Local supersymmetry enhancement and the entropy of three-charge black holes

Iberian Strings 2023, Murcia



Max-Planck-Institut für Physik (Werner-Heisenberg-Institut)

Based on [2211.14326] with I. Bena, S. Hampton, A. Houppe and D. Toulikas

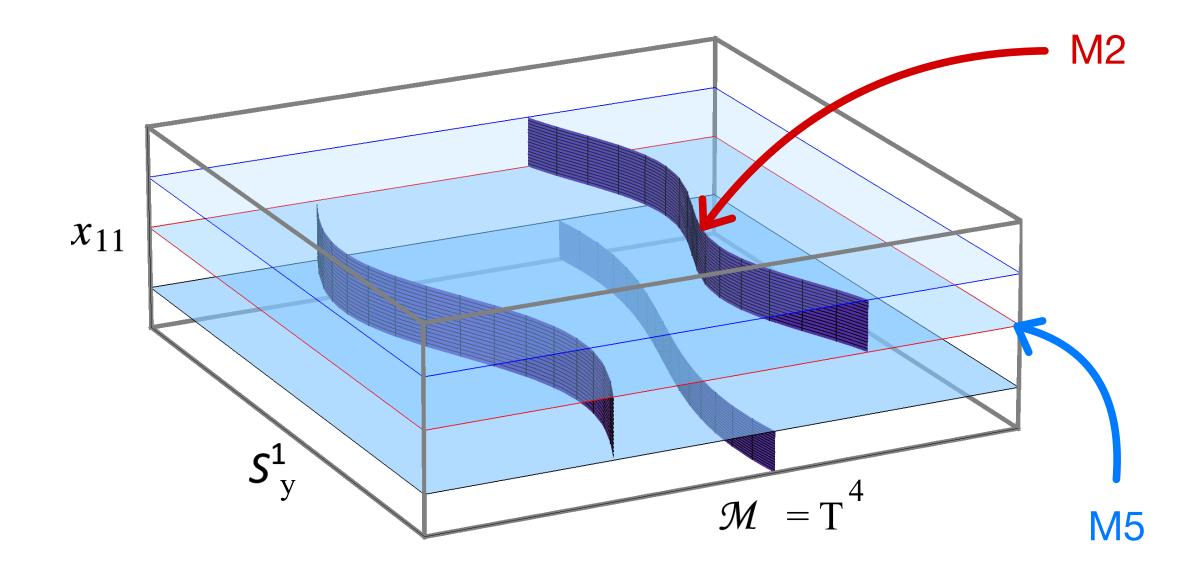
January 12th, 2023

Yixuan Li

MPI Munich



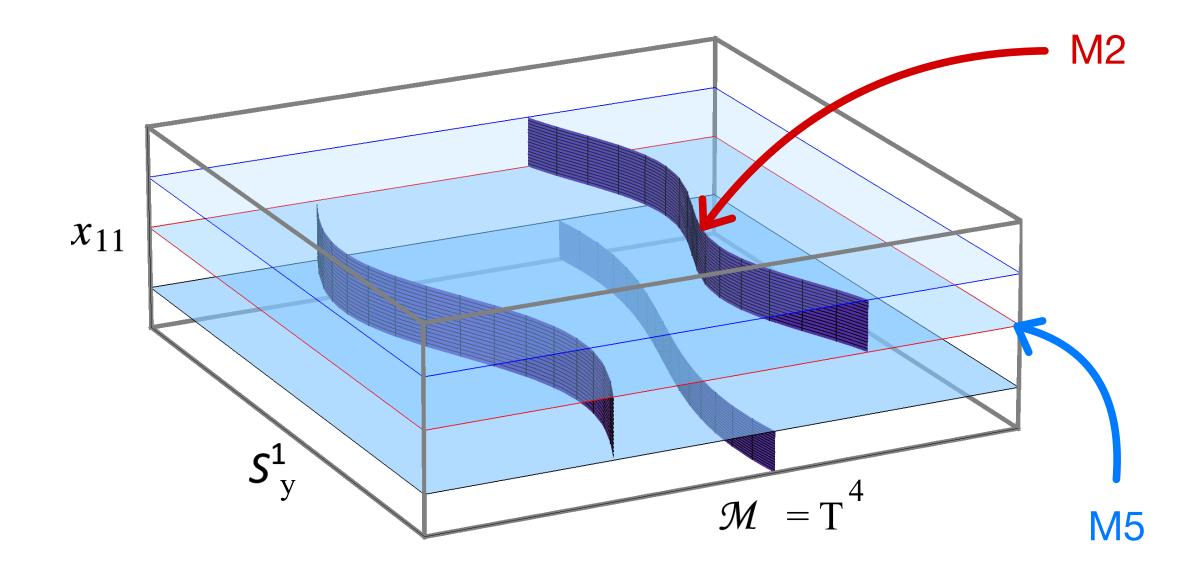
What this talk is about



« Dijkgraaf-Verlinde-Verlinde -Maldacena microstates »

- M5-M2(-P) black hole: The microstates that are made of **fractionated M2 branes** account for the entropy.
- We found: They can transition into microstates with 16 local supersymmetries.

What this talk is about



« Dijkgraaf-Verlinde-Verlinde -Maldacena microstates »

- M5-M2(-P) black hole: The microstates that are made of **fractionated M2 branes** account for the entropy.
- We found: They can transition into microstates with 16 local supersymmetries.

Microstates with 16 local susys account for the black-hole entropy!

• We expect their backreaction to be horizonless microstates.



1. Local supersymmetry enhancement and black-hole microstates

2. The new M5-M2-P microstates with 16 local supersymmetries



1. Local supersymmetry enhancement and black-hole microstates

2. The new M5-M2-P microstates with 16 local supersymmetries

- Type IIA/IIB: $\mathbb{R}^{4,1} \times S_v^1 \times T^4$
- Take brane system with 3 charges:
 D5(y, T⁴), D1(y), P(y)
 or NS5(y, T⁴), F1(y), P(y)
 - naively, 1/8-BPS everywhere



- Type IIA/IIB: $\mathbb{R}^{4,1} \times S_v^1 \times T^4$
- Take brane system with 3 charges:
 D5(y, T⁴), D1(y), P(y)
 or NS5(y, T⁴), F1(y), P(y)
 - naively, 1/8-BPS everywhere
 - Use harmonic function rule

$$ds^{2} = -\frac{2}{\sqrt{H_{1}H_{5}}} \left[dt^{2} + dy^{2} + (H_{P} - 1)^{-1} (dy - dt) + \sqrt{H_{1}H_{5}} ds^{2}_{\mathbb{R}^{4}} + (H_{1}H_{5})^{-1/2} ds^{2}_{T^{4}} \right]$$

with
$$H_{1,5,P} = 1 + rac{Q_{1,5,P}}{r^2}$$
 .

 $[t)^2$

- Type IIA/IIB: $\mathbb{R}^{4,1} \times S_v^1 \times T^4$
- Take brane system with 3 charges:
 D5(y, T⁴), D1(y), P(y)
 or NS5(y, T⁴), F1(y), P(y)
 - naively, 1/8-BPS everywhere
 - Use harmonic function rule
 ⇒ develops horizon in
 supergravity

$$ds^{2} = -\frac{2}{\sqrt{H_{1}H_{5}}} \left[dt^{2} + dy^{2} + (H_{P} - 1)^{-1} (dy - dt) + \sqrt{H_{1}H_{5}} ds^{2}_{\mathbb{R}^{4}} + (H_{1}H_{5})^{-1/2} ds^{2}_{T^{4}} \right]$$

with
$$H_{1,5,P} = 1 + rac{Q_{1,5,P}}{r^2}$$
 .

 $[t)^2$

- Type IIA/IIB: $\mathbb{R}^{4,1} \times S_v^1 \times T^4$
- Take brane system with 3 charges:
 D5(y, T⁴), D1(y), P(y)
 or NS5(y, T⁴), F1(y), P(y)
 - naively, 1/8-BPS everywhere
 - Use harmonic function rule
 ⇒ develops horizon in
 supergravity

Possible conclusion:

Global charges and supersymmetries seem to control near-horizon geometry.

Therefore all brane systems develop the same horizon: To have access the information about the microstates, probe singularity region, where supergravity breaks down.

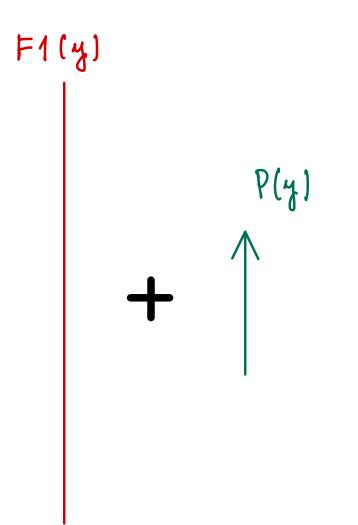


- Type IIA/IIB: $\mathbb{R}^{4,1} \times S_v^1 \times T^4$
- Take brane system with 3 charges:
 D5(y, T⁴), D1(y), P(y)
 or NS5(y, T⁴), F1(y), P(y)
 - naively, 1/8-BPS everywhere
 - Use harmonic function rule
 ⇒ develops horizon in
 supergravity

Local supersymmetry enhancement: String-theory excitations (branes,

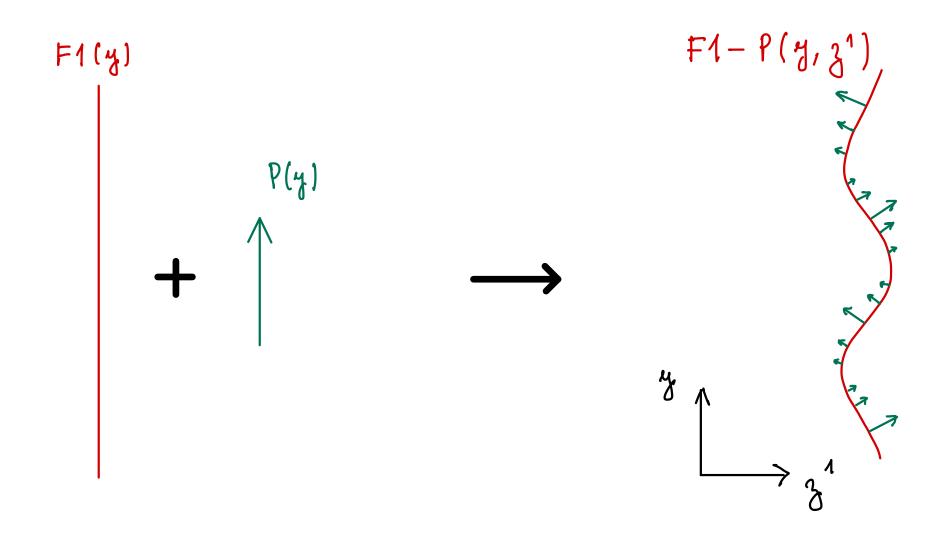
strings) combine together to form a *bound state* that is locally 1/2-BPS (*16 susies*).

• Ex.: F1(y) and parallel P(y):



- F1or P preserve 16 real supercharges
- Together, F1-P preserve 8

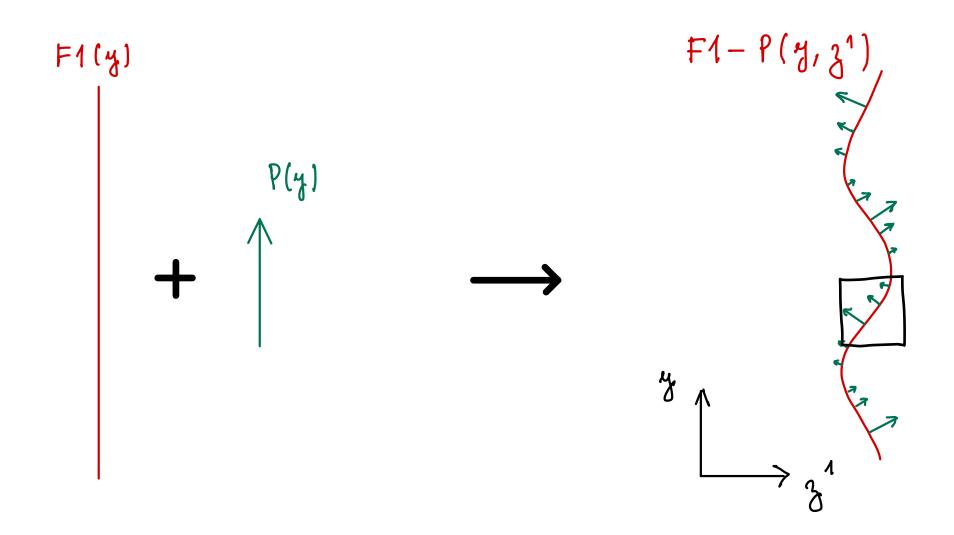
• Ex.: F1(y) and parallel P(y):



- F1or P preserve 16 real supercharges
- Together, F1-P preserve 8

- Actually the string can carry momentum: profile
- The F1-P profile preserves the same global supersymmetries...

• Ex.: F1(y) and parallel P(y):



- F1or P preserve 16 real supercharges
- Together, F1-P preserve 8

- Actually the string can carry momentum: profile
- The F1-P profile preserves the same global supersymmetries...
 - ...but locally it is a F1(\hat{y}) boosted by orthogonal P(\hat{y}^{\perp})
 - $F1(\hat{y})$ -P (\hat{y}^{\perp}) preserves 16 supercharges

Local supersymmetry → information on microstate?



• Branes, strings \rightarrow constraint on ϵ :

$\prod_{t} \epsilon \equiv \frac{1}{2}(1+P)\epsilon = 0$ Projector Traceless involution $\Pi^2 = \Pi$ $P^{2} = 1$

 $\operatorname{tr}(P) = 0$

Local VS global supersymmetries

See e.g. [Bena, Hampton, Houppe, YL, Toulikas '22]



Constraint halves number of supersymmetries



- Branes, strings \rightarrow constraint on ϵ : $\Pi \epsilon \equiv \frac{1}{2}(1+P) \epsilon = 0$
- Combine k different excitations: $\epsilon \in \ker(\Pi_1) \cap \ldots \cap \ker(\Pi_k)$ \rightarrow # (global supersymmetries)



- Branes, strings \rightarrow constraint on ϵ : $\Pi \epsilon \equiv \frac{1}{2}(1+P) \epsilon = 0$
- Combine k different excitations: $\epsilon \in \ker(\Pi_1) \cap \ldots \cap \ker(\Pi_k)$ \rightarrow \ddagger (global supersymmetries)
- Add other involutions $(P_{k+1}, ..., P_n)$ and weights $(\alpha_1, \ldots, \alpha_n)$ s.t. $\alpha_1 + \ldots + \alpha_n = 1.$

See e.g. [Bena, Hampton, Houppe, YL, Toulikas '22]

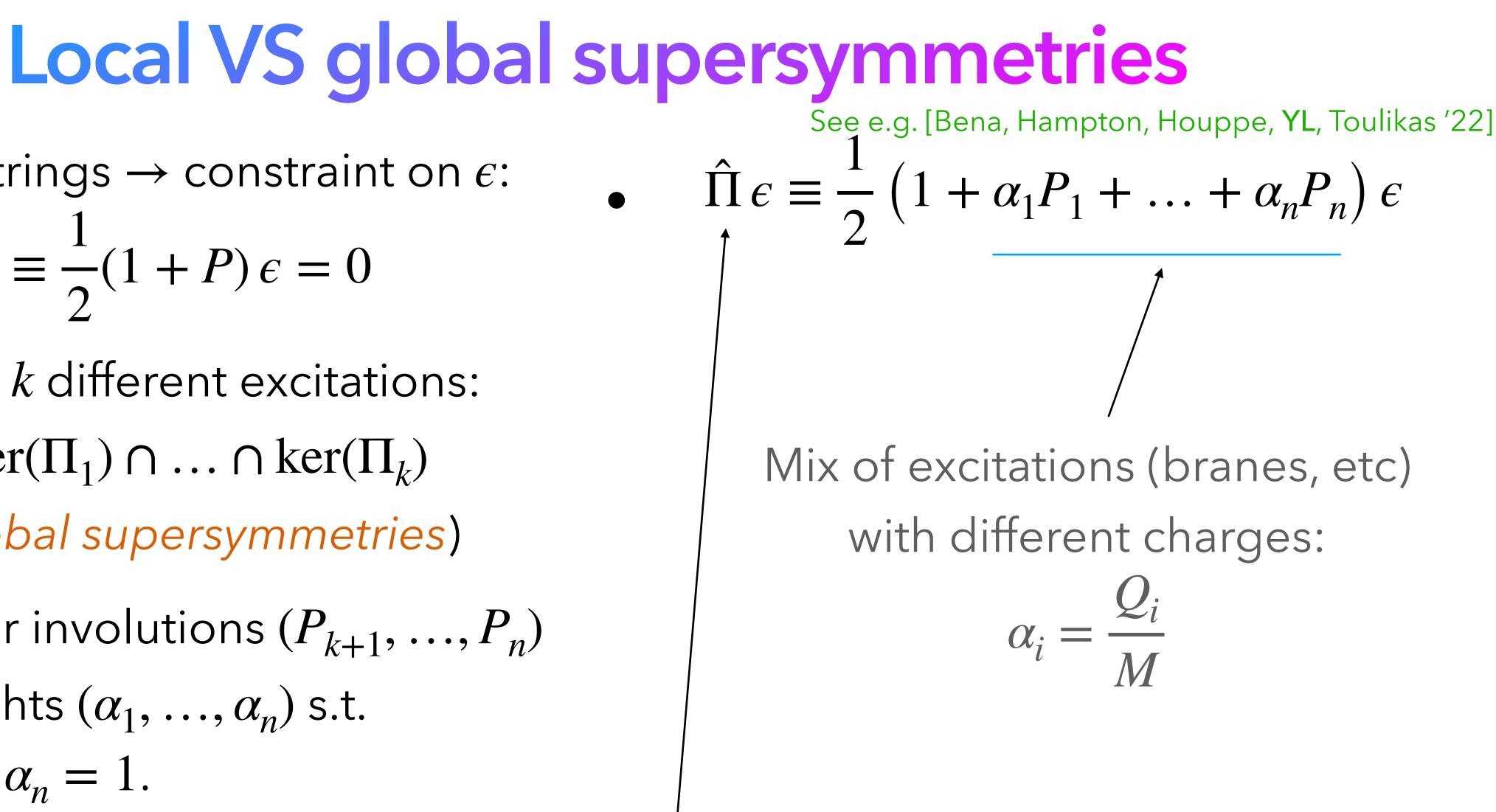
$$\hat{\Pi} \epsilon \equiv \frac{1}{2} \left(1 + \alpha_1 P_1 + \dots + \alpha_n P_n \right) \epsilon$$

Mix of excitations (branes, etc) with different charges:

$$\alpha_i = \frac{Q_i}{M}$$



- Branes, strings \rightarrow constraint on ϵ : $\Pi \epsilon \equiv \frac{1}{2}(1+P) \epsilon = 0$
- Combine k different excitations: $\epsilon \in \ker(\Pi_1) \cap \ldots \cap \ker(\Pi_k)$ \rightarrow \ddagger (global supersymmetries)
- Add other involutions $(P_{k+1}, ..., P_n)$ and weights $(\alpha_1, \ldots, \alpha_n)$ s.t. $\alpha_1 + \ldots + \alpha_n = 1.$



Not necessarily a projector!



- Branes, strings \rightarrow constraint on ϵ : $\Pi \epsilon \equiv \frac{1}{2}(1+P) \epsilon = 0$
- Combine k different excitations: $\epsilon \in \ker(\Pi_1) \cap \ldots \cap \ker(\Pi_k)$ \rightarrow \ddagger (global supersymmetries)
- Add other involutions $(P_{k+1}, ..., P_n)$ and weights $(\alpha_1, \ldots, \alpha_n)$ s.t. $\alpha_1 + \ldots + \alpha_n = 1.$

See e.g. [Bena, Hampton, Houppe, YL, Toulikas '22]

•
$$\hat{\Pi} \epsilon \equiv \frac{1}{2} \left(1 + \alpha_1 P_1 + \dots + \alpha_n P_n \right) \epsilon$$

• $\hat{\Pi}^2 = \hat{\Pi}$ iff the system has 16 susies.

One can enhance supersymmetries



- Branes, strings \rightarrow constraint on ϵ : $\Pi \epsilon \equiv \frac{1}{2}(1+P) \epsilon = 0$
- Combine k different excitations: $\epsilon \in \ker(\Pi_1) \cap \ldots \cap \ker(\Pi_k)$ \rightarrow \ddagger (global supersymmetries)
- Add other involutions $(P_{k+1}, ..., P_n)$ and weights $(\alpha_1, \ldots, \alpha_n)$ s.t. $\alpha_1 + \ldots + \alpha_n = 1.$

See e.g. [Bena, Hampton, Houppe, YL, Toulikas '22]

•
$$\hat{\Pi} \epsilon \equiv \frac{1}{2} \left(1 + \alpha_1 P_1 + \dots + \alpha_n P_n \right) \epsilon$$

- $\hat{\Pi}^2 = \hat{\Pi}$ iff the system has 16 susies.
- $\{\alpha_i\}$ not unique $\rightarrow \{\alpha_i(x)\}$ $\hat{\Pi}(x)\,\epsilon(x)=0^{n}$

along the bound state

 ϵ promoted to be a function



- Branes, strings \rightarrow constraint on ϵ : $\Pi \epsilon \equiv \frac{1}{2}(1+P) \epsilon = 0$
- Combine k different excitations: $\epsilon \in \ker(\Pi_1) \cap \ldots \cap \ker(\Pi_k)$ \rightarrow \ddagger (global supersymmetries)
- Add other involutions $(P_{k+1}, ..., P_n)$ and weights $(\alpha_1, \ldots, \alpha_n)$ s.t. $\alpha_1 + \ldots + \alpha_n = 1.$

See e.g. [Bena, Hampton, Houppe, YL, Toulikas '22]

•
$$\hat{\Pi} \epsilon \equiv \frac{1}{2} \left(1 + \alpha_1 P_1 + \dots + \alpha_n P_n \right) \epsilon$$

- $\hat{\Pi}^2 = \hat{\Pi}$ iff the system has 16 susies.
- $\{\alpha_i\}$ not unique $\rightarrow \{\alpha_i(x)\}$ $\hat{\Pi}(x)\,\epsilon(x)=0$ At $x, \epsilon(x) \in \ker(\widehat{\Pi}(x)) \to |oca|$ supersymmetry
- While for *global supersymmetry*: $\epsilon \in \bigcap \ker \left(\hat{\Pi}(x) \right).$ X

 $\ddagger (global susies) \le \ddagger (local susies)$



- Branes, strings \rightarrow constraint on ϵ : $\Pi \epsilon \equiv \frac{1}{2}(1+P) \epsilon = 0$
- Combine k different excitations: $\epsilon \in \ker(\Pi_1) \cap \ldots \cap \ker(\Pi_k)$ \rightarrow # (global supersymmetries)

Local supersymmetry enhancement

See e.g. [Bena, Hampton, Houppe, YL, Toulikas '22]

•
$$\hat{\Pi} \epsilon \equiv \frac{1}{2} \left(1 + \alpha_1 P_1 + \dots + \alpha_n P_n \right) \epsilon$$

- $\hat{\Pi}^2 = \hat{\Pi}$ iff the system has 16 susies.
- $\{\alpha_i\}$ not unique $\rightarrow \{\alpha_i(x)\}$ $\hat{\Pi}(x)\,\epsilon(x)=0$ At $x, \epsilon(x) \in \ker(\widehat{\Pi}(x)) \to |oca|$ supersymmetry
- While for *global supersymmetry*: $\epsilon \in \bigcap \ker \left(\hat{\Pi}(x) \right).$

 $\ddagger(global susies) \leq \ddagger(local susies)$



Local supersymmetry enhancement

Local supersymmetry enhancement:

Given a set of global supersymmetries, there sometimes exists a whole moduli space of brane/string systems, parameterised by $\{\alpha_i(x)\}$, preserving those same global supersymmetries, but whose number of local supersymmetries is enhanced.

- identify the additional excitations (« glues ») to make a bound state

– determine the charge-to-mass ratios $\{\alpha_i(x)\}$.



Local supersymmetry enhancement

Local supersymmetry enhancement:

Given a set of global supersymmetries, there sometimes exists a whole moduli space of brane/string systems, parameterised by $\{\alpha_i(x)\}$, preserving those same global supersymmetries, but whose number of local supersymmetries is enhanced.

- identify the additional excitations (« glues ») to make a bound state

– determine the charge-to-mass ratios $\{\alpha_i(x)\}$.



Local supersymmetry enhancement

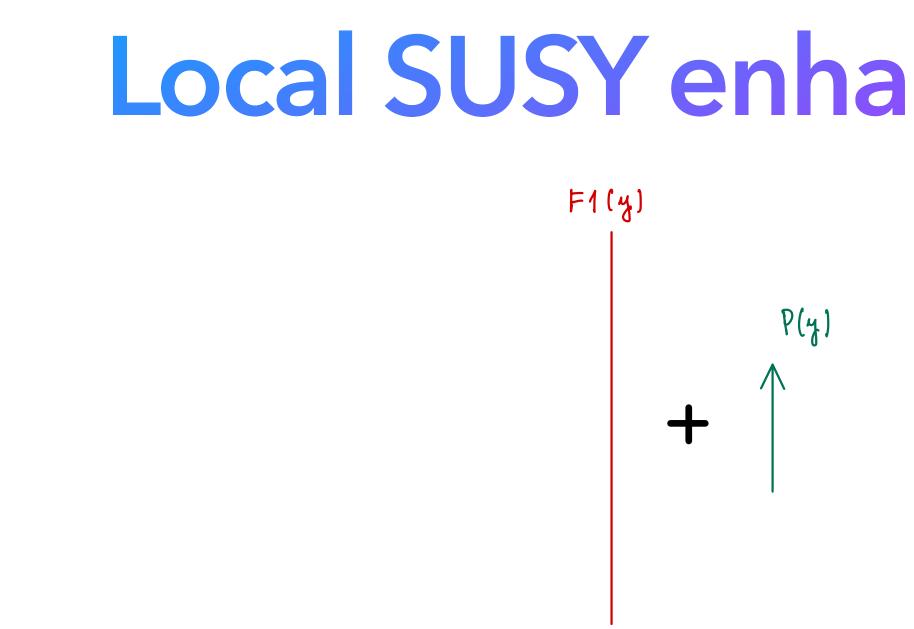
Local supersymmetry enhancement:

Given a set of global supersymmetries, there sometimes exists a whole moduli space of brane/string systems, parameterised by $\{\alpha_i(x)\}$, preserving those same global supersymmetries, but whose number of local supersymmetries is enhanced.

- identify the additional excitations (« glues ») to make a bound state

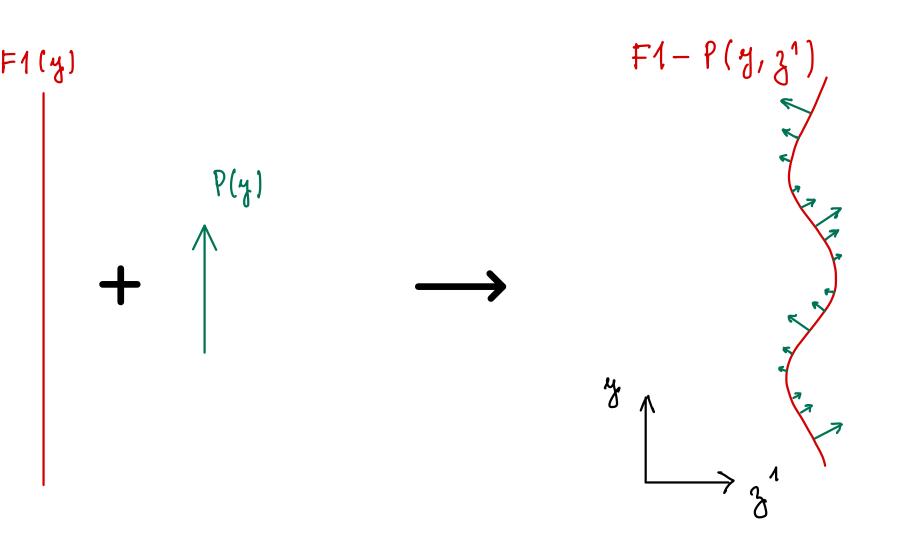
– determine the charge-to-mass ratios $\{\alpha_i(x)\}$.





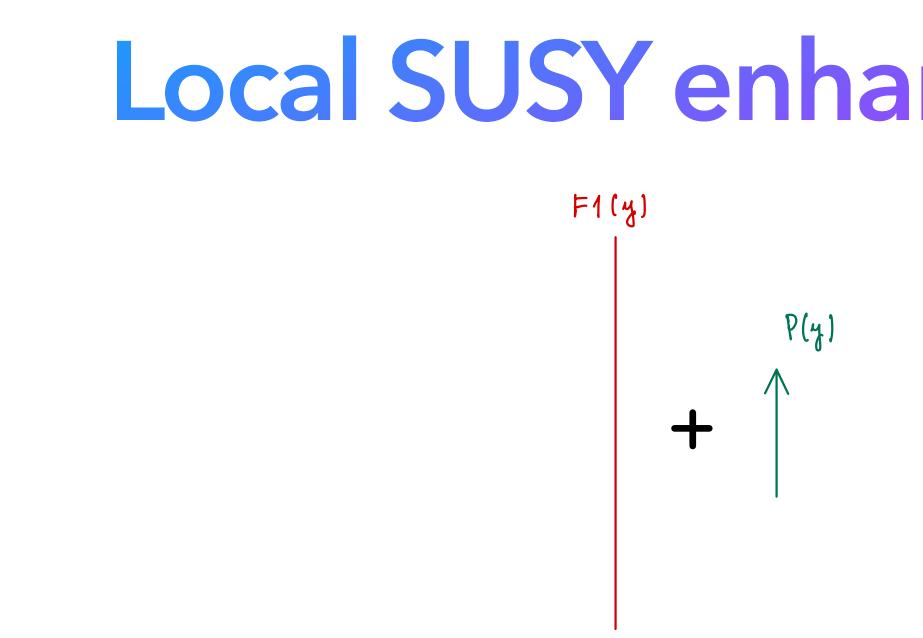
•
$$\Pi_{F1(y)} = \frac{1}{2}(1 + P_{F1(y)}), P_{F1(y)} = \Gamma^{0y}\sigma_3$$

• $\Pi_{P(y)} = \frac{1}{2}(1 + P_{P(y)}), P_{P(y)} = \Gamma^{0y}$



• $\hat{\Pi} = \frac{1}{2} \left(1 + \alpha_1 P_{F1(y)} + \alpha_2 P_{P(y)} + \alpha_3 P_{F1(1)} + \alpha_4 P_{P(1)} \right)$

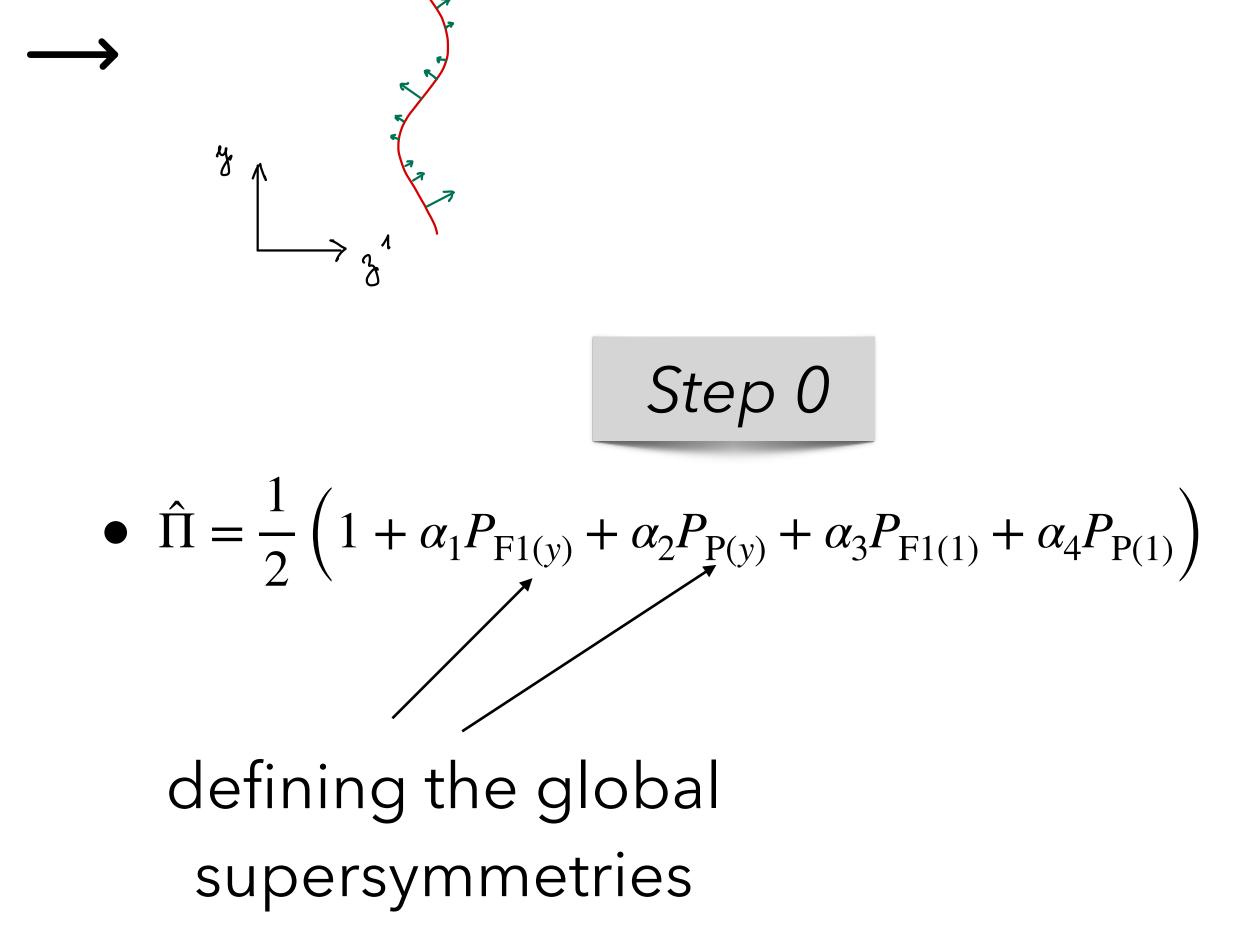


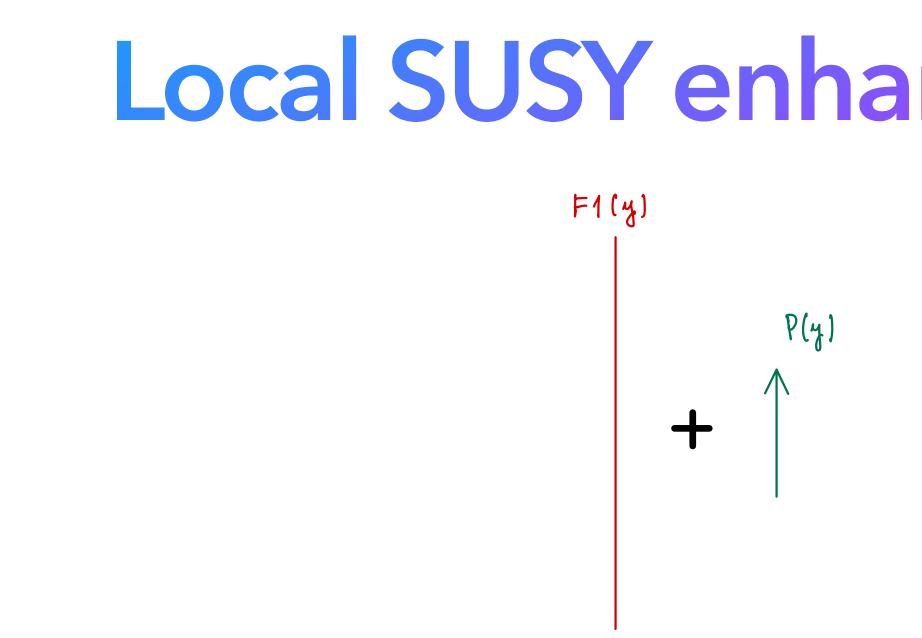


•
$$\Pi_{F1(y)} = \frac{1}{2}(1 + P_{F1(y)}), P_{F1(y)} = \Gamma^{0y}\sigma_3$$

• $\Pi_{P(y)} = \frac{1}{2}(1 + P_{P(y)}), P_{P(y)} = \Gamma^{0y}$

F1-P(y,z1)

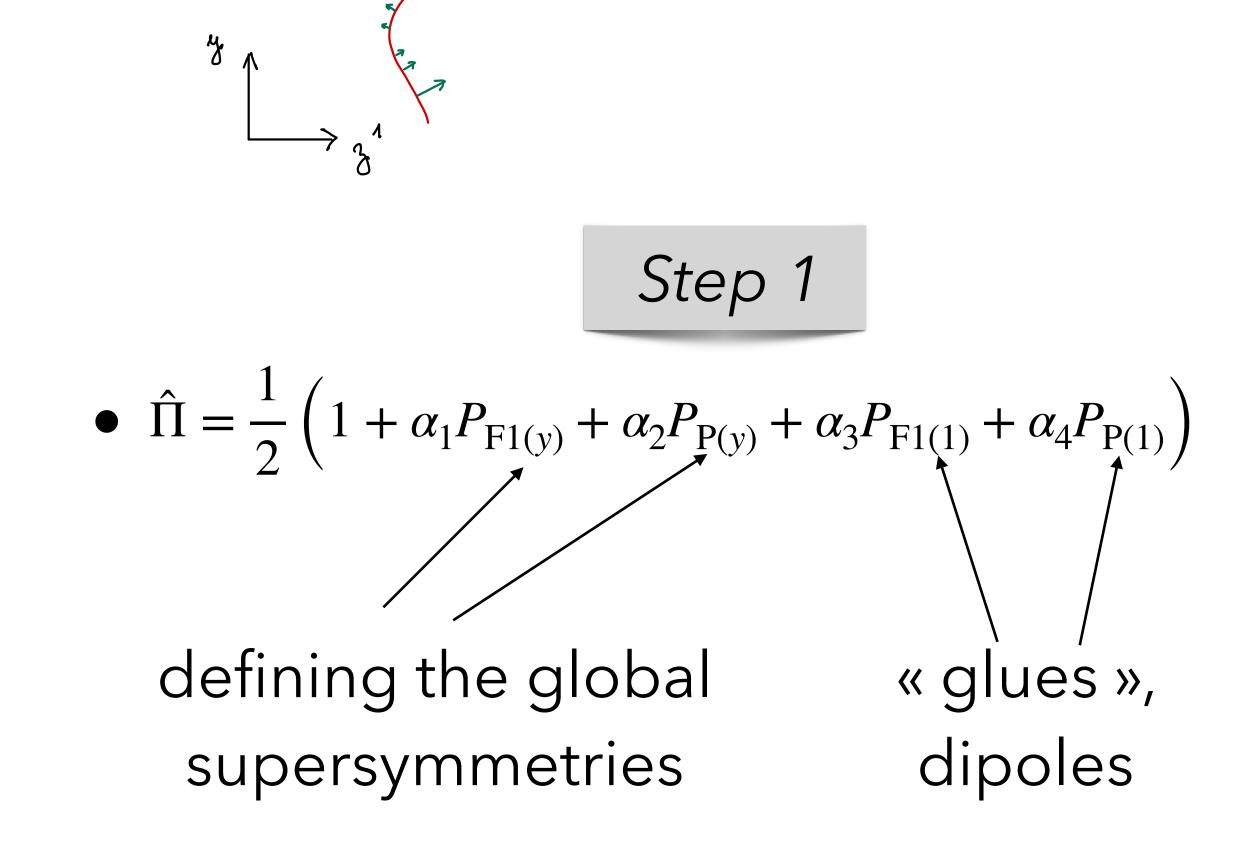


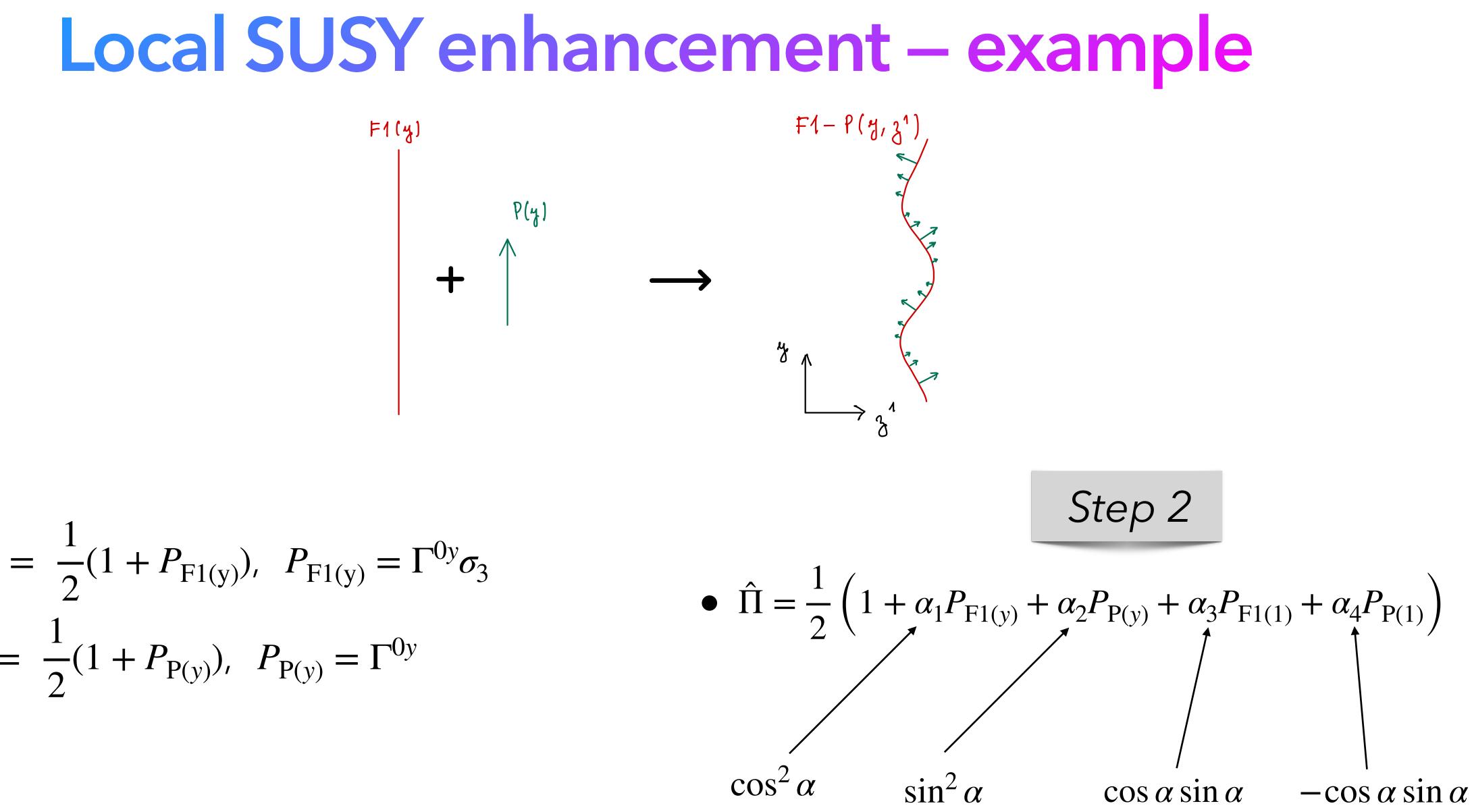


•
$$\Pi_{F1(y)} = \frac{1}{2}(1 + P_{F1(y)}), P_{F1(y)} = \Gamma^{0y}\sigma_3$$

• $\Pi_{P(y)} = \frac{1}{2}(1 + P_{P(y)}), P_{P(y)} = \Gamma^{0y}$

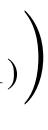
F1-P(y, 21)



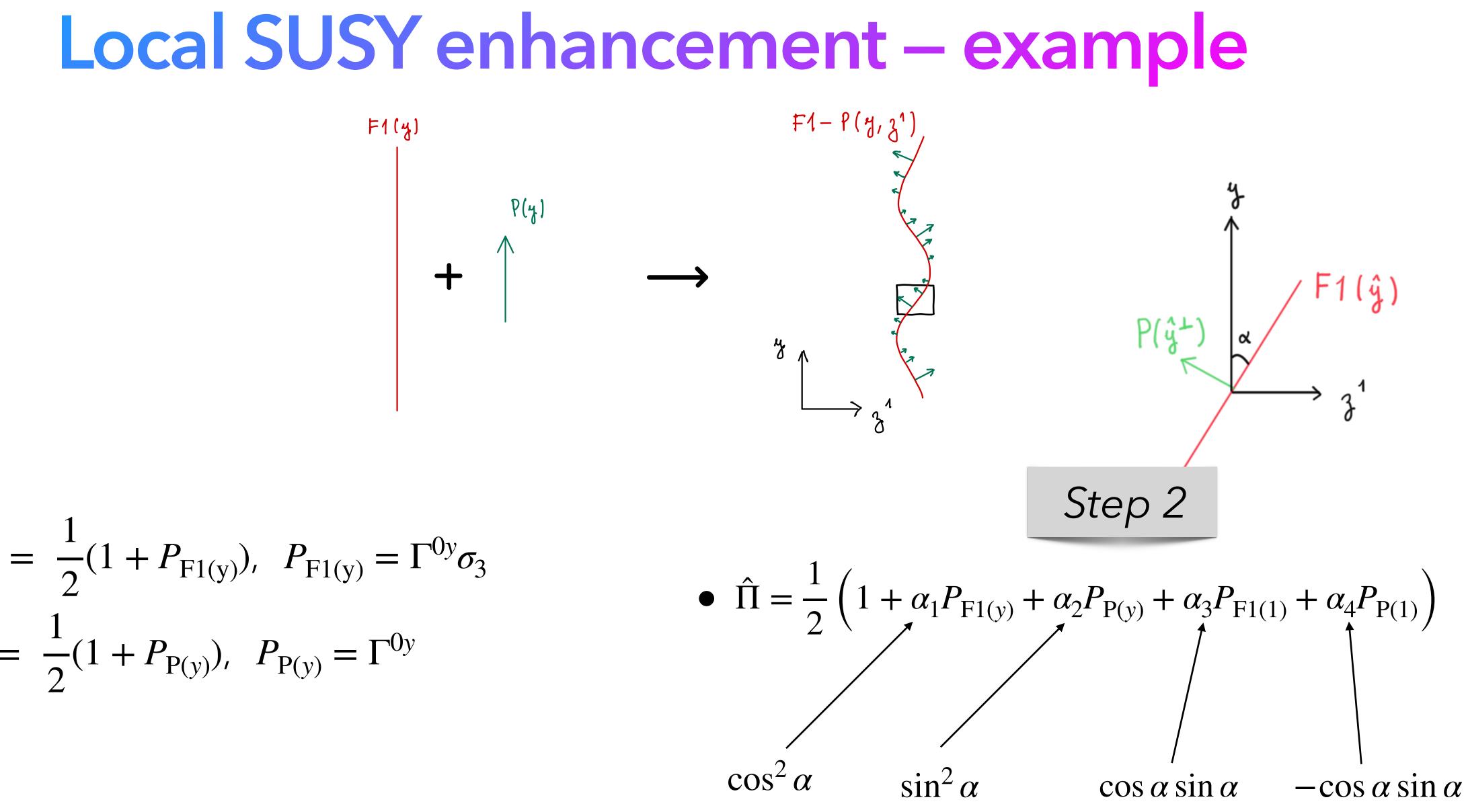


•
$$\Pi_{F1(y)} = \frac{1}{2}(1 + P_{F1(y)}), P_{F1(y)} = \Gamma^{0y}\sigma_3$$

• $\Pi_{P(y)} = \frac{1}{2}(1 + P_{P(y)}), P_{P(y)} = \Gamma^{0y}$

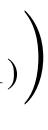






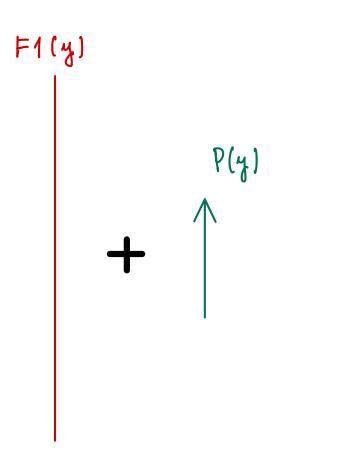
•
$$\Pi_{F1(y)} = \frac{1}{2}(1 + P_{F1(y)}), P_{F1(y)} = \Gamma^{0y}\sigma_3$$

• $\Pi_{P(y)} = \frac{1}{2}(1 + P_{P(y)}), P_{P(y)} = \Gamma^{0y}$





Microstates of the F1-P black hole



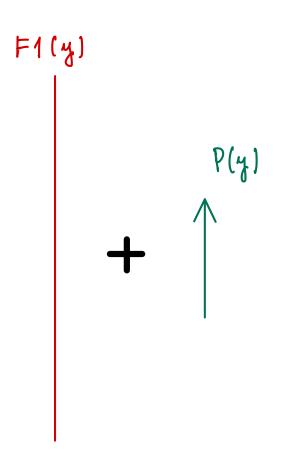
• Harmonic rule:

$$ds_{string}^{2} = H[-dudv + Kdv^{2}] + \sum_{i=1}^{4} dx_{i}dx_{i} + \sum_{a=1}^{4} dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{a}dz_{$$

 \rightarrow black hole with horizon at r = 0.

 z_a

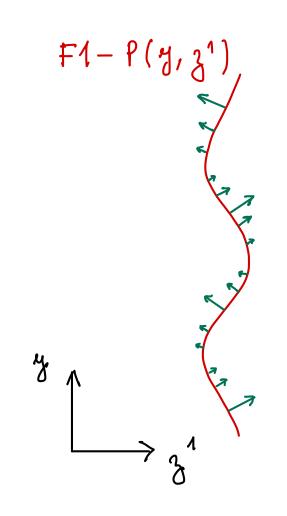
Microstates of the F1-P black hole



• Harmonic rule:

$$ds_{string}^{2} = H[-dudv + Kdv^{2}] + \sum_{i=1}^{4} dx_{i}dx_{i} + \sum_{a=1}^{4} dz_{a}dz_{a}$$

 \rightarrow black hole with horizon at r = 0.



• Metric sourced by the string:

 $ds_{string}^{2} = H[-dudv + Kdv^{2} + 2A_{i}dx_{i}dv] + \sum_{i=1}^{4} dx_{i}dx_{i} + \sum_{i=1}^{4} dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i}dz_{i$

[Dabhorkar, Gauntlett, Harvey, Waldram '95]

 \rightarrow smooth, horizonless solution.





Microstates of the F1-P black hole

The massive string states accounting for the F1-P blackhole entropy can be described in supergravity.

• Harmonic rule:

$$ds_{string}^{2} = H[-dudv + Kdv^{2}] + \sum_{i=1}^{4} dx_{i}dx_{i} + \sum_{a=1}^{4} dz_{a}dz_{a} \qquad ds_{string}^{2} = H[-dudv + Kdv^{2} + 2A_{i}dx_{i}dv] + \sum_{i=1}^{4} dx_{i}dx_{i} + \sum_{a=1}^{4} dz_{a}dz_{a}$$

 \rightarrow black hole with horizon at r = 0.

• Metric sourced by the string:

[Dabhorkar, Gauntlett, Harvey, Waldram '95]

 \rightarrow smooth, horizonless solution.





2-charge VS 3-charge black holes

 Such « classical » string profiles, thr the F1-P black-hole entropy:

• Such « classical » string profiles, through geometric quantization, account for

[Lunin, Mathur '01], [Rychkov '05]



2-charge VS 3-charge black holes

the F1-P black-hole entropy:

• However: 2-charge black holes: singularity, horizon at r = 0.

 \rightarrow Is the stringy structure resolving the horizon or the singularity?

• Such « classical » string profiles, through geometric quantization, account for

[Lunin, Mathur '01], [Rychkov '05]

 $S = 2\pi \sqrt{N_1 N_P}$

2-charge VS 3-charge black holes

the F1-P black-hole entropy:

• However: 2-charge black holes: singularity, horizon at r = 0.

 \rightarrow Is the stringy structure resolving the horizon or the singularity?

• 3-charge black holes: singularity and horizon separated. \rightarrow in particular D1-D5-P or F1-NS5-P

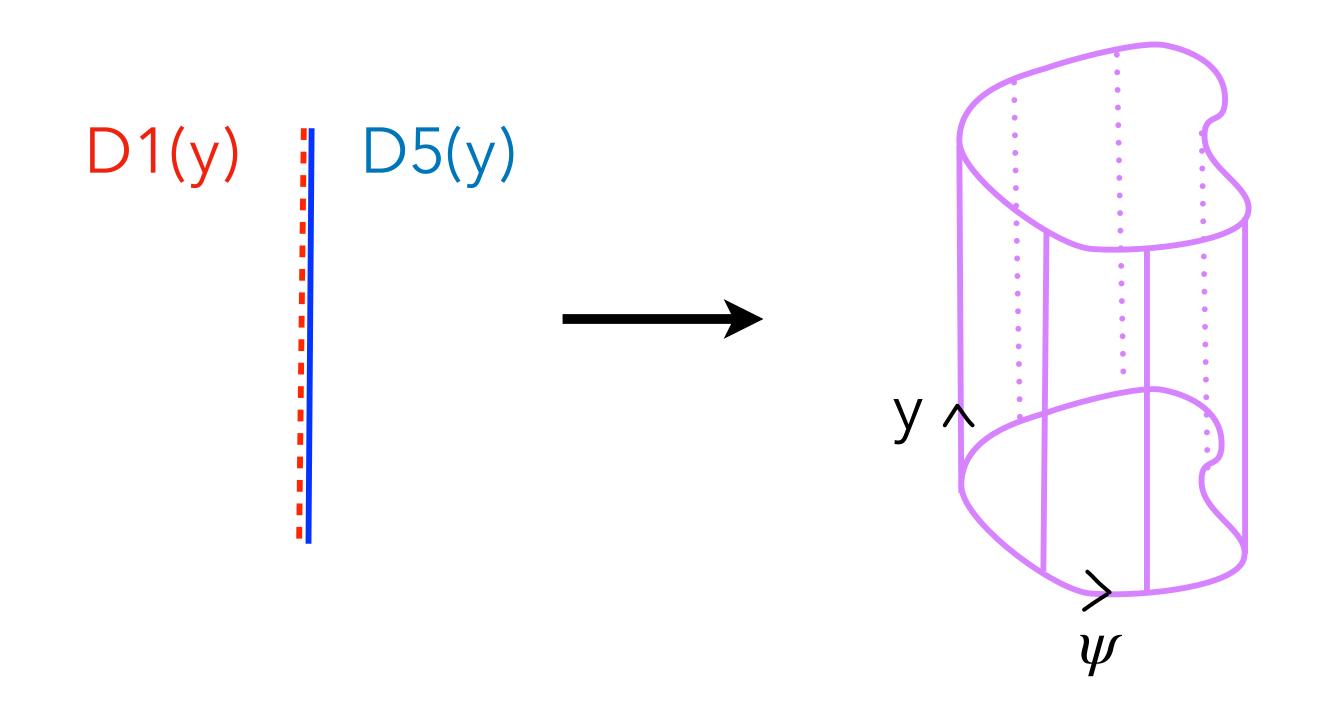
• Such « classical » string profiles, through geometric quantization, account for

[Lunin, Mathur '01], [Rychkov '05]

 $S = 2\pi \sqrt{N_1 N_P}$

1st approach: enhancing D1-D5 through KKM

- D1(y), D5(y1234) \longrightarrow KKM(1234 ψ , y), P(ψ) dipoles
 - The D1-D5 brane system gains a dimension through the KKM

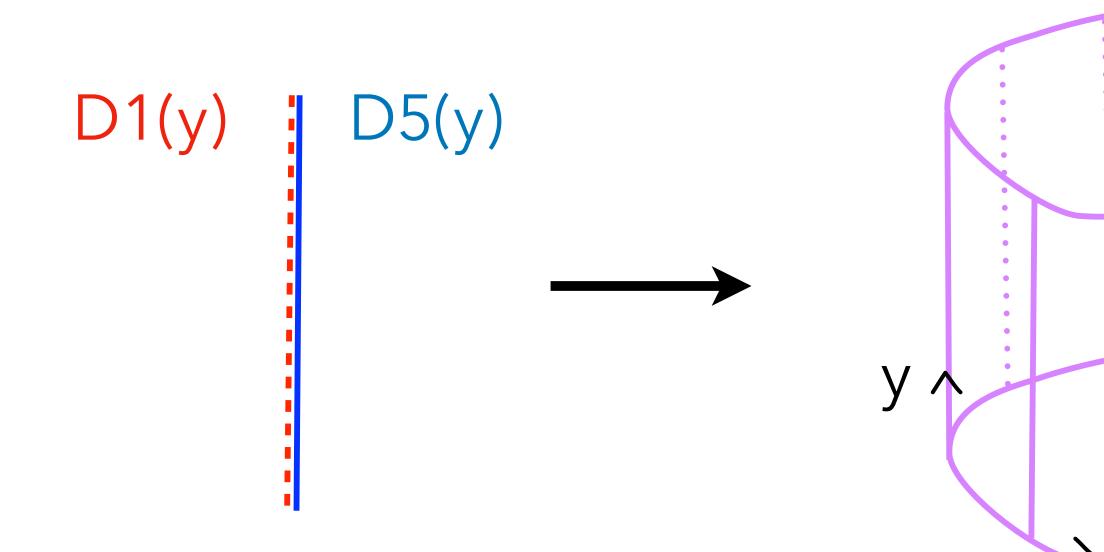


 \rightarrow « supertube »

[Emparan, Mateos, Townsend '01]

1st approach: enhancing D1-D5 through KKM

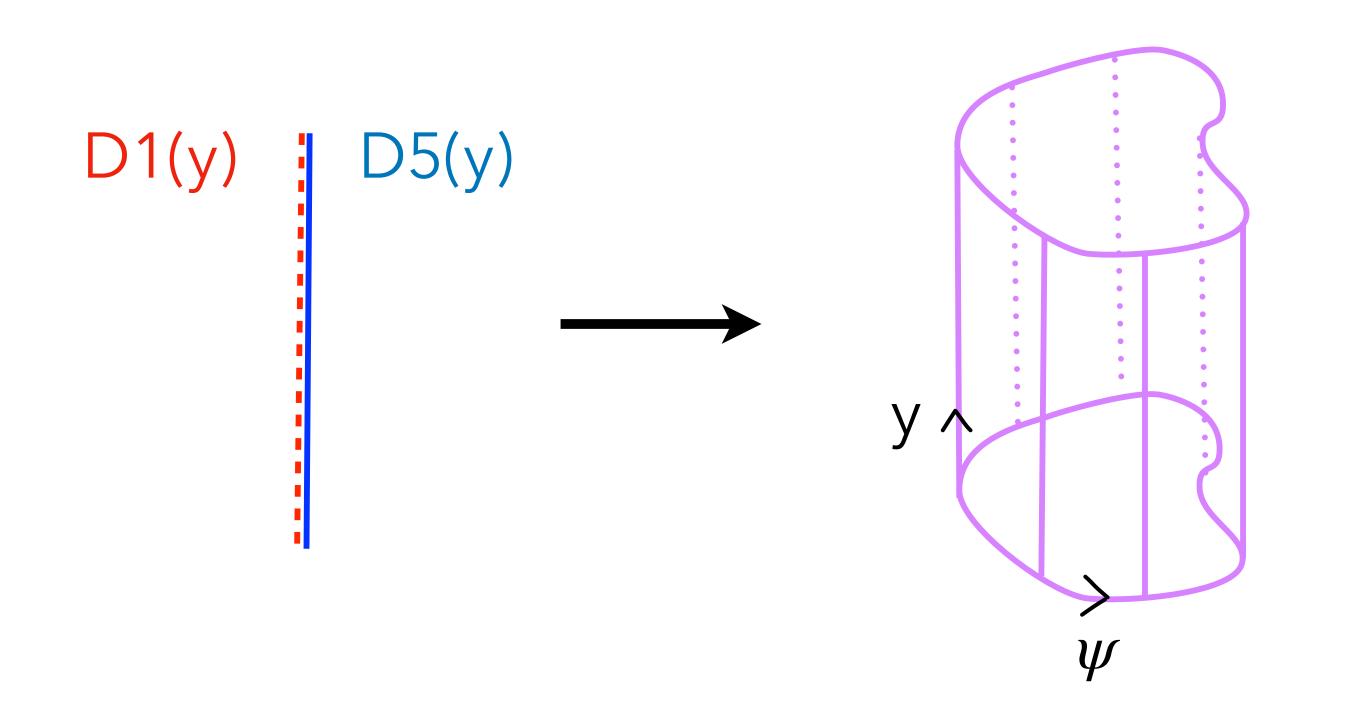
- D1(y), D5(y1234) \longrightarrow KKM(1234 ψ , y), P(ψ) dipoles
 - The D1-D5 brane system gains a dimension through the KKM
 - The (angular) momentum $P(\psi)$ stabilises the size of the supertube.



→ replace the deltafunction brane singularity by a source extended in the non-compact dimensions

1st approach: enhancing D1-D5 through KKM

- D1(y), D5(y1234) \longrightarrow KKM(1234 ψ , y), P(ψ) dipoles
 - The D1-D5 brane system gains a dimension through the KKM
 - The (angular) momentum $P(\psi)$ stabilises the size of the supertube.



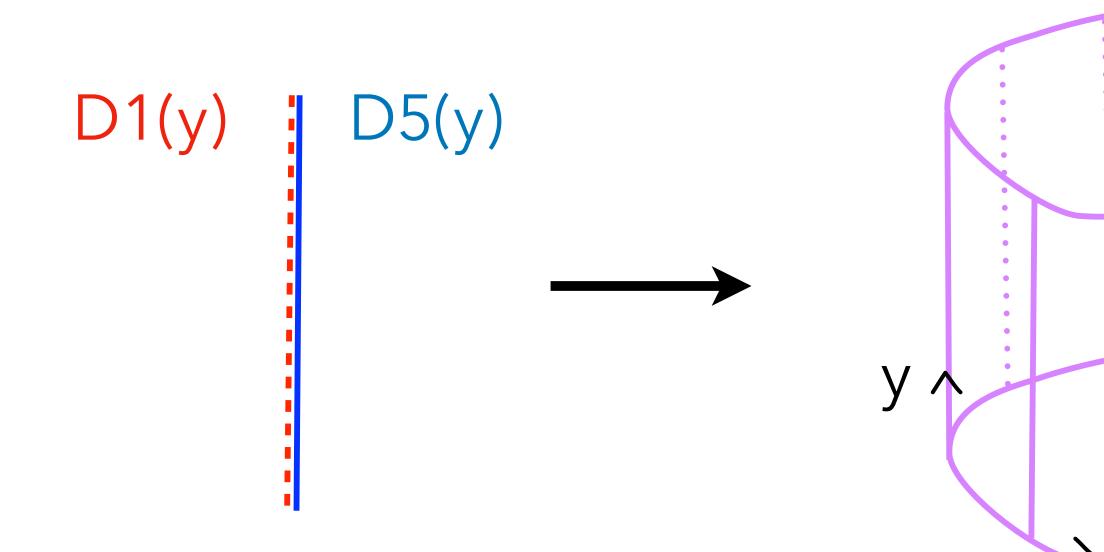
• The bound state is globally 1/4-BPS, but locally 1/2-BPS

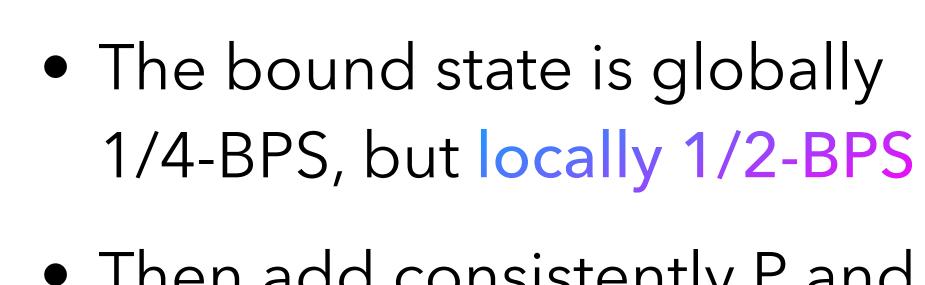


1st approach: enhancing D1-D5 through KKM

- D1(y), D5(y1234) \longrightarrow KKM(1234 ψ , y), P(ψ) dipoles
 - The D1-D5 brane system gains a dimension through the KKM
 - The (angular) momentum $P(\psi)$ stabilises the size of the supertube.

Ψ





 Then add consistently P and keep locally 1/2-BPS

\rightarrow « superstrata »

[Bena, de Boer, Shigemori, Warner '11] [Bena, Giusto, Martinec, Russo,

Shigemori, Turton, Warner '16]





• In supergravity, superstrata are horizonless solutions with same charges as the D1-D5-P black hole [Bena, Giusto, Martinec, Russo, Shigemori, Turton, Warner '16]

• Part of the

Fuzzball hypothesis:

Individual black-hole microstates differ from themselves and from the BH solution at the horizon scale.

Superstrata and their limits



• In supergravity, superstrata are horizonless solutions with same charges as the D1-D5-P black hole [Bena, Giusto, Martinec, Russo, Shigemori, Turton, Warner '16]

Part of the

Fuzzball hypothesis:

Individual black-hole microstates differ from themselves and from the BH solution at the horizon scale.

Superstrata and their limits

Drawbacks:

1. $S \sim \sqrt{N_1 N_5} N_P^{1/4} \ll \sqrt{N_1 N_5 N_P}$

[Shigemori '19]

2. Have a non-vanishing angular momentum in \mathbb{R}^4 \Rightarrow could be atypical ↑ are not exactly spherically symmetric

See also [Lin, Maldacena, Rozenberg, Shan '22]

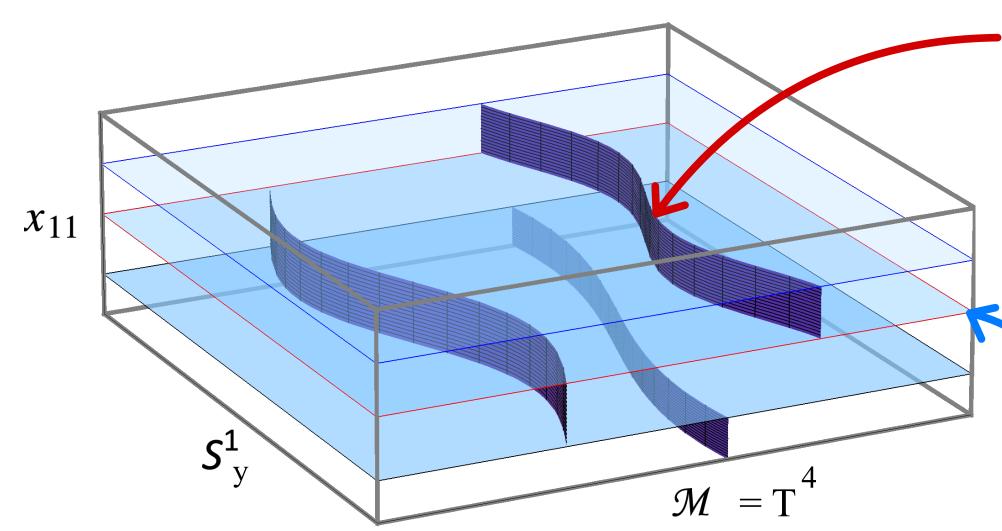




1. Local supersymmetry enhancement and black-hole microstates

2. The new M5-M2-P microstates with 16 local supersymmetries

• For the NS5-F1-P black hole (IIA), we know where the entropy is coming from: Little strings / fractionated (M2) branes



« Dijkgraaf-Verlinde-Verlinde -Maldacena microstates »

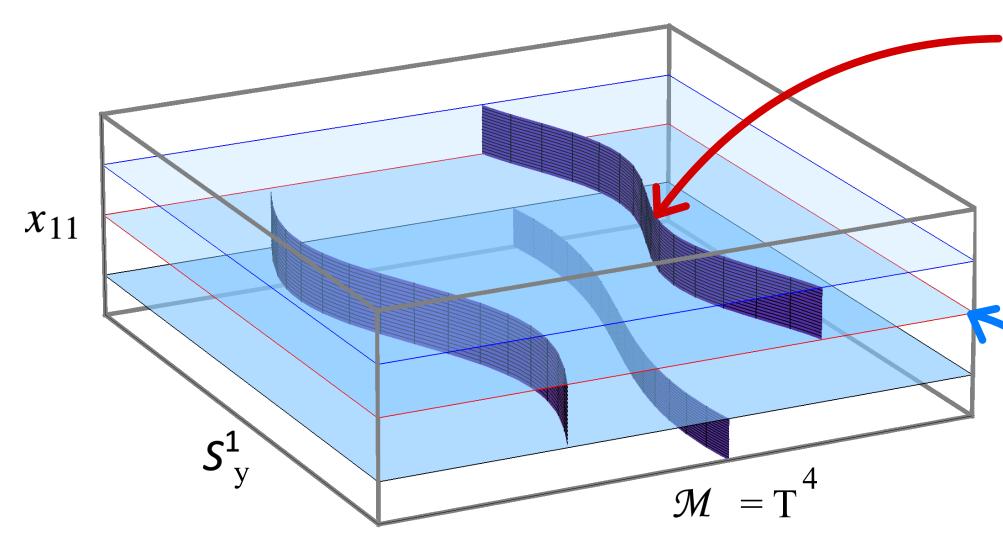
[Dijkgraaf-Verlinde-Verlinde '96], [Maldacena,'96]

See e.g. [Martinec, Massai, Turton '19]

M2

M5

from: Little strings / fractionated (M2) branes



« Dijkgraaf-Verlinde-Verlinde -Maldacena microstates »

[Dijkgraaf-Verlinde-Verlinde '96], [Maldacena,'96]

• For the NS5-F1-P black hole (IIA), we know where the entropy is coming

See e.g. [Martinec, Massai, Turton '19]

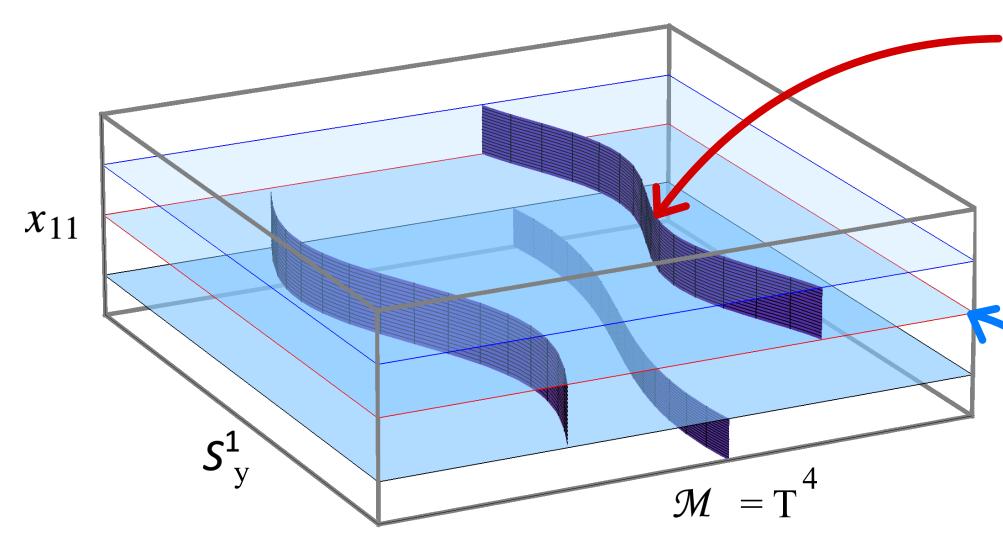
M2

• The momentum is carried by the fractionated M2's through their motion in the T^4

 \rightarrow reproduce entropy.

M5

from: Little strings / fractionated (M2) branes



« Dijkgraaf-Verlinde-Verlinde -Maldacena microstates »

[Dijkgraaf-Verlinde-Verlinde '96], [Maldacena,'96]

• For the NS5-F1-P black hole (IIA), we know where the entropy is coming

See e.g. [Martinec, Massai, Turton '19]

M2

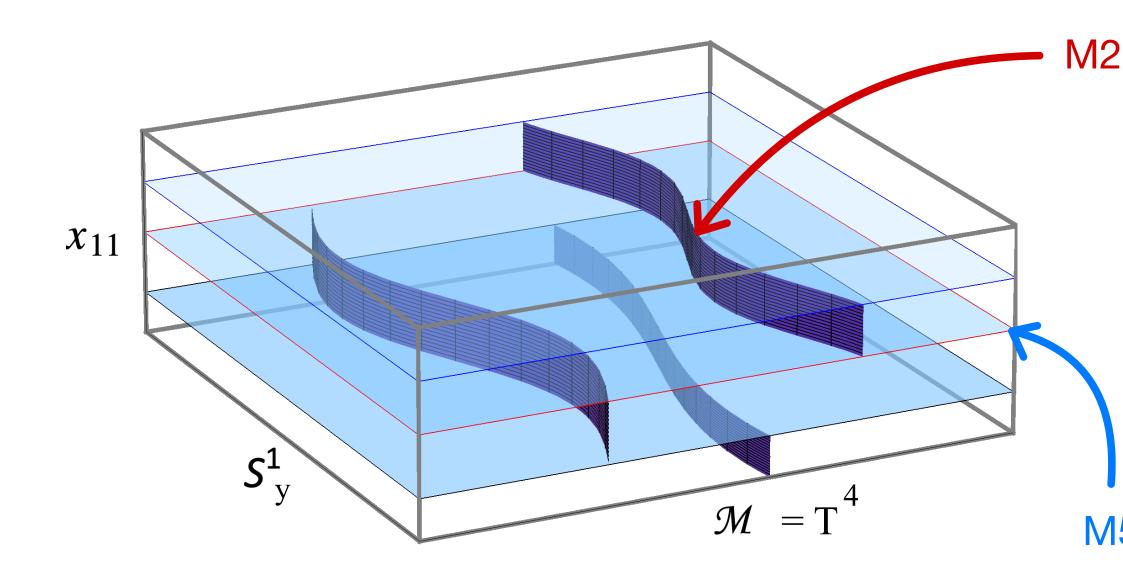
• The momentum is carried by the fractionated M2's through their motion in the T^4

 \rightarrow reproduce entropy.

M5

 $S = 2\pi \sqrt{cN_P/6}$, $c = 6N_1N_5$

• For the NS5-F1-P black hole (IIA), we know where the entropy is coming from: Little strings / fractionated (M2) branes



« Dijkgraaf-Verlinde-Verlinde -Maldacena microstates »

[Dijkgraaf-Verlinde-Verlinde '96], [Maldacena,'96]

See e.g. [Martinec, Massai, Turton '19]

• The momentum is carried by the fractionated M2's through their motion in the T^4

 \rightarrow reproduce entropy.

• The brane system is point-like in M5 the non-compact spatial dimensions

 \rightarrow exact spherical symmetry.

Enhancing the DVVM microstates

Maldacena (DVVM) microstates.

[Bena, Hampton, Houppe, YL, Toulikas '22]

• We enhanced the local supersymmetries of the Dijkgraaf-Verlinde-Verlinde-



Enhancing the DVVM microstates

- Maldacena (DVVM) microstates.
- We found the *supersymmetric projector*
 - preserving the supersymmetries of NS5(y, T^4), F1(y), P(y) (||A)
 - corresponding to an object with 16 local supersymmetries:

$$\Pi_{\text{NS5-F1-P}} = \frac{1}{2} \left[1 + a^2 P_{\text{NS5}(y1234)}^{\text{IIA}} + b^2 P_{\text{F1}(y)} + c^2 P_{\text{P}(y)} + ab \left(P_{\text{D4}(y234)} - P_{\text{D2}(y1)} \right) + bc \left(P_{\text{P}(1)} - P_{\text{F1}(1)} \right) + ca \left(P_{\text{D4}(1234)} - P_{\text{D0}} \right) \right]$$

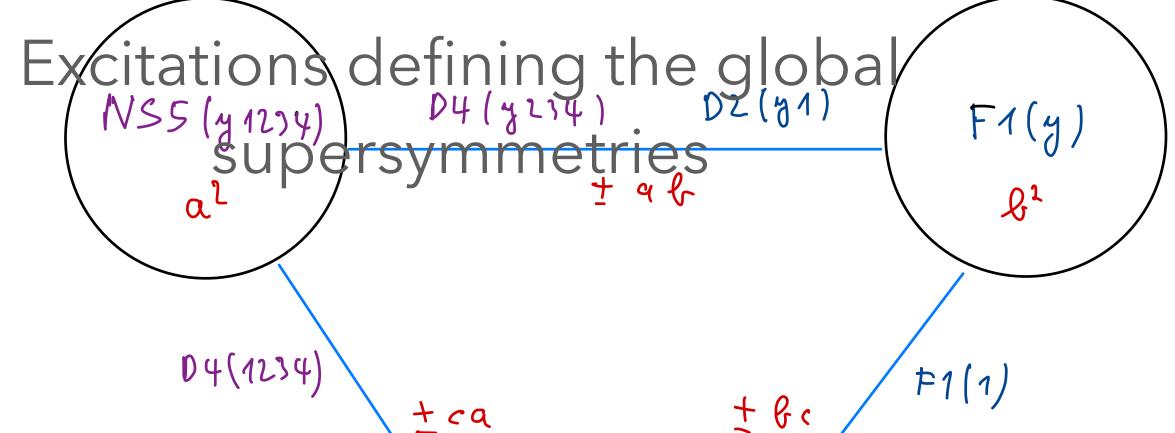
[Bena, Hampton, Houppe, YL, Toulikas '22] • We enhanced the local supersymmetries of the Dijkgraaf-Verlinde-Verlinde-



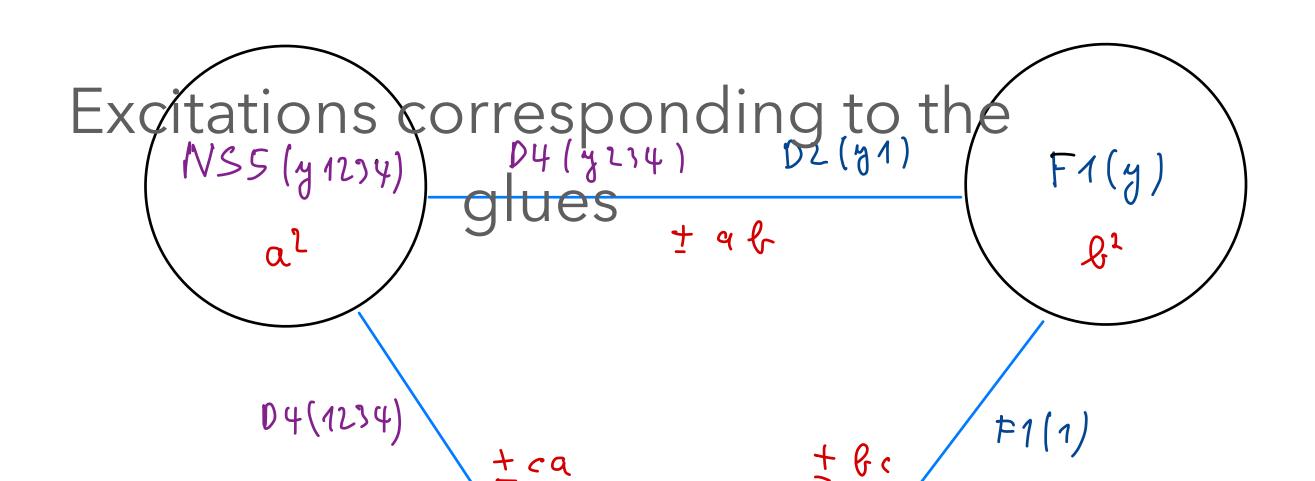
 $\Pi_{\text{NS5-F1-P}} = \frac{1}{2} \left| 1 + a^2 P_{\text{NS5}(y_{1234})}^{\text{IIA}} + b^2 P_{\text{F1}(y)} + c^2 P_{\text{P}(y)} \right|$ 04(1234)+ ca

First look at the projector

 $+ ab \left(P_{\mathrm{D4}(y234)} - P_{\mathrm{D2}(y1)} \right) + bc \left(P_{\mathrm{P}(1)} - P_{\mathrm{F1}(1)} \right) + ca \left(P_{\mathrm{D4}(1234)} - P_{\mathrm{D0}} \right) \right| .$

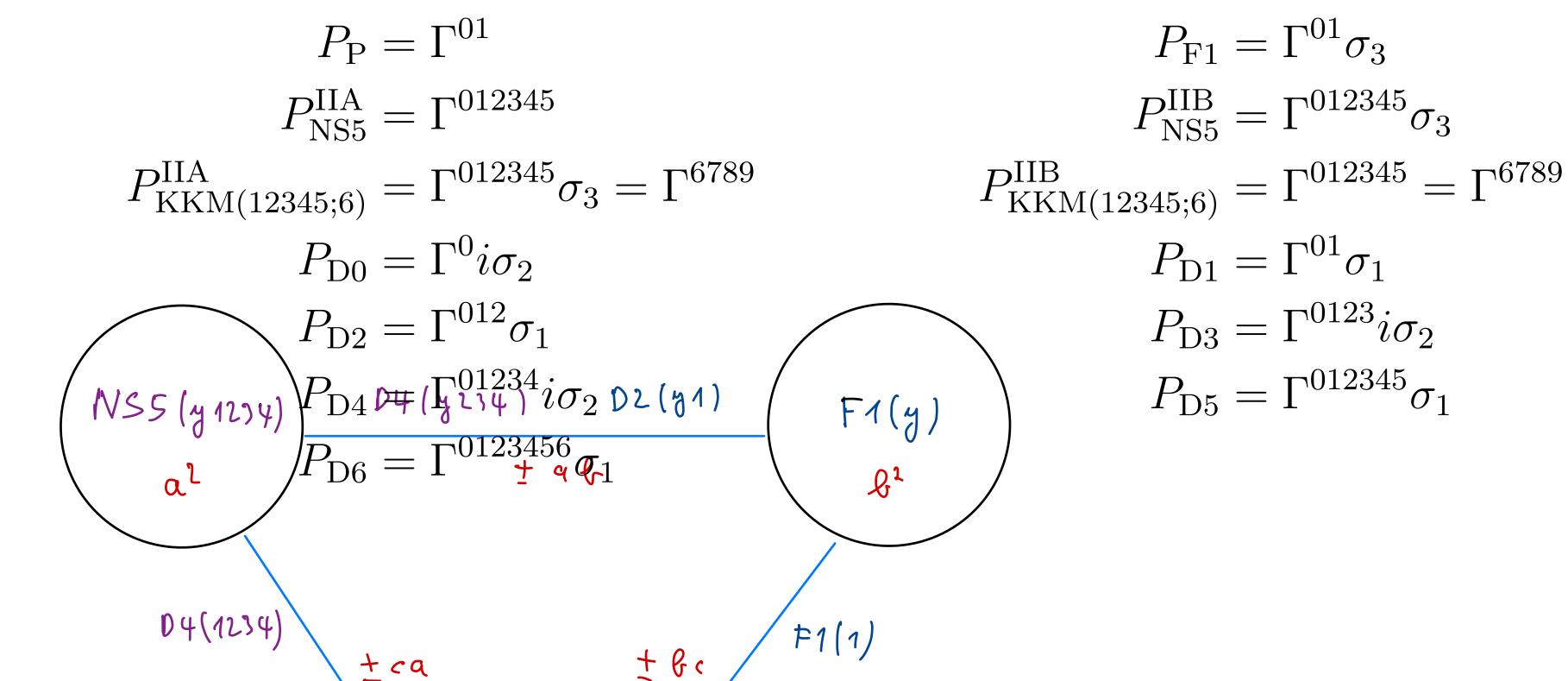


First look at the projector $\Pi_{\text{NS5-F1-P}} = \frac{1}{2} \left| 1 + a^2 P_{\text{NS5}(y_{1234})}^{\text{IIA}} + b^2 P_{\text{F1}(y)} + c^2 P_{\text{P}(y)} \right|$



 $+ ab \left(P_{\mathrm{D4}(y234)} - P_{\mathrm{D2}(y1)} \right) + bc \left(P_{\mathrm{P}(1)} - P_{\mathrm{F1}(1)} \right) + ca \left(P_{\mathrm{D4}(1234)} - P_{\mathrm{D0}} \right) \right| .$

$$\begin{aligned} & \mathsf{First \ look \ at \ the \ projector} \\ \Pi_{\text{NS5-F1-P}} &= \frac{1}{2} \bigg[1 + a^2 P_{\text{NS5}(y1234)}^{\text{IIA}} + b^2 P_{\text{F1}(y)} + c^2 P_{\text{P}(y)} \\ &\quad + ab \left(P_{\text{D4}(y234)} - P_{\text{D2}(y1)} \right) + bc \left(P_{\text{P}(1)} - P_{\text{F1}(1)} \right) + ca \left(P_{\text{D4}(1234)} - P_{\text{D0}} \right) \bigg] \end{aligned}$$



BPS condition:

$$a^2 + b^2 + c^2 = 1$$

Glueing NS5 and F1 $\Pi_{\text{NS5-F1-P}} = \frac{1}{2} \left| 1 + a^2 P_{\text{NS5}(y1234)}^{\text{IIA}} + b^2 P_{\text{F1}(y)} + c^2 P_{\text{P}(y)} \right|$

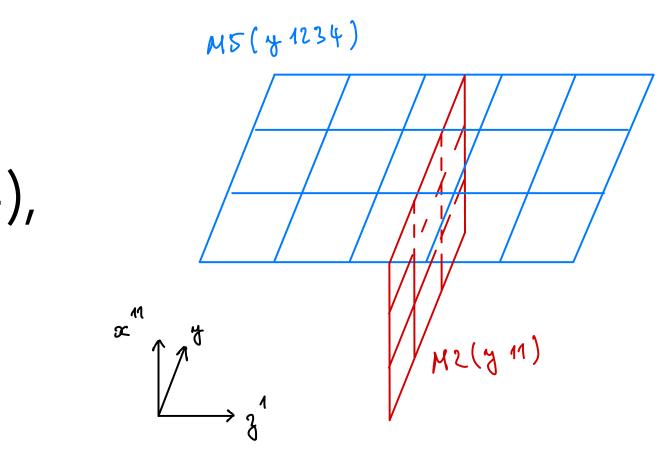
- Put c = 0
- NS5(y, T^4), F1(y) \longrightarrow local D4(y234), D2(y1)

 $+ ab \left(P_{\mathrm{D4}(y234)} - P_{\mathrm{D2}(y1)} \right) + bc \left(P_{\mathrm{P}(1)} - P_{\mathrm{F1}(1)} \right) + ca \left(P_{\mathrm{D4}(1234)} - P_{\mathrm{D0}} \right) \right| \,.$

Glueing NS5 and F1 $\Pi_{\text{NS5-F1-P}} = \frac{1}{2} \left| 1 + a^2 P_{\text{NS5}(y1234)}^{\text{IIA}} + b^2 P_{\text{F1}(y)} + c^2 P_{\text{P}(y)} \right|$

- Put c = 0
- NS5(y, T^4), F1(y) \longrightarrow local D4(y234), D2(y1)

 $+ ab \left(P_{\mathrm{D4}(y234)} - P_{\mathrm{D2}(y1)} \right) + bc \left(P_{\mathrm{P}(1)} - P_{\mathrm{F1}(1)} \right) + ca \left(P_{\mathrm{D4}(1234)} - P_{\mathrm{D0}} \right) \right| .$



$$\begin{aligned} &\Pi_{\text{NS5-F1-P}} = \frac{1}{2} \bigg[1 + a^2 P_{\text{NS5}(y1234)}^{\text{IIA}} + b^2 P_{\text{F1}} \\ &+ ab \left(P_{\text{D4}(y234)} - P_{\text{D2}(y24)} \right) \bigg] \end{aligned}$$

• Put
$$c = 0$$

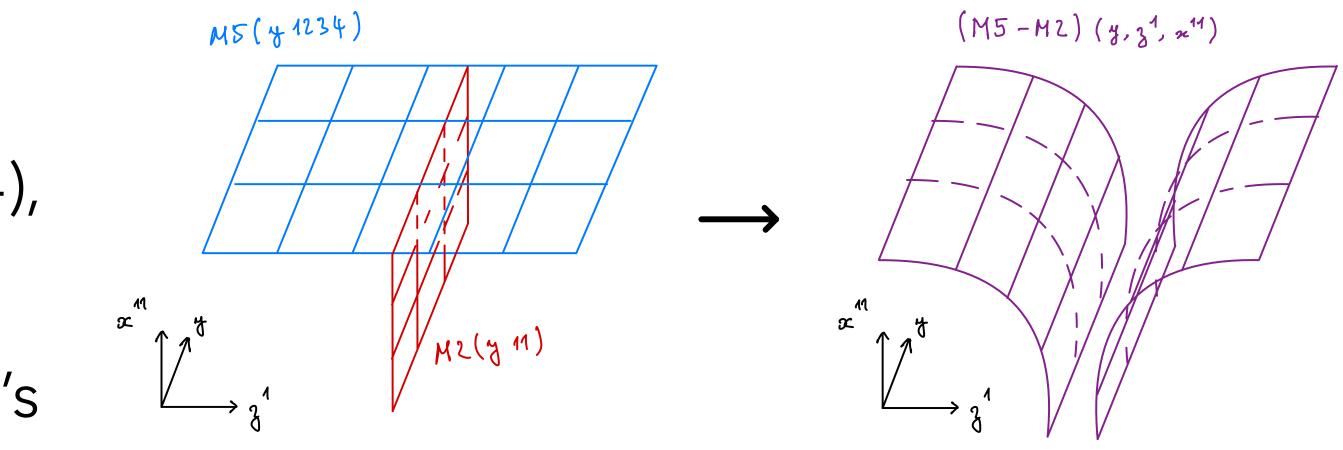
- NS5(y, T^4), F1(y) \longrightarrow local D4(y234), D2(y1)
- ROUGH ANGLES between M5's and M2's become *smooth*:

 \rightarrow new brane system looks like a *furrow* along *y*.

NS5 and F1

 $\mathbf{L}(y) + c^2 P_{\mathbf{P}(y)}$

 $(y_{1}) + bc \left(P_{P(1)} - P_{F1(1)} \right) + ca \left(P_{D4(1234)} - P_{D0} \right) \right|.$



$$\begin{aligned} &\Pi_{\text{NS5-F1-P}} = \frac{1}{2} \bigg[1 + a^2 P_{\text{NS5}(y1234)}^{\text{IIA}} + b^2 P_{\text{F1}} \\ &+ ab \left(P_{\text{D4}(y234)} - P_{\text{D2}(y24)} \right) \bigg] \end{aligned}$$

• Put
$$c = 0$$

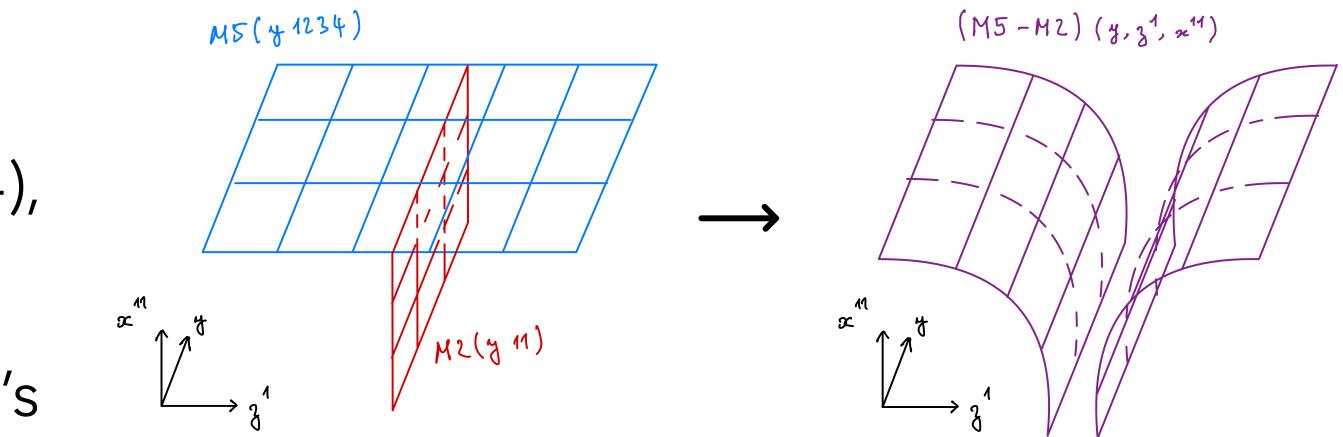
- NS5(y, T^4), F1(y) \longrightarrow local D4(y234), D2(y1)
- ROUGH ANGLES between M5's and M2's become *smooth*:

 \rightarrow new brane system looks like a *furrow* along *y*.

NS5 and F1

 $l(y) + c^2 P_{\mathrm{P}(y)}$

 $(P_{P(1)}) + bc \left(P_{P(1)} - P_{F1(1)} \right) + ca \left(P_{D4(1234)} - P_{D0} \right) \right|.$



↑ This M5-M2 furrow is dual to aD4-F1 Callan-Maldacena spike

$$\begin{aligned} & \Pi_{\text{NS5-F1-P}} = \frac{1}{2} \bigg[1 + a^2 P_{\text{NS5}(y1234)}^{\text{IIA}} + b^2 P_{\text{F1}} \\ & + ab \left(P_{\text{D4}(y234)} - P_{\text{D2}(y234)} + b^2 P_{\text{D2}(y234)} \right) \bigg] \end{aligned}$$

 The furrow interpolates between M5 and M2:

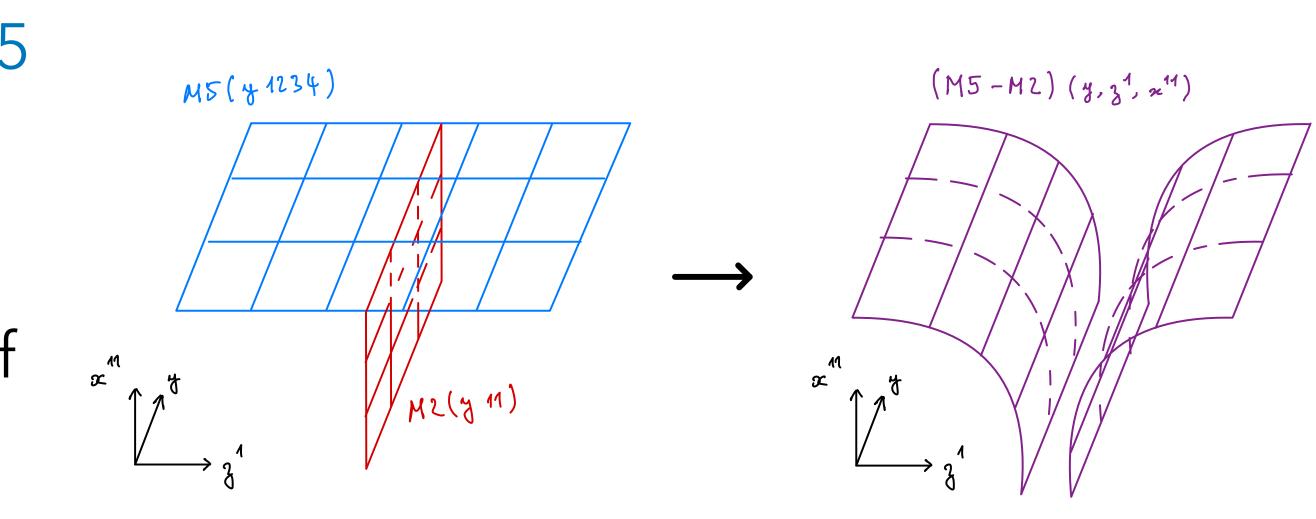
$$a = \cos \beta$$
, $b = \sin \beta$

⇒ The orientation of a local piece of the furrow determines the ratio between M5 and M2 charges.

NS5 and F1

 $\mathbf{l}(y) + c^2 P_{\mathbf{P}(y)}$

 $(D_{2(y1)}) + bc \left(P_{P(1)} - P_{F1(1)} \right) + ca \left(P_{D4(1234)} - P_{D0} \right) \right|.$



Transition of a M5-M2 black-hole microstate

M5

M5

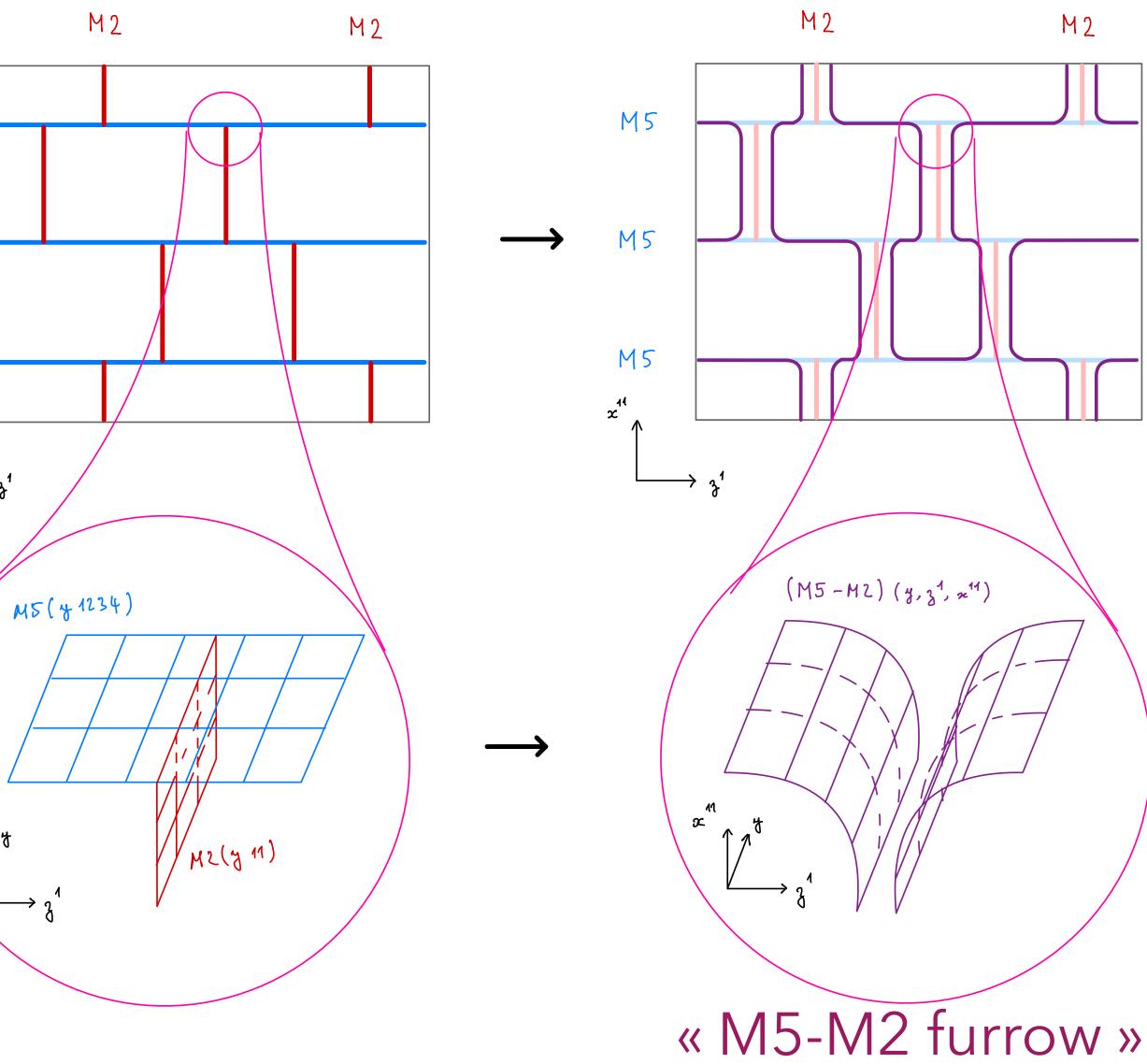
M5

x

31

• Local transition \Rightarrow a M5-M2 black-hole microstate will transition into a « labyrinth/maze »

 \rightarrow « super-maze »





$$\begin{aligned} & \Pi_{\text{NS5-F1-P}} = \frac{1}{2} \bigg[1 + a^2 P_{\text{NS5}(y1234)}^{\text{IIA}} + b^2 P_{\text{F1}} \\ & + ab \left(P_{\text{D4}(y234)} - P_{\text{D2}(y1234)} + b^2 P_{\text{D2}(y1234)} \right) \end{vmatrix}$$

 The M5-M2 furrow carries momentum through *rípples* modulated orthogonally to its surface

$$a = \cos \alpha \cos \beta$$

$$b = \cos \alpha \sin \beta \left(\sum_{N \le 5 (y + 12) y} \right)$$

$$c = \sin \alpha$$

$$a^{1}$$

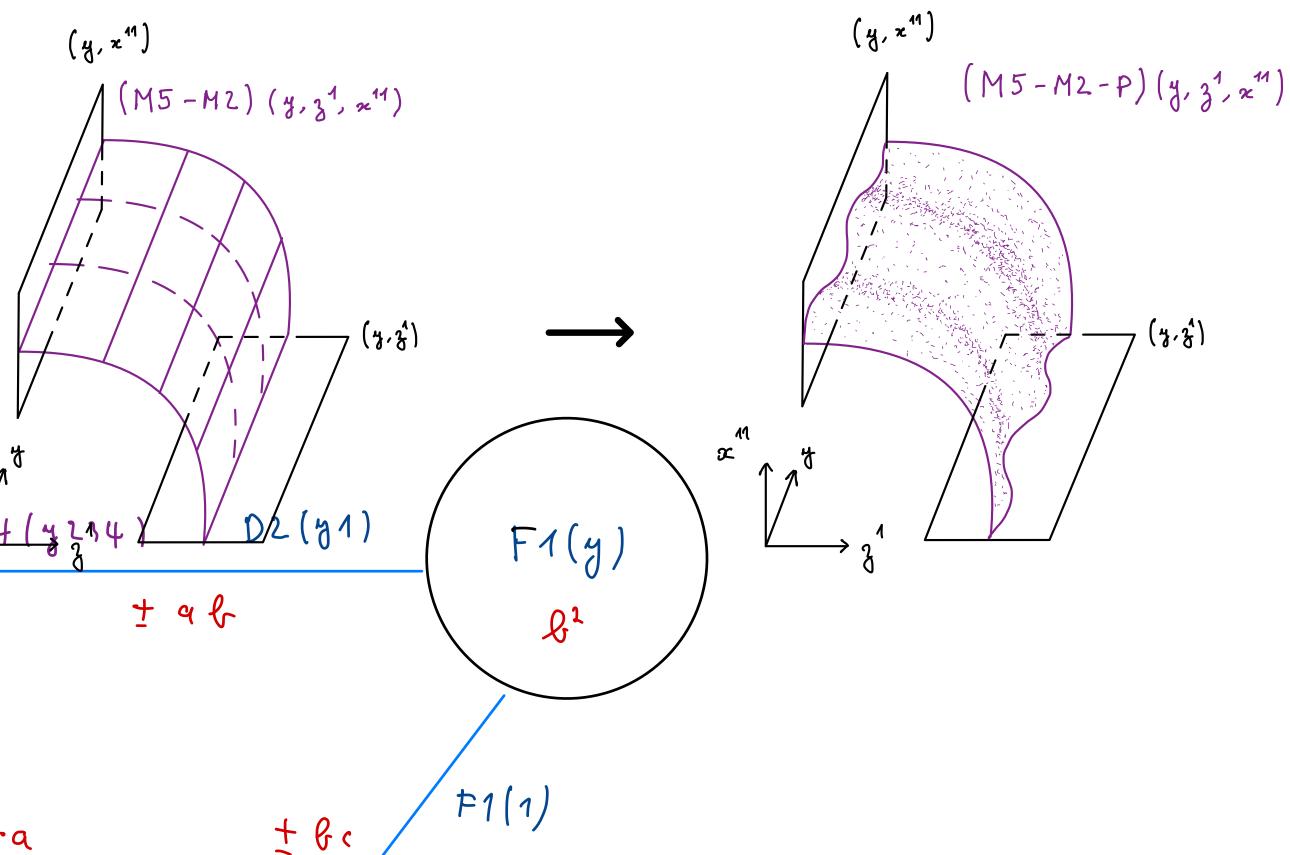
04(1234)

+ ca

JS5, F1 and P

 $P_{\mathrm{P}(y)} + c^2 P_{\mathrm{P}(y)}$

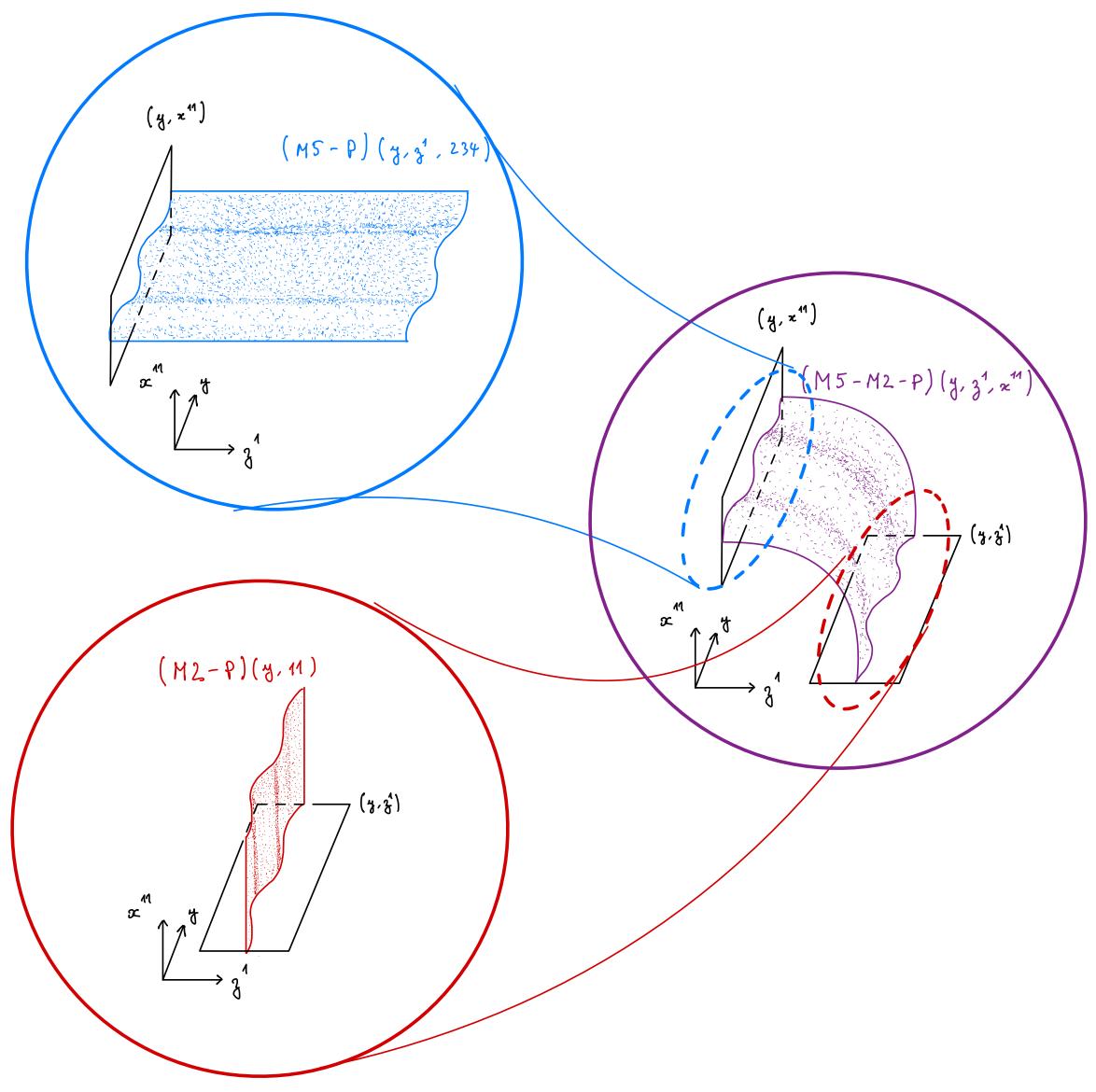
 $_{2(y1)}) + bc\left(P_{P(1)} - P_{F1(1)}\right) + ca\left(P_{D4(1234)} - P_{D0}\right)\right).$



Glueing NS5, F1 and P

- The M5-M2 furrow carries momentum through *rípples* modulated orthogonally to its surface
 - $a = \cos \alpha \cos \beta$ $b = \cos \alpha \sin \beta$ $c = \sin \alpha$
- β controls the bending

 angle of the furrow; α
 controls the angle of ripples
 orthogonal to the furrow.



Consequence on a M5-M2-P microstate

M5

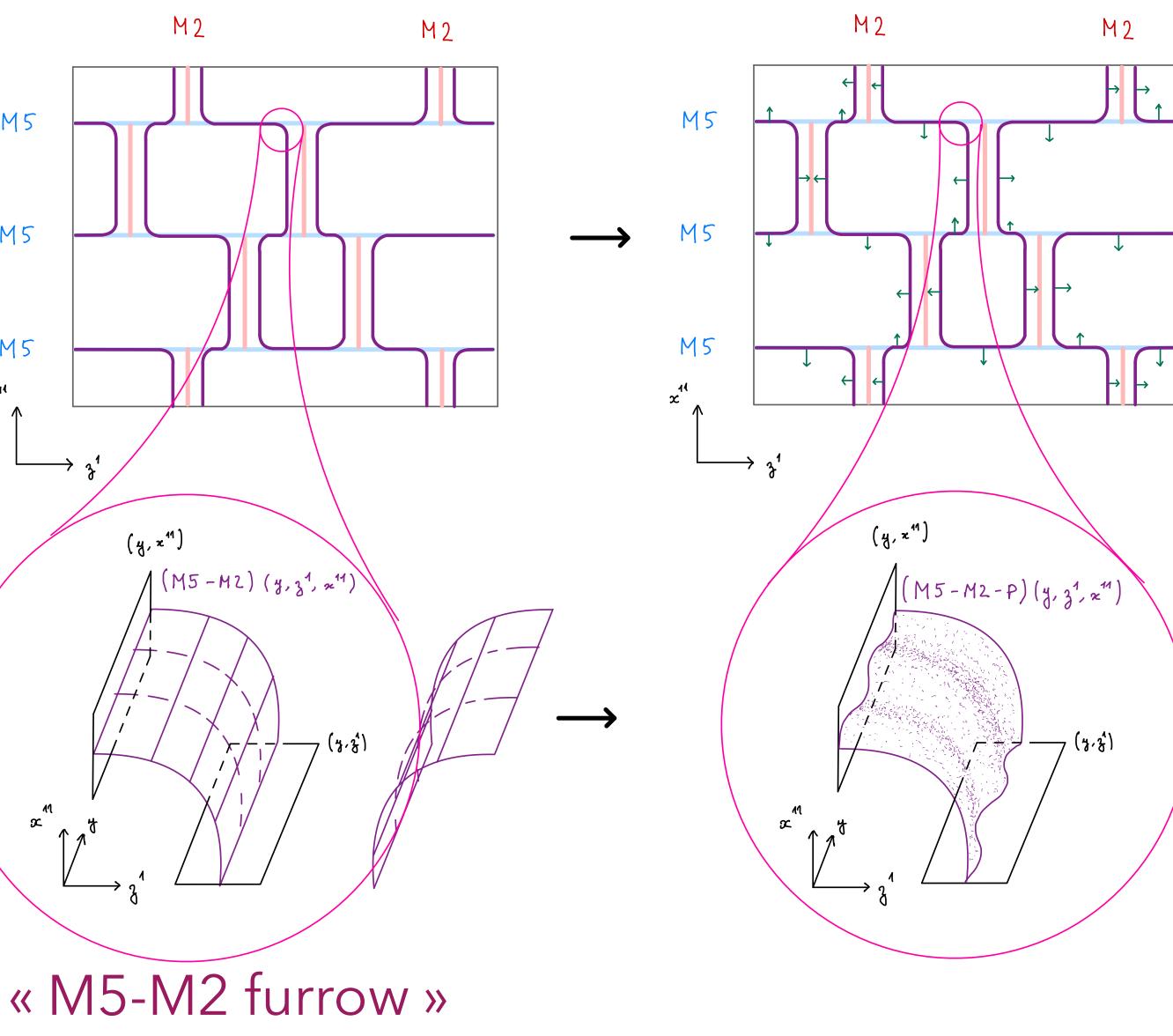
M5

M5

11 72

11 X

- The *ripples* of the furrow correspond to shape modes of the M5-M2 labyrinth
- The shape modes are the way 16-susy microstates carry momentum.







Consequence on a M5-M2-P microstate

M5

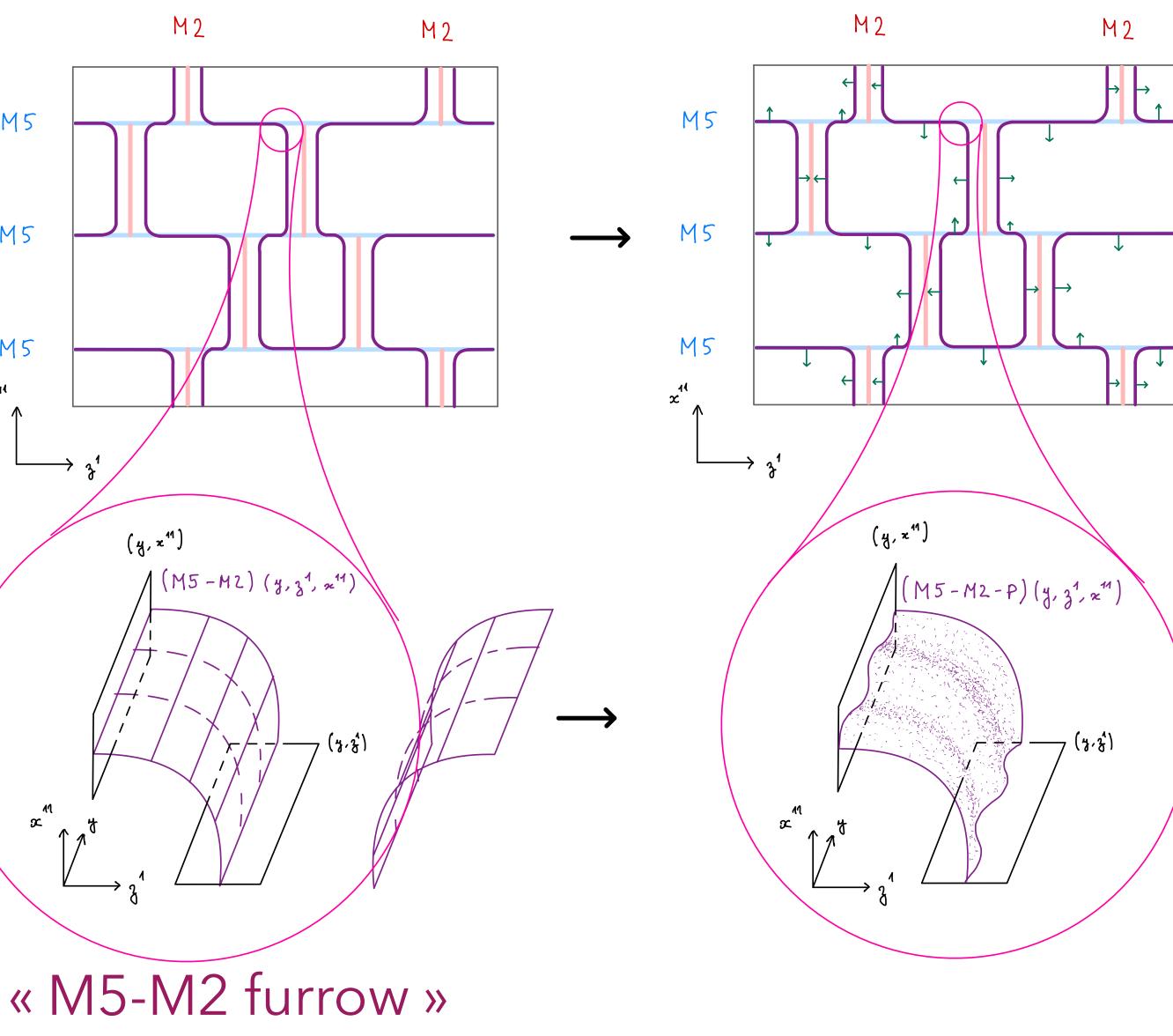
M5

M5

11 92

11 X

• The *ripples* of the furrow correspond to shape modes of the M5-M2 maze







Consequence on a M5-M2-P microstate

M5

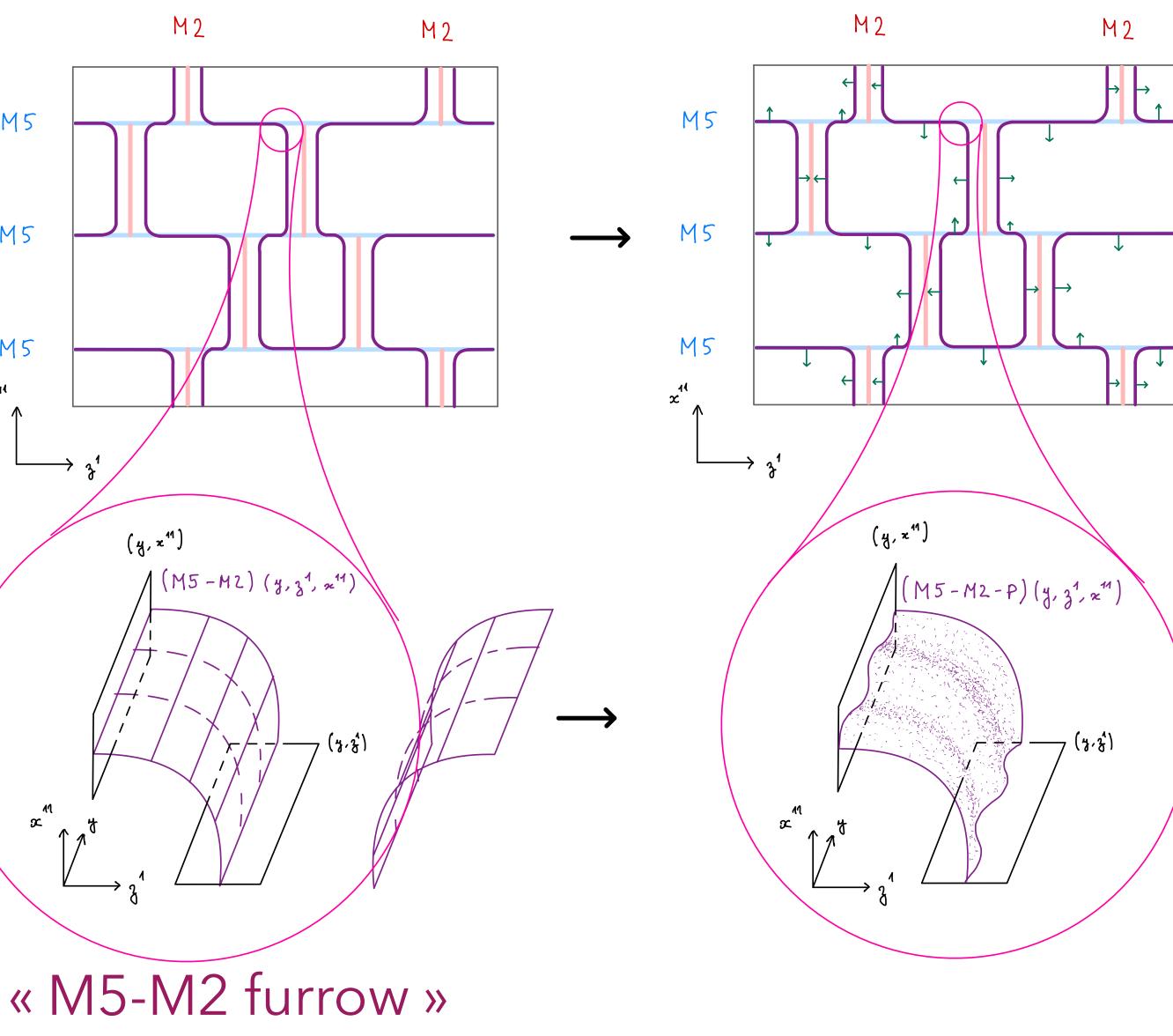
M5

M5

41 92 1

- The *ripples* of the furrow correspond to shape modes of the M5-M2 maze
- The shape modes are the way 16-local-susy microstates carry momentum.

 \Rightarrow The microstates are ensured to have *exact* spherical symmetry.







Horizonless geometries in supergravity?

• Can they described in supergravity?

י?

Horizonless geometries in supergravity?

- Can they described in supergravity?
- In Type IIA, one can only separate the NS5 branes in the non-compact dilaton gets a large value, so supergravity breaks down.

spatial dimensions. There exists a region close to the branes where the

e.g. [Martinec, Massai, Turton '22]

Horizonless geometries in supergravity?

- Can they described in supergravity?
- In Type IIA, one can only separate the NS5 branes in the non-compact dilaton gets a large value, so supergravity breaks down.
- the branes as well.

spatial dimensions. There exists a region close to the branes where the

e.g. [Martinec, Massai, Turton '22]

• But in the M2-M5-P frame, the basic ingredient of the super-maze is a M5 brane with M2 flux on it. The supergravity description of it is valid close to



• Global charges and supersymmetries control the near-horizon geometry.

Local supersymmetries are a means to get information on the microstates.

- Global charges and supersymmetries control the near-horizon geometry.
- Local supersymmetries are a means to get information on the microstates.
- 1/8-BPS systems (3-charge BHs) have a large moduli space of solutions that have more supersymmetries locally

theory resolve the singularity or the horizon.

↑ This is crucial in order to understand whether microstates in string

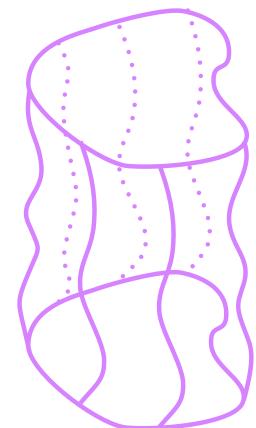
- Global charges and supersymmetries control the near-horizon geometry.
- Local supersymmetries are a means to get information on the microstates.
- 1/8-BPS systems (3-charge BHs) have a large moduli space of solutions that have more supersymmetries locally

theory resolve the singularity or the horizon.

• The microstate geometries programme used to replace D1-D5-P horizons with brane systems that extend in \mathbb{R}^4

↑ But this approach seems to have limits: entropy, typicality...

↑ This is crucial in order to understand whether microstates in string



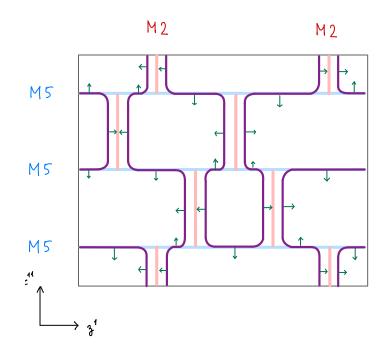


• New approach: microstates can carry momentum by having motion in the *internal dimensions* \Rightarrow exactly spherical symmetry

Conclusion

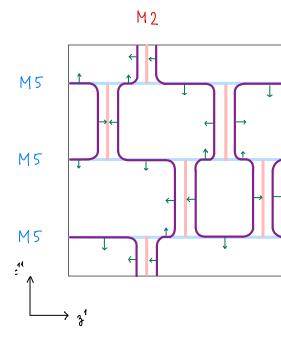


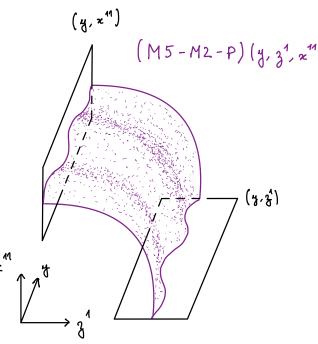
- New approach: microstates can carry momentum by having motion in the internal dimensions \Rightarrow exactly spherical symmetry
- The DVVM microstates account for the black-hole entropy... ... and we have identified what they become when the branes start interacting.





- New approach: microstates can carry momentum by having motion in the internal dimensions \Rightarrow exactly spherical symmetry
- The DVVM microstates account for the black-hole entropy... ... and we have identified what they become when the branes start interacting.
- These « super-mazes » have 16 local susys, just like the superstrata, but without having their drawbacks.





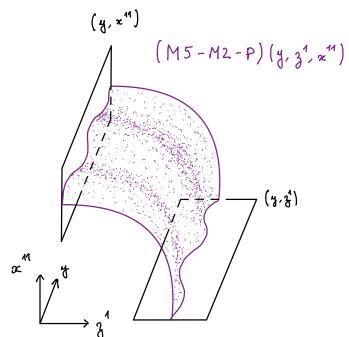




7 (y,ż)

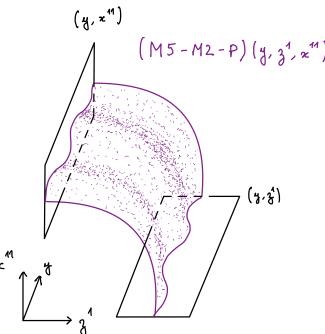


 16 local susys is a smoking gun for horizonless microstate solutions



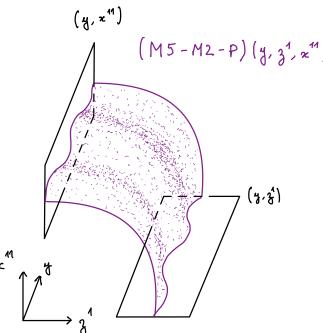


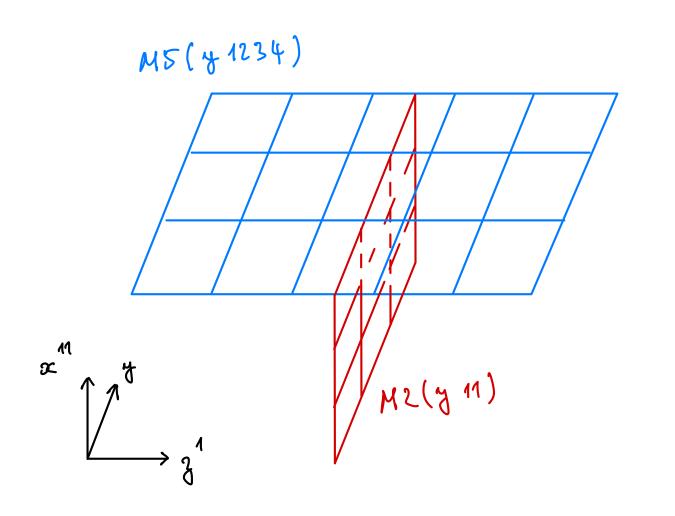
- 16 local susys is a smoking gun for horizonless microstate solutions
 - \Rightarrow seems to support Fuzzball hypothesis for M2-M5-P blackhole microstates
 - ↑ Construct the fully backreacted supergravity solutions
 - **†** Apply geometric quantization to them.

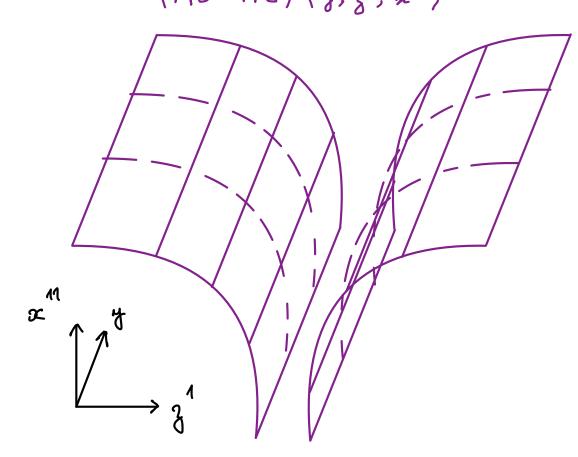




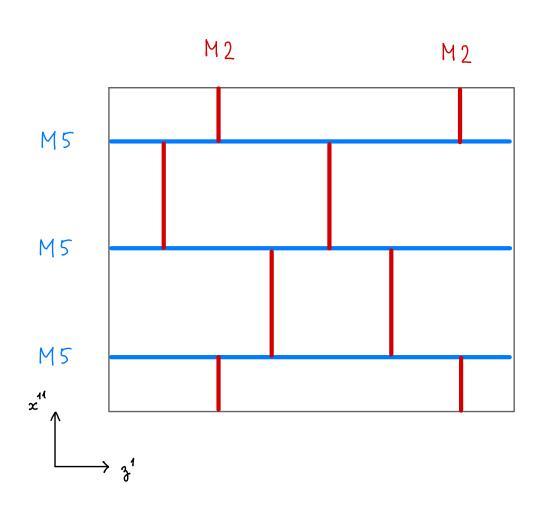
- 16 local susys is a smoking gun for horizonless microstate solutions
 - \Rightarrow seems to support Fuzzball hypothesis for M2-M5-P blackhole microstates
 - ↑ Construct the fully backreacted supergravity solutions
 - ↑ Apply geometric quantization to them.
- End goal:
 - «Where » is the information about the black-hole microstate?

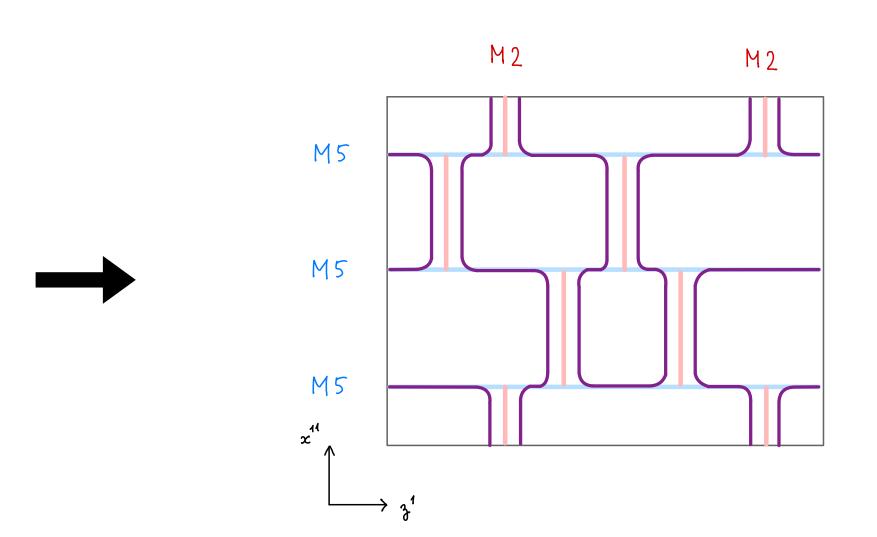


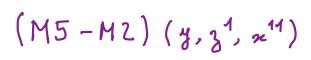


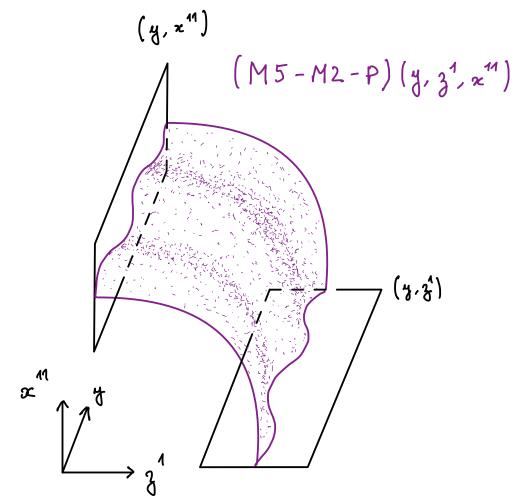


Thank you



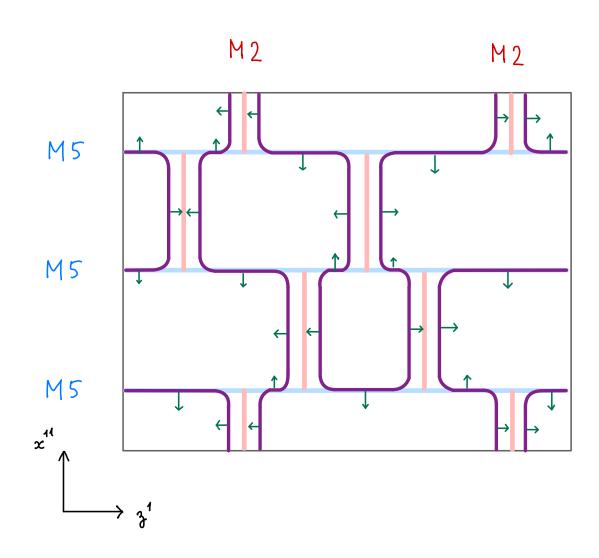






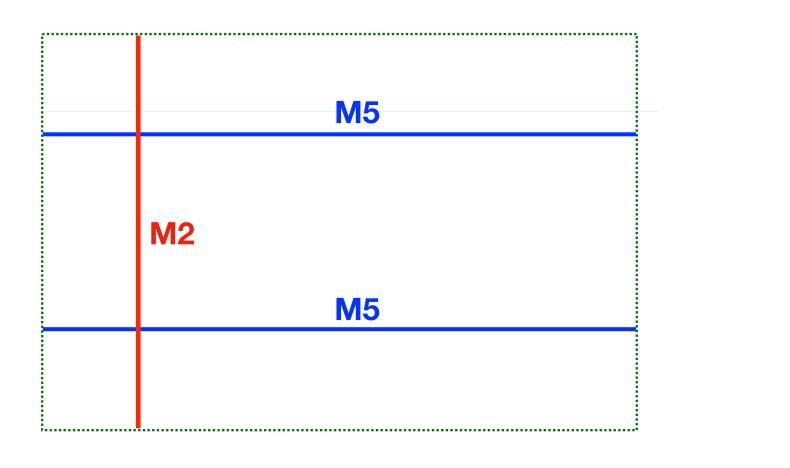
for your

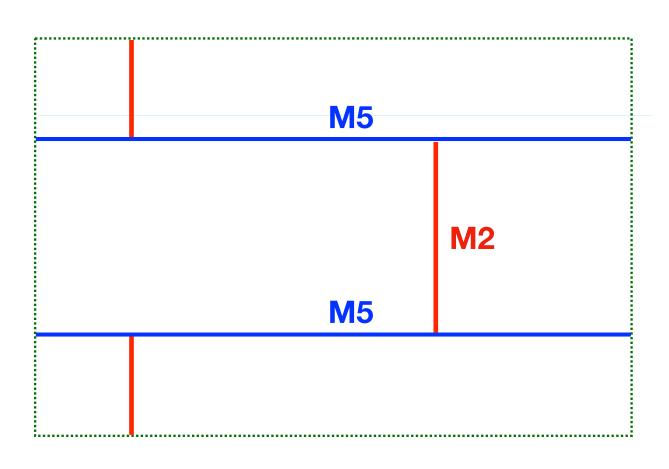


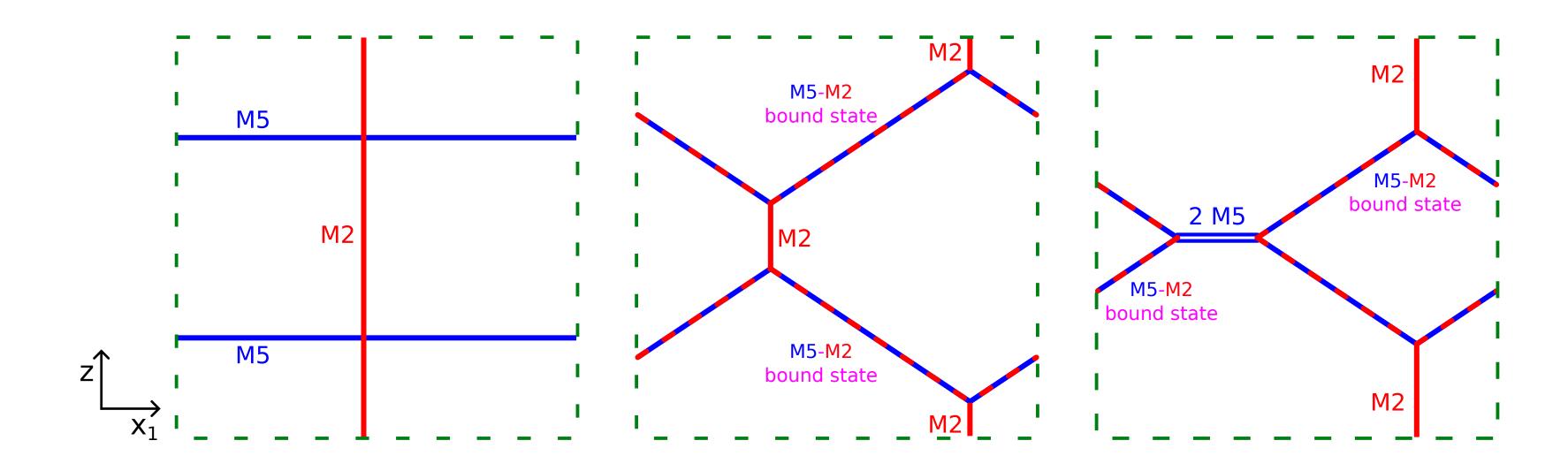












Back-up slide

