# Local supersymmetry enhancement and the entropy of three-charge black holes 

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Based on [2211.14326] with I. Bena, S. Hampton, A. Houppe and D. Toulikas

## What this talk is about



- M5-M2(-P) black hole: The microstates that are made of fractionated $\mathbf{M}$ 2 branes account for the entropy.
- We found: They can transition into microstates with 16 local supersymmetries.
« Dijkgraaf-Verlinde-Verlinde -Maldacena microstates»


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- We found: They can transition into microstates with 16 local supersymmetries.

Microstates with 16 local susys account for the black-hole entropy!

- We expect their backreaction to be horizonless microstates.


## Outline

1. Local supersymmetry enhancement and black-hole microstates
2. The new M5-M2-P microstates with 16 local supersymmetries

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## The 3-charge black hole and near-horizon geometry

- Type IIA/IIB: $\mathbb{R}^{4,1} \times S_{y}^{1} \times T^{4}$
- Take brane system with 3 charges:

D5(y, $\left.T^{4}\right), ~ D 1(y), ~ P(y)$
or NS5(y, $\left.T^{4}\right)$, F1 (y), P(y)
$\Rightarrow$ naively, 1/8-BPS everywhere

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- Use harmonic function rule

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d s^{2}= & -\frac{2}{\sqrt{H_{1} H_{5}}}\left[d t^{2}+d y^{2}+\left(H_{P}-1\right)^{-1}(d y-d t)^{2}\right] \\
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with

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H_{1,5, P}=1+\frac{Q_{1,5, P}}{r^{2}}
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## Possible conclusion:

Global charges and
supersymmetries seem to control near-horizon geometry.

Therefore all brane systems develop the same horizon:
To have access the information about the microstates, probe singularity region, where supergravity breaks down.

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- Use harmonic function rule
$\Rightarrow$ develops horizon in supergravity
- Local supersymmetry enhancement:
String-theory excitations (branes, strings) combine together to form a bound state that is locally 1/2-BPS (16 susies).


## Local SUSY enhancement - example

- Ex.: F1 (y) and parallel P(y):

- F1or P preserve 16 real supercharges
- Together, F1-P preserve 8


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- Ex.: F1 (y) and parallel P(y):

- F1or P preserve 16 real supercharges
- Together, F1-P preserve 8
- Actually the string can carry momentum: profile
- The F1-P profile preserves the same global supersymmetries...
- ...but locally it is a $\mathrm{F} 1(\hat{y})$ boosted by orthogonal $P\left(\hat{y}^{\perp}\right)$
- F1 $(\hat{y})-\mathrm{P}\left(\hat{y}^{\perp}\right)$ preserves 16 supercharges

Local supersymmetry $\rightarrow$ information on microstate?

## Local VS global supersymmetries

- Branes, strings $\rightarrow$ constraint on $\epsilon$;


Killing spinor

Projector

$$
\Pi^{2}=\Pi
$$

Traceless involution

$$
\begin{gathered}
P^{2}=1 \\
\operatorname{tr}(P)=0
\end{gathered}
$$

Constraint halves number
of supersymmetries

## Local VS global supersymmetries

- Branes, strings $\rightarrow$ constraint on $\epsilon$ :

$$
\Pi \epsilon \equiv \frac{1}{2}(1+P) \epsilon=0
$$

- Combine $k$ different excitations:

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\begin{aligned}
& \epsilon \in \operatorname{ker}\left(\Pi_{1}\right) \cap \ldots \cap \operatorname{ker}\left(\Pi_{k}\right) \\
& \rightarrow \sharp(\text { global supersymmetries })
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- Add other involutions $\left(P_{k+1}, \ldots, P_{n}\right)$ and weights $\left(\alpha_{1}, \ldots, \alpha_{n}\right)$ s.t. $\alpha_{1}+\ldots+\alpha_{n}=1$.


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Mix of excitations (branes, etc) with different charges:

$$
\alpha_{i}=\frac{Q_{i}}{M}
$$

Not necessarily a projector!

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- $\left\{\alpha_{i}\right\}$ not unique $\rightarrow\left\{\alpha_{i}(x)\right\}$

along the bound state
$\epsilon$ promoted to be a function


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$$
\hat{\Pi}(x) \epsilon(x)=0
$$

At $x, \epsilon(x) \in \operatorname{ker}(\hat{\Pi}(x)) \rightarrow$ local supersymmetry

- While for global supersymmetry:

$$
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$\#$ (global susies) $\leq \#$ (local susies)

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## Local supersymmetry enhancement

Local supersymmetry enhancement:
Given a set of global supersymmetries, there sometimes exists a whole moduli space of brane/string systems, parameterised by $\left\{\alpha_{i}(x)\right\}$, preserving those same global supersymmetries, but whose number of local supersymmetries is enhanced.

- identify the additional excitations («glues ») to make a bound state
- determine the charge-to-mass ratios $\left\{\alpha_{i}(x)\right\}$.


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## Local SUSY enhancement - example



- $\Pi_{\mathrm{FI}(y)}=\frac{1}{2}\left(1+P_{\mathrm{FI}(y)}\right), \quad P_{\mathrm{FI}(y)}=\Gamma^{0 y} \sigma_{3}$
- $\hat{\Pi}=\frac{1}{2}\left(1+\alpha_{1} P_{\mathrm{F} 1(y)}+\alpha_{2} P_{\mathrm{P}(y)}+\alpha_{3} P_{\mathrm{F}(1)}+\alpha_{4} P_{\mathrm{P}(1)}\right)$
- $\Pi_{\mathrm{P}(y)}=\frac{1}{2}\left(1+P_{\mathrm{P}(y)}\right), \quad P_{\mathrm{P}(y)}=\Gamma^{0 y}$


## Local SUSY enhancement - example



## Step 0

- $\Pi_{\mathrm{Fl}(\mathrm{y})}=\frac{1}{2}\left(1+P_{\mathrm{Fl}(\mathrm{y})}\right), \quad P_{\mathrm{Fl}(\mathrm{y})}=\Gamma^{0 y} \sigma_{3}$
- $\Pi_{\mathrm{P}(y)}=\frac{1}{2}\left(1+P_{\mathrm{P}(y)}\right), \quad P_{\mathrm{P}(y)}=\Gamma^{0 y}$
- $\hat{\Pi}=\frac{1}{2}\left(1+\alpha_{1} P_{\mathrm{F}(y)}+\alpha_{2} P_{\mathrm{P}(y)}+\alpha_{3} P_{\mathrm{F}(1)}+\alpha_{4} P_{\mathrm{P}(1)}\right)$
defining the global
supersymmetries


## Local SUSY enhancement - example



## Step 1

- $\Pi_{\mathrm{Fl}(\mathrm{y})}=\frac{1}{2}\left(1+P_{\mathrm{Fl}(\mathrm{y})}\right), \quad P_{\mathrm{Fl}(\mathrm{y})}=\Gamma^{0 y} \sigma_{3}$
- $\Pi_{\mathrm{P}(y)}=\frac{1}{2}\left(1+P_{\mathrm{P}(y)}\right), \quad P_{\mathrm{P}(y)}=\Gamma^{0 y}$
- $\hat{\Pi}=\frac{1}{2}\left(1+\alpha_{1} P_{\mathrm{Fl}(y)}+\alpha_{2} P_{\mathrm{P}(y)}+\alpha_{3} P_{\mathrm{Fl}(1)}+\alpha_{4} P_{\mathrm{P}(1)}\right)$ dipoles


## Local SUSY enhancement - example



## Step 2

- $\Pi_{\mathrm{FI}(y)}=\frac{1}{2}\left(1+P_{\mathrm{FI}(y)}\right), \quad P_{\mathrm{FI}(y)}=\Gamma^{0 y} \sigma_{3}$
- $\Pi_{\mathrm{P}(y)}=\frac{1}{2}\left(1+P_{\mathrm{P}())}\right), P_{\mathrm{P}(y)}=\Gamma^{0 y}$



## Local SUSY enhancement - example



## Microstates of the F1-P black hole



- Harmonic rule:
$d s_{s t r i n g}^{2}=H\left[-d u d v+K d v^{2}\right]+\sum_{i=1}^{4} d x_{i} d x_{i}+\sum_{a=1}^{4} d z_{a} d z_{a}$
$\rightarrow$ black hole with horizon at $r=0$.


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$\rightarrow$ black hole with horizon at $r=0$.

- Metric sourced by the string:

$$
d s_{\text {string }}^{2}=H\left[-d u d v+K d v^{2}+2 A_{i} d x_{i} d v\right]+\sum_{i=1}^{4} d x_{i} d x_{i}+\sum_{a=1}^{4} d z_{a} d z_{a}
$$

[Dabhorkar, Gauntlett, Harvey, Waldram '95]
$\rightarrow$ smooth, horizonless solution.

## Microstates of the F1-P black hole

The massive string states accounting for the F1-P blackhole entropy can be described in supergravity.

- Harmonic rule:
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## 2-charge VS 3-charge black holes

- Such «classical» string profiles, through geometric quantization, account for the F1-P black-hole entropy:

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$\rightarrow$ Is the stringy structure resolving the horizon or the singularity?


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- However: 2-charge black holes: singularity, horizon at $r=0$.
$\rightarrow$ Is the stringy structure resolving the horizon or the singularity?
- 3-charge black holes: singularity and horizon separated.
$\rightarrow$ in particular D1-D5-P or F1-NS5-P


## 1st approach: enhancing D1-D5 through KKM

- D1(y), D5(y1234) $\longrightarrow$ KKM(1234 $\psi, y), ~ P(\psi)$ dipoles
- The D1-D5 brane system gains a dimension through the KKM



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- The (angular) momentum $\mathrm{P}(\psi)$ stabilises the size of the supertube.

$\rightarrow$ replace the deltafunction brane singularity by a source extended in the non-compact dimensions


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- The bound state is globally 1/4-BPS, but locally 1/2-BPS
- Then add consistently P and keep locally 1/2-BPS
$\rightarrow$ «superstrata »
[Bena, de Boer, Shigemori, Warner '11]
[Bena, Giusto, Martinec, Russo,
Shigemori, Turton, Warner '16]


## Superstrata and their limits

- In supergravity, superstrata are horizonless solutions with same charges as the D1-D5-P black hole

[Bena, Giusto, Martinec, Russo, Shigemori, Turton, Warner '16]

- Part of the


## Fuzzball hypothesis: <br> Individual black-hole microstates differ from themselves and from the BH solution at the horizon scale.

## Superstrata and their limits

- In supergravity, superstrata are horizonless solutions with same charges as the D1-D5-P black hole [Bena, Giusto, Martinec, Russo, Shigemori, Turton, Warner '16]
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## Fuzzball hypothesis:

Individual black-hole microstates differ from themselves and from the BH solution at the horizon scale.

Drawbacks:

1. $S \sim \sqrt{N_{1} N_{5}} N_{P}^{1 / 4} \ll \sqrt{N_{1} N_{5} N_{P}}$
[Shigemori '19]
2. Have a non-vanishing angular momentum in $\mathbb{R}^{4}$
$\Rightarrow$ could be atypical
$\uparrow$ are not exactly spherically symmetric

See also [Lin, Maldacena, Rozenberg, Shan '22]

## Outline

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## 2nd approach: internal dimensions

- For the NS5-F1-P black hole (IIA), we know where the entropy is coming from: Litlle strings / frackionaled (M2) branes

«Dijkgraaf-Verlinde-Verlinde -Maldacena microstates»


## 2nd approach: internal dimensions

- For the NS5-F1-P black hole (IIA), we know where the entropy is coming from: little strings / fractionated (M2) branes

- The momentum is carried by the fractionated M2's through their motion in the $T^{4}$
$\rightarrow$ reproduce entropy.
«Dijkgraaf-Verlinde-Verlinde -Maldacena microstates»
[Dijkgraaf-Verlinde-Verlinde '96], [Maldacena,'96]


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$\rightarrow$ reproduce entropy.

$$
S=2 \pi \sqrt{c N_{P} / 6}, \quad c=6 N_{1} N_{5}
$$

«Dijkgraaf-Verlinde-Verlinde -Maldacena microstates»

## 2nd approach: internal dimensions

- For the NS5-F1-P black hole (IIA), we know where the entropy is coming from: Little strings / fractionated (M2) branes

See e.g. [Martinec, Massai, Turton '19]

«Dijkgraaf-Verlinde-Verlinde -Maldacena microstates»
[Dijkgraaf-Verlinde-Verlinde '96], [Maldacena,'96]

- The momentum is carried by the fractionated M2's through their motion in the $T^{4}$
$\rightarrow$ reproduce entropy.
- The brane system is point-like in the non-compact spatial dimensions
$\rightarrow$ exact spherical symmetry.


## Enhancing the DVVM microstates

[Bena, Hampton, Houppe, YL, Toulikas '22]

- We enhanced the local supersymmetries of the Dijkgraaf-Verlinde-VerlindeMaldacena (DVVM) microstates.


## Enhancing the DVVM microstates

[Bena, Hampton, Houppe, YL, Toulikas '22]

- We enhanced the local supersymmetries of the Dijkgraaf-Verlinde-VerlindeMaldacena (DVVM) microstates.
- We found the supersymmetric projector
- preserving the supersymmetries of NS5 (y, $T^{4}$ ), F1 $(y), P(y)$
- corresponding to an object with 16 local supersymmetries:

$$
\begin{aligned}
\Pi_{\mathrm{NS} 5-\mathrm{F} 1-\mathrm{P}}=\frac{1}{2}[1 & +a^{2} P_{\mathrm{NS} 5(y 1234)}^{\mathrm{IIA}}+b^{2} P_{\mathrm{F} 1(y)}+c^{2} P_{\mathrm{P}(y)} \\
& \left.+a b\left(P_{\mathrm{D} 4(y 234)}-P_{\mathrm{D} 2(y 1)}\right)+b c\left(P_{\mathrm{P}(1)}-P_{\mathrm{F} 1(1)}\right)+c a\left(P_{\mathrm{D} 4(1234)}-P_{\mathrm{D} 0}\right)\right] .
\end{aligned}
$$

## First look at the projector

$$
\Pi_{\mathrm{NS} 5-\mathrm{F} 1-\mathrm{P}}=\frac{1}{2}\left[1+a^{a^{2} P_{\mathrm{NS} 5(y 1234)}^{\mathrm{IIA}}+b^{2} P_{\mathrm{F} 1(y)}+c^{2} P_{\mathrm{P}(y)}}\right.
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\left.+a b\left(P_{\mathrm{D} 4(y 234)}-P_{\mathrm{D} 2(y 1)}\right)+b c\left(P_{\mathrm{P}(1)}-P_{\mathrm{F} 1(1)}\right)+c a\left(P_{\mathrm{D} 4(1234)}-P_{\mathrm{D} 0}\right)\right]
$$

Excitations defining the global supersymmetries

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$$

$$
+\underline{\left.a b\left(P_{\mathrm{D} 4(y 234)}-P_{\mathrm{D} 2(y 1)}\right)+b c\left(P_{\mathrm{P}(1)}-P_{\mathrm{F} 1(1)}\right)+c a\left(P_{\mathrm{D} 4(1234)}-P_{\mathrm{D} 0}\right)\right] .}
$$



Excitations corresponding to the glues

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BPS condition:

$$
a^{2}+b^{2}+c^{2}=1
$$

$$
\begin{aligned}
P_{\mathrm{P}} & =\Gamma^{01} \\
P_{\mathrm{NS} 5}^{\mathrm{IIA}} & =\Gamma^{012345}
\end{aligned}
$$

$$
P_{\mathrm{KKM}(12345 ; 6)}^{\operatorname{IIA}}=\Gamma^{012345} \sigma_{3}=\Gamma^{6789}
$$

$$
P_{\mathrm{D} 0}=\Gamma^{0} i \sigma_{2}
$$

$$
P_{\mathrm{D} 2}=\Gamma^{012} \sigma_{1}
$$

$$
P_{\mathrm{D} 4}=\Gamma^{01234} i \sigma_{2}
$$

$$
P_{\mathrm{D} 6}=\Gamma^{0123456} \sigma_{1}
$$

$$
\begin{aligned}
P_{\mathrm{F} 1} & =\Gamma^{01} \sigma_{3} \\
P_{\mathrm{NS} 5}^{\mathrm{II}} & =\Gamma^{012345} \sigma_{3} \\
P_{\mathrm{KKM}(12345 ; 6)}^{\mathrm{IIB}} & =\Gamma^{012345}=\Gamma^{6789} \\
P_{\mathrm{D} 1} & =\Gamma^{01} \sigma_{1} \\
P_{\mathrm{D} 3} & =\Gamma^{0023} \sigma_{2} \\
P_{\mathrm{D} 5} & =\Gamma^{012345} \sigma_{1}
\end{aligned}
$$

## Glueing NS5 and F1

$$
\begin{aligned}
\Pi_{\mathrm{NS} 5-\mathrm{F} 1-\mathrm{P}}=\frac{1}{2}[1 & +a^{2} P_{\mathrm{NS} 5(y 1234)}^{\mathrm{IIA}}+b^{2} P_{\mathrm{F} 1(y)}+c^{2} P_{\mathrm{P}(y)} \\
& \left.+a b\left(P_{\mathrm{D} 4(y 234)}-P_{\mathrm{D} 2(y 1)}\right)+b c\left(P_{\mathrm{P}(1)}-P_{\mathrm{F} 1(1)}\right)+c a\left(P_{\mathrm{D} 4(1234)}-P_{\mathrm{D} 0}\right)\right] .
\end{aligned}
$$

- Put $c=0$
- NS5(y, $\left.T^{4}\right)$, F1 $(y) \longrightarrow$ local D4(y234), D2(y1)


## Glueing NS5 and F1

$\Pi_{\mathrm{NS} 5-\mathrm{F} 1-\mathrm{P}}=\frac{1}{2}\left[1+a^{2} P_{\mathrm{NS}(y 1234)}^{\mathrm{IIA}}+b^{2} P_{\mathrm{F} 1(y)}+c^{2} P_{\mathrm{P}(y)}\right.$

$$
\left.+a b\left(P_{\mathrm{D} 4(y 234)}-P_{\mathrm{D} 2(y 1)}\right)+b c\left(P_{\mathrm{P}(1)}-P_{\mathrm{F} 1(1)}\right)+c a\left(P_{\mathrm{D} 4(1234)}-P_{\mathrm{D} 0}\right)\right] .
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 become smooth:
$\rightarrow$ new brane system looks like a furrow along $y$.


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$\rightarrow$ new brane system looks like a furrow along $y$.
$\uparrow$ This M5-M2 furrow is dual to a
D4-F1 Callan-Maldacena spike


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\end{aligned}
$$

- The furrow interpolates between M5 and M2:

$$
a=\cos \beta, \quad b=\sin \beta
$$

$\Rightarrow$ The orientation of a local piece of the furrow determines the ratio
 between M5 and M2 charges.

## Transition of a M5-M2 black-hole microstate

- Local transition $\Rightarrow$ a M5M2 black-hole microstate will transition into a «labyrinth/maze»
$\rightarrow$ «super-maze »



## Glueing NS5, F1 and P

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$$

- The M5-M2 furrow carries momentum through ripples modulated orthogonally to its surface

$$
\begin{gathered}
a=\cos \alpha \cos \beta \\
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c=\sin \alpha
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- $\beta$ controls the bending angle of the furrow; $\alpha$ controls the angle of ripples orthogonal to the furrow.



## Consequence on a M5-M2-P microstate

- The ripples of the furrow correspond to shape modes of the M5-M2 labyrinth
- The shape modes are the way 16 -susy microstates carry momentum.



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## Consequence on a M5-M2-P microstate

- The ripples of the furrow correspond to shape modes of the M5-M2 maze
- The shape modes are the way 16 -local-susy microstates carry momentum.
$\Rightarrow$ The microstates are ensured to have exact spherical symmetry.



## Horizonless geometries in supergravity?

- Can they described in supergravity?


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- In Type IIA, one can only separate the NS5 branes in the non-compact spatial dimensions. There exists a region close to the branes where the dilaton gets a large value, so supergravity breaks down.
e.g. [Martinec, Massai, Turton '22]


## Horizonless geometries in supergravity?

- Can they described in supergravity?
- In Type IIA, one can only separate the NS5 branes in the non-compact spatial dimensions. There exists a region close to the branes where the dilaton gets a large value, so supergravity breaks down.
e.g. [Martinec, Massai, Turton '22]
- But in the M2-M5-P frame, the basic ingredient of the super-maze is a M5 brane with M2 flux on it. The supergravity description of it is valid close to the branes as well.


## Conclusion

- Global charges and supersymmetries control the near-horizon geometry.
- Local supersymmetries are a means to get information on the microstates.


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$\uparrow$ This is crucial in order to understand whether microstates in string theory resolve the singularity or the horizon.


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- Local supersymmetries are a means to get information on the microstates.
- $1 / 8$-BPS systems ( 3 -charge BH ) have a large moduli space of solutions that have more supersymmetries locally
$\uparrow$ This is crucial in order to understand whether microstates in string theory resolve the singularity or the horizon.
- The microstate geometries programme used to replace D1-D5P horizons with brane systems that extend in $\mathbb{R}^{4}$
$\uparrow$ But this approach seems to have limits: entropy, typicality...



## Conclusion

- New approach: microstates can carry momentum by having motion in the internal dimensions $\Rightarrow$ exactly spherical symmetry


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... and we have identified what they become when the branes start interacting.



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- New approach: microstates can carry momentum by having motion in the internal dimensions $\Rightarrow$ exactly spherical symmetry
- The DVVM microstates account for the black-hole entropy...
... and we have identified what they become when the branes start interacting.

- These «super-mazes »have 16 local susys, just like the superstrata, but without having their drawbacks.



## Outlook

- 16 local susys is a smoking gun for horizonless microstate solutions



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- 16 local susys is a smoking gun for horizonless microstate solutions
$\Rightarrow$ seems to support Fuzzball hypothesis for M2-M5-P blackhole microstates
$\uparrow$ Construct the fully backreacted supergravity solutions
$\uparrow$ Apply geometric quantization to them.


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$\Rightarrow$ seems to support Fuzzball hypothesis for M2-M5-P black-
 hole microstates
$\uparrow$ Construct the fully backreacted supergravity solutions
$\uparrow$ Apply geometric quantization to them.
- End goal:
«Where » is the information about the black-hole microstate?



## Back-up slide

|  |  |  |
| :--- | :--- | :--- |
|  | M2 |  |
|  |  |  |
|  |  | M5 |
|  |  |  |





