





Impact of non-perturbative Effects in Simplified *t*-Channel Dark Matter Models

in collaboration with

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Motivation I



 \Rightarrow How impactful are non-perturbative Effects for experimental exclusion limits







What is new?

- Bound State Formation (BSF) effects on the relic density in a *realistic* model of colored coannihilation
- Study of the interplay of a large variety of experimental searches, considering the SE and BSF.
- We point out the possibilities of Bound state searches at the LHC
- Correct implemenation of BSF in micrOMEGAs (including three-gauge boson vertex contribution and an estimate of bound state three-body decays)







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- Bound State Formation (BSF) effects on the relic density in a *realistic* model of colored coannihilation
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- We point out the possibilities of Bound state searches at the LHC
- Correct implemenation of BSF in micrOMEGAs (including three-gauge boson vertex contribution and an estimate of bound state three-body decays)
- \rightarrow A flat correction factor for non-perturbative effects is unapplicable
- \rightarrow Corrections on the exclusion limits can be as large as $\mathcal{O}\left(100\%\right)$
- \rightarrow Bound State searches close gap between prompt and long-lived searches!







Simplified t-Channel Dark Matter

Universal framework for t-channel DM models [Arina, Fuks, Mantani (2020)]

S3M-uR t-channel Dark Matter

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{kin,BSM} + g_{DM}\overline{\chi}(u_R)_i(X^{\dagger})_i + h.c.$$

$$\chi = (\mathbf{1}, \mathbf{1})_0 \qquad X_i = (\mathbf{3}, \mathbf{1})_{2/3}$$







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$$egin{aligned} \mathcal{L} &= \mathcal{L}_{\mathsf{SM}} + \mathcal{L}_{\mathsf{kin},\mathsf{BSM}} + g_{\mathsf{DM}} \overline{\chi}(u_R)_i (X^\dagger)_i + h.c. \ \chi &= (\mathbf{1},\mathbf{1})_0 \qquad X_i = (\mathbf{3},\mathbf{1})_{\mathtt{2/3}} \end{aligned}$$

- Discrete \mathcal{Z}_2 : SM fields even, dark sector fields odd
- Majorana fermion DM χ
- 3 generations of mediators X_i

Parameters:
$$(m_{\chi} = m_{\text{DM}}, \Delta m = m_X - m_{\text{DM}}, g_{\text{DM}})$$







Dark Matter Freeze-Out

Assumptions during DM freeze-out:

- Dark sector in *kinetic* eq. with the SM.
- Dark sector particles in *chemical* eq. with themselves.

Coannihilation

$$\frac{dn}{dt} + 3Hn = -\langle \sigma_{\text{eff}} \mathbf{v} \rangle \left(n^2 - (n^{\text{eq}})^2 \right)$$
$$\langle \sigma_{\text{eff}} \mathbf{v} \rangle = \sum_{i,j} \langle \sigma_{ij} \mathbf{v}_{ij} \rangle \frac{n_i^{\text{eq}}}{n^{\text{eq}}} \frac{n_j^{\text{eq}}}{n^{\text{eq}}}$$

$$n = \sum_i n_i$$
 and $i, j = \{\chi, X_1, X_2, X_3\}$

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n-gluon exchanges contribute with $\left(\frac{\alpha}{v}\right)^n$ for $\alpha \sim v$

- \rightarrow Resummation required since $\alpha_{\rm s} \sim {\it v} \sim 0.1$
- ightarrow Reduces to Schrödinger Equation for $v \ll 1$. For details [Petraki,Postma,Wiechers(2015)]

Figure from Talk by J.Harz @ DM Working Group































SE vs BSF

Modified Coannihilation [Ellis,Luo,Olive(2015)]

$$\langle \sigma_{\text{eff}} \mathbf{v} \rangle = \sum_{i,j \in \{\chi,X\}} \langle S\left(\alpha/v_{ij}\right) \cdot \sigma_{ij} \mathbf{v}_{ij} \rangle \frac{n_i^{\text{eq}}}{n^{\text{eq}}} \frac{n_j^{\text{eq}}}{n^{\text{eq}}} + \langle \sigma_{\text{BSF}} \mathbf{v} \rangle_{\text{eff}} \left(\frac{n_X^{\text{eq}}}{n^{\text{eq}}}\right)^2$$

$\langle \sigma_{\rm eff} \mathbf{v} \rangle$	Sommerfeld Effect	Bound State Formation
$g_{ extsf{DM}} \gg g_s$	_	0
$g_{ extsf{DM}} \ll g_s$	+	++

 \rightarrow No flat factor







BSF vs SE

$$\langle \sigma_{\text{eff}} \mathbf{v} \rangle = \sum_{i,j \in \{\chi, X\}} \langle S\left(\alpha / v_{ij}\right) \cdot \sigma_{ij} \mathbf{v}_{ij} \rangle \frac{n_i^{\text{eq}}}{n^{\text{eq}}} \frac{n_j^{\text{eq}}}{n^{\text{eq}}} + \langle \sigma_{\text{BSF}} \mathbf{v} \rangle_{\text{eff}} \left(\frac{n_X^{\text{eq}}}{n^{\text{eq}}}\right)^2$$

Sommerfeld Effect:

- Has an effect independently of the hierarchy between g_{DM} and g_s
- Tends to lower $\langle \sigma_{
 m eff} {\it v}
 angle$ for $g_{
 m DM} > g_s$
- Tends to increase $\langle \sigma_{
 m eff} v
 angle$ for $g_{
 m DM} < g_s$

Bound State Formation:

- BSF is purely mediated by $g_s
 ightarrow$ less important for $g_{ extsf{DM}} \gg g_s$
- Always increases $\langle \sigma_{\rm eff} {\it v} \rangle$
- For $g_{\text{DM}} < g_s$ more sizable than the Sommerfeld effect.







Calculation of the Relic Density

We adjusted micrOMEGAs 5.2.7 such that

- the Sommerfeld Effect is included for colored scalars up to the adjoint representation
- Bound State effects are included for colored scalars up to the adjoint representation

Determine $g_{DM,0}$ for each data point $(m_{DM}, \Delta m)$ such that DM is *not* overproduced.









Figure from [MB,Copello,Harz,Mohan,Sengupta(2022)]









 \rightarrow Bound State Formation increases the area where the strong interaction deplete relic density significantly!







Direct Detection Constraints



- SI stronger than SD for Δm < m_{DM}
- Strong constraints on the coannihilating area
- Inclusion of BSF opens up parameter space in this region







Prompt Collider Searches mono-jet + ETmiss search by ATLAS multi-jets + ETmiss search by CMS [arXiv:1711.03301] [arXiv:1704.07781] perturbative only +Sommerfeld Effect +Bound State Formation 10^{2} 10³ 10^{4} 10^{2} 10³ 10^{4} 102 10^{3} 10^{4} Control Control Control a start of the second second a a sa contra de la contra de 103 103 103 103 10 103 Unitarity Unitarity Initarit Atlas 139 fb tlas 139 fb tias 139 fb⁻ ∆m[GeV] Dm[GeV] ∆m[GeV] 10 10^{2} 10² 10² 10² SI XENON1T SI XENON1T SLXENON1T SD Pico-60 SD Pico-60 SD Pico-60 10^{1} 101 10^{1} 10^{1} 0 101 Underabundant DV Underabundant DM Underabundant DM 10^{0} 1.00 100 109 100 104 10^{4} 103 10^{4} 10^{3} 10^{3} m_{DM}[GeV] m_{DM}[GeV] m_{DM}[GeV]

- mainly constrains larger Δm
- Thus non-perturbative effects are mild/absent.







How to constrain the "gray" area ?

Freeze-out leads to underabundant $\text{DM} \rightarrow \text{correct}$ abundance requires alternative production

Out-of-chemical equilibrium estimate

$$rac{\Gamma_{X\leftrightarrow\chi}\left(ilde{g}_{ ext{DM}}
ight)}{H} \lesssim 1 \,, \, ext{at freeze-out}$$

 $ightarrow ilde{g}_{ ext{DM}} \lesssim \sqrt{rac{m_{ ext{DM}}}{GeV}} \left(10^{-9} + 6.8 \cdot 10^{-11} rac{\Delta m}{m_{ ext{DM}}}
ight)$

For $g_{\rm DM} < \tilde{g}_{\rm DM}$ DM production is non-thermal

Long-Lived-Particle (LLP) searches constrain large lifetimes $\rightarrow g_{\text{DM}} \ge g_{\text{DM}}^{\text{LLP}}$

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 \rightarrow Region can not be fully tested when including non-perturbative effects

 \rightarrow A precise treatment, assuming conversion driven freeze-out, has been analyzed in $_{[\text{Garny,Heisig}\,(2021)]}$







Bound State Formation at the LHC

Production Cross Section[Martin(2008)]

$$\sigma(pp \to \mathcal{B}(XX^{\dagger})) = \frac{\pi^2}{8m_B^3} \Gamma(\mathcal{B}(XX^{\dagger}) \to gg) \mathcal{P}_{gg}\left(\frac{m_B}{13 \text{ TeV}}\right)$$

 \rightarrow try to observe the bound state resonance in $\gamma\gamma$ final state. <code>ATLAS (2017)</code>

Efficient for all g_{DM} small enough such that $\Gamma_X < E_B$, roughly speaking $g_{DM} \leq g_s$.











Limits at 37 $\rm fb^{-1}$ relatively weak in mass (\sim 300 $\rm GeV)$ But huge potential: Closes the gap between prompt and LLP searches







Expected Future Limits



- Highly testable: Parameter space almost completely probed
- Remember: HSCP not a strict exclusion here (BSF@LHC is!)
- Bound State effects enlarge the area still necessary to test







When is BSF relevant?

No coannihilation required!

ightarrow Expect potentially large non-perturbative effects for $lpha \sim {\it v} \sim$ 0.1

Case I: Massless/light mediator (for instance colored annihilation)

$$\sigma_{\rm ann} \sim \frac{\alpha^2}{m^2} \quad \xrightarrow{\Omega_{\rm DM} \sim 1/\sigma_{\rm ann}} \quad \alpha \sim 0.1 \frac{m}{{
m TeV}}$$







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Case II: Massive mediator with mass *M* (Yukawa potential)

$$\sigma_{\rm ann} \sim \alpha^2 \frac{m^2}{\left(m^2 + M^2\right)^2} \quad \xrightarrow{\Omega_{\rm DM} \sim 1/\sigma_{\rm ann}} \quad \alpha \sim 0.1 \left(\frac{m}{\rm TeV}\right) \left(1 + \frac{M^2}{m^2}\right)$$

But Yukawa suppression sizable if $\alpha m \lesssim M$







Conclusion

- Non-perturbative Effects can increase or decrease the annihilation cross section of DM
 - \rightarrow Cannot be handled by a flat correction factor!
- Non-perturbative Effects are non-neglible in scenarios of colored coannihilation and open up small mass parameter space:

Viable Parameter space shifts from $(m_{DM}, \Delta m) < (1 TeV, 30 GeV)$ to (1.4 TeV, 40 GeV) (Sommerfeld Effect) and (2.4 TeV, 50 GeV) (Bound State Formation)

- \rightarrow Sommerfeld Effect alone not a good approximation!
- Bound State searches at colliders close the gap in "coupling space" between prompt and long-lived-particle searches







Annihilation Channels

NPE = Non-Perturbative Effects



 \Rightarrow No NPE \Rightarrow No NPE

 \Rightarrow Subject to NPE







Color Decomposition

Process: $(X_1)_{\mathbf{R_1}} + (X_2)_{\mathbf{R_2}} \rightarrow SM + SM$

Color Potential

$$V(r) = -\frac{\alpha_s}{2r} \left[C_2(\mathbf{R_1}) + C_2(\mathbf{R_2}) - C_2(\mathbf{R}) \right] = -\frac{\alpha_{\text{eff},[\mathbf{R}]}}{r}$$







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Color Configurations

$$\begin{aligned} \mathbf{3} \times \overline{\mathbf{3}} &= \mathbf{1} + \mathbf{8} \to \alpha_{\text{eff},[\mathbf{1}]} = \frac{4}{3}, \ \alpha_{\text{eff},[\mathbf{8}]} = -\frac{1}{6} \\ \mathbf{3} \times \mathbf{3} &= \overline{\mathbf{3}} + \mathbf{6} \to \alpha_{\text{eff},[\overline{\mathbf{3}}]} = \frac{2}{3}, \ \alpha_{\text{eff},[\mathbf{6}]} = -\frac{1}{3} \end{aligned}$$

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n-gluon exchanges contribute with $\left(\frac{\alpha}{\nu}\right)^n$ for $\alpha \sim \nu$

Sommerfeld Effect

$$\sigma(X_1X_2 \to SMSM) = S\left(\frac{\alpha_{\text{eff}}}{v}\right)\sigma_{\text{pert.}}$$

Sommerfeld Factor

$$S\left(\frac{\alpha_{\text{eff}}}{v}\right) = \begin{cases} 1 & , \text{ if } |\frac{\alpha_{\text{eff}}}{v}| \ll 1, \\ \frac{\alpha_{\text{eff}}}{v} & , \text{ if } |\frac{\alpha_{\text{eff}}}{v}| \gg 1 \land \alpha_{\text{eff}} > 0 \\ \exp\left(2\pi\frac{\alpha_{\text{eff}}}{v}\right) & , \text{ if } |\frac{\alpha_{\text{eff}}}{v}| \gg 1 \land \alpha_{\text{eff}} < 0 \end{cases}$$

Figure from Talk by J.Harz @ DM Working Group

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Annihilation Channels II



As a rule of thumb, we find:

 $g_{\text{DM}} > g_s \rightarrow$ Sommerfeld effect reduces annihilation cross section $g_{\text{DM}} < g_s \rightarrow$ Sommerfeld effect increases annihilation cross section







SE vs BSF

Modified Coannihilation [Ellis, Luo, Olive(2015)]

$$\left\langle \sigma_{\mathsf{eff}} \mathbf{V} \right\rangle = \sum_{i,j \in \{\chi, X\}} \left\langle S\left(\alpha / \mathsf{V}_{ij}\right) \cdot \sigma_{ij} \mathbf{V}_{ij} \right\rangle \frac{n_i^{\mathsf{eq}}}{n^{\mathsf{eq}}} \frac{n_j^{\mathsf{eq}}}{n^{\mathsf{eq}}}$$

$\langle \sigma_{\sf eff} {m v} angle$	Sommerfeld Effect	Bound State Formation
$g_{ extsf{DM}} \gg g_{s}$	—	
$g_{ extsf{DM}} \ll g_{s}$	+	







Bound State Formation



Bound State Formation (BSF)

$$\sigma(X_1X_2 \to \mathcal{B}(X_1X_2) \ g) = \sigma_{\mathsf{BSF}} \sim \frac{\alpha_s^2}{m_X^2} S_{\mathsf{BSF}}\left(\frac{\alpha}{v}\right)$$

Bound state as an additional particle in the thermal bath. \Rightarrow Boltzmann Equation needs to be modified

Figures from [Harz,Petraki (2018)]

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Modified Coannihilation [Ellis, Luo, Olive (2015)]

$$\langle \sigma_{\text{eff}} \boldsymbol{\nu} \rangle = \sum_{i,j \in \{\chi,X\}} \langle \sigma_{ij} \boldsymbol{\nu}_{ij} \rangle \frac{n_i^{\text{eq}}}{n^{\text{eq}}} \frac{n_j^{\text{eq}}}{n^{\text{eq}}} + \langle \sigma_{\text{BSF}} \boldsymbol{\nu} \rangle_{\text{eff}} \frac{n_X^{\text{eq}}}{n^{\text{eq}}} \frac{n_X^{\text{eq}}}{n^{\text{eq}}}$$

Bound states effectively provide an additional annihilation channel.







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Bound states effectively provide an additional annihilation channel.

Bound State contribution to $\langle \sigma_{\rm eff} v \rangle$

$$\left\langle \sigma_{\text{BSF}} \nu \right\rangle_{\text{eff}} = \left\langle \sigma_{\text{BSF}} \nu \right\rangle \frac{\Gamma_{\mathcal{B} \to SM}}{\Gamma_{\mathcal{B}, \text{ion}} + \Gamma_{\mathcal{B} \to SM}}$$

 \rightarrow BSF only contributes to the annihilation cross section of DM if the bound states decay into SM particles!