

Late-time Accelerating Universe in Teleparallel Gravity

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Outline

- The **H_0 tension** problem
- The **Teleparallel Gravity** (TG) formalism
- Constraining parameters in **TG models** using **Markov Chain Monte Carlo (MCMC) technique** and the **updated version of Pantheon, Pantheon+ data set**
- Conclusion

Motivation

Tensions in the Hubble expansion

- The Λ CDM model has been called into question – due to open problems such as the H_0 tension

		H_0 [km s ⁻¹ Mpc ⁻¹]
Indirect	P18	67.4 ± 0.5
	Atacama Cosmology Telescope	67.9 ± 1.5
↕		~5σ tension
Direct	R22 (SHOES Team)	73.04 ± 1.04
	H0LiCOW Collaboration	73.3 ^{+1.7} _{-1.8}

Theories beyond GR

Teleparallel Gravity

Einstein 1915: **General Relativity (GR)**,

- Metric tensor ($g_{\mu\nu}$)
- **Levi-Civita Connection:** Curvature, Torsion-free
- Einstein-Hilbert action

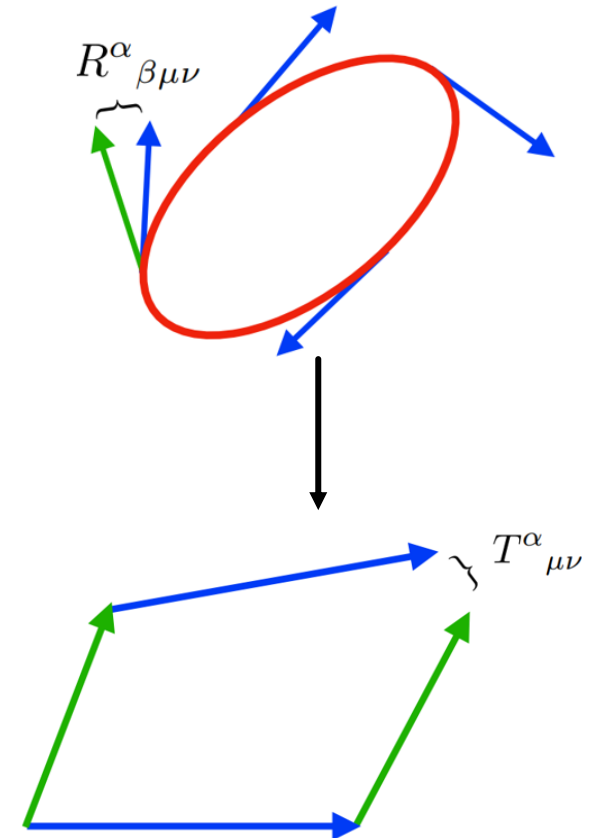


Einstein 1928: **Teleparallel Equivalent of GR**

- Tetrad (e^a_μ): relates the tangent space
- **Weitzenböck Connection:** Curvatureless, Torsion

$$S = -\frac{1}{16\pi G} \int d^4x e [T]$$

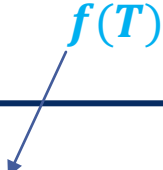
T = Torsion Scalar



Modified Teleparallel Gravity, $f(T)$

•Curvature-Torsion
Relation: $R = -T + B$

- Inspired by $f(R)$ gravity: **$f(T)$ gravity**

$$S = \frac{1}{16\pi G} \int d^4x e[-T + F(T)] + S_{\text{matter}}$$


- Taking a flat FRLW cosmology
- Friedmann equations:

$$H^2 = \frac{8\pi G}{3} \rho_m - \frac{F(T)}{6} + \frac{T}{3} F_T$$

$$T \equiv -6H^2$$

$$\dot{H} = -\frac{4\pi G(\rho_m + p_m)}{1 - F_T - 2TF_{TT}}$$

The Data



Observational Expansion Data Sets (1)

- Cosmic Chronometers (CC): Spectroscopic dating that depend on **stellar evolution** and **differential ageing** but **independent of cosmological models**, redshift range $z < 2$
- Baryonic Acoustic Oscillations (BAO): 10 model dependent points up to redshift $z < 2.4$

Observational Expansion Data Sets (2)

- Type Ia Supernovae (SNe Ia)
 - **Pantheon Compilation (SN):**
 - ❖ 1048 data points with redshift range, $0.01 < z < 2.3$
 - **Pantheon+ Compilation (SN+):** A successor of the original Pantheon
 - ❖ 1701 data points with redshift range, $0.001 < z < 2.3$
 - **Pantheon+SH0ES Compilation (SN+ & SH0ES):**
 - ❖ Uses the Pantheon+ compilations and includes SH0ES Cepheids (i.e constraints on H_0 with R22)

Our work

- Probe $f(T)$ gravity – Use recent observational **Hubble data**
- In conjunction with **Markov Chain Monte Carlo** (MCMC) algorithm

Explore **two** models in $f(T)$ gravity by constraining the model parameters

Analyse the **impact** of different data sets and **compare** results with Λ CDM model

Power Law Model



f_1 – Power Law Model

$$S = \frac{1}{16\pi G} \int d^4x e^{-T + F(T)} + S_{\text{matter}}$$

$f(T)$

$$F(T) = \alpha_1 T^{b_1}$$

- To obtain α_1 – evaluate the Friedmann equation at current times

$$\alpha_1 = (6H_0^2)^{1-b_1} \frac{1 - \Omega_{m_0} - \Omega_{r_0}}{2b_1 - 1}$$

- Thus, the Friedmann equation for this model could be written as

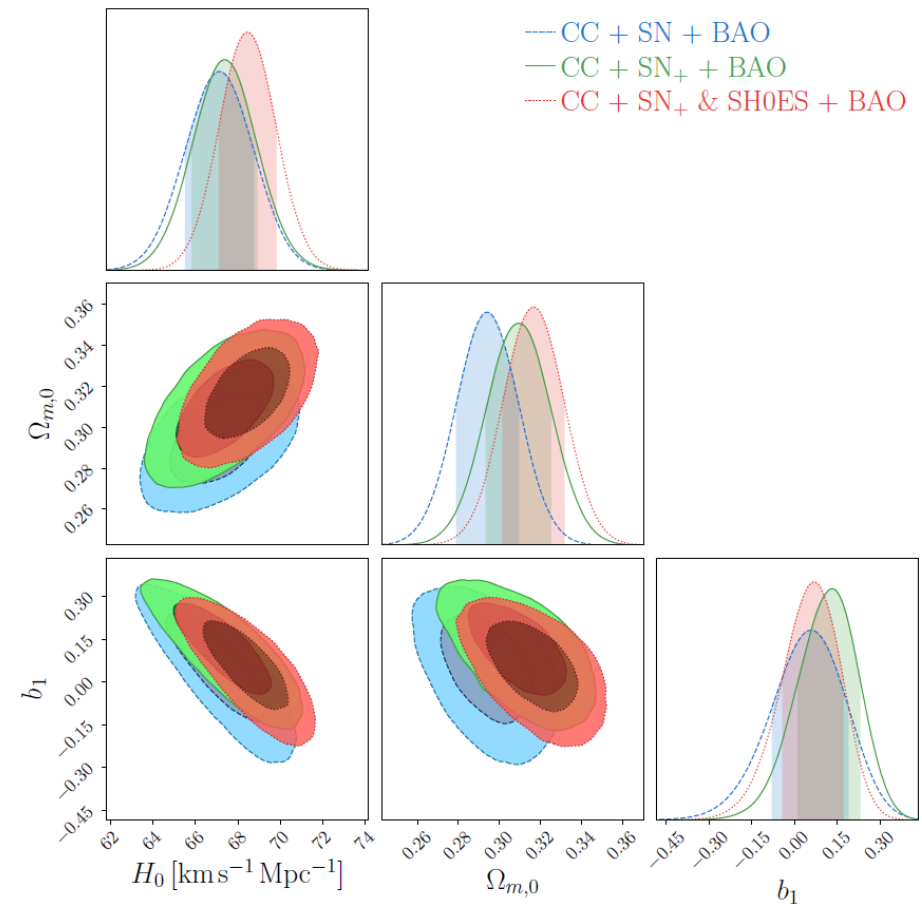
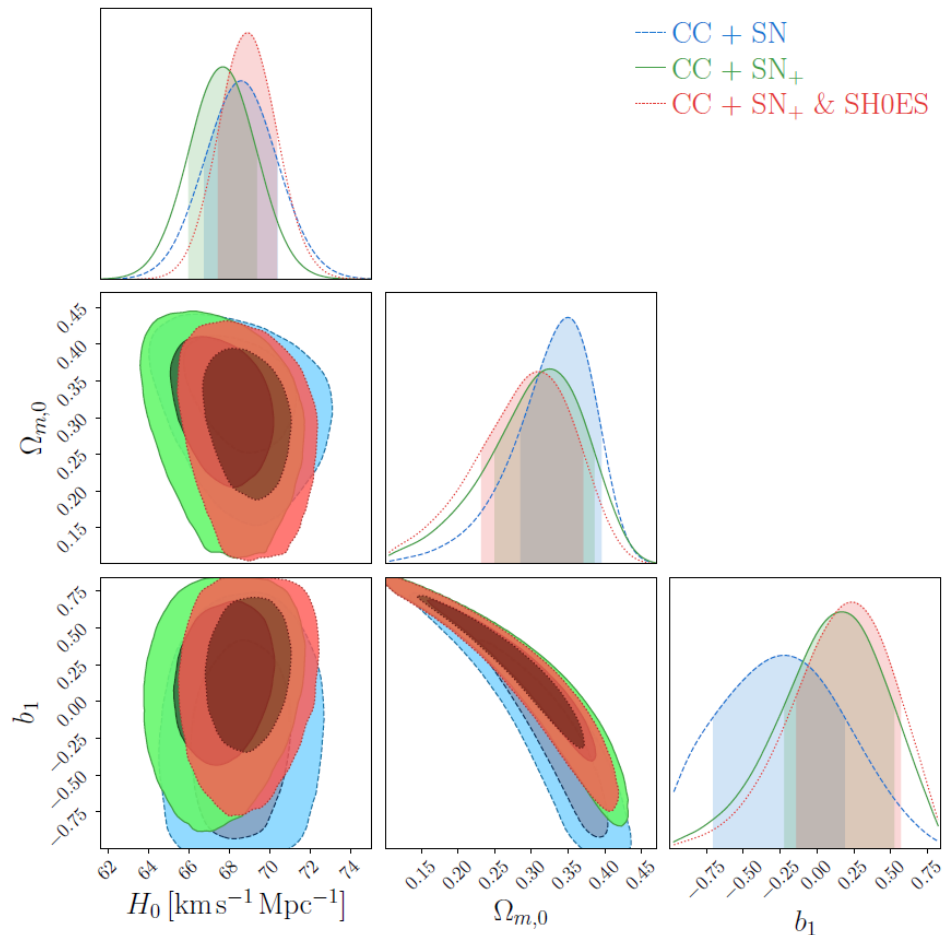
$$E^2(z) = \Omega_{m_0}(1+z)^3 + \Omega_r(1+z)^4 + (1 - \Omega_{m_0} - \Omega_{r_0})E^{2b_1}(z)$$

$$E = \frac{H(z)}{H_0}$$

$$\Omega_r = 4.15 \times 10^{-5} / h_0$$

f_1 – Power Law Model Results

$$F(T) = \alpha_1 T^{b_1}$$



f_1 – Power Law Model

$$F(T) = \alpha_1 T^{b_1}$$

Data Sets	H_0 [km s ⁻¹ Mpc ⁻¹]	$\Omega_{m,0}$	b_1	AIC	BIC	Δ AIC	Δ BIC
CC + SN	68.6 ± 1.8	0.350 ^{+0.045} _{-0.064}	-0.22 ^{+0.41} _{-0.48}	1048.94	1068.88	1.45	6.43
CC + SN ₊	67.7 ± 1.7	0.324 ^{+0.061} _{-0.074}	0.16 ^{+0.36} _{-0.38}	1426.81	1446.83	1.72	6.72
CC + SN₊ & SHOES	68.9 ± 1.4	0.310^{+0.059}_{-0.078}	0.24^{+0.33}_{-0.37}	1420.02	1440.04	1.47	6.47
CC + SN + BAO	67.1 ± 1.6	0.294 ± 0.015	0.06 ± 0.13	1065.13	1085.13	1.68	6.68
CC + SN ₊ + BAO	67.4 ± 1.5	0.309 ± 0.016	0.132 ^{+0.097} _{-0.120}	1434.83	1454.91	0.54	5.66
CC + SN ₊ & SHOES + BAO	68.4 ± 1.3	0.317 ± 0.015	0.069 ^{+0.100} _{-0.111}	1428.80	1448.88	1.55	6.57

$$\text{AIC} = 2k - \ln(L)$$

No. of parameters in model

Maximum likelihood

$$\text{BIC} = k \ln(n) - 2 \ln(L)$$

No. of data points

Difference between Λ CDM and PLM in AIC & BIC

Linder Model

f_2 – Linder Model

$$S = \frac{1}{16\pi G} \int d^4x e^{-T + F(T)} + S_{\text{matter}}$$

$f(T)$ →

$$F(T) = \alpha_2 T_0 \left(1 - \text{Exp} \left[-b_2 \sqrt{T/T_0} \right] \right)$$

- To obtain α_2 – evaluate the Friedmann equation at current times

$$\alpha_2 = \frac{1 - \Omega_{m_0} - \Omega_{r_0}}{1 - (1 + b_2)e^{-b_2}}$$

- Thus, the Friedmann equation for this model could be written as

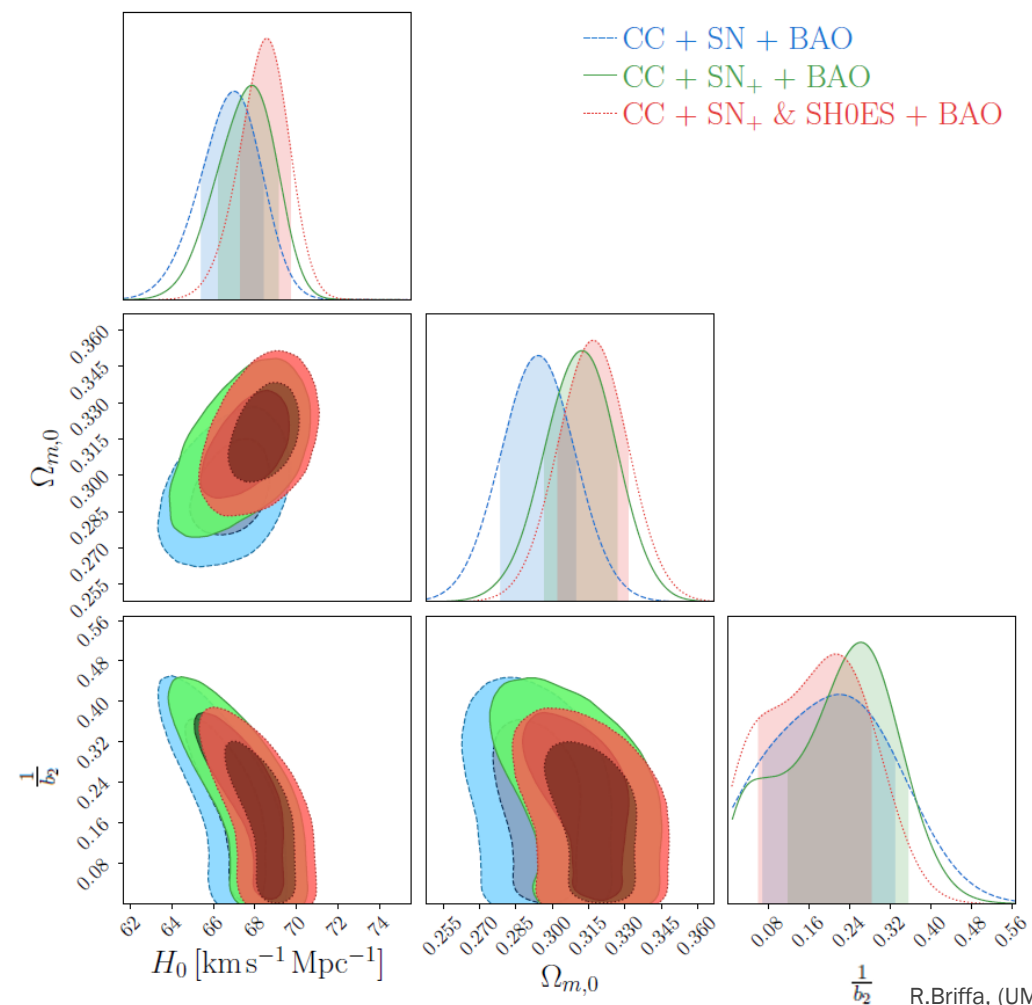
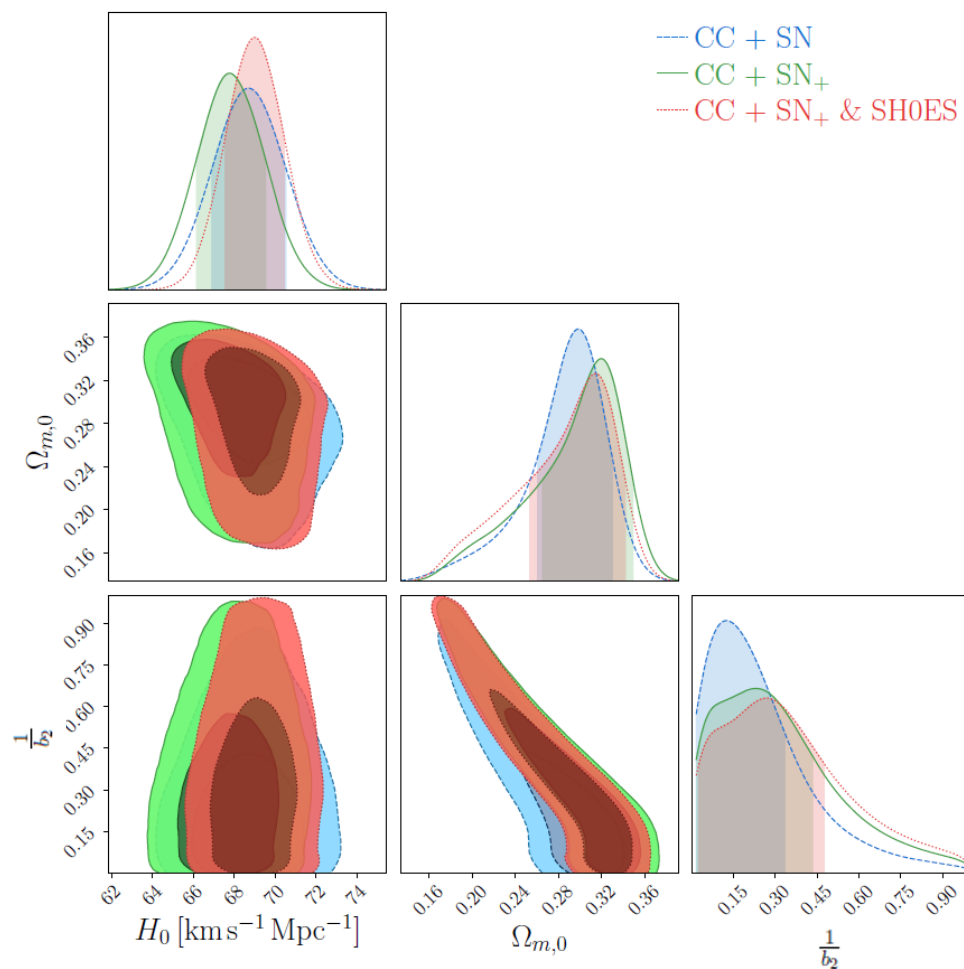
$$E^2(z) = \Omega_{m_0}(1+z)^3 + \Omega_r(1+z)^4 - \frac{1 - \Omega_{m_0} - \Omega_{r_0}}{1 - (1 + b_2)e^{-b_2}} \left[(1 + b_2 E(z)) \text{Exp}[-b_2 E(z)] - 1 \right]$$

$$\Omega_r = 4.15 \times 10^{-5} / h_0$$

$$E = \frac{H(z)}{H_0}$$

f_2 – Linder Model Results

$$F(T) = \alpha_2 T_0 \left(1 - \text{Exp} \left[-b_2 \sqrt{T/T_0} \right] \right)$$



f_2 – Linder Model

$$F(T) = \alpha_2 T_0 \left(1 - \text{Exp} \left[-b_2 \sqrt{T/T_0} \right] \right)$$

Data Sets	H_0 [km s ⁻¹ Mpc ⁻¹]	$\Omega_{m,0}$	$\frac{1}{b_2}$	AIC	BIC	ΔAIC	ΔBIC
CC + SN	68.7 ^{+1.8} _{-1.7}	0.298 ^{+0.031} _{-0.035}	0.12 ^{+0.21} _{-0.11}	1049.49	1069.43	2.00	6.98
CC + SN ₊	67.7 ^{+1.8} _{-1.6}	0.319 ^{+0.030} _{-0.054}	0.23 ^{+0.20} _{-0.21}	1426.83	1446.85	2.74	6.74
CC + SN ₊ & SHOES	69.8 ^{+1.4} _{-1.5}	0.315 ^{+0.027} _{-0.061}	0.27 ^{+0.21} _{-0.24}	1420.05	1440.07	1.50	6.51
CC + SN + BAO	67.0 ^{+1.4} _{-1.6}	0.294 ± 0.016	0.22 ^{+0.12} _{-0.15}	1064.62	1084.52	1.06	6.06
CC + SN ₊ + BAO	67.9 ^{+1.3} _{-1.6}	0.312 ^{+0.014} _{-0.015}	0.262 ^{+0.094} _{-0.142}	1434.52	1454.60	0.23	5.35
CC + SN ₊ & SHOES + BAO	68.6 ^{+1.1} _{-1.3}	0.317 ^{+0.014} _{-0.015}	0.212 ^{+0.070} _{-0.151}	1428.50	1448.58	1.26	6.27

$$\text{AIC} = 2k - \ln(L)$$

No. of parameters in model

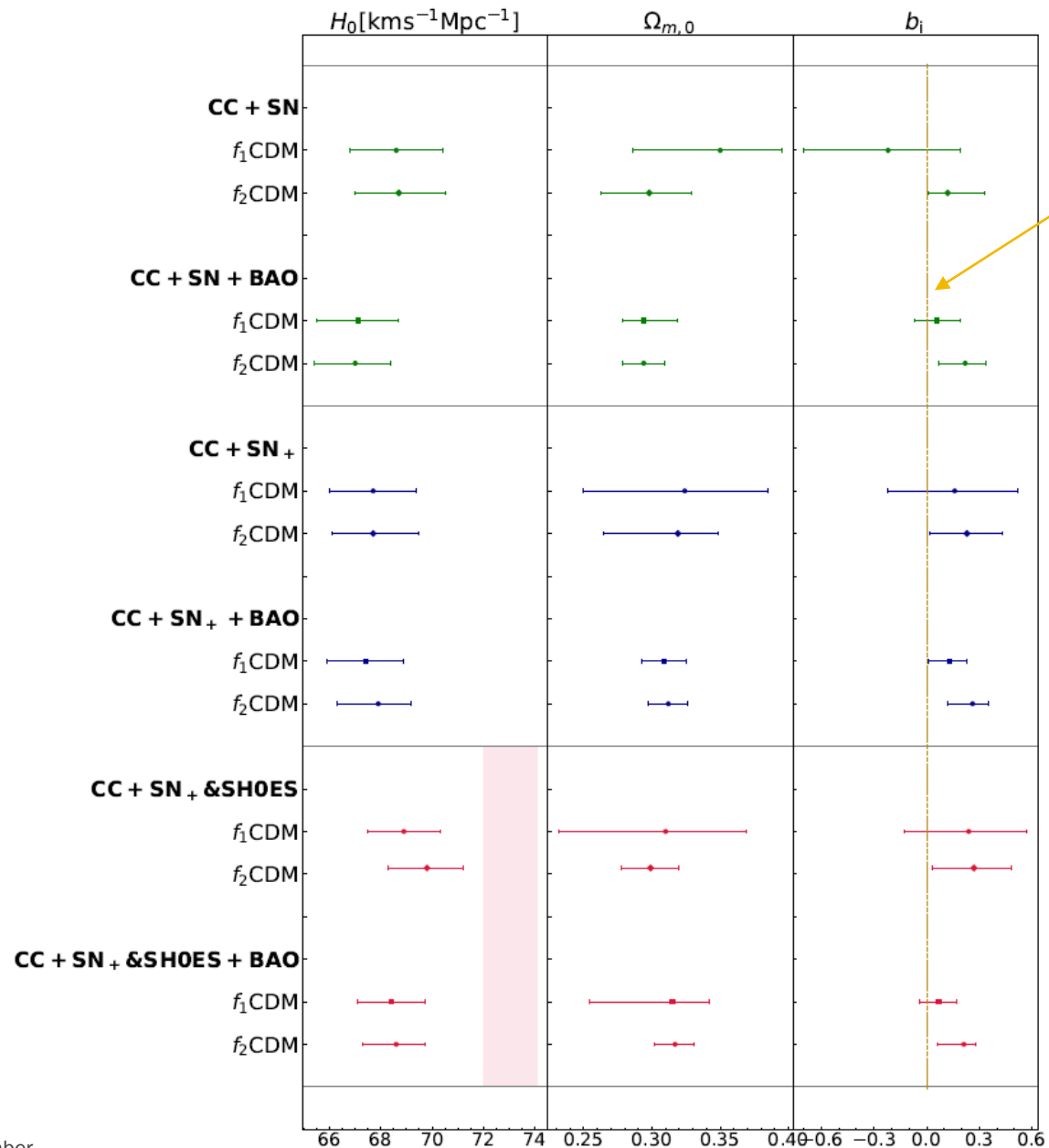
Maximum likelihood

$$\text{BIC} = k \ln(n) - 2 \ln(L)$$

No. of data points

Difference between Λ CDM and PLM in AIC & BIC

Conclusion



$b = 0$
 Λ CDM model

■ R22 value ($H_0 = 73.04 \pm 104 \text{ km s}^{-1} \text{ Mpc}^{-1}$)

Conclusion and prospects

- We have tested **2 $f(T)$ models**
- **MCMC** allows us to **constrain the parameters** in these models
- Our interest: assess how the **cosmological parameters** are **altered**
- This analysis can be **extended** to other theories such as **$f(T, B)$**



Thank You
