No more trouble with Hubble

Solving the H_0 -tension à la Ellis & Stoeger

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Please look into the light...



...and forget all cosmology you have known



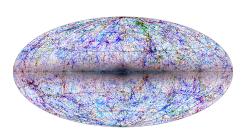
Image credits: VectorStock

Our universe in all its complexity



 $z \approx 0$

highly non-linear structures

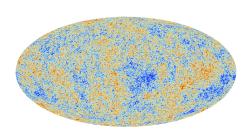




z = 1100

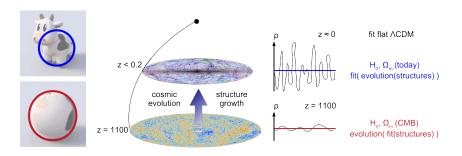
almost homogeneous & isotropic





The fitting problem in cosmology (Ellis & Stoeger 1987)

find the best-fit (flat) Λ CDM models for observations today and at CMB:



evolution(fit(structures)) ≠ fit(evolution(structures))

ightarrow best-fit flat Λ CDM from today and CMB can be different!

Independent best-fit Λ CDM model parameters

today (Cepheids + SNe)

$$h = 0.7304 \pm 0.0104$$

 $\Omega_m = 0.285 \pm 0.013$

(Riess et al. 2022)

(Wagner & Meyer 2019)



CMB

$$h = 0.6727 \pm 0.0060$$

$$\Omega_m = 0.3166 \pm 0.0084$$

(Planck 2020, CMB-only) (Planck 2020, CMB-only)

Can the fitting problem account for the H_0 -tension?

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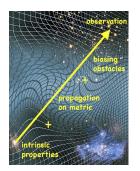
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assume astro & biases are consistent

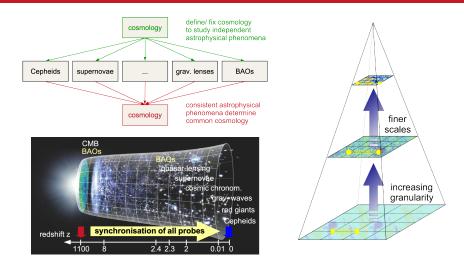
$$D_{\rm com}(z) = \frac{c}{H_0} \int_0^z \frac{\mathrm{d}x}{\sqrt{\Omega_m (1+x)^3 + \Omega_\Lambda}}$$

$$D_{\rm com}(z) = D_{\rm com}(z)$$

$$\Omega_m \approx \Omega_m \left(\frac{H_0}{H_0}\right)^2 \stackrel{!}{=} 0.269$$

 \rightarrow yes, right trend in $(H_0, \Omega_m)!$ clumpy universe difficult to fit!

The solution to the H_0 -tension – consistency & synchronisation



→ synchronise all probes over time and scale (using BAOs)

The H_0 -tension as a fitting problem (arXiv:2203.11219)

H_0 -tension caused by different model fits across scales and time



Advantages

- no need for additional (dark) physics
- no ad-hoc changes in evolution
- easy to fix by tracing BAOs across cosmic time
- may solve other tensions as well (e.g. σ_8)

Foundations in Ellis 1984, Ellis et al. 1985, Ellis & Stoeger 1987

THANK YOU!



Our white paper on H_0 :

arXiv:2207.05765

Cosmo-of-69 web blog: cosmoprinciple.wordpress.com