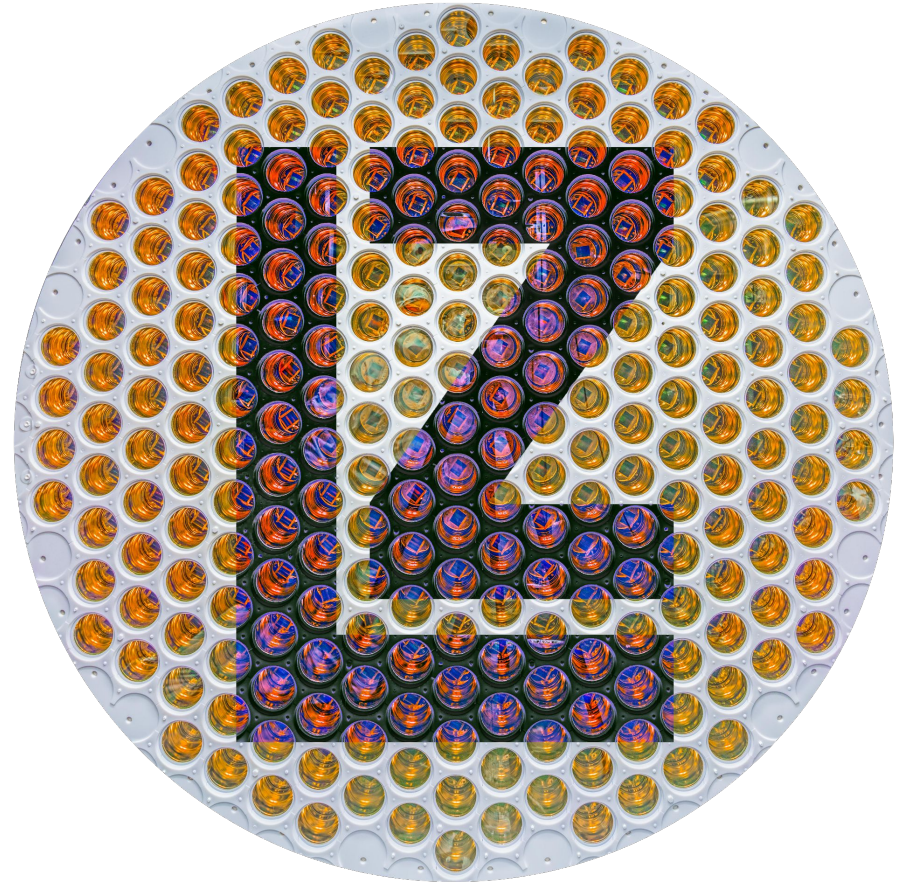


# First WIMP Search Results from LZ and Outlook

Hao Chen (LBNL)

On behalf of LZ Collaboration



# Dark matter

Evidence for dark matter:

- galaxy rotation curves
- cosmic microwave background
- Bullet cluster collision
- And more

All from **gravitational** and **collective** effects.



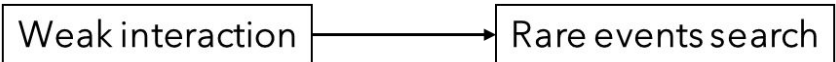
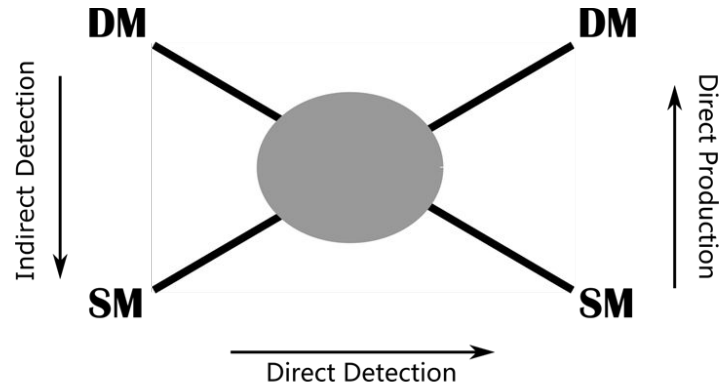
Strongest evidence for beyond standard model



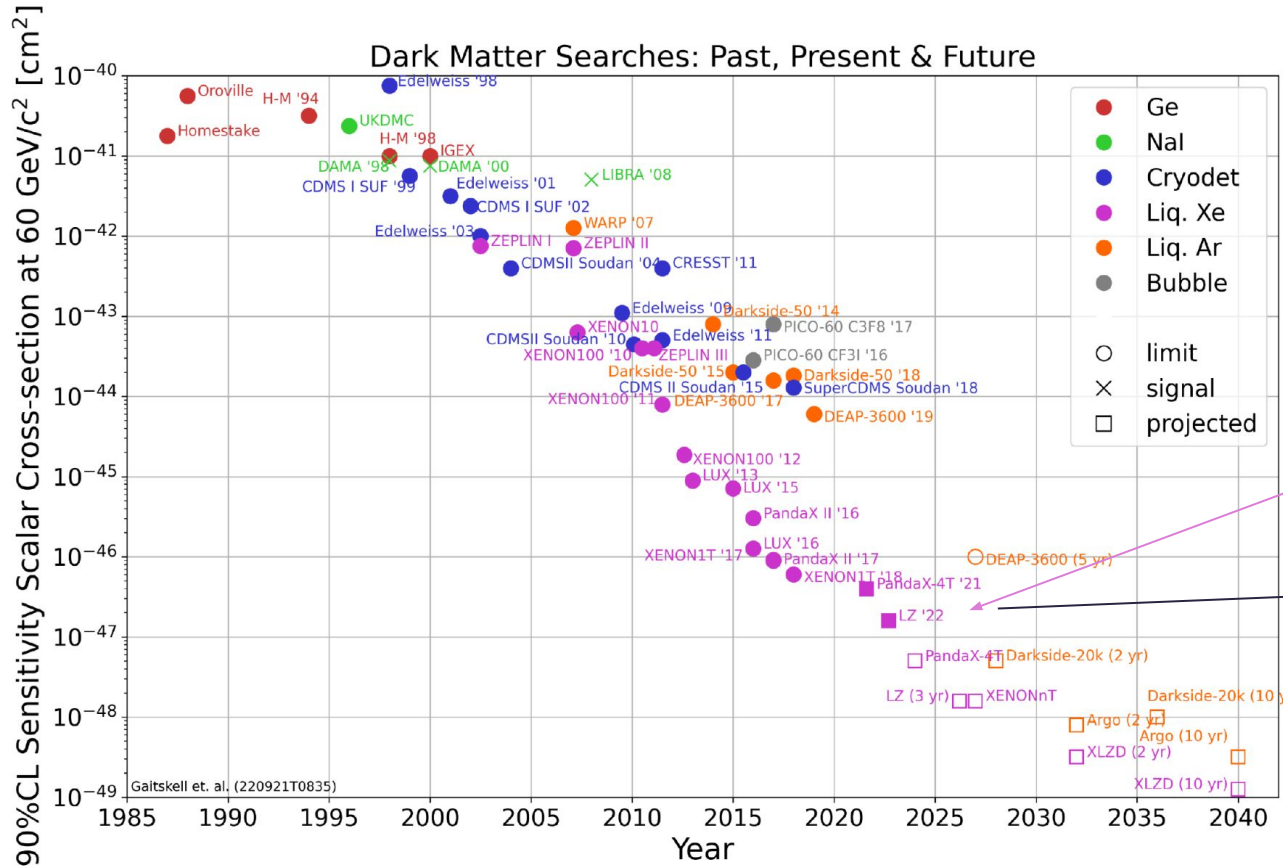
Many theoretic **candidates** with mass range from  $10^{-22}$  eV to 10 times sun mass.

- **Weakly interacting massive particles (WIMP)** ← LZ's main target
- Axion
- Many others, see arXiv:2211.09978

What is it? In the sense of elementary particle



# Direct detection of WIMP



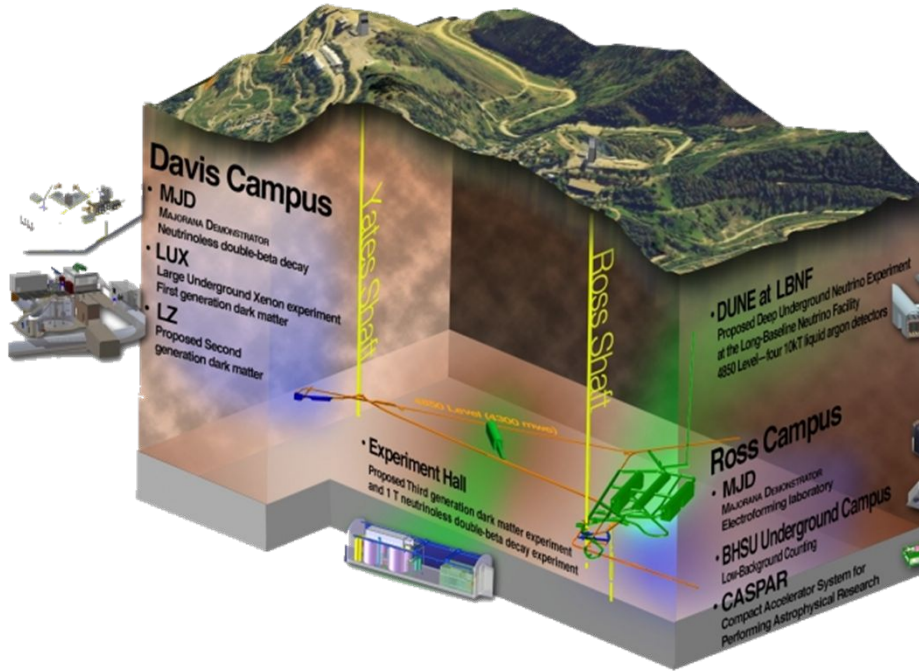
This talk

1 kg xenon needs 150 years to see one 60 GeV WIMP

# Sanford Underground Research Facility (SURF)

LZ at SURF in Lead, South Dakota, USA

Shielded by an overburden of 4300m water-equivalent  
**Muon flux reduced by  $O(10^6)$**



# LZ (LUX-ZEPLIN) Collaboration, 37 Institutions

250 scientists, engineers, and technical staff

- Black Hills State University
- Brookhaven National Laboratory
- Brown University
- Center for Underground Physics
- Edinburgh University
- Fermi National Accelerator Lab.
- Imperial College London
- King's College London
- Lawrence Berkeley National Lab.
- Lawrence Livermore National Lab.
- LIP Coimbra
- Northwestern University
- Pennsylvania State University
- Royal Holloway University of London
- SLAC National Accelerator Lab.
- South Dakota School of Mines & Tech
- South Dakota Science & Technology Authority
- STFC Rutherford Appleton Lab.
- Texas A&M University
- University of Albany, SUNY
- University of Alabama
- University of Bristol
- University College London
- University of California Berkeley
- University of California Davis
- University of California Los Angeles
- University of California Santa Barbara
- University of Liverpool
- University of Maryland
- University of Massachusetts, Amherst
- University of Michigan
- University of Oxford
- University of Rochester
- University of Sheffield
- University of Sydney
- University of Texas at Austin
- University of Wisconsin, Madison



US UK Portugal Korea Australia

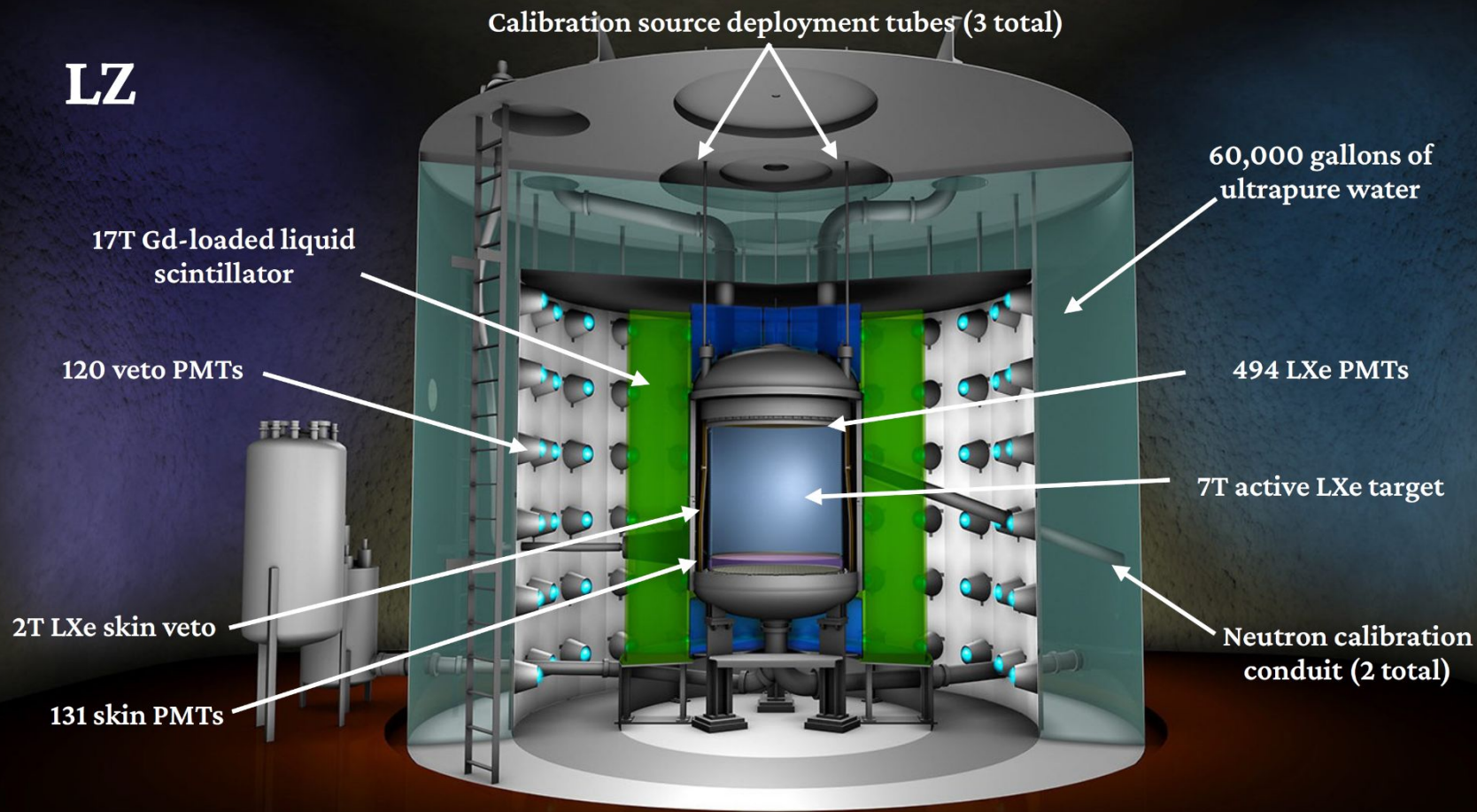


Science and  
Technology  
Facilities Council

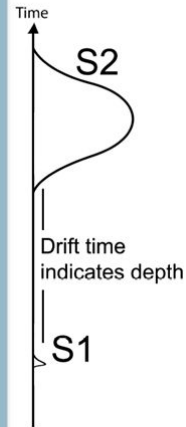
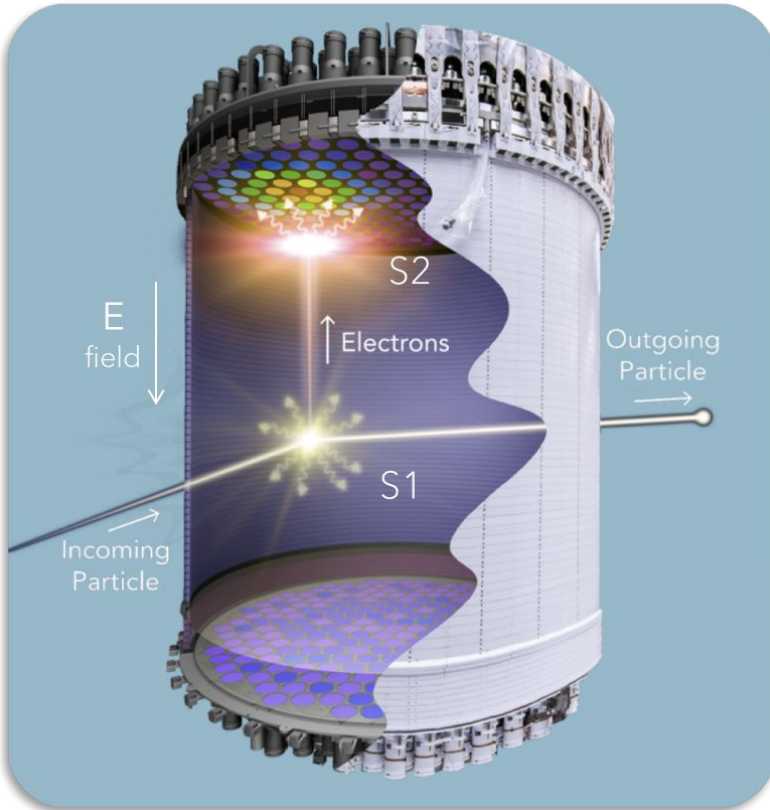


Thanks to our sponsors and participating institutions!

# LZ



# Main detector: the dual-phase time projection chamber

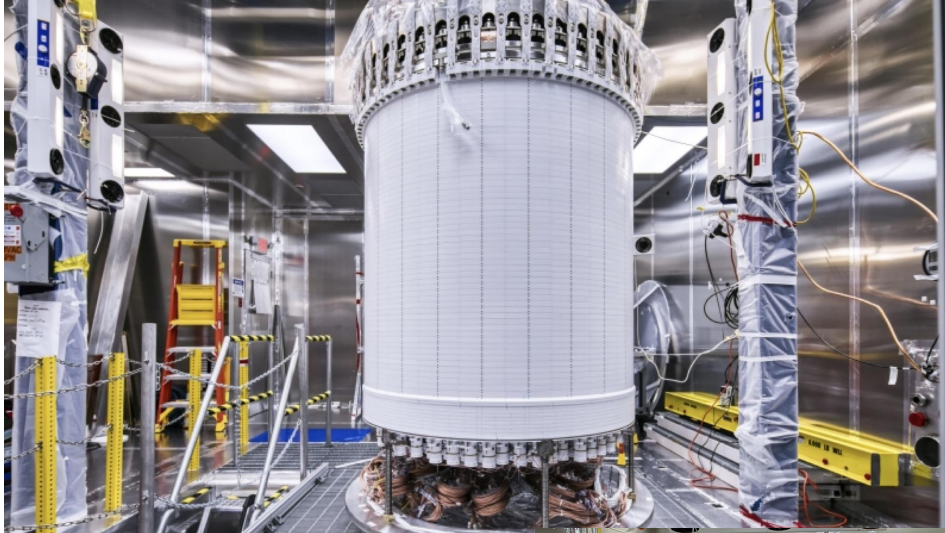


## Detector mechanism:

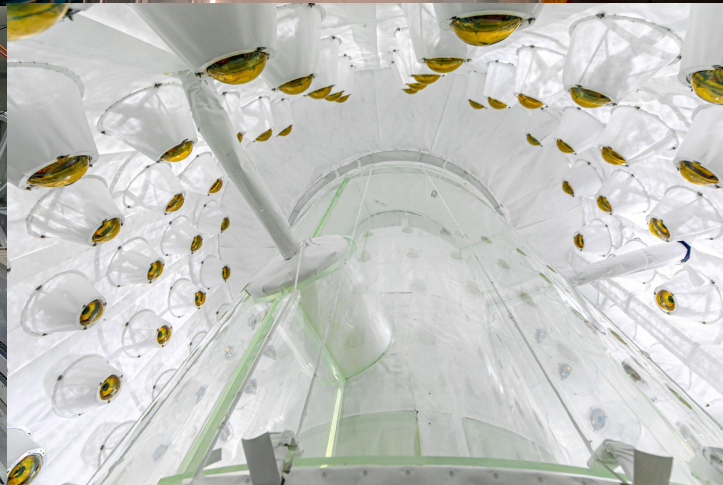
- Prompt scintillation photons → “S1”
- Ionization → electroluminescence → “S2”
- Energy reconstruction from S1/S2 size
- Position reconstruction from drift time and PMT hit pattern

## LZ specs:

- 1.5 m in diameter and height
- 7 t active xenon (5.5 fiducial)
- 494x 3” PMTs in two arrays
- 4 wire mesh electrodes + Ti field cage for uniform electric fields

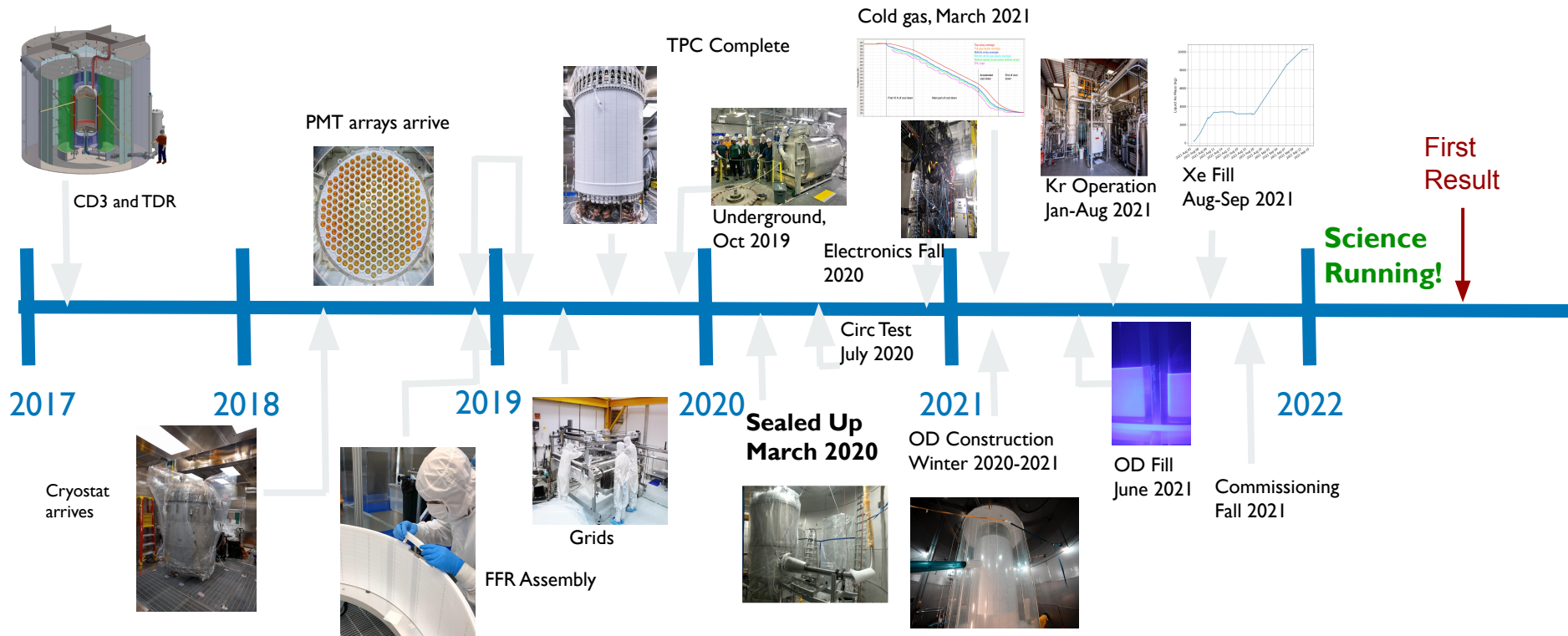


LZ for real





# LZ timeline



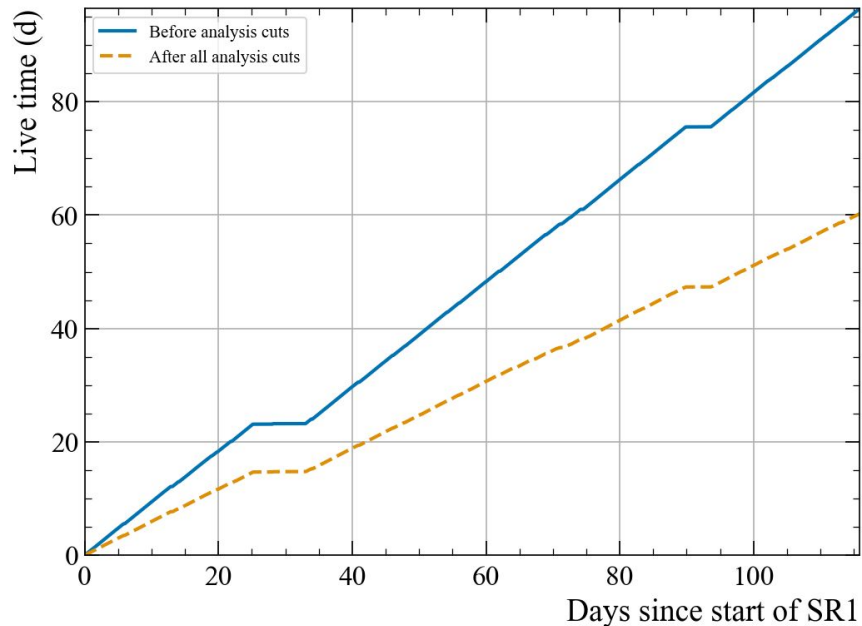
# First science run (SR1)

**Goal:** Demonstrate physics capability of the LZ detector

Key parameters during this run:

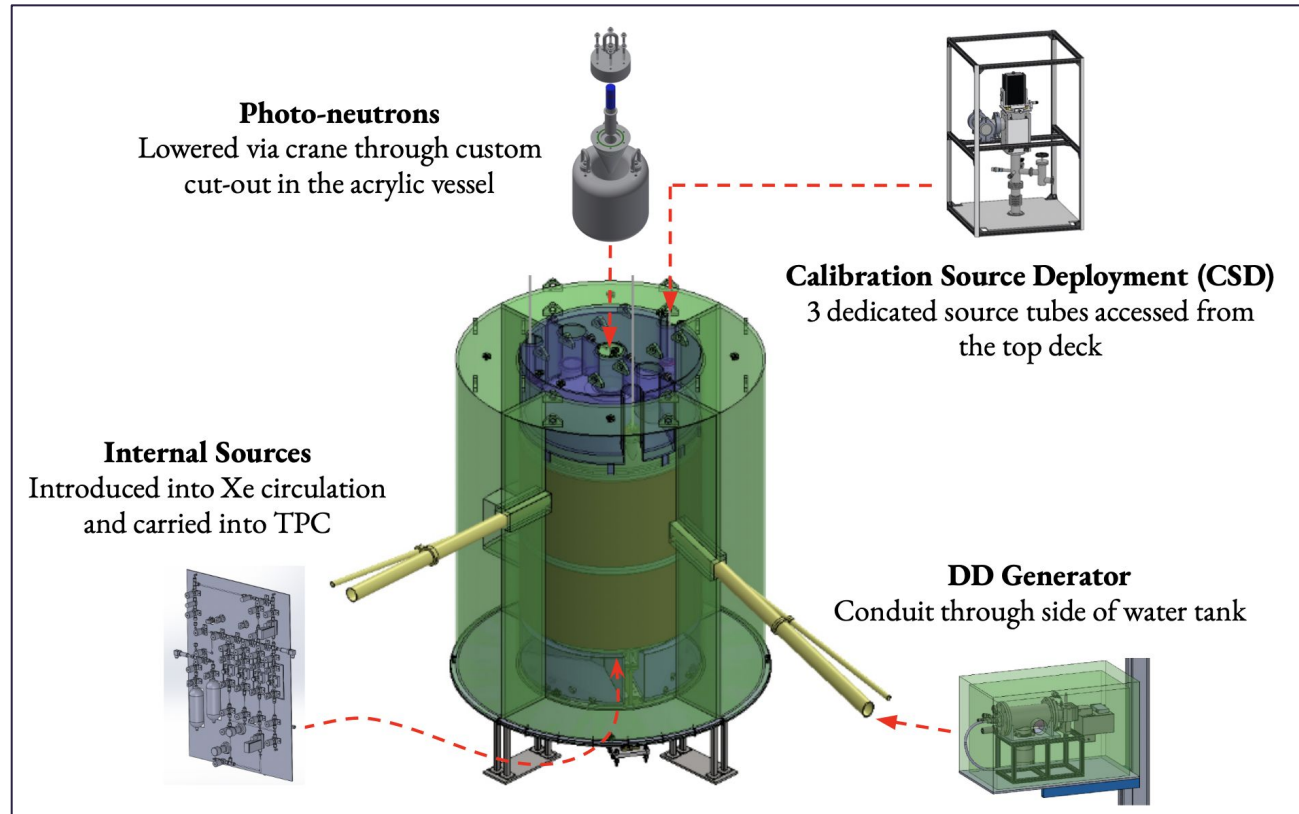
- Data taken starts 23 Dec 2021
- 5.5 t fiducial mass
- Drift field: 193 V/cm
- Gas extraction field: 7.3 kV/cm at center
- 174.1 K (at the TPC bottom) and 1.791 bar with <0.2% fluctuation
- Continuously purified at 3.3 t/day through a hot getter system
- Electron lifetime between 5000  $\mu\text{s}$  and 8000  $\mu\text{s}$ , much longer than the 951  $\mu\text{s}$  maximum drift time in the TPC

LZ detector reached the expected condition, it was stable during SR1.

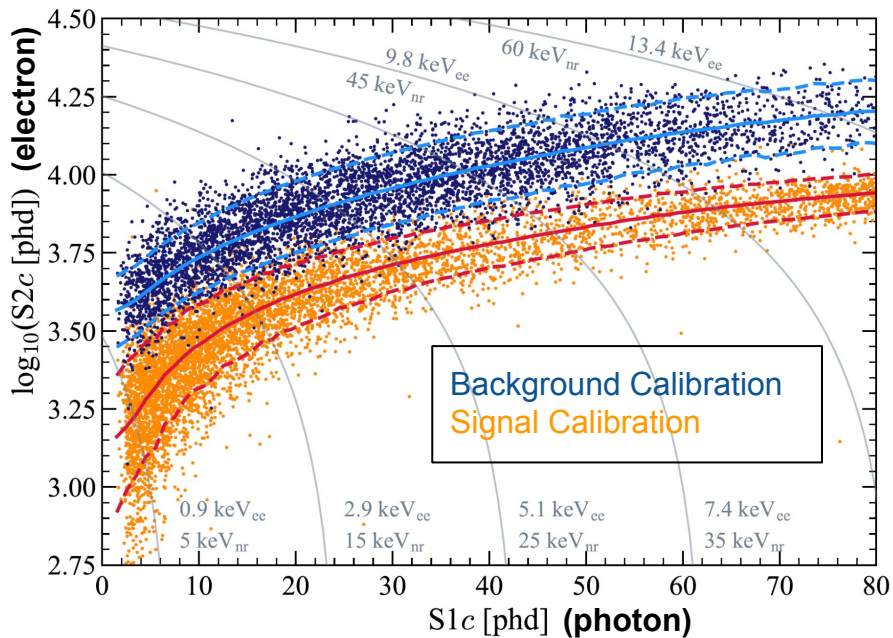


# Understand the tool → calibration

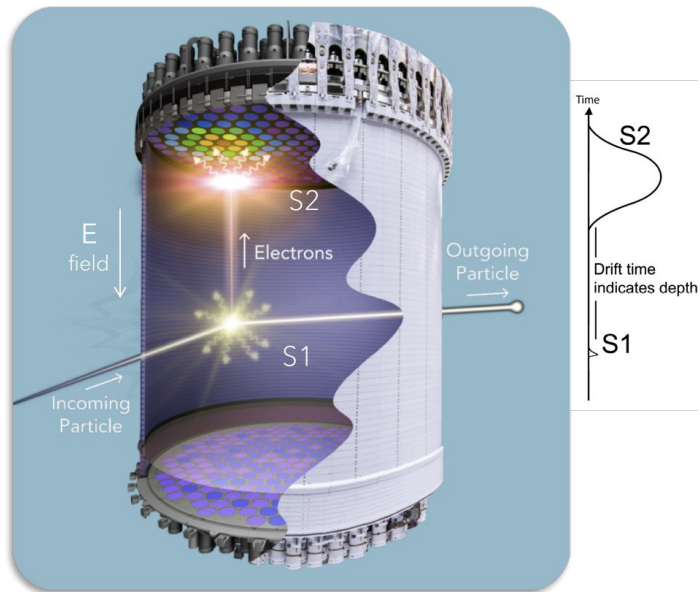
- **Neutron sources:** DD, AmLi, AmBe, YBe
- **Injected sources:**  $\text{CH}_3\text{T}$ ,  $^{83\text{m}}\text{Kr}$ ,  $^{131\text{m}}\text{Xe}$ ,  $^{220}\text{Rn}$
- **External gamma sources:**  $^{57}\text{Co}$ ,  $^{22}\text{Na}$ ,  $^{228}\text{Th}$ ,  $^{54}\text{Mn}$
- Additionally, background (e.g. alphas and cosmics), used for: energy scale calibration and other purposes



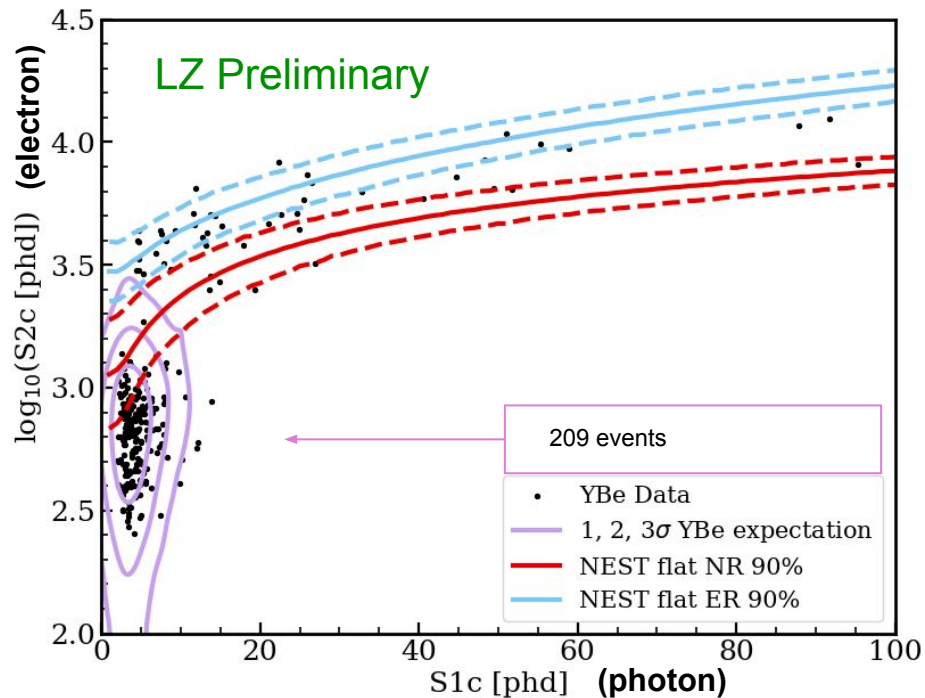
# Detector response



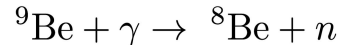
## Powerful tool to remove beta/gamma background



# Low-energy signal calibration



Yttrium-beryllium  
photoneutron source:



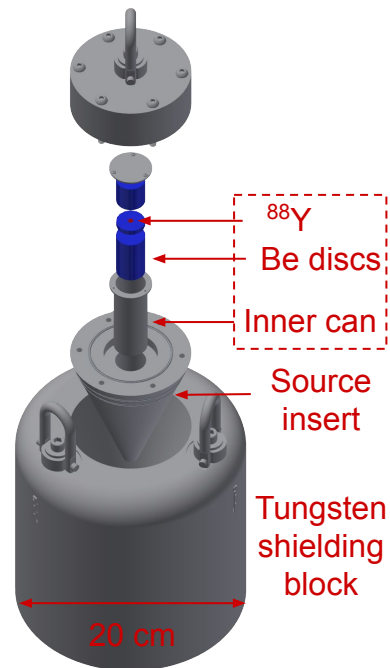
Q value = -1.667 MeV

Neutron energy: ~152 keV

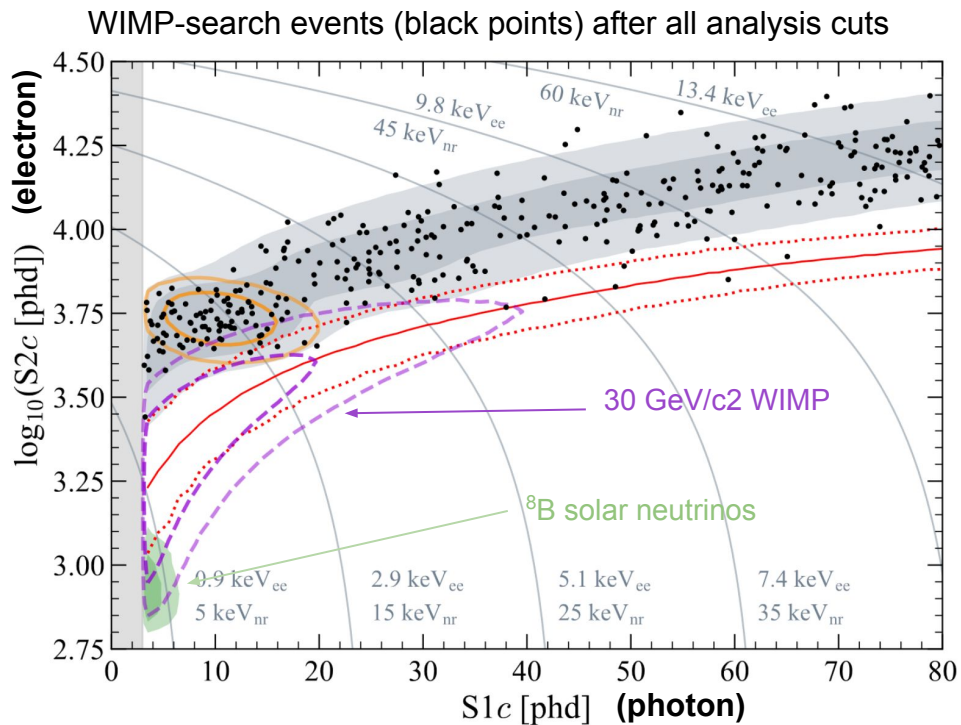
Energy deposited in  
liquid xenon < 4.6 keV

What we get:

- Demonstrate LZ's sensitivity to low energy WIMP signals.
- Calibrate the detector response for  ${}^8\text{B}$  solar neutrinos and < 10 GeV dark matter.



# WIMP-search data



Source	Expected Events	Fit Result
$^{214}\text{Pb}$	$164 \pm 35$	-
$^{212}\text{Pb}$	$18 \pm 5$	-
$^{85}\text{Kr}$	$32 \pm 5$	-
Det. ER	$1.4 \pm 0.4$	-
$\beta$ decays + Det. ER	$215 \pm 36$	$222 \pm 16$
$\nu$ ER	$27.1 \pm 1.6$	$27.2 \pm 1.6$
$^{127}\text{Xe}$	$9.2 \pm 0.8$	$9.3 \pm 0.8$
$^{124}\text{Xe}$	$5.0 \pm 1.4$	$5.2 \pm 1.4$
$^{136}\text{Xe}$	$15.1 \pm 2.4$	$15.2 \pm 2.4$
$^8\text{B}$ CE $\nu$ NS	$0.14 \pm 0.01$	$0.15 \pm 0.01$
Accidentals	$1.2 \pm 0.3$	$1.2 \pm 0.3$
Subtotal	$273 \pm 36$	$280 \pm 16$
$^{37}\text{Ar}$	[0, 288]	$52.5^{+9.6}_{-8.9}$
Detector neutrons	$0.0^{+0.2}$	$0.0^{+0.2}$
30 GeV/ $c^2$ WIMP	-	$0.0^{+0.6}$
Total	-	$333 \pm 17$

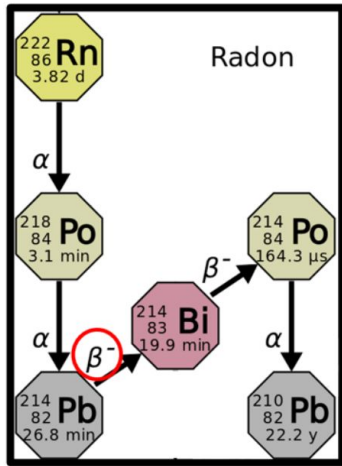
"Background Determination for the LUX-ZEPLIN (LZ) Dark Matter Experiment", arXiv:2211.17120

# Understand the enemy → background in SR1

## Radon daughters:

- Emanate from detector material
- Hard to be tagged

**Mitigation:** clean, purification, discrimination



Source	Expected Events	Fit Result
<sup>214</sup> Pb	164 ± 35	-
<sup>212</sup> Pb	18 ± 5	-
<sup>85</sup> Kr	32 ± 5	-
Det. ER	1.4 ± 0.4	-
<hr/>		
$\beta$ decays + Det. ER	215 ± 36	222 ± 16
$\nu$ ER	27.1 ± 1.6	27.2 ± 1.6
<sup>127</sup> Xe	9.2 ± 0.8	9.3 ± 0.8
<sup>124</sup> Xe	5.0 ± 1.4	5.2 ± 1.4
<sup>136</sup> Xe	15.1 ± 2.4	15.2 ± 2.4
<sup>8</sup> B CE $\nu$ NS	0.14 ± 0.01	0.15 ± 0.01
Accidentals	1.2 ± 0.3	1.2 ± 0.3
<hr/>		
Subtotal	273 ± 36	280 ± 16
<sup>37</sup> Ar	[0, 288]	52.5 <sup>+9.6</sup> <sub>-8.9</sub>
Detector neutrons	0.0 <sup>+0.2</sup>	0.0 <sup>+0.2</sup>
30 GeV/c <sup>2</sup> WIMP	-	0.0 <sup>+0.6</sup>
<hr/>		
Total	-	333 ± 17

**Neutron:** same detector response as WIMP! <0.2 in SR1

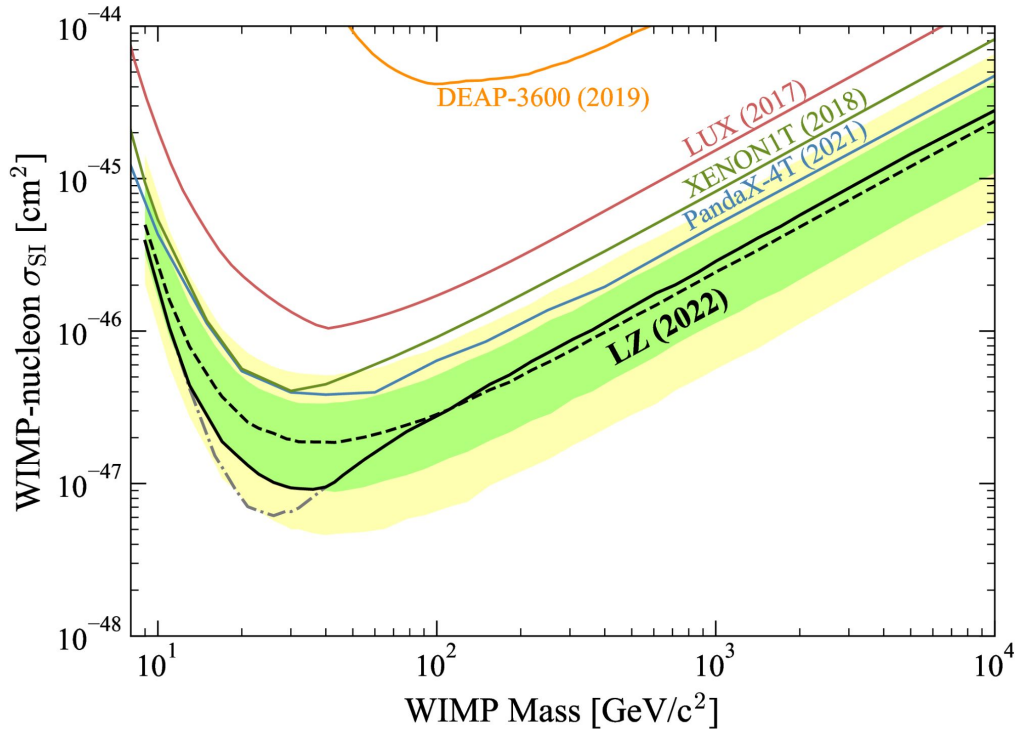
## Mitigation:

- Passive shielding
- Active neutron veto: measures 89+/-3% tagging efficiency

## Other leading backgrounds:

<sup>85</sup>Kr, <sup>37</sup>Ar, solar neutrinos

# First WIMP search result from LZ



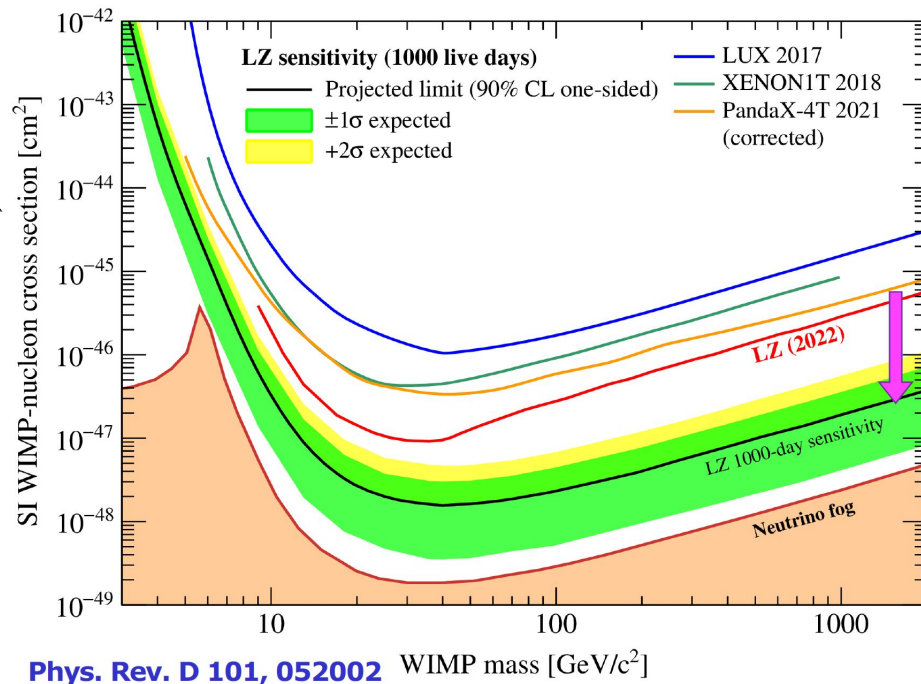
- Consistent with a background-only hypothesis,
- Setting new limits on spin-independent WIMP-nucleon for WIMP masses above  $9 \text{ GeV}/c^2$
- arXiv:2207.03764 (accepted in PRL)

- Frequentist, two-sided profile-likelihood-ratio (PLR) test statistic
- Power constrained
- Followed conventions of Eur.Phys.J.C 81 (2021) 10, 907

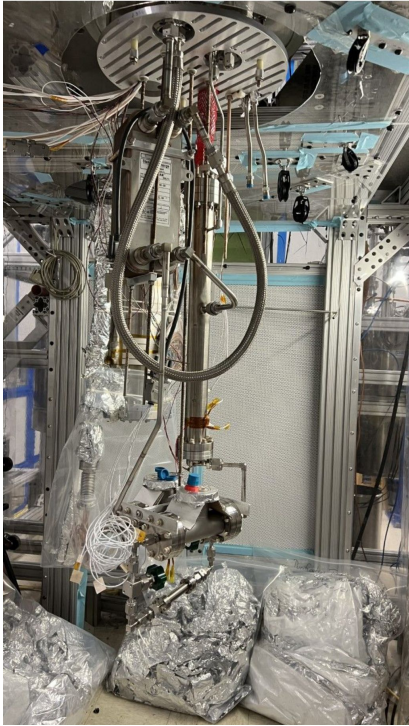


# The next step of LZ

- Next science run started, with real discovery potential !
- Broad science program:
  - Low-threshold searches
  - WIMP EFT searches
  - Non-WIMP DM candidates
  - Astrophysical neutrinos
  - $0\nu\beta\beta$  of  $^{136}\text{Xe}$
  - And more
- XLZD: LZ and XENONnT formed a consortium to build the ultimate liquid xenon dark matter detector

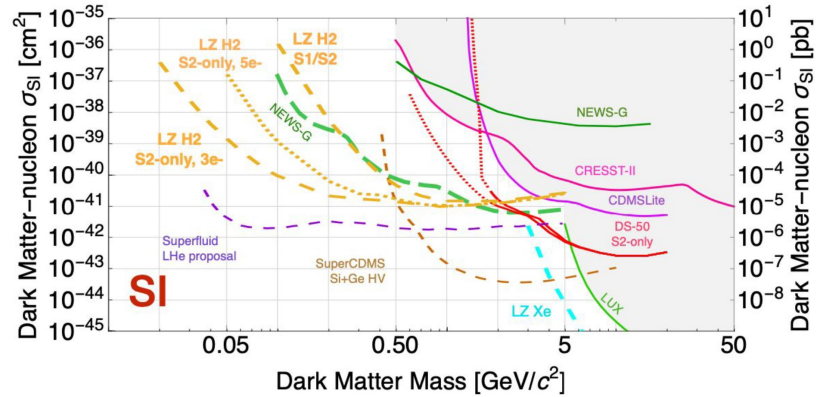


# R&D for possible upgrade

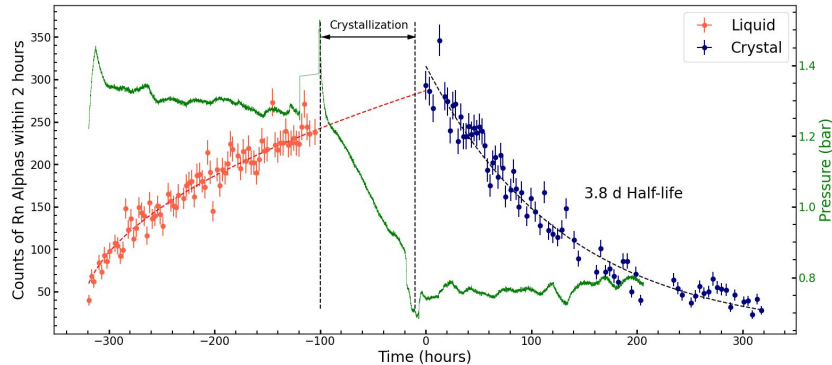


Improved Rn removal@SLAC

HydroX: H or He doped xenon for low mass detection@LBNL, SLAC and UCSB

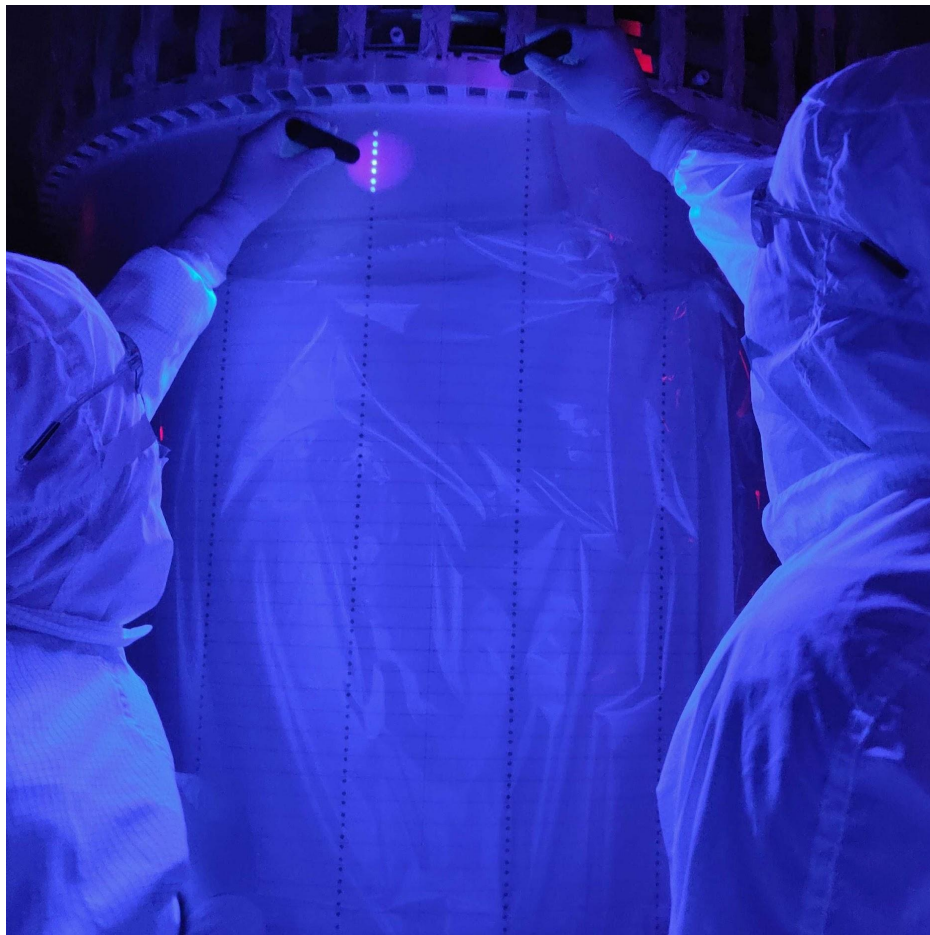


CrystaLiZe: freeze xenon to exclude radon@LBNL



# Summary

- LZ is operating and taking high quality physics data
  - All detectors are performing well
  - Backgrounds are within expectation
- With its first run, LZ has achieved world-leading WIMP sensitivity
- Broad physics program still lies ahead for LZ
- Exciting R&D in progress to improve xenon dark matter searches



# Backup

# Data Selection

**Single scatter:** 1 S1 + 1 S2, one interaction vertex

**S1 & S2 Accidentals Cuts:** remove events whose S1 or S2 is not consistent with a real event ( shape, timing)

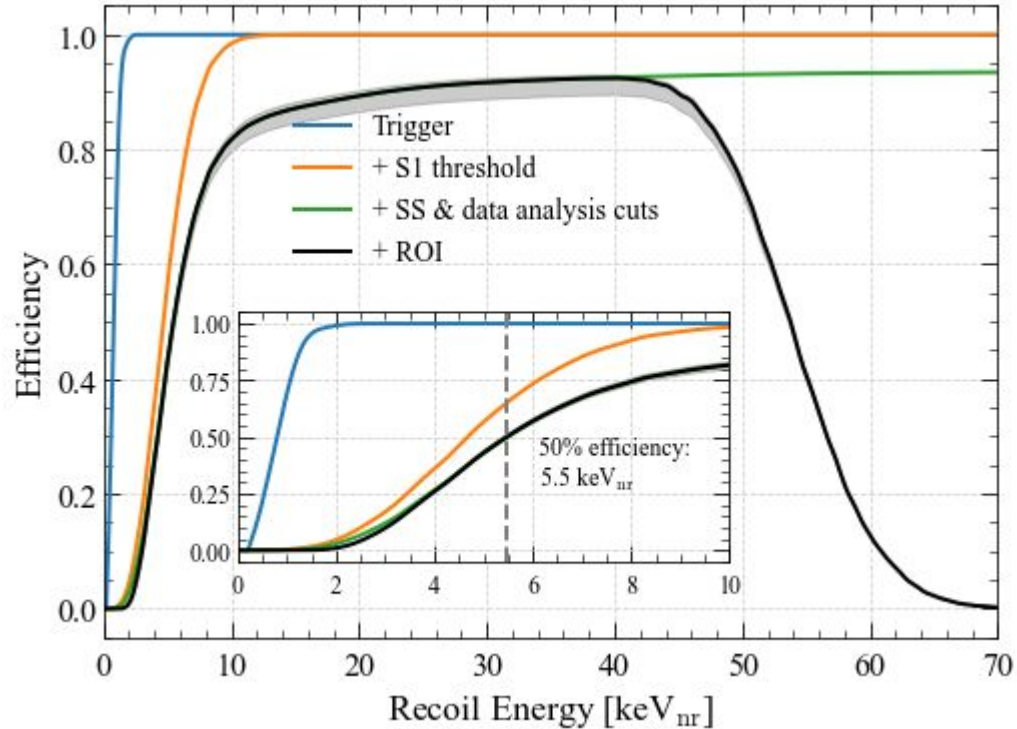
**S1 and S2 thresholds:**

- S1c > 3 phd in at least 3 PMTs, S1c < 80 phd
- S2 > 600 phd & S2c < 10000

**Fiducial Volume:** remove external & wall backgrounds

**Skin & OD Vetoes:** remove events with corresponding signal in veto detector

**Data quality:** Muon holdoff, electron/photon train holdoff, hot spot exclusion, etc.

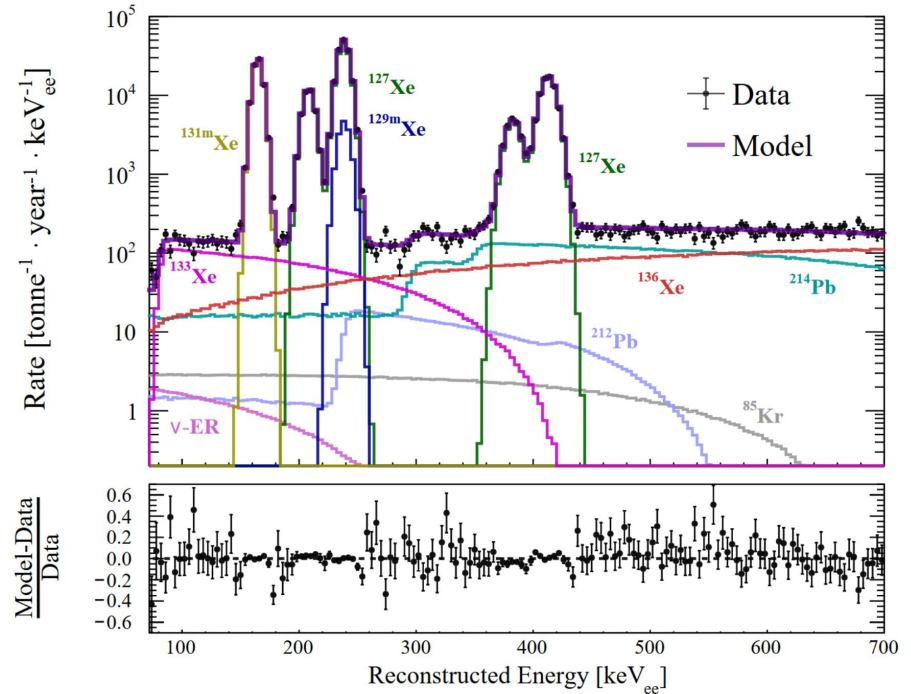


# OUTLINE

- ***Dark matter search and LZ experiment***
- LZ experiment
- Results from the first science run
- Outlook

# Background in LZ

- Dissolved beta emitters:  $^{214}\text{Pb}$  ( $^{222}\text{Rn}$  daughter),  $^{212}\text{Pb}$  ( $^{220}\text{Rn}$  daughter),  $^{85}\text{Kr}$ ,  $^{136}\text{Xe}$  (2 beta)
- Dissolved e-captures (monoenergetic x-ray/Auger cascades):  $^{127}\text{Xe}$ ,  $^{124}\text{Xe}$  (2 e-capture),  $^{37}\text{Ar}$
- Long-lived gamma emitters in detector materials:  $^{238}\text{U}$  chain,  $^{232}\text{Th}$  chain,  $^{40}\text{K}$ ,  $^{60}\text{Co}$
- Neutron emission from spontaneous fission and  $(\alpha, n)$
- Solar neutrinos:  $^8\text{B}$  (NR), pp (ER)
- Accidental coincidences.



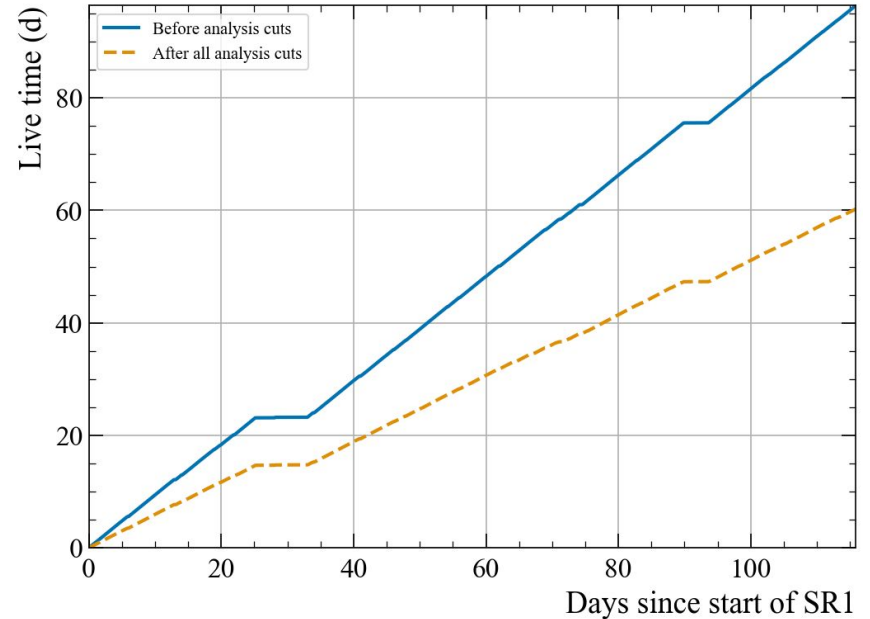
Background model

# First Science Run (SR1)

**Goal:** Demonstrate physics capability of the LZ detector

Key parameters during this run:

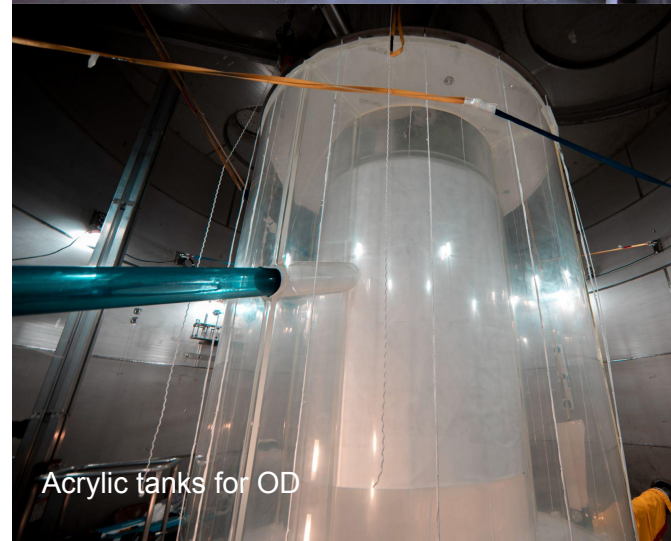
- Data taken between 23 Dec 2021 to 11 May 2022 → 60 live days for WIMP search
- 5.5 t fiducial mass
- Drift field: 193 V/cm
- Gas extraction field: 7.3 kV/cm at center
- 174.1 K (at the TPC bottom) and 1.791 bar with <0.2% fluctuation
- Continuously purified at 3.3 t/day through a hot getter system
- electron lifetime between 5000  $\mu\text{s}$  and 8000  $\mu\text{s}$ , much longer than the 951  $\mu\text{s}$  maximum drift time in the TPC





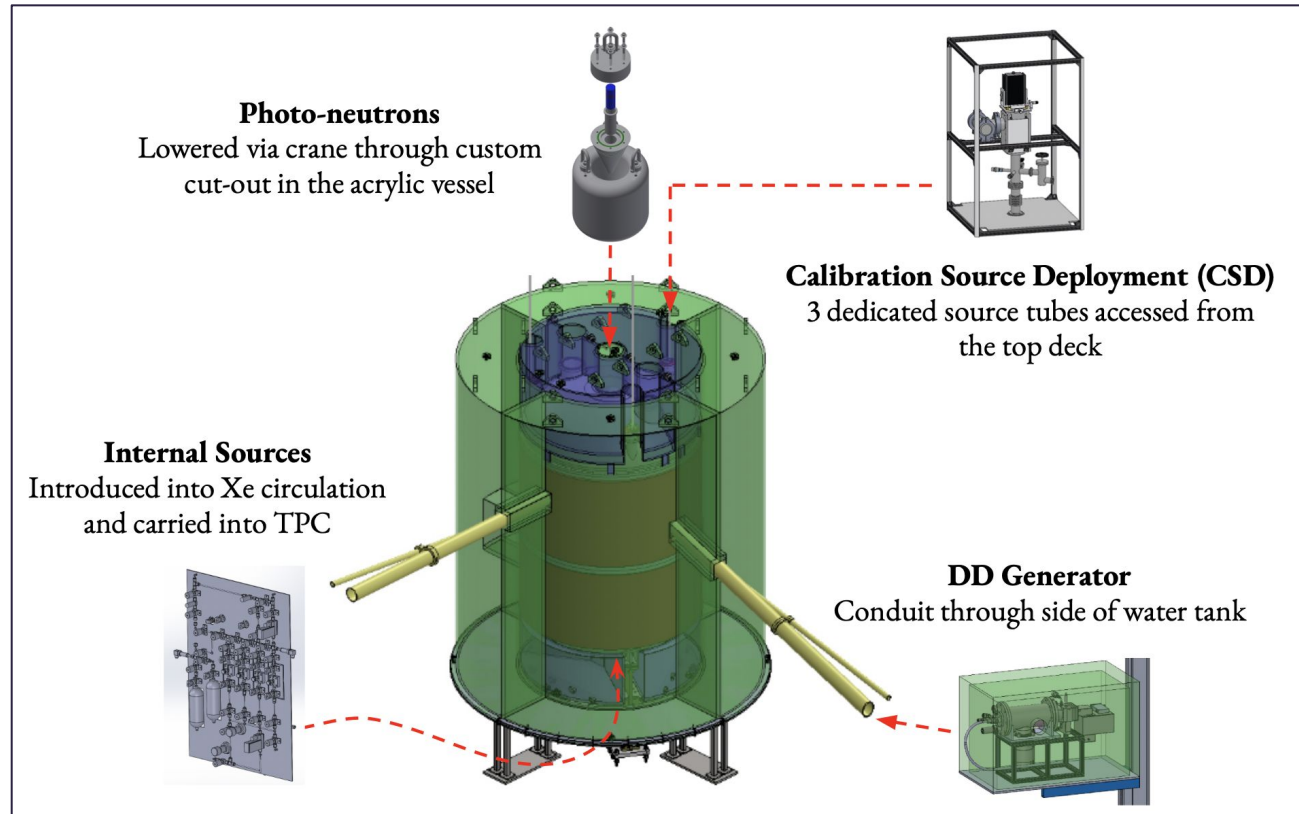
# Beat the enemy→Background mitigation

- Cleaning:
  - rigorous program of cleanliness management
  - Xenon circulation and purification
- Shielding:
  - Underground operation
  - Water tank for shielding
  - liquid xenon self-shielding
- Veto system
  - Liquid xenon skin: gamma
  - OD (gadolinium-loaded liquid scintillator (GdLS) : neutron
- Event selection:
  - 3D position reconstruction → background-light fiducial volume
  - Electronic/nuclear recoil discrimination
  - Various analysis cuts



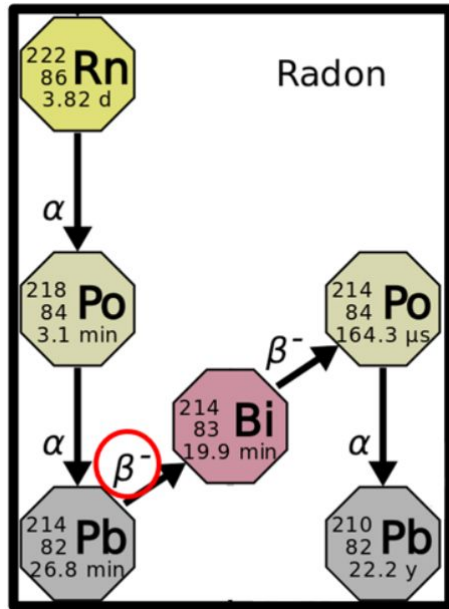
# Understand the tool → Calibration

- **DD neutron generator:** 2.45 MeV neutrons, collimated, used for: NR band, trigger efficiency, S1 cut acceptance
- **Am-Li:** continuum neutrons, isotropic, used for: Outer Detector (OD), neutron-tagging efficiency, S2 cut acceptance
- **CH<sub>3</sub>T:** continuum betas up to 18.6 keV, used for: ER band, fiducial volume, S1 cut acceptance
- **<sup>83m</sup>Kr:** monoenergetic ERs, 32.1 and 9.4 keV, used for: energy scale, xy spatial corrections
- **<sup>131m</sup>Xe:** monoenergetic ER, 164 keV, used for: energy scale, electron lifetime
- Additional background sources (e.g. alphas and cosmics), used for: energy scale, electron lifetime



# Background in SR1

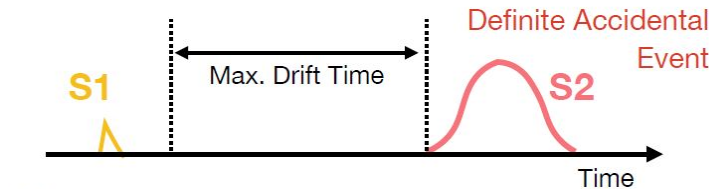
**Radon daughters:**  $^{214}\text{Pb}$ ,  $^{212}\text{Pb}$ , flat band in low energy, hard to tag  
Mitigation: clean, purification



**Accidentals:** Unrelated S1s & S2s can accidentally combine to produce single scatter events.

Analysis cuts developed to combat observed pulse/event pathologies

- >99.5% efficiency in removing accidentals
- SR1 WIMP search counts:  $1.2 \pm 0.3$



**Neutron:** same detector response as WIMP! <0.2 in SR1

Mitigation:

- Passive shielding
- Active OD veto: measures 89+-3% tagging efficiency

**Other leading backgrounds:**

$^{85}\text{Kr}$ ,  $^{37}\text{Ar}$ , solar neutrinos

# Detector response

