# Dark Fifth Forces in the Universe

## Non-linear Cosmology

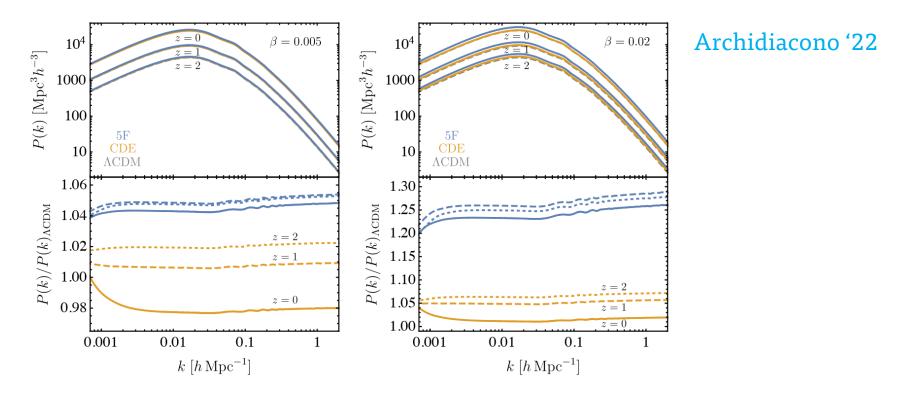
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# Why going to Non-linear order?

1. No new phases in the power spectrum. Bounds from full shape.



2. Effects emerging only at non-linear order (e.g. squeezed bispectrum).

# Lagrangian Perturbation Theory

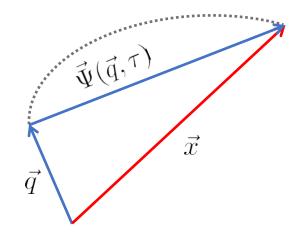
White '14 Bernardeau '02

Follow the fluid element:

$$\vec{x} = \vec{q} + \vec{\Psi}(\vec{q}, \tau)$$

Equations for displacements from geodesic equation  $\tilde{m} \equiv \frac{\partial \log m}{\partial s}$ 

$$\begin{cases} \vec{\Psi}_{\chi}^{"} + (\mathcal{H} + \bar{s}^{'}\tilde{m})\vec{\Psi}_{\chi}^{'} = -\vec{\nabla}_{\chi}\Phi_{\chi} - \tilde{m}\vec{\nabla}_{\chi}s \\ \vec{\Psi}_{b}^{"} + \mathcal{H}\vec{\Psi}_{b}^{'} = -\vec{\nabla}_{b}\Phi_{b} \end{cases}$$



Poisson's equations

$$\begin{cases} \nabla_x^2 \Phi = \frac{3}{2} \mathcal{H}^2 \Omega_m (f_\chi \delta_\chi + (1 - f_\chi) \delta_b) \equiv \frac{3}{2} \mathcal{H}^2 \Omega_m \delta_m \\ \nabla_x^2 s = \frac{3}{2} \mathcal{H}^2 \Omega_m \beta f_\chi \tilde{m} \delta_\chi \end{cases}$$

Energy conservation

$$\rho(\vec{x})d^3x = \bar{\rho}(1 + \delta(\vec{x}))d^3x = \bar{\rho}d^3q \longrightarrow \delta(\vec{x}) = \left(\det\frac{\partial\vec{x}}{\partial\vec{q}}\right)^{-1} - 1 = -\vec{\nabla}\vec{\Psi} + \frac{1}{2}\left((\vec{\nabla}\vec{\Psi})^2 + \partial_i\Psi_j\partial_j\Psi_i\right) + \mathcal{O}(\Psi^3)$$

# Lagrangian Perturbation Theory

#### Convenient variables:

$$\begin{cases} \vec{\Psi}_m = f_{\chi} \vec{\Psi}_{\chi} + (1 - f_{\chi}) \vec{\Psi}_b \\ \vec{\Psi}_r = \vec{\Psi}_{\chi} - \vec{\Psi}_b \end{cases} \longrightarrow \begin{cases} \vec{\Psi}''_r + \mathcal{H} \vec{\Psi}'_r = -(\vec{\Psi}_r \cdot \vec{\nabla}) \vec{\nabla} \Phi - \tilde{m} \vec{\nabla}_{\chi} s - \bar{s}' \tilde{m} \vec{\Psi}'_m \propto \beta \\ \vec{\Psi}''_m + \mathcal{H} \vec{\Psi}'_m = -\vec{\nabla}_m \Phi_m - f_{\chi} \tilde{m} \vec{\nabla}_{\chi} s - f_{\chi} \bar{s}' \tilde{m} \vec{\Psi}'_m \end{cases}$$

Solve order by order in Fourier space:

$$\vec{\Psi}(\vec{k},\tau) = \vec{\Psi}^{(1)}(\vec{k},\tau) + \vec{\Psi}^{(2)}(\vec{k},\tau) + \cdots$$

Structure of the solutions up to second order:

$$\vec{\Psi}_{m,r}^{(1)}(\vec{k},\tau) = \frac{i\vec{k}}{k^2} \delta_0(\vec{k}) D_{1m,1r}(\tau)$$

$$\vec{\Psi}_{m}^{(2)}(\vec{k},\tau) = \frac{i\vec{k}}{2k^2} \int \frac{\mathrm{d}^3k_1}{(2\pi)^3} \frac{\mathrm{d}^3k_2}{(2\pi)^3} (2\pi)^3 \delta(\vec{k} - \vec{k}_1 - \vec{k}_2) \delta_0(\vec{k}_1) \delta_0(\vec{k}_2) \left(1 - (\hat{k}_1 \cdot \hat{k}_2)^2\right) D_{2m}(\tau)$$

$$\vec{\Psi}_{r}^{(2)}(\vec{k},\tau) = \frac{i\vec{k}}{2k^2} \int \frac{\mathrm{d}^3k_1}{(2\pi)^3} \frac{\mathrm{d}^3k_2}{(2\pi)^3} (2\pi)^3 \delta(\vec{k} - \vec{k}_1 - \vec{k}_2) \delta_0(\vec{k}_1) \delta_0(\vec{k}_2) \left[\left(1 - (\hat{k}_1 \cdot \hat{k}_2)^2\right) D_{2r}^{np}(\tau) + \left(1 + \frac{\hat{k}_1 \cdot \hat{k}_2}{2} \left(\frac{k_2}{k_1} + \frac{k_1}{k_2}\right)\right) D_{2r}^{p}(\tau)\right]$$

### Growth Factors and Green's Functions

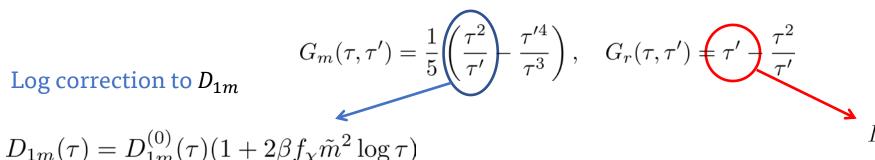
General equations for growth factors:

$$\begin{cases}
L_{\tau m} D_{nm} \equiv \left(\partial_{\tau}^{2} + \frac{2}{\tau} \partial_{\tau} - \frac{6}{\tau^{2}}\right) D_{nm} = S_{m}(\tau) \equiv S_{0}(D_{(n-1)m}^{(0)}, \cdots, D_{1m}^{(0)}) + \beta \tilde{m}^{2} f_{\chi}^{2} S_{1}(D_{nm}^{(0)}, \cdots, D_{1m}^{(0)}) \\
L_{\tau r} D_{nr} \equiv \left(\partial_{\tau}^{2} + \frac{2}{\tau} \partial_{\tau}\right) D_{nr} = \beta \tilde{m}^{2} f_{\chi}^{2} S_{r}(D_{nm}^{(0)}, \cdots, D_{1m}^{(0)})
\end{cases}$$

Introduce the Green's functions:

$$\begin{cases} L_{\tau}G(\tau,\tau') = \delta(\tau - \tau') \\ G(\tau - \tau') = 0 \text{ if } \tau < \tau' \end{cases} \longrightarrow D(\tau) = \int_{-\tau}^{\tau} d\tau G(\tau,\tau') S(\tau')$$

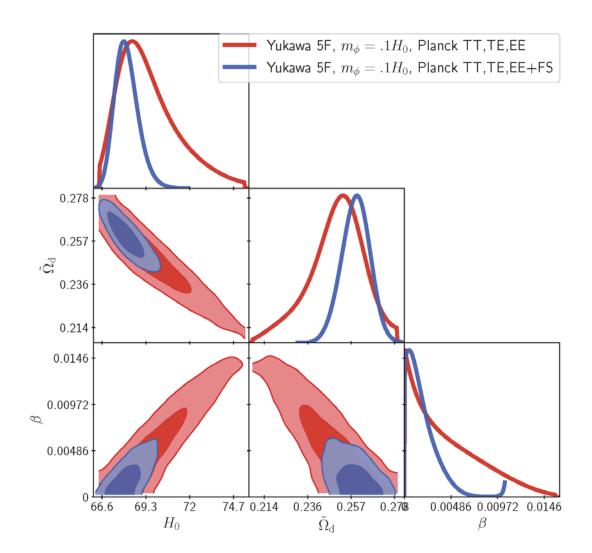
We only need the Green's function in  $\Lambda$ CDM:



Non-decaying relative growth factor

$$D_{nr}(\tau) \propto \beta f_{\chi}^2 \tilde{m}^2 (D_{1m}^{(0)}(\tau))^n$$

# Power Spectrum: Full Shape Analysis

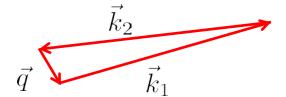


# EP violations: squeezed bispectrum

Peloso '13 Creminelli '13 Lewandowski '19

Equal-time squeezed bispectrum:

$$\lim_{q\to 0} \langle \delta(\vec{q},\tau)\delta(\vec{k}_1,\tau)\delta(\vec{k}_2,\tau)\rangle$$



- Equivalence principle holds
   Adiabatic initial conditions

Existence of free-falling frame for the long mode



Consistency relations:

$$\lim_{q \to 0} \langle \delta(\vec{q}, \tau) \delta(\vec{k}_1, \tau) \delta(\vec{k}_2, \tau) \rangle' = O_a(q, k) P(q) P(k)$$

Analytic function of q

Poles are generated in q: 
$$\lim_{q\to 0} \langle \delta(\vec{q},\tau) \delta(\vec{k}_1,\tau) \delta(\vec{k}_2,\tau) \rangle' \propto \frac{\vec{q} \cdot \vec{k}}{q^2} P(q) P(k)$$

# EP violations: squeezed bispectrum

Fifth forces induce violation of the EP in the growth of  $\delta_r$ 

$$\lim_{q \to 0} \langle \delta_{\chi}(\vec{q}) \delta_{\chi}(\vec{k}_{1}) \delta_{b}(\vec{k}_{2}) \rangle = (1 - f_{\chi}) \lim_{q \to 0} \langle \delta_{m}^{(1)}(\vec{q}) \delta_{m}^{(2)}(\vec{k}_{1}) \delta_{m}^{(1)}(\vec{k}_{2}) \rangle - f_{\chi} \lim_{q \to 0} \langle \delta_{m}^{(1)}(\vec{q}) \delta_{m}^{(1)}(\vec{k}) \delta_{r}^{(2)}(\vec{k}_{2}) \rangle$$

Only term with the correct spatial structure

$$\sim \alpha(\vec{k}_1, \vec{k}_2) = 1 + \frac{\hat{k}_1 \cdot \hat{k}_2}{2} \left( \frac{k_1}{k_2} + \frac{k_2}{k_1} \right)$$

# EP violations: squeezed bispectrum

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$$\downarrow \qquad \qquad \downarrow$$

$$\lim_{q \to 0} \langle \delta_{\chi}(\vec{q}) \delta_{\chi}(\vec{k}_{1}) \delta_{b}(\vec{k}_{2}) \rangle' = \lim_{q \to 0} \frac{D_{2r}^{p}}{D_{1m}^{2}} P_{m}(q) P_{m}(k) \left( f_{\chi} \alpha(q, k_{1}) - (1 - f_{\chi}) \alpha(q, k_{2}) \right) = \frac{\beta f_{\chi} \tilde{m}^{2}}{2} P_{m}(q) P_{m}(k) \frac{\vec{q} \cdot \vec{k}}{q^{2}}$$

Smoking gun!

To be compared with  $\Delta \langle \delta_m(\vec{k}_1) \delta_m(\vec{k}_2) \delta_m(\vec{k}_3) \rangle \propto \beta \log \tau$ 

#### Outlook

- Precise full-shape analysis: include BAO reconstrunction
- Bispectrum:  $O(\beta \log \tau)$  deviations from  $\Lambda$ CDM vs  $O(\beta)$  EP violation
- Bispectrum: non-adiabatic initial conditions?

Thanks for the attention