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CHICAGO



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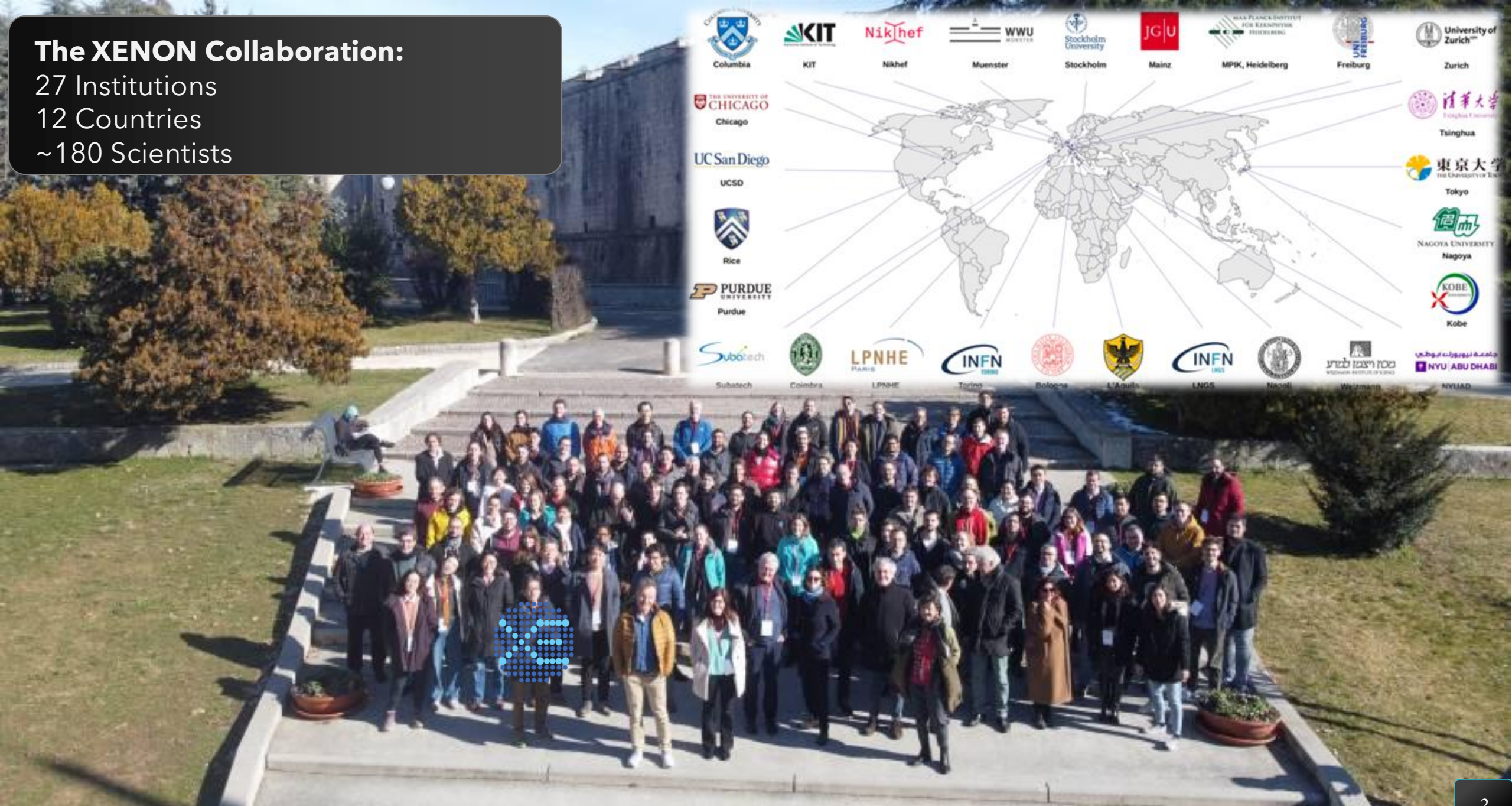
# XENONnT The First WIMP Results

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University of Chicago  
XENON Collaboration

26 June 2023

# The XENON Collaboration:

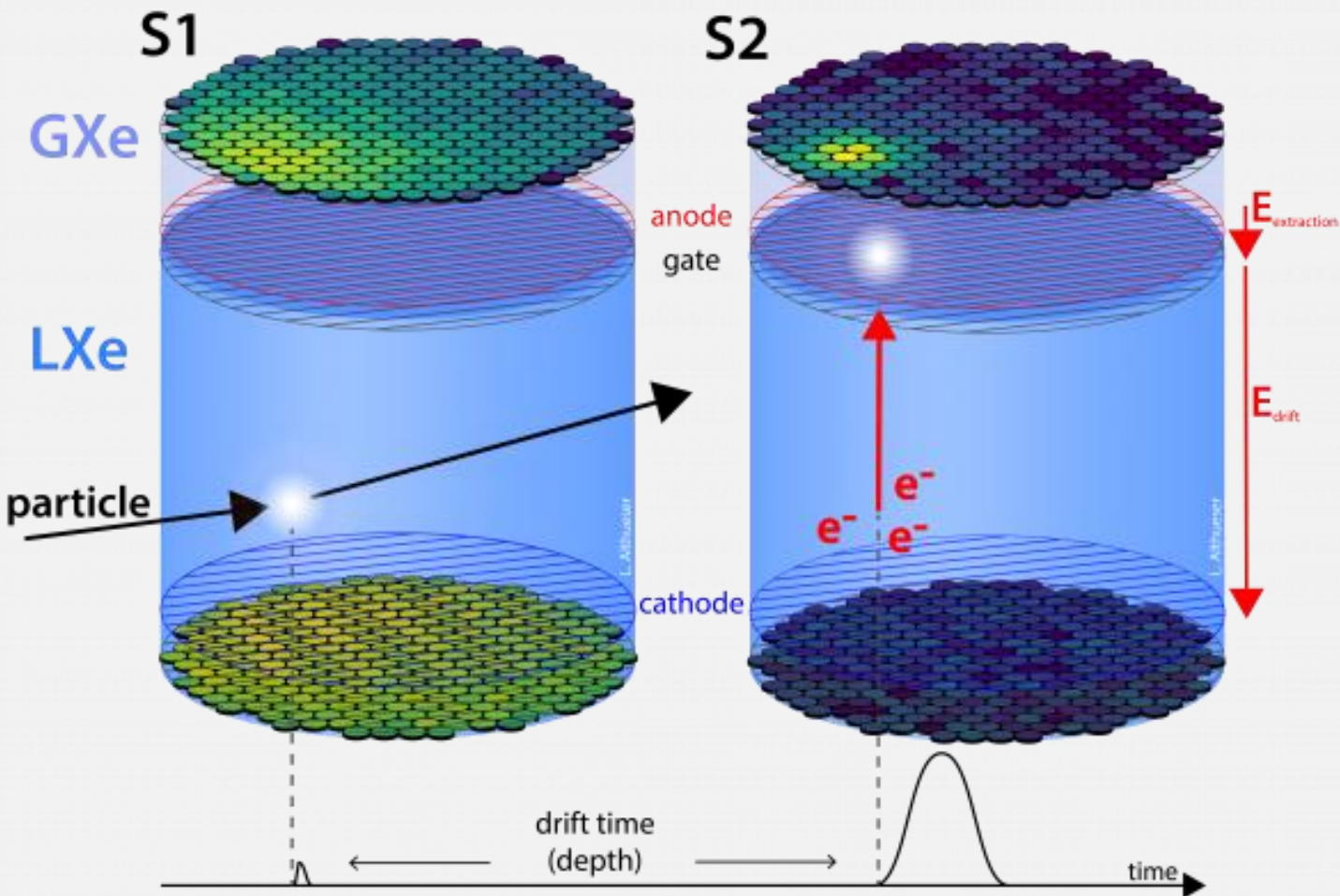
27 Institutions  
12 Countries  
~180 Scientists





# XENONnT: A Two-phase Time Projection Chamber

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## Signal detection after particle interaction

- Prompt scintillation light (**S1**)
- Delayed signal from charge extracted into gas phase (**S2**)

## Full 3D position reconstruction

- x-y from PMT response to S2
- Depth (z) from S1-S2 delay

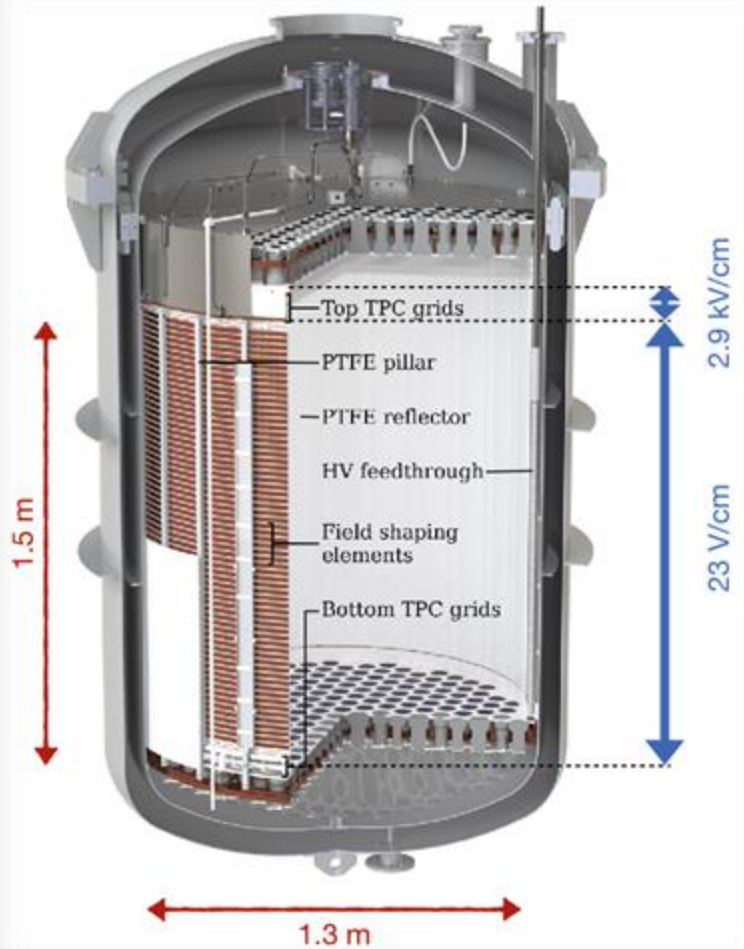
## Particle discrimination through S1/S2 signal ratio

- Neutrons, WIMPs: **Nuclear Recoil**
- Gammas, Betas: **Electronic Recoil**



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# The XENONnT TPC in its First Science Run



## TPC Design

- 8.5 tonnes of Liquid Xenon (LXe) total
- 5.9 tonnes LXe inside TPC

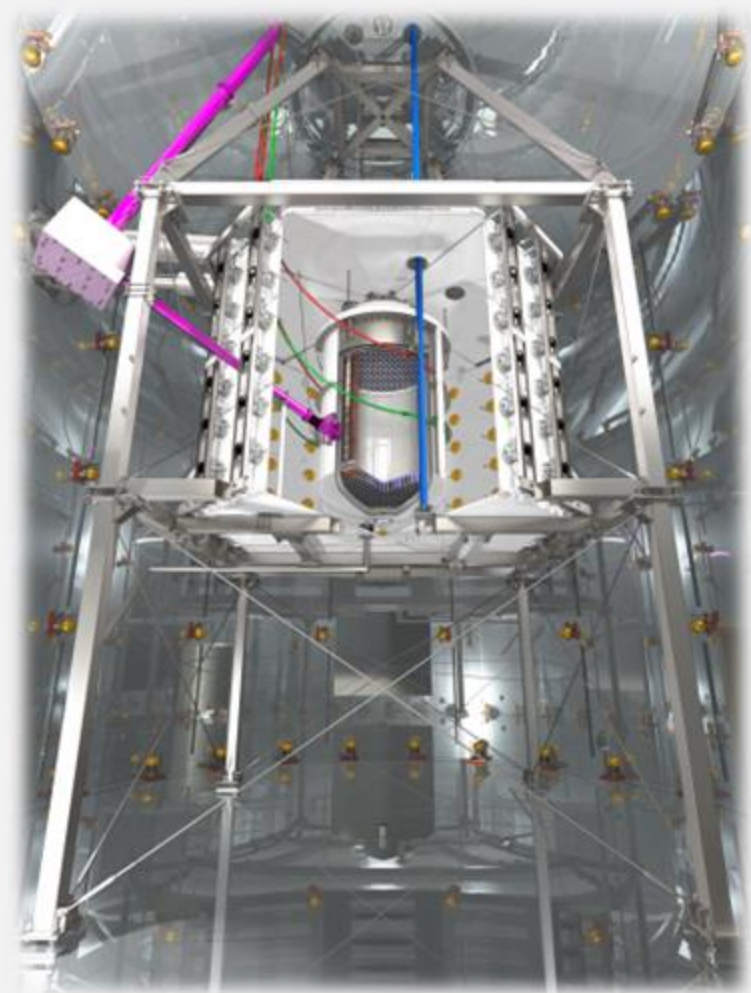
## First Science Run (SR0) Configuration

- Drift field 23 V/cm
- 477 out of 494 PMTs operational
- Extraction field of 2.9 kV/cm (in Liquid)  
~50% Extraction efficiency



# Important new Subsystems

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### Neutron Veto

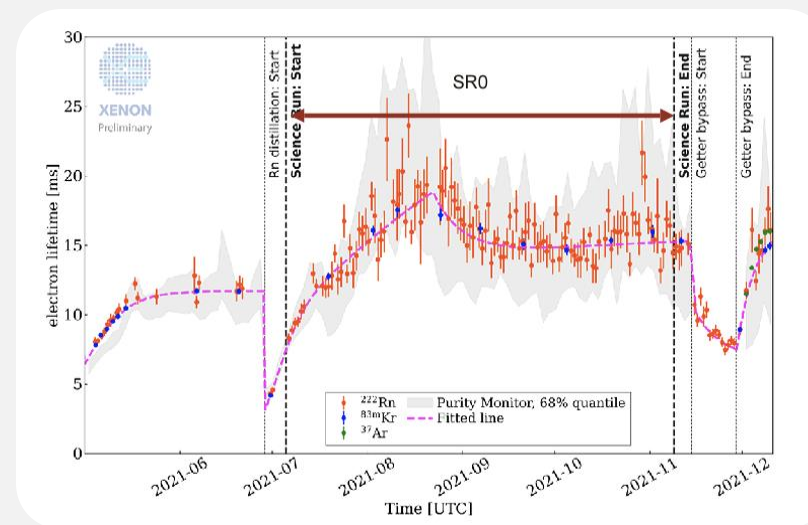
- Built inside muon veto (700 tonne water tank)
- 120 PMTs optically separated from muon veto by PTFE panels
- Tagging efficiency for neutrons of 53.3% (lifetime loss of 1.6%)

### Radon Distillation Column

- Allows for continuous Rn removal via cryogenic distillation
- Rn activity reduced to  $<2 \mu\text{Bq/kg}$
- Electronic recoil background 1/5 of XENON1T

### Liquid Xenon Purification

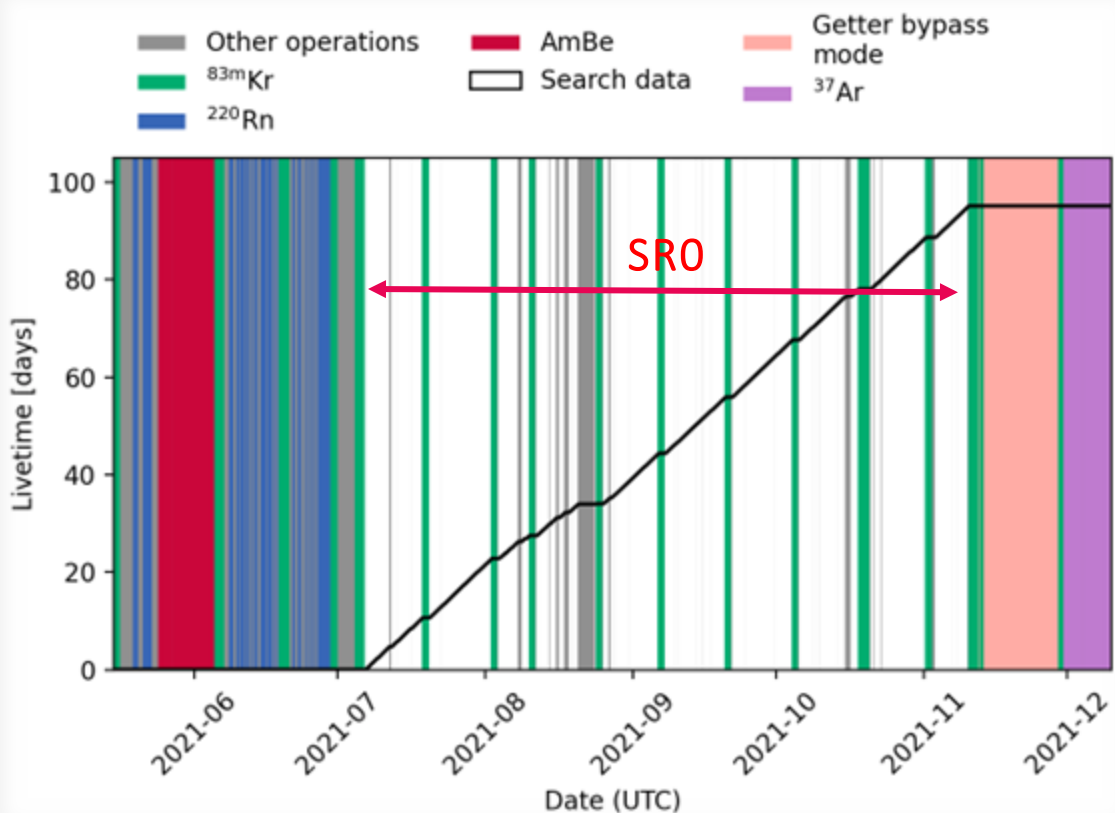
- Electronegative impurities removed from liquid phase
- 2 liters of LXe/min (Entire inventory in 18h)
- Achieved electron lifetime  $\sim 15 \text{ ms}$





# Data Taking for SR0

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## Calibration before/after dark matter search

- Nuclear Recoil Response: **AmBe**
- Electronic Recoil Response: **Rn**
- Light yield and Corrections: **Kr** and **Ar**

## SR0

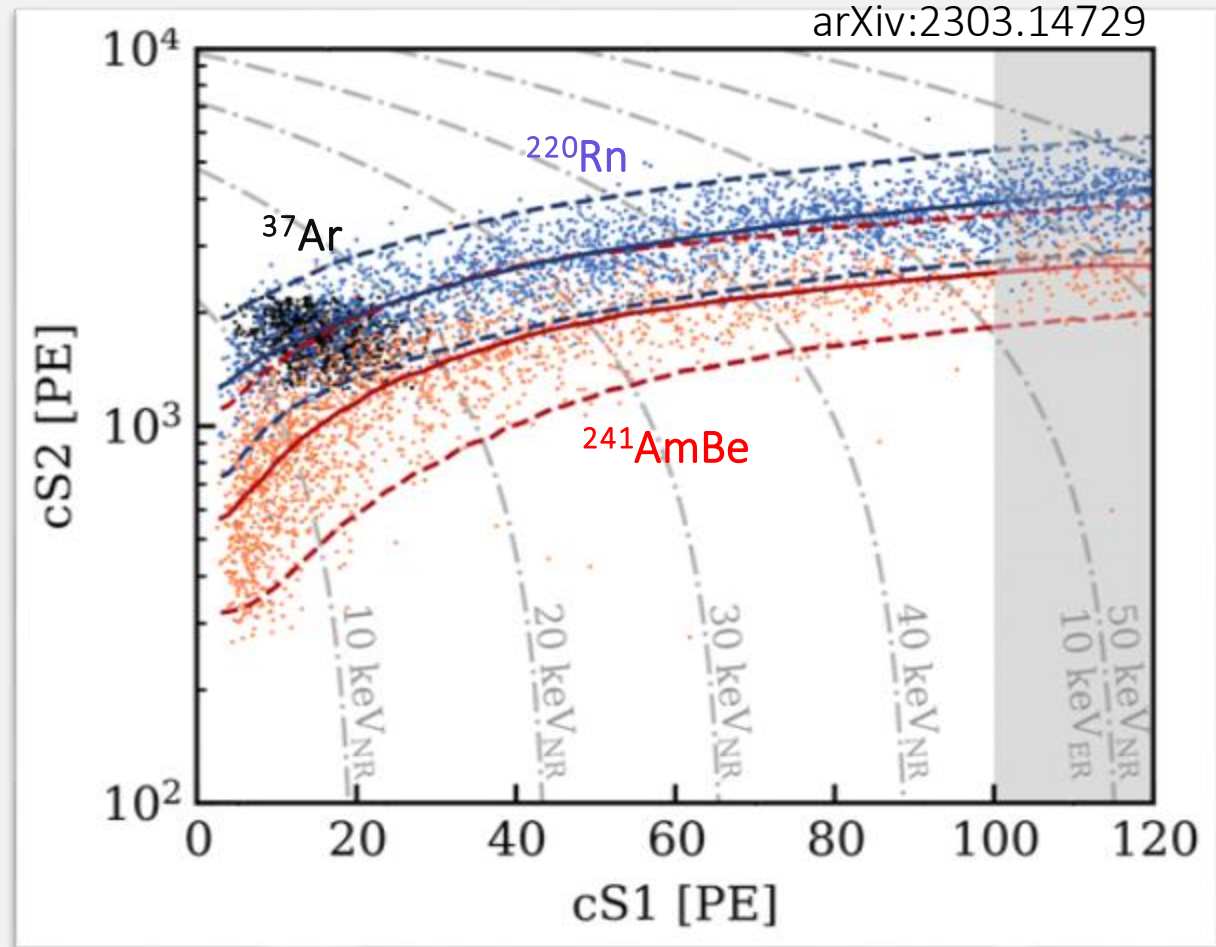
- July - November (97.1 real live days)
- **95.1** live days after corrections
- **(4.18 ± 0.13) tonnes** Fiducial Volume
- Total exposure of **1.1 tonne.year**



# Calibration of the Detector Response

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arXiv:2303.14729



## Nuclear Recoil

- Neutrons from external  $^{241}\text{AmBe}$  source
- Selected via coincident 4.4 MeV gamma in NV

## Electronic Recoil

- $^{220}\text{Rn}$  has flat energy spectrum in region of interest
- Validates cut acceptances

## Low energy

- $^{37}\text{Ar}$  has monoenergetic signal at  $2.8 \text{ keV}_{\text{ER}}$
- Validate response at lowest energies

## Detector response model

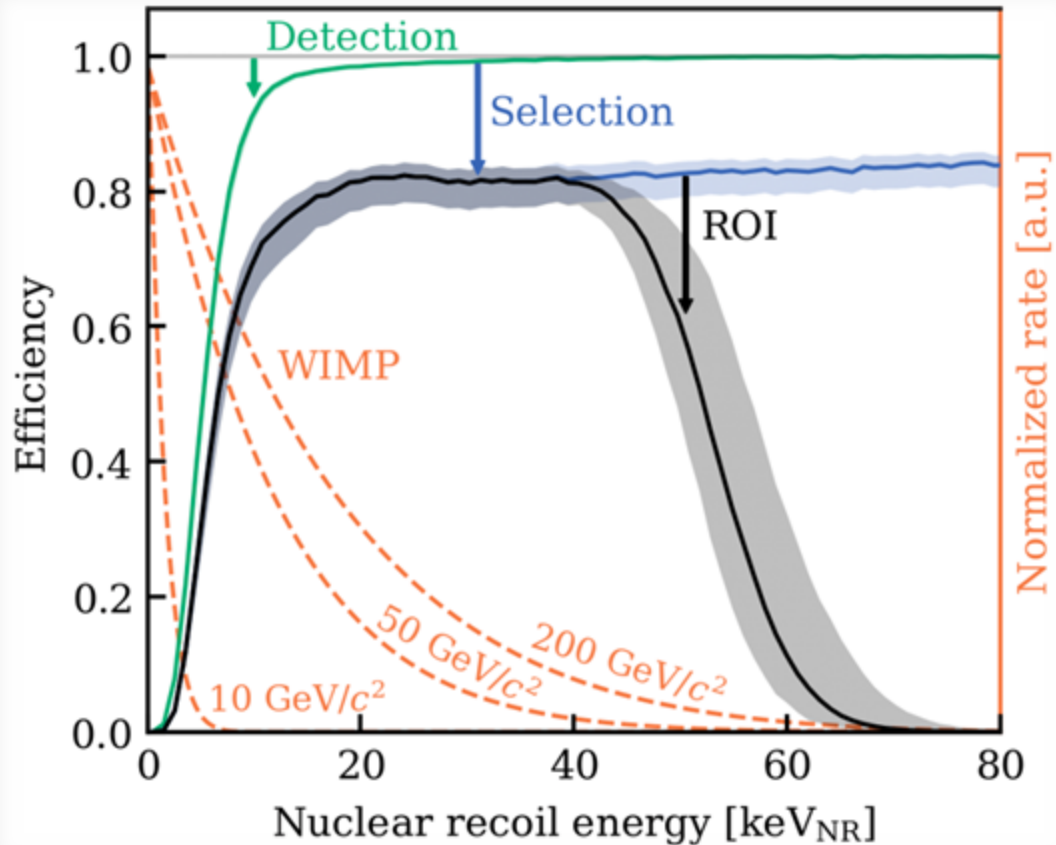
- LXe NR response derived from fit to  $^{241}\text{AmBe}$
- Combined fit to  $^{37}\text{Ar}$  and  $^{220}\text{Rn}$  data to derive ER response model



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# Detection and Selection Efficiencies

arXiv:2303.14729



## Detection Efficiency

- Dominated by 3-fold PMT coincidence for S1s
- Derived from waveform simulation, verified by data driven approach

## Selection Efficiency

- Removal of unphysical and multi-site events
- Plateaus at 80%

## Region of Interest

- Corrected S1 (cS1) range: 0-100 PE
- Corrected S2 (cS2) range: 10<sup>2.1</sup> - 10<sup>4.1</sup> PE
- Covers recoil energies of ~GeV WIMPs

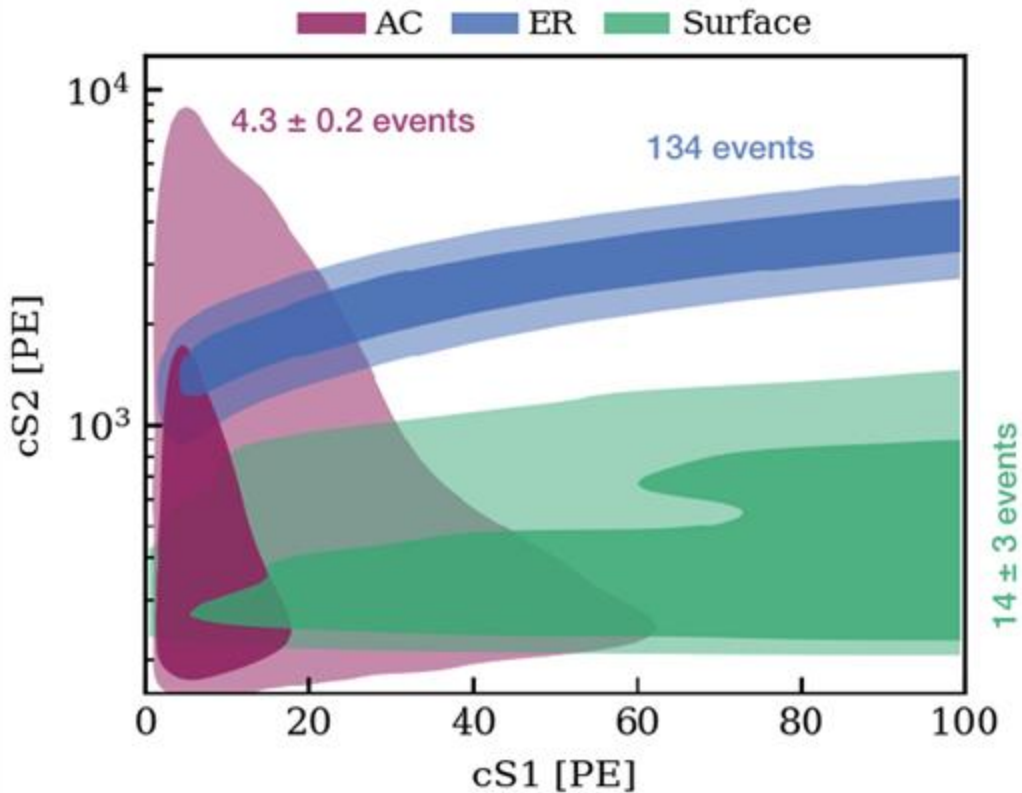




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# Backgrounds in WIMP ROI

1 $\sigma$  and 2 $\sigma$  contours of BG components:



Total Expectation: 154 Events

## Accidental Coincidence (AC)

- Random pairing of lone S2s and S1s
- Use GBDT cut, applied to S2 features to suppress

## Electronic Recoils (ER)

- Major contribution from beta decay of <sup>214</sup>Pb
- Sub dominant contribution from <sup>85</sup>Kr

## Surface Background

- <sup>210</sup>Pb is known to plate out on PTFE walls
- Decays at walls result in charge loss before S2 is observed
- Suppressed by FV cut

## Nuclear Recoil backgrounds

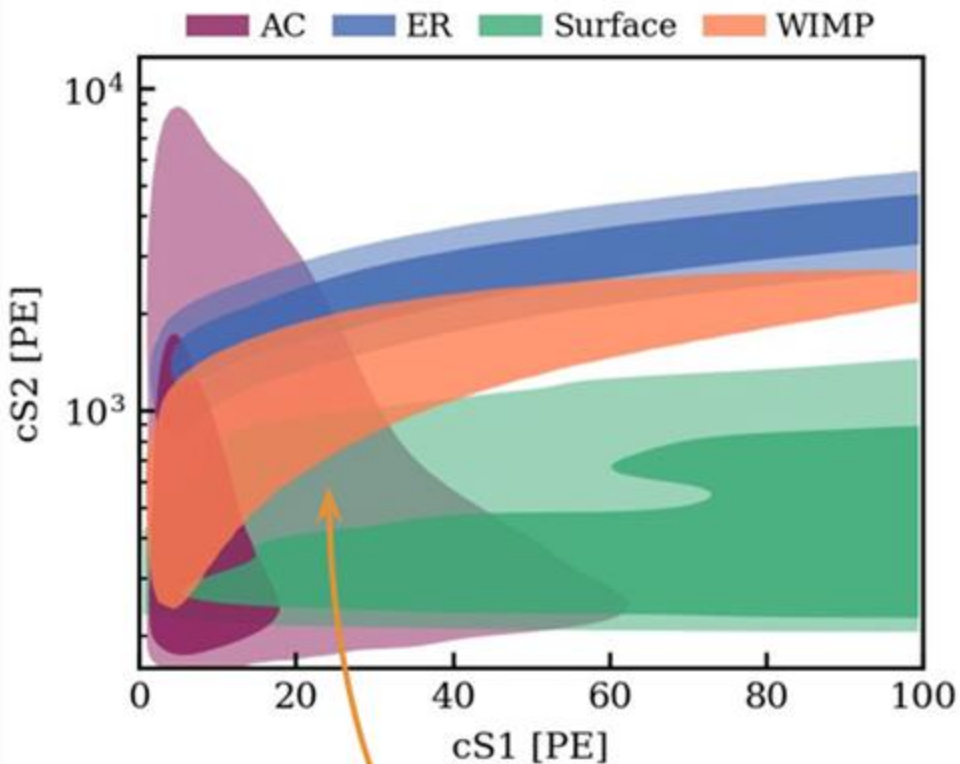
- Radiogenic neutrons constrained by NV tagging (~1.1 events)
- CEvNS expectation of 0.2 events



# Blinded WIMP Search in SR0

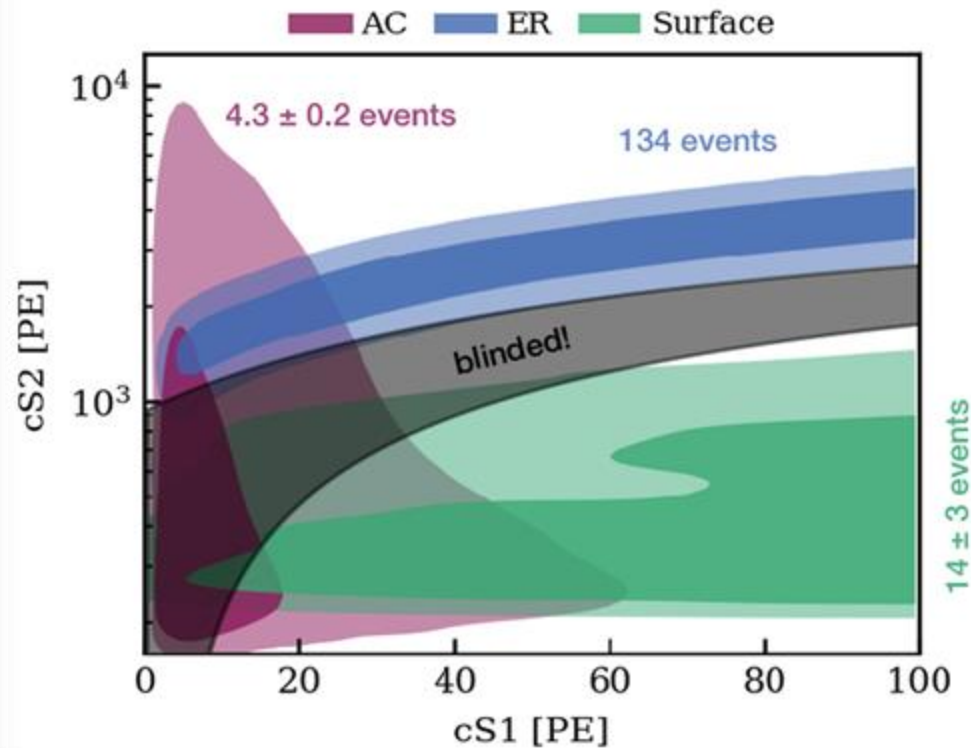
- WIMP signature is single interaction site nuclear recoil
- Analysis was blinded in lower portion of NR band

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2σ contour of 200 GeV/c<sup>2</sup> WIMP

1σ and 2σ contours of BG components:





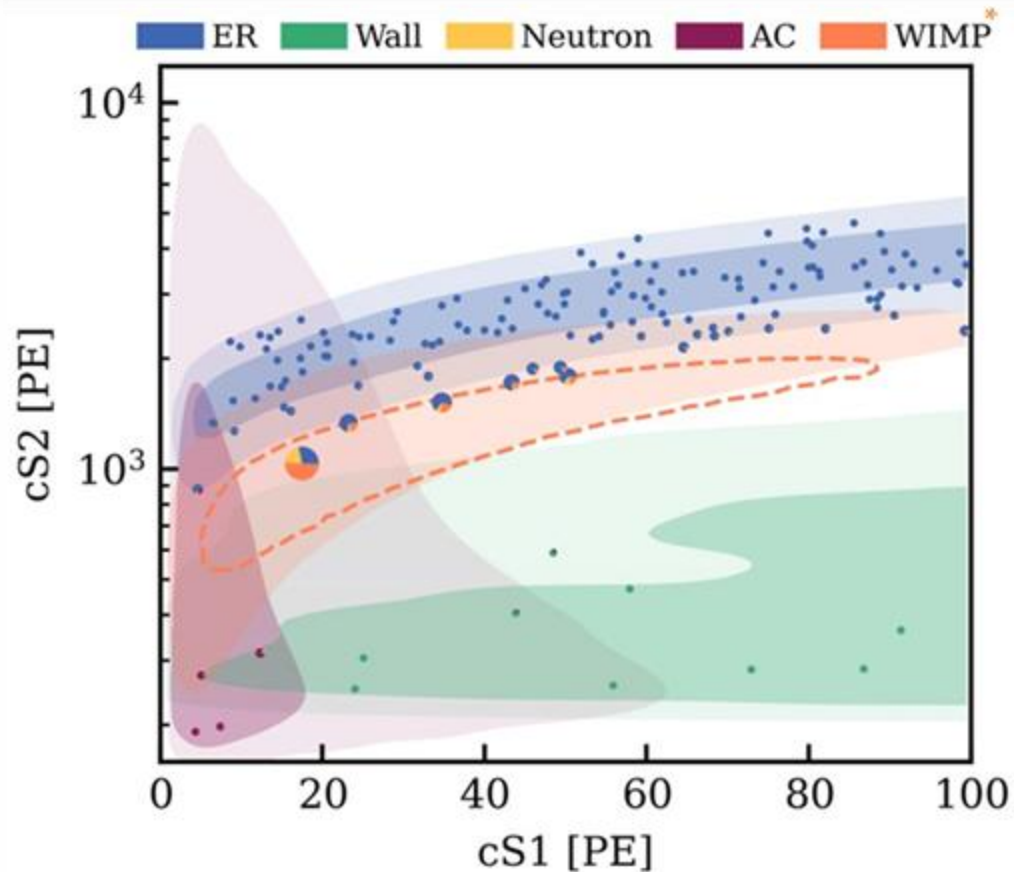
# Events in search region

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	Nominal	Best Fit (200 GeV/c <sup>2</sup> WIMP)
ER	134	135(+12)(-11)
Neutrons	1.1(+0.6)(-0.5)	1.1±0.4
CEvνS	0.23±0.06	0.23±0.06
AC	4.3±0.2	4.32±0.15
Surface	14±3	12(+0)(-4)
<b>Background Total</b>	<b>154</b>	<b>152±12</b>
WIMP	-	2.6
Observed	-	152

- No significant excess
- 152 events in ROI, 16 in blinded region

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\*Assuming a 200 GeV/c<sup>2</sup> WIMP

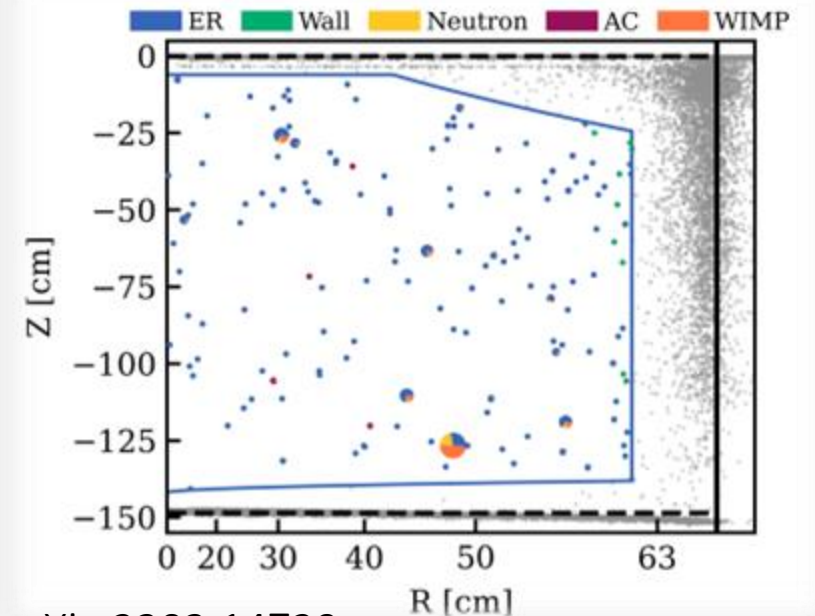
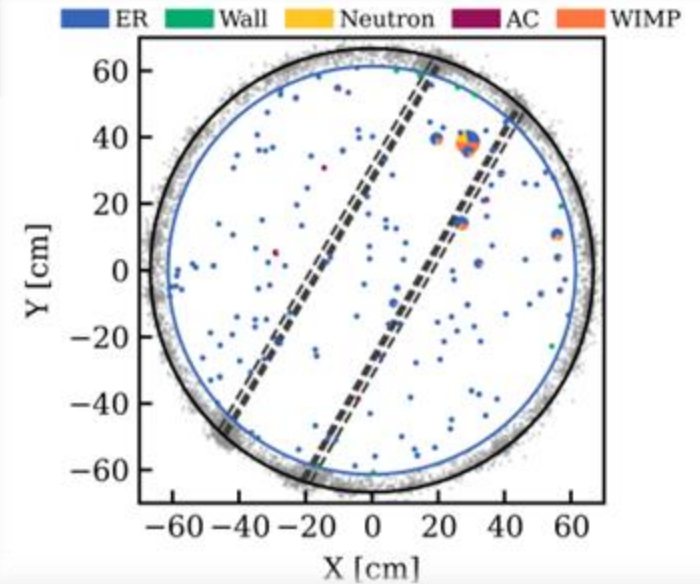
# WIMP Results

## Limit Setting

- Performed unbinned maximum likelihood
- Considered range of WIMP masses up to several hundred  $\text{GeV}/c^2$
- No significant excess for any WIMP mass

## Events in search region

- Represented as pie charts, displaying contribution of each background component in best fit model at event position

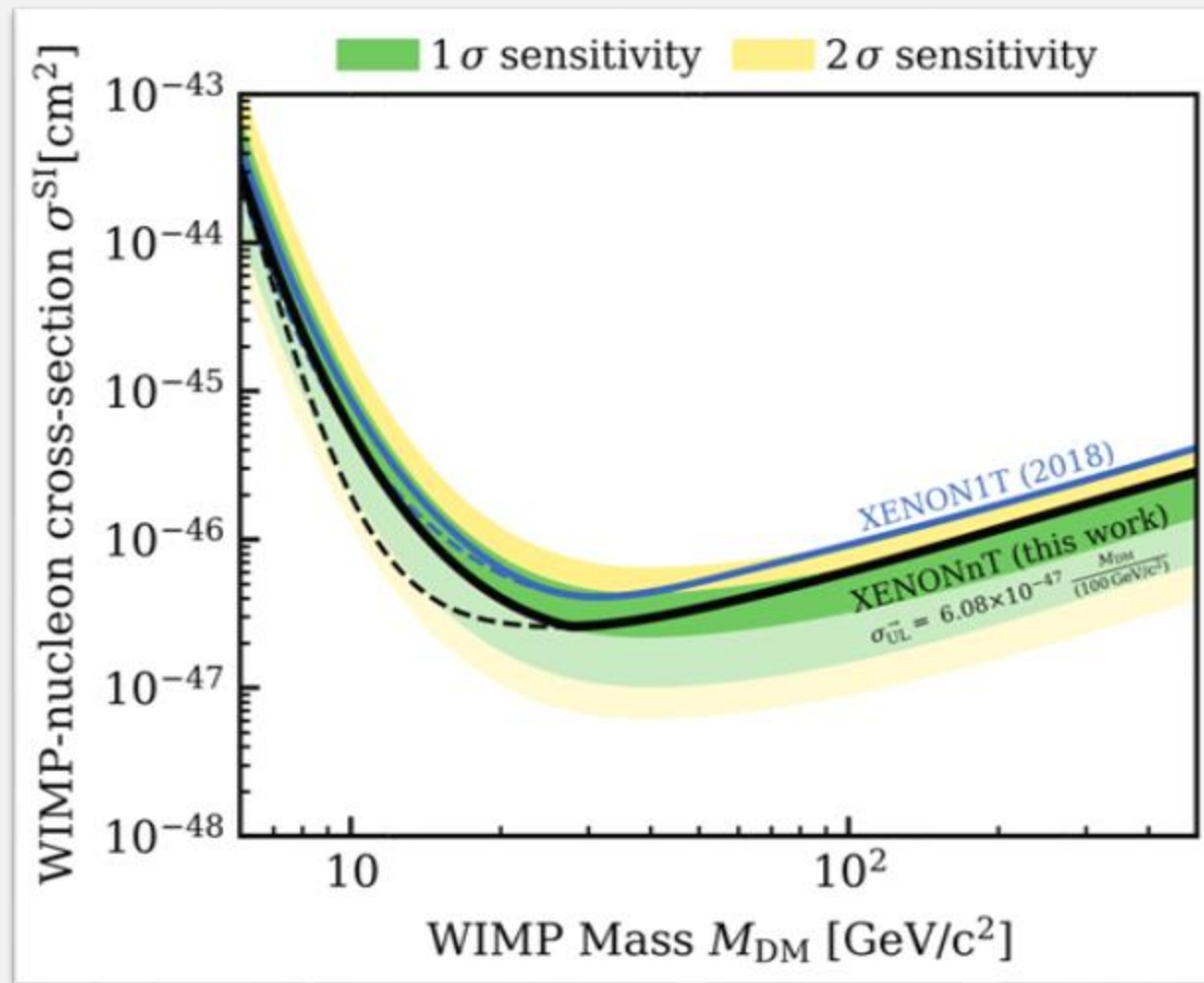




# Limit Setting Procedure in XENONnT

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- Power constrained limits to avoid spurious exclusion limits
- Minimum rejection power set to 50%
- Limit in effect constrained to median of sensitivity band
- Choice is conservative within community
- Strongest exclusion limit at **28 GeV/c<sup>2</sup>** :  
 $2.6 \cdot 10^{-47} \text{ cm}^2$  (90% CL)



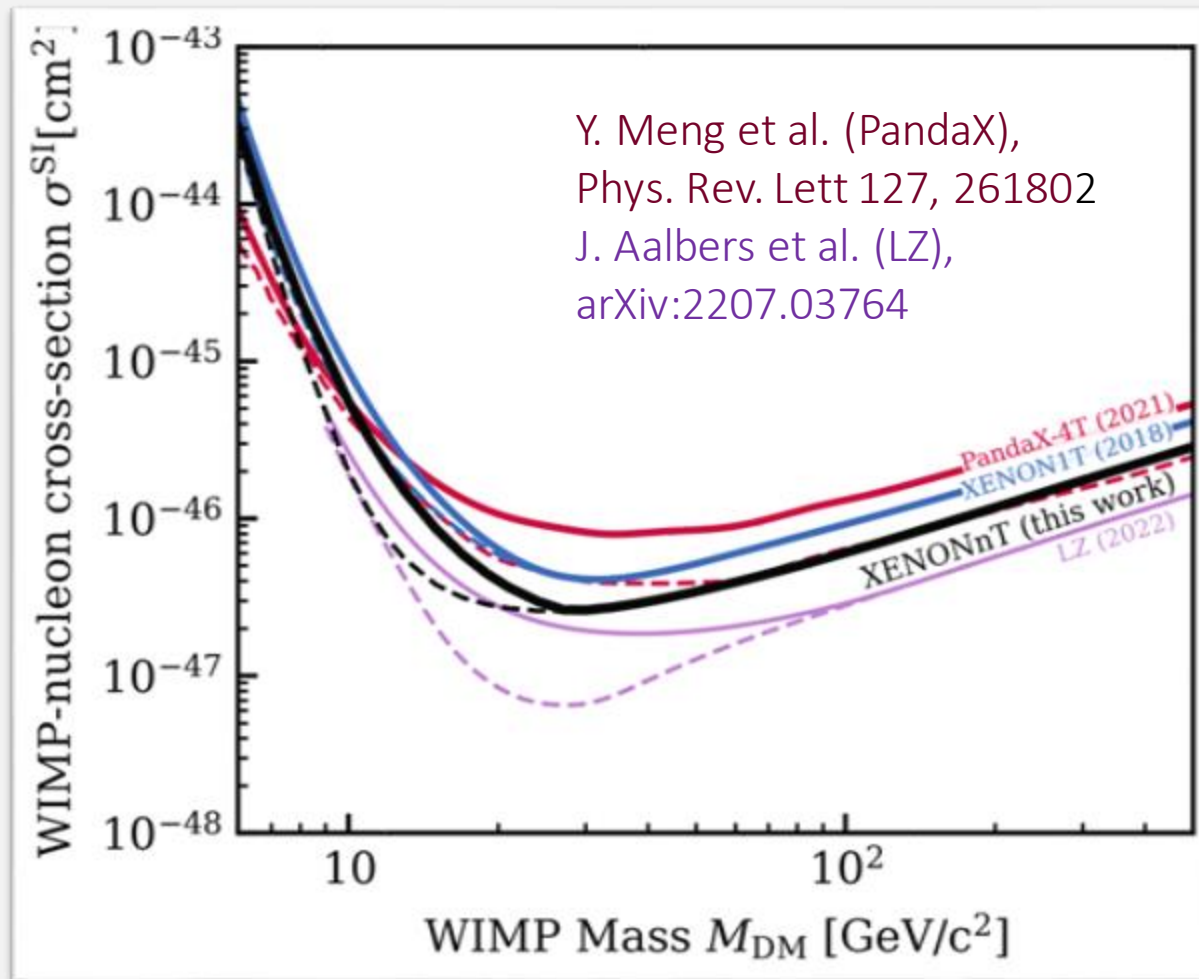
arXiv:2303.14729



# Limits on SI WIMP-nucleon cross section

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- Same procedure can be applied to previous limits
- Other Limits shown here are not blinded analysis
- XENONnT limit represents factor 1.6 improvement w.r.t XENON1T upper limit
- Data taken in considerably shorter time than XENON1T

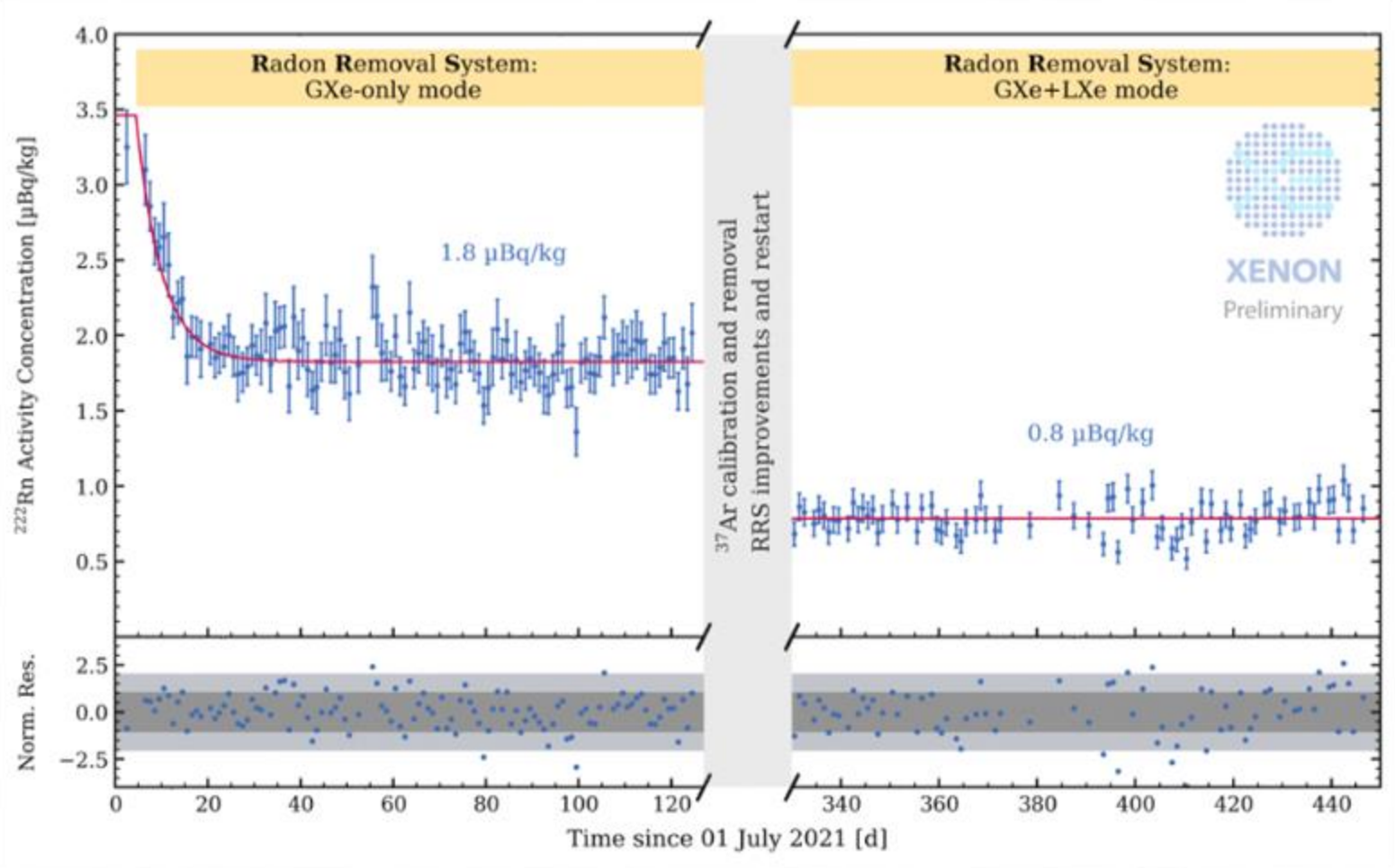


arXiv:2303.14729

Looking into the future!



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Detector Improvements

- SR1 data taking after further radon reduction
- Gd loading of Neutron Veto

Upcoming analyses

- Solar neutrino analyses
- S2-only studies



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