

Current and Future Constraints on H_0 from Infrared SBF

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We recently published a new measurement of the Hubble constant H_0 determined from distances to over 60 early-type galaxies out to 100 Mpc measured by the Surface Brightness Fluctuations (SBF) method using data from the WFC3 Infrared Channel on the Hubble Space Telescope (HST). More than a third of these galaxies have hosted well-measured Type Ia supernovae, and we use these to do detailed comparisons of SBF and SNIa distances. If we recalibrate SNIa using SBF, we find a steeper dependence of supernova peak luminosity on decline rate in early-type galaxies, but the resulting value of H_0 is indistinguishable from the values given independently by SNIa and SBF when each are calibrated from Cepheids. We discuss these recent SBF results from HST/WFC3 as well as the excellent prospects for a fully independent TRGB-calibrated SBF precision measurement of H_0 using JWST.

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