

Coupled maps for electron and ion clouds

Ubaldo Iriso and Steve Peggs

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Contents:

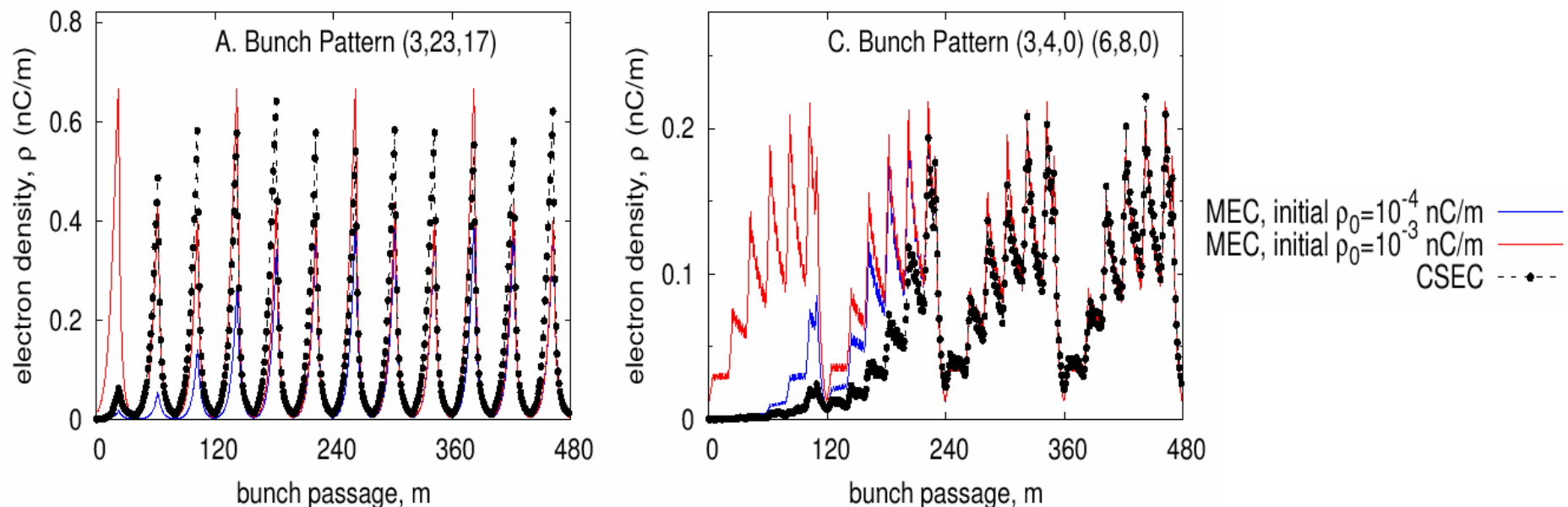
- 1. Introduction: Maps for Electron Clouds**
- 2. Experimental observations:**
 - a) Phase transitions**
 - b) Ion clouds**
- 3. Coupled electron and ion clouds**
- 4. Results using an "proof-of-principle" coupling**
 - a) Stability and importance of initial conditions**
 - b) Hysteresis and additional dynamical phases**
- 5. Conclusions**

1. Introduction: Maps for Electron Clouds

- A cubic iterative map model was presented to study e-clouds*, where the average e-density ρ at bunch passage $m+1$ depends on the e-density at previous bunch passage, m :

$$\rho_{m+1} = a\rho_m + b\rho_m^2 + c\rho_m^3$$

- Good agreement between maps (**MEC**) vs long computer sim codes (**CSEC**) for different bunch patterns, but map simulations run ~ 7 orders of magnitude faster:



- Maps also provide a level of abstraction to tackle electron clouds that can render fruitful conclusions. For instance, the way to minimize e-cloud density for different bunch patterns.

*U. Iriso and S. Peggs, *PRST-AB*, 8, 024403, 2005

Holy grail: calculation of maps coefficients

But, the map coefficients were inferred after fitting results obtained from the detailed simulation codes (like *ECLLOUD*, *CSEC*...)

→ Ideally: calculate map coefficients analytically.

With some simplifications, Ref* shows the linear map coefficient **a** is interpreted as the effective SEY – δ_{eff} :

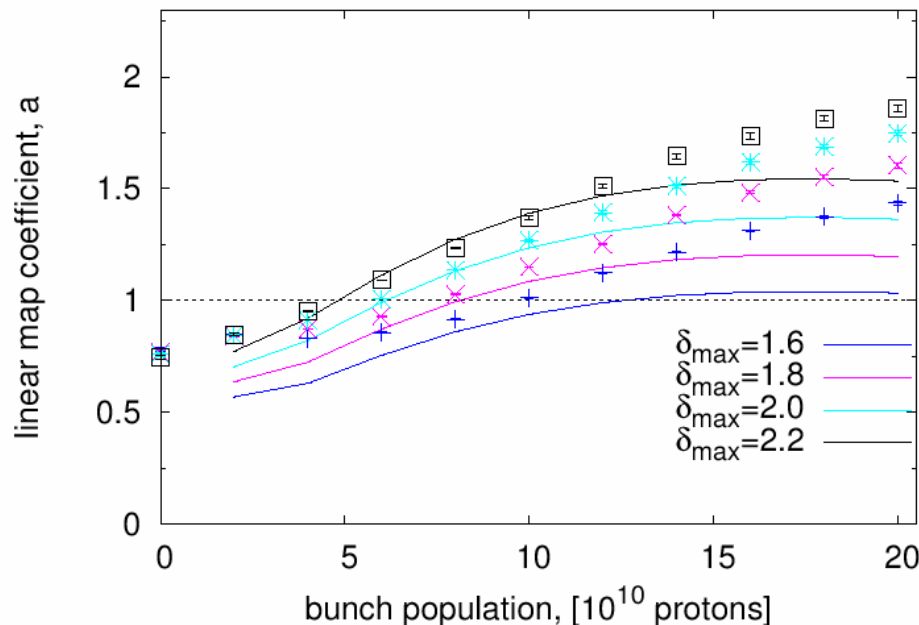
$$a = \int_0^\infty \left[\delta_r(E)^{n(E)} + \delta_t(E) \delta_{sec}^{\xi(E)} \frac{\delta_{sec}^{n(E)\xi(E)} - \delta_r^{n(E)}}{\delta_{sec}^{\xi(E)} - \delta_r(E)} \right] h(E) dE$$

h(E) = energy spectrum after bunch passage

n(E) = number of oscillations of an e- at energy E between two bunches

δ_r = secondary electrons coming from elastic reflections

δ_t = “true” secondary electrons



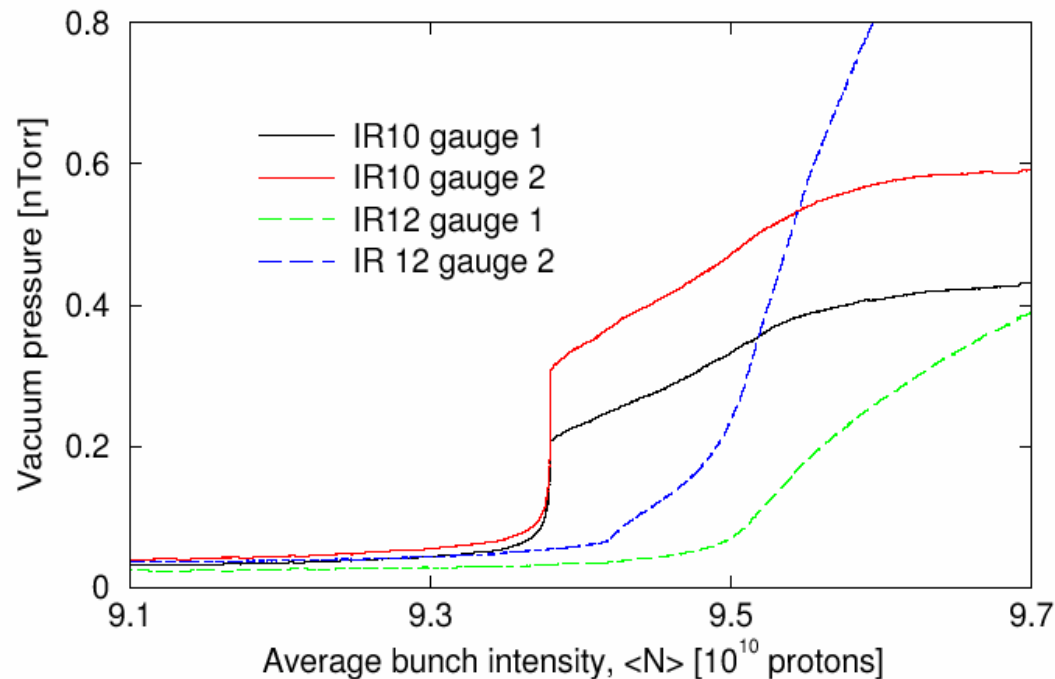
Acceptable agreement between fitting (points) and the analytical solution for **a**

*U. Iriso, PhD Thesis, BNL/CAD-228

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2. Experimental observations

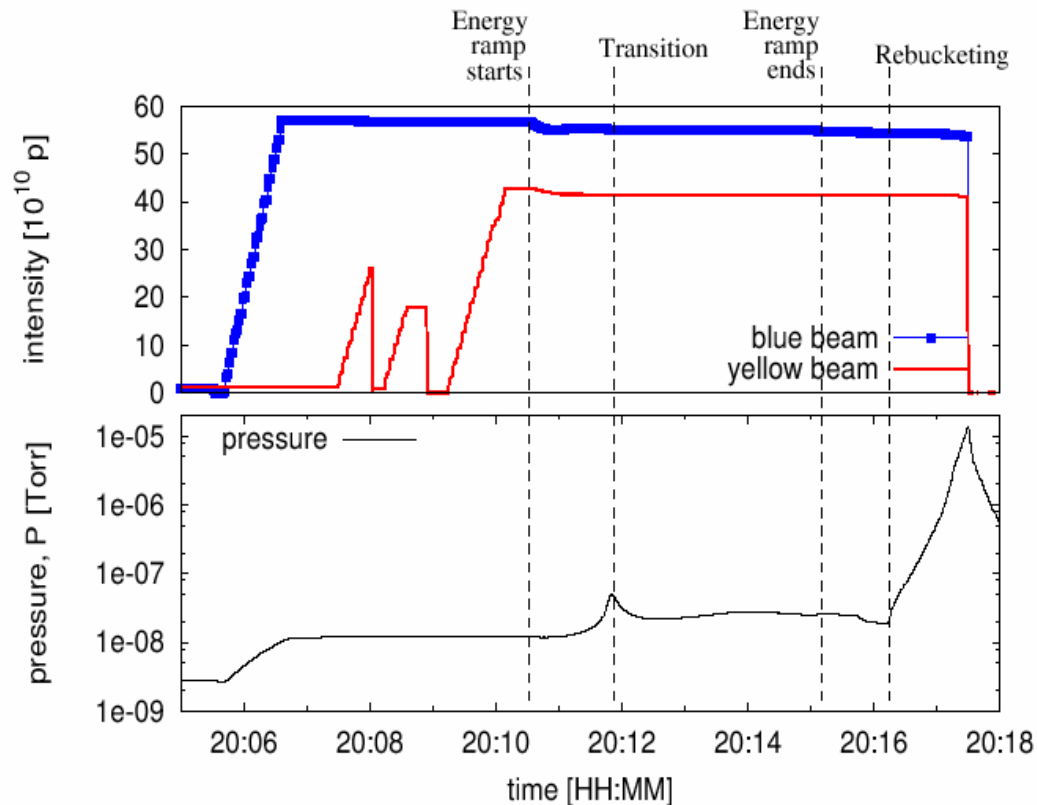
2.a. First and second order phase transitions



- Pressure (due to e-clouds) smoothly decays in IR12, it shows an abrupt decay in IR10, as the bunch intensity threshold for e-clouds is crossed.
- But contemporary simulation codes only reproduce a smooth transition from “cloud off” \leftrightarrow “cloud on” *
- How can both first and second order phase transition occur in e-clouds?

*S. Peggs and U. Iriso, *Proceedings of ELOUD'04, 2004*

2.b. A vacuum instability driven by e-clouds: ion clouds?



Proposed explanation*:

e-clouds and beam-gas collisions create ions, leading to a vacuum instability

- So, e-clouds can trigger an ion cloud (see Refs. * and **)
- Significant number of parameters to determine ions behaviour: different cross sections for different gases, backscattering probability, vacuum pumping...

*W. Fischer, U. Iriso, E. Mustafin, 33rd ICFA, 2004

** O. Grobner, CARE HHH, 2004

→ But ion lifetimes are ~3-6 order of magnitude larger than e- → rather complex to introduce into the contemporary e-cloud codes (*CSEC*, *E-CLOUD*) because of their prohibitively large CPU times

→ **Can maps circumvent this prohibition?**

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3. Coupled electron and ion clouds

Assume ion clouds can be formed and “couple” them to the electron cloud using maps:

$$\begin{aligned} \rho_{m+1} &= f(\rho_m, R_m) \longrightarrow \text{for electron density} \\ R_{m+1} &= g(\rho_m, R_m) \longrightarrow \text{for ion density} \end{aligned}$$

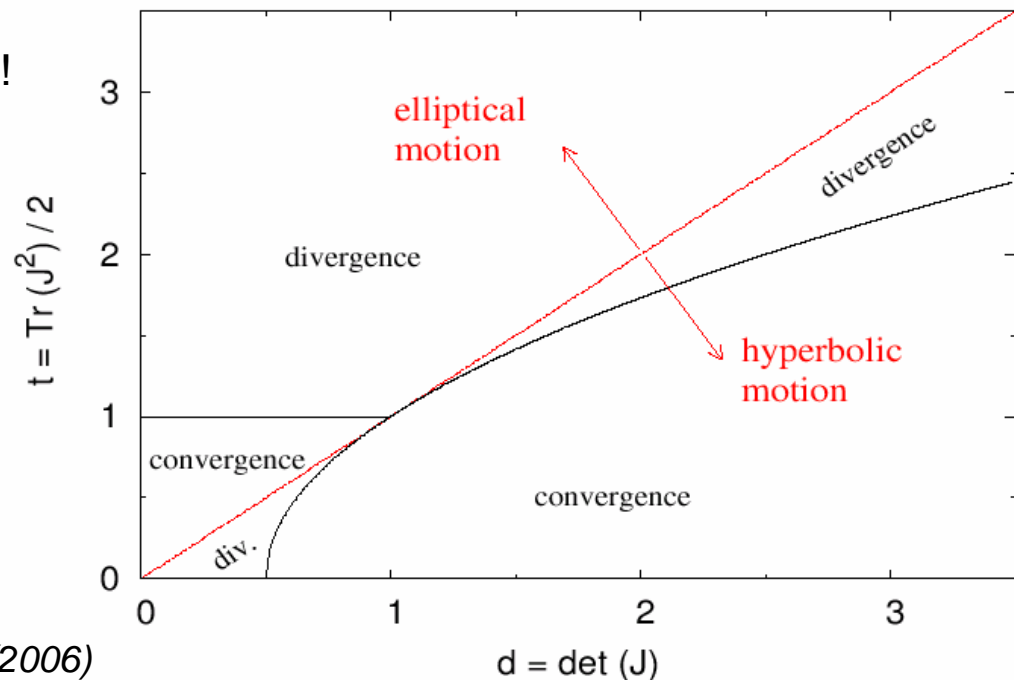
The 2-system is characterized by the vector $\vec{r}_m = \begin{pmatrix} \rho_m \\ R_m \end{pmatrix}$

Equilibrium is found if $\vec{r}_{m+1} = \vec{r}_m \equiv \vec{r}^*$ at the so-called, “fixed points”

→ But we need the fixed points to be **stable!!**

• Stability condition of the fixed points depend on the Jacobian matrix:

$$J = \begin{pmatrix} \frac{\partial f}{\partial \rho_m} & \frac{\partial f}{\partial R_m} \\ \frac{\partial g}{\partial \rho_m} & \frac{\partial g}{\partial R_m} \end{pmatrix}_{\vec{r}^*}$$



*see U. Iriso and S. Peggs, PRST-AB, 9, 071002 (2006)

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 - b) Hysteresis and additional dynamical phases
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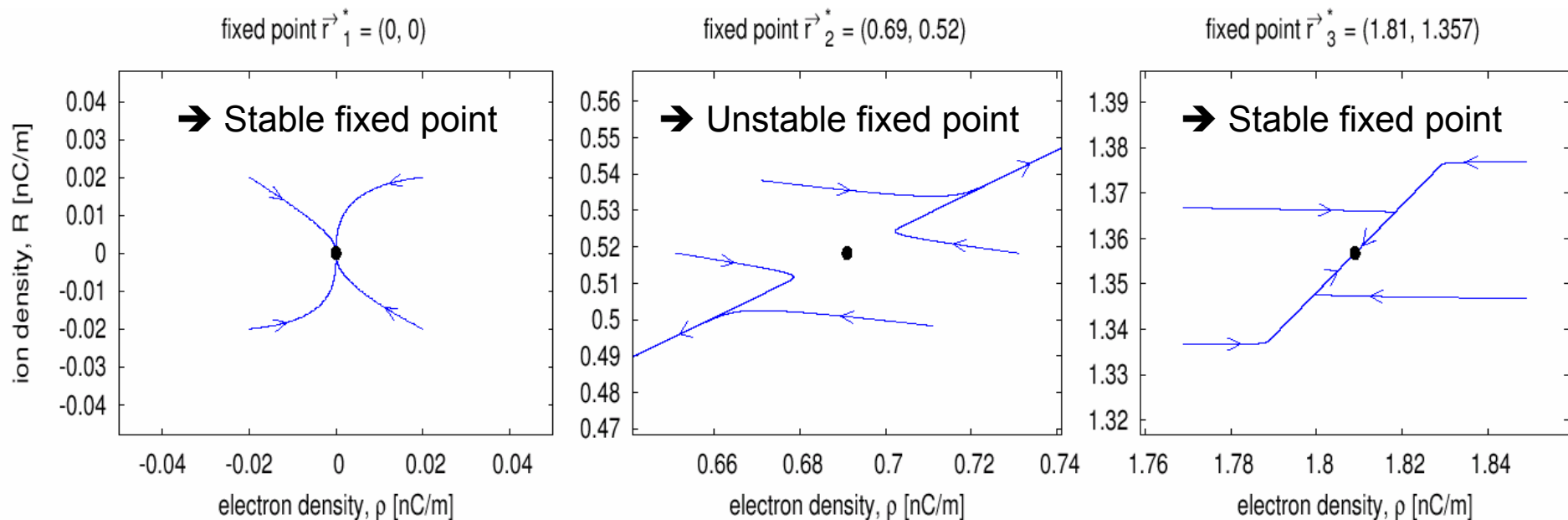
4. “Proof-of-principle” coupled maps

$$\begin{aligned} \rho_{m+1} &= (a + b\rho_m + yR_m)\rho_m + c\rho_m^3 \\ R_{m+1} &= AR_m + Y\rho_m . \end{aligned}$$

- Electrons collide with rest gas and create ions: $Y\rho_m$
- Ions enlarge a (enlarge e- survival between bunches): $b\rho_m R_m$

- Assume now $a(N)$ changes linearly and keep rest of parameters constants
- For a given bunch population N , more than one solution can be found;

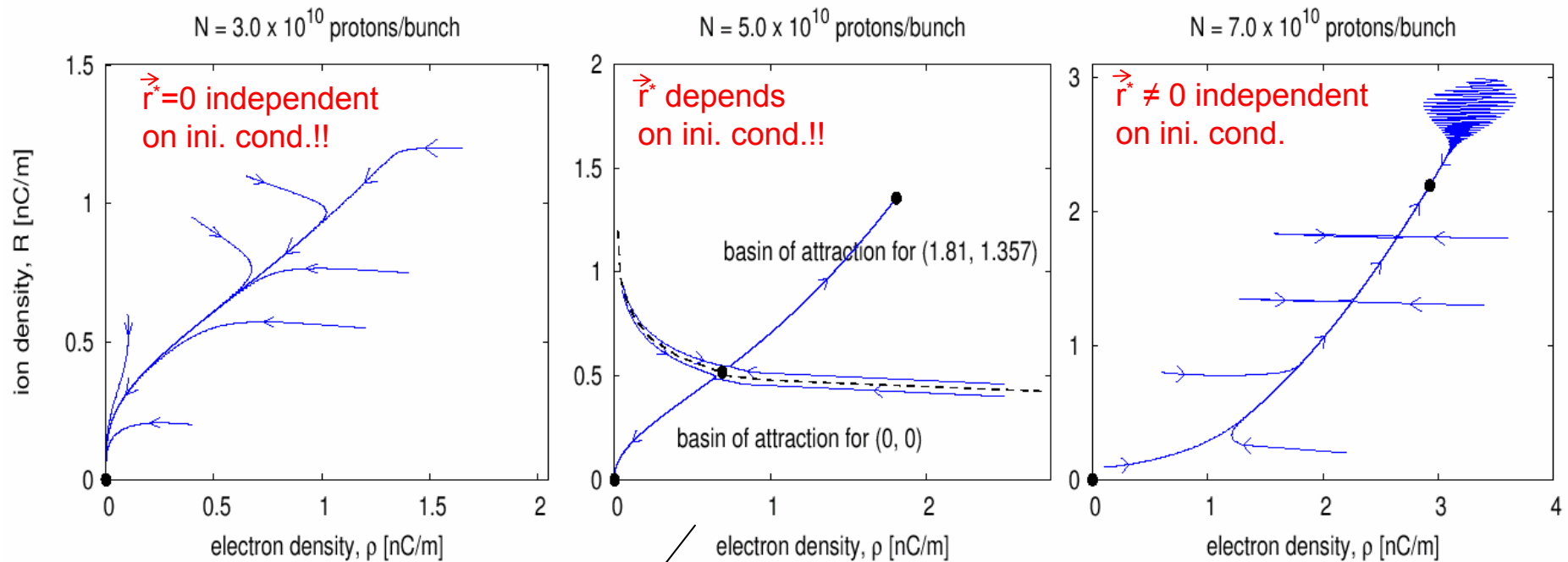
→ Example: the 3 fixed points are found for $N=5 \cdot 10^{10}$ protons/bunch*. Their stability depends on the jacobian matrix and can be checked in the (ρ, R) space:



*see values of the coefficients at U. Iriso and S. Peggs, PRST-AB, 9, 071002 (2006)

4.a. Stability and importance of the initial conditions

- \vec{r}^* depends on the initial conditions (memory effects) for different N
- (ρ, R) space behaviour differs with different bunch populations:

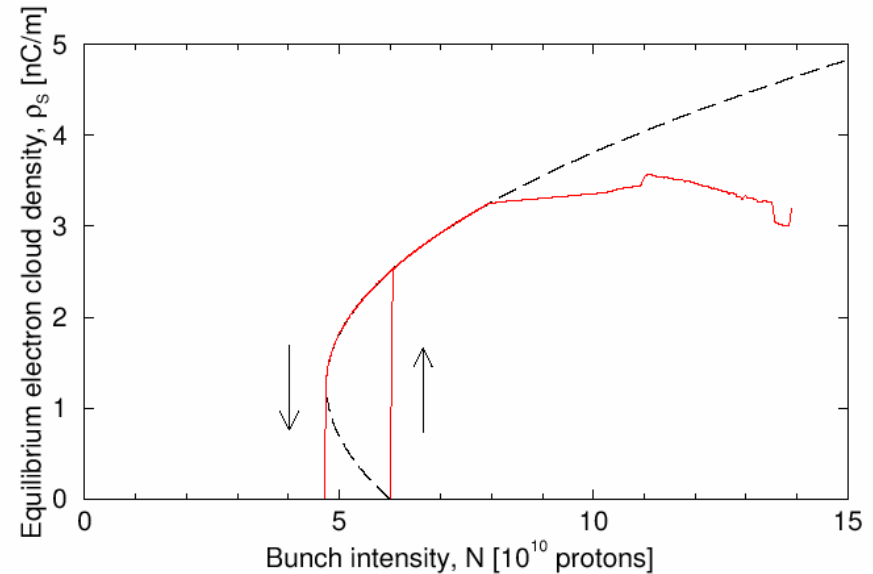


- The two basins of attraction in the (ρ, R) space produce the hysteresis

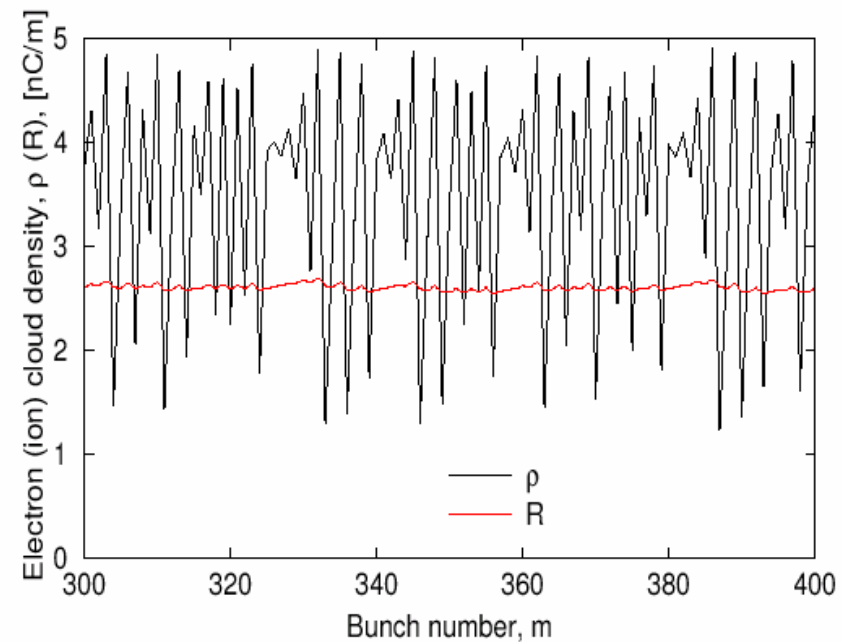
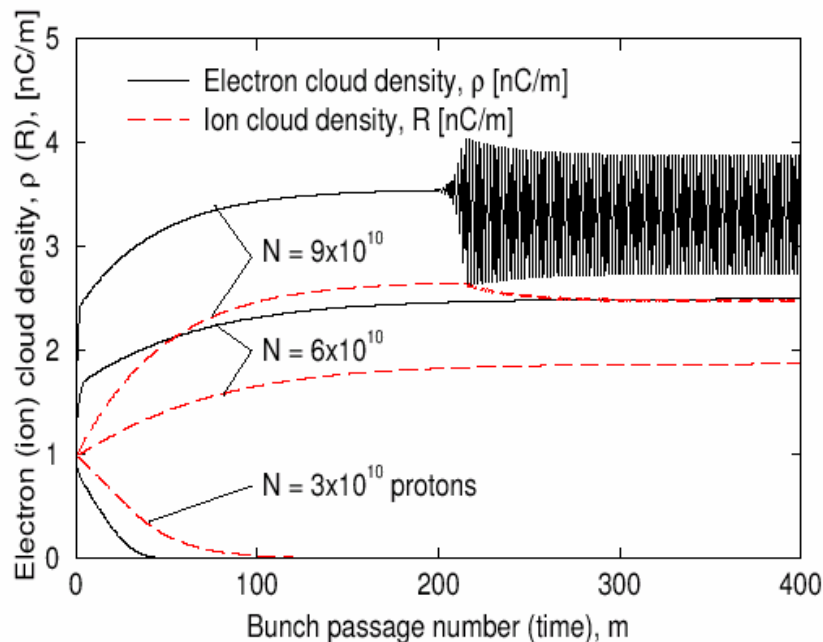
4.b. Hysteresis and additional dynamical phases

Results from a dynamical simulation based on the coupled maps first as N is slowly increased, then as N is slowly decreased

Hysteresis is observed because the final state depends on the initial conditions for some bunch intensities.



...chaotic regimes can be found...



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5. Conclusions

- Maps are a suitable tool to overcome CPU limitations presented by possible electron and ion clouds coupling.
- The development of stability conditions for a broad spectrum of potential coupling mechanisms is presented.
- Final solution \vec{r}^* for e- and ion densities can depend on the initial conditions \rightarrow hysteresis effects.
- The model reproduce the first order phase transitions seen in practice for the pressure.
- They also predict that chaotic regimes may appear near machine operating conditions.