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DarkPACK: A modular software to compute BSM squared amplitudes for particle physics and dark matter observables

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How to verify if a given BSM model can describe some dark matter observables

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- Decay rates
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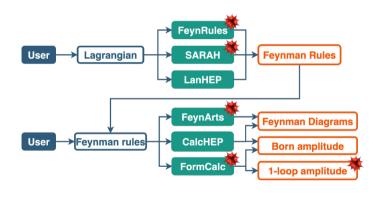
- Cross sections
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To compute

- Relic density
- Direct and indirect detection observables

Some solutions



- Many codes are required
- Several passages of input
- Mathematica dependencies

DarkPack's phylosophy

DarkPACK is conceived to have a unique and modular workflow

Unique

Lagrangian density \rightarrow amplitudes, ... \rightarrow DM observables

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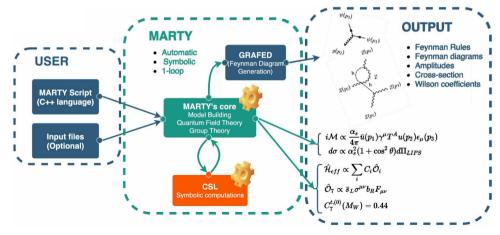
Lagrangian density \rightarrow amplitudes, ... \rightarrow DM observables

Modular

- Possibility of stopping at any point of the chain...
- ...to link it with external software
- More ease in writing custom functionalities ← Object-oriented structure

website: https://marty.in2p3.fr

manual: 2011.02478



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 - $\sum |\overline{M}|^2$, Γ
 - Wilson coefficients
 - Feyman diagrams

With MARTY the user can

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 - By defining the gauge symmetries of the model

 \rightarrow up to 1 loop level

- By defining the **fields** of the model
- By adding potential terms
- By performing **SSB** if that's in the model
- Symbolically get quantities such as
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 - Feyman diagrams
- Output those results in a numerical C++ library

DarkPACK and its documentation can be downloaded at

https://gitlab.in2p3.fr/darkpack/darkpack-public

(2211.10376 Palmiotto, Arbey, Mahmoudi)

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 - The model-specific source code for the programs (in script_mssm2to2)

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Why the MSSM?

- Numerical tests
- Performance check

Setup of DarkPACK

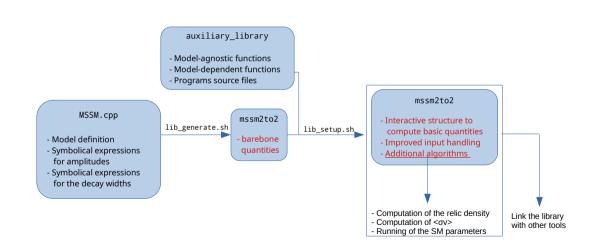
It relies on two script

- lib_generate.sh to generate the library
- lib_setup.sh to copy the files in auxiliary_library in the needed paths and to compile the final library

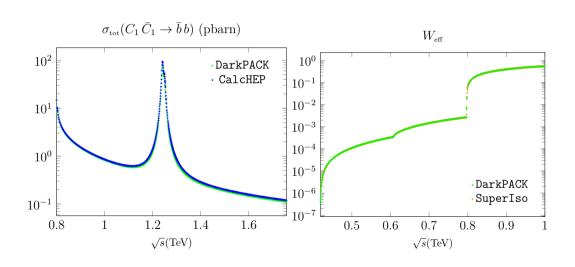
Detailed instructions on the scripts can be found in the README.md

You need to have MARTY installed, and define the environmental variable INSTALLMARTYPATH as the path where it is built

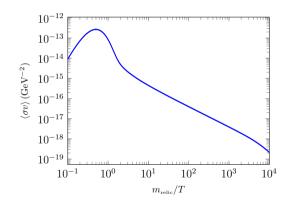
How it works



Some output



Preliminary results



$$\mathcal{L}\supset -g_{\chi}ar{\chi}\chi-rac{\phi}{\mathsf{v}}\sum_{\mathrm{fermions}}g_{\mathsf{f}}m_{\mathsf{f}}ar{\mathsf{f}}\mathsf{f}$$

- χ is a Dirac fermion and DM candidate
- $\bullet \ \phi$ is a scalar mediator

Capabilities

Generating a user-friendly library containing

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√ performance

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Performance and portability

- Avoiding global variables
- Parallelization using C++ STL

Present and next development

- Improving the model-agnostic algorithms
- Releasing new models
- Include coscattering
- Solving one Boltzmann equation for every species
- Supporting models with multiple DM candidates
- Native functions to compute direct and indirect detection observables
- Considering more general scenarios, i. e. freeze-in

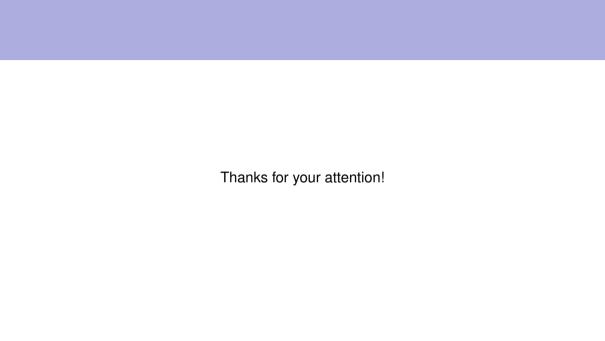
Conclusions

Today DarkPACK allows to

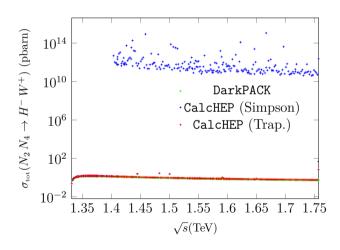
- Compute $\sum |M|^2$ and Γ at LO in any NP scenario
- Compute $\langle \sigma v \rangle$, Ωh^2 for coannihilation
- Have a library easy to link with other software
- Have a framework portable and performance-oriented
- → validated in the MSSM

Next, we will

- Validate with a new model and relese it
- Follow the development roadmap
- use DarkPack to see if specific NP models can help to explain DM observables



Simpson rule vs trapezoidal rule pt. 1



Simpson rule vs trapezoidal rule pt. 2

