Generalizing the Friedmann Model in Light of Cosmological Tensions

Timothy Clifton (Queen Mary, University of London)

in collaboration with Theo Anton

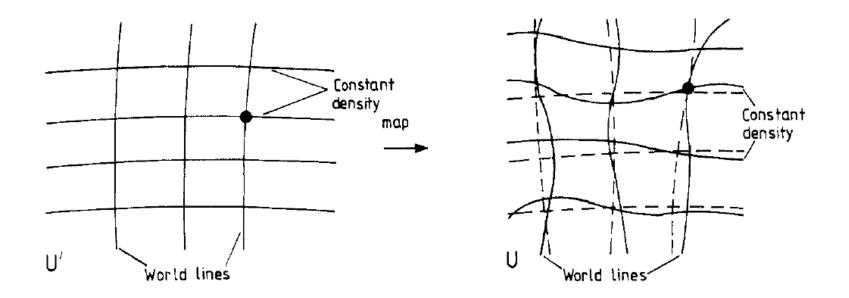
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[[]Ellis & Stoeger CQG 4, 1697 (1987)]

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➤ à la Gasperini et al.

• Define an 'average' scale factor: $a_{\mathcal{D}}(a)$

$$(t) \equiv \left(\frac{\int_{\mathcal{D}} d^3 X \sqrt{{}^{(3)}g(t, X^i)}}{\int_{\mathcal{D}} d^3 X \sqrt{{}^{(3)}g(t_0, X^i)}}\right)^{\frac{1}{3}}$$

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where:
$$\langle \psi \rangle_{\mathcal{D}}(t) \equiv \frac{\int_{\mathcal{D}} d^3 X \sqrt{(3)} g(t, X^i)}{\int_{\mathcal{D}} d^3 X \sqrt{(3)} g(t, X^i)}$$
, $\mathcal{Q}_{\mathcal{D}} \equiv \frac{2}{3} \left(\langle \Theta^2 \rangle_{\mathcal{D}} - \langle \Theta \rangle_{\mathcal{D}}^2 \right) - 2 \langle \sigma^2 \rangle_{\mathcal{D}}$

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see e.g. MacCallum, arXív:2001.11387)

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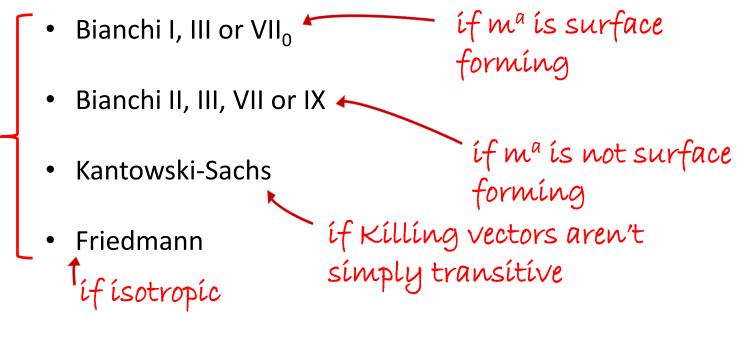


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where, e.g.,

$$\mathcal{Q}_{3} = \frac{1}{3} \operatorname{Cov}\left(\Theta, \Sigma\right) + \frac{2}{3} \operatorname{Var} \mathcal{A} - \frac{1}{3} \operatorname{Cov}\left(\phi, \mathcal{A}\right) + \frac{2}{3} \left\langle m^{a} D_{a} \mathcal{A} \right\rangle - \frac{1}{2} \operatorname{Var} \Sigma - \frac{1}{3} \left\langle M^{ab} D_{a} \mathcal{A}_{b} \right\rangle \\ - \frac{1}{3} \left\langle \Sigma_{a} \Sigma^{a} \right\rangle + \frac{1}{3} \left\langle \mathcal{A}_{a} \mathcal{A}^{a} \right\rangle + \frac{1}{3} \left\langle \Sigma_{ab} \Sigma^{ab} \right\rangle + 2 \left\langle \alpha_{a} \Sigma^{a} \right\rangle - \frac{2}{3} \left\langle a_{a} \mathcal{A}^{a} \right\rangle.$$

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publication to appear!

Thank you