Gaia parallaxes of Galactic open clusters calibrate Cepheid luminosity to 0.9% and raise H₀ tension with Planck+ Λ CDM to 5.3σ

Richard I. Anderson

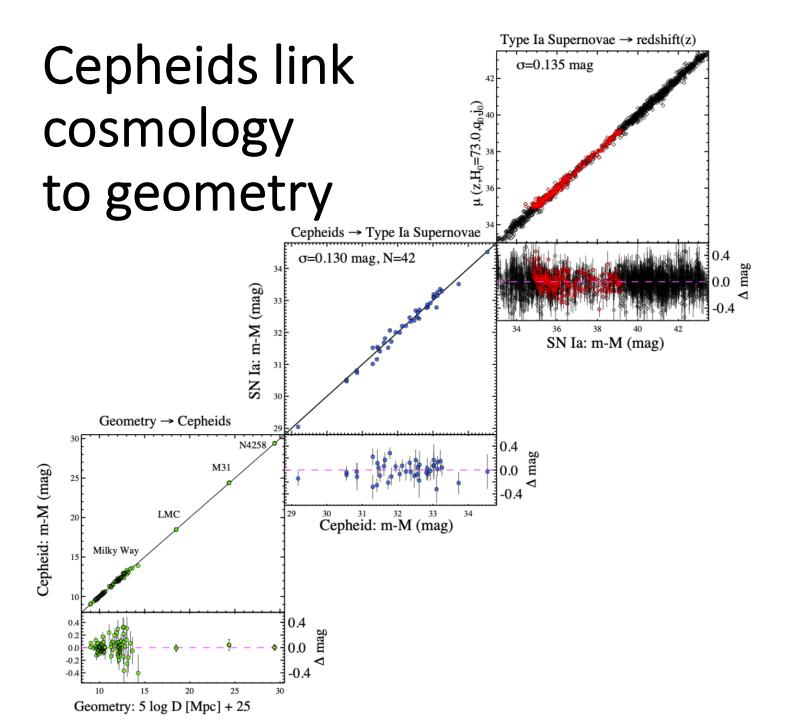


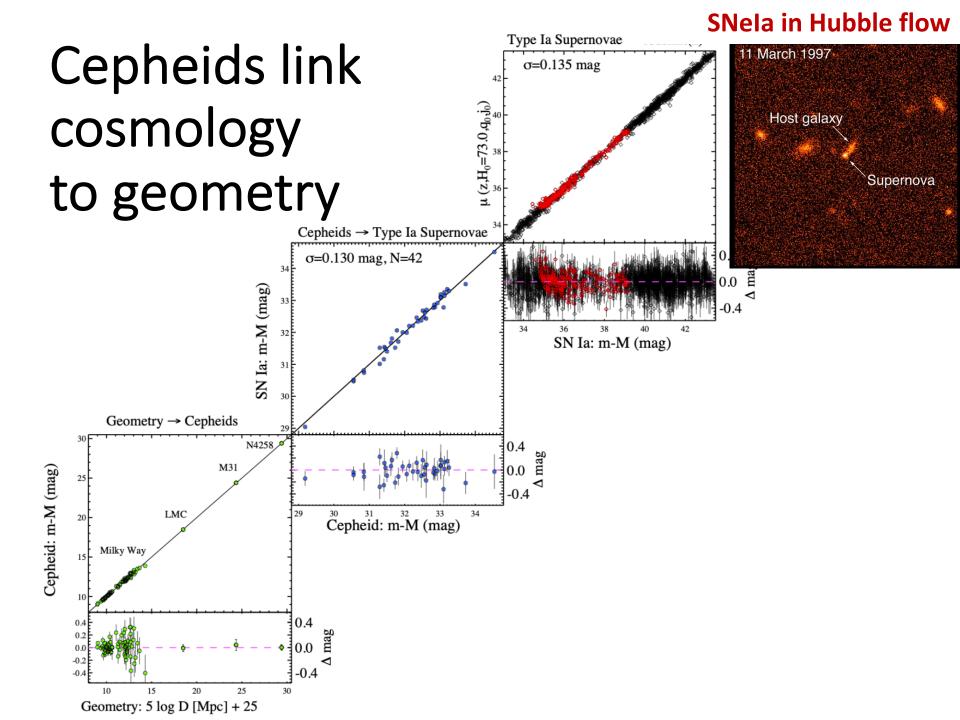
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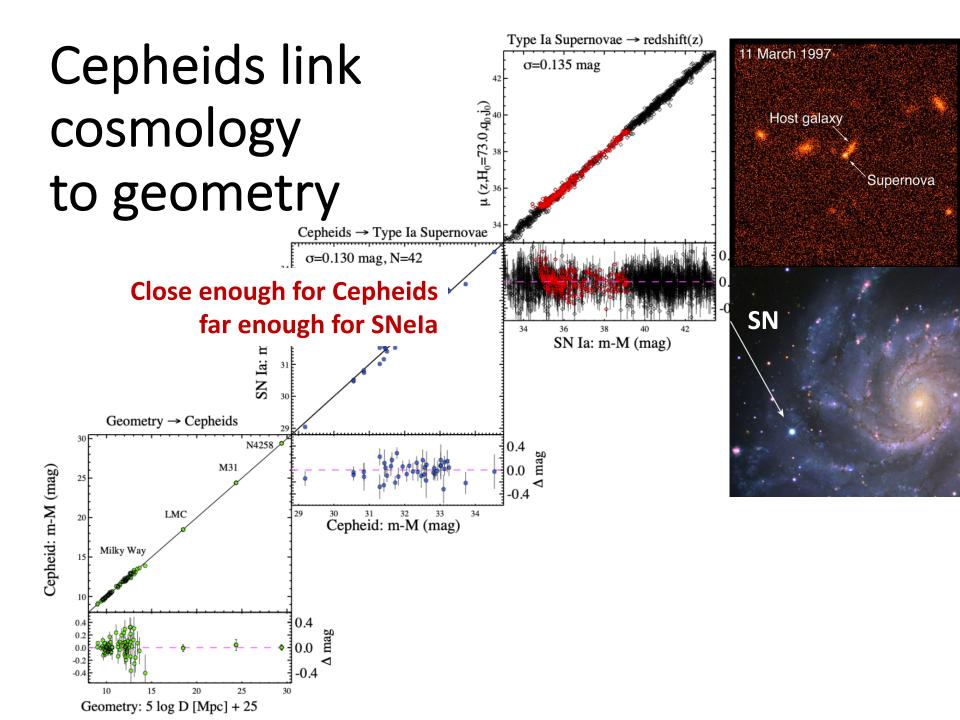
with Mauricio Cruz Reyes, Adam Riess, Stefano Casertano, and Louise Breuval

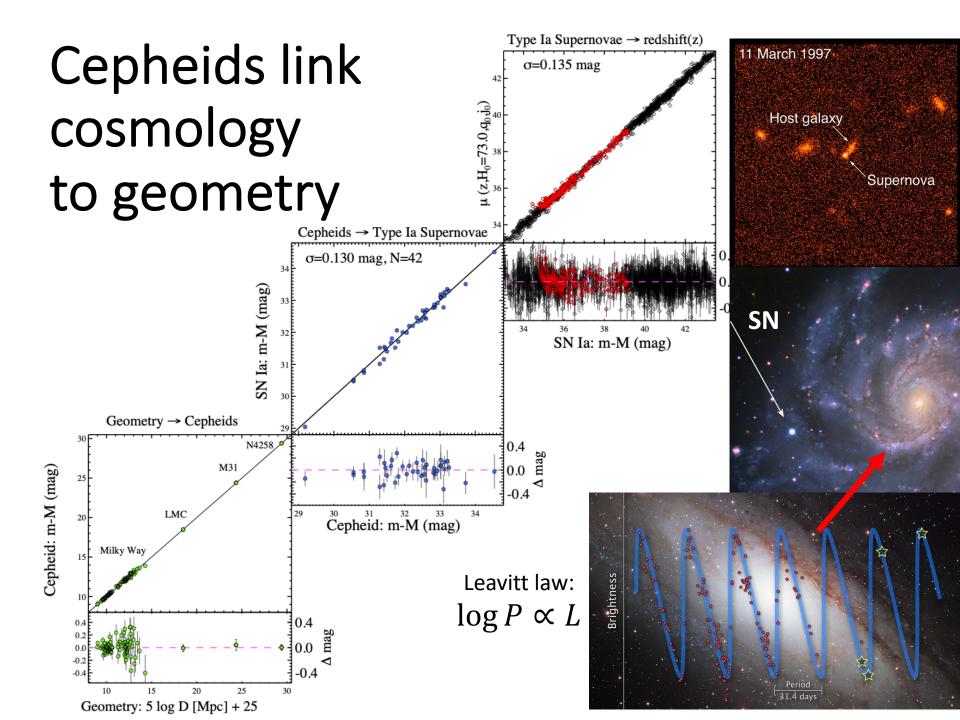
Please see arXiv IDs: 2208.01045 : Riess et al., ApJ accepted 2208.09403 : Cruz Reyes & Anderson, A&A submitted

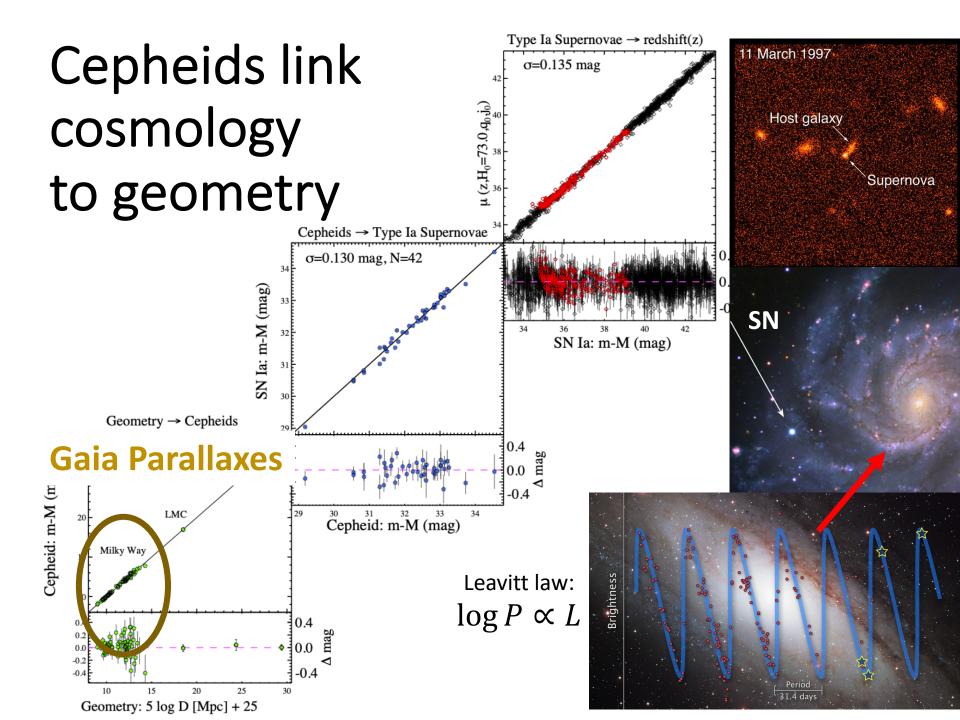












Classical Cepheids are great for this!

- Each Cepheid a standard candle
- Characteristic variability identifies
 Cepheids
- observed PL-sequences
 - Minimal contamination by non-Cepheids
 - Observed scatter contrains uncertainties
- Well understood by stellar evolution (e.g. RIA+2016, A&A 591, A8)

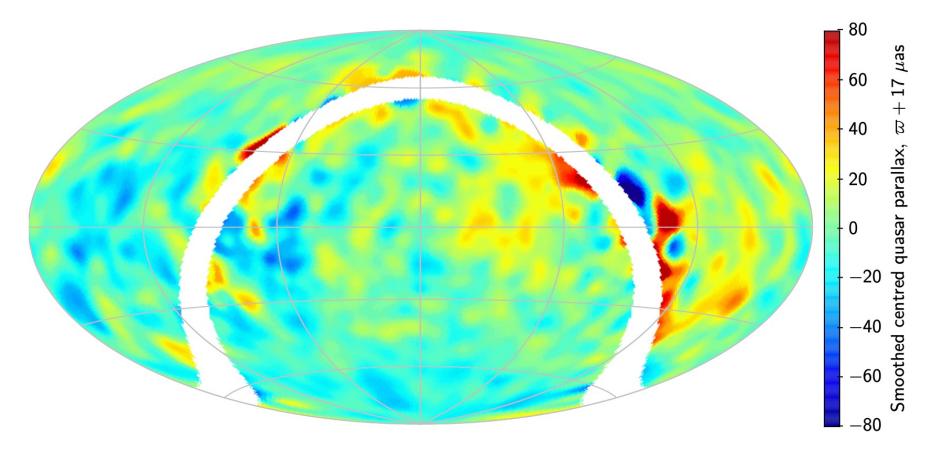
Cepheids improved <u>a lot</u> since 2016

- Improved geometric anchors
- Doubled SN-hosts with identified Cepheids
- Homogeneous HST photometry
- Metallicity effect measured by high-res spectroscopy

cf. Louise Breuval's talk on Saturday @ 10:20

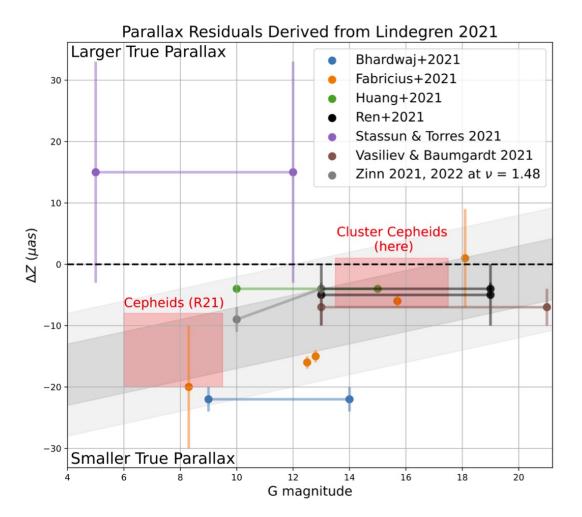
- Systematics checked extensively
- Gaia parallaxes : gold standard
- Precision & systematics improved : tension grows

Gaia parallax systematics by magnitude, color, position

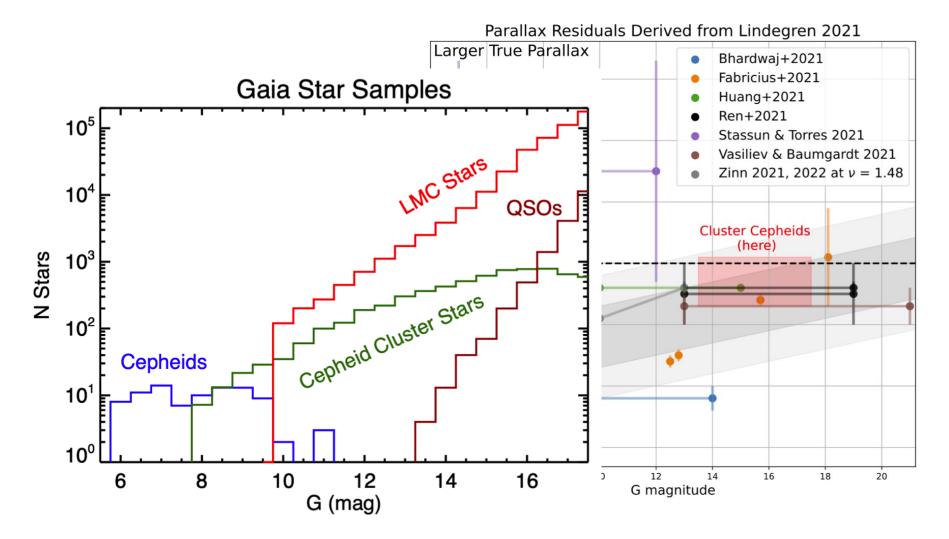


Lindegren+2021, A&A 649, A4

Systematics correction adequate at magnitudes where it is constrained



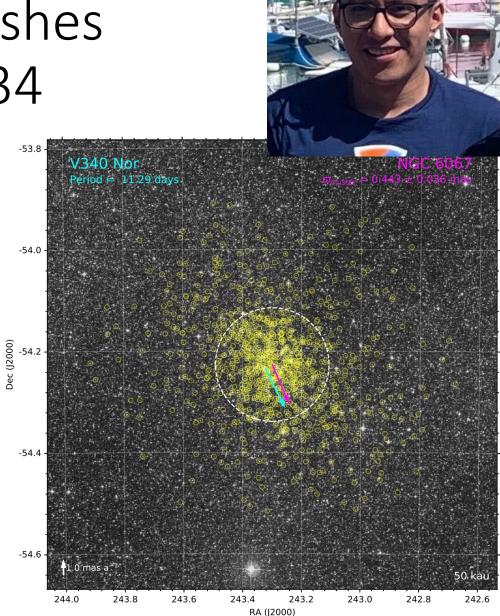
Systematics correction adequate at magnitudes where it is constrained



New census of Cluster Cepheids establishes Gold sample of 34

2208.09403 : Cruz Reyes & RIA

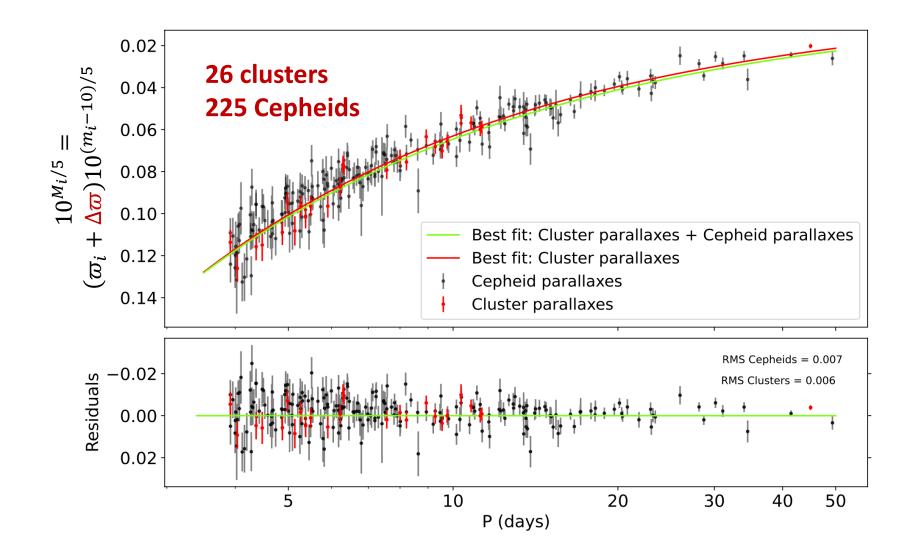
- Details: Mauricio Cruz Reyes, Friday 16:30 PS A
- New approach: find open clusters near Cepheids using Gaia DR3 data
- Typical total uncertainty : $\sim 7 \ \mu$ as



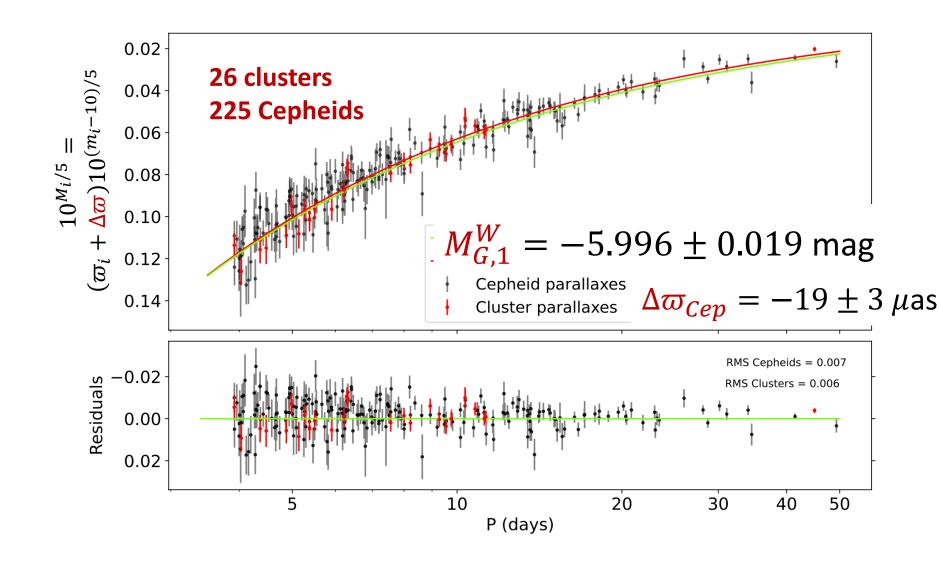
Gaia-only Cepheid luminosity calibration (no HST!) 2208.09403 : Cruz Reyes & RIA

- $M_i^W = \alpha \left(\log P 1 \right) + \frac{M_1^W}{H} + \gamma \left[\frac{\text{Fe}}{\text{H}} \right]$
- $10^{M_i/5} = (\varpi_i + \Delta \varpi) 10^{(m_i 10)/5}$
- $\Delta \varpi_{Cl} = 0$, $\Delta \varpi_{Cep} \neq 0$ (Lindegren+2021, Flynn+2022, Maiz Apellaniz 2022)
- L21 corrections applied to all parallaxes
- New Cepheid photometry from Gaia DR3 SOS (Ripepi+22)
- Combined 3-param fit to 26 cluster and 225 field Cepheids

0.9% calibration of Cepheid luminosity

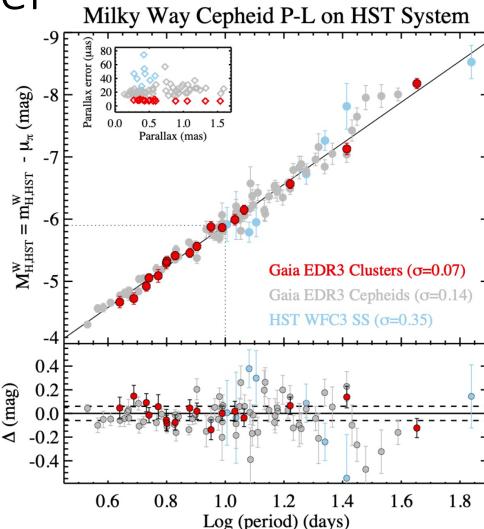


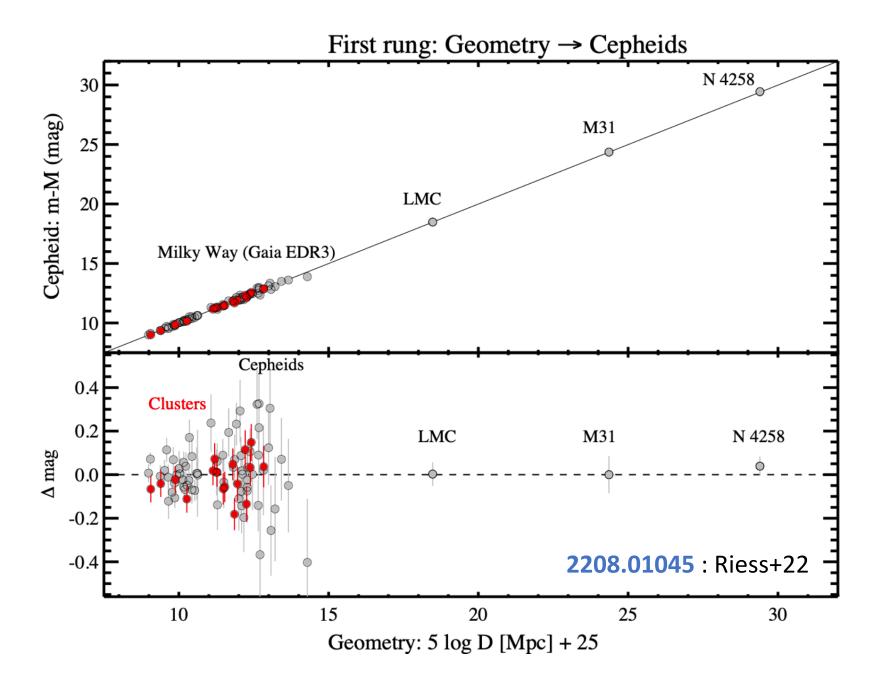
0.9% calibration of Cepheid luminosity



New HST photometry: 2208.01045 : Riess+22 LMC-like LL scatter Milky Way Cepheid P-L on HST S

- 17 cluster Cepheids with
 3-band HST photometry
- 3x better parallax errors
 9x the weight for LL calibration
- 0.06 0.045 mag scatter constrains parallax & reddening systematics
- 0.9% calibration of $M_{H,1}^W$





Constraining $M_{H,1}^W$ for SHOES H₀ analysis

•
$$\chi^2 = \left(\overline{\omega}_{EDR3,i} - \overline{\omega}_{phot,i} + \Delta \overline{\omega}_{Cl} \right)^2 \sigma_i^{-2}$$

•
$$\varpi_{phot,i} = 10^{-0.2(\mu_{o,i}-10)}$$

• $\mu_{o,i} = m_{i,H}^{W} - \left(M_{H,1}^{W} + \alpha \left(\log P - 1\right) + \gamma \Delta \left[\frac{0}{H}\right]_{i}\right)$

- σ_i combines photometric & astrometric uncertainties, PLR intrinsic width, uncertainty for offset if needed
- Free parameters: $M_{H,1}^W$ and (optionally) $\Delta \varpi_{Cl}$
- α, γ fixed to SH0ES baseline values : use $M_{H,1}^W$ as constraint for SH0ES H₀

$H_0 = 73.15 \pm 0.97 \text{ km s}^{-1} \text{ Mpc}^{-1}$ differs by **5**. **3** σ from *Planck+* Λ CDM

- 5-7% improvement on H₀ from cluster parallaxes
- $M_{H,1}^W$ and $\Delta \varpi_{Cl}$ fully consistent MCR & RIA (2022)
- MW alone : $H_0 = 73.3 \pm 1.1 \text{ km s}^{-1} \text{Mpc}^{-1}$

Fit^a	Ν	$M^W_{H,1}$	zp	b_W	Z_W	σ_{IS}	$H_0 \ CC \ only$	H ₀ all anchors
		[mag]	$[\mu \mathrm{as}]$	[mag/dex]		[mag]	$[\mathrm{kms^{-}}$	$^{1}\mathrm{Mpc}^{-1}]$
2, G+S	17	-5.902 ± 0.026	-3 ± 4	-3.299 ± 0.015^{b}	-0.217 ± 0.046^{b}	0.060	72.9 ± 1.3	$\textbf{73.04} \pm \textbf{0.99}$
2, G	14	-5.907 ± 0.024	-4 ± 4	-3.299 ± 0.015^{b}	-0.217 ± 0.046^{b}	0.047	72.7 ± 1.3	72.98 ± 0.99
$2, \mathrm{G+S}^c$	17	-5.893 ± 0.018	0^d	$-3.36 ext{ }\pm 0.07 ext{ }$	-0.217 ± 0.046^{b}	0.060		
$2, \mathrm{G}^{c}$	14	-5.907 ± 0.018	0^d	-3.44 ± 0.08	-0.217 ± 0.046^{b}	0.047		
1, G+S	17	-5.890 ± 0.018	0^d	-3.299 ± 0.015^{b}	-0.217 ± 0.046^{b}	0.060	73.3 ± 1.1	73.16 ± 0.97
$1, \mathrm{G}$	14	-5.892 ± 0.017	0^d	-3.299 ± 0.015^{b}	-0.217 ± 0.045^{b}	0.047	73.2 ± 1.1	73.14 ± 0.97

$H_0 = 73.15 \pm 0.97 \text{ km s}^{-1} \text{ Mpc}^{-1}$ differs by **5**. **3** σ from *Planck+* Λ CDM

MCR & RIA 22, fixed α, γ : $M_{H,1}^{W} = -5.915 \pm 0.017 \text{ mag}$ $\Delta \varpi_{Cep} = -13 \pm 5 \mu \text{as}$

- 5-7% improvement on H_0 from cluster parallaxes
- $M_{H,1}^W$ and $\Delta \varpi_{Cl}$ fully consistent MCR & RIA (2022)
- MW alone : $H_0 = 73.3 \pm 1.1 \text{ km s}^{-1} \text{Mpc}^{-1}$

Fit^a	Ν	$M^W_{H,1}$	zp	b_W	Z_W	σ_{IS}	$H_0 \ CC \ only$	H_0 all anchors
		[mag]	$[\mu as]$	[mag/dex]		[mag]	$[{\rm kms^{-1}Mpc^{-1}}]$	
2, G+S	17	-5.902 ± 0.026	-3 ± 4	-3.299 ± 0.015^{b}	-0.217 ± 0.046^{b}	0.060	72.9 ± 1.3	73.04 ± 0.99
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1, G+S	17	-5.890 ± 0.018	0^d	-3.299 ± 0.015^{b}	-0.217 ± 0.046^{b}	0.060	73.3 ± 1.1	73.16 ± 0.97
1, G	14	-5.892 ± 0.017	0^d	-3.299 ± 0.015^{b}	-0.217 ± 0.045^{b}	0.047	73.2 ± 1.1	73.14 ± 0.97

Conclusions

2208.01045 : Riess et al., ApJ accepted2208.09403 : Cruz Reyes & Anderson, A&A submitted

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- New cluster parallaxes combine statistical precision (more stars) and accuracy (systematics adequately corrected by Lindegren+2021)
- Cepheid luminosity scale calibrated to 0.9%, with or without HST
- Consistent $\Delta \varpi_{Cep}$ for both (Gaia only, Gaia+HST)
- Two independently determined sets of Gaia EDR3 cluster parallaxes yield close agreement for $M_{H,1}^W$ and $\Delta \varpi_{Cep}$
- Tight LL scatter severely limits margin for parallax uncertainties
- New HST photometry plus cluster parallaxes improves H₀ by 5-7%
- $H_0 = 73.15 \pm 0.97 \text{ km s}^{-1} \text{ Mpc}^{-1}$ is **5**.3 σ from Planck+ Λ CDM
- Exciting times: the tension rises as systematics improve