

STATISTICALLY LEARNING THE NEXT STANDARD MODEL FROM LHC DATA



Wolfgang Waltenberger
(ÖAW and Uni Wien)

Seminar for HEPHY AI group leader position
Vienna, September 2022

WHO AM I

Priv. Doz. Dr. Wolfgang Waltenberger

Born december 10th 1974

Father of Fabian (2000) and Yannick (2005)

2001 - now: CMS member

2004: Finished PhD at TU Wien

2008 - now: staff member at HEPHY Wien

2014/15: Guest Professor at USP São Paulo (Brazil)

since 2016: lecturing at TU Wien, Uni Wien

2018: Habilitation at Uni Wien

2021/22: Guest Professor in Grenoble (France)

Worked on: tracking, vertexing, b -tagging, search for SUSY, interpretation,
simplified models, SModelS, ml/stats

inverse problem of particle physics



MY CURRENT DATA SCIENCE INVOLVEMENTS

At the Austrian Academy of Sciences

Founded **ÖAW AI Graduate School**, directed first installment (2019), now preparing to be again **director** of third iteration (january 2023)

At HEPHY

(Co-)supervising **three PhD students in three different groups** [see next slide]
Work on learning (as opposed to postulating) the Next Standard Model from data

At TU Wien and Uni Wien

Teaching **regular courses in statistics** (see youtube.com/WolfgangWaltenberger)

Generally – as AI/ML is still trending: currently many invited lectures, lecture series, and seminars on AI/ML/stats related topics

Eg. WMLQ2022, September 2022, Poland, Vietnam School of Physics, July 2022, Vietnam
In the past: MCNet Summer School, Sweden, VDSP Winter School, Austria, DKPI Graduate School, Austria,

MY CURRENT DATA SCIENCE INVOLVEMENTS

Co-supervising three PhD students:

	Felix Wagner	Mark Matthewman	Sahana Narasimha
group	CRESST/ COSINUS	CMS (HGCAL)	phenomenology
co-supervising with	Florian Reindl (TU Wien)	Erica Brondolin (CERN)	-
physically at	Vienna	CERN	Vienna
Data scientific highlights	Reinforcement learning of detector control	Graph neural network for HGCAL energy regression	
Financed by	FFG	CERN PhD programme	Bilateral French-Austrian ANR-FWF
Started	July 2020	March 2022	October 2022

common denominator is data science!

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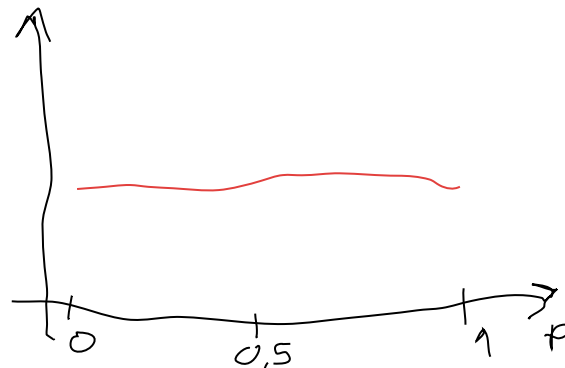
PRELUDE:

FUN WITH META-STATISTICS

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- Our **SModelS v2.2.0 database** summarizes the results of **almost 1000 signal regions** of **about 100 CMS and ATLAS publications** of searches for new physics
- For each signal region, we know the number of observed events ending up in this signal region, alongside with the number of expected Standard Model “background” events and its error. Assuming a simplified statistical model, **we can compute p -values for the Standard Model hypothesis**
- **If there is no new physics in the data, the distribution of p -values should look like this:**

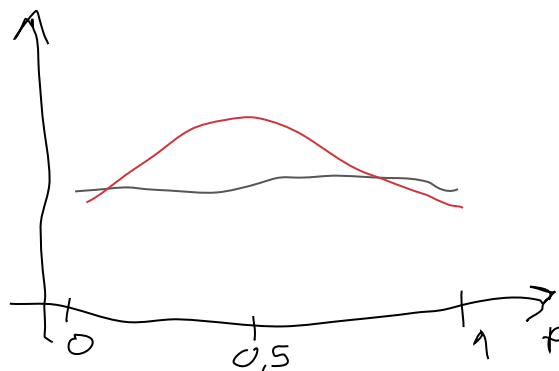


($p = 0$ means huge excess of observed events)

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- **Given that the background errors are conservatively, systematically overestimated by $\sim 30\%$, we expect the following distribution for the p -values:**

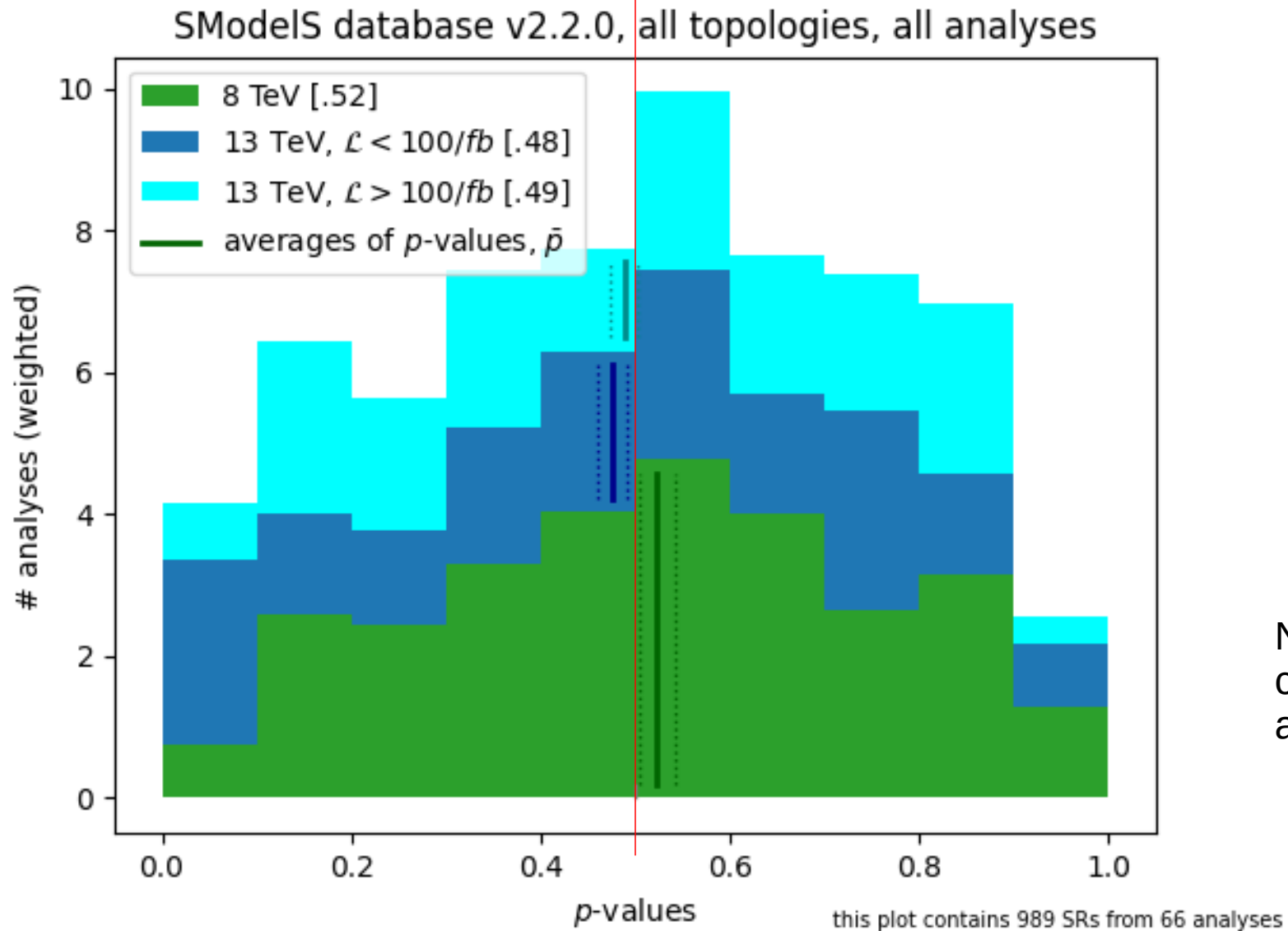


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FUN WITH META-STATISTICS

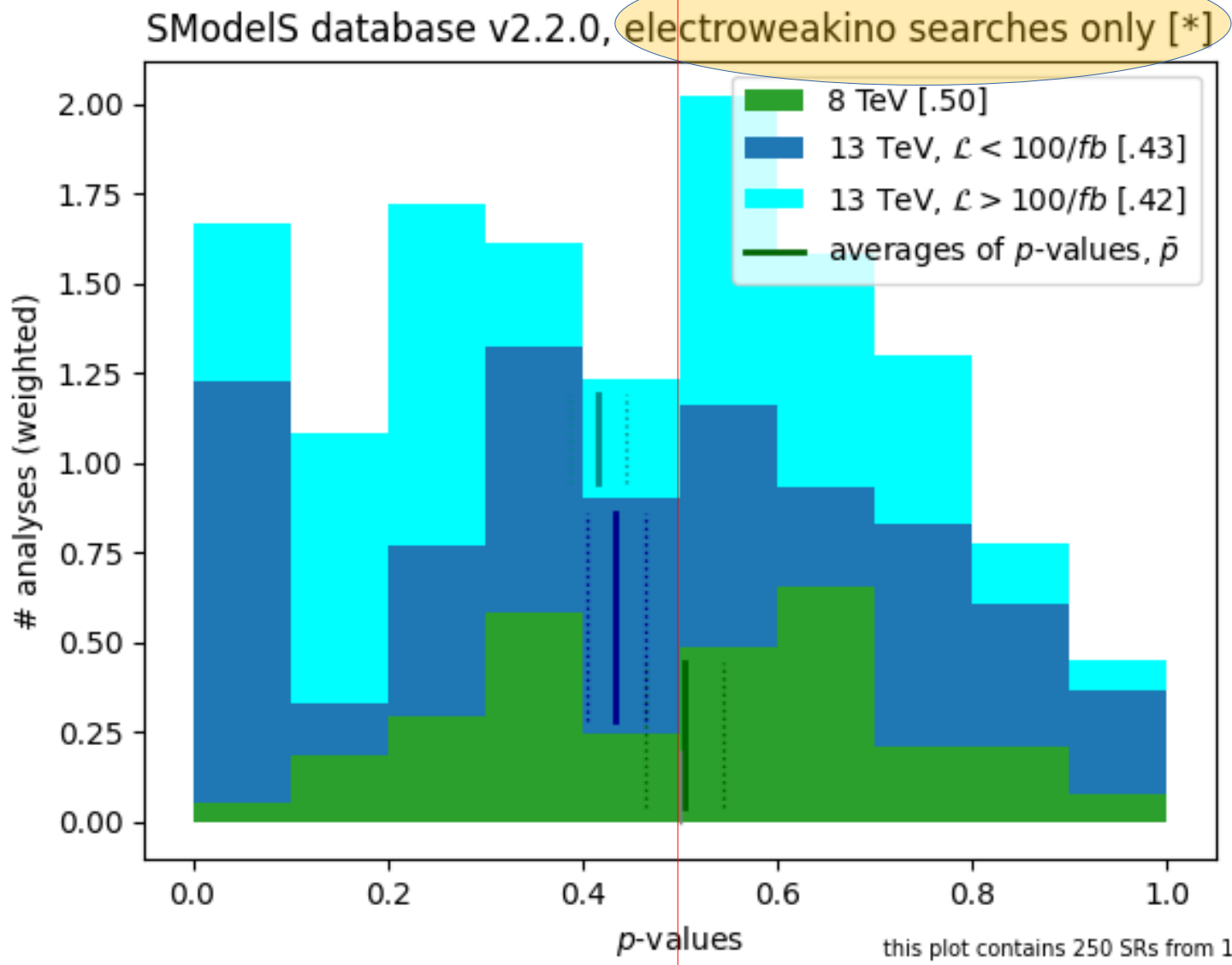


Here's with the actual, **real data**:



Lookin' good! No obvious p -hacking in our search programme.

FUN WITH META-STATISTICS

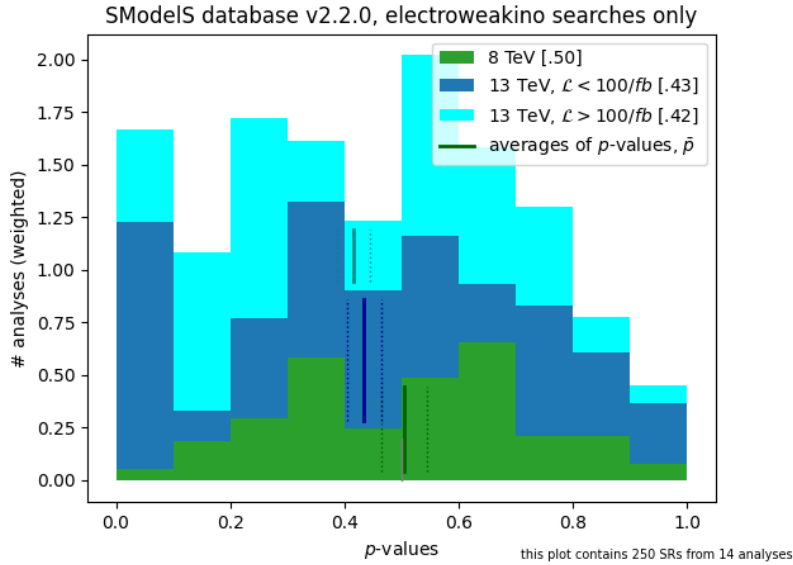


N.B: the color bars are stacked

Random fluke? Selection bias? New physics slowly seeping in?

[*] searches that target chargino/neutralino productions in RPC SUSY scenarios. Decays via W,Z,h bosons + dark matter candidate
[**] some data are used more than once in this plot. We cannot – and do not pretend to -- make too serious frequentist statements

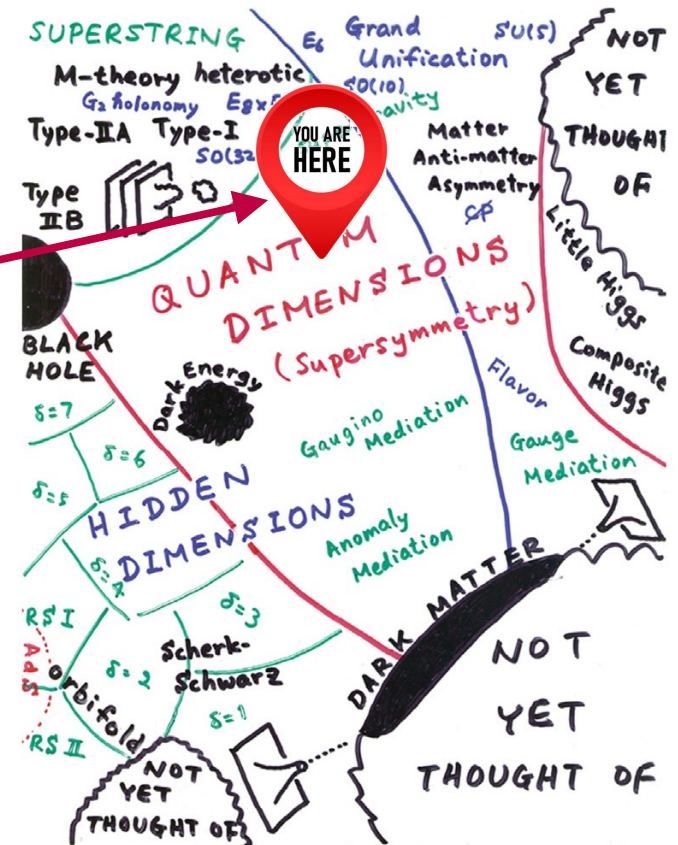
FUN WITH META-STATISTICS



If the trend were (conjunctive form!) due to new physics, **how would we go from such dispersed signals**

... to identifying the One True Point [*] (a.k.a. **the Next Standard Model**) in Hitoshi Murayama's landscape of theories?

→ **our Inverse Problem!**



[*] obviously in reality we would actually want a posterior density, not one point

STATISTICALLY LEARNING THE NEXT STANDARD MODEL FROM LHC DATA



image courtesy of Jon Butterworth, Chris Wormell

Wolfgang Waltenberger (ÖAW, Uni Wien)

[presenting work in collaboration with Andre Lessa and Sabine Kraml]

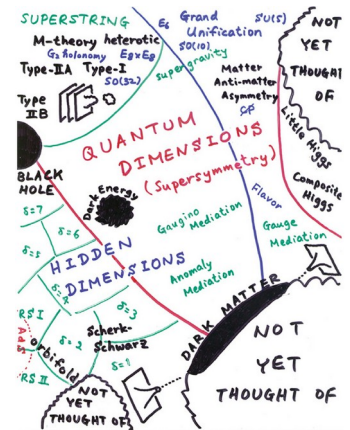
<https://arxiv.org/abs/2012.12246>

Seminar, HEPHY AI group leader
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OUR INVERSE PROBLEM



Model	Mass (TeV)	95% CL Lower Limits	ATLAS Primary Reference
Minimal Supergravity	100	100	ATLAS-CONF-2015-025
Gravitino GMSUGRA	100	100	ATLAS-CONF-2015-025
Universal Extra Dimensions	100	100	ATLAS-CONF-2015-025
...

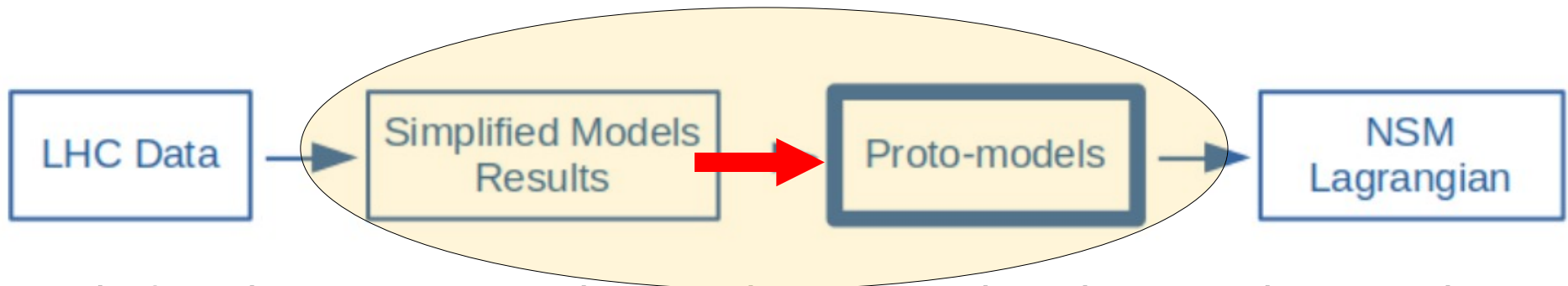


- too difficult a task for us humans (we are neither smart nor creative enough)
- **Let the machine solve it!**

Our mission statement:

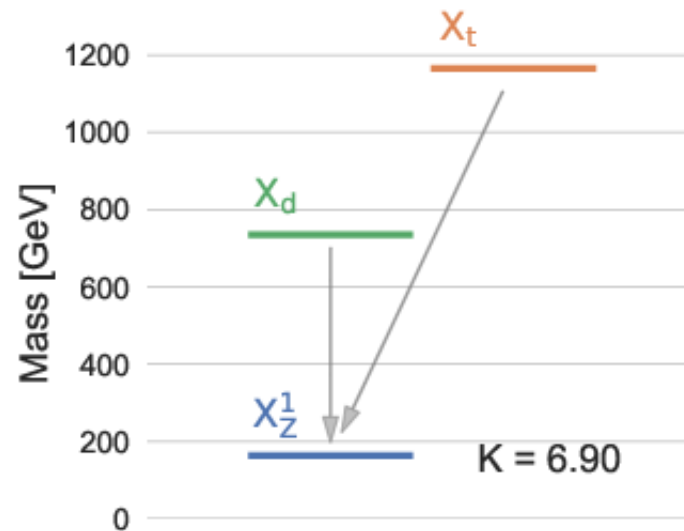
- Given the SModelS database of simplified models results, we **let a machine find the simplest possible model** that identifies **the largest possible violation of the Standard Model hypothesis** in the results, while evading all constraints from the negative search results in our database.
- Actually, we don't just want a single model, we want posteriori probabilities in these theory landscapes.
- The **models are allowed to be "incomplete"**, we want precursor theories, "proto-models", whose construction is driven by data, not by abstract principles. UV-completing such models will be a separate step (of course also partly executed by machines)

PROTOMODELS



Instead of testing BSM scenarios one-by-one against the experimental data:

- identify potential dispersed signals in the slew of published LHC analyses
- build candidate “protomodels” from them.



PROTOMODELS

- Protomodels can be thought of as **consistent sets of simplified models**.
- **Caveat:** The (variable-, or trans-dimensional) protomodels space is **restricted** by the SModelS software and database: currently restricted to **models exhibiting a Z_2 symmetry** (i.e. SUSY- and UED-like):

particle	decay channels	particle	decay channels
X_q	$qX_Z^j, q'X_W^i, qX_g$	X_W^1	WX_Z^j
X_t^1	$tX_Z^j, bX_W^i, WX_b^1, tX_g$	X_W^2	WX_Z^j, ZX_W^1, hX_W^1
X_b^1	$bX_Z^j, tX_W^i, WX_t^1, bX_g$	$X_Z^{j \neq 1}$	WX_W^i, ZX_Z^k, hX_Z^k
X_t^2	$tX_Z^j, bX_W^i, ZX_t^1, WX_b^1, tX_g$	X_ℓ	$\ell X_Z^j, \nu_\ell X_W^i$
X_b^2	$bX_Z^j, tX_W^i, ZX_b^1, WX_t^1, bX_g$	X_{ν_ℓ}	$\nu_\ell X_Z^j, \ell X_W^i$
X_g	$q\bar{q}X_Z^i, q\bar{q}'X_W^i, b\bar{b}X_Z^i, t\bar{t}X_Z^j, btX_W^i, qX_q, bX_b, tX_t$		

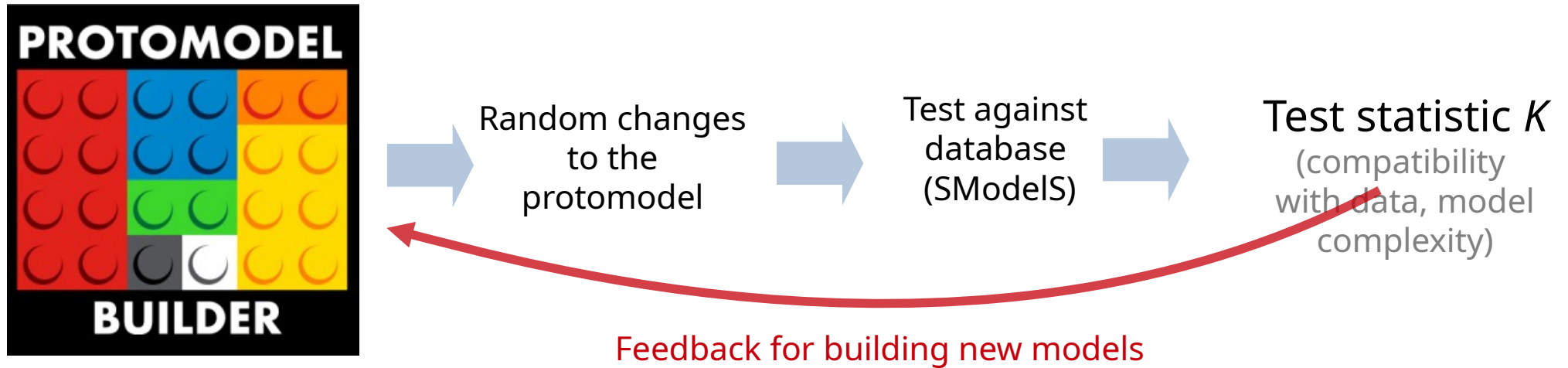
X (“Xeno”-)
particles, X_q
is squark-like,
 X_Z is
neutralino-
like, etc

Ongoing work:

Next iteration will likely not anymore require a Z_2 symmetry

[Andre Lessa, UFABC Brazil]

BUILDING THE NEXT STANDARD MODEL

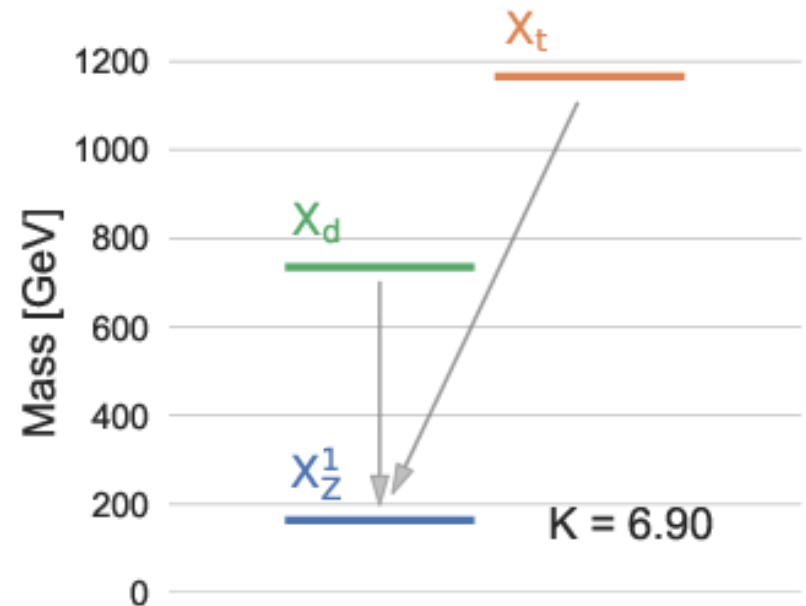


➔ “MCMC-type walk” over model+parameter space

➔ After many iterations/steps, the builder “learns” the best BSM model

BUILDING THE NEXT STANDARD MODEL

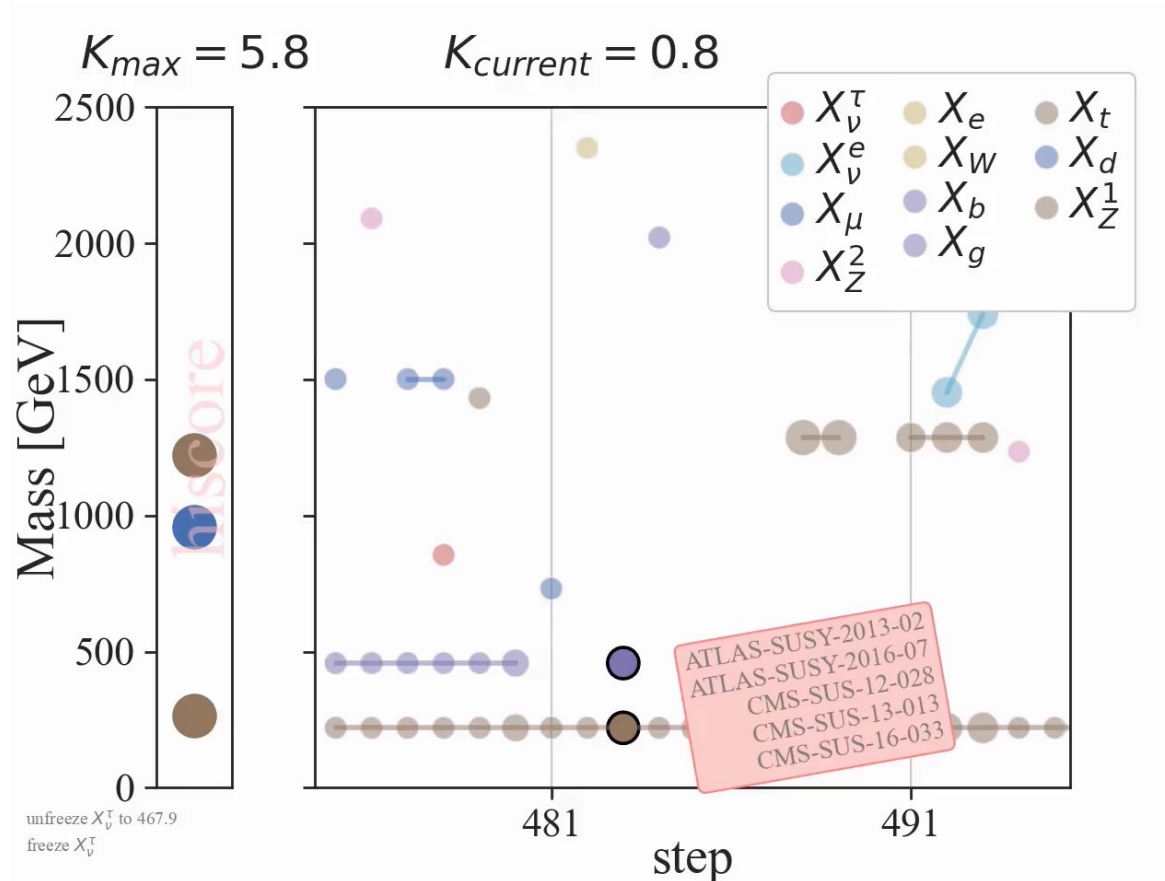
- In each step of this random walk, the following changes to the existing protomodel are allowed:
 - randomly add or remove a particle
 - randomly change a branching ratio, or the mass of a particle
 - randomly change a production cross section of a particle
- after each step a **test statistic K** is computed that quantifies how well the protomodel describes the data.
- K got much worse? → Revert to old protomodel
- K stayed the same or got better? → keep new protomodel



OUR APPROACH

the test statistic

Particle spectra



A hiscore
protomodel

Random
modifications

potential
dispersed
signals

an MCMC-like walk

THE TEST STATISTIC

The test statistic K^c is a likelihood-ratio test that quantifies how much better the proto-model describes the data than the Standard-Model (plus a penalty for model complexity).

Quantifies violation of Standard Model hypothesis

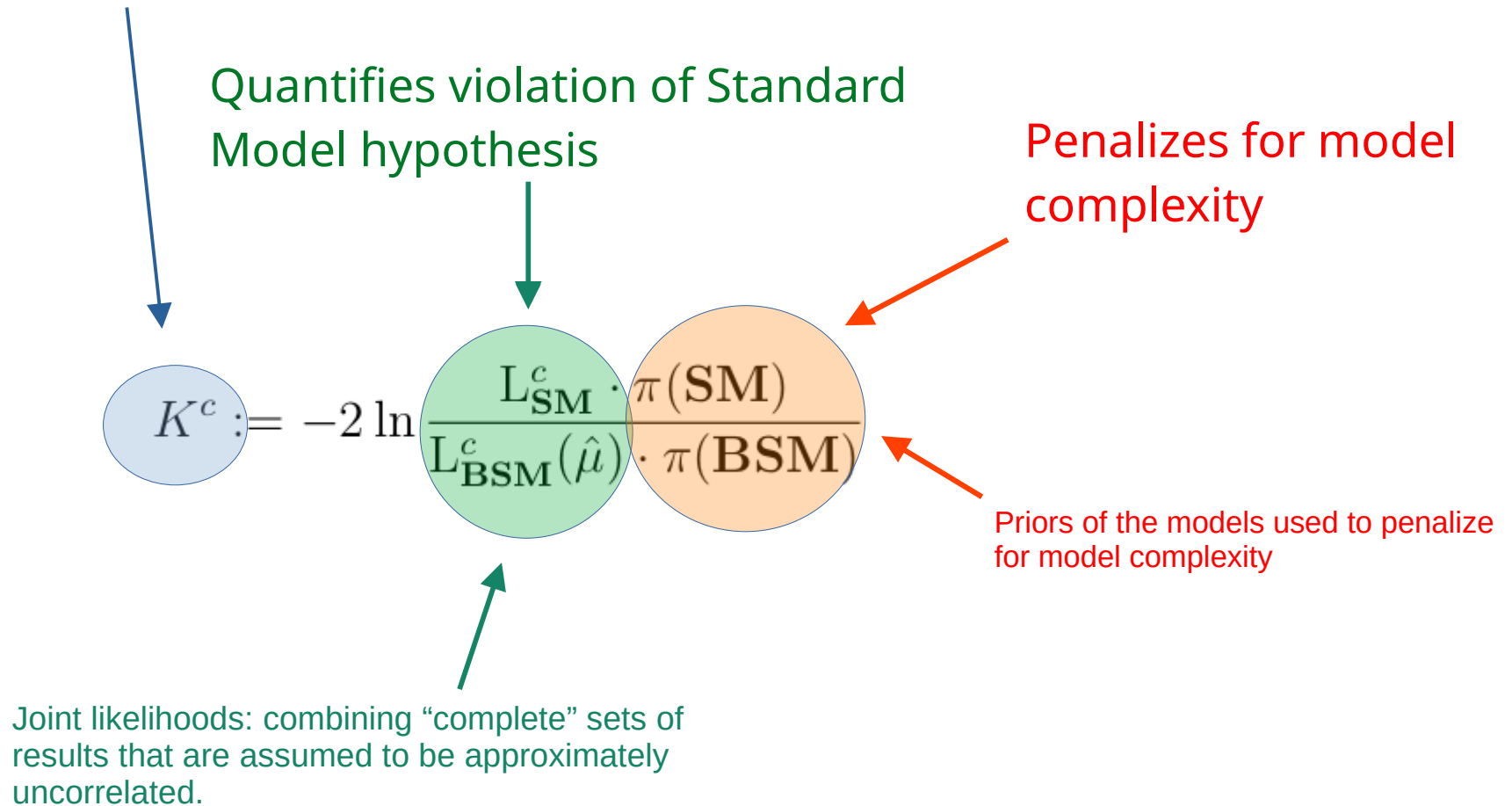
Penalizes for model complexity

$$K^c := -2 \ln \frac{L_{\text{SM}}^c \cdot \pi(\text{SM})}{L_{\text{BSM}}^c(\hat{\mu}) \cdot \pi(\text{BSM})}$$

We search for proto-models and combinations of results / likelihoods that maximize K^c *while remaining compatible with all negative results in our database.*

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INPUT DATA

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- likelihood computation based on simplified models results in SModels database
- vast number efficiency and upper limit maps from **~ 50 CMS and ~ 50 ATLAS publications.**
- Assume simplified statistical models “behind” the data → simplified likelihoods

#	ID	Short Description	Type	\mathcal{L} [fb ⁻¹]
1	CMS-PAS-EXO-16-036	hscp search	ul, eff	12.9
2	CMS-PAS-SUS-16-052	soft l, <= 2 jets	ul, eff	35.9
3	CMS-SUS-16-009	multijets + \cancel{E}_T , top tagging	ul	2.3
4	CMS-SUS-16-032	Sbottom and compressed stop	ul	35.9
5	CMS-SUS-16-033	0 ℓ + jets + \cancel{E}_T	ul, eff	35.9
6	CMS-SUS-16-034	2 OSSF l's	ul	35.9
7	CMS-SUS-16-035	2 SS l's	ul	35.9
8	CMS-SUS-16-036	0 ℓ + jets + \cancel{E}_T	ul	35.9
9	CMS-SUS-16-037	1 ℓ + jets + \cancel{E}_T with MJ	ul	35.9
10	CMS-SUS-16-039	multi-l EWK searches	ul	35.9
11	CMS-SUS-16-041	multi-ls + jets + \cancel{E}_T	ul	35.9
12	CMS-SUS-16-042	1 ℓ + jets + \cancel{E}_T	ul	35.9
13	CMS-SUS-16-043	EWK WH	ul	35.9
14	CMS-SUS-16-045	Sbottom to bHbH and H → $\gamma\gamma$	ul	35.9
15	CMS-SUS-16-046	γ + \cancel{E}_T	ul	35.9
16	CMS-SUS-16-047	γ + HT	ul	35.9
17	CMS-SUS-16-049	All hadronic stop	ul	35.9
18	CMS-SUS-16-050	0 ℓ + top tag	ul	35.9
19	CMS-SUS-16-051	1 ℓ stop	ul	35.9
20	CMS-SUS-17-001	Stop search in dil + jets + \cancel{E}_T	ul	35.9
21	CMS-SUS-17-003	2 taus + \cancel{E}_T	ul	35.9
22	CMS-SUS-17-004	EW-ino combination	ul	35.9
23	CMS-SUS-17-005	1 ℓ + multijets + \cancel{E}_T , top tagging	ul	35.9
24	CMS-SUS-17-006	jets + boosted H(bb) + \cancel{E}_T	ul	35.9
25	CMS-SUS-17-009	SFOS l's + \cancel{E}_T	ul	35.9
26	CMS-SUS-17-010	2L stop	ul	35.9
27	CMS-SUS-18-002	γ , jets, b-jets + \cancel{E}_T , top tagging	ul	35.9
28	CMS-SUS-19-006	0 ℓ + jets, MHT	ul	137.0
18	CMS-SUS-14-021	soft l's, low n _{jets} , high \cancel{E}_T	ul	19.7

#	ID	Short Description	Type	\mathcal{L} [fb ⁻¹]
1	ATLAS-SUSY-2015-01	2 b-jets + \cancel{E}_T	ul	3.2
2	ATLAS-SUSY-2015-02	single l stop	ul, eff	3.2
3	ATLAS-SUSY-2015-06	0 l's + 2-6 jets + \cancel{E}_T	eff	3.2
4	ATLAS-SUSY-2015-09	jets + 2 SS l's or >=3 l's	ul	3.2
5	ATLAS-SUSY-2016-07	0 ℓ + jets + \cancel{E}_T	ul, eff	36.1
6	ATLAS-SUSY-2016-14	2 SS or 3 l's + jets + \cancel{E}_T	ul	36.1
7	ATLAS-SUSY-2016-15	0 ℓ stop	ul	36.1
8	ATLAS-SUSY-2016-16	1 ℓ stop	ul, eff	36.1
9	ATLAS-SUSY-2016-17	2 opposite sign l's + \cancel{E}_T	ul	36.1
10	ATLAS-SUSY-2016-19	stops to staus	ul	36.1
11	ATLAS-SUSY-2016-24	2-3 l's + \cancel{E}_T , EWino	ul, eff	36.1
12	ATLAS-SUSY-2016-26	>=2 c jets + \cancel{E}_T	ul	36.1
13	ATLAS-SUSY-2016-27	jets + γ + \cancel{E}_T	ul, eff	36.1
14	ATLAS-SUSY-2016-28	2 b-jets + \cancel{E}_T	ul	36.1
15	ATLAS-SUSY-2016-33	2 OSSF l's + \cancel{E}_T	ul	36.1
16	ATLAS-SUSY-2017-01	EWK WH(bb) + \cancel{E}_T	ul	36.1
17	ATLAS-SUSY-2017-02	0 ℓ + jets + \cancel{E}_T	ul	36.1
18	ATLAS-SUSY-2017-03	multi-l EWK searches	ul	36.1
19	ATLAS-SUSY-2018-04	2 hadronic taus	ul, eff	139.0
20	ATLAS-SUSY-2018-06	3 l's EW-ino	ul	139.0
21	ATLAS-SUSY-2018-31	2b + 2H(bb) + \cancel{E}_T	ul, eff	139.0
22	ATLAS-SUSY-2018-32	2 OS l's + \cancel{E}_T	ul	139.0
23	ATLAS-SUSY-2019-08	1 ℓ + higgs + \cancel{E}_T	ul, eff	139.0
14	ATLAS-SUSY-2013-19	2 OS l's + (b-)jets + \cancel{E}_T	ul	20.3
15	ATLAS-SUSY-2013-21	monojet or c-jet + \cancel{E}_T	eff	20.3
16	ATLAS-SUSY-2013-23	1 ℓ + 2 b-jets (or 2 γ s) + \cancel{E}_T	ul	20.3
17	ATLAS-SUSY-2014-03	>= 2(c-)jets + \cancel{E}_T	eff	20.3

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Ongoing work for next iteration:
 moving to full likelihoods, **machine-learning surrogate models** (multi-layer perceptrons, normalizing flows) as accelerators

[Humberto Gonzales, PostDoc in Genova]

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27	CMS-SUS-18-002	γ , jets, b-jets + \cancel{E}_T , top tagging	ul	35.9		16	ATLAS-SUSY-2013-23	1 ℓ + 2 b-jets (or 2 γ s) + \cancel{E}_T	ul	20.3	
28	CMS-SUS-19-006	0 ℓ + jets, MHT	ul	137.0		17	ATLAS-SUSY-2014-03	≥ 2 (c-)jets + \cancel{E}_T	eff	20.3	
18	CMS-SUS-14-021	soft l's, low n_{jets} , high \cancel{E}_T	ul	19.7							

Ongoing work for next iteration:
 moving to full likelihoods, **machine-learning surrogate models** (multi-layer perceptrons, normalizing flows) as accelerators
 [Humberto Gonzales, PostDoc in Genova]

Also: started work to include measurements (via rivet) in addition to these searches
 [Jamie Yellen, PhD student (Glasgow)]

THE COMBINER

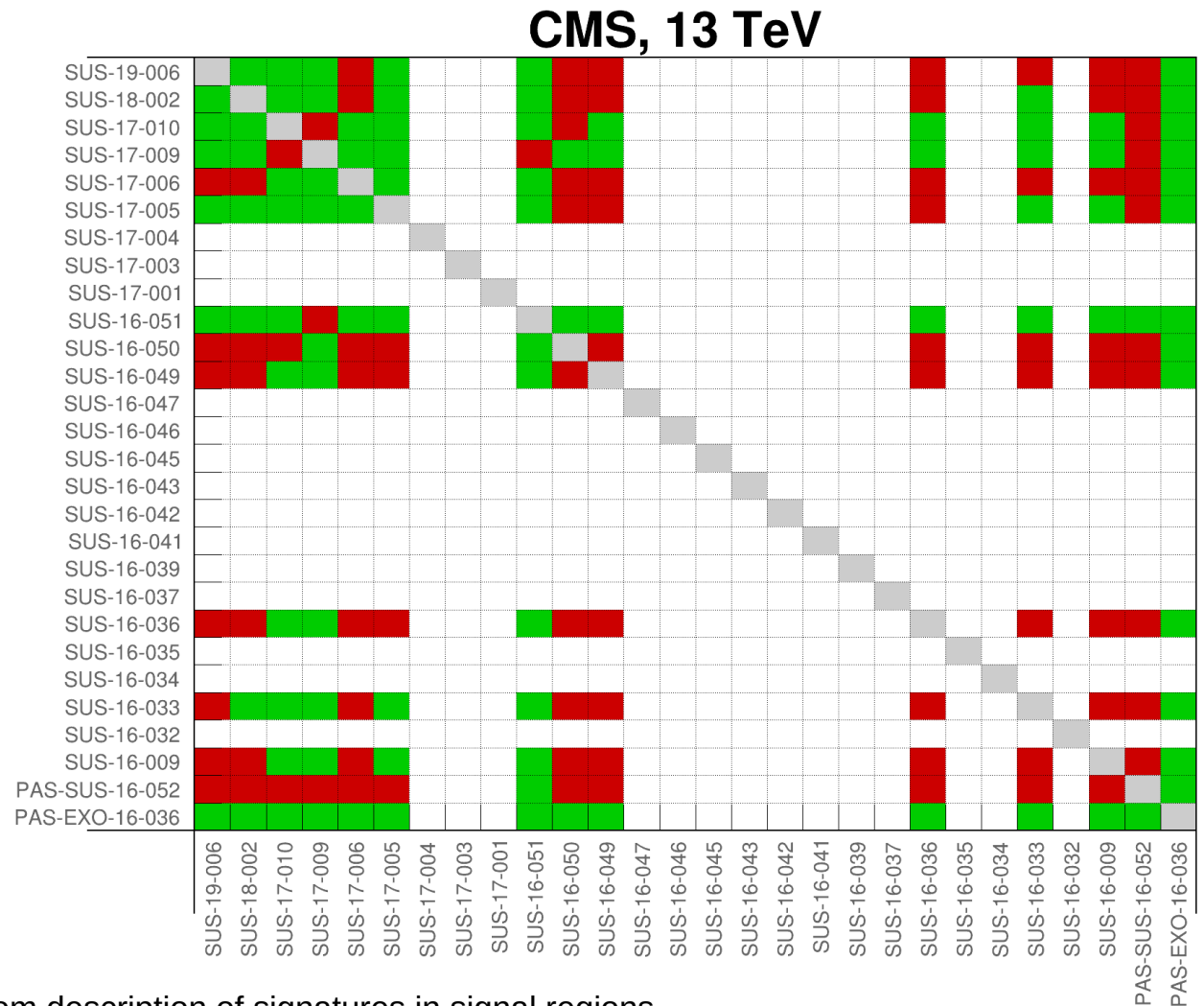
As we are chasing dispersed signals, we need to allow the machine to combine (i.e. multiply) likelihoods. Simplified, binaric “inter-analyses *exclusivity* matrix”:

green:
approximately
uncorrelated
→ combinable

red: correlated,
not combinable

White: cannot
construct a
likelihood

Signal regions
within each
analysis:
correlated



In this publication: “educated guesses” from description of signatures in signal regions.

THE COMBINER

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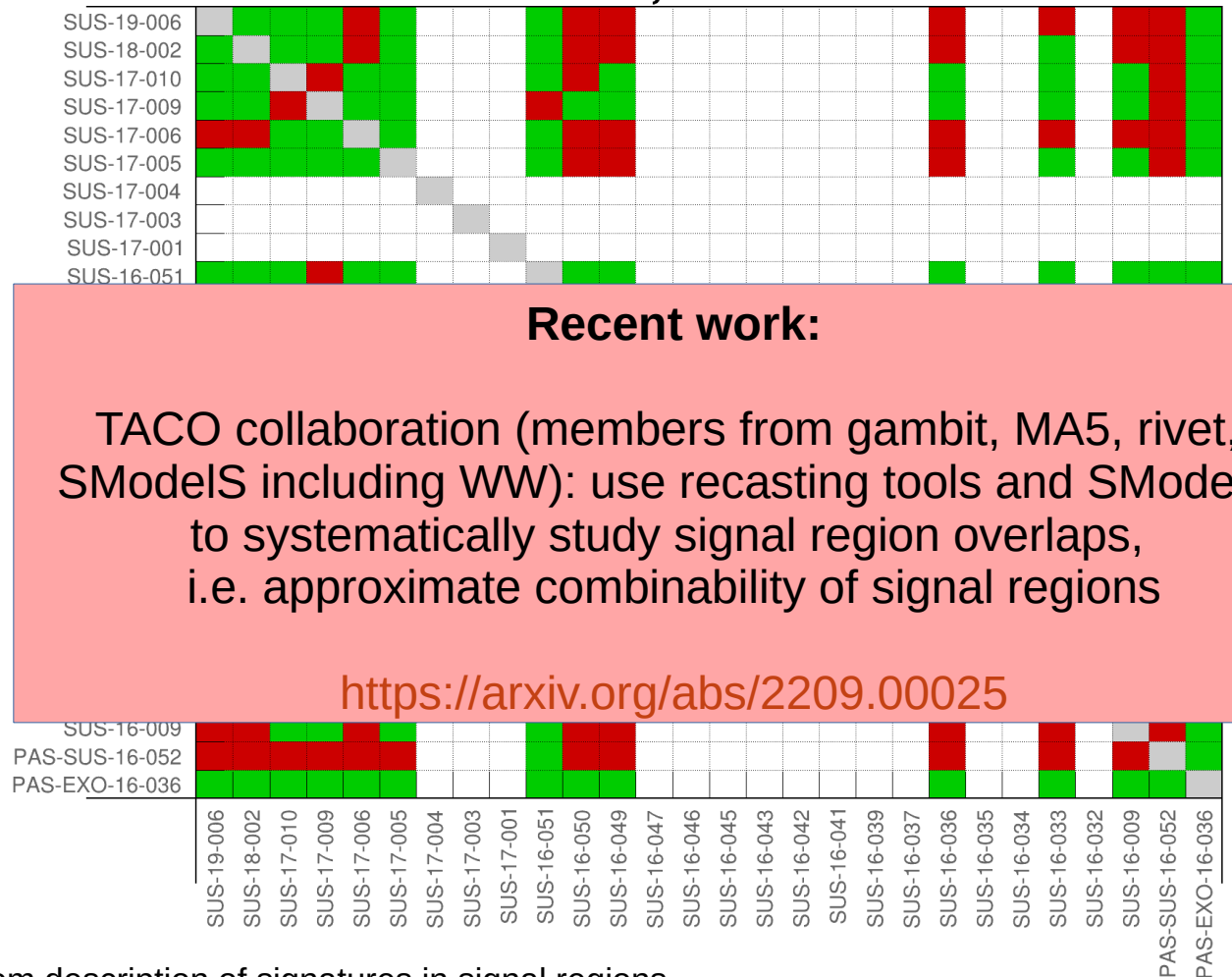
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construct a
likelihood

Signal regions
within each
analysis:
correlated

CMS, 13 TeV



In this publication: “educated guesses” from description of signatures in signal regions.

THE PENALTY TERM

For every legal combination, we define a test statistic K^c

$$K^c := -2 \ln \frac{L_{\text{SM}}^c \cdot \pi(\text{SM})}{L_{\text{BSM}}^c(\hat{\mu}) \cdot \pi(\text{BSM})}$$

$\pi(\text{BSM})$ is the prior of the BSM model. We use it to “regularize” the model, i.e. impose the *law of parsimony*:

$$\pi(\text{BSM}) \approx \exp[-n_{\text{particles}}^{\text{BSM}}]$$

Resulting in a test statistic that resembles an “Akaike Information Criterion” (AIC):

$$K^c \approx \Delta\chi^2 - 2n_{\text{particles}}^{\text{BSM}}$$

An additional BSM particle will have to increase the (delta-)chi-square by approximately two units.

THE PENALTY TERM

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Resulting in a test statistic that r

**An additional BSM particle
approximately two units.**

Ongoing work:

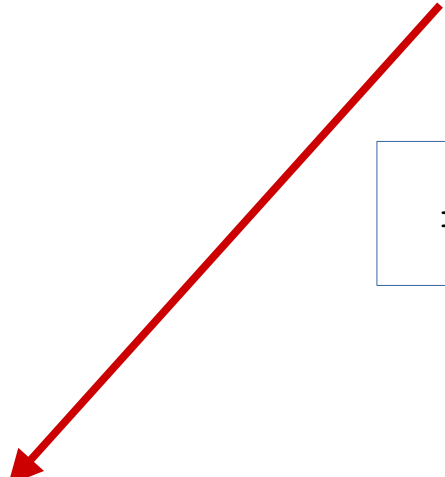
Extend penalty term of test statistic to not only correctly penalize for model complexity, but also for number of experimental results

With: Jamie Yellen (PhD student, Glasgow)

THE CRITIC

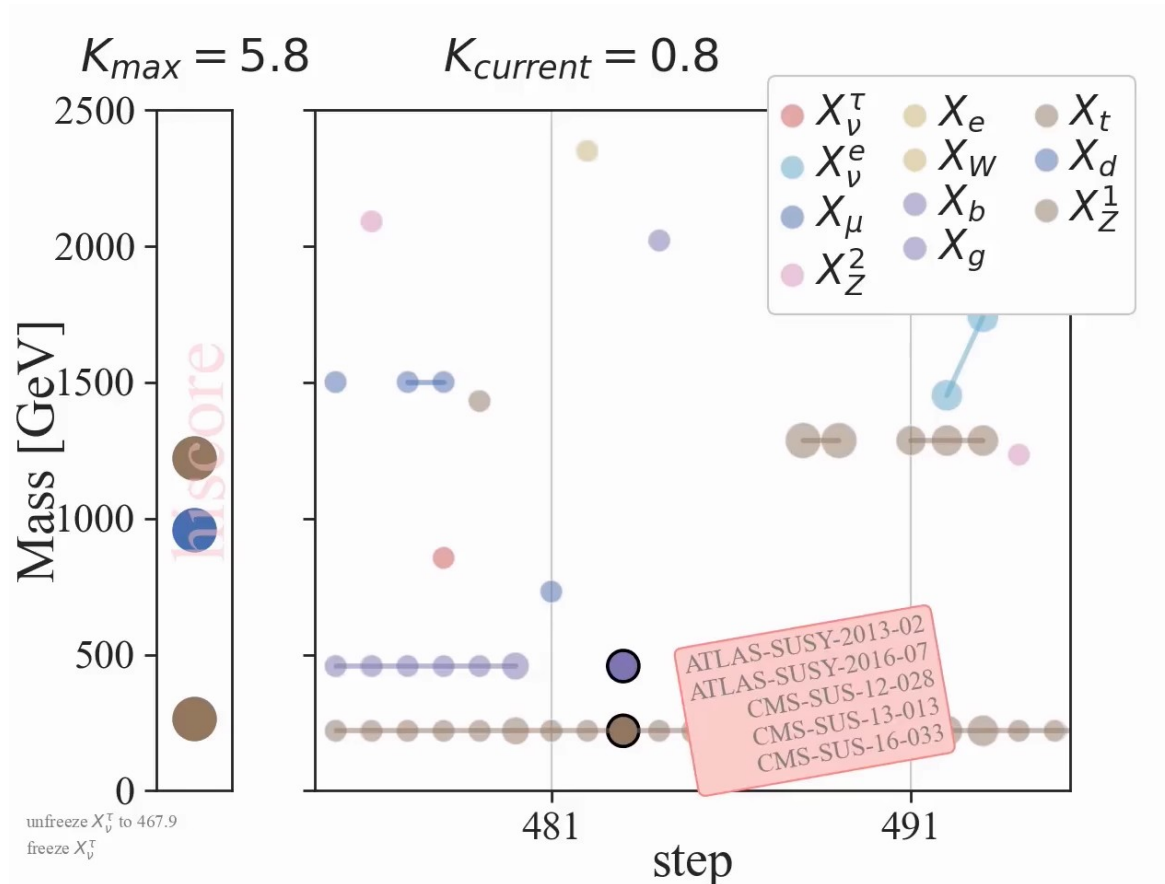
For every legal combination c , we define a test statistic K^c

$$K^c := -2 \ln \frac{L_{\text{SM}}^c \cdot \pi(\text{SM})}{L_{\text{BSM}}^c(\hat{\mu}) \cdot \pi(\text{BSM})}$$


$$\Rightarrow K = \max\{K^c \mid \forall \text{ combinations}\}$$

- $\hat{\mu}$ is the signal strength of the model that maximizes the likelihood.
- By limiting its support we guarantee **compatibility** with **all negative results** in the SModelS database.
- In allusion to adversarial setups, we also call this feature **the critic**

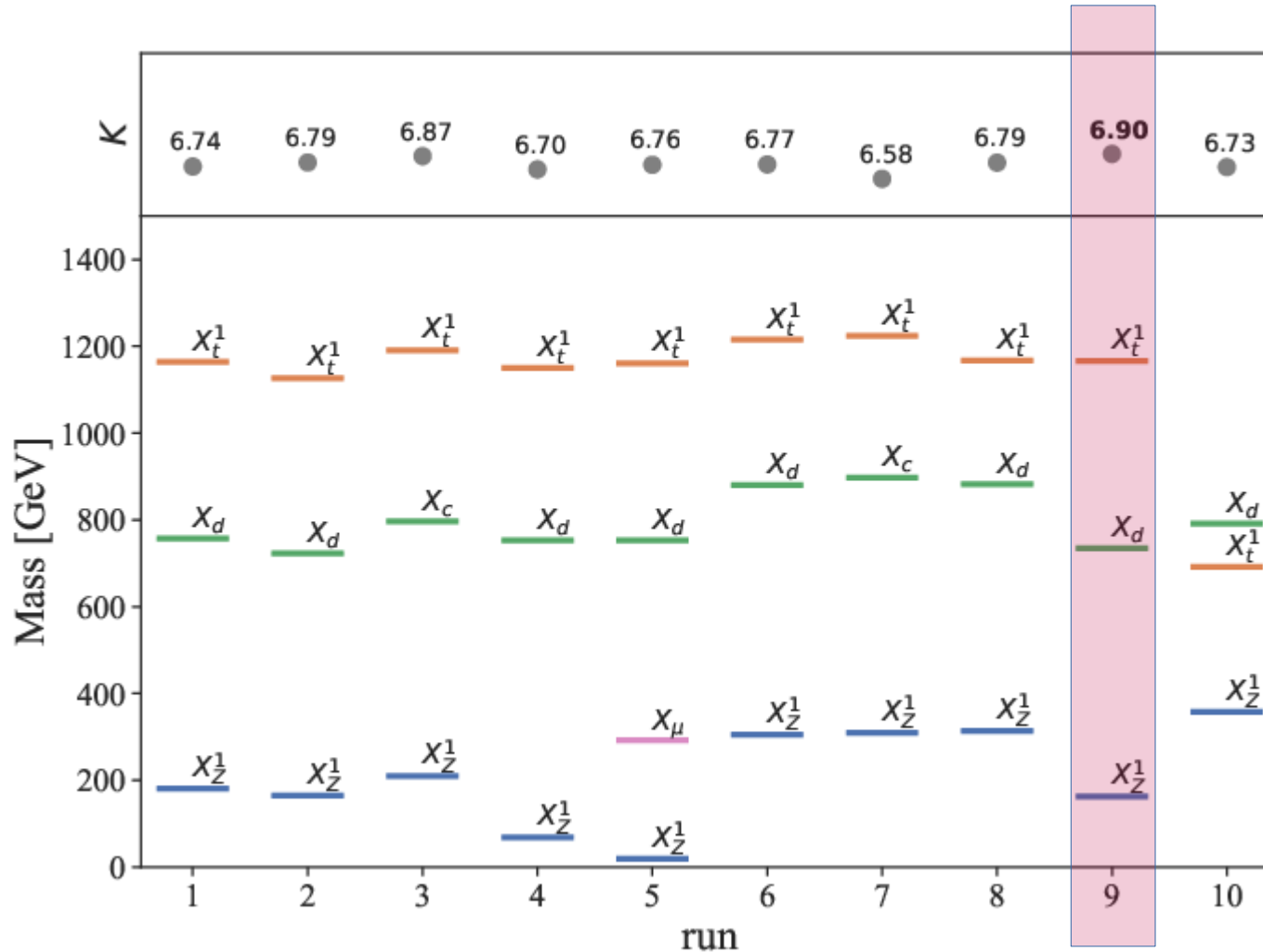
AND THEN WE RAN THE ALGORITHM ...



We defined a “run” as 50 parallel walkers, making 1,000 steps each.
We performed 10 such runs on the SModelS database.
Total computing resources spent: ~ 1,000,000 CPU hours

... AND OBTAINED RESULTS

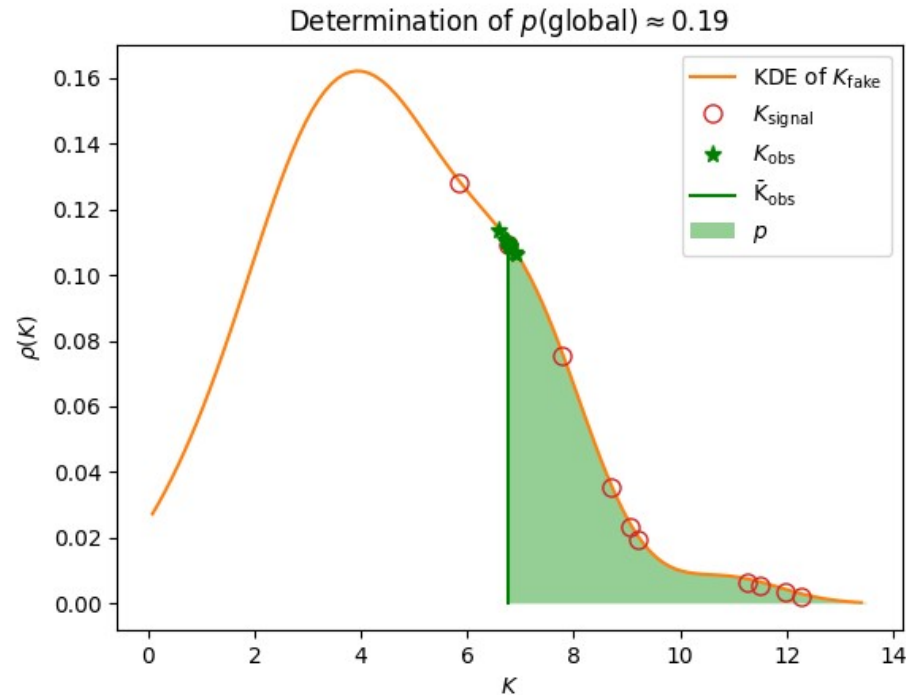
We performed 10 such runs on the SModelS database:



All 10 runs introduced a **top partner** as well as a **light quark partner**. The cross sections are compatible with values expected from the MSSM. The best test statistic was $K=6.9$.

GLOBAL p -VALUE

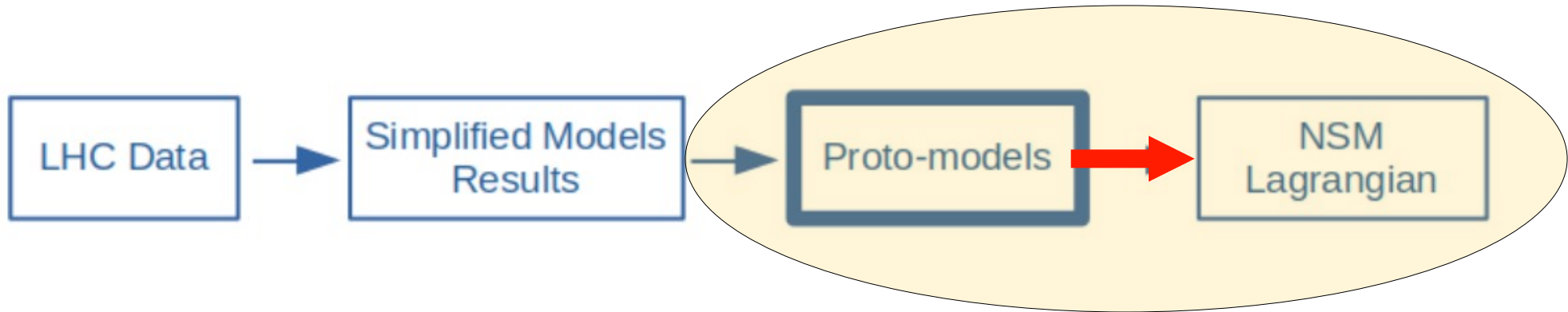
- Because we have statistical models of the search results, we can synthesize statistically correct databases of results that are “typical”, if no new physics is in the data.
- From this we can compute a p -value for the Standard Model hypothesis: that is the chances that – under the SM hypothesis – we would obtain a result as extreme as ours or more extreme.



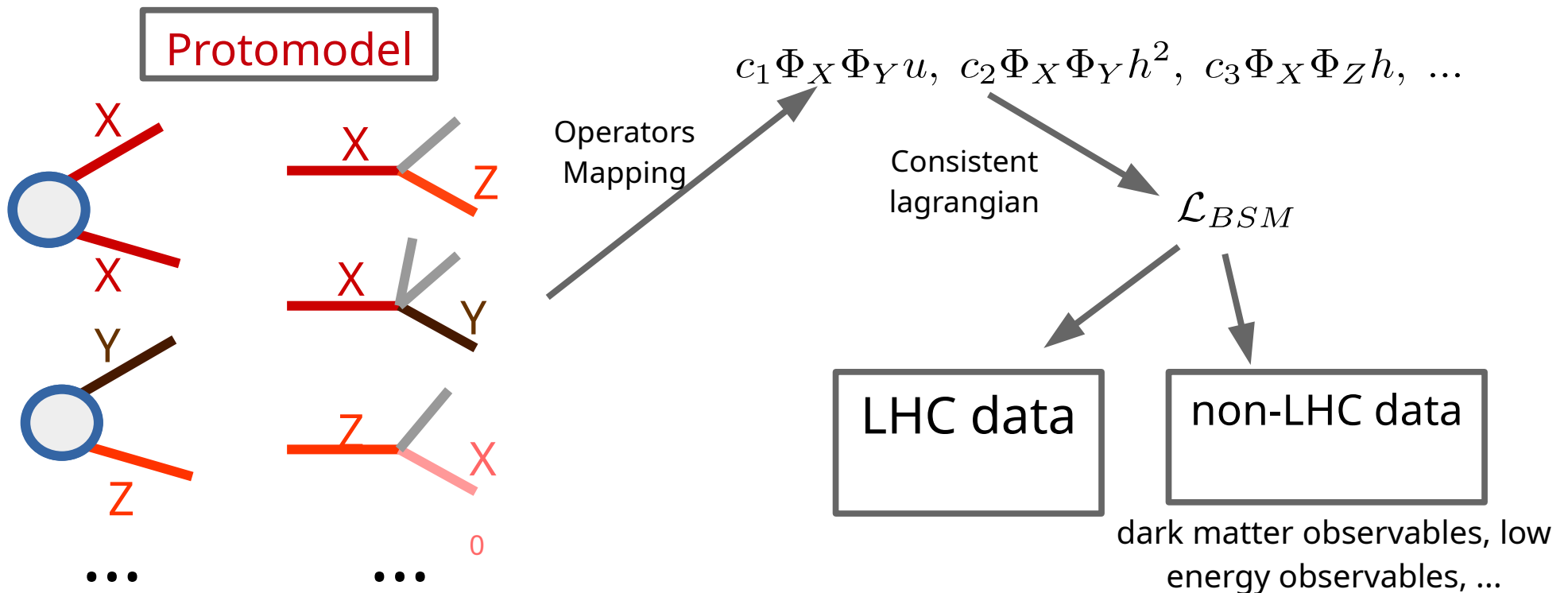
Since we did not correct for the conservativeness of the experimental results, we assume our result to also be conservative.

By construction, no Look-Elsewhere Effect applies. (within the database, the machine does look everywhere)

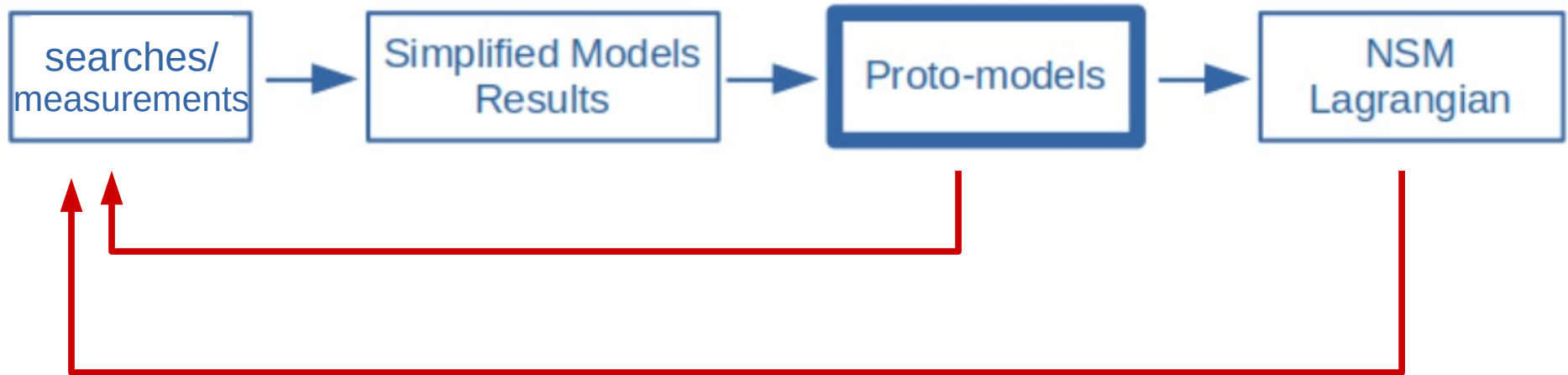
TOWARDS UV COMPLETION



Work from protomodels to UV complete theories has recently begun (John Gargalionis, PostDoc in Valencia) – lot's of combinatorics!



ITERATIONS

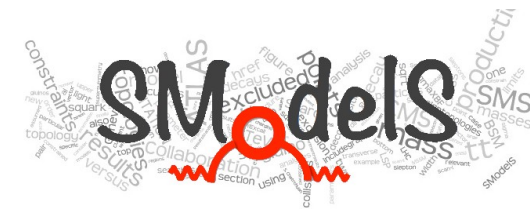


We of course envisage the procedure to be iterative:

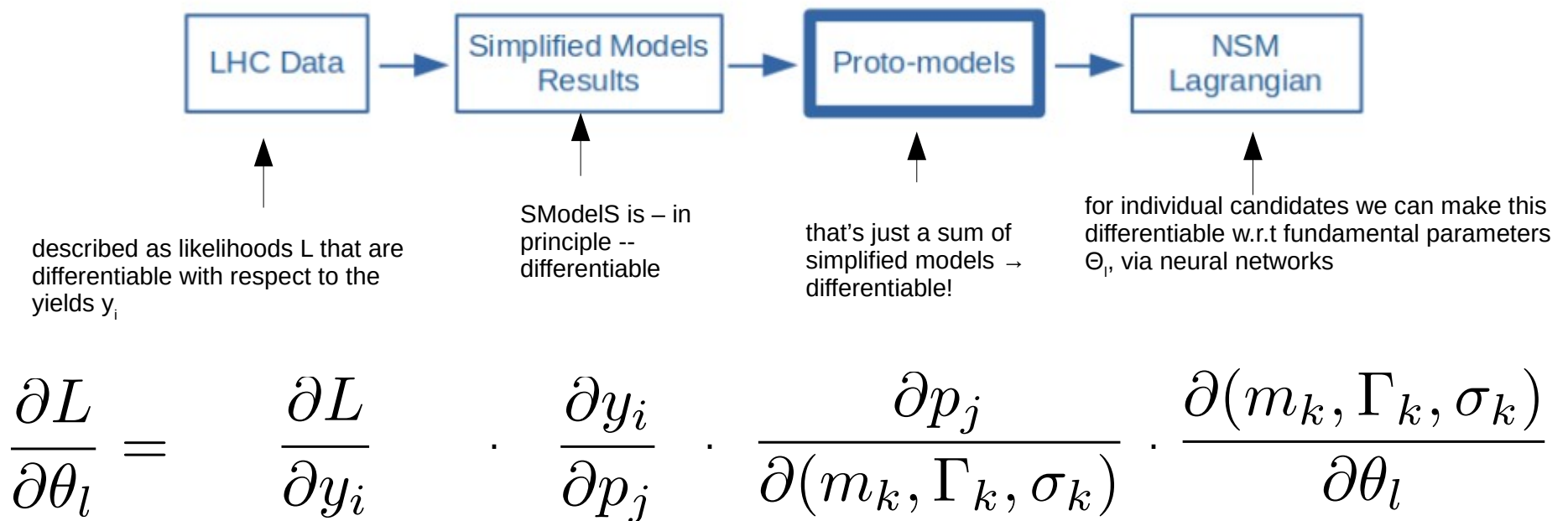
The **pheno** community can point out models and combinations of interest, interpret results, make predictions about potential channels, **provide feedback to the experiments**

Only the experimental communities can claim potential (dispersed) discoveries.

FUTURE: FULLY DIFFERENTIABLE CHAIN OF INDUCTIVE REASONING



If we had gradients we could perform gradient descent to find the best model, and we could use e.g. the Fisher information to infer the error on its parameters (or, alternatively we can then MCMC-sample).

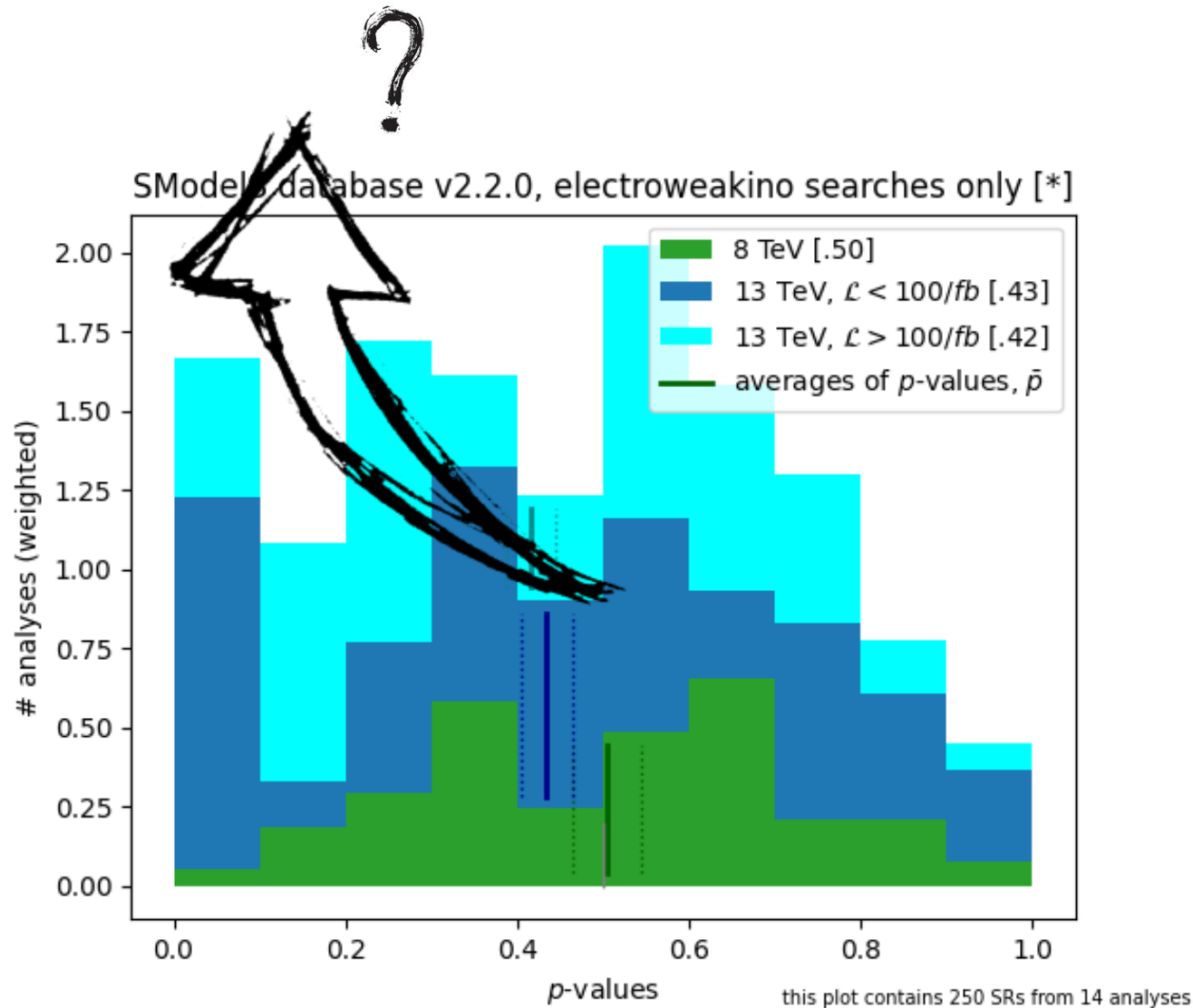


Not yet required (theory space as well as space of measurements are still low-dimensional enough).

SUMMARY, OUTLOOK

- In light of no clear evidence for new physics in the individual channels/results, a more global attempt at finding new physics seems appropriate
- **First prototype run** of a machine that builds protomodels with results from ~ 100 **analyses** resulted in **p-value of SM hypothesis of ~ 0.2** : a very small tension with the Standard Model hypothesis (but also some tension between some results)
- Working on next iteration with **more results, better likelihoods, surrogate models as accelerators, covering more signatures, larger protomodels space, UV completion**
- **Extend to measurements**, not only searches
- Can we make the entire chain **differentiable**?

RUN-3 WILL CONTINUE THE TRENDS OR REFUTE THEM



THE JURY IS OUT [*]

[*] no pun intended

BACKUP

A SITUATION UNLIKE IN THE PAST

ATLAS SUSY Searches* - 95% CL Lower Limits

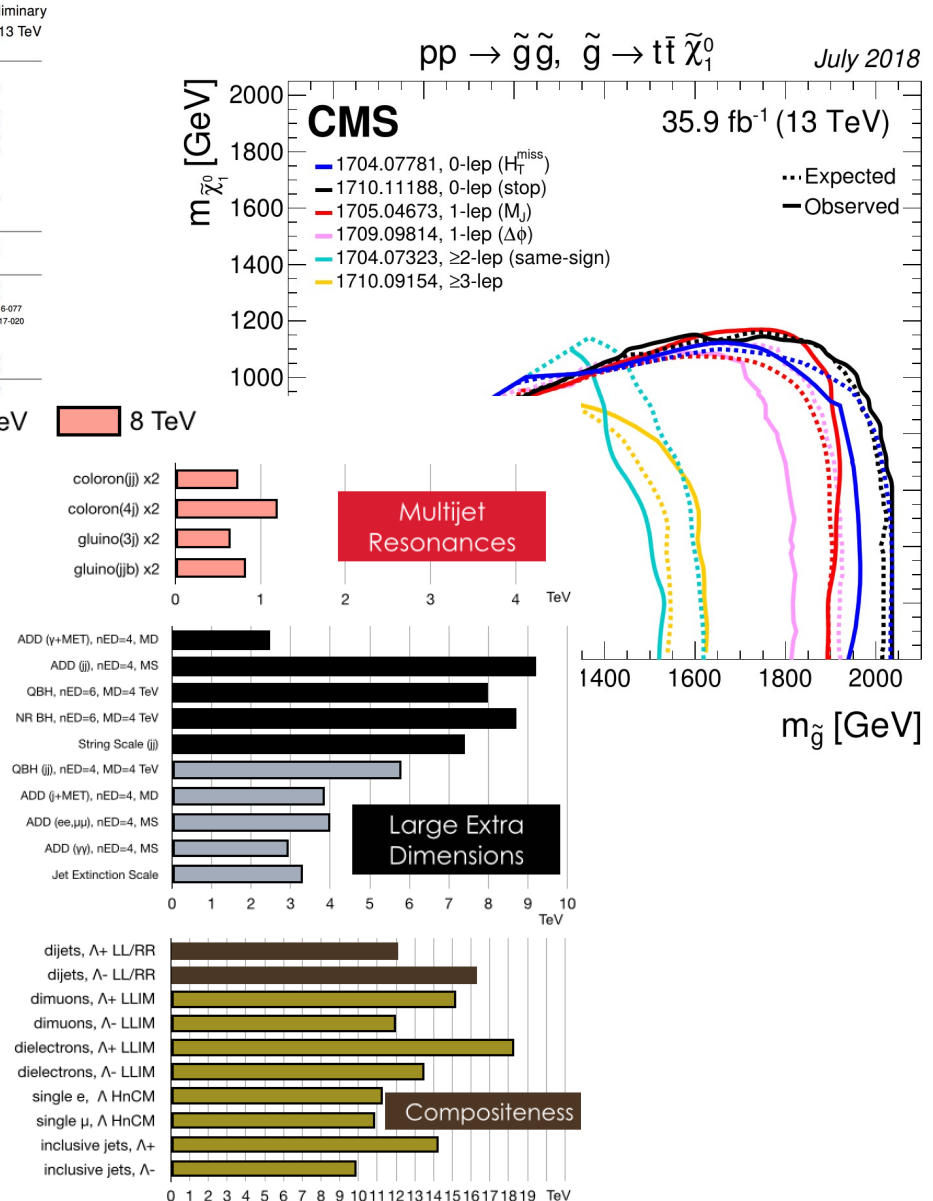
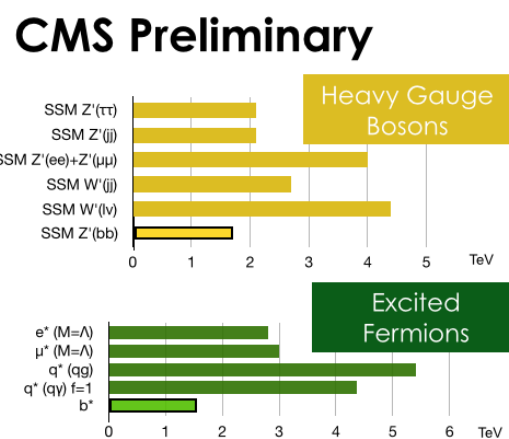
May 2017

ATLAS Preliminary $\sqrt{s} = 7, 8, 13 \text{ TeV}$

Model	e, μ, τ, γ	Jets	$E_{\text{miss}}^{\text{max}}$ [$\mathcal{L} d\Gamma(\text{fb}^{-1})$]	Mass limit	Reference
MSUGRA/CMSSM	$0-3 e, \mu, 1-2 \tau$	2-10 jets/3 b	Yes 20.3	1.85 TeV	m(\tilde{g})=m(\tilde{u})
$\tilde{g}, \tilde{u}, \tilde{d}$	0	2-6 jets	Yes 36.1	1.57 TeV	m(\tilde{t})<200 GeV, m($\tilde{1}^{\text{st}}$ gen. \tilde{g})+m($\tilde{2}^{\text{nd}}$ gen. \tilde{g})
$\tilde{g}, \tilde{u}, \tilde{d}$ (compressed)	mono-jet	1-6 jets	Yes 3.2	608 GeV	m(\tilde{g})=m(\tilde{t})-5 GeV
$\tilde{g}, \tilde{u}, \tilde{d}$	0	2-6 jets	Yes 36.1	2.02 TeV	m(\tilde{t})<200 GeV
$\tilde{g}, \tilde{u}, \tilde{d}$	0	2-6 jets	Yes 36.1	2.01 TeV	m(\tilde{t})<200 GeV, m($\tilde{1}^{\text{st}}$)>0.5m(\tilde{t})+m(\tilde{g})
$\tilde{g}, \tilde{u}, \tilde{d}$	3 e, μ	4 jets	Yes 36.1	1.925 TeV	m(\tilde{t})<400 GeV
$\tilde{g}, \tilde{u}, \tilde{d}$	0	7-11 jets	Yes 36.1	1.8 TeV	m(\tilde{t})<400 GeV
GMSB (\tilde{L} NLSP)	$1-2 \tau + 0-1 \ell$	0-2 jets	Yes 3.2	2.0 TeV	c(NLSP)<0.1 mm
GGM (bino NLSP)	2 γ	Yes 3.2	1.85 TeV	c(NLSP)<0.1 mm, $\mu < 0$	
GGM (higgsino-bino NLSP)	γ	1 b	Yes 20.3	1.37 TeV	m($\tilde{0}$)<680 GeV, c(NLSP)<0.1 mm, $\mu < 0$
GGM (higgsino-bino NLSP)	γ	2 jets	Yes 13.3	1.8 TeV	m(NLSP)<430 GeV
GGM (higgsino NLSP)	$2 e, \mu$ (Z)	2 jets	Yes 20.3	900 GeV	m($\tilde{0}$)>1.8 x 10 ⁻⁴ eV, m(\tilde{g})=m(\tilde{g})+1.5 TeV
Gravitino LSP	0	mono-jet	Yes 20.3	865 GeV	$\mu < 0$
$\tilde{g}, \tilde{u}, \tilde{d}$	0	3 b	Yes 36.1	1.92 TeV	m(\tilde{t})<600 GeV
$\tilde{g}, \tilde{u}, \tilde{d}$	0-1 e, μ	3 b	Yes 36.1	1.97 TeV	m(\tilde{t})<200 GeV
$\tilde{g}, \tilde{u}, \tilde{d}$	0-1 e, μ	3 b	Yes 20.1	1.37 TeV	m(\tilde{t})<300 GeV
$\tilde{b}_1, \tilde{b}_1, \tilde{b}_1 \rightarrow \tilde{b}_1^0 b$	0	2 b	Yes 36.1	950 GeV	m(\tilde{t})<420 GeV
$\tilde{b}_1, \tilde{b}_1, \tilde{b}_1 \rightarrow \tilde{b}_1^0 b$	2 e, μ (SS)	1 b	Yes 36.1	275-700 GeV	m(\tilde{t})<200 GeV, m($\tilde{1}^{\text{st}}$)>100 GeV
$\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{b}_1^0 b$	0-2 e, μ	1-2 b	Yes 4.7/13.3	117-170 GeV	m(\tilde{t})>200 GeV, m($\tilde{1}^{\text{st}}$)>100 GeV
$\tilde{t}_1, \tilde{t}_1 \rightarrow W b$ or $\tilde{b}_1^0 b$	0-2 e, μ	0-2 jets/1-2 b	Yes 20.3/36.1	90-198 GeV	m(\tilde{t})>200 GeV, m($\tilde{1}^{\text{st}}$)>100 GeV
$\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{c} b$	0	mono-jet	Yes 3.2	90-323 GeV	m(\tilde{t})>1 GeV
$\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{c} b$	0	mono-jet	Yes 3.2	150-600 GeV	m(\tilde{t})>5 GeV
\tilde{t}_1, \tilde{t}_1 (natural GMSB)	2 e, μ (Z)	1 b	Yes 20.3	290-790 GeV	m(\tilde{t})>150 GeV
$\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{t} + Z$	3 e, μ (Z)	1 b	Yes 36.1	320-880 GeV	m(\tilde{t})>0 GeV
$\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{t} + h$	1-2 e, μ	1 b	Yes 36.1	320-880 GeV	m(\tilde{t})>0 GeV
$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1 \rightarrow \tilde{t} b$	2 e, μ	0	Yes 36.1	90-440 GeV	m(\tilde{t})>0
$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1 \rightarrow \tilde{t} b$	2 e, μ	0	Yes 36.1	90-440 GeV	m(\tilde{t})>0
$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1 \rightarrow \tilde{t} b$	2 τ	-	Yes 36.1	270 GeV	
$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1 \rightarrow \tilde{t} b$	3 e, μ	0	Yes 36.1	270 GeV	
$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1 \rightarrow \tilde{t} b$	2-3 e, μ	0-2 jets	Yes 36.1	270 GeV	
$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1 \rightarrow \tilde{t} b$	e, μ, τ	0-2 b	Yes 20.3	270 GeV	
$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1 \rightarrow \tilde{t} b$	4 e, μ	0	Yes 20.3	270 GeV	
GGM (wino NLSP) weak prod. $\tilde{X}_1^0 \rightarrow \gamma \tilde{G}$	1 e, $\mu + \gamma$	-	Yes 20.3	115-371	
GGM (bino NLSP) weak prod. $\tilde{X}_1^0 \rightarrow \gamma \tilde{G}$	2 γ	-	Yes 20.3	115-371	
Direct $\tilde{X}_1^0 \tilde{X}_1^0$ prod., long-lived \tilde{X}_1^0	Disapp. trk	1 jet	Yes 36.1		
Direct $\tilde{X}_1^0 \tilde{X}_1^0$ prod., long-lived \tilde{X}_1^0	dE/dx trk	-	Yes 18.4		
Stable, stopped \tilde{g} R-hadron	0	1-5 jets	Yes 27.9		
Stable \tilde{g} R-hadron	trk	-	Yes 3.2		
Metastable \tilde{g} R-hadron	dE/dx trk	-	Yes 3.2		
GMSB, stable \tilde{g} R-hadron $\tilde{X}_1^0 \rightarrow \tilde{g} \tilde{G}$, long-lived \tilde{X}_1^0	1-2 μ	-	Yes 19.1		
GMSB, $\tilde{X}_1^0 \rightarrow \tilde{g} \tilde{G}$, long-lived \tilde{X}_1^0	2 γ	-	Yes 20.3		
GGM $\tilde{X}_1^0 \rightarrow \tilde{g} \tilde{G}$, long-lived \tilde{X}_1^0	displ. ee/ $\mu\mu$	-	Yes 20.3		
GGM $\tilde{X}_1^0 \rightarrow \tilde{g} \tilde{G}$, long-lived \tilde{X}_1^0	displ. vtx + jets	-	Yes 20.3		
LFV $\tilde{g} \tilde{g} \rightarrow \tilde{g} + X, \tilde{g} \tilde{g} \rightarrow \tilde{g} \mu \tau / \mu \tau$	e, μ, τ	-	Yes 3.2		
Bilinear RPV CMSSM	2 e, μ (SS)	0-3 b	Yes 20.3		
$\tilde{X}_1^0 \tilde{X}_1^0, \tilde{X}_1^0 \rightarrow W \tilde{X}_1^0, \tilde{X}_1^0 \rightarrow \tilde{e} \nu_e, \tilde{g} \mu \nu, \mu \nu$	4 e, μ	-	Yes 13.3		
$\tilde{X}_1^0 \tilde{X}_1^0, \tilde{X}_1^0 \rightarrow W \tilde{X}_1^0, \tilde{X}_1^0 \rightarrow \tilde{e} \nu_e, \tilde{g} \mu \nu, \mu \nu$	3 e, $\mu + \tau$	-	Yes 20.3		
$\tilde{g}, \tilde{u}, \tilde{d}$	0	4-5 large-R jets	Yes 14.8		
$\tilde{g}, \tilde{u}, \tilde{d}$	0	4-5 large-R jets	Yes 14.8		
$\tilde{g}, \tilde{u}, \tilde{d}$	1 e, μ	8-10 jets/0-4 b	Yes 36.1		
$\tilde{g}, \tilde{u}, \tilde{d}$	1 e, μ	8-10 jets/0-4 b	Yes 36.1		
$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1 \rightarrow \tilde{t} b$	0	2 jets + 2 b	Yes 15.4		
$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1 \rightarrow \tilde{t} b$	2 e, μ	2 b	Yes 36.1		
Other Scalar charm, $\tilde{c} \rightarrow \tilde{c}^0 \tilde{G}$	0	2 c	Yes 20.3		

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10⁻¹



CMS Exotica Physics Group Summary – ICHP 2016

Not a handful of experimental signatures. Hundreds of publications with a wide range of signatures!

THE INVERSE PROBLEM

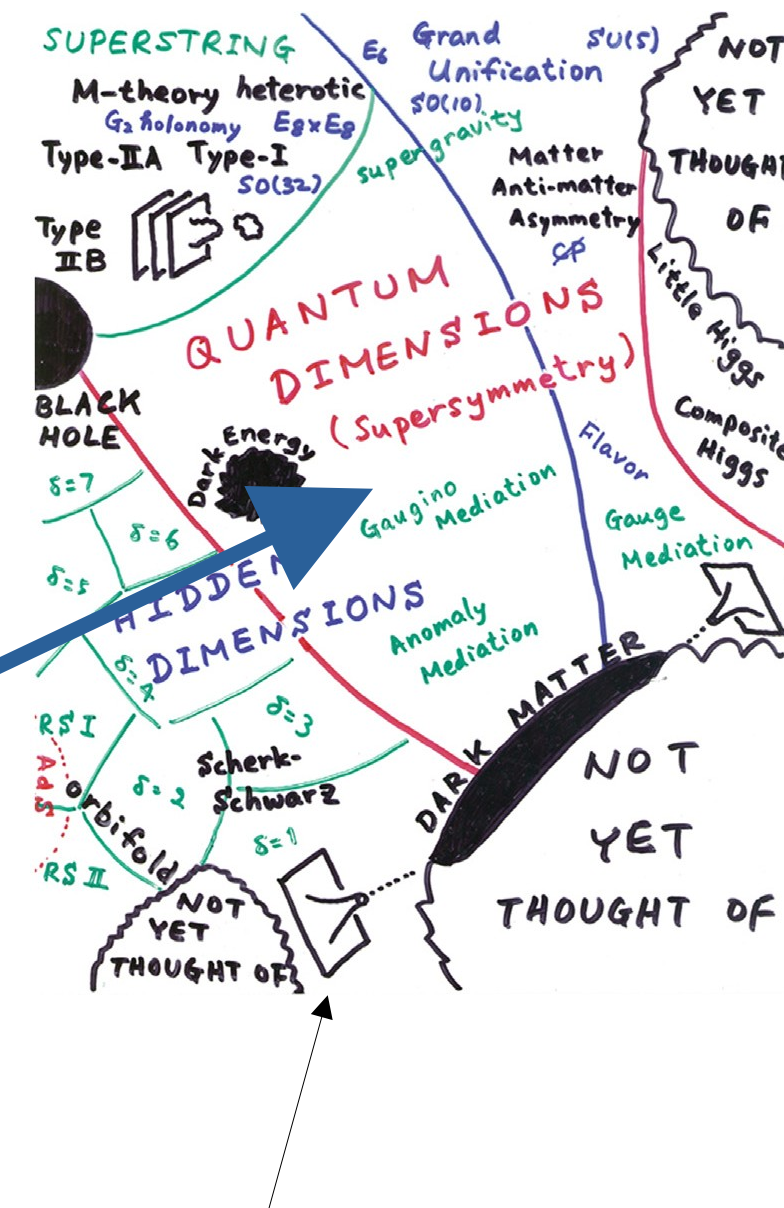
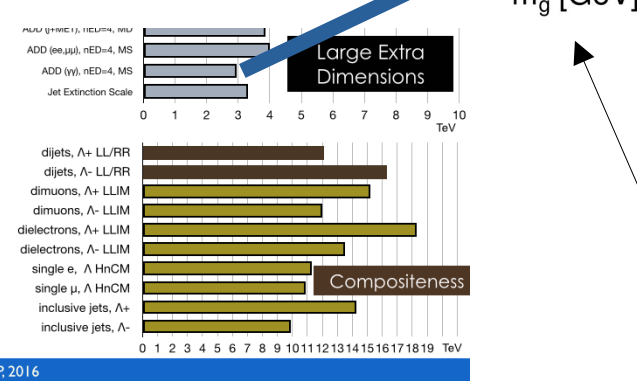
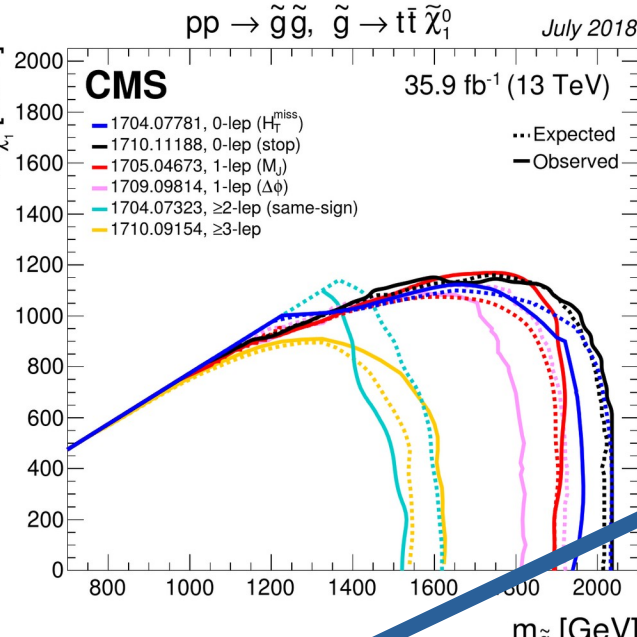
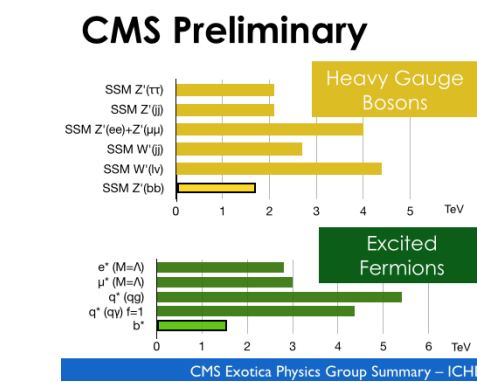
ATLAS SUSY Searches* - 95% CL Lower Limits
May 2017

Model	$\epsilon, \mu, \tau, \gamma$	Jets	E_{T}^{miss}	Mass limit	$\sqrt{s} = 7.8 \text{ TeV}$	$\sqrt{s} = 13 \text{ TeV}$
MSSUGRA/CMSSM	$0.3, \mu(1,2) \times 2.10 \text{ jets} \times 3$	Yes	20.3	1.36 TeV	$m_0 = m_{1/2}$	1507.6565
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	0	2 jets	Yes	36.1	1.57 TeV	ATLAS CONF-2017-022
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	0	2 jets	Yes	36.1	1.57 TeV	1604.47779
monojet	1-3 jets	Yes	3.2	2.02 TeV	$m_0 = m_{1/2} = 0.5 \text{ GeV}$	ATLAS CONF-2017-022
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	3- μ	4 jets	Yes	36.1	2.01 TeV	ATLAS CONF-2017-022
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	3- μ	4 jets	Yes	36.1	1.855 TeV	ATLAS CONF-2017-022
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	0	2 jets	Yes	36.1	1.8 TeV	ATLAS CONF-2017-022
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	0	2 jets	Yes	36.1	1.8 TeV	1607.55979
GGM (bino NLSP)	$1.2 \times 0.1 \times 1 \times 0.2 \text{ jets}$	Yes	3.2	2.5 TeV	$m_0 = m_{1/2} = 0.1 \text{ mm}$	1605.09150
GGM (bino NLSP)	2	1 jet	Yes	36.1	1.87 TeV	1507.25483
GGM (higgsino-bino NLSP)	7	2 jets	Yes	36.1	1.8 TeV	1507.25483
GGM (higgsino-bino NLSP)	7	2 jets	Yes	36.1	1.8 TeV	1503.03290
GGM (higgsino NLSP)	$2 \times \mu(2)$	2 jets	Yes	36.1	1.8 TeV	1503.03290
Grafitino LSP	0	monojet	Yes	20.3		

ATLAS Preliminary
 $\sqrt{s} = 7.8, 13 \text{ TeV}$

Reference
1507.6565
1604.47779
1507.25483
1503.03290
1605.09150
1507.25483
1503.03290

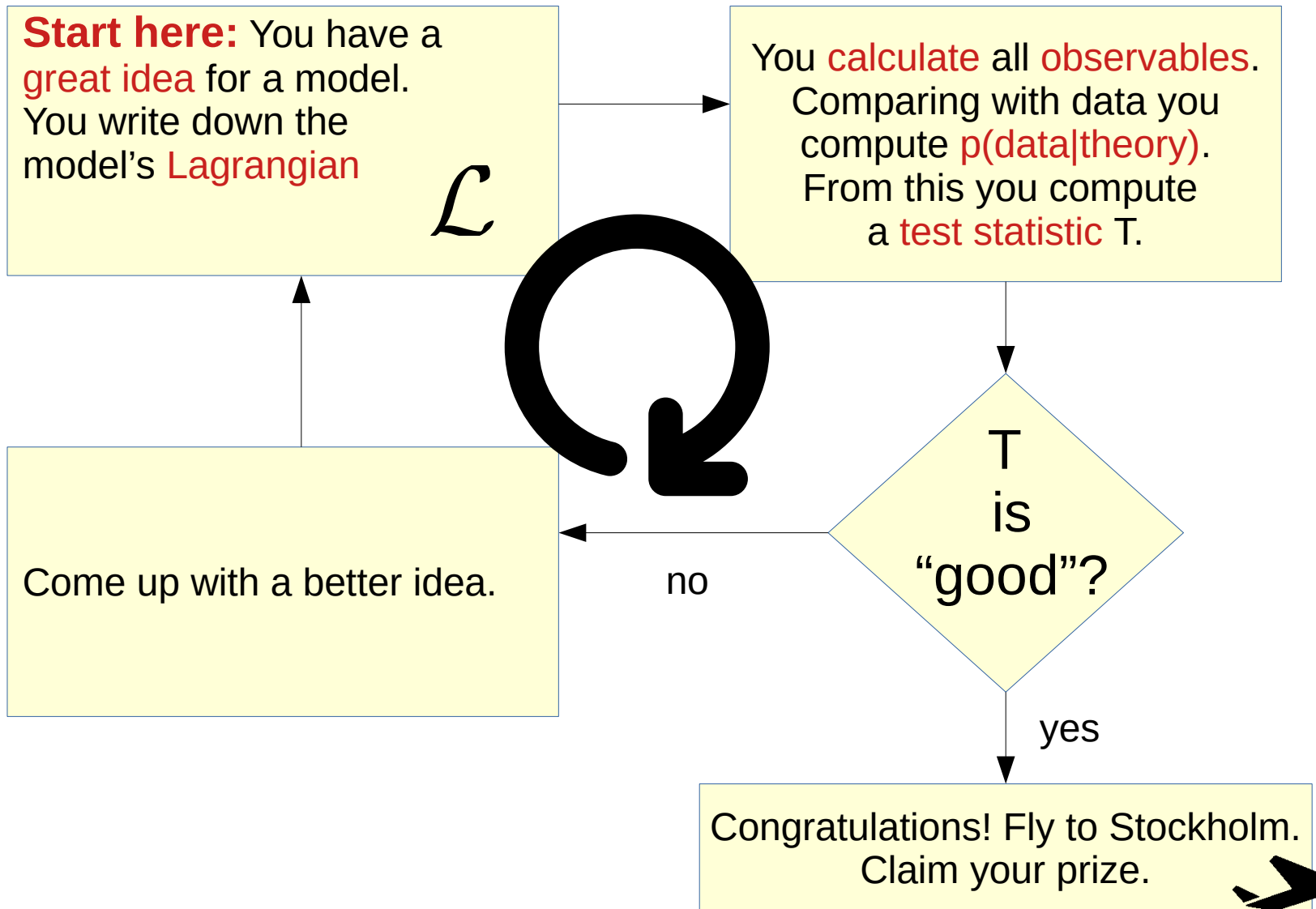
3 σ gluon annihilation direct production	$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	Mass limit
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	0	2 jets	Yes	36.1	1117.170 GeV
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	0	2 jets	Yes	36.1	90-198 GeV
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	2- μ (SS)	1 jet	Yes	36.1	275-700 GeV
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	2- μ (SS)	1 jet	Yes	36.1	200-700 GeV
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	0.2- μ	0.2 jets	Yes	20.3/36.1	205
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	0.2- μ	0.2 jets	Yes	20.3/36.1	205
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	2- μ (2)	1 jet	Yes	36.1	90-323 GeV
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	2- μ (2)	1 jet	Yes	36.1	150-600 GeV
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	2- μ (2)	1 jet	Yes	36.1	200-700 GeV
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	2- μ (2)	1 jet	Yes	36.1	200-700 GeV
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	2- μ (2)	1 jet	Yes	36.1	90-440 GeV
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	2- μ (2)	1 jet	Yes	36.1	710 GeV
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	0	0 jets	Yes	36.1	380 GeV
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	0	0 jets	Yes	36.1	380 GeV
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	0.2- μ	0.2 jets	Yes	20.3	270 GeV
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	0.2- μ	0.2 jets	Yes	20.3	270 GeV
$\tilde{g}, \tilde{u}_L, \tilde{u}_R$	0.2- μ	0.2 jets	Yes	20.3	270 GeV
$\tilde{g}, \tilde{d}_L, \tilde{d}_R$	0.2- μ	0.2 jets	Yes	20.3	270 GeV
GGM (bino NLSP) weak prod. $\tilde{g}, \tilde{u}_L, \tilde{u}_R$	1- μ	2 jets	Yes	20.3	115-370 GeV
GGM (bino NLSP) weak prod. $\tilde{g}, \tilde{d}_L, \tilde{d}_R$	1- μ	2 jets	Yes	20.3	590 GeV
GGM (higgsino NLSP) weak prod. $\tilde{g}, \tilde{u}_L, \tilde{u}_R$	1- μ	2 jets	Yes	20.3	430 GeV
GGM (higgsino NLSP) weak prod. $\tilde{g}, \tilde{d}_L, \tilde{d}_R$	1- μ	2 jets	Yes	20.3	430 GeV
Stable \tilde{g} R-hadron	1 jet	Yes	36.1	430 GeV	
Stable \tilde{u} R-hadron	1 jet	Yes	36.1	880 GeV	
Metastable \tilde{g} R-hadron	1 jet	Yes	36.1	880 GeV	
GMSB, $\tilde{g}, \tilde{u}_L, \tilde{u}_R$	1 jet	Yes	36.1	337 GeV	
GMSB, $\tilde{g}, \tilde{d}_L, \tilde{d}_R$	1 jet	Yes	36.1	440 GeV	
GMSB, $\tilde{g}, \tilde{u}_L, \tilde{u}_R$	1 jet	Yes	36.1	440 GeV	
GMSB, $\tilde{g}, \tilde{d}_L, \tilde{d}_R$	1 jet	Yes	36.1	440 GeV	
GGM $\tilde{g}, \tilde{u}_L, \tilde{u}_R$	1 jet	Yes	36.1	440 GeV	
GGM $\tilde{g}, \tilde{d}_L, \tilde{d}_R$	1 jet	Yes	36.1	440 GeV	



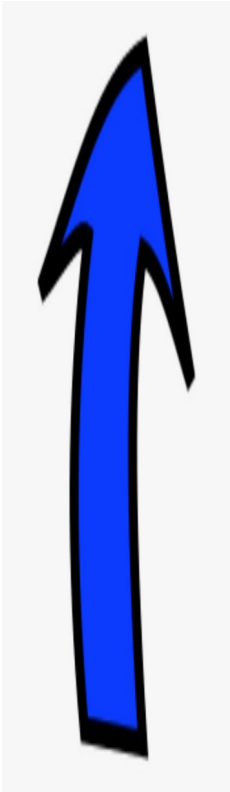
So how do we get from here to here?

Top-down versus bottom-up

Top-Down:



Top-down versus bottom-up



**Bottom-
Up:**

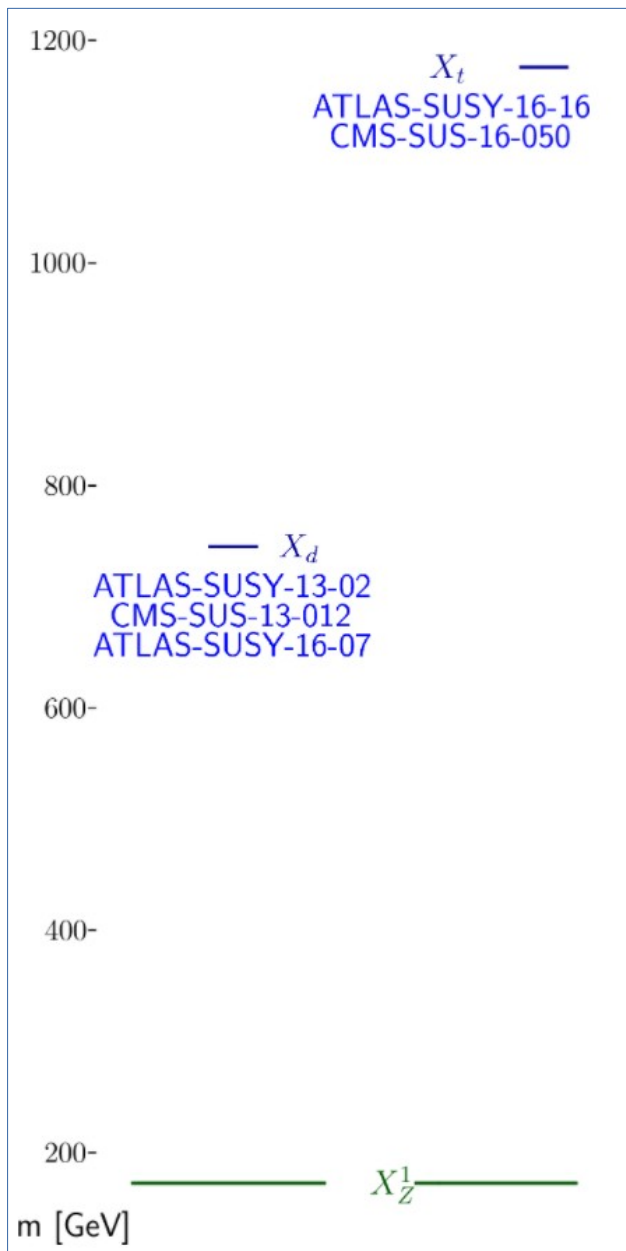
Only now do you think about symmetries, gauge groups, etc that may underlie all observations. Construct your Lagrangian.

\mathcal{L}

From the descriptions you try and **construct precursor theories** to the NSM that **describe everything you really know** about TeV-scale (and below) physics

Start here: You **describe your experimental findings** in a language amenable to theoretical physics, e.g. **simplified models** for on-shell effects (“searches”), **effective field theories** for off-shell effects (“measurements”).

THE HISCORE PROTO-MODEL



Analysis	Dataset	Obs	Exp	Z	P	Signal
ATL multijet, 8 TeV [54]	SR6jtp	6	4.9 ± 1.6	0.4σ	X_d	0.25
ATL multijet, 13 TeV [55]	2j.Me ...	611	526 ± 31	2.2σ	X_d	44.18
ATL 1ℓ stop, 13 TeV [48]	tN_high	8	3.8 ± 1	1.9σ	X_t	3.93
CMS multijet, 8 TeV [56]		30.8 fb	19.6 fb	1.1σ	X_d	2.66 fb
CMS 0ℓ stop, 13 TeV [49]		4.5 fb	2.5 fb	1.6σ	X_t	2.62 fb

Tension!

Table 3: Analyses contributing to the K value of the highest score proto-model

the dispersed excess

Analysis (all CMS 13 TeV)	Prod	σ_{XX} (fb)	$\sigma_{\text{obs}}^{\text{UL}}$ (fb)	$\sigma_{\text{exp}}^{\text{UL}}$ (fb)	r_{obs}
CMS multijet, M_{HT} , 137 fb^{-1} [15]	(\bar{X}_d, X_d)	23.96	18.45	21.57	1.30
CMS multijet, M_{HT} , 137 fb^{-1} [15]	(\bar{X}_t, X_t)	2.62	2.04	2.08	1.28
CMS multijet, M_{HT} , 36 fb^{-1} [57]	(\bar{X}_d, X_d)	23.96	19.26	28.31	1.24
CMS multijet, M_{T2} , 36 fb^{-1} [58]	(\bar{X}_d, X_d)	23.96	26.02	31.79	0.92
CMS 1ℓ stop, 36 fb^{-1} [59]	(\bar{X}_t, X_t)	2.62	2.91	4.44	0.90

Table 4: List of the most constraining results for the highest score proto-model. The

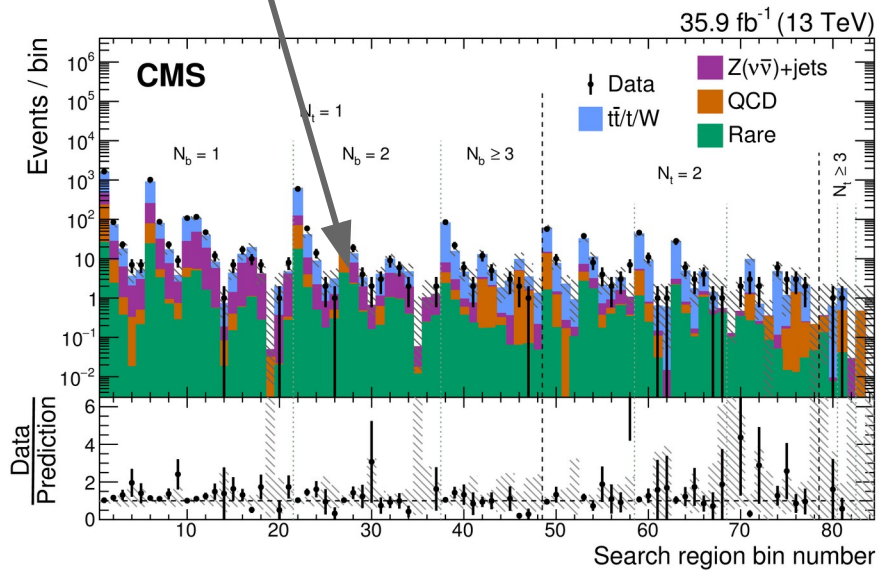
what is driving the “critic”

Signal strength multipliers: $(\bar{X}_t, X_t) = 1.2$; $(\bar{X}_d, X_d), (X_d, X_d^1), (\bar{X}_d, X_d^1) = 0.49$

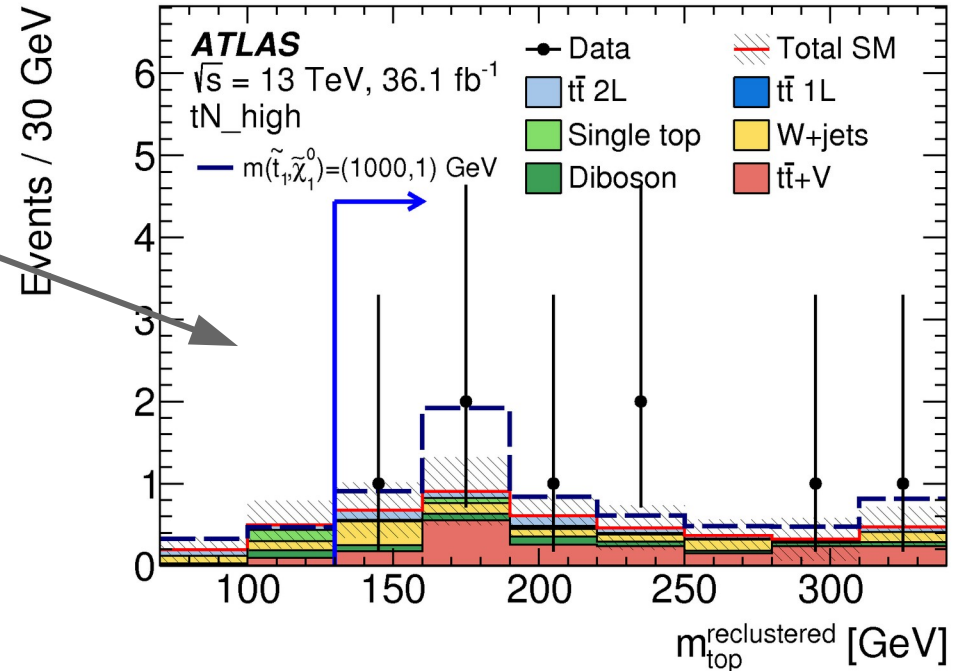
Contributions by particles: $X_t : K_{\text{without}} = 2.59(59\%), X_d : K_{\text{without}} = 3.90(41\%)$
Last updated: Mon Dec 14 20:08:06 2020

Data driving the protomodel

Nothing unusual is visible in the individual results



CMS 0l stop (CMS-SUS-16-050)



ATLAS 1l stop (ATLAS-SUSY-2016-16)

But when interpreted as a correlated signal originating from the same protomodel, things may seem different!

LIKELIHOODS



- **Only exclusion lines**

If only exclusion lines are given, without upper limits, we can do nothing

- **Observed 95% CL upper limits only:**

cannot construct likelihood, binary decision “excluded” / “not-excluded” only (“critic”)

- **Expected and observed 95% CL upper limits**

can construct an approximate likelihood with truncated Gaussian, cannot combine topologies, very crude approximation

- **Efficiency maps**

can construct a likelihood as Gaussian (for the nuisances) * Poissonian (for yields), can work per SR, and combine topologies in each SR [*]

- **Efficiency maps + correlation matrices**

can combine signal regions via multivariate Gaussian * Poissonians

- **Efficiency maps + full likelihoods**

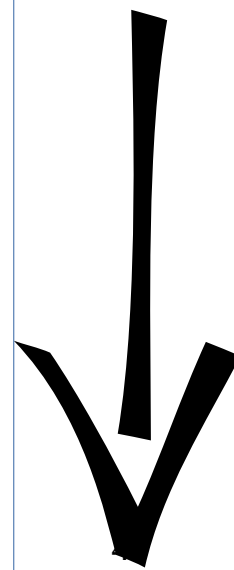
full realism, correct statistical model



Compos

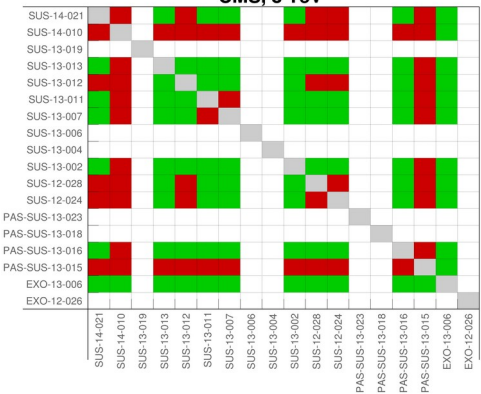
Likelihoods

BETTER



[*] if efficiency maps are not supplied, we can try to produce them with recasting frameworks

CMS, 8 TeV



THE COMBINER

we allow the machine to combine likelihoods.

Approximately uncorrelated are analyses that are:

- from different runs, and/or
- from different experiments, and/or
- looking for (clearly) different signatures

ATLAS, 8 TeV

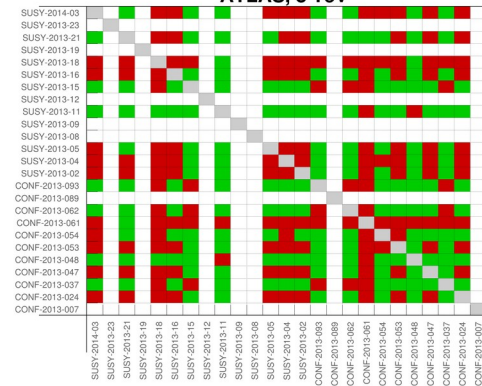
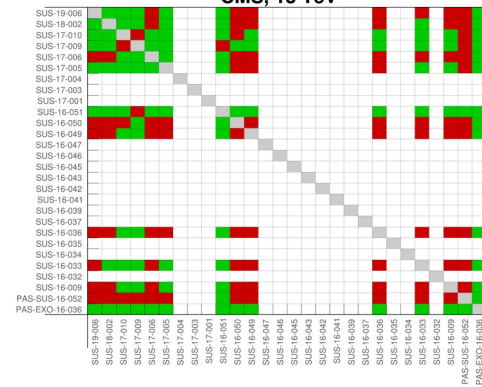


Fig. 2

CMS, 13 TeV

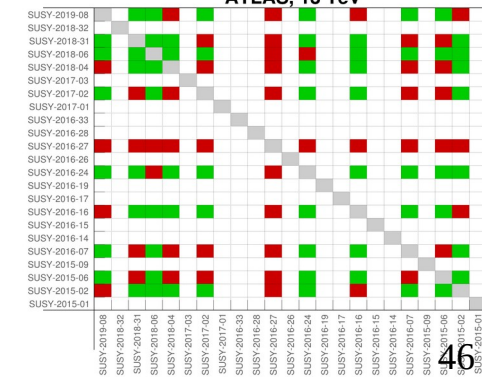


A combination “c” of analyses is “legal” if the following conditions are met:

- all results are mutually uncorrelated (= “combinable”)
- if a result can be added, it has to be added (any subset of a legal combination is not itself legal)

- combined likelihood: $L_c = \prod_{i \in c} L_i$

ATLAS, 13 TeV



THE TEST STATISTIC

For every legal combination, we define a test statistic K

$$K^c := -2 \ln \frac{L_{\text{SM}}^c \cdot \pi(\text{SM})}{L_{\text{BSM}}^c(\hat{\mu}) \cdot \pi(\text{BSM})} \quad \text{Eq. 6}$$

(Remember, we have a database of results from ~ 100 CMS+ATLAS searches. We want to find the most interesting combinations of these results, i.e. the ones that maximally violate the SM hypothesis)

Of all “legal” combinations of experimental results, the builder chooses the one combination “c” that maximizes K :

$$K := \max_{\forall c \in C} K^c \quad \text{Eq. 7}$$

μ denotes an global signal strength multiplier – the production cross sections are free parameters

$$\forall i, j : \sigma(pp \rightarrow X_i X_j) = \mu \bar{\sigma}(pp \rightarrow X_i X_j)$$

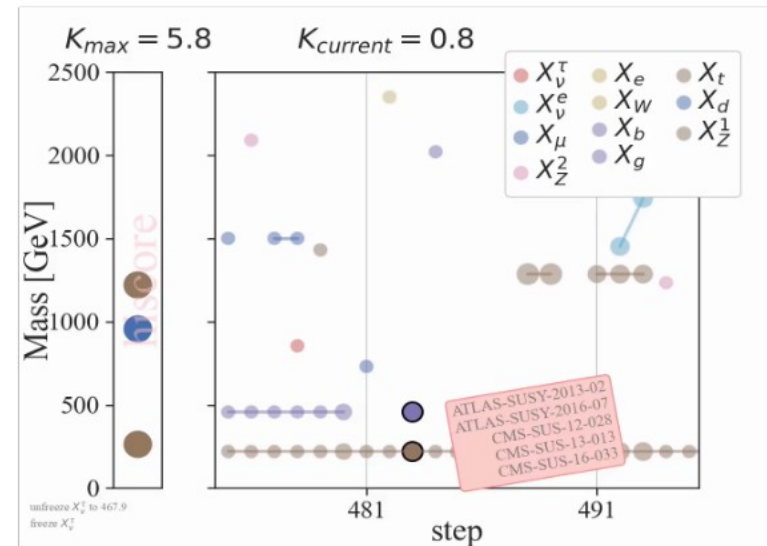
It is maximized in the denominator, but its support is confined such that no limits in the SModelS database are violated (the “critic”),

$$\hat{\mu} \in [0, \mu_{\text{max}}]$$

THE WALKER

The Walker takes care of moving in the protomodel space with varying dimensionality by performing the following types of modifications to the protomodel:

- **add or remove particles** from the protomodel
- **change the masses** of particles
- **change the signal strengths** of production modes
- **change decay channels and branching ratios**



At each step the test statistic K is computed. An MCMC-like procedure[*] is then applied in the sense that the step is reverted with a probability of

$$\exp \left[\frac{1}{2} (K_i - K_{i-1}) \right]$$

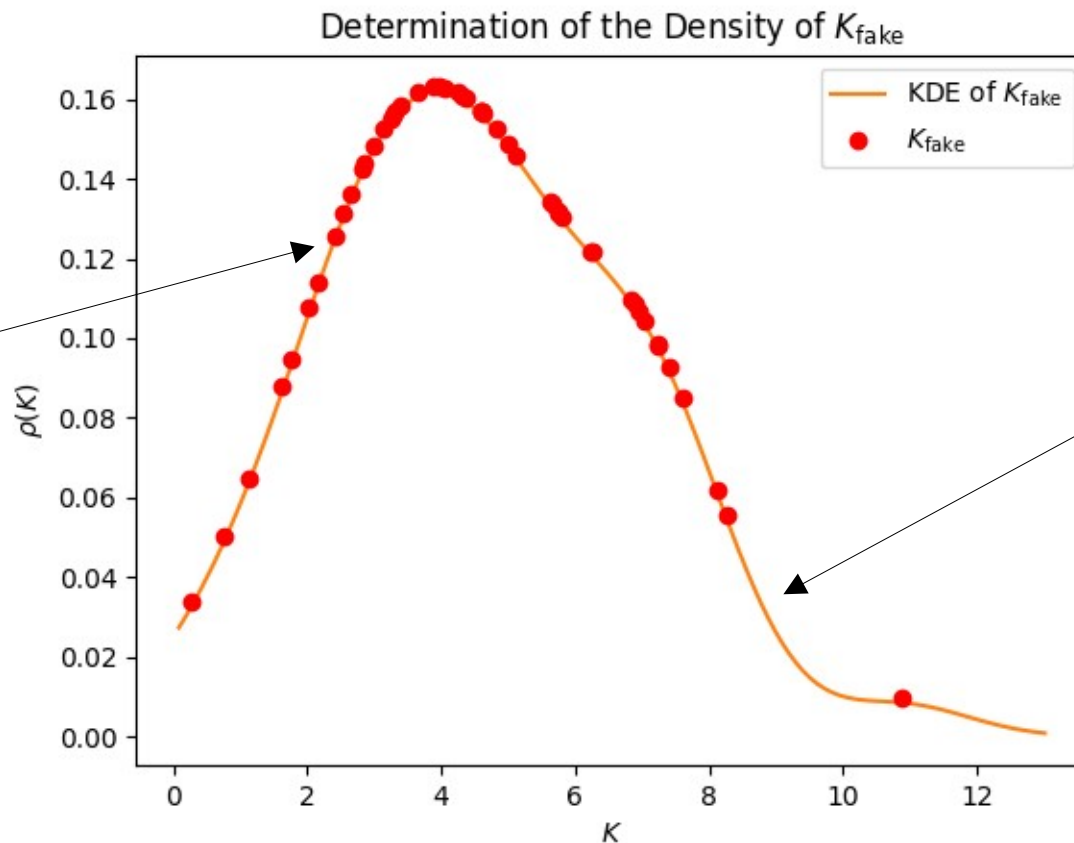
if and only if K_i is smaller than K_{i-1}

* (note however, instead of ratios of unnormalized posteriors we have ratios of ratios of unnormalized posteriors)

WALKING OVER FAKE STANDARD MODEL DATABASES

- Produced 50 “fake” SModelS databases by sampling background models
- Corresponds to typical LHC results if no new physics is in data
- Determine 50 “fake” K values by running 50 walkers on each of the 50 databases (50 x 50 walkers in total) → density of K under null SM-only hypothesis

K for one “fake” background-only database.



Density of K estimated via a simple Kernel density estimator.

THE WALKS

We define a “run” as 50 parallel walks, each taking 1000 steps.

We performed

- 10 runs on the SModelS database (Sec. 5.2)
- 50 runs on fake “Standard Model-like” databases (Sec 5.1)
to be able to determine a global p -value under the SM hypothesis
- 2x10 runs on fake “Signal-like” databases (Sec 5.3)
to show closure of the method

WALKING OVER DATABASES WITH FAKE SIGNALS

To show closure of our method, we inject the winning protomodel as a signal in fake databases, and see if the algorithm can reconstruct the injected signal.

Sec 5.3

Technical closure test

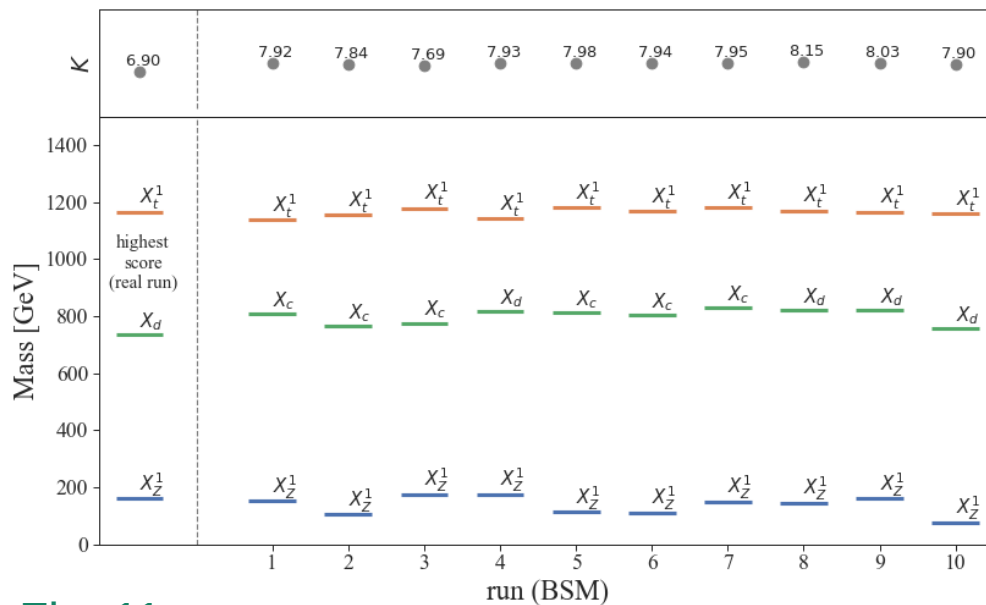


Fig. 11

No sampling of the models for the SRs, i.e. observed events := expected SM + expected signal events

Physics closure test

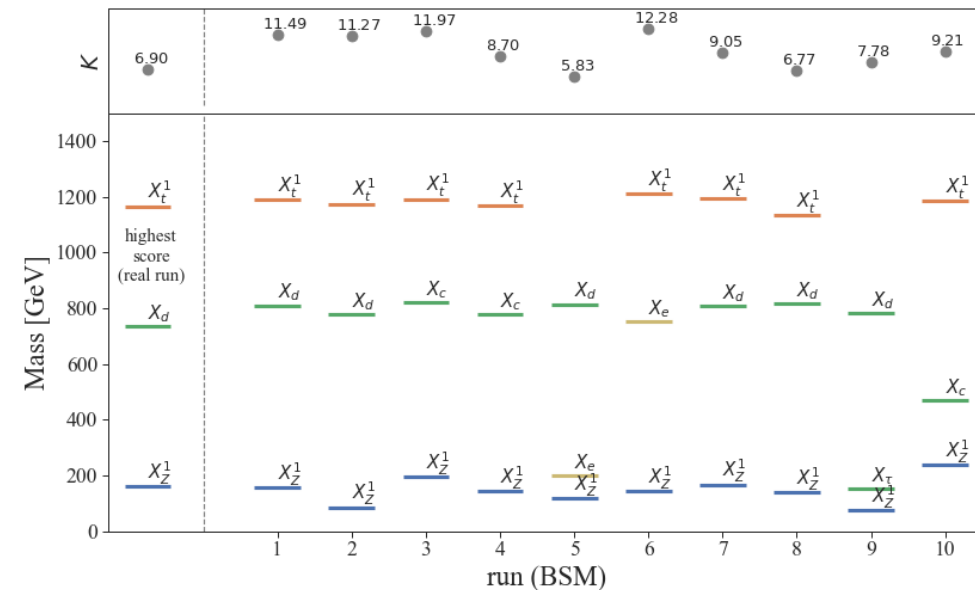


Fig. 10

Sampling turned on

FUTURE IMPROVEMENTS

Improvements of the SModelS database:

- add latest full run-2 CMS and ATLAS publications (Moriond!)
- produce efficiency maps for existing results
- enlarge mass range of older efficiency maps

Improvements in speed:

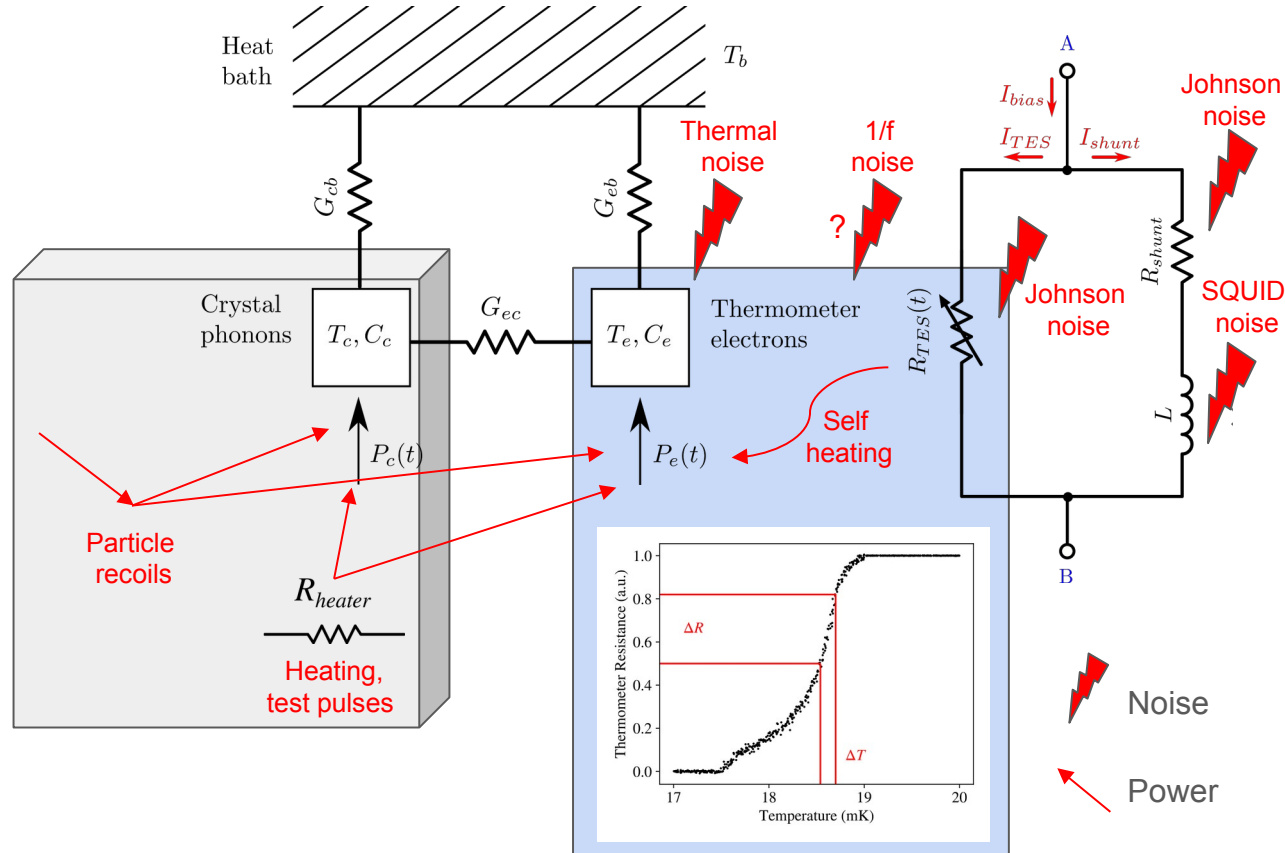
- learn the SModelS database
- make everything differentiable

Improvements in procedure:

- improve the “analyses correlation matrix”, automate the determination
- ponder relationship between proto-models and effective field theories
- connect proto-models with complete theories

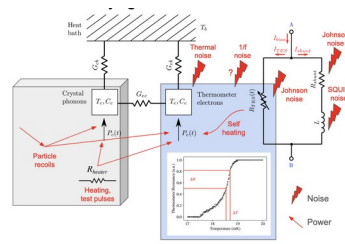
REINFORCEMENT LEARNING, CRYOGENIC DETECTORS

The situation with cryogenic detectors

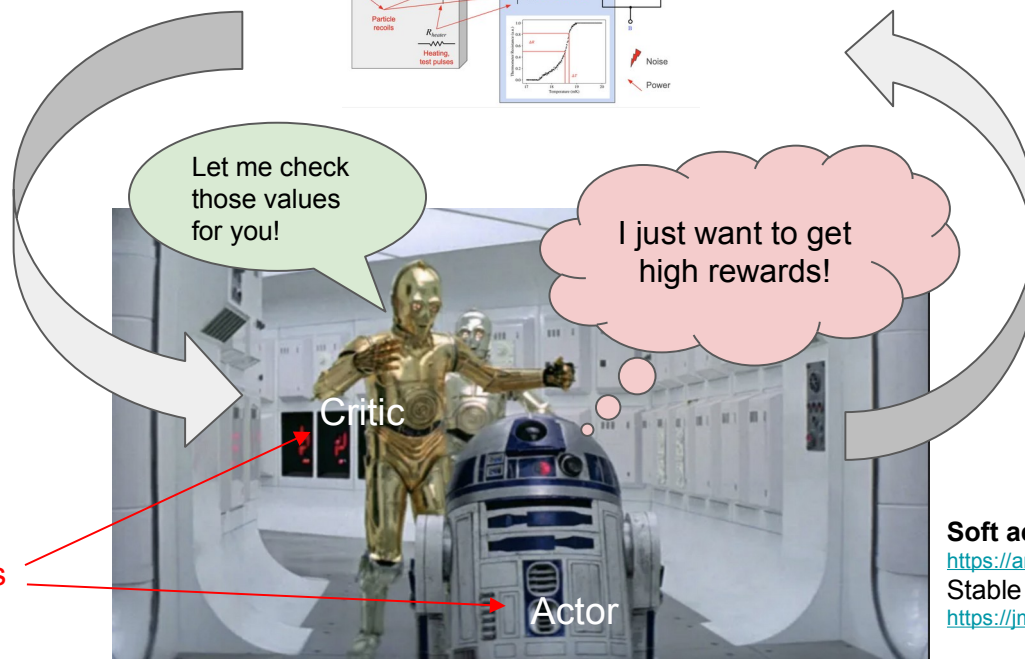


REINFORCEMENT LEARNING, CRYOGENIC DETECTORS

A framework for policy optimization: *reinforcement learning*



Observation
(PH, RMS) and
Reward (LT
objective)



Action
(setting DAC
and Ib)

2 neural networks

Critic

Actor

Soft actor critic (SAC)

<https://arxiv.org/abs/1801.01290v2>

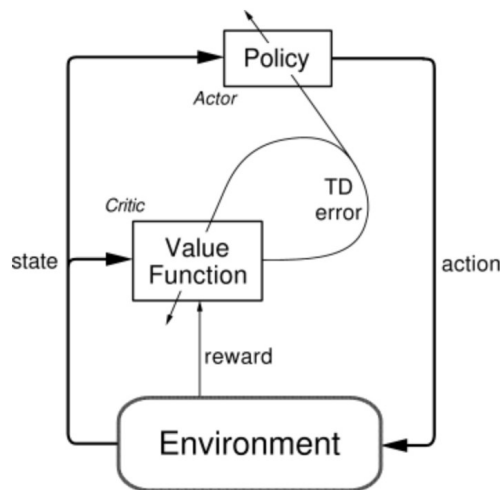
Stable baselines 3 implementation

<https://jmlr.org/papers/volume22/20-1364/20-1364.pdf>

54

REINFORCEMENT LEARNING, CRYOGENIC DETECTORS

Actor critic



$$\delta_t = r_{t+1} + \gamma V(s_{t+1}) - V(s_t),$$

Soft Actor-Critic: Off-Policy Maximum Entropy Deep Reinforcement Learning with a Stochastic Actor

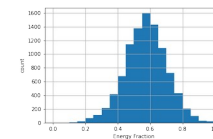
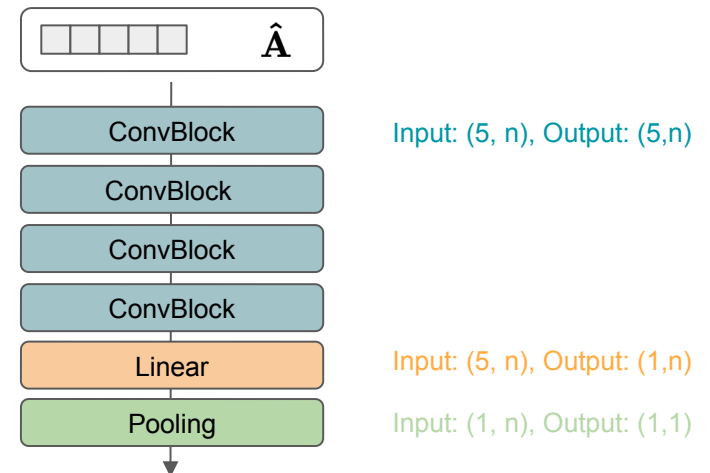
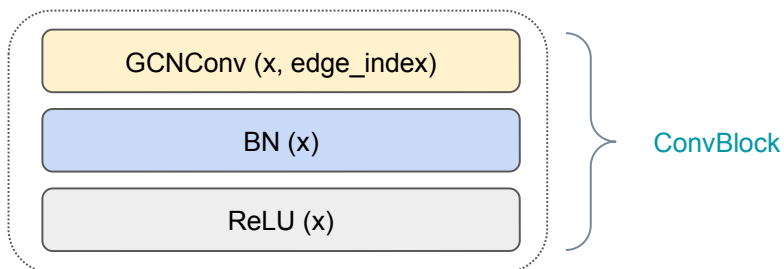
Tuomas Haarnoja, Aurick Zhou, Pieter Abbeel, Sergey Levine

Model-free deep reinforcement learning (RL) algorithms have been demonstrated on a range of challenging decision making and control tasks. However, these methods typically suffer from two major challenges: very high sample complexity and brittle convergence properties, which necessitate meticulous hyperparameter tuning. Both of these challenges severely limit the applicability of such methods to complex, real-world domains. In this paper, we propose soft actor-critic, an off-policy actor-critic deep RL algorithm based on the maximum entropy reinforcement learning framework. In this framework, the actor aims to maximize expected reward while also maximizing entropy. [That is, to succeed at the task while acting as randomly as possible.](#) Prior deep RL methods based on this framework have been formulated as Q-learning methods. By combining off-policy updates with a stable stochastic actor-critic formulation, our method achieves state-of-the-art performance on a range of continuous control benchmark tasks, outperforming prior on-policy and off-policy methods. Furthermore, we demonstrate that, in contrast to other off-policy algorithms, our approach is very stable, achieving very similar performance across different random seeds.

$$\pi^* = \arg \max_{\pi} \mathbb{E}_{\pi} \left[\sum_{t=1}^T \underbrace{R(s_t, a_t)}_{\text{reward}} + \alpha \underbrace{H(\pi(\cdot | s_t))}_{\text{entropy}} \right]$$

GRAPH NEURAL NETWORKS

- ❖ using pytorch geometric
- ❖ consists of ConvBlocks, a linear model and a pooling layer
- ❖ GCNConv
 - updates nodes according to neighbors via adjacency matrix
 - symmetric normalization



$$\mathbf{X}' = \hat{\mathbf{D}}^{-1/2} \hat{\mathbf{A}} \hat{\mathbf{D}}^{-1/2} \mathbf{X} \Theta$$

<https://arxiv.org/abs/1609.02907>

GRAPH NEURAL NETWORKS

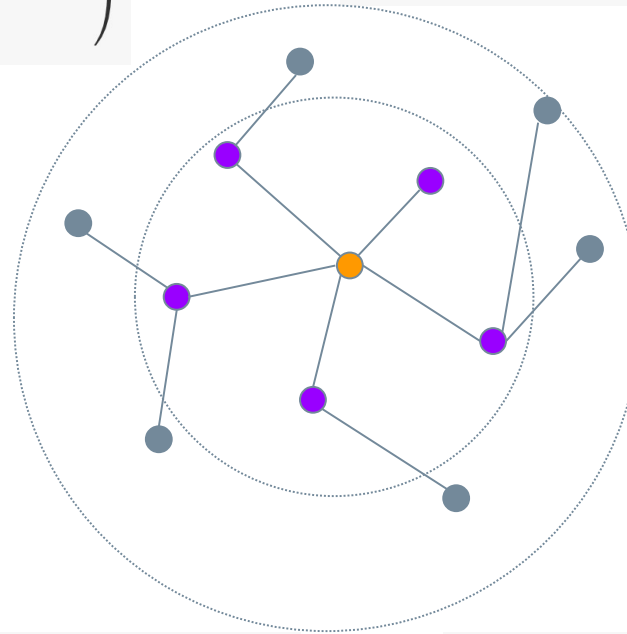
Hackathon - Convolution Operations

$$h_v^{(k)} = f^{(k)} \left(W^{(k)} \cdot \frac{\sum_{u \in \mathcal{N}(v)} h_u^{(k-1)}}{|\mathcal{N}(v)|} + B^{(k)} \cdot h_v^{(k-1)} \right)$$

GCNConv

$$h_v^{(k)} = f^{(k)} \left(W^{(k)} \cdot \left[\text{AGG}_{u \in \mathcal{N}(v)} (\{h_u^{(k-1)}\}), h_v^{(k-1)} \right] \right)$$

SAGEConv



GATConv

GINConv

$$h_v^{(k)} = f^{(k)} \left(W^{(k)} \cdot \left[\sum_{u \in \mathcal{N}(v)} \alpha_{vu}^{(k-1)} h_u^{(k-1)} + \alpha_{vv}^{(k-1)} h_v^{(k-1)} \right] \right)$$

$$h_v^{(k)} = f^{(k)} \left(\sum_{u \in \mathcal{N}(v)} h_u^{(k-1)} + (1 + \epsilon^{(k)}) \cdot h_v^{(k-1)} \right)$$

GRAPH NEURAL NETWORKS

Hackathon - Results

