

Energy of anti-de Sitter black holes in odd-dimensional Quadratic Curvature Gravity

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in collaboration with: O. Miskovic and R. Olea

Yoel Parra-Cisterna
Instituto de Física PUCV

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Outline

1 Motivation

- AdS/CFT Correspondence
- AdS Gravity
- Kounterterms in AdS Gravity

2 Quadratic Curvature Gravity

- Quadratic Curvature Gravity Action
- Renormalized Action
- Energy for Static Black Holes

3 Discussions

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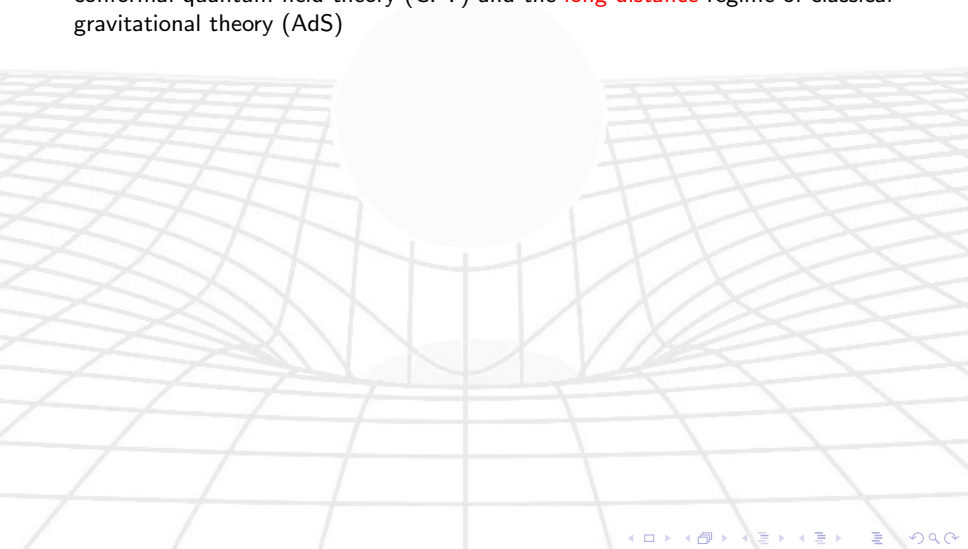
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- In particular, the CFT **stress tensor**

$$\langle T_{ij} \rangle_{\text{CFT}} = \frac{2}{\sqrt{-g(0)}} \frac{\delta I_{\text{gr}}(g_{\mu\nu})}{\delta g_{ij}(0)},$$

which satisfy Ward identities $\nabla^i \langle T_{ij} \rangle_{\text{CFT}} = 0$ and $\langle T^i_i \rangle_{\text{CFT}} = \mathcal{A}$ (**Weyl anomaly**)

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- However, $I_{\text{gr}}(\text{AdS})|_{\text{on-shell}} \rightarrow \infty \implies$ Holographic Renormalization (**Infrared Renormalization**)

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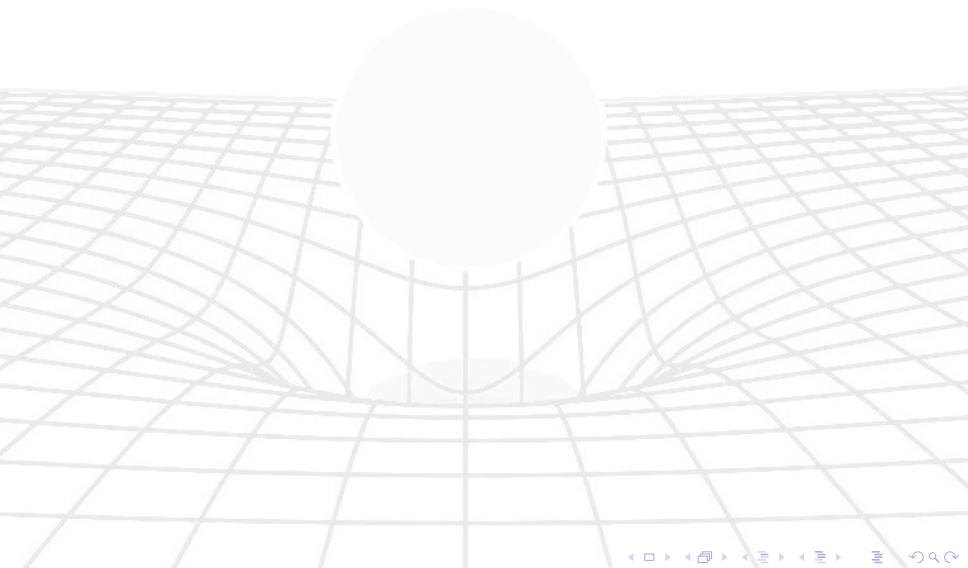
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- In normal coordinates $x^\mu = (z, x^i)$, $ds^2 = N^2(z)dz^2 + h_{ij}(z, x^i)dx^i dx^j$



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with $\Lambda = -\frac{d(d-1)}{2\ell^2}$ and K is the trace of **extrinsic curvature** $K_{ij} = -\frac{1}{2}n^\mu \partial_\mu h_{ij}$, and

Counterterms: $L_{\text{ct}} = L_{\text{ct}}(h_{ij}, \mathcal{R}^i{}_{jkl}, D_m \mathcal{R}^i{}_{jkl}) \iff$ **Holographic Renormalization**

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- Holographic information, in FG coordinates

$$T^{ij}[g_{(0)}] = \lim_{z \rightarrow 0} \left(\frac{1}{z^{d-2}} T_{\text{ren}}^{ij}[h] \right) \equiv \langle T_{ij} \rangle_{\text{CFT}}$$

Holographic stress-tensor

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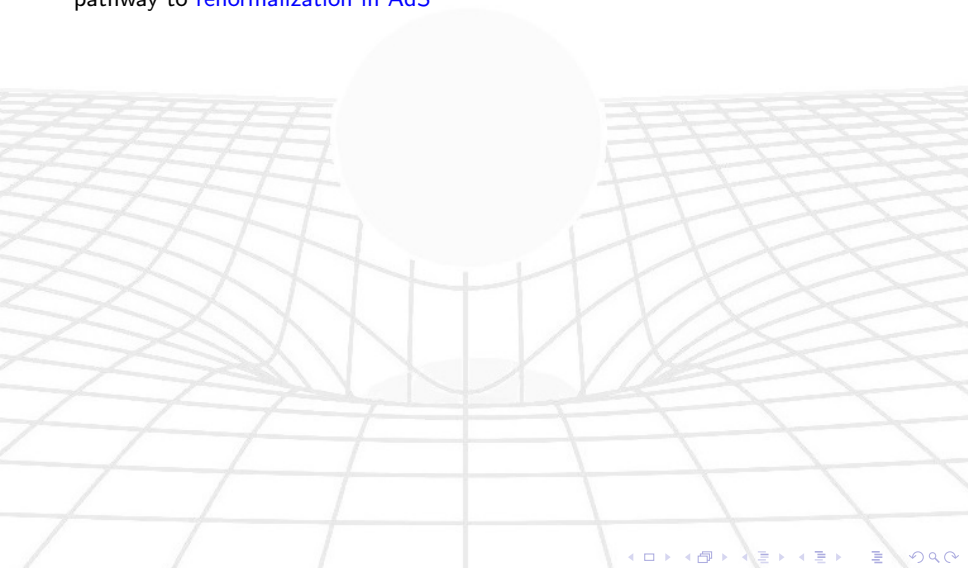
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$$\tilde{I}_{\text{ren}} = I_{\text{bulk}} + c_d \int_{\partial M} d^d x \sqrt{-h} B_d(f(h), K)$$

\implies Well posed variational principle for **Dirichlet boundary conditions** only for $g_{(0)ij}$

$$\delta \tilde{I}_{\text{ren}} = \delta \tilde{I}_{\text{ren}}(\delta h, \delta K) = \frac{1}{2} \int_{\partial M} d^d x \sqrt{-g_{(0)}} \tau^{ij} \delta g_{(0)ij}, \quad \text{Action } \mathbf{holographically\ finite\ !!!}$$

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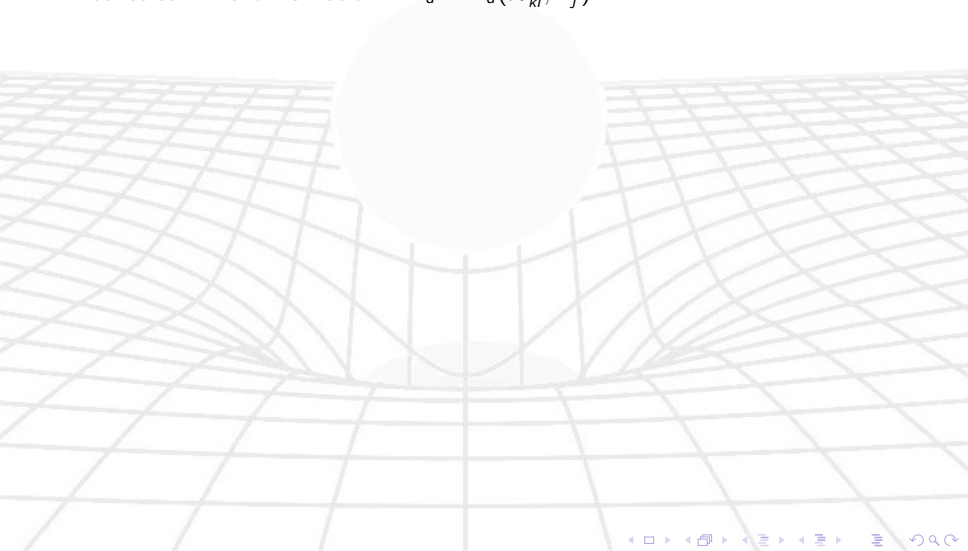
- Another observation: in $d + 1 = 5$

$$T_{\text{ren}j}{}^i = - \left(\mathcal{R}_j^i - K_j^i K + K_k^i K_j^k + \frac{3}{\ell^2} \delta_j^i \right) + \Delta_j^i,$$

which can be truncated as $T_{\text{ren}j}{}^i = \pi_j^i + \mathcal{O}(K^2) \implies B_d$ **non-linear** in K .

Kounterterms in AdS Gravity

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- **Kounterterm Renormalization:** $B_d = B_d(\mathcal{R}_{kl}^j, K_j^i)$
- In **even** dimensions $d + 1 = 2n$

$$B_{2n-1} = 2n \int_0^1 dt \delta_{i_1 \dots i_{2n-1}}^{j_1 \dots j_{2n-1}} K_{j_1}^{i_1} \left(\frac{1}{2} \mathcal{R}_{j_2 j_3}^{i_2 i_3} - t^2 K_{j_2}^{i_2} K_{j_3}^{i_3} \right) \times \dots$$

$$\dots \times \left(\frac{1}{2} \mathcal{R}_{j_{2n-2} j_{2n-1}}^{i_{2n-2} i_{2n-1}} - t^2 K_{j_{2n-2}}^{i_{2n-2}} K_{j_{2n-1}}^{i_{2n-1}} \right).$$

In this case, the Kounterterms are proportional to the n -th Chern form.

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- In **odd** dimensions $d + 1 = 2n + 1$

$$B_{2n} = 2n \int_0^1 dt \int_0^t ds \delta_{i_1 \dots i_{2n}}^{j_1 \dots j_{2n}} K_{j_1}^{i_1} \delta_{j_2}^{i_2} \left(\frac{1}{2} \mathcal{R}_{j_3 j_4}^{i_3 i_4} - t^2 K_{j_3}^{i_3} K_{j_4}^{i_4} + \frac{s^2}{\ell^2} \delta_{j_3}^{i_3} \delta_{j_4}^{i_4} \right) \times \dots$$

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$$I_{\text{QCG}} = \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left[\frac{1}{\kappa} (R - 2\Lambda_0) + \alpha R_{\mu\nu} R^{\mu\nu} + \beta R^2 + \gamma GB \right]$$

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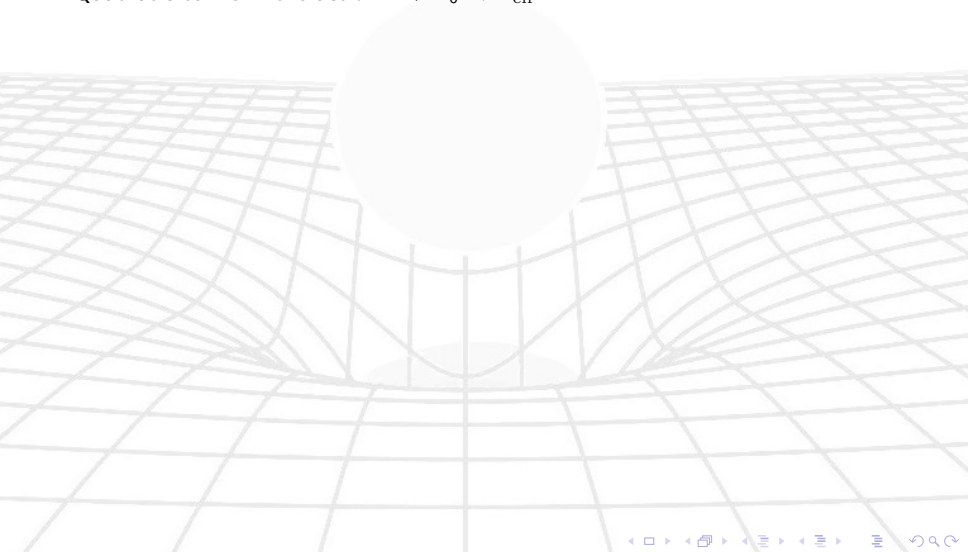
Einstein's Tensor: $G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R + \Lambda_0 g_{\mu\nu}$

Lanczos tensor: $H_{\mu\nu} = -\frac{1}{8} g_{\mu\lambda} \delta^{\lambda\mu_1 \dots \mu_4}_{\nu\nu_1 \dots \nu_4} R^{\nu_1 \nu_2}_{\mu_1 \mu_2} R^{\nu_3 \nu_4}_{\mu_3 \mu_4}$.

Quadratic terms: $P_{\mu\nu} = 2\beta R \left(R_{\mu\nu} - \frac{1}{4} g_{\mu\nu} R \right) + (\alpha + 2\beta) (g_{\mu\nu} \square - \nabla_{\mu} \nabla_{\nu}) R$
 $+ \alpha \square G_{\mu\nu} + 2\alpha \left(R_{\mu\sigma\nu\lambda} - \frac{1}{4} g_{\mu\nu} R_{\sigma\lambda} \right) R^{\sigma\lambda}$

Effective cosmological constant

- Quadratic terms in the action $\implies \Lambda_0 \rightarrow \Lambda_{\text{eff}}$.



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- (e.o.m.) $\implies \Lambda_{\text{eff}} = \Lambda_{\text{eff}}(\Lambda_0, \alpha, \beta, \gamma)$

$$\frac{1}{\Lambda_{\text{eff}}^{\pm}} = \frac{1}{2\Lambda_0} \left[1 \pm \sqrt{1 + 8\kappa\Lambda_0 \frac{d-3}{d-1} \left(\frac{\alpha + (d+1)\beta}{d-1} + \gamma \frac{d-2}{d} \right)} \right],$$

Einstein branch: Λ_{eff}^+ ,

non-Einstein branch: Λ_{eff}^-

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- Asymptotically anti-de Sitter (AAdS) spacetimes:** $\Lambda_{\text{eff}} = -\frac{d(d-1)}{2\ell_{\text{eff}}^2}$,

$$\frac{1}{\ell_0^2} = \frac{1}{\ell_{\text{eff}}^2} - \frac{\kappa d(d-3)}{\ell_{\text{eff}}^4} \left(\frac{\alpha + (d+1)\beta}{d-1} + \gamma \frac{d-2}{d} \right)$$

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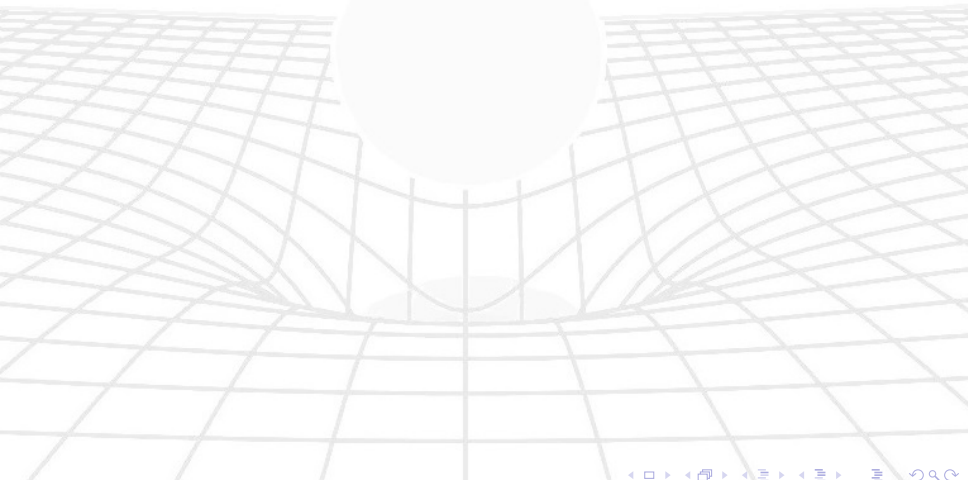
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- In order to determine c_d , we use normal coordinates $x^\mu = (z, x^i)$, the bulk boundary term becomes

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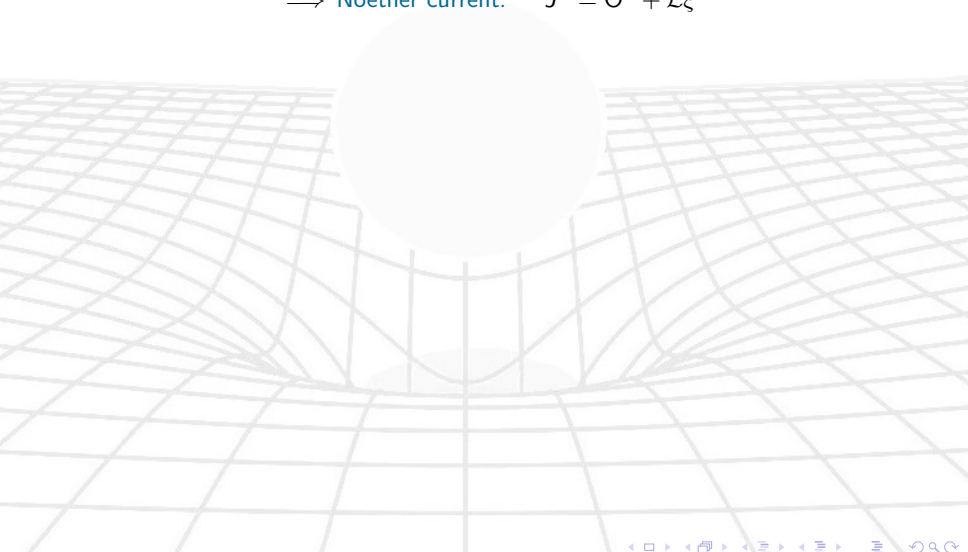
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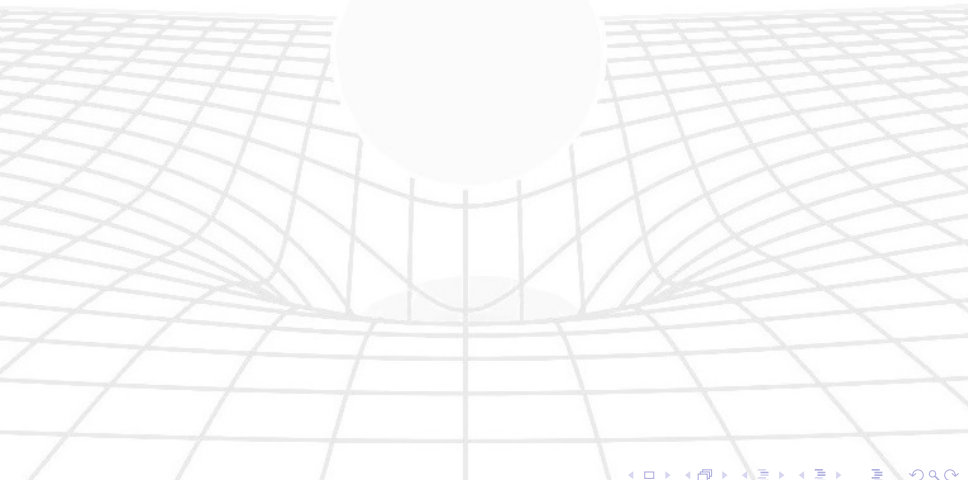


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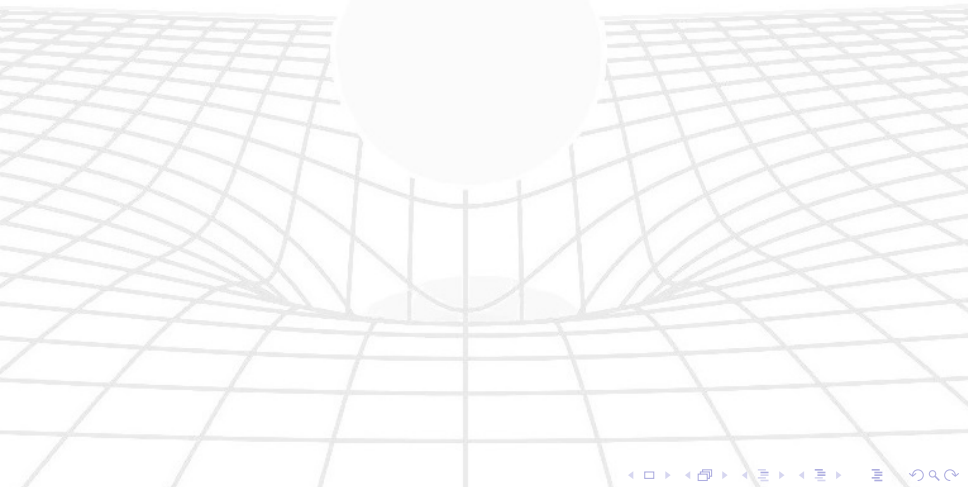
$$Q[\xi] = \int_{\Sigma} d^{d-1}x \sqrt{\sigma} u_i q_j^i \xi^j,$$

where $u_i u^i = -1$ and $\xi = \xi^i \partial_i$ are asymptotics isometries.

Renormalized conserved charge in odd dimensions

- In the **odd** dimensional case

$$Q[\xi] = \int_{\Sigma} d^{d-1}x \sqrt{\sigma} u_i (q_j^i + q_{(0)j}^i) \xi^j = q[\xi] + \underbrace{q_{(0)}[\xi]}_{\sim \text{AdS vacuum energy}}$$



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$$q_i^j = \frac{1}{2^{n-2}} \delta_{k_1 \dots i_{2n-1}}^{j_1 \dots j_{2n-1}} K_i^k \delta_{j_1}^{i_1} \left[nc_{2n} \int_0^1 du \left(R_{j_2 j_3}^{i_2 i_3} + \frac{u^2}{\ell_{\text{eff}}^2} \delta_{j_2 j_3}^{i_2 i_3} \right) \dots \left(R_{j_{2n-2} j_{2n-1}}^{i_{2n-2} i_{2n-1}} + \frac{u^2}{\ell_{\text{eff}}^2} \delta_{j_{2n-2} j_{2n-1}}^{i_{2n-2} i_{2n-1}} \right) \right. \\ \left. - \frac{1}{(2n-1)!} \left(\frac{1}{\kappa} \delta_{j_2 j_3}^{i_2 i_3} + \alpha R_z^z \delta_{j_2 j_3}^{i_2 i_3} + 2\beta R \delta_{j_2 j_3}^{i_2 i_3} + 2(2n-1)(2n-2)\gamma R_{j_2 j_3}^{i_2 i_3} \right) \delta_{j_4 j_5}^{i_4 i_5} \dots \delta_{j_{2n-2} j_{2n-1}}^{i_{2n-2} i_{2n-1}} \right] \\ - 2N \left[\alpha \left(\nabla^z R_i^j - \nabla^j R_i^z + \nabla^z R_z^z \delta_i^j + \nabla^k R_k^z \delta_i^j \right) + 2\beta \nabla^z R \delta_i^j \right] - 2\alpha K_k^j R_i^k \\ - N \left[\alpha \left(\nabla^k R_k^z \delta_i^j - \nabla^j R_i^z \right) + 4\gamma \left(\nabla^j R_i^z - \delta_i^j \nabla^k R_k^z + \nabla^k R_{ik}^z \right) \right],$$

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- AdS/CFT Correspondence
- AdS Gravity
- Counterterms in AdS Gravity

2 Quadratic Curvature Gravity

- Quadratic Curvature Gravity Action
- Renormalized Action
- **Energy for Static Black Holes**

3 Discussions

Topological Black Holes

- Static ansatz, in coordinates $x^\mu = (t, r, \varphi^n)$,

$$ds^2 = -f^2(r)dt^2 + f^{-2}(r)dr^2 + r^2\gamma_{mn}(\varphi) d\varphi^m d\varphi^n, \quad \varphi^m \in \Sigma^{d-1},$$

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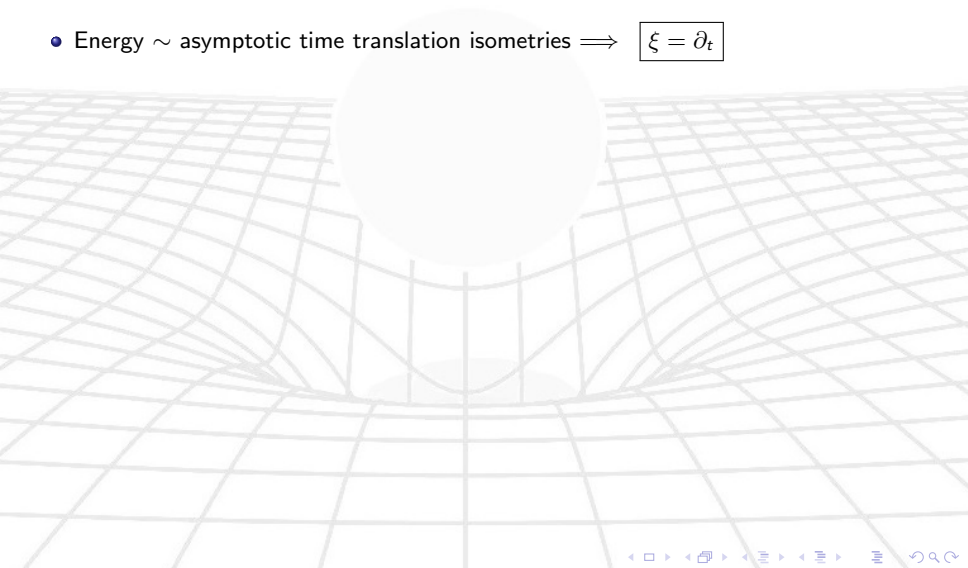
- In AAdS in QCG, one can assume an asymptotic behavior of the form

$$f^2(r) \approx k - \left(\frac{r_0}{r}\right)^{d-2} + \frac{r^2}{\ell_{\text{eff}}^2} + \dots,$$

where the constant r_0^{d-2} is related with the mass of the TBH.

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- And the **vacuum energy** $q_{(0)}[\partial_t] = E_{\text{vac}}$,

$$E_{\text{vac}} = k^n (2n-1)! c_{2n} \text{Vol}(\Sigma^{2n-1})$$

$$= (-k)^n \frac{2(2n-1)!!^2}{(2n)!} \frac{\text{Vol}(\Sigma^{2n-1})}{\kappa} \ell_{\text{eff}}^{2n-2} \left[1 - \frac{4n\kappa}{\ell_{\text{eff}}^2} \left(\alpha + (2n+1)\beta + \gamma \frac{(2n-1)(2n-2)}{2n} \right) \right]$$

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- Unlike Lovelock, in QCG case, $\Xi_d = 0$ is not the degenerate vacuum condition, then we can have a well-defined charge, even in the degenerate case

$$\Xi_d - \Delta' = -4\kappa d \frac{\alpha + (d+1)\beta}{(d-1)\ell_{\text{eff}}^2}.$$

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Muchas Gracias a Todos y Todas !!!