

Dark Matter p-wave Annihilation Flux Enhancement Near the Supermassive Black Hole at the Center of M87*

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Outline

- Dark Matter Profile
- Particle Model
- Results

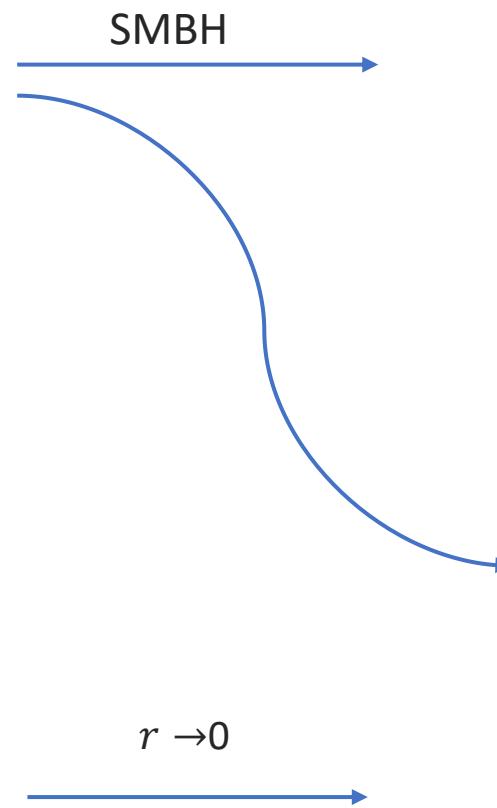
Dark Matter Profile

$$\rho_{\text{DM}}(r) = \begin{cases} 0 & r < r_m \\ \rho_{\text{sp}}(r) & r_m \leq r < r_{\text{sp}} \\ \rho_h(r) & r_{\text{sp}} \leq r \leq R_{200} \end{cases}$$

$$\rho_h(r) = \begin{cases} \rho_{\text{Burkert}}(r) & 0 \leq r < 200 \\ \rho_{\text{NFW}}(r) & 200 \leq r \leq R_{200} \end{cases}$$

$$\rho_{\text{NFW}} = \frac{\rho_s}{\frac{r}{r_s} \left(1 + \frac{r}{r_s}\right)^2}$$

$$\rho_{\text{Burkert}} = \frac{\rho_0}{\left(1 + \frac{r}{r_0}\right) \left(1 + \frac{r}{r_0}\right)^2}$$



$$\rho_{\text{sp}} \propto r^{-\gamma_{\text{sp}}}$$

$$\gamma_{\text{sp}} = (3 + a)/4$$

$$\sigma \propto v^{-a}$$

$$v_h(r_{\text{spike}})^2 = GM_\bullet/r_{\text{spike}}$$

$$\rho(r) \propto r^0$$

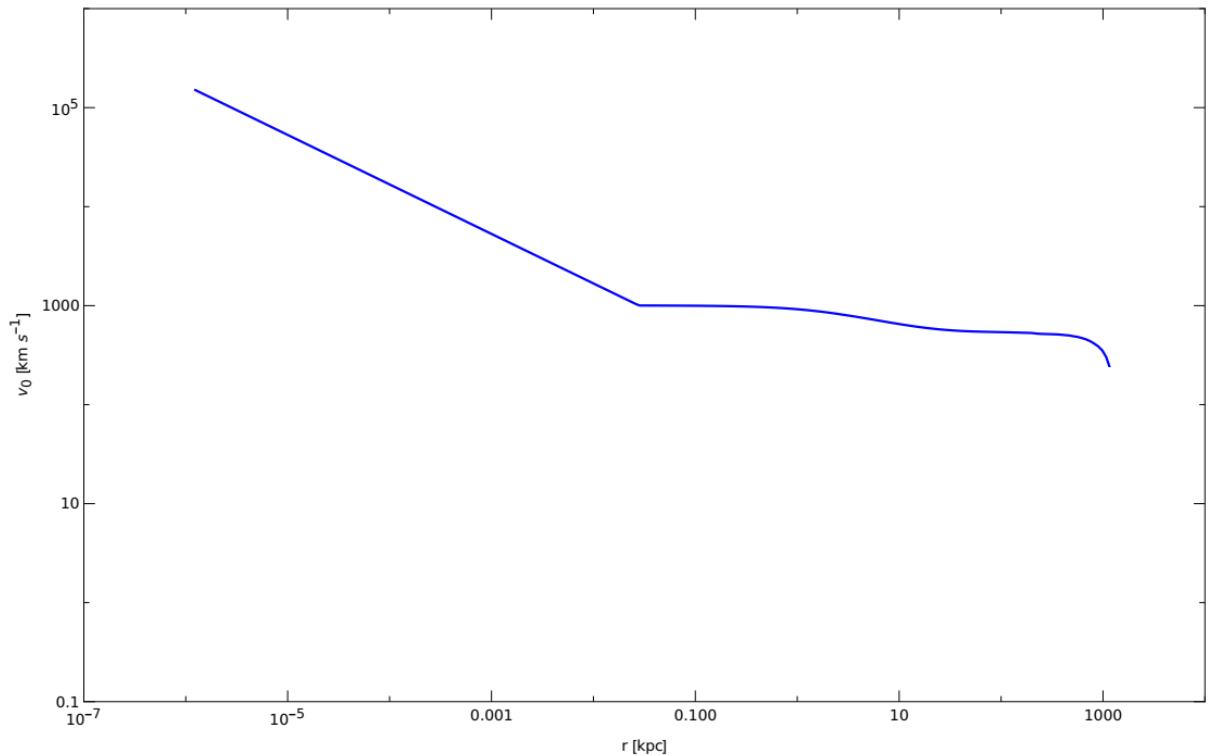
Velocity Dispersion

$$v_0(r) = \begin{cases} v_h(r_{\text{sp}}) \left(\frac{r_{\text{sp}}}{r}\right)^{\frac{1}{2}} & r_m \leq r < r_{\text{sp}} \\ v_h(r) & r_{\text{sp}} \leq r \leq R_{200} \end{cases}$$

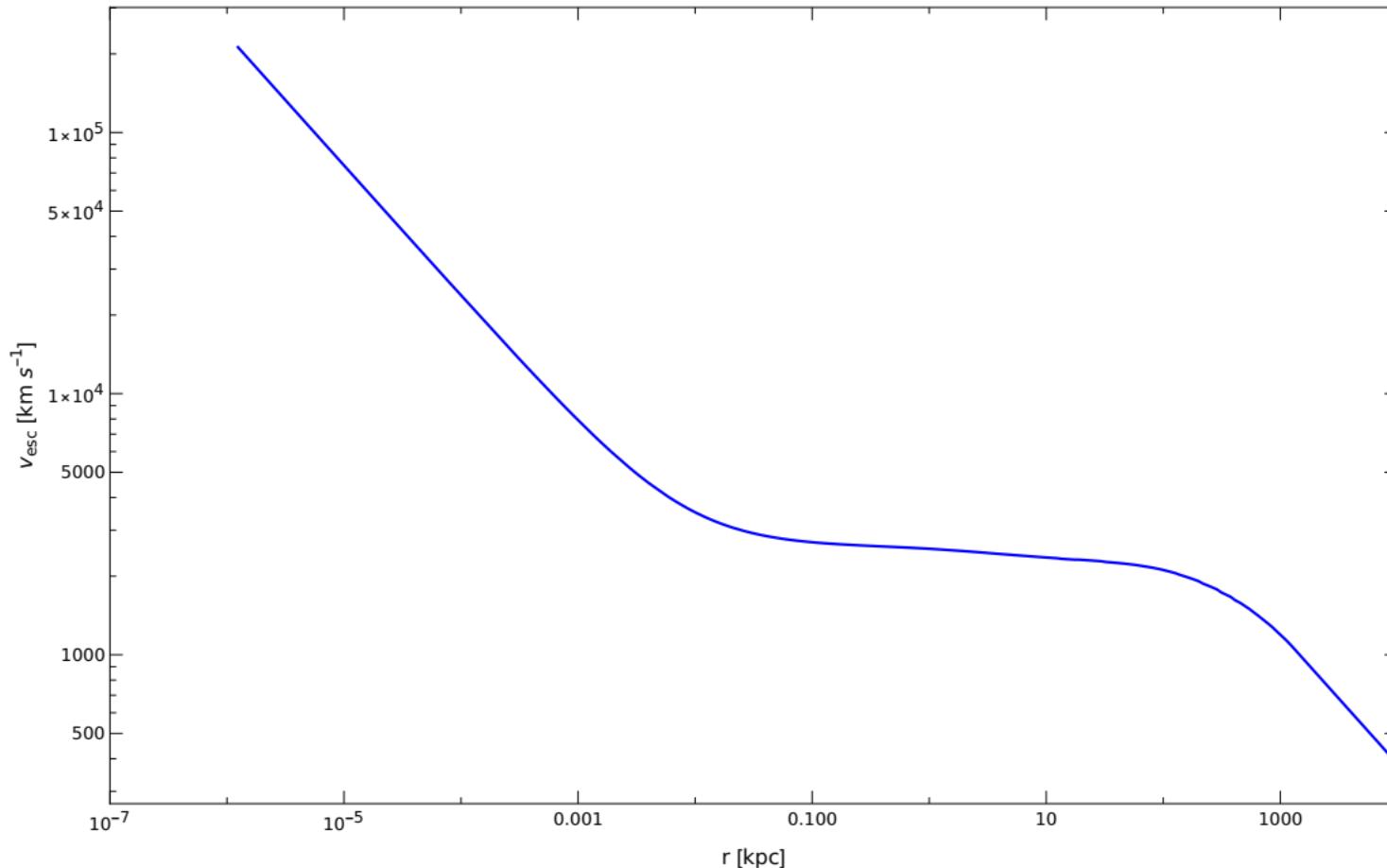
Jeans equation

$$v_h^2(r) = \frac{1}{\rho_h(r)} \int_r^\infty dr' \rho_h(r') \frac{d\Phi}{dr'}$$

SMBH $\longrightarrow v_0 \propto r^{-1/2}$



Escape Velocity



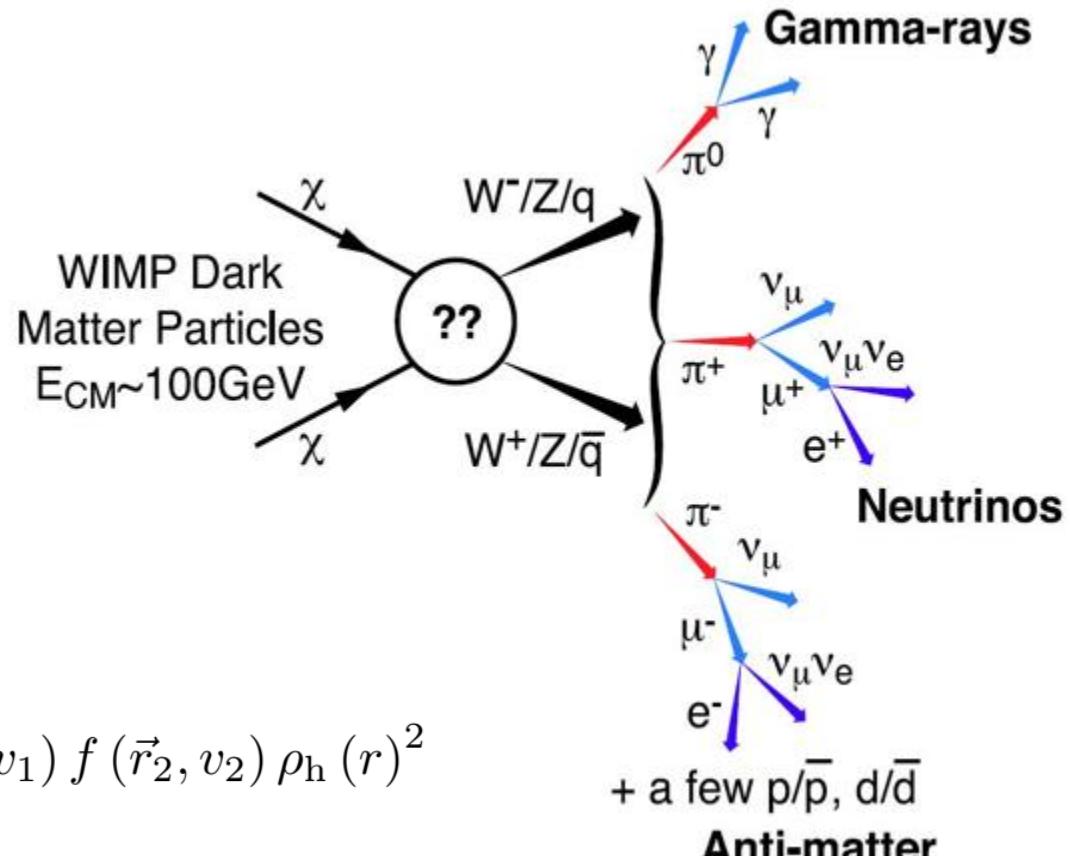
$$\Phi(r) = -4\pi G \left[\frac{1}{r} \int_0^r dr' r'^2 (\rho_{\text{DM}}(r') + \rho_{\star}(r')) + \int_r^R dr' r' (\rho_{\text{DM}}(r') + \rho_{\star}(r')) \right] - \frac{GM_{\bullet}}{r}$$

Annihilation Signals

J-Factor and Q-Factor

$$J(\Omega) = \int d\ell \int d^3v_1 \int d^3v_2 (\sigma v) f(\vec{r}_1, v_1) f(\vec{r}_2, v_2) \rho_h(r)^2$$

$$Q(\Omega) = \frac{1}{D^2} \int_{r_m}^{r_{\text{spike}}} dr r^2 \rho_{\text{spike}}(r)^2 \int d^3v_1 \int d^3v_2 (\sigma v) f(\vec{r}_1, v_1) f(\vec{r}_2, v_2)$$



Particle Model

Particle Model (SIDM)

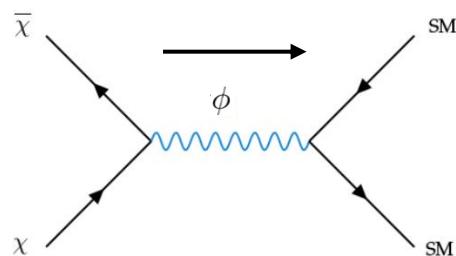
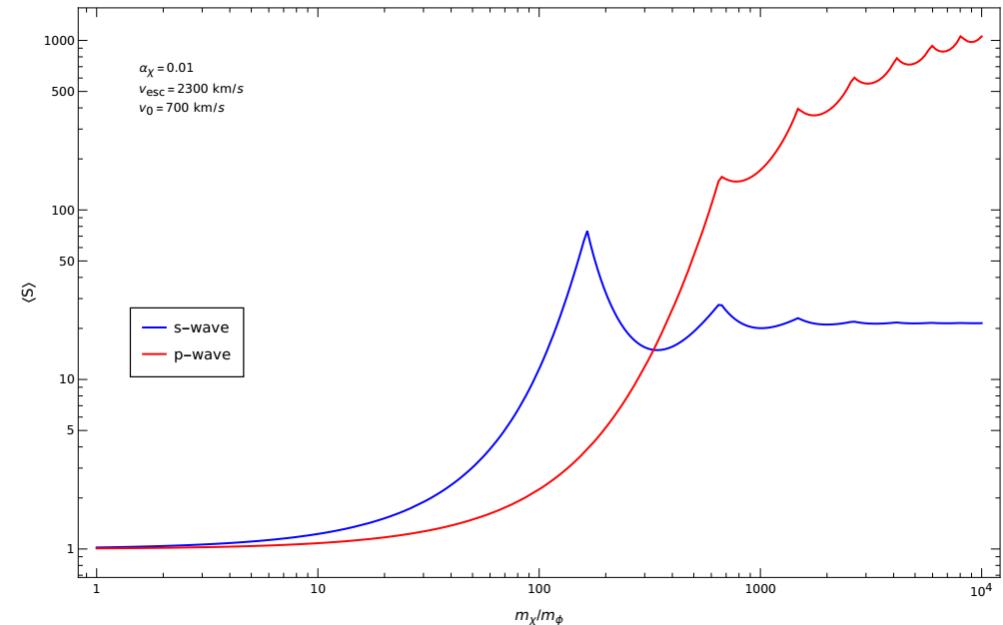
$$\mathcal{L}_{int} = \begin{cases} g_\chi \bar{\chi} \gamma^\mu \chi \phi_\mu & \text{vector mediator} \\ g_\chi \bar{\chi} \chi \phi & \text{scalar mediator} \end{cases}$$

$$(\sigma v)^{\text{tree}} = \begin{cases} \frac{\pi \alpha_\chi^2}{m_\chi^2} \sqrt{1 - \frac{m_\phi^2}{m_\chi^2}} & \text{vector mediator (s-wave)} \\ \frac{3}{4} \frac{\pi \alpha_\chi^2}{m_\chi^2} v^2 \sqrt{1 - \frac{m_\phi^2}{m_\chi^2}} & \text{scalar mediator (p-wave)} \end{cases}$$

$$S_n(v) = \begin{cases} \frac{\pi}{a} \frac{\sinh(2\pi ab)}{\cosh(2\pi ab) - \cos(2\pi \sqrt{b-(ab)^2})} & n=0 \text{ (s-wave)} \\ \frac{\pi}{a} \frac{\sinh(2\pi ab)}{\cosh(2\pi ab) - \cos(2\pi \sqrt{b-(ab)^2})} \frac{(b-1)^2 + 4(ab)^2}{1+4(ab)^2} & n=2 \text{ (p-wave)} \end{cases}$$

$$a = v/2\alpha_\chi \quad b = 6\alpha_\chi m_\chi / \pi^2 m_\phi$$

time



Dark Matter Fine Structure

Boltzmann Equation

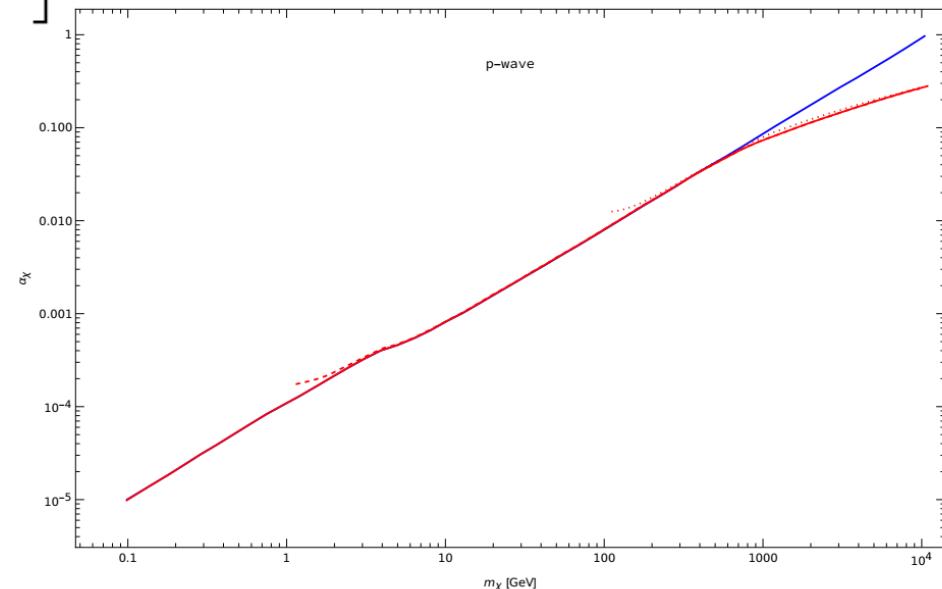
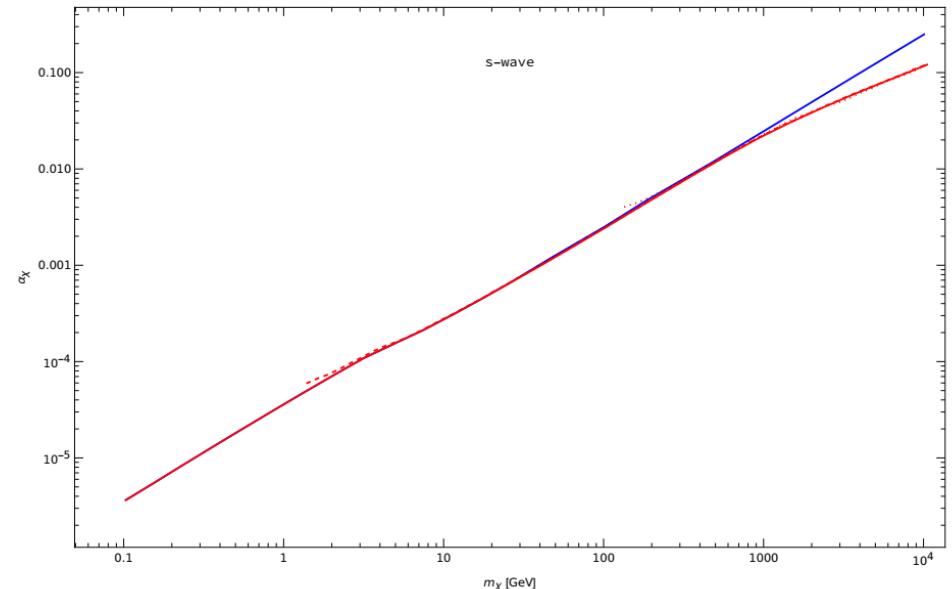
$$\frac{63\sqrt{5}x_f^{-1/2}e^{-x_f}g}{32\pi^3}\frac{g_*^{1/2}(x_f)}{g_{*s}(x_f)}m_\chi m_{\text{pl}}\langle\sigma v\rangle = 1$$

$$\Omega_\chi h^2 = 8.77 \times 10^{-11} \text{ GeV}^{-2} \left[\int_{x_f}^{x_{\text{kd}}} \frac{g_{*s}(x)}{g_*^{1/2}(x)} \frac{\langle\sigma v\rangle}{x^2} dx + \int_{x_{\text{kd}}}^{x_0} \frac{g_{*s}(x)}{g_*^{1/2}(x)} \frac{\langle\sigma v\rangle}{x^2} dx \right]^{-1}$$

$$\langle\sigma v\rangle = (2\pi v_0^2)^{-\frac{3}{2}} \int d^3v e^{-\frac{v^2}{2v_0^2}} S_n(v) \times (\sigma v)^{\text{tree}}$$

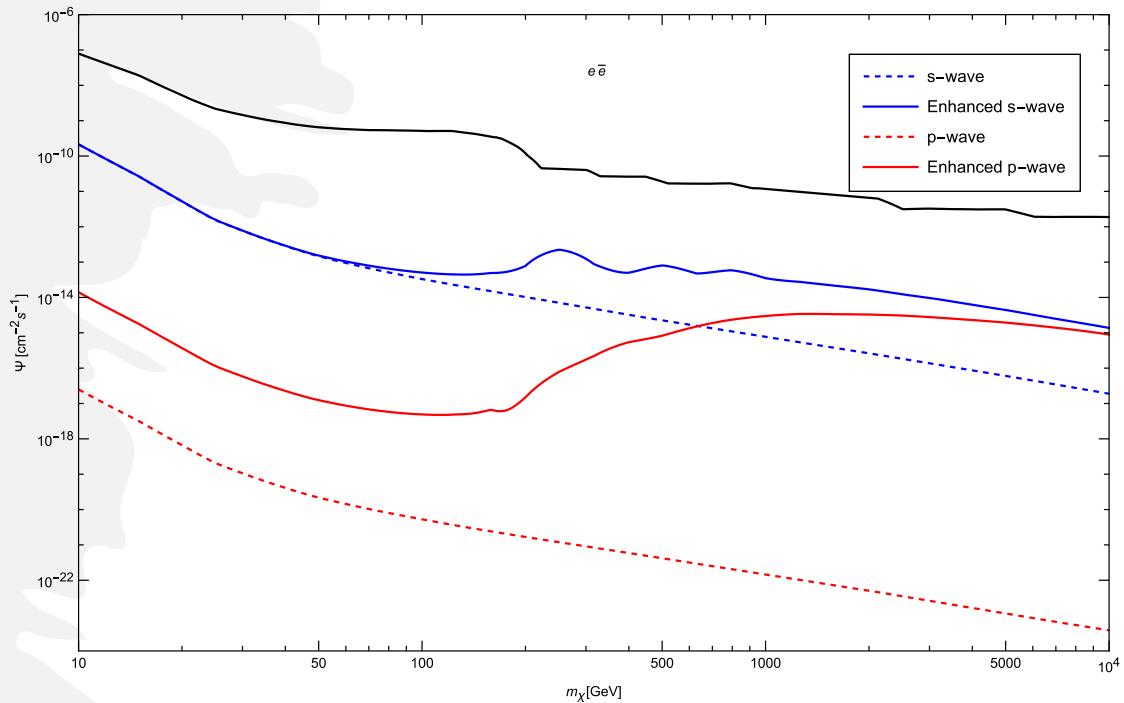
$$v_0 = \sqrt{2/x}$$

$$v_0 = \sqrt{2m_\chi/T_{\text{kd}}}/x$$

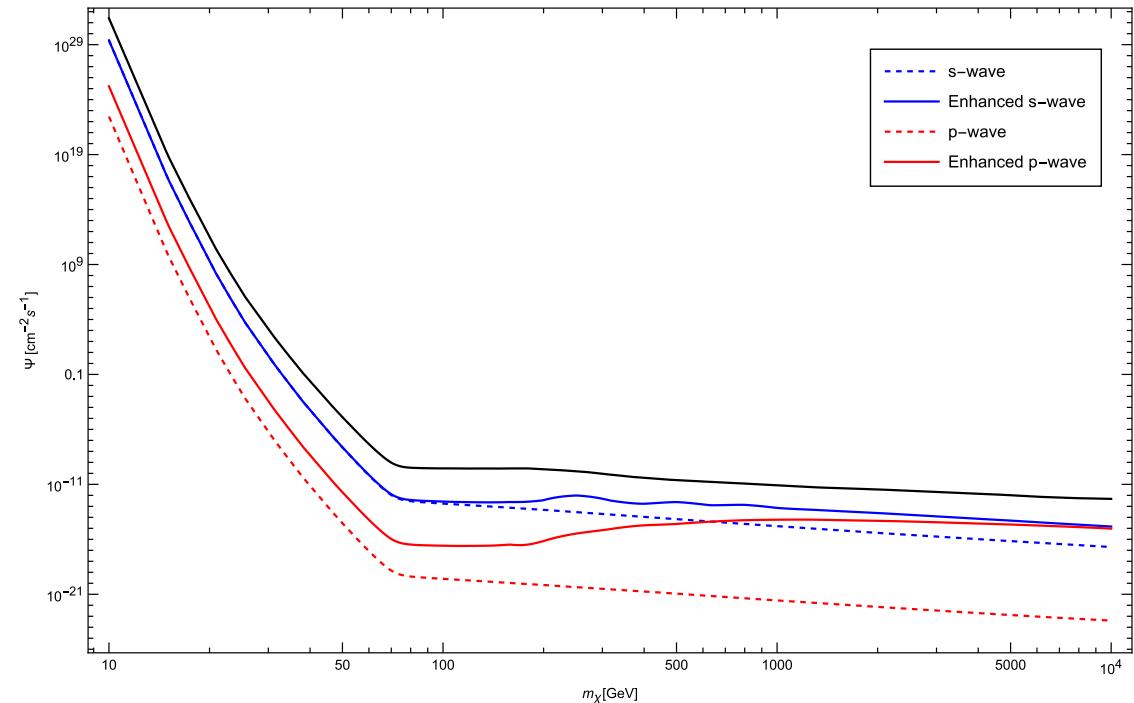


Flux Comparison

$$m_\phi = 5 \text{ GeV}$$



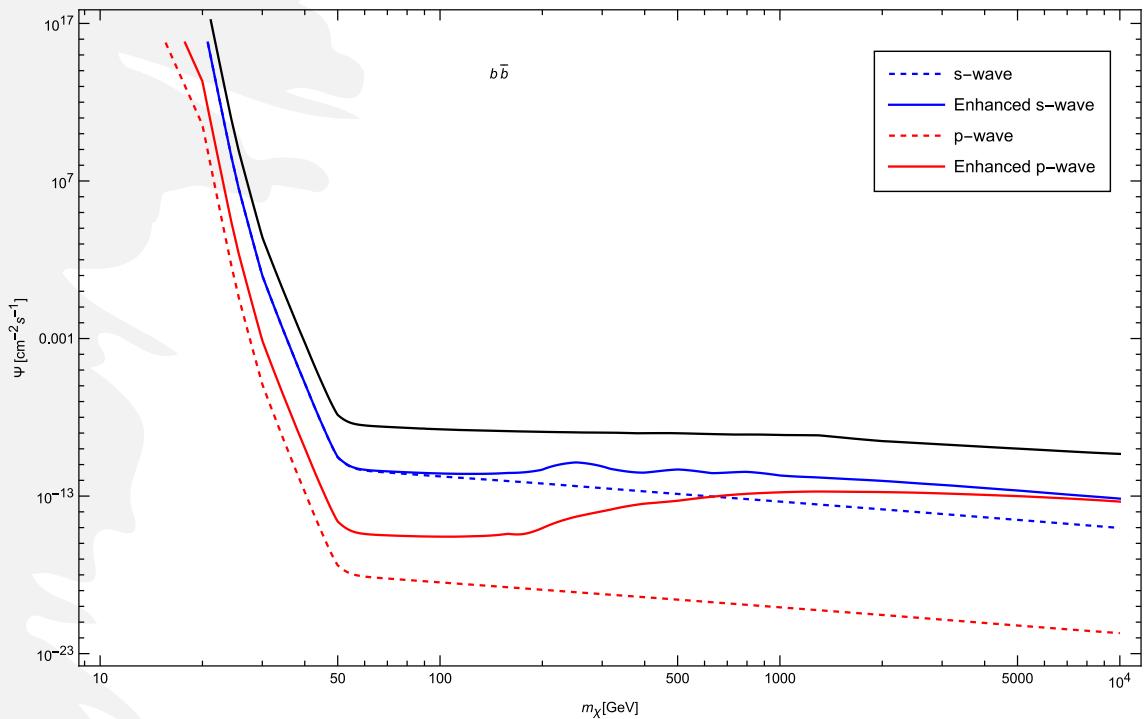
$e\bar{e}$



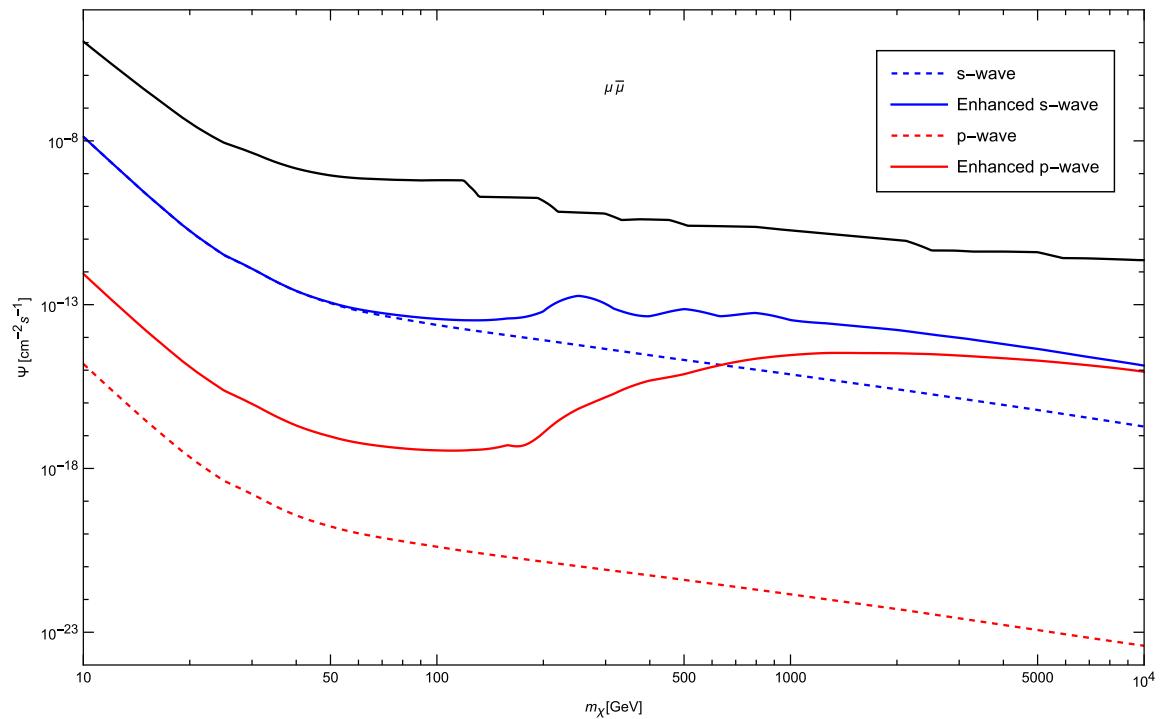
$\tau\bar{\tau}$

Flux Comparison

$$m_\phi = 5 \text{ GeV}$$



$b\bar{b}$



$\mu\bar{\mu}$

Summary

- Existence of Supermassive black hole redistributes Dark Matter halo on the region of influence, creating a spike.
- The spike boosts the annihilation cross section.
- In the light mediator model, Sommerfeld enhancement further enhances the annihilation cross section.
- Dark matter spike and Sommerfeld enhancement boost the p-wave annihilation flux, making it comparable to the s-wave annihilation flux.



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Thank You