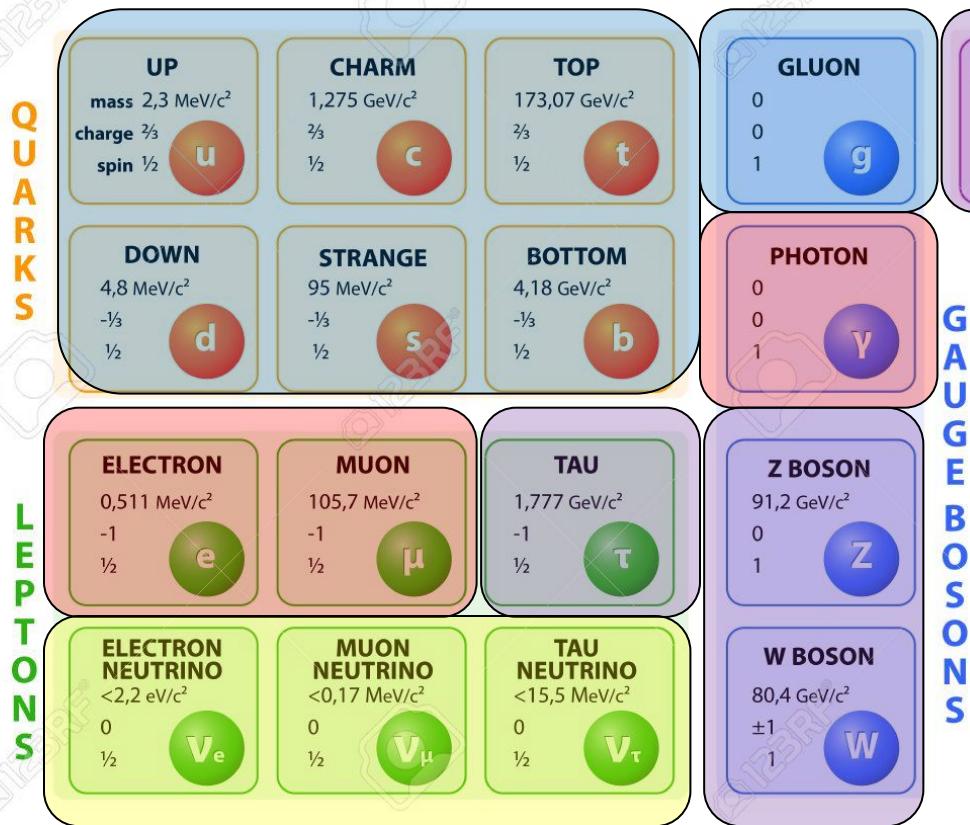


Elektron, foton i mion u CMS detektoru



Čestice Standardnog modela

STANDARD MODEL OF ELEMENTARY PARTICLES



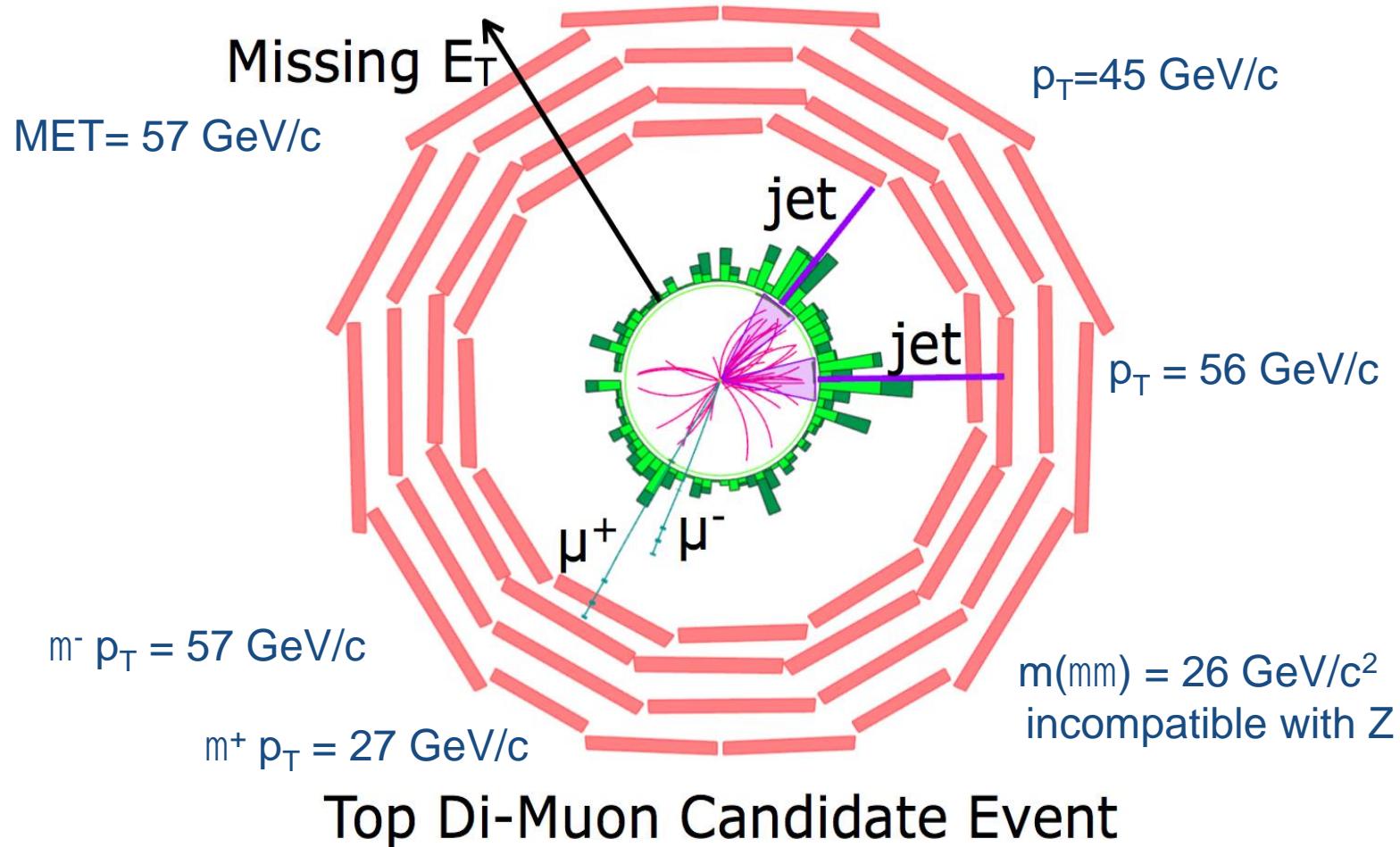
Vidimo ih
u detektoru

Vidimo ih
Grupirane
u hadronima

Vidimo produkte
Njihovog raspada

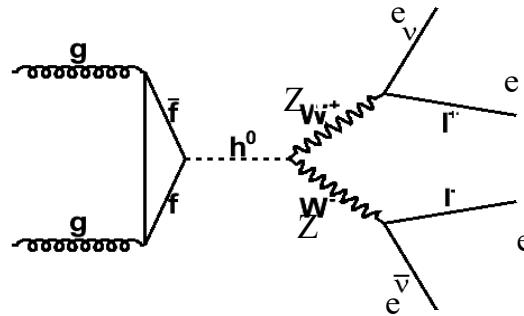
Ne vidimo ih

Fizikalni objekti u CMS detektoru

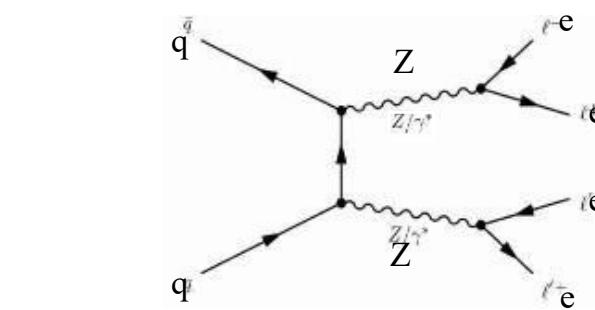


Signal i pozadina

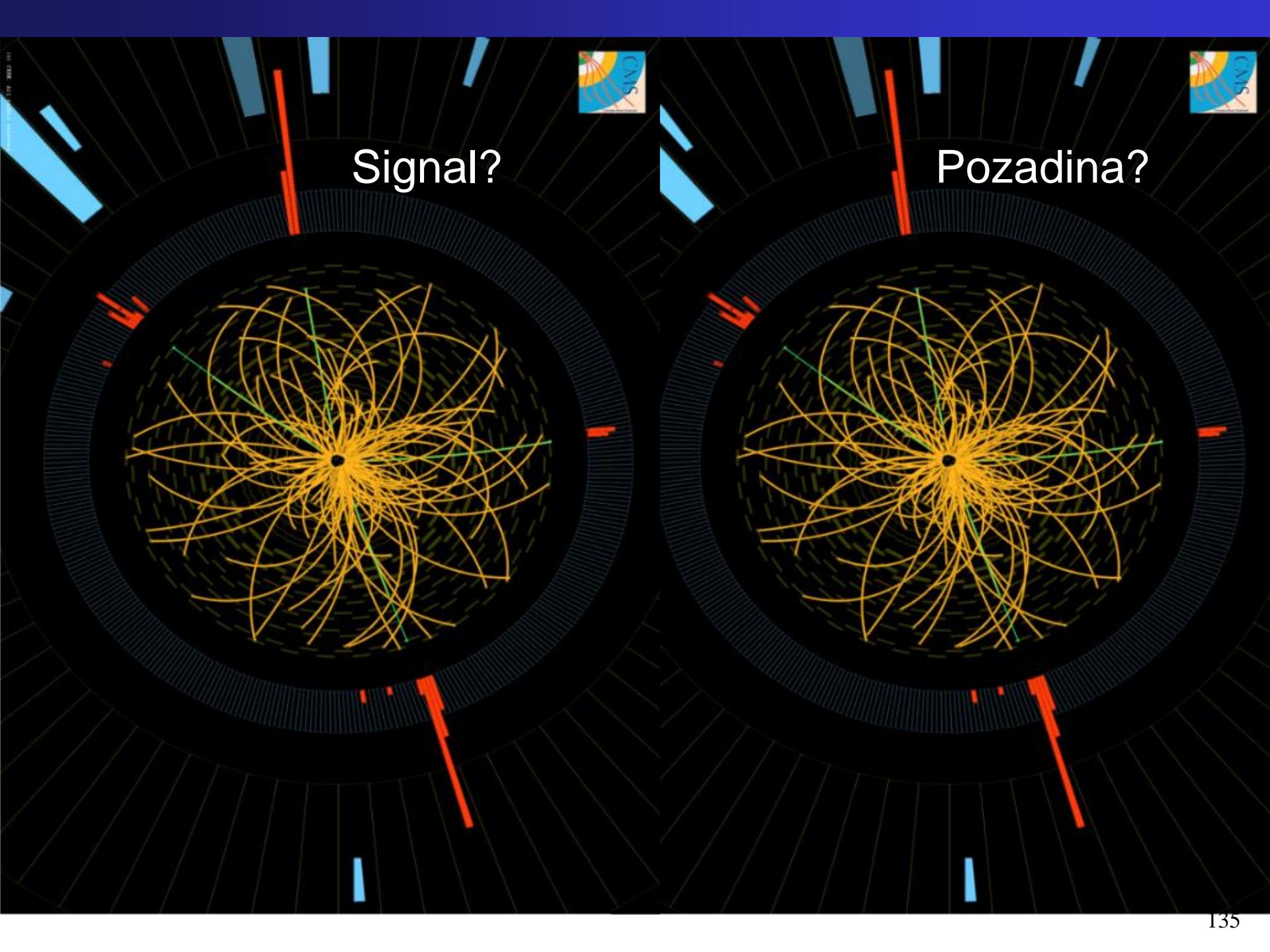
- **Signal:** Događaj koji dolazi iz traženog procesa
 - Example: $H \rightarrow ZZ \rightarrow e^+e^-e^+e^-$
- **Pozadina:** sve ostalo
 - Opasno: može izgledati upravo kao signal



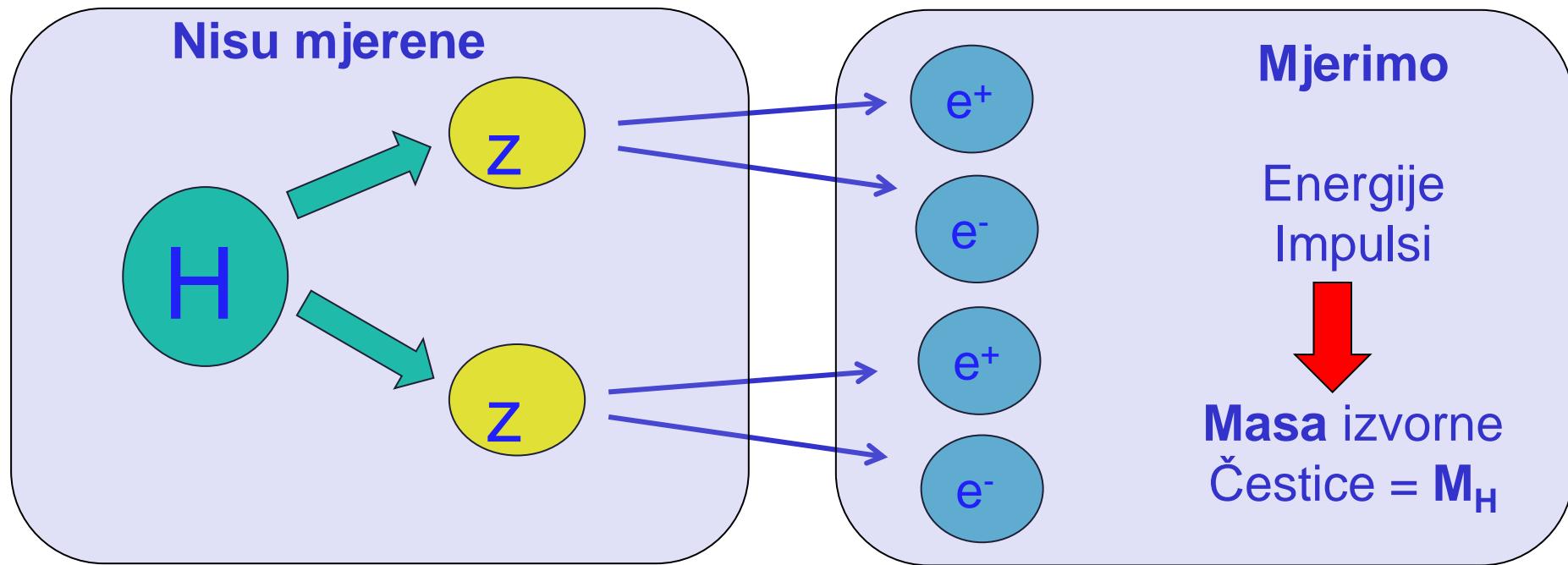
Signal: $pp \rightarrow H \rightarrow ZZ \rightarrow 4e$



'Opasna' pozadina: $pp \rightarrow ZZ \rightarrow 4e$



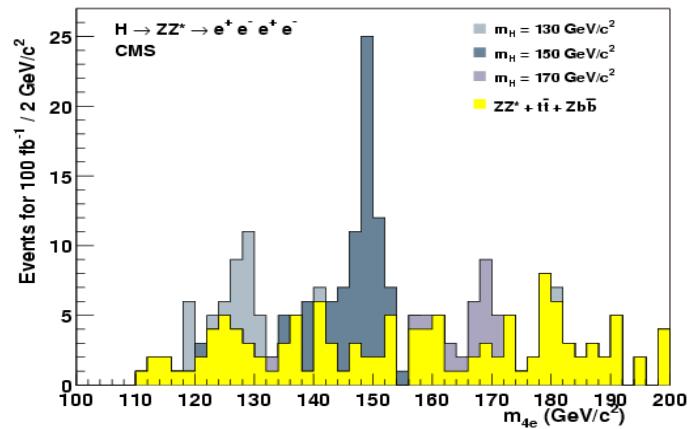
Kako identificirati signal



- “Invarijantna masa” produkta raspada:
 - Masa Higgsovog bozona za signal
 - Proizvoljna vrijednost za pozadinu

Potraga za raspadom $H \rightarrow ZZ^ \rightarrow 4l$ ($l=e,\mu$)*

"Zlatni kanal"



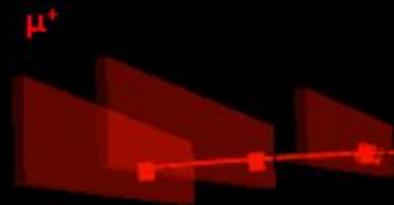


4 leptons candidate : 4 μ

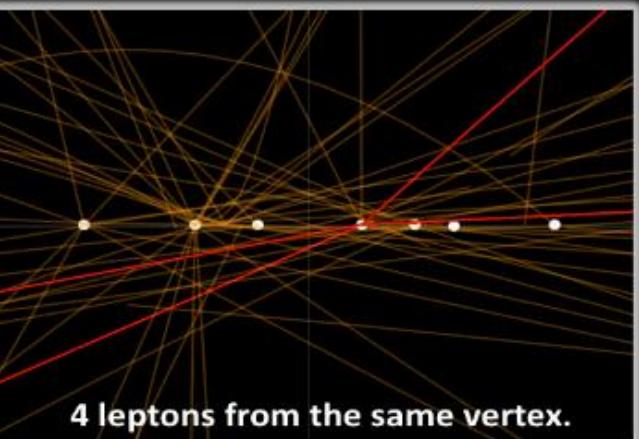
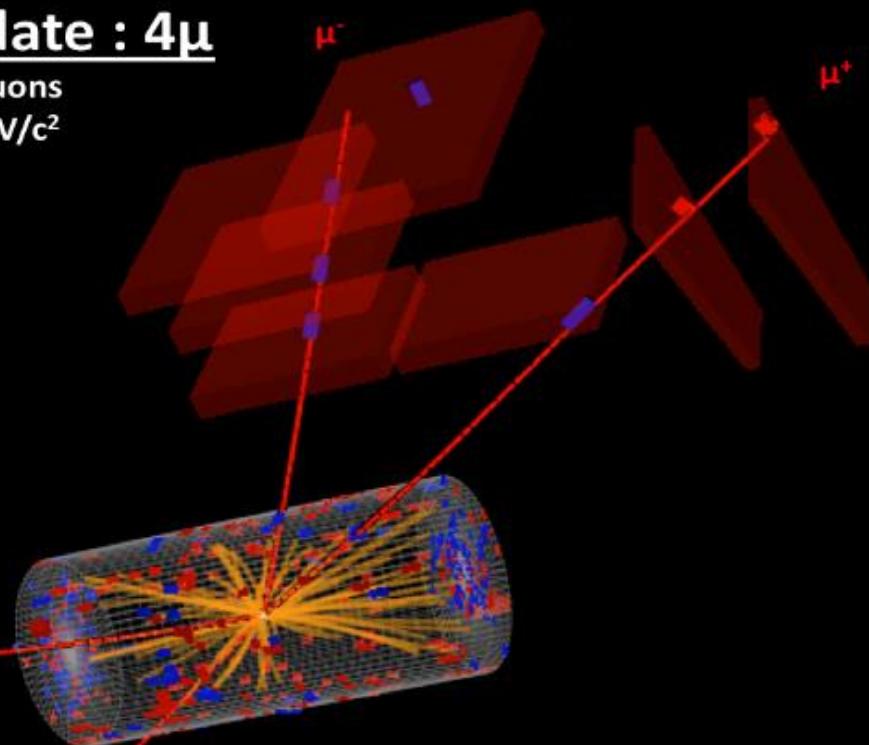
4 isolated muons

$M_{4l} = 145 \text{ GeV}/c^2$

CMS Experiment at LHC, CERN
Data recorded: Mon May 2 07:05:01 2011 CEST
Run/Event: 163817 / 155679852
Lumi section: 174
Orbit/Crossing: 45568654 / 469



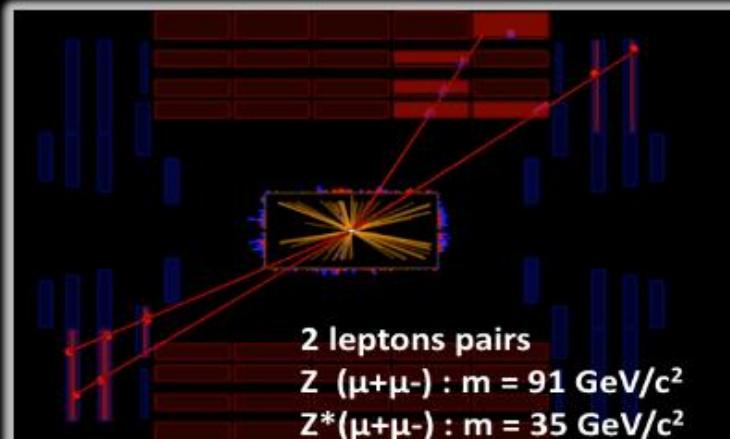
μ^+



4 leptons from the same vertex.



μ^-



2 leptons pairs

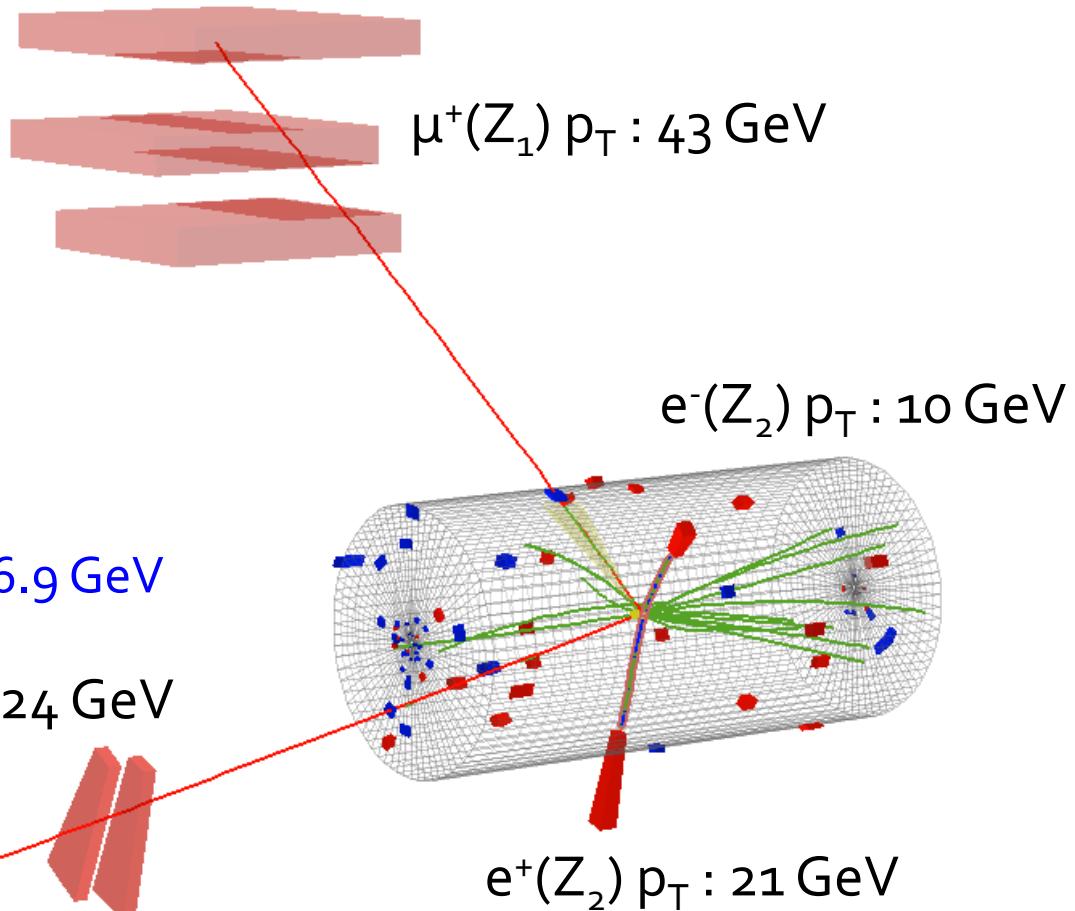
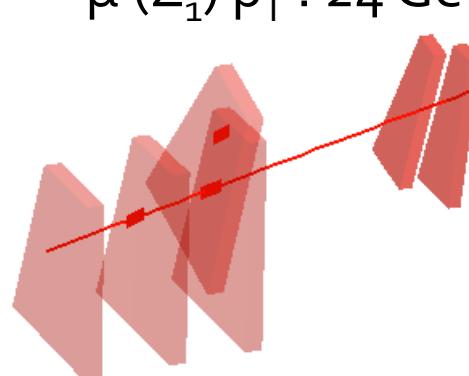
$Z (\mu^+\mu^-) : m = 91 \text{ GeV}/c^2$

$Z^*(\mu^+\mu^-) : m = 35 \text{ GeV}/c^2$



8 TeV DATA

4-lepton Mass : 126.9 GeV



CMS Experiment at LHC, CERN
Data recorded: Mon May 28 01:35:47 2012 CEST
Run/Event: 195099 / 137440354
Lumi section: 115

CMS Experiment at LHC, CERN
Data recorded: Thu Oct 13 03:39:46 2011 CEST
Run/Event: 178421 / 87514902
Lumi section: 86



γ TeV DATA

$4\mu + \gamma$ Mass : 126.1 GeV

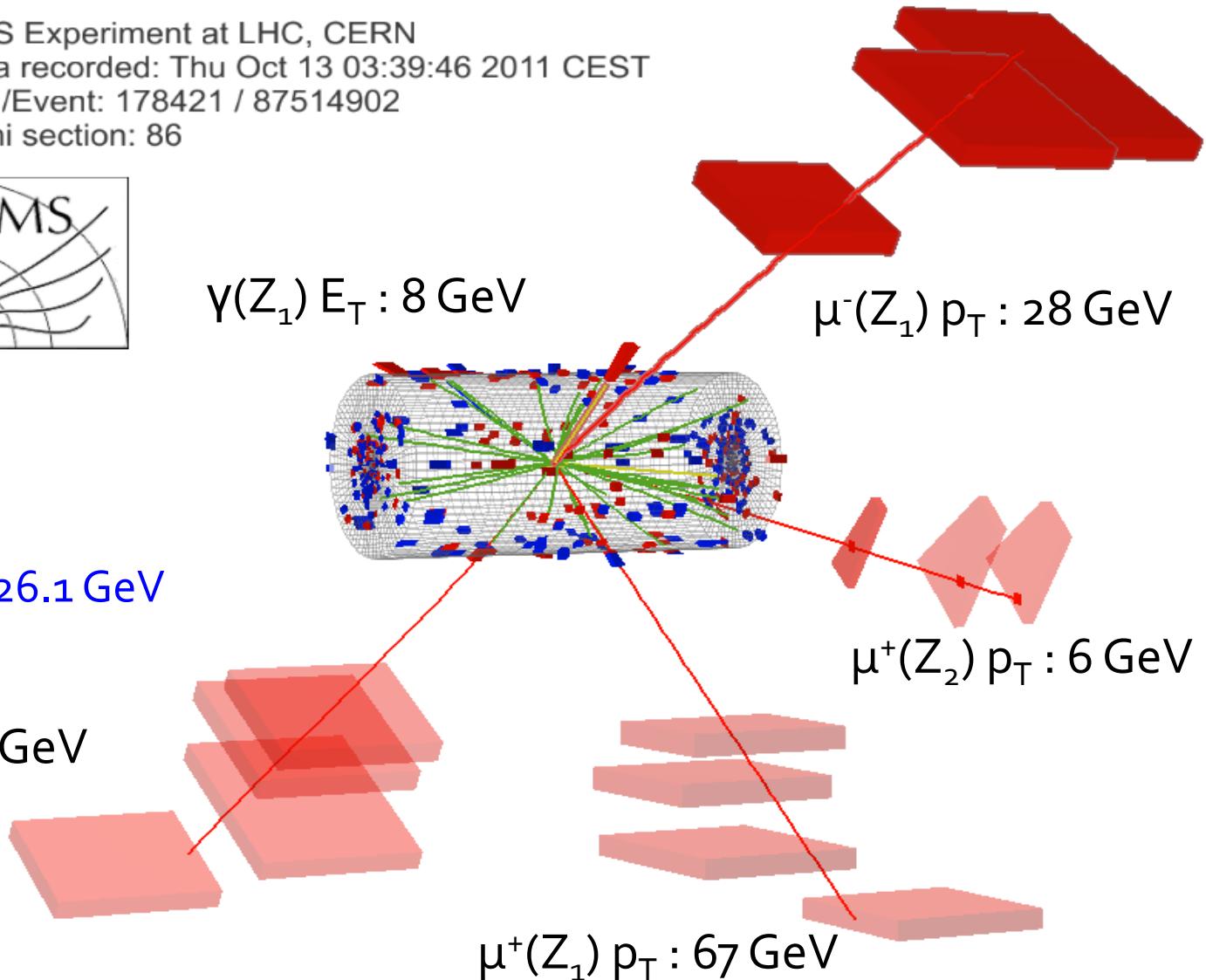
$\mu^-(Z_2) p_T : 14$ GeV

$\gamma(Z_1) E_T : 8$ GeV

$\mu^-(Z_1) p_T : 28$ GeV

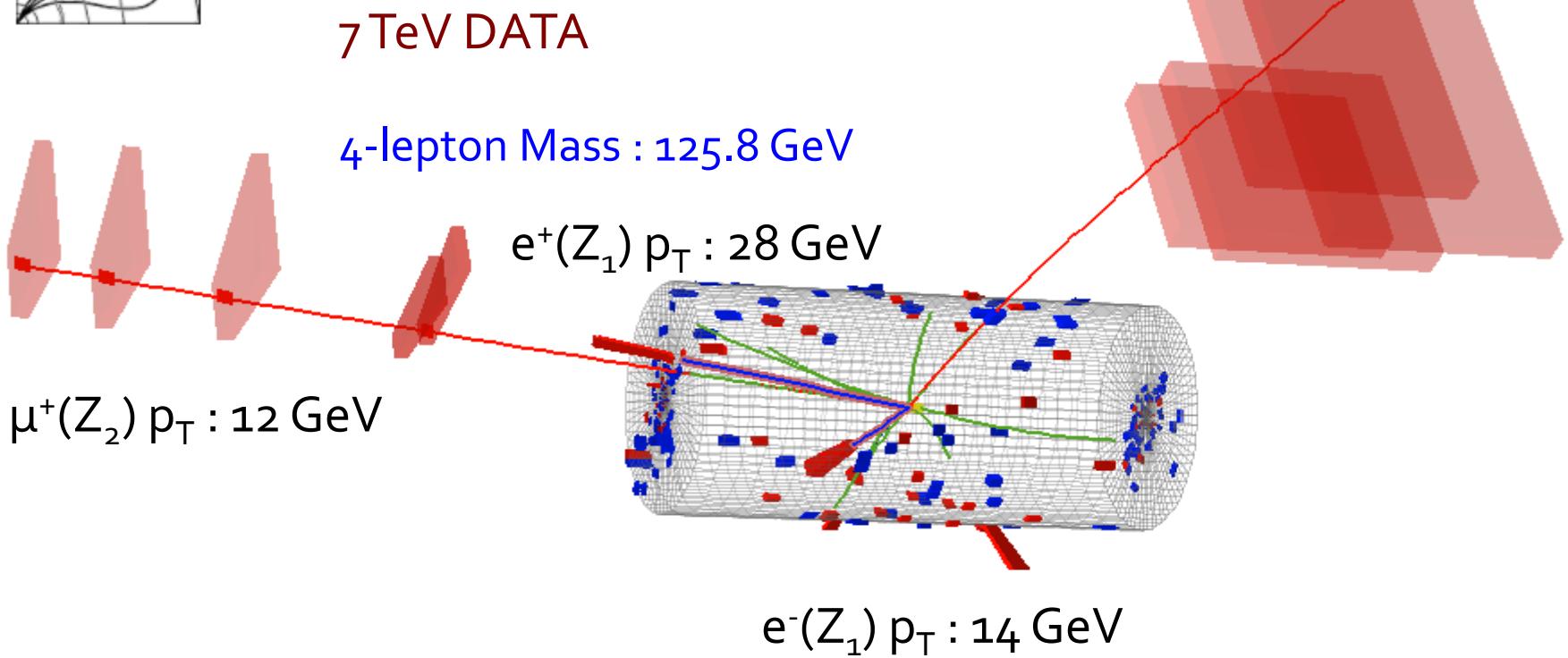
$\mu^+(Z_2) p_T : 6$ GeV

$\mu^+(Z_1) p_T : 67$ GeV



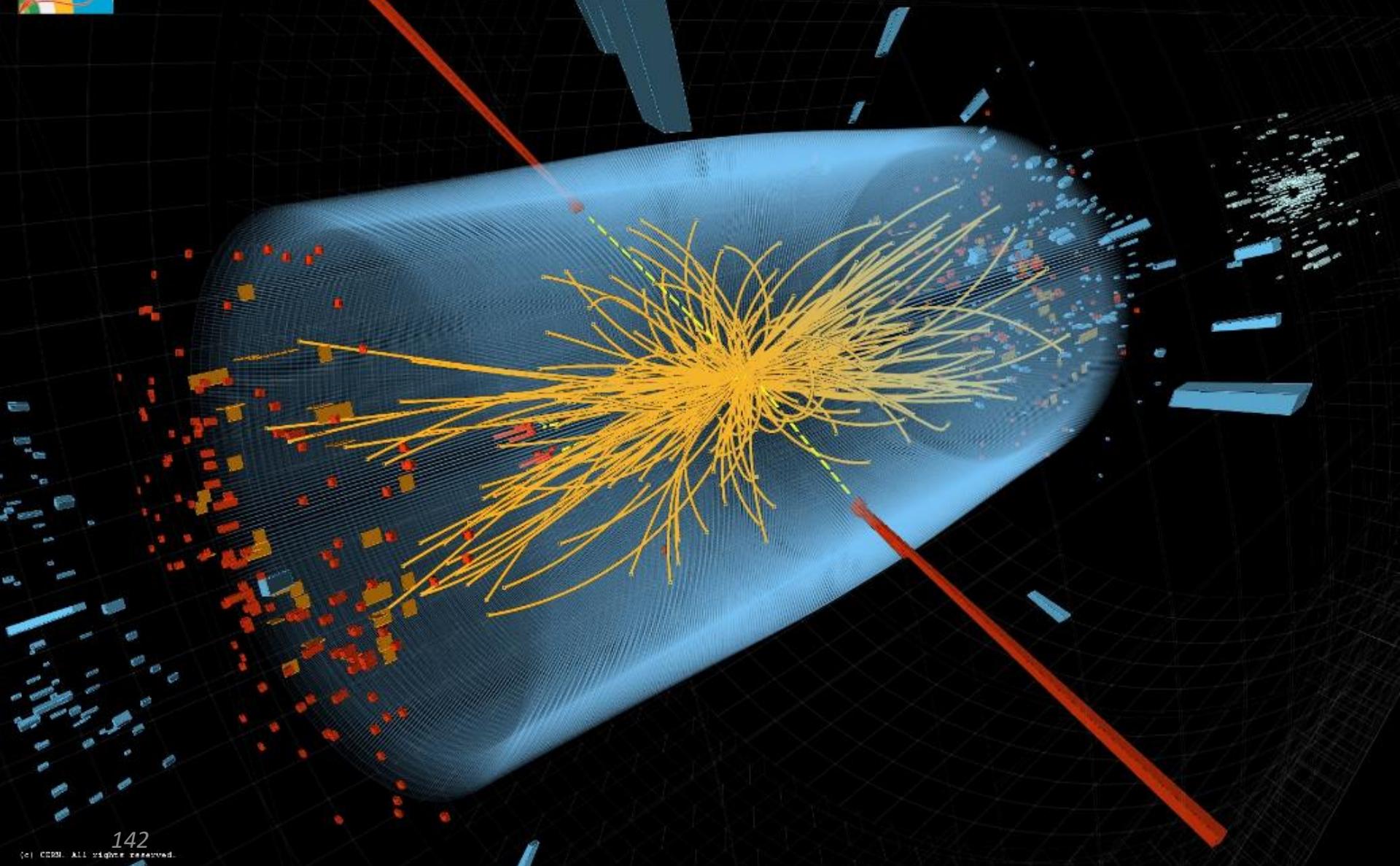


CMS Experiment at LHC, CERN
Data recorded: Tue Oct 4 00:10:13 2011 CEST
Run/Event: 177782 / 72158025
Lumi section: 99

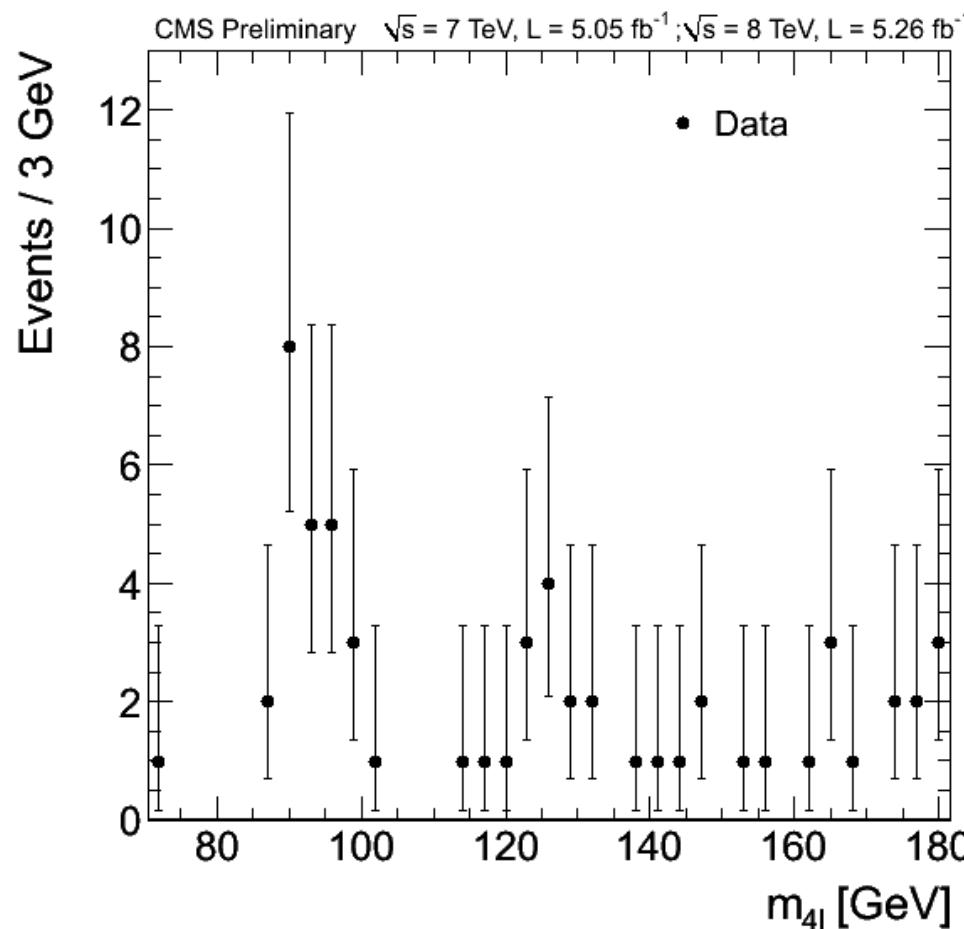




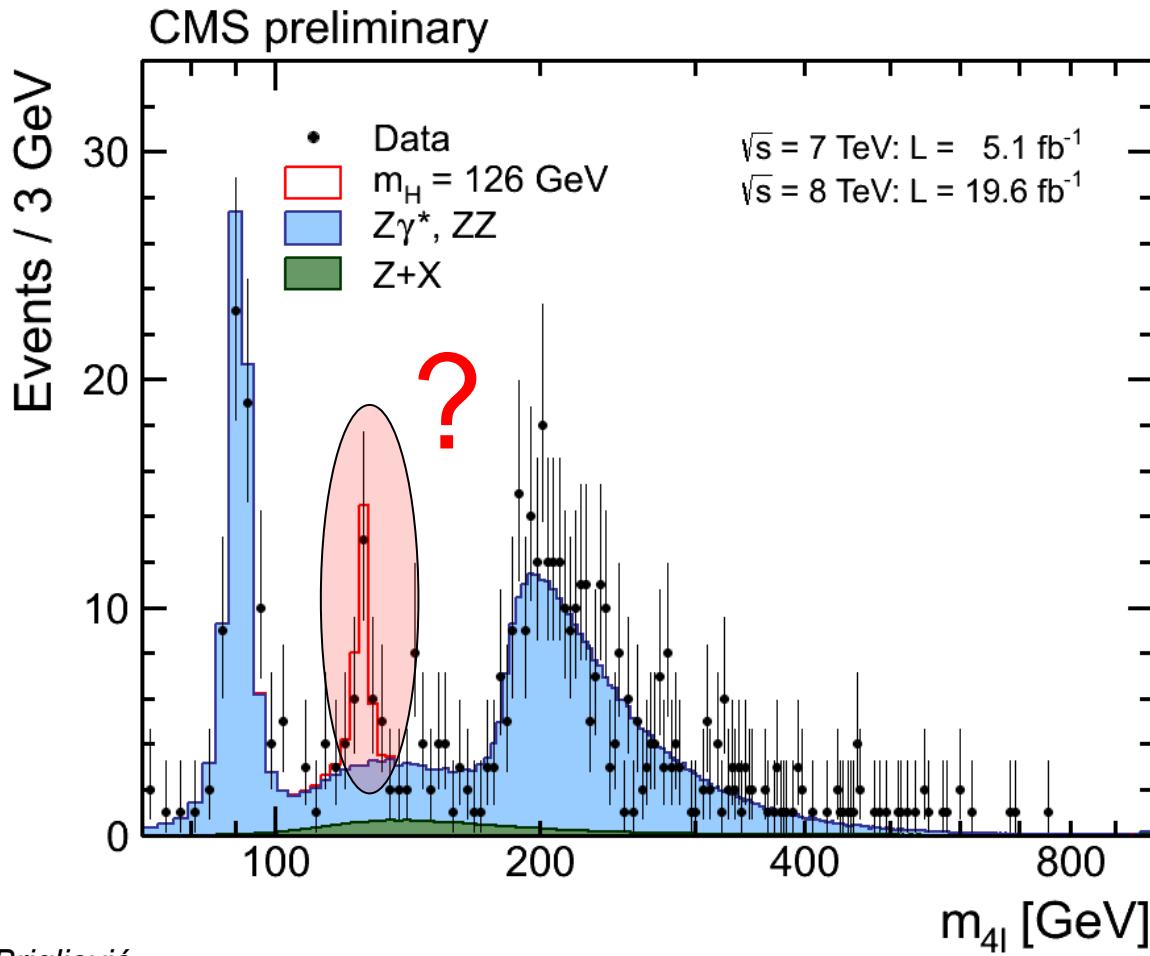
$H \rightarrow \gamma\gamma$ kandidat



Mjereni spektar mase 2 Z bozona (srpanj 2012.)



Zaključak: otkriće ili fluktuacija?



Pronašli novu
česticu?

ili

Pozadina se igra
s nama?

Prihvatanje ili odbijanje teorija?

- Dobili smo podatke s jednog eksperimenta
 - Teorija 1 se slaže s podacima
 - Teorija 2 se također slaže
 - Teorija 3 također
 - ...
 - Teorija n također
 - Tvrđnja "*Teorija 1 je prihvatljiva*" nema veliku težinu
 - Nije ni pogrešna
- Ali pretpostavimo drugi scenarij
 - Teorija 1 daje precizno predviđanje
 - Eksperiment nije u slaganju s predviđanjem
 - Onda tvrdnja "*Teorija 1 nije prihvatljiva*" ima veliku težinu
 - Zaključak: bolje je odbijati nego prihvatiti teorije / hipoteze

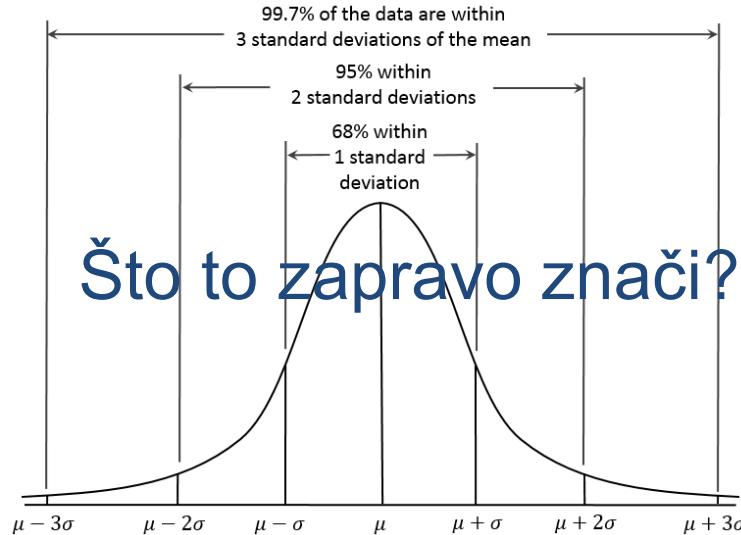
Kada najaviti otkriće?

- Tvrđiti da ste otkrili nešto je ozbiljna stvar
 - Ostati će dugo s nama (ako ne i zauvijek ☺)
- Dakle, kada tvrdimo da smo nešto otkrili?
 - Kad smo sigurni.
 - Da, ali nismo nikada sigurni!
 - U redu, ali možete biti prilično sigurni ☺
 - ‘Prilično’ nije znanstvena riječ!?
 - U redu, hajdemo se dakle dogovoriti:
 - Napravite hipotezu da je dobiveni rezultat posljedica fluktuacije pozadine (već poznatih procesa)
 - Izračunajte vjerojatnost za tu hipotezu
 - Odbacite hipotezu ako je < 0.000000287
(significance > 5)

Kada najaviti otkriće?

Konvencija je da se otkriće najavljuje kada je vjerojatnost da je opaženi signal samo fluktuacija pozadine

1 u 3 500 000
(=signifikantnost od 5σ)
ili manja



Kada najaviti otkriće?



Vjerojatnost dobivanja jedne 6-ice: $1/6$



Vjerojatnost dobivanja osam
6-ica zaredom: $(1/6)^8 = 1/1679616$

Manja od 5σ

Vjerojatnost dobivanja devet
6-ica zaredom: $(1/6)^9 = 1/10077696$

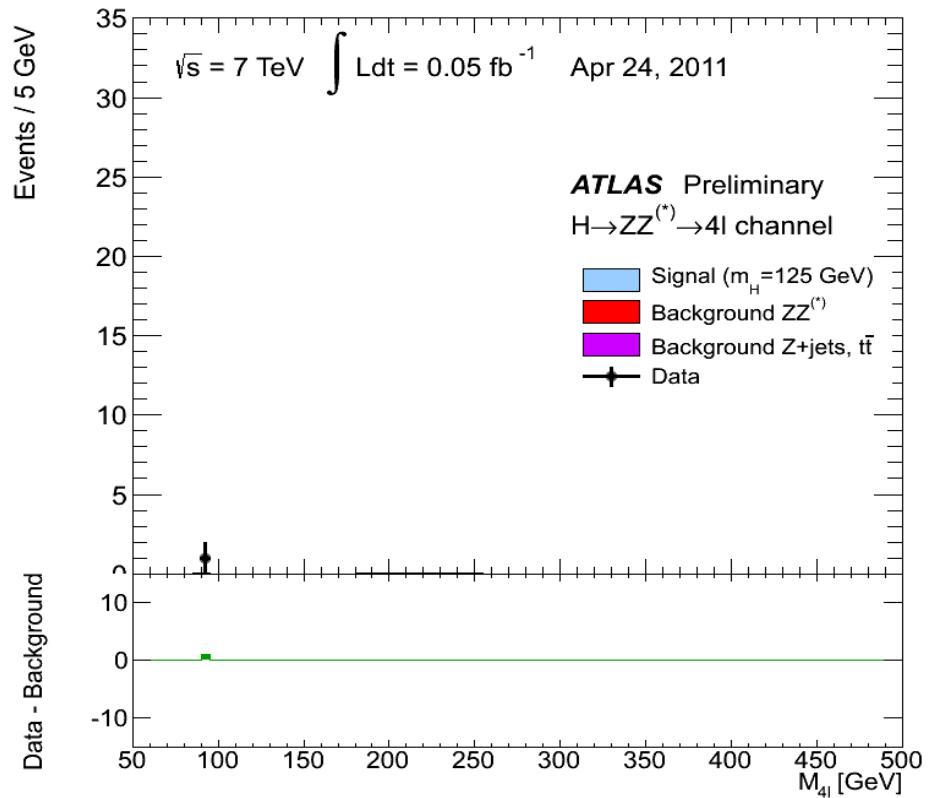
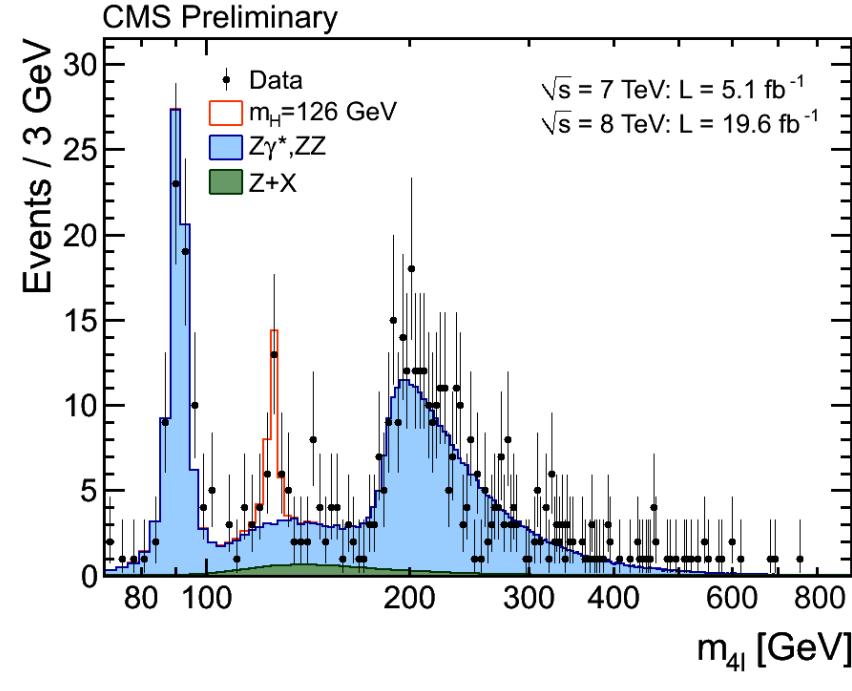
Veća od 5σ

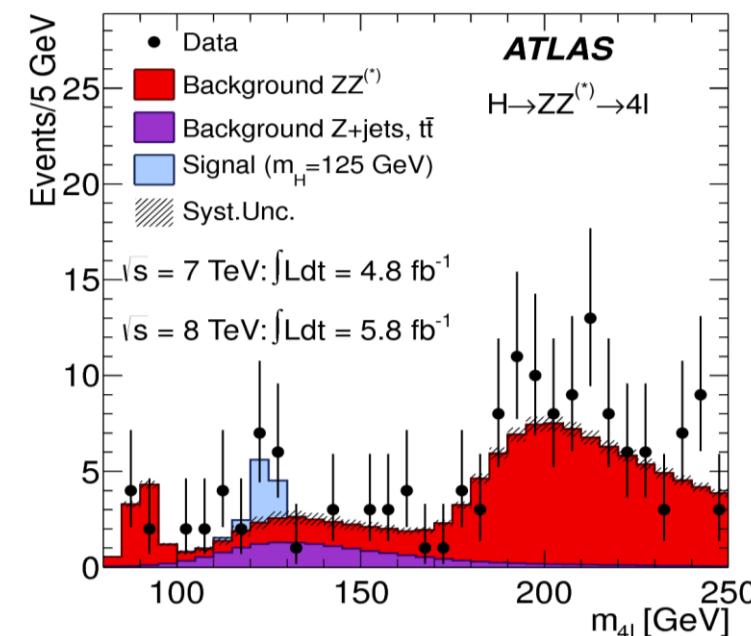
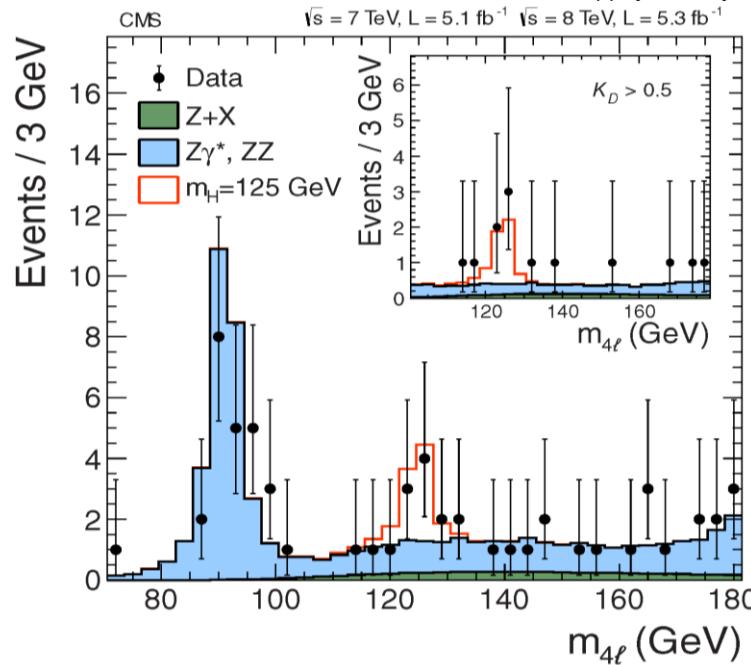
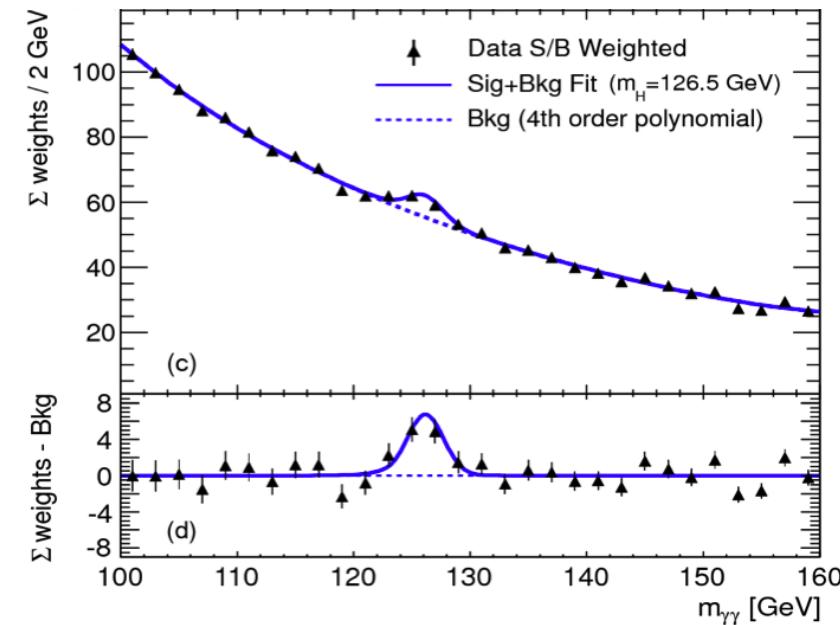
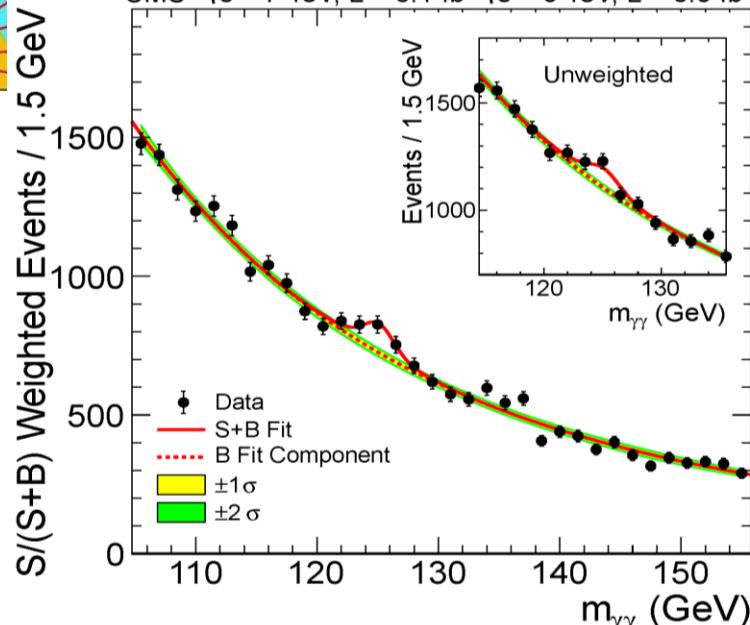
Pa još kad se takva situacija dogodi dvaput!
(npr. Higgsov bozon u ATLAS-u i CMS-u)

Kombinacija i interpretacija

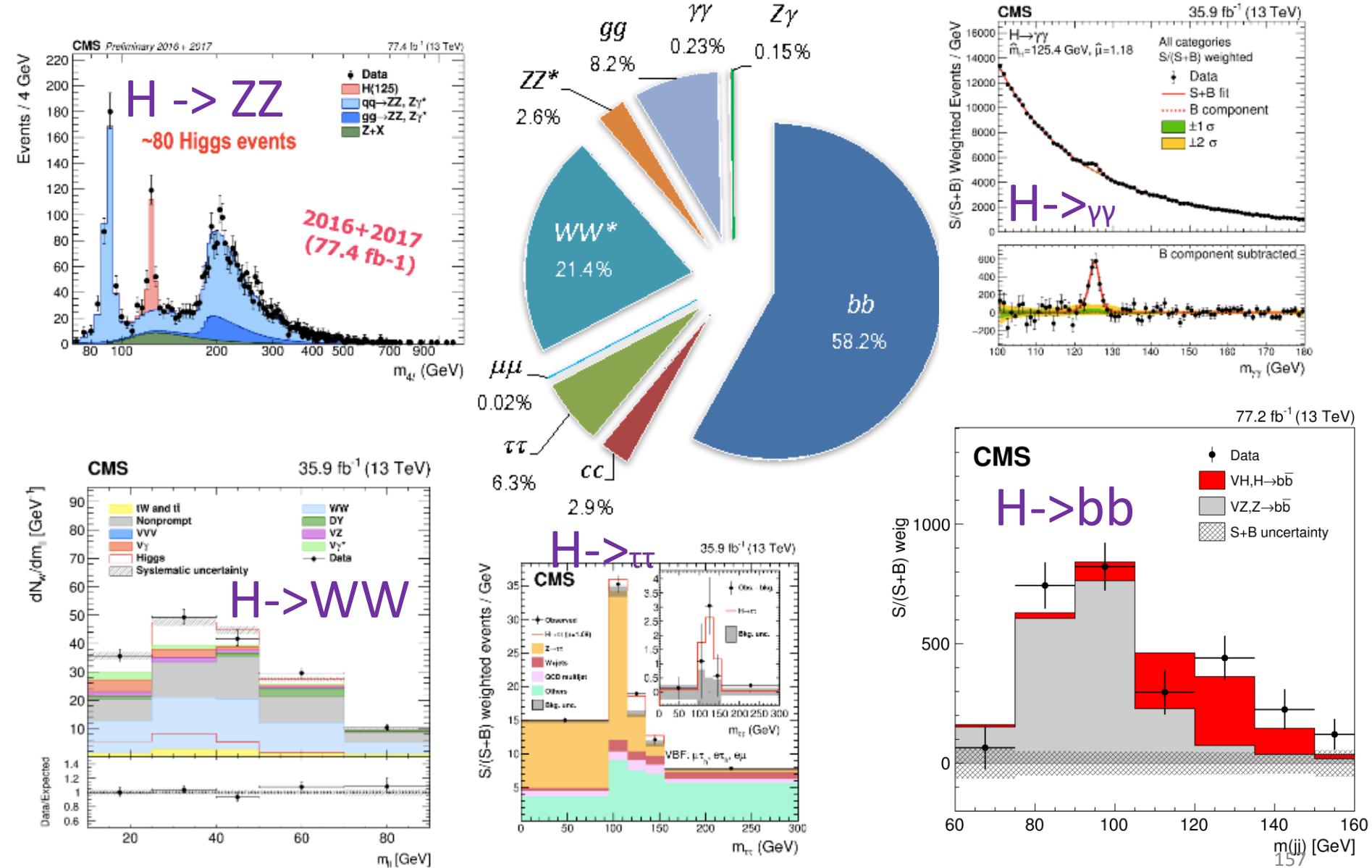


Čekamo, mjerimo i brojimo...

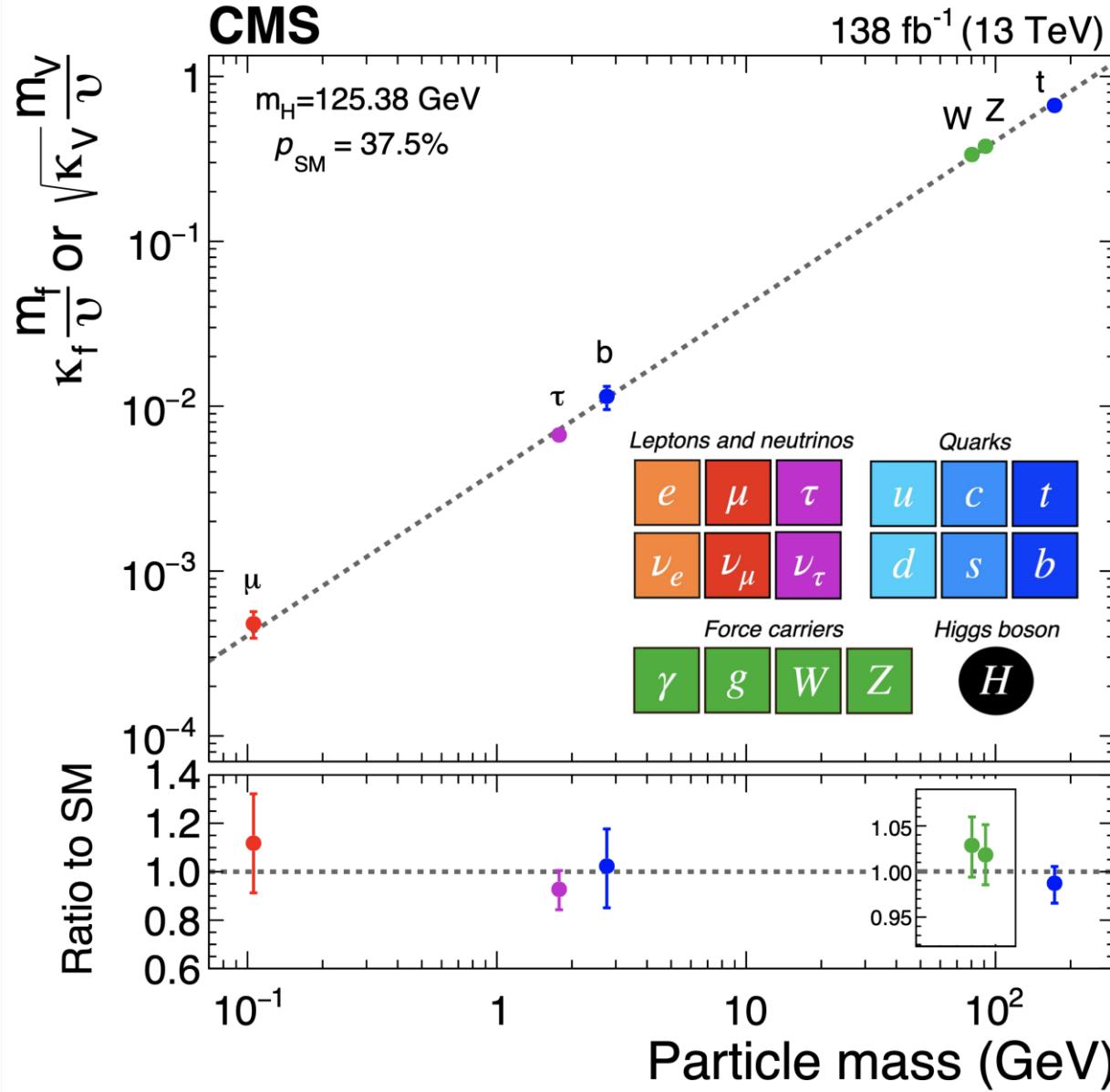




Higgsov bozon 2018.



Higgs se stvarno veže za masu



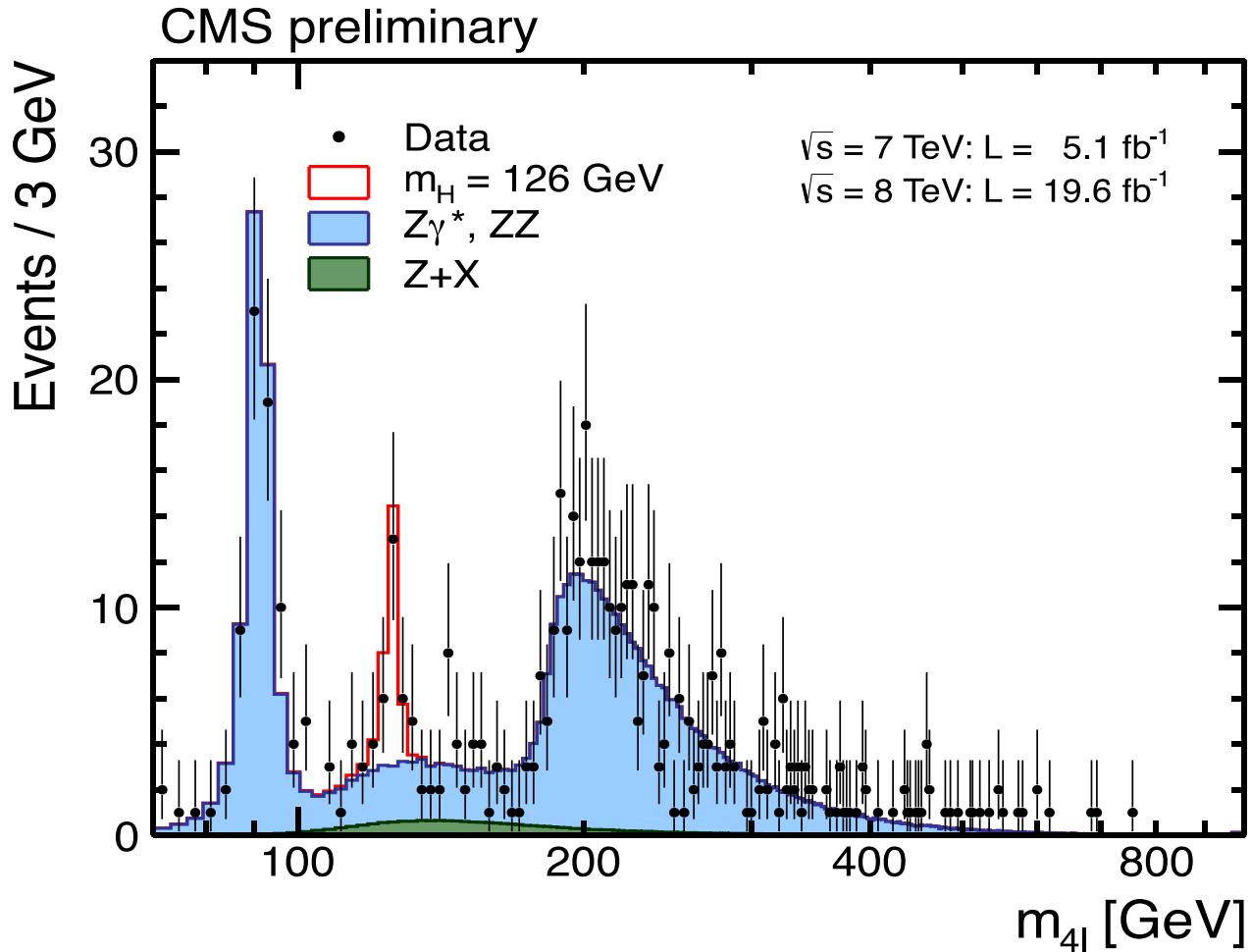


Zaključak

CERN-ovi eksperimenti su uočili novu česticu - bozon, s masom od

~ 125 GeV

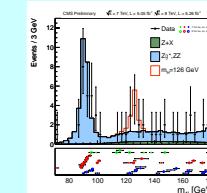
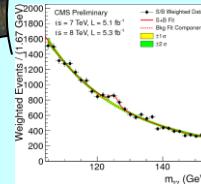
Ovo je bilo Kako smo ga otkrili?



Pitanje se promijenilo

Postoji li Higgsov
bozon?

Je li



Higgsov bozon?

The New York Times

Wednesday, July 4, 2012 Last Update: 8:54 AM ET

Discovery of New Particle Could Redefine Physical World

By DENNIS OVERBYE
21 minutes ago

The discovery by physicists at CERN's Large Hadron Collider, if confirmed to be the Higgs boson particle, could lead to a new understanding of how the universe began.

- The Lede Blog: What in the World Is a Higgs Boson?

4:16 AM ET



Fabrice Coffrini/Agence France-Presse — Getty Images

CERN officials held a press conference near Geneva on Wednesday.

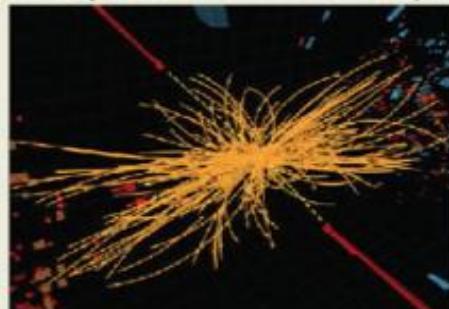


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04.07.2012, 12:13:02

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Изображение с сайта CERN



Physicists discover a candidate for the boson Higgs

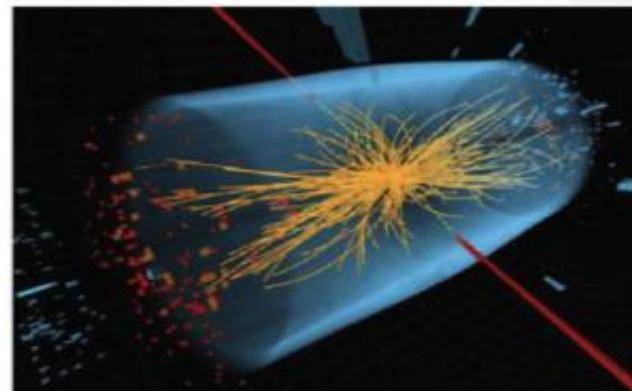
Физики обнаружили претендента на роль бозона Хиггса

LA NEWS DEL GIORNO | CRONACA | POLITICA | 10:21 - Roma, 4 lug 2012



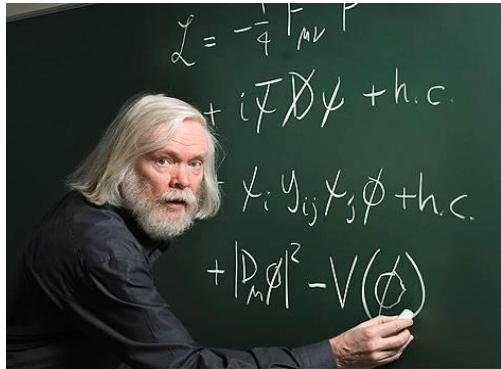
Il Bosone di Higgs esiste, oggi l'annuncio del Cern a Ginevra

Tanti indizi per il "Santo Graal" della fisica quantistica teorizzato nel 1964. E' l'ultima particella ancora da scoprire



Roma, 4 lug. (TMNews) - L'enigma relativo all'esistenza del "bosone di Higgs", il "Santo Graal" della fisica delle particelle elementari, potrebbe essere ormai vicino alla soluzione: la conferenza stampa in programma oggi al Cern potrebbe dissipare gli ultimi dubbi.





John Ellis^{1,2} and Tevong You¹

*Particle Physics and Cosmology Group, Physics Department,
King's College London, London WC2R 2LS, UK
Physics Department, CERN, CH-1211 Geneva 23, Switzerland*

Abstract

There are many indirect and direct experimental indications that the new particle H discovered by the ATLAS and CMS Collaborations has spin zero and (mostly) positive parity, and that its couplings to other particles are correlated with their masses. Beyond any reasonable doubt, it is a Higgs boson, and here we examine the extent to which its couplings resemble those of the single Higgs boson of the Standard Model. Our global analysis of its couplings to fermions and massive bosons determines that they have the same relative sign as in the Standard Model. We also show directly that these couplings are highly consistent with a dependence on particle masses that is linear to within a few %, and scaled by the conventional electroweak symmetry-breaking scale to within 10%. We also give constraints



The Nobel Prize in Physics 2013
François Englert, Peter Higgs

The Nobel Prize in Physics 2013

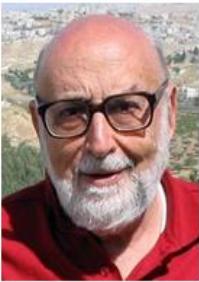


Photo: Pnicot via
Wikimedia Commons
François Englert

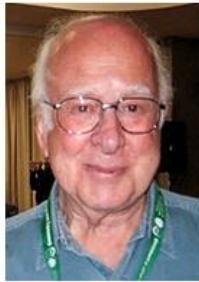
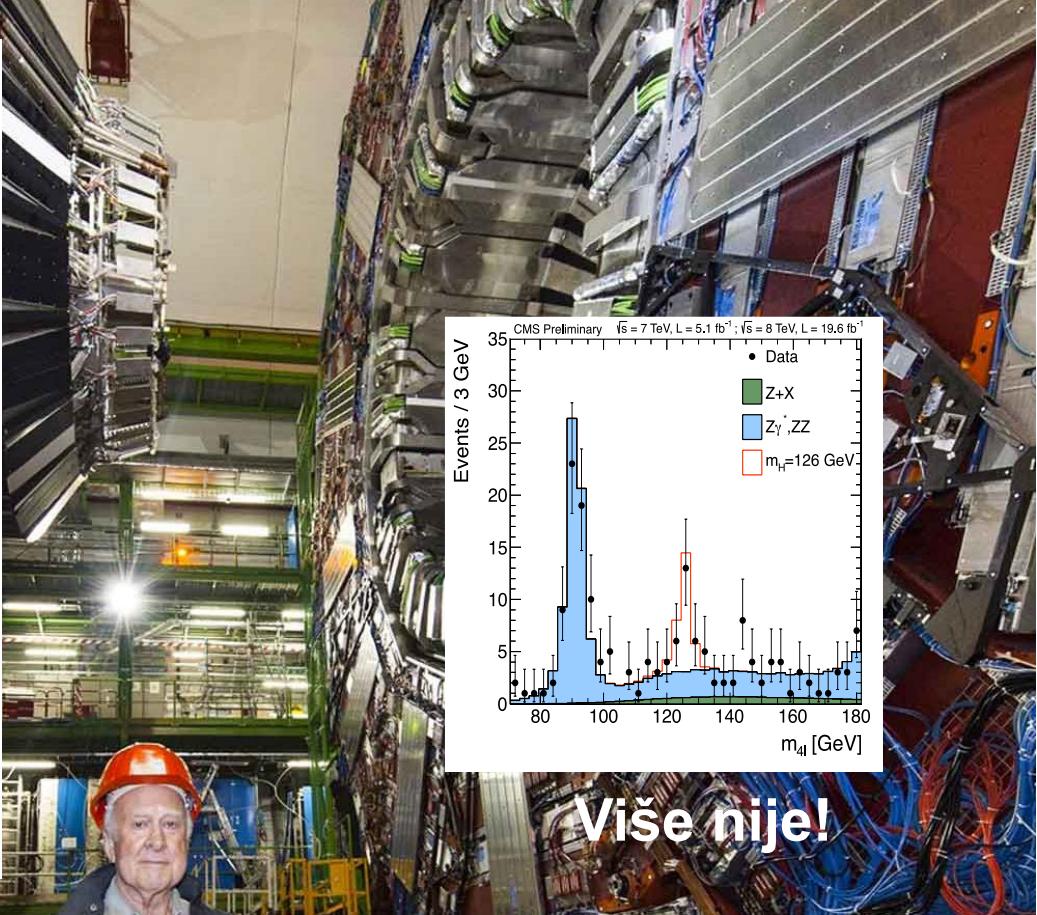


Photo: G-M Greuel via
Wikimedia Commons
Peter W. Higgs

The Nobel Prize in Physics 2013 was awarded jointly to François Englert and Peter W. Higgs "for the theoretical discovery of a mechanism that contributes to our understanding of the origin of mass of subatomic particles, and which recently was confirmed through the discovery of the predicted fundamental particle, by the ATLAS and CMS experiments at CERN's Large Hadron Collider"



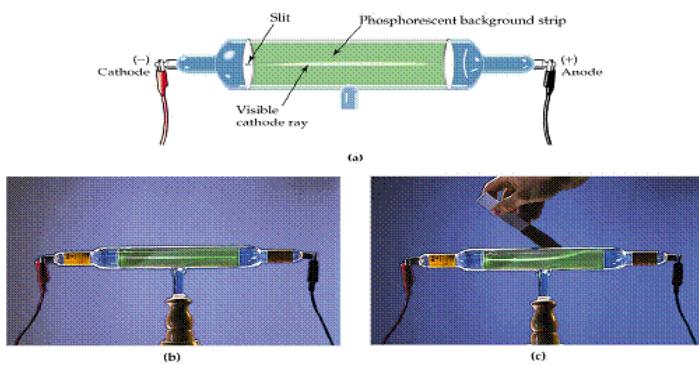
Više nije!

Jedini sigurno opaženi
Higgs u CMS
eksperimentu do 2012.
Dr. Peter Higgs



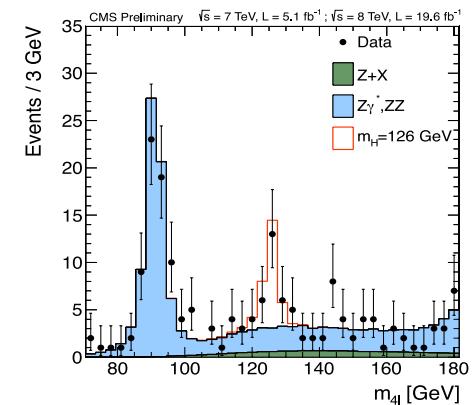
STANDARDNI MODEL JE KOMPLETAN I POTVRĐEN

Discovery of the electron
J.J. Thomson (1897)

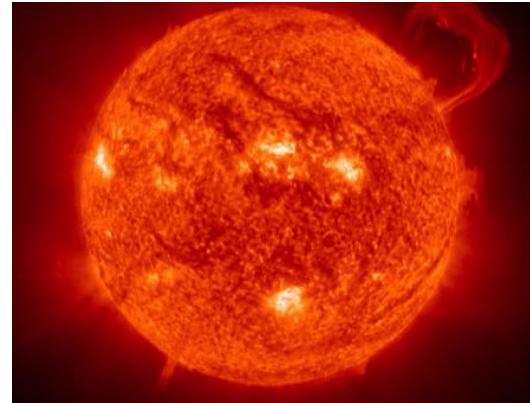


Discovery of the Higgs boson
LHC (2012)

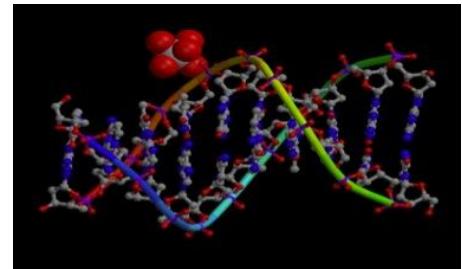
$$\begin{aligned}\mathcal{L} = & -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\ & + i \bar{\psi} D^\mu \psi + h.c. \\ & + \bar{\chi}_i Y_{ij} \chi_j \phi + h.c. \\ & + |\partial_\mu \phi|^2 - V(\phi)\end{aligned}$$

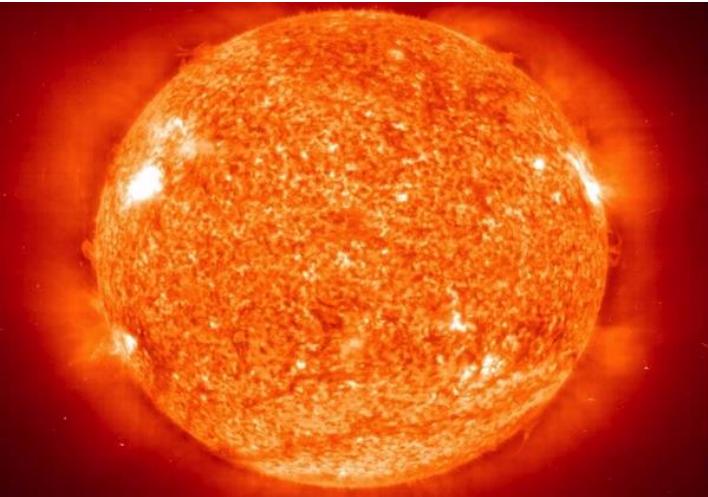


I objašnjava mnogo toga

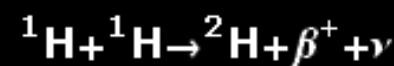
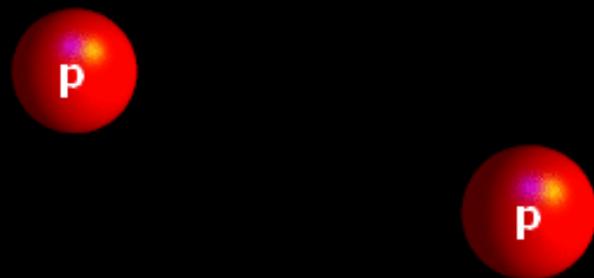
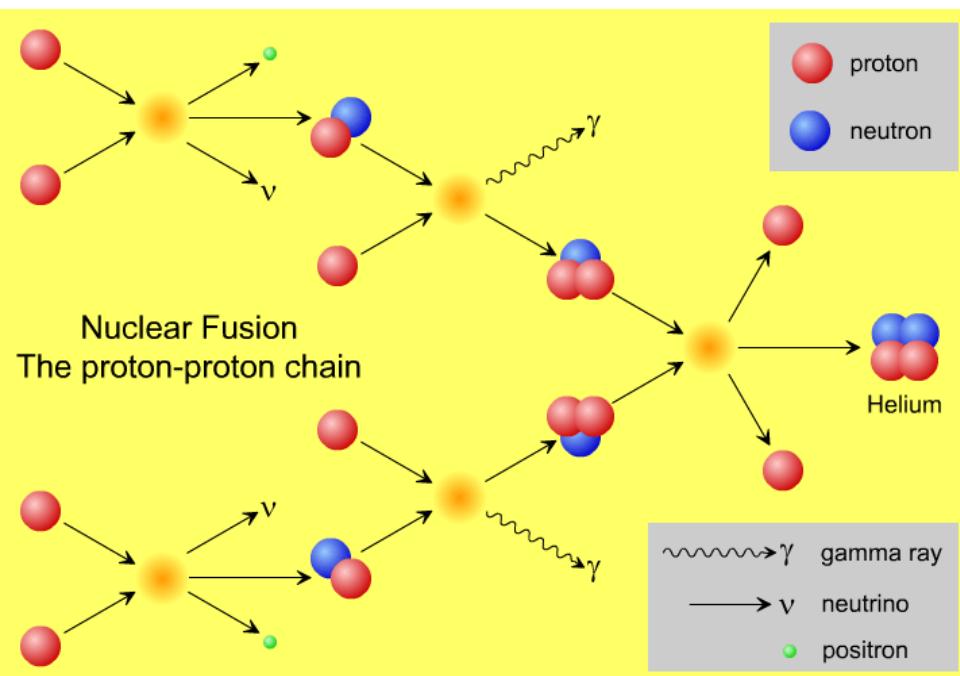


$$\begin{aligned}\mathcal{L} = & -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\& + i \bar{\psi} \not{D} \psi + h.c. \\& + \chi_i Y_{ij} \chi_j \phi + h.c. \\& + |\partial_\mu \phi|^2 - V(\phi)\end{aligned}$$



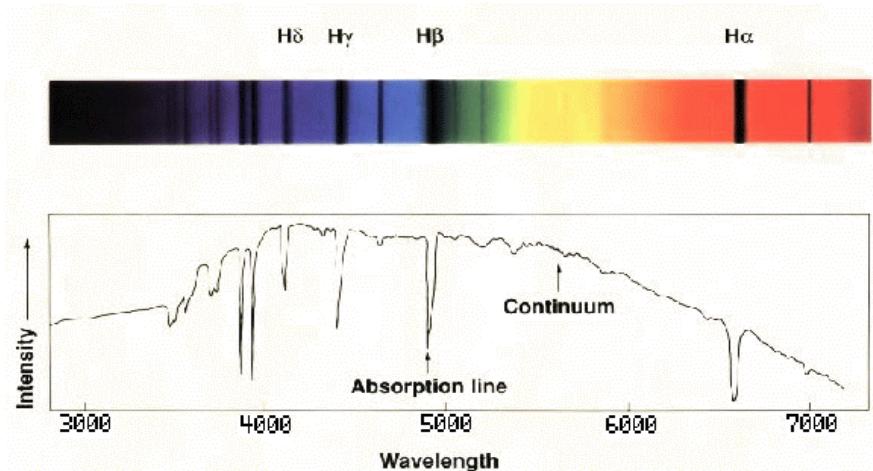
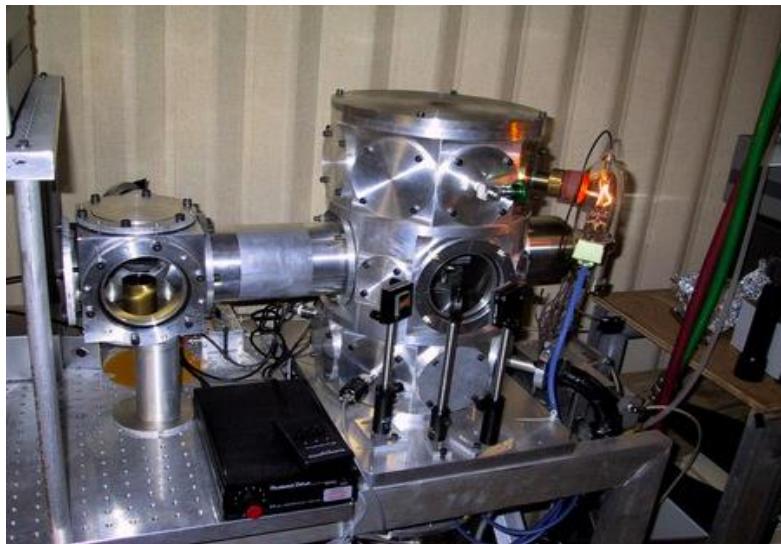


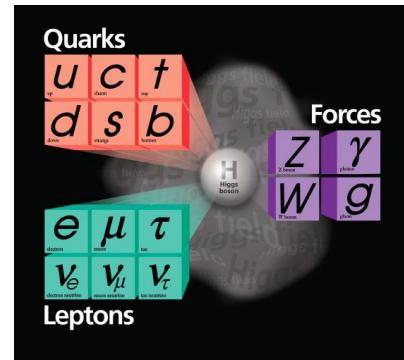
Primjer: fuzija u suncu



Na zemlji i u cijelom svemiru

Stella spectra tell us stars are made of the same stuff we can study in an atomic physics lab.



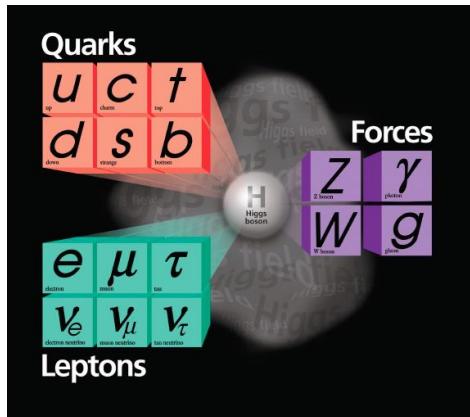


*Zašto većina nas mislimo da
SM ne može biti kraj priče*

There are things about
the known particles and forces
the SM does not explain
(or at least not satisfactorily)

SM NE objašnjava

- Zašto 3 generacije?
- Zašto je tako fino podešeno?
- Zašto su mase čestica tako različite
- > 20 free parameters
 - masses, couplings, ...
- Where did all the antimatter go?
- Does the Unification go on?



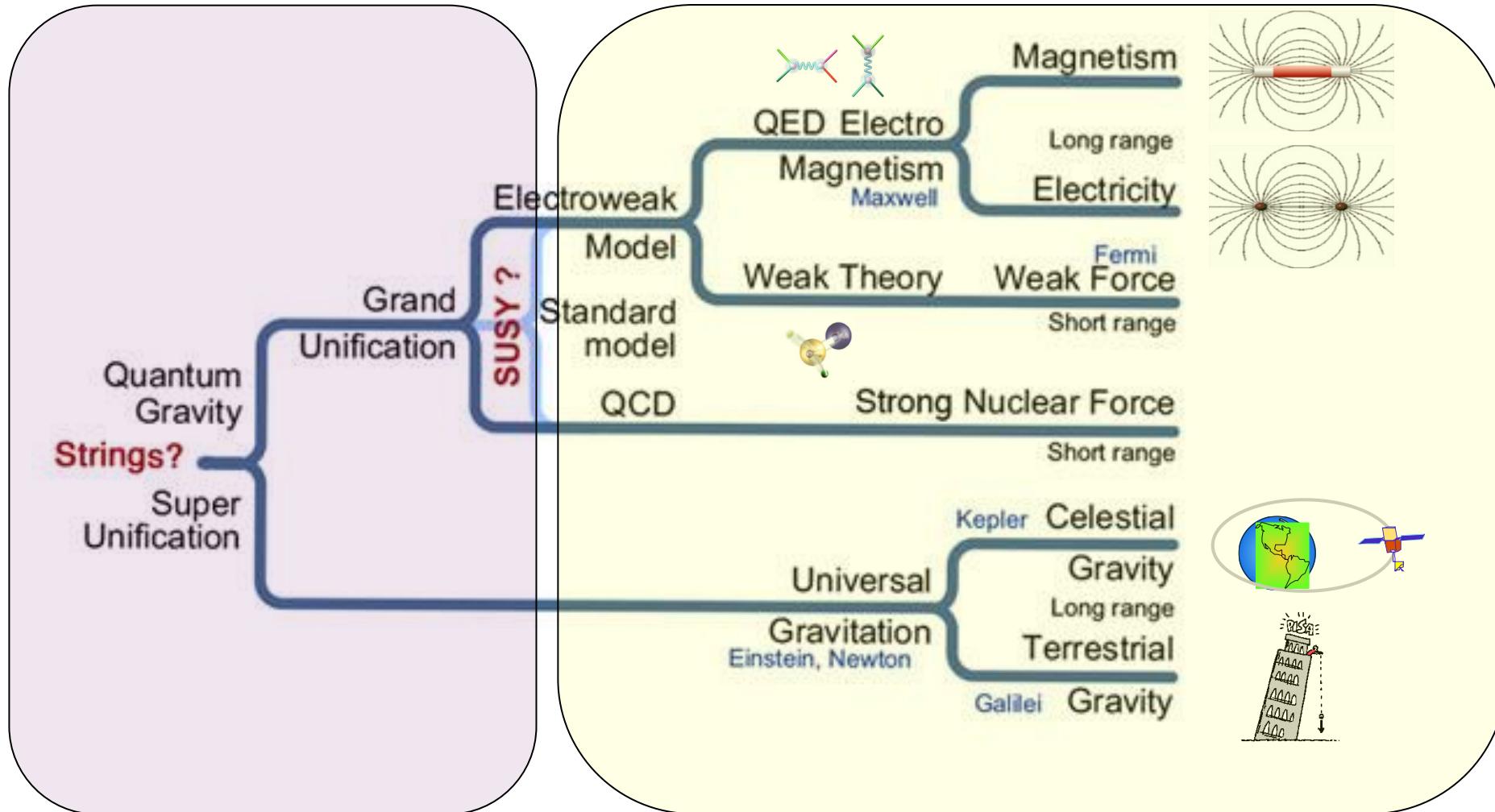
And most importantly:

- **WHAT ABOUT GRAVITY?**

Fizika voli UJEDINITI Ide li to dalje?

SPEKULACIJA

ČINJENICE

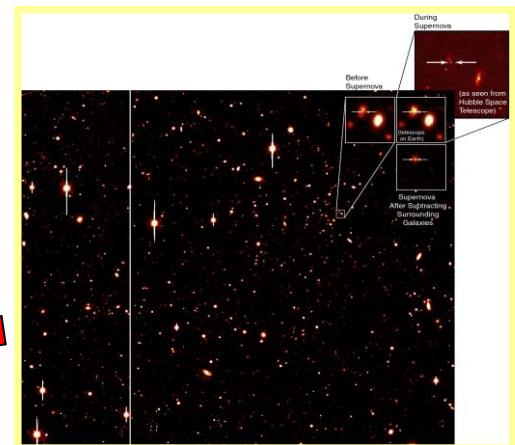
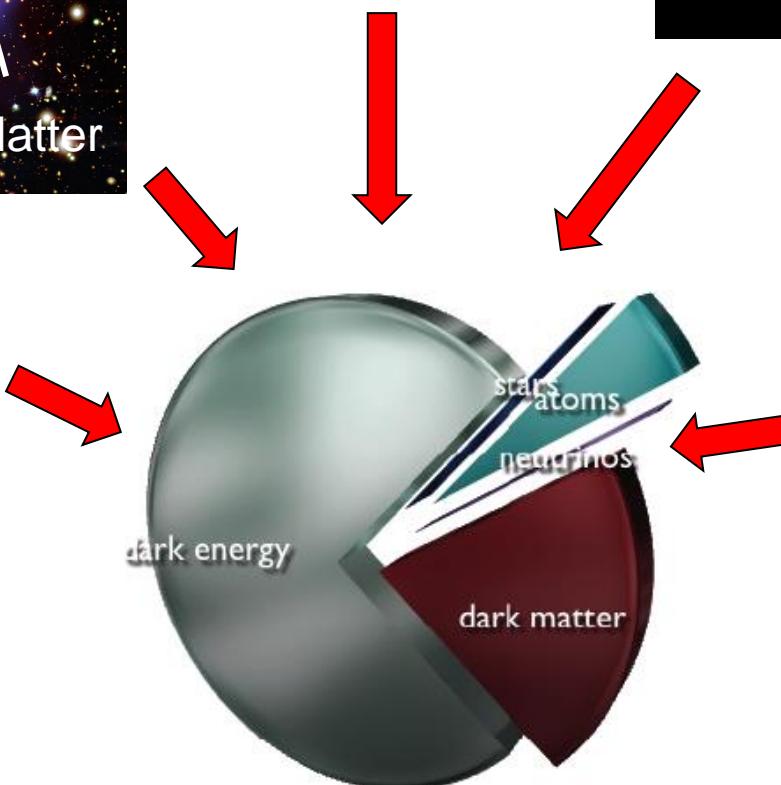
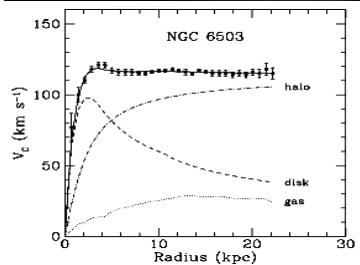
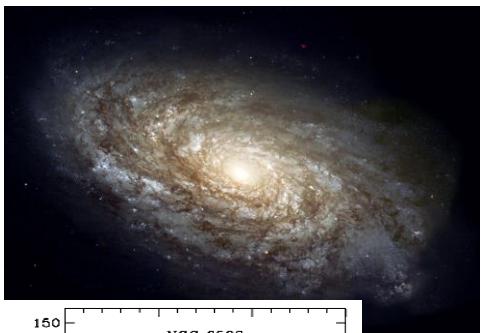
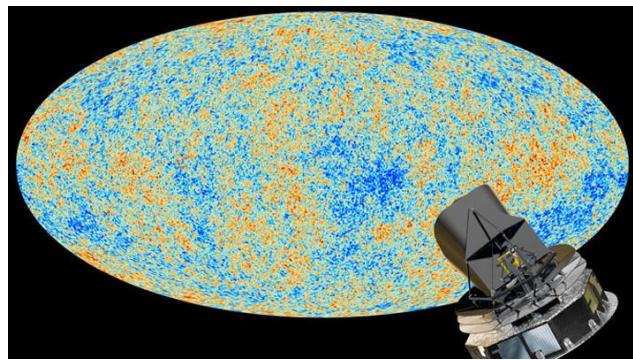
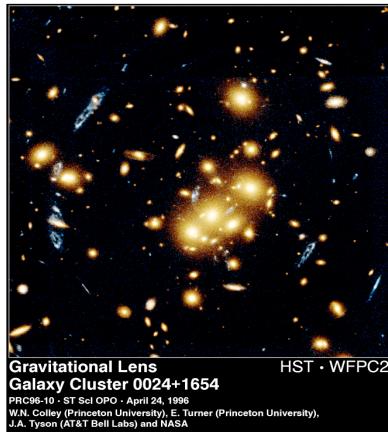
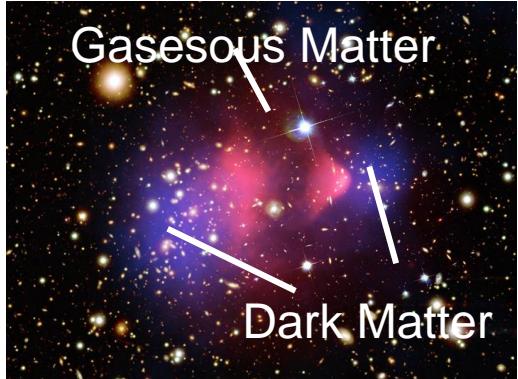


Zašto ZNAMO da SM nikako ne može biti kraj priče

Postoje opažanja u vidljivom
svemiru koje SM nikako ne može
objasniti

Huge amounts of invisible matter:

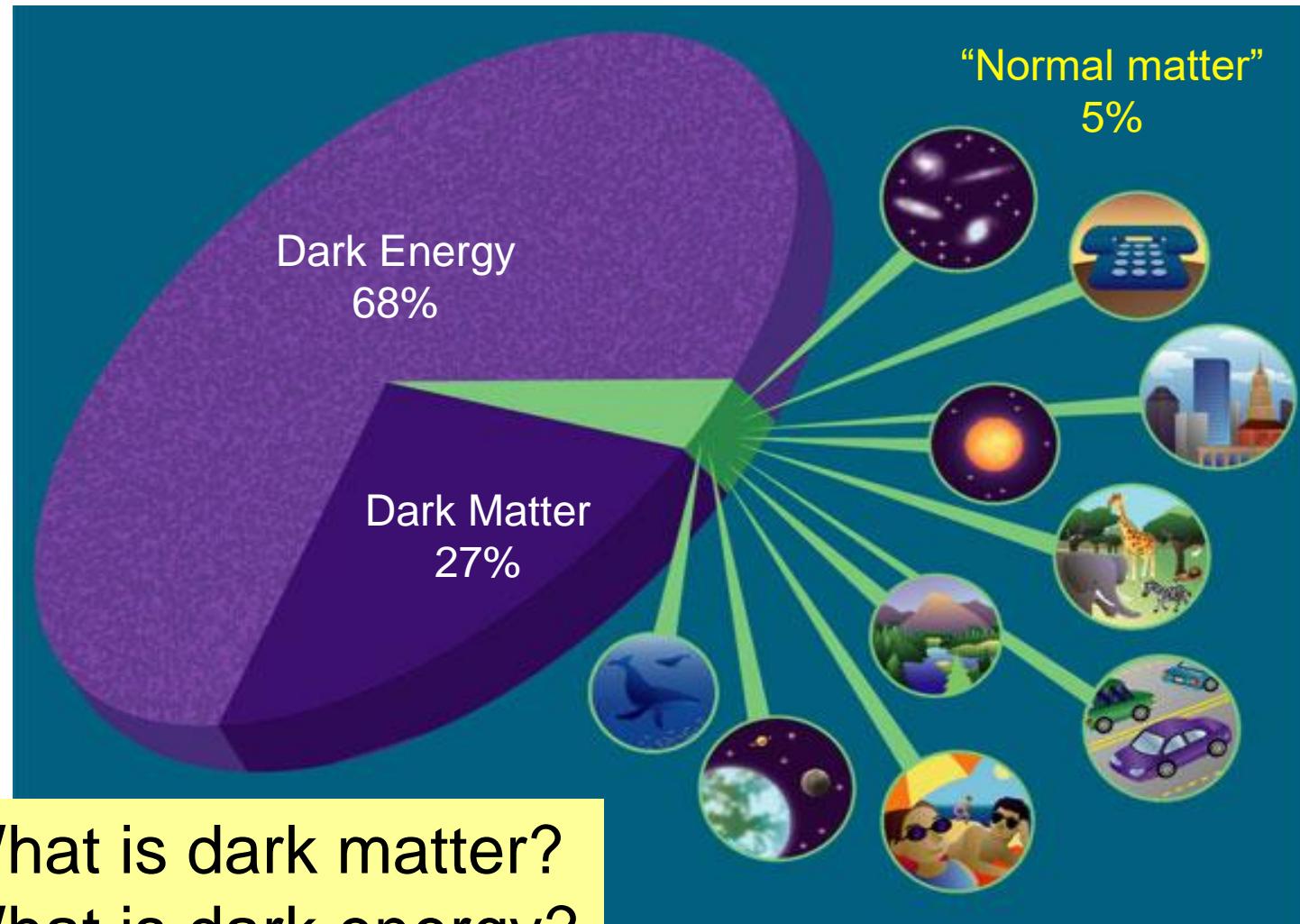
“DARK MATTER”



The expansion of the universe is accelerating:

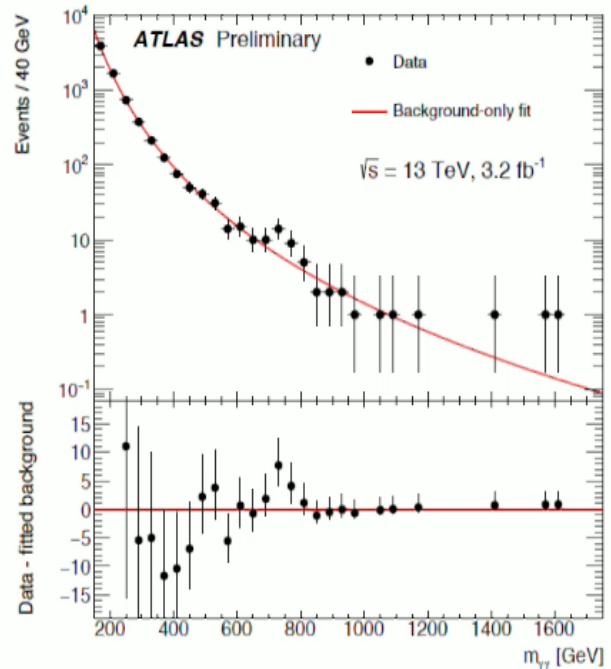
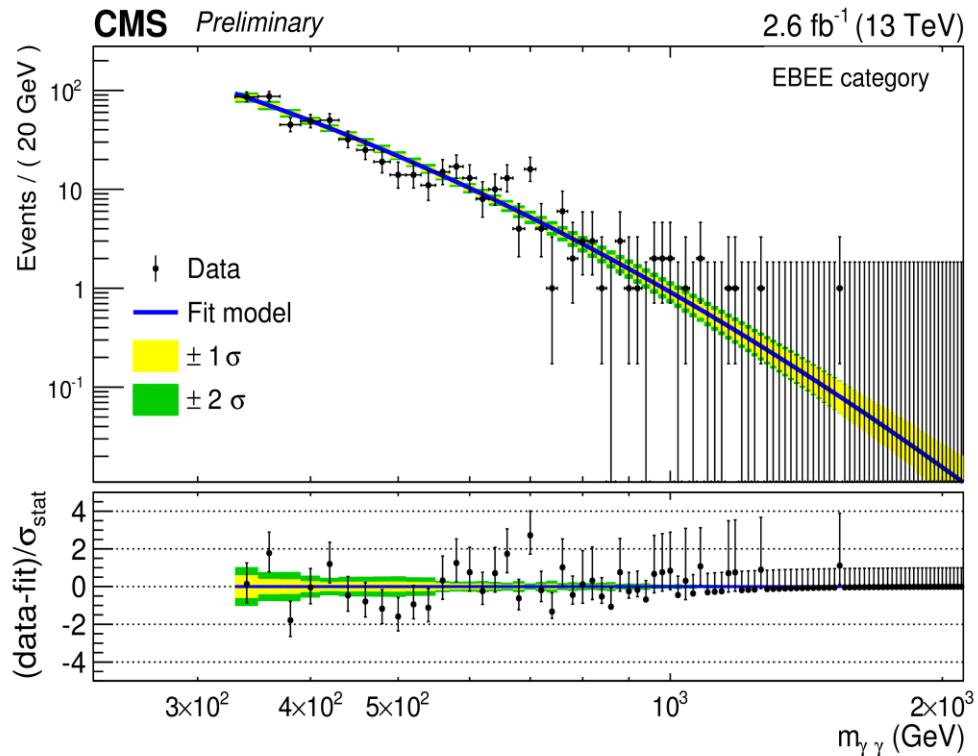
“DARK ENERGY”

What is the Universe made of?



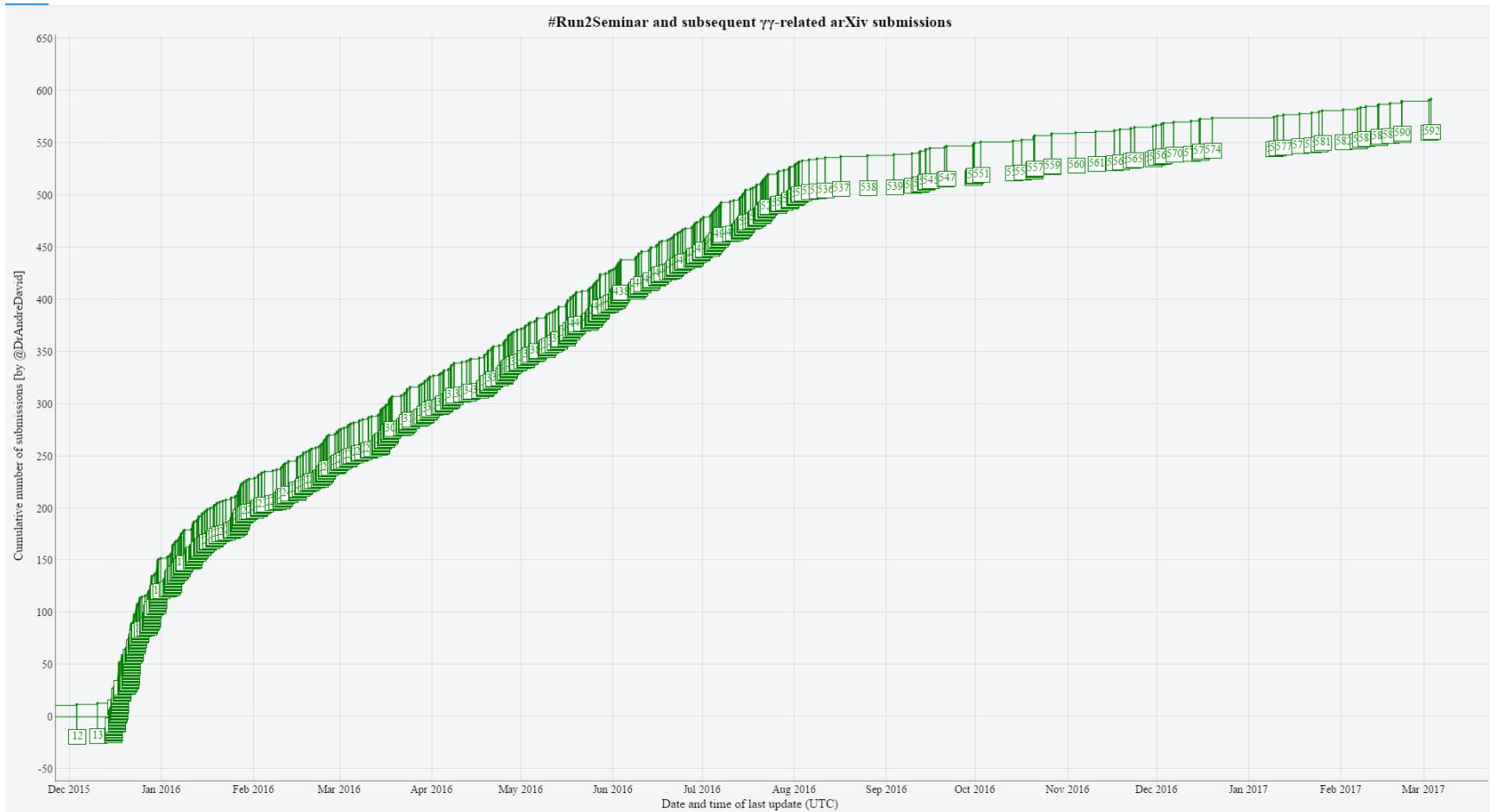
What is dark matter?
What is dark energy?

2015: Naznake nove čestice?

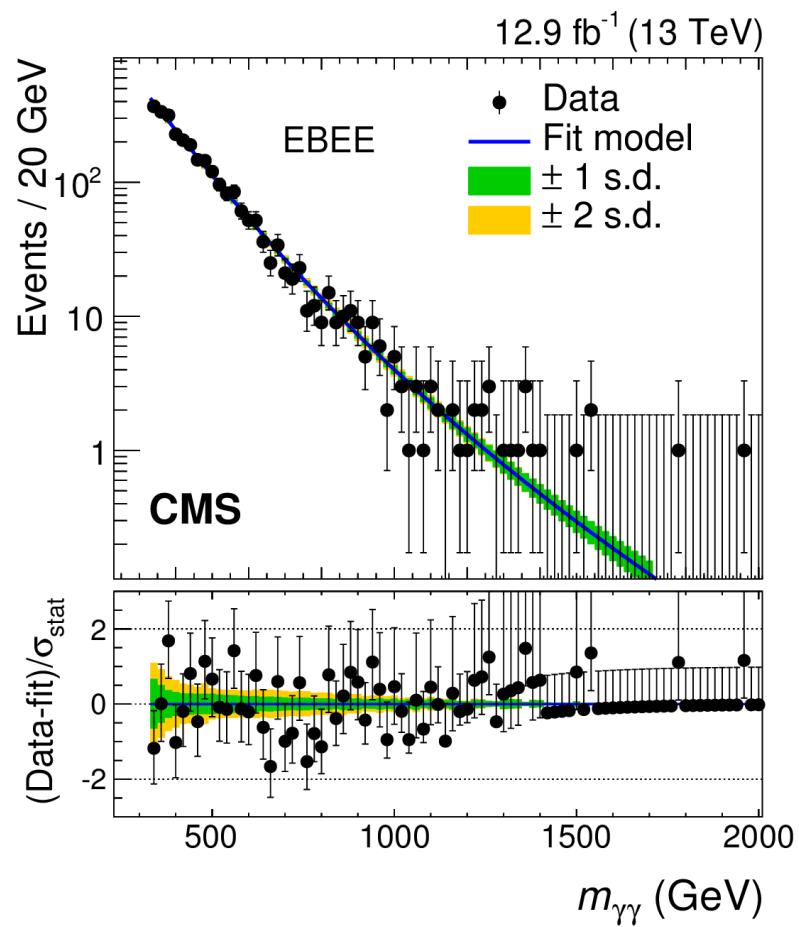
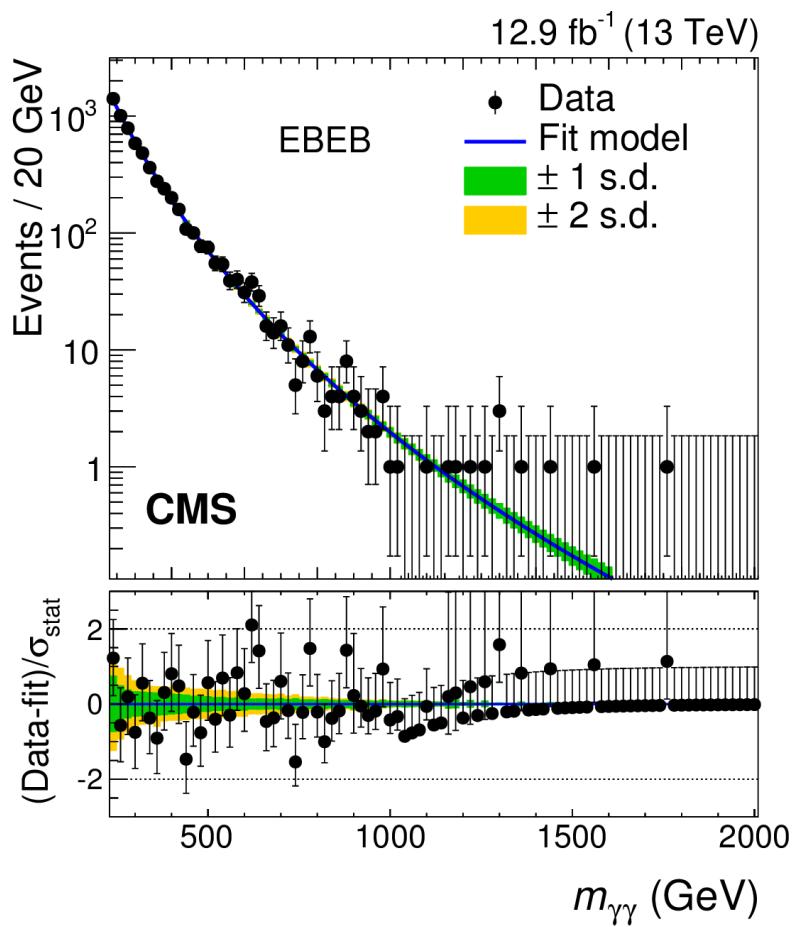


Signifikantnost viška događaja između 3 i 4 σ u oba eksperimenta

2015: Naznake nove čestice?



...ipak se statistika poigrala s nama

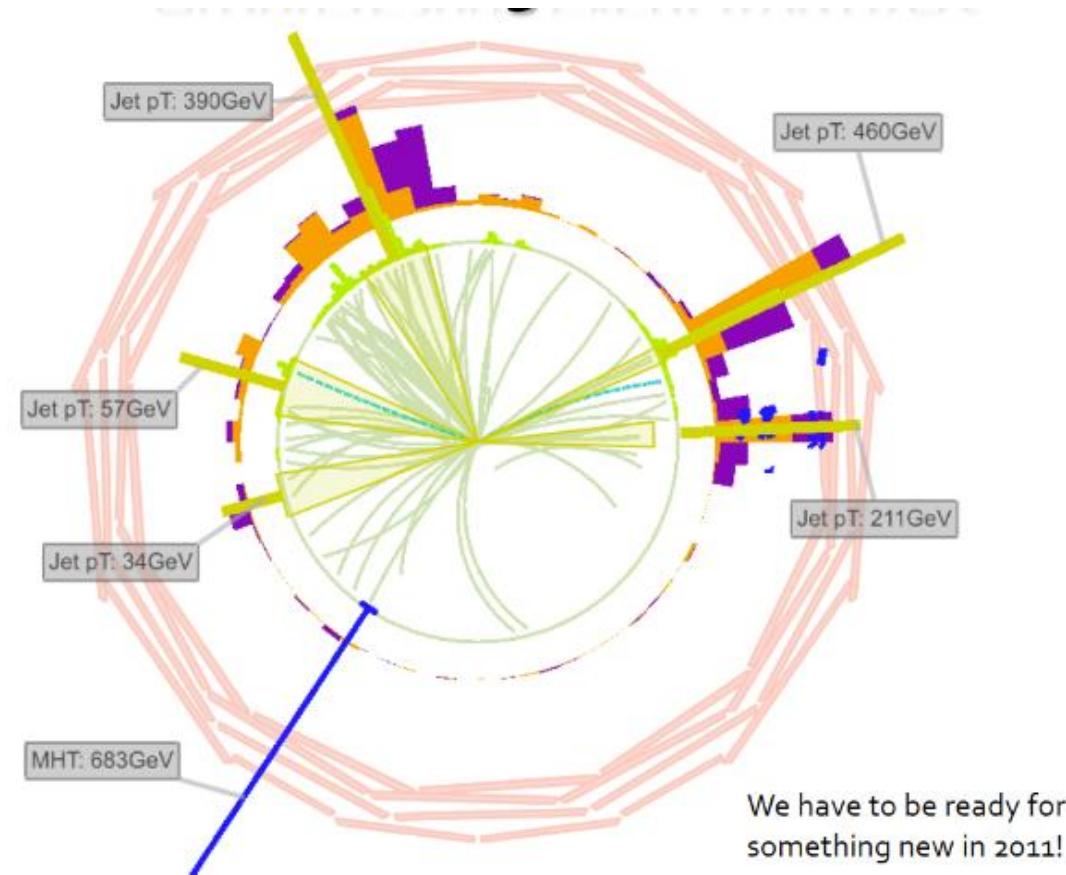


Može li nam LHC nešto reći o
prirodi tamne materije?

Može

Tamna tvar na LHC-u?

$$M\vec{ET} = - \sum_i \vec{p}_{T,i}$$



A tamne energije?

???



ALI TRAŽIMO DALJE

Što zahtijevaju nova otkrića ...



- **Sudarivače čestica:** Moćne strojeve koji ubrzavaju čestice na ekstremno velike brzine i ‘natjeraju’ da se međusobne sudare
- **Detektore:** Gigantske instrumente za snimanje rezultirajućih čestica koje izlaze iz točke interakcije
- **Računala:** Za prikupljanje, snimanje, distribuciju i analizu ogromnih količina podataka koje proizvode detektori
- **Ljude:** Samo kolaboracije na svjetskoj razini, od po nekoliko tisuća znanstvenika, mogu izvoditi ovakve eksperimente

CMS Collaboration



Hvala na pažnji!