Exotic turbulence opportunities in superfluid helium

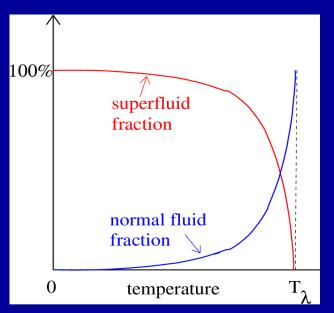


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Helium II

- Normal fluid component:
 - (related to thermal excitations)
 - Viscous. Vortices can have any size and strength
- Superfluid component:
 - (related to quantum ground state)
 - Inviscid. Only vortex lines of fixed circulation & core radius

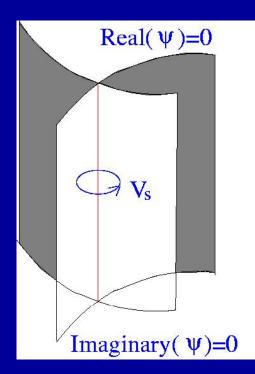


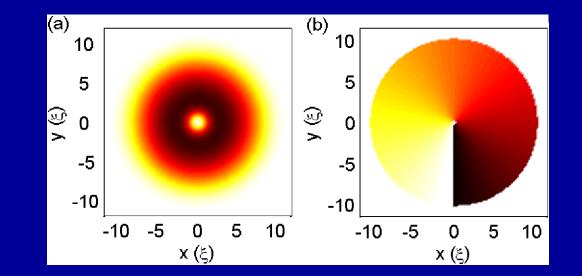
Relative proportion of superfluid and normal fluid depends on T

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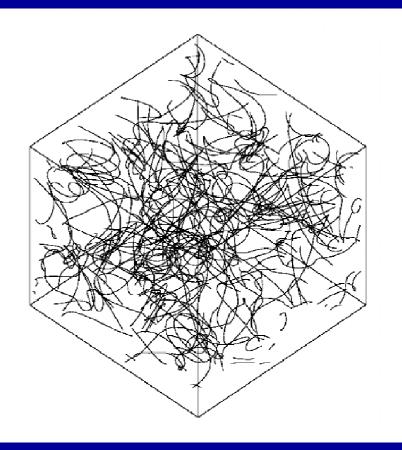


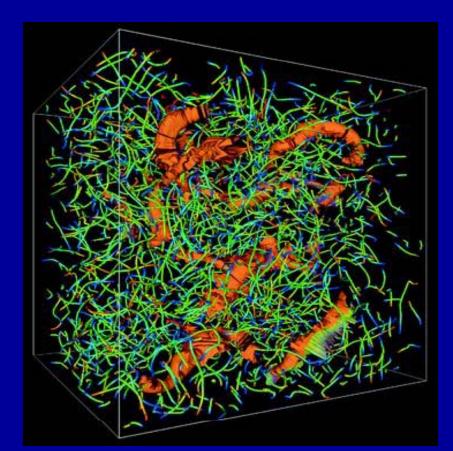


Quantised vortices as topological defects No classical vortex core stretching

Superfluid turbulence

Ordinary turbulence







(Newcastle)

KNOWN SUPERFLUIDS ⁴He ³He atomic Bose-Einstein condensates neutron stars

TURBULENCE Relevant to all known superfluids

MANY FORMS OF TURBULENCE

1) ⁴He at low T, ³He at low T: Single quantised turbulence

 ³He at high T: Single quantised turbulence with simple dissipation against stationary normal fluid

3) ⁴He at high T: Double turbulence (one quantised, one not)

4) ⁴He counterflow (heat transfer):

5) 1-dim turbulence along vortices in ⁴He and ³He at low T: Kelvin wave cascade

1) ⁴He at low T, ³He at low T: Single quantised turbulence

Experiments show that superfluid turbulence decays (Davis et al, Physica B 280, 43) and that it diffuses (Fisher et al, PRL 86, 244, 2001).

Numerical simulations (NLSE model: Nore et al, PRL 78, 3896, 1997; vortex dynamics model: Araki et al, PRL 89, 145301, 2002) predict k^{-5/3} energy spectrum.

Energy sink ? Sound emission at vortex reconnections (Leadbeater et al PRL 86, 1410, 2001) and by high-k Kelvin waves radiation (Vinen, PRB 64, 134520, 2001)

2) ³He at high T: Single quantised turbulence with simple dissipation against stationary normal fluid

Experiments Finne et al, Nature 424, 1022, 2003

Prediction of spectrum L'vov, Nazarenko & Volovik, JETP Lett 80, 479, 2004, Vinen, J Phys Chem Solids 66, 1493, 2005

3) ⁴He at high T: Double turbulence (one quantised, one not)

Experiments: k^{-5/3} spectrum for energy (Maurer & Tabeling, Europhysics Lett 43, 29, 1998) and vortex line density (Roche et al 2007); t^{-3/2} temporal decay (Smith et al, PRL 71, 2583, 1993), drag crisis past sphere (Smith et al PoF 11, 751, 1999)

Superfluid has Kolmogorov spectrum driven by imposed Kolmogorov normal fluid (Kivotides et al PRL 57, 845 2002)

Prediction of unusual pressure spectrum (Kivotides et al PRL 87, 275301, 2001)

Coupled, self-consistent 2-fluid calculations done only for vortex ring (Kivotides et al, Science 290, 777, 2000)

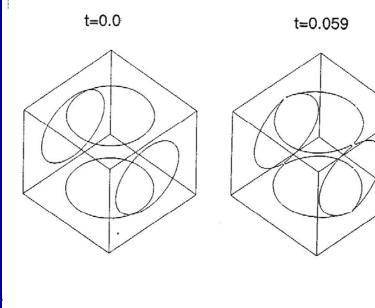
4) ⁴He counterflow (heat transfer):

The most studied case (from Vinen, Donnelly, Tough and Schwarz, to eg Blaztkova et al, PRE 75, 025302, 2007)

Experiments show t^{-3/2} decay (Barenghi, Godeev, Skrbek, PRE 74, 026309, 2006)

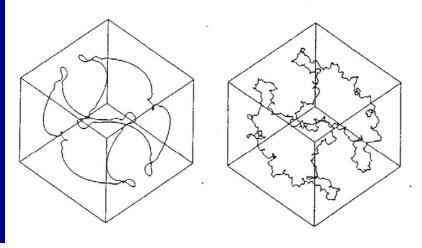
Large scale motion V_n and V_s must be different

5) 1-dim Kelvin wave cascade along vortices in ⁴He and ³He at low T



t=0.069

t=0.129



-Kivotides, Vassilicos, Samuels & Barenghi, PRL 86, 3080, 2001)
-Vinen, Tsubota & Mitani, PRL 91, 135301, 2003
-Kozik & Svistunov PRL 92, 035301, 2004

-Nazarenko, JETP Lett 84, 585, 2007

AN ANALOGY

Mars Thin atmosphere (Almost all CO₂ in ground) Average temperature : - 50°C



Venus Thick atmosphere containing 96% of CO₂ Average temperature : + 420°C

We know more about the Earth's atmosphere because we can test our toolkit of ideas and methods on other planets

GRID () Areadal UNEP

Examples: -greenhouse effect of Venus vs global warming -dust storms on Mars vs volcanic eruptions / nuclear winter

Planets and atmospheres

HOW TO WRITE A SUCCESSFUL PROPOSAL?

A successful EU proposal requires a PUNCHLINE strong enough to be the focus and to attract the attention

Achieving relatively large Rayleigh / Reynolds numbers in He gas, Hel or high-T Hell is not strong enough

It is the rich VARIETY of turbulence problems available in superfluid ³He and ⁴He which provides us with a punchline

CONCLUSIONS

- Labs involved in ³He and low-T ⁴He are essential nodes

- Crucial role of theoretical / numerical modelling